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Quintessential inflation in Palatini F(R,X) gravity

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Palatini F(R,X) gravity, with X the inflaton kinetic term, proved to be a powerful framework for generating asymptotically flat inflaton potentials. Here we show that the general form of the F(R,X) potential also provides a tail that generates, in principle, quintessential behavior for dark energy. We describe a few examples that fit the observed value of the cosmological constant and discuss the pros and cons of those models.

Author: DIOGUARDI, Christian (NICBP)

Co-author: RACIOPPI, Antonio (National Institute of Chemical Physics and Biophysics (EE))

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