



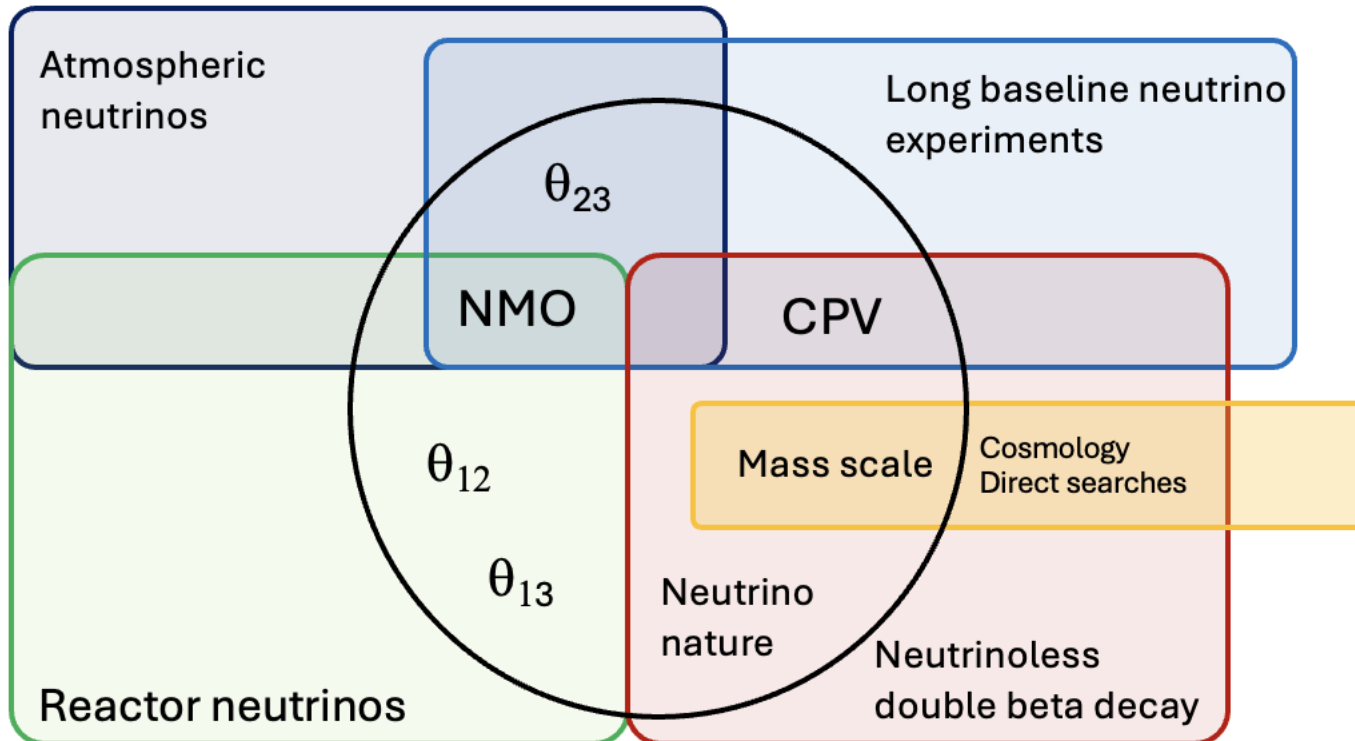
Experimental neutrino physics in Finland

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Particle Physics Days 27.-28.11.2024, Lammi



What we want to know

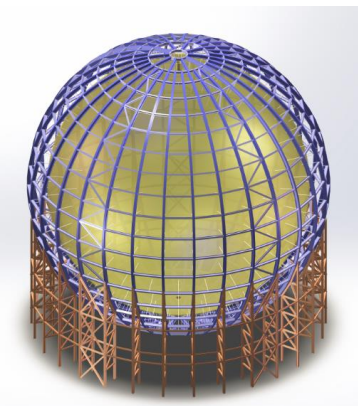


- Three mixing angles (θ_{ij})
- Neutrino mass ordering (NMO)
- CP-violation phase (CPV)
=> oscillation experiments
- Neutrino absolute mass
=> beta decays
=> constraints from cosmological measurements
- Nature: Dirac vs. Majorana
=> double beta decay

+ using neutrinos as messengers: Earth, Sun, stars, other sources...



Jianmen Underground Neutrino Observatory



Main goal: Neutrino mass ordering ($>3\sigma$)

Additional goals: solar, supernova, atmospheric, geo neutrinos, proton decay (P2: double beta)

Host: China (IHEP)

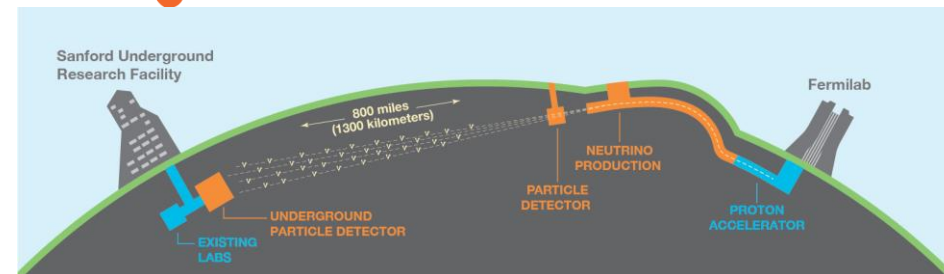
Source: 2x reactor complexes 53 km away providing high flux of electron antineutrinos with 0-10 MeV

Detectors:

- 20 kton spherical liquid scintillator, ~17k + 32k photomultipliers.
- Near detector JUNO-TAO, **Radiopurity detector OSIRIS**

Status: Central Detector under instrumentation. Filling of Central detector 15th of Decemeber 2024.

Collaboration: ~700 collaborators, 71 institutes, 17 countries (mainly China, Italy, France, Germany)



Main goal: Neutrino mass ordering (5σ) and CP violation ($>3\sigma$)

Additional: supernova, atmospheric, proton decay

Host: US, Fermilab + SURF

Source: 1.2 MW neutrino beam 1-10 GeV muon (anti)neutrinos

Detectors: Far: 2 (4) x 10 kton liquid Argon Time projection chambers. Near: multi-technique.

Status: Excavation started, starts with 2 modules, Beam upgrade(?), First data expected 2029 (2031).

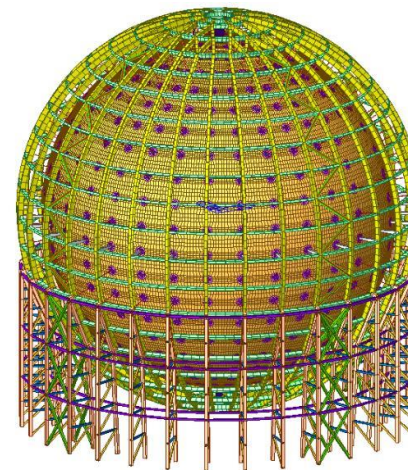
Collaboration: ~1400 collaborators, ~200 institutes, ~30 countries

Neutrino Mass Ordering with **JUNO**
Screening of JUNOs target material with **OSIRIS**

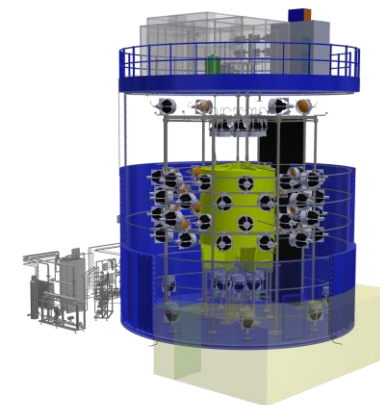
JUNO's RUN2: **neutrinoless double beta decay**
OSIRIS upgrade to R&D test bed
Slow Scintillator development and characterization

Proof-of-concept experiment for **beta spectral shapes** and
effective g_A

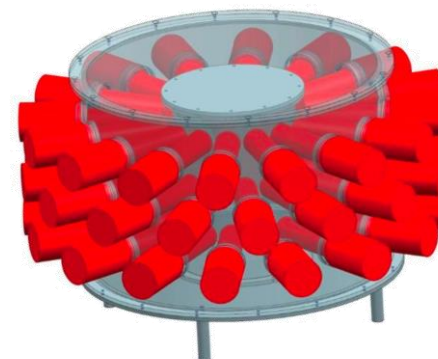
JUNO: ~20 000 tons



OSIRIS: ~20 tons



Lavatrice: 20 litres



Betashape: ~1 desilitre



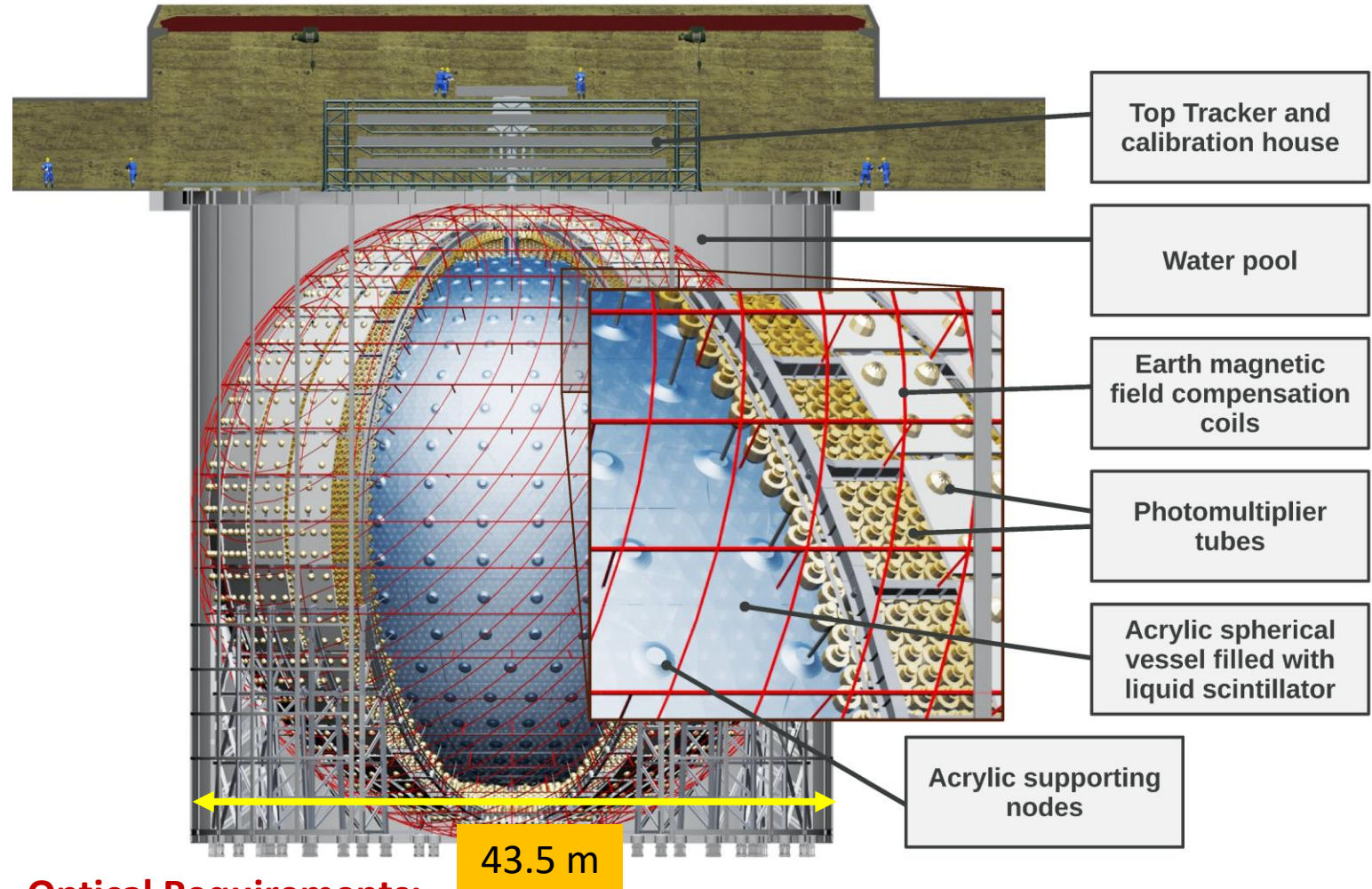
Jiangmen **U**nderground **N**eutrino **O**bservatory (**JUNO**)

Multi-purpose experiment but with a main focus:
Measurement of the Neutrino Mass Ordering using reactor anti-electron neutrinos



JUNO Detector

- Target
 - Spherical acrylic vessel
 - 20 kton of LAB-based liquid scintillator
 - High transparency
 - High radiopurity
- Light detection
 - ~17 000 20-inch PMTs
 - ~26 000 small PMTs
 - < 3% energy resolution
- 700m underground
- Biggest LSc detector in the world
- Under construction in China



Optical Requirements:

Light output: ~ 10.000 Photons / MeV $\rightarrow \sim 1200$ p.e

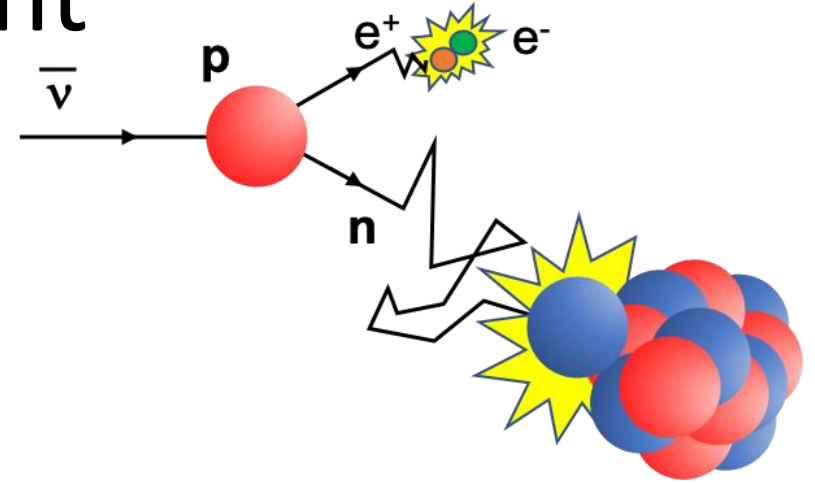
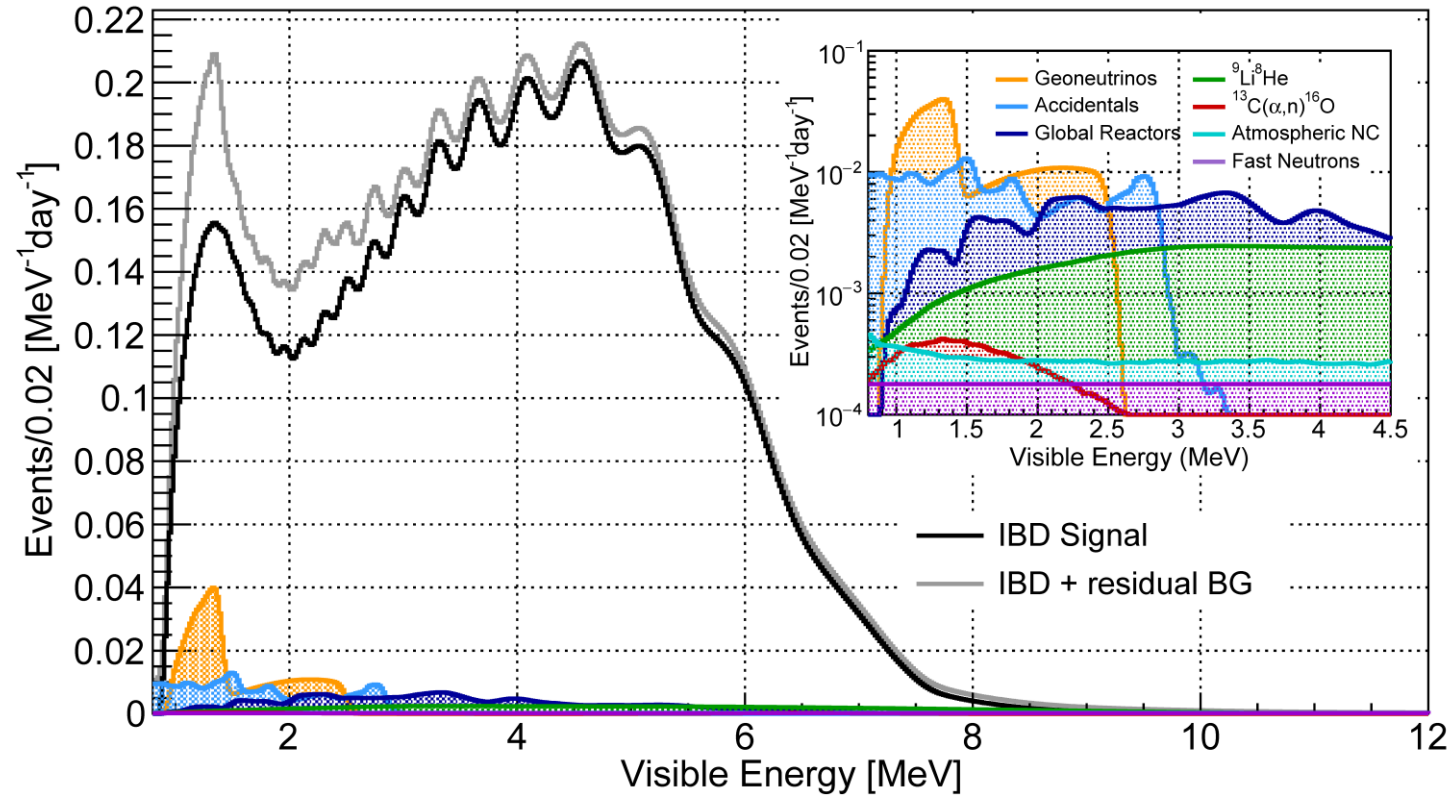
Attenuation length: > 20 m @ 430 nm

Required Radiopurity:

Reactor neutrinos: $^{238}\text{U} / ^{232}\text{Th} < 10^{-15}$ g/g, $^{40}\text{K} < 10^{-16}$ g/g

Solar neutrinos: $^{238}\text{U} / ^{232}\text{Th} < 10^{-17}$ g/g, $^{40}\text{K} < 10^{-18}$ g/g, $^{14}\text{C} < 10^{-18}$ g/g⁶

Reactor neutrino measurement



	Rate [/day]	JUNO	TAO
Reactor IBD signal		47	2000
Geo- ν 's		1.2	-
Accidental signals		0.8	155
Fast-n		0.1	92
${}^9\text{Li}/{}^8\text{He}$		0.8	54
${}^{13}\text{C}(\alpha, n){}^{16}\text{O}$		0.05	-
Global reactors		1.0	-
Atmospheric ν 's		0.16	- 7

JUNO Collaboration:

Sub-percent precision measurement of neutrino oscillation parameters with JUNO *Chin.Phys.C* 46 (2022) 12, 123001



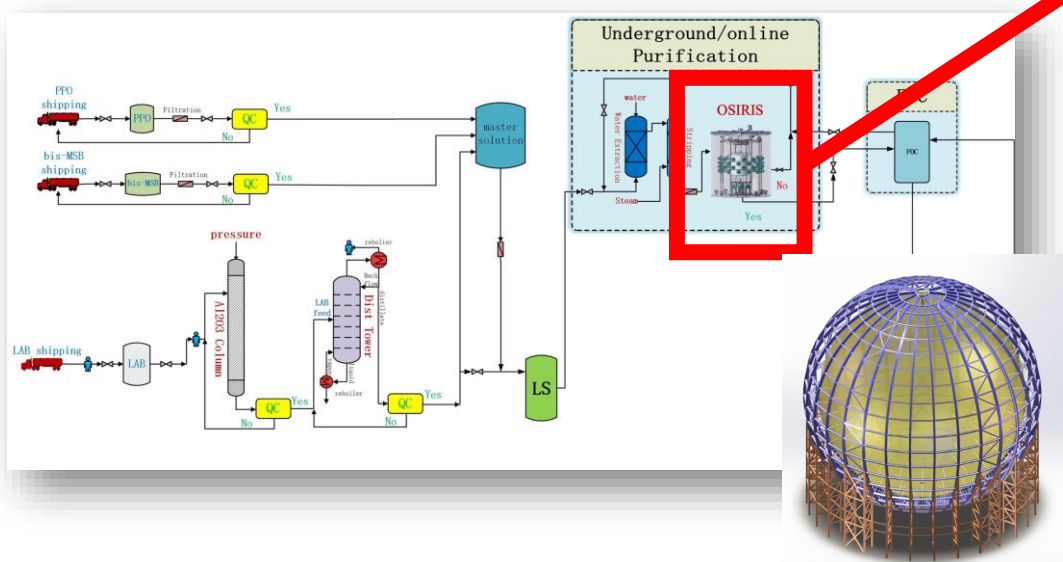
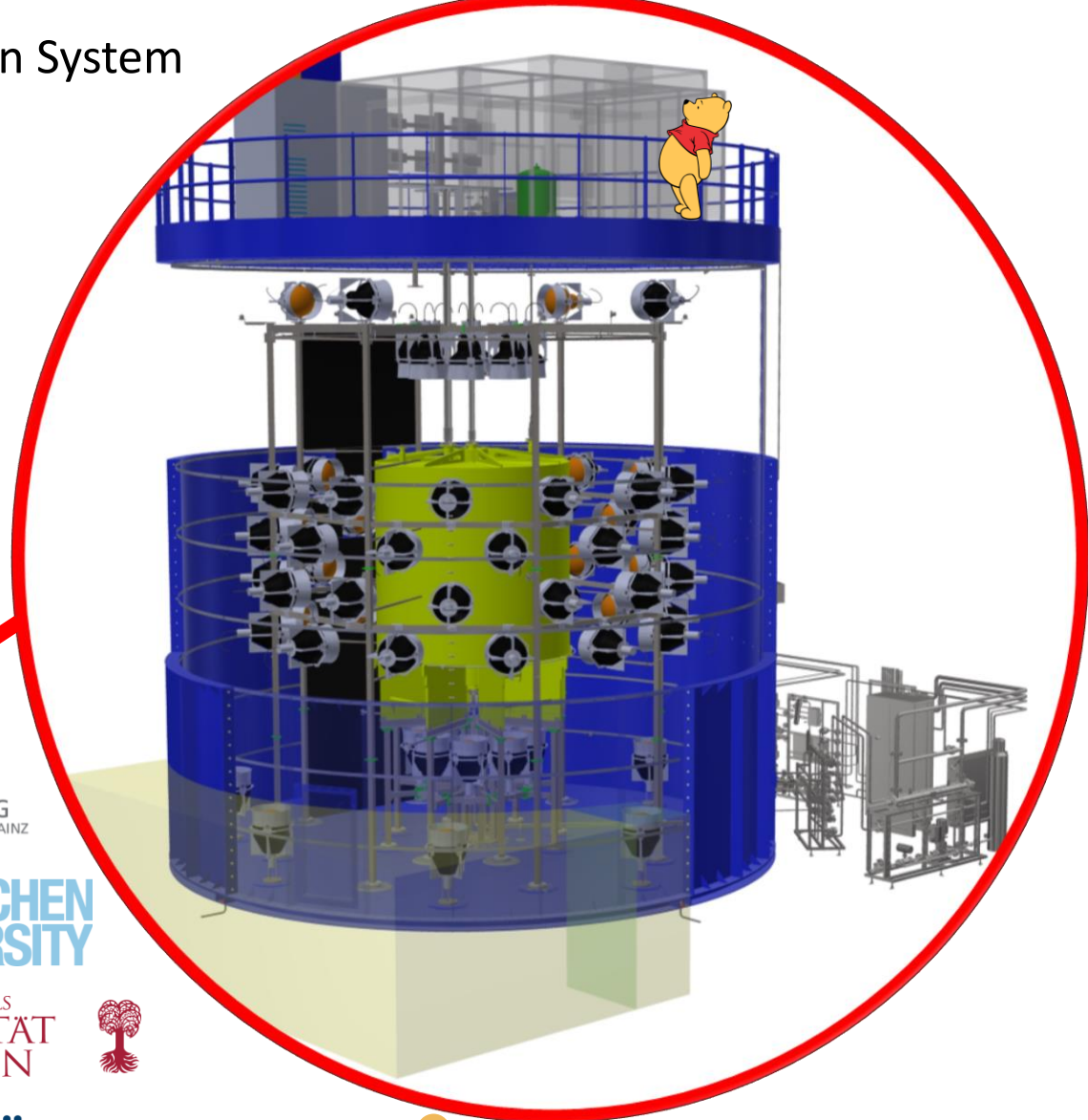
Online Scintillator Internal Radioactivity Investigation System

OSIRIS

The **OSIRIS** detector is a subsystem of the liquid scintillator filling chain of the **JUNO neutrino experiment** and its purpose is

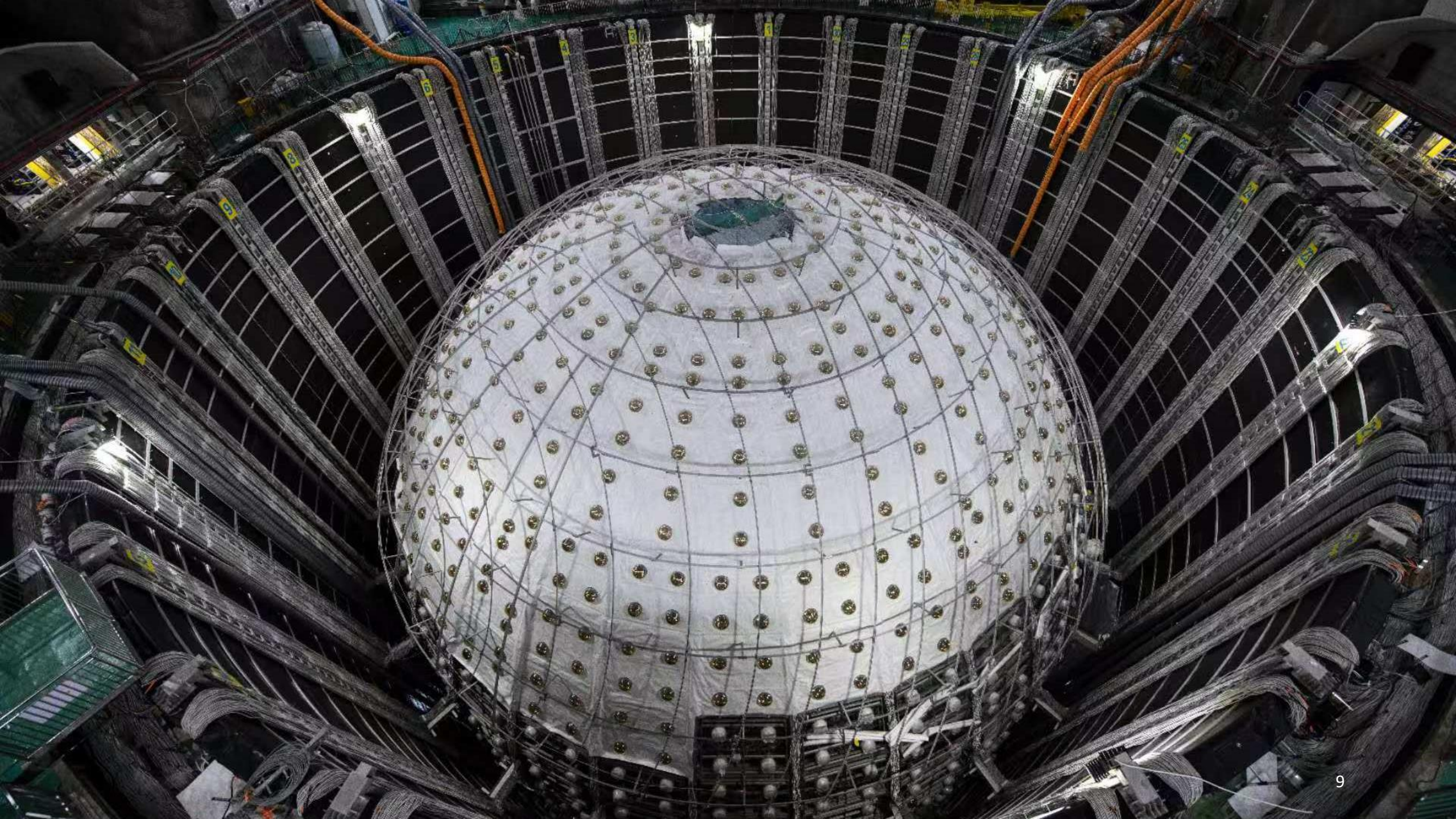
- ✓ **To validate the radiopurity of the scintillator**
- ✓ **To assure that all components of the JUNO scintillator system work to specifications and**
- ✓ **To verify that only neutrino-grade scintillator is filled into the JUNO Central Detector.**

The aspired radiopurity level of 10^{-15} to 10^{-17} g/g (reactor to solar) of ^{238}U and ^{232}Th requires a large ($\sim 20 \text{ m}^3$) detection volume and ultralow background levels.



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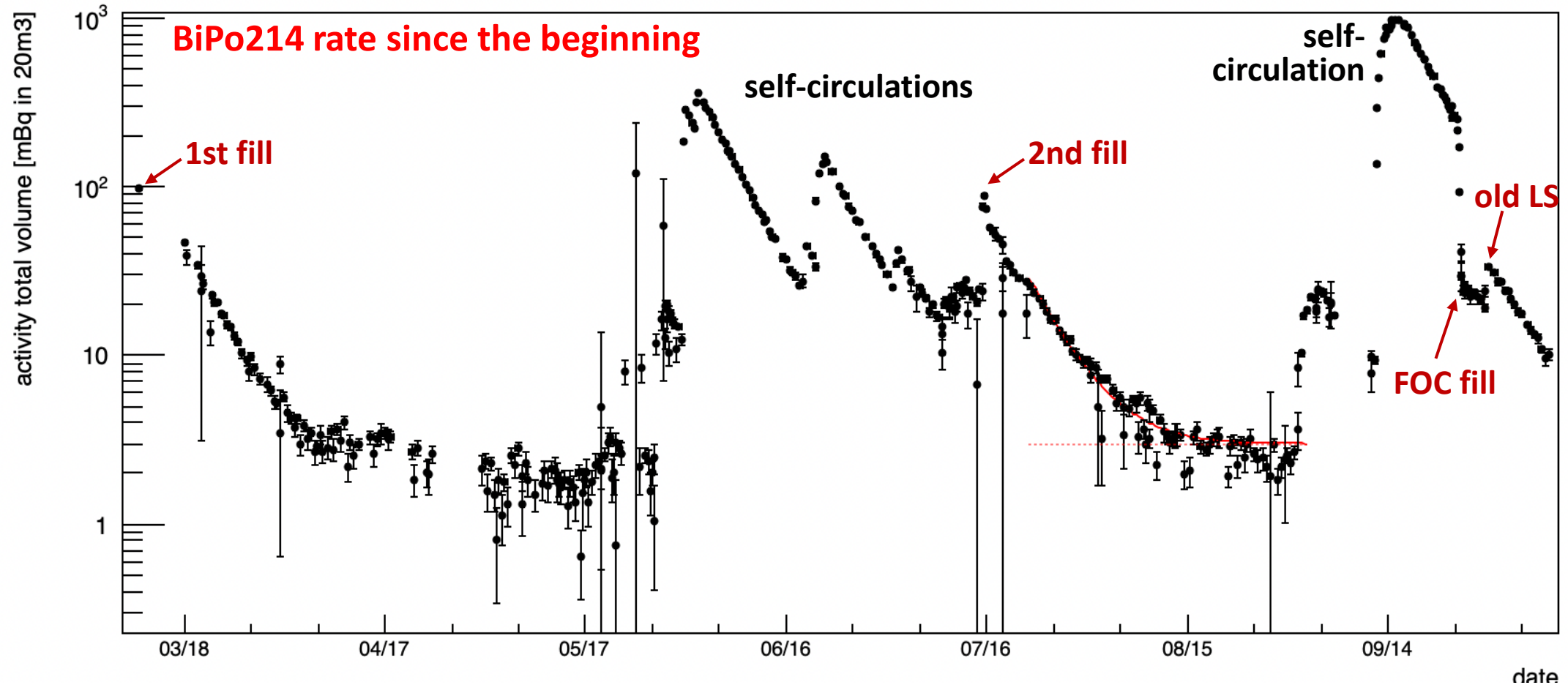




May '22



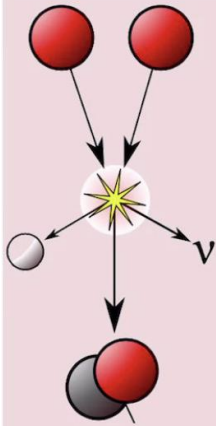
Data taking journey so far with OSIRIS



FOC filling was the lowest in **initial Rn-222** level that we have achieved so far, ~ 25 mBq in total volume

OSIRIS: After JUNO filling

After JUNO filling, purification infrastructure including OSIRIS will become available

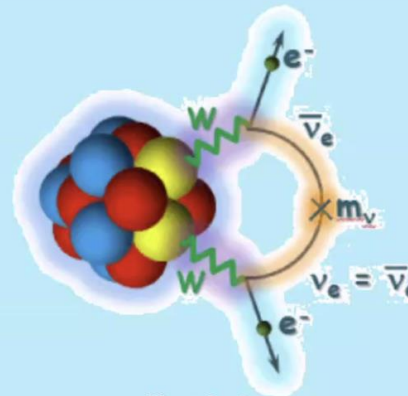


Precision measurement of solar pp neutrinos

- improve Borexino measurement (10%) to 3% level or better
→ θ_{12} for ν_e , for 1% accuracy: hadronic axions, hidden photons
- OSIRIS Upgrade could achieve
 - low ^{14}C scintillator
 - better energy resolution than OSIRIS (and maybe JUNO)
 - slow scintillator for integrated directionality (CID)

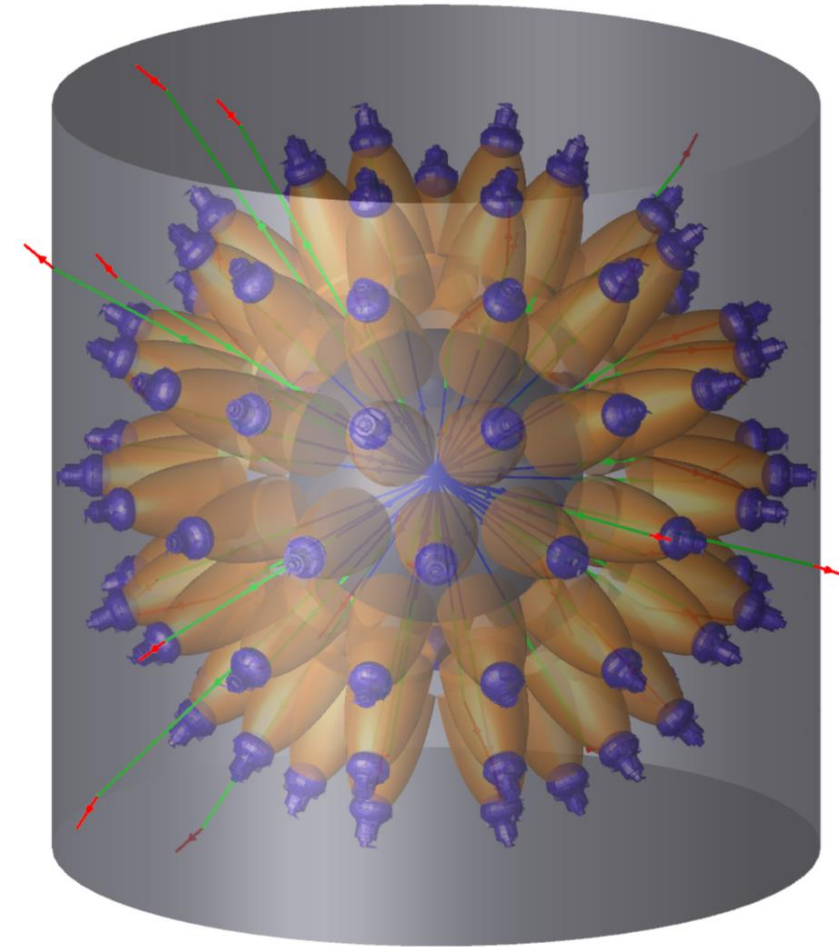
Demonstrator for neutrino-less double-beta decay ($0\nu 2\beta$)

- test Te/Xe-loaded scintillators for JUNO- $\beta\beta$ phase
- develop new background discrimination techniques essential for JUNO sensitivity, e.g. **particle ID based on Cherenkov/scintillation ratio**
- **very sensitive $2\beta^+$ decay measurements (10^{24-25} yrs)** with $^{78}\text{Kr}/^{124}\text{Xe}$ -loaded slow scintillators



Foreseen upgrade:

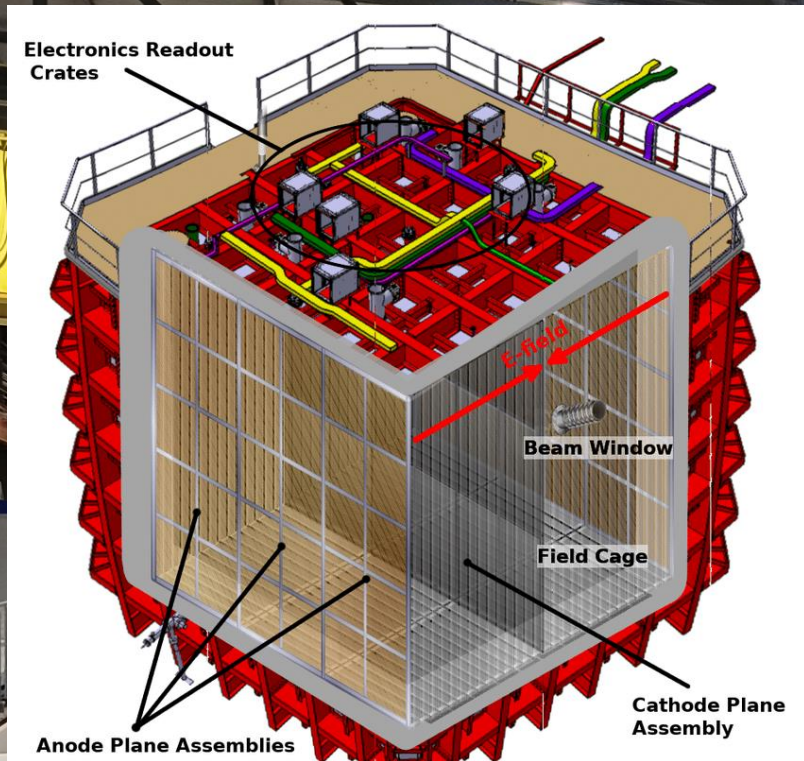
- Rearranging PMTs
- Attaching light collector cones
- Adding external shielding against γ 's



Long baseline neutrinos



- Represents an effort of CERN to foster fundamental research in the field of Neutrino Accelerator Physics as decided by CERN Council in framework of the 2013 European Strategy.
- Experimental Hall North 1 (EHN1) at Prevezin site: cryogenics, magnets, beamlines H2, H4 protons SPS 10-400 GeV/c, secondary mixed hadrons
- Focus in detector development:
 - DUNE prototyping: ProtoDUNE-SinglePhase (NP04), **ProtoDune-DualPhase (NP02)**
 - T2K near detector (NP07)
 - Prototyping for ENUBET and NuTag collaborations
- Filling and beam to protoDUNEs 2024-2025
- JYU contribution declined since the beginning of 2019. Reactivating.



Other closely related activities

- OSIRIS Upgrade planning
 - Standalone test bed for double beta decays (JUNO-II)
- LS technology development
 - Characterisation of slow LS, Cerenkov/Scintillation separation.
 - Improved particle identification, directionality, background discrimination
 - Results included to OSIRIS-upgrade sensitivity studies and THEIA@DUNE
- Measurement of beta decay spectral shapes
 - Important for all low energy rare-event search (dark Matter, double beta, neutrino) in understanding the radioactive background.

Summary

- Within Finland the current main activity is JUNO. Operational within one year. Main contribution through OSIRIS radiopurity screening
- ProtoDUNE activities at CERN Neutrino Platform has been on hold. Definitely room for improvement. (C/S)hould be reactivated.
- Other indirect nuclear physics and R&D activities: Liquid scintillator R&D, background measurements.
- + theoretical studies: neutrino phenomenology, nuclear physics.

Backups

Schedule

V.6.6

V.6.4 established in July 2024

		Start	End	Condition
1	Underground lab construction	2015.1.1	2021.11.25	
2	Water pool cleaning and CD construction preparation	2021.11.26	2021.12.10	1
3	Acrylic construction/bottom acrylic chimney	2021.12.11	2024.9.30	2
4	CD bottom 4 layers steel structure	2024.10.1	2024.10.15	3
5	CD & VETO PMT installation	2022.10.1	2024.11.15	4
6	CD dust settlement, water washing and film removal	2024.10.1	2024.10.15	3,7
7	TT bridge installation	2024.8.1	2024.9.30	
8	CD top chimney installation	2024.10.16	2024.10.20	6,7
9	pole PMT installation/Calib. House (sealed with chimney)	2024.10.16	2024.11.15	7,8
10	pool cleaning, bottom instrumentation, concrete door construction, cover installation	2024.10.16	2024.11.30	3,4,9
11	VETO & CD water filling	2024.12.1	2025.1.31	10
12	LS filling/water exchange	2025.2.1	2025.7.31	11
13	TT module installation/commissioning	2024.12.1	2025.7.30	10
14	TAO installation	2024.8.1	2025.7.30	
15	Test run	2025.8.1		

← Filling of the big detector starts in December

We plan to start the water filling on Dec. 15

From neutrino mixing to oscillation

Flavor states

PMNS matrix

Mass states

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = \begin{bmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{bmatrix} \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix}$$

$$\begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

$$= \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \begin{bmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

Notation: $s_{12} = \sin(\theta_{12})$, $c_{12} = \cos(\theta_{12})$

Probability to observe ν_α as ν_β after distance L

$$P_{\nu_\alpha \rightarrow \nu_\beta}(t) = |\langle \nu_\beta | \nu_\alpha(t) \rangle|^2$$

$$= \delta_{\alpha\beta} - 4 \sum_{i>j} \Re(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin^2\left(\frac{\Delta m_{ij}^2 L}{4E}\right)$$

$$+ 2 \sum_{i>j} \Im(U_{\alpha i}^* U_{\beta i} U_{\alpha j} U_{\beta j}^*) \sin\left(\frac{\Delta m_{ij}^2 L}{4E}\right),$$

where $\Delta m_{ij}^2 = m_i^2 - m_j^2$

Parameters

- Three mixing angles
- Two mass-squared splittings
- CP-violation phase δ
- (Two Majorana phases α_i)

The Neutrino Mass Ordering with Reactor ν 's

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta m_{21}^2 \frac{L}{4E} - \sin^2 2\theta_{13} \left(\cos^2 \theta_{12} \sin^2 \Delta m_{31}^2 \frac{L}{4E} + \sin^2 \theta_{12} \sin^2 \Delta m_{32}^2 \frac{L}{4E} \right)$$

$$\approx 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta m_{21}^2 \frac{L}{4E} - \sin^2 \theta_{13} \sin^2 \Delta m_{ee}^2 \frac{L}{4E} \quad \text{for } \Delta m_{12}^2 \ll \Delta m_{32}^2$$

Δm_{ee}^2
effective ν -mass-squared difference
(beat frequency)

$$\Delta m_{31}^2 = \Delta m_{32}^2 + \Delta m_{21}^2$$

NO: $|\Delta m_{31}^2| = |\Delta m_{32}^2| + |\Delta m_{21}^2|$

IO: $|\Delta m_{31}^2| = |\Delta m_{32}^2| - |\Delta m_{21}^2|$

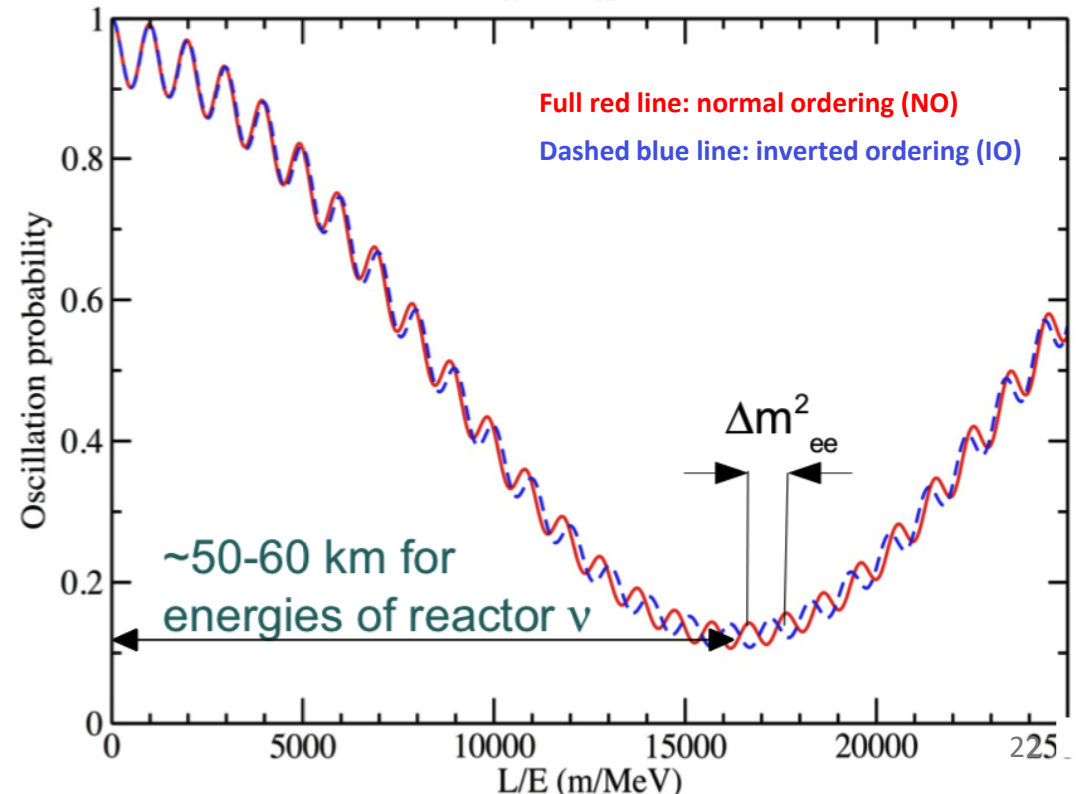
With:

$$\Delta m_{12}^2 \ll \Delta m_{32}^2$$

Different beat frequency Δm_{ee}^2 for both orderings!

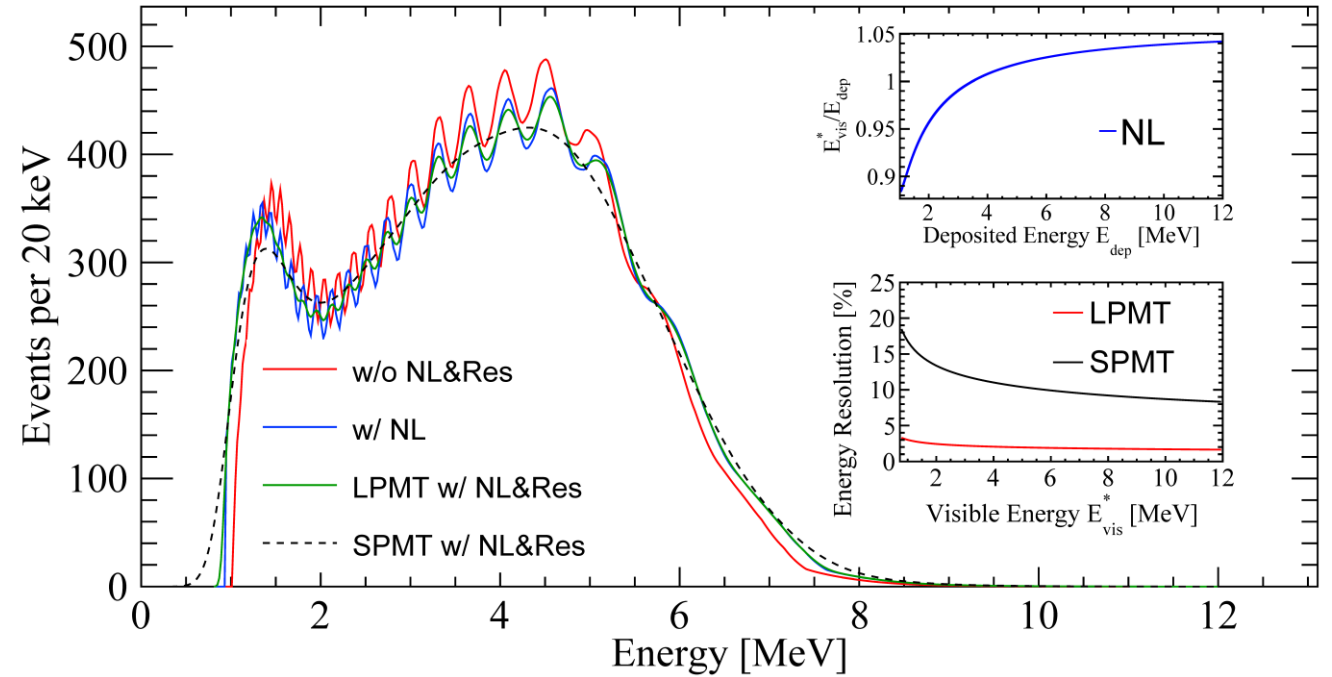
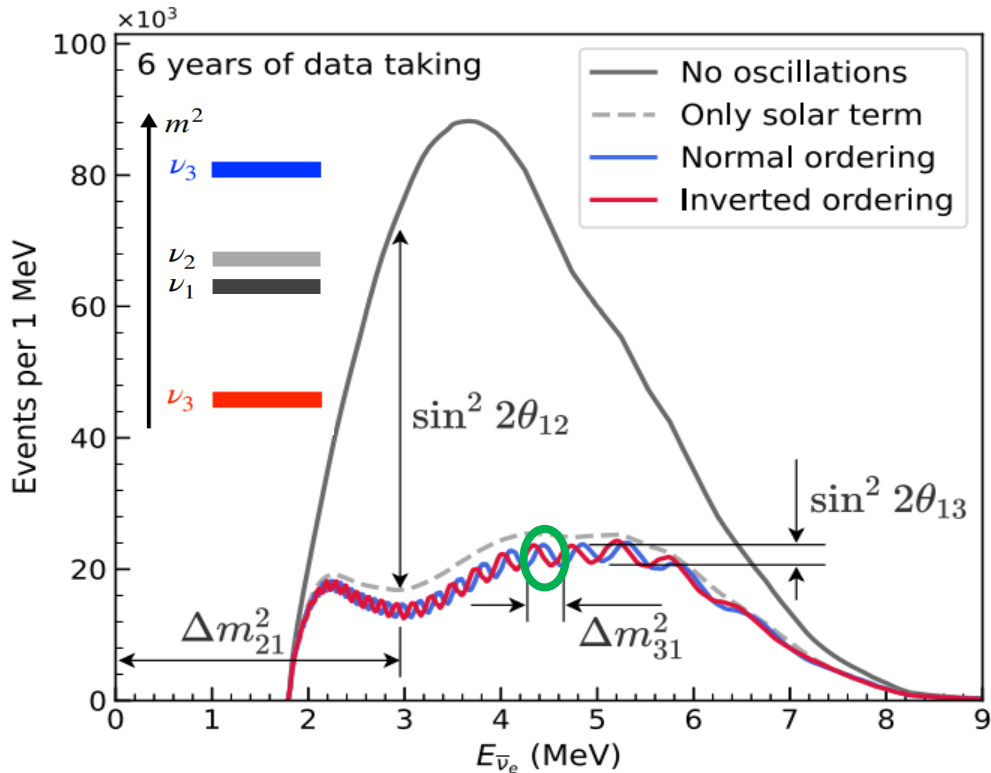
Vacuum oscillation probability $P(\nu_e \rightarrow \nu_e)$

Here for $\Delta m_{31}^2 + \Delta m_{32}^2 = 2 \times 2.49 \times 10^{-3} \text{ eV}^2$



Precision measurement of mixing parameters

JUNO 6 years data taking



Precision of $\sin^2\theta_{12}$, Δm_{21}^2 , $|\Delta m_{31}^2|/|\Delta m_{32}^2| < 0.5\%$ in 6 yrs

	Central Value	PDG2020	100 days	6 years	20 years
Δm_{31}^2 ($\times 10^{-3}$ eV ²)	2.5283	± 0.034 (1.3%)	± 0.021 (0.8%)	± 0.0047 (0.2%)	± 0.0029 (0.1%)
Δm_{21}^2 ($\times 10^{-5}$ eV ²)	7.53	± 0.18 (2.4%)	± 0.074 (1.0%)	± 0.024 (0.3%)	± 0.017 (0.2%)
$\sin^2 \theta_{12}$	0.307	± 0.013 (4.2%)	± 0.0058 (1.9%)	± 0.0016 (0.5%)	± 0.0010 (0.3%)
$\sin^2 \theta_{13}$	0.0218	± 0.0007 (3.2%)	± 0.010 (47.9%)	± 0.0026 (12.1%)	± 0.0016 (7.2%)

Sensitivity for NMO determination

