



Probing Physics Beyond the Standard Model

also in Finland?

Wladyslaw H. Trzaska



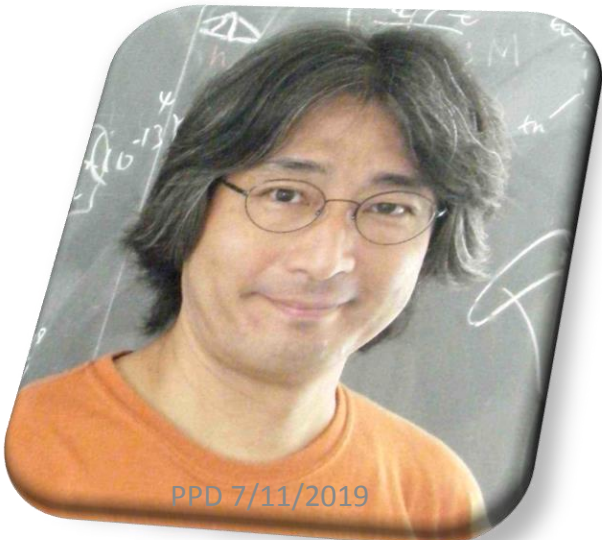
UNIVERSITY OF JYVÄSKYLÄ



- 540 participants
- 2 Nobel laureates



Plenary talk by
Hitoshi Murayama
at the opening session
of TAUP 2019



PPD 7/11/2019

Five empirical evidences for physics beyond SM

- *at least five missing pieces in the SM:*
 - *dark matter*
 - *neutrino mass*
 - $0\nu\beta\beta$*
 - gravitational wave*
 - *dark energy*
 - galaxy surveys*
 - DESI, PFS, Euclid, LSST, WFIRST*
 - *apparently acausal density fluctuations*
 - Simons Array*
 - CMB S4*
 - LiteBIRD*
 - *baryon asymmetry*
 - HyperK, DUNE*
 - LHCb, Belle II*



Implicit conclusions → Get more data; especially interesting would be extragalactic neutrinos

- There is a broad consensus that we are at the threshold of a major breakthrough discovery in physics → keep looking!
- We need more data → the first key steps towards the solution are bound to come from experiments!
- Since neutrinos are the only known particles disobeying the predictions of the SM → focus on neutrinos!
- Significance of cosmological aspects → importance of extragalactic messengers

Global response to the BSM challenge

Just a few examples



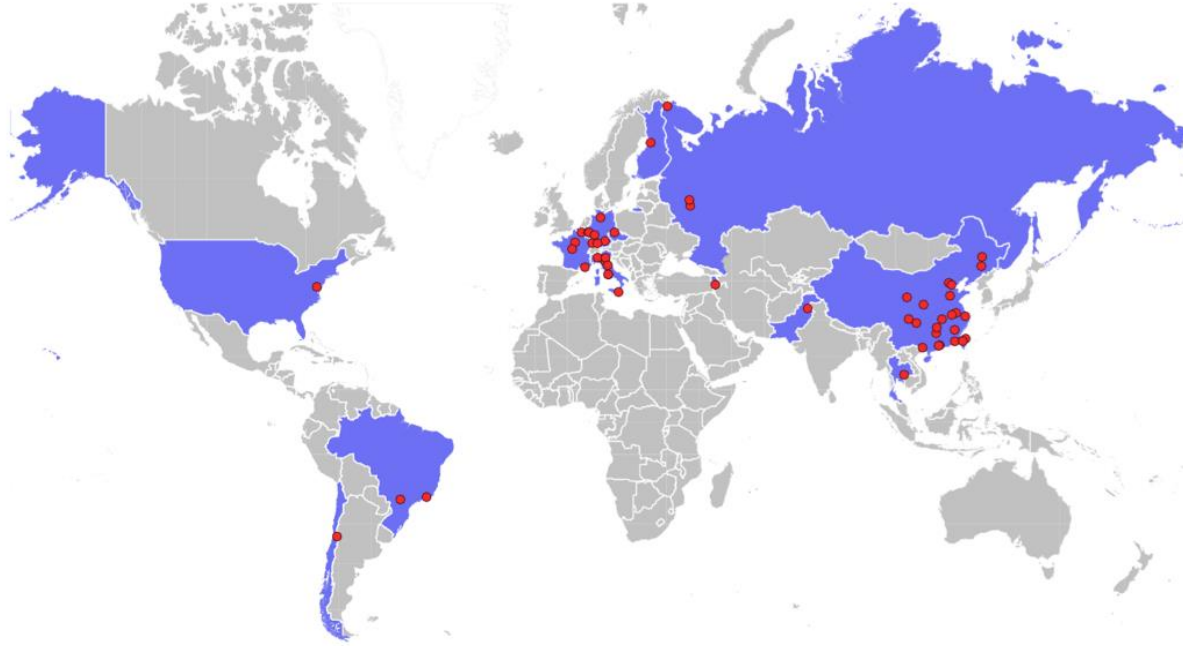


IceCube

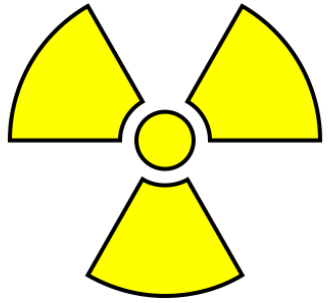


Jiangmen Underground Neutrino Observatory

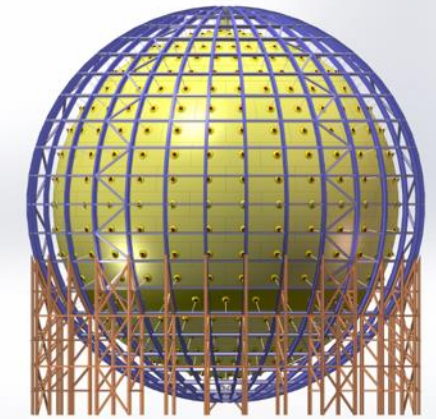
Multipurpose, medium-baseline reactor neutrino experiment



36 GW_{th}

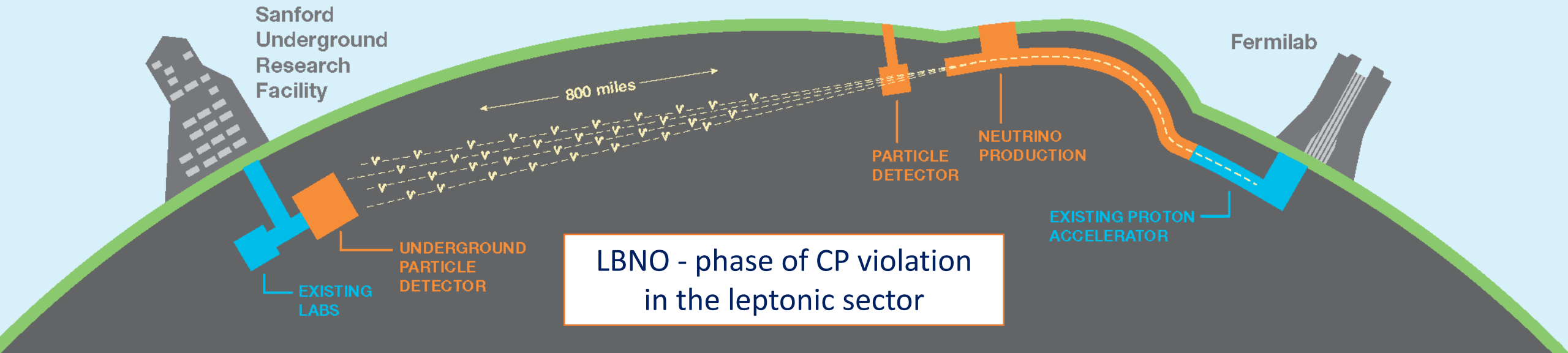


20 kton LS



~550 participants from 72 institutions in 17 countries

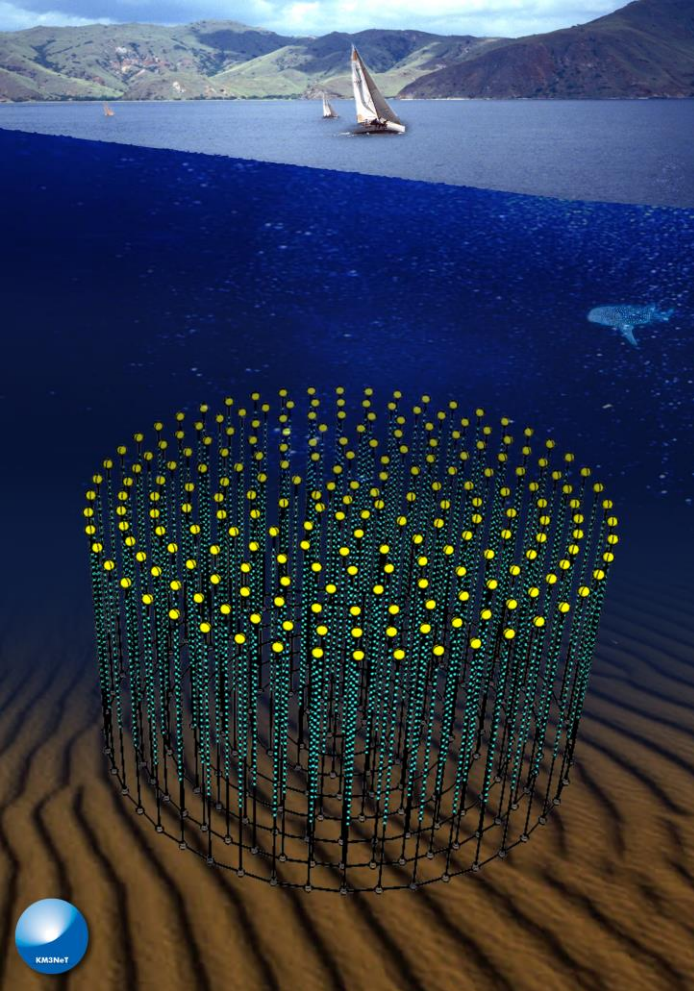




DUNE Collaboration

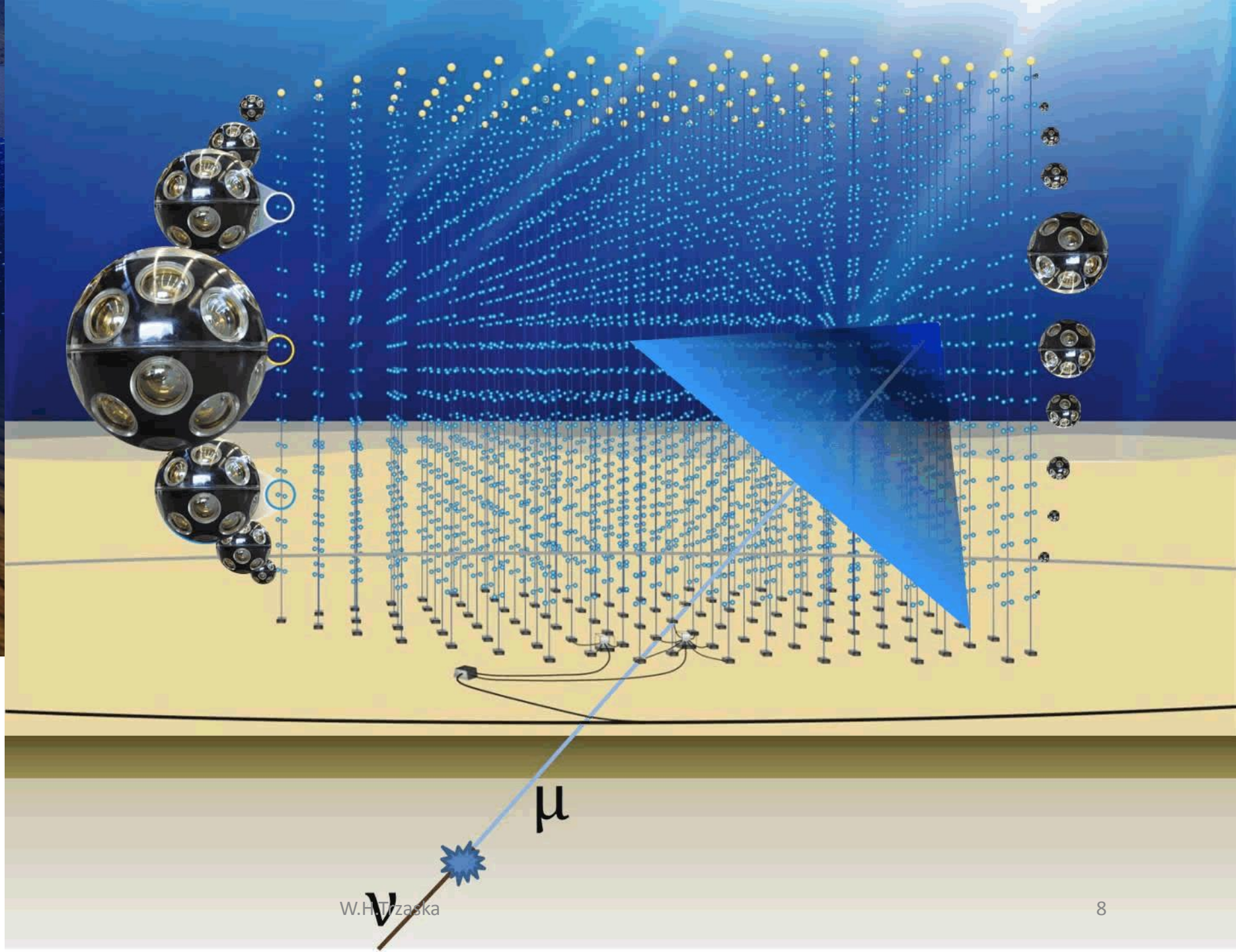
~1000 members, 184 inst. 32 countries + CERN





KM3NeT




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What could be done in Finland now?

1. No financial resources → a novel idea is needed 
2. No new constructions → use the available infrastructure 
3. No new positions → rely on the network of experts 

Background for the idea



- Ultra high-energy neutrinos are extragalactic & rare
 - giant target mass + long measurements
- Size of water and ice-based experiments is limited $\sim \text{km}^3$
- The density of water and ice is only 1 g/cm^3
- IceCube optical sensors are pointing downwards
 - not optimized for top-to-bottom events
- EeV neutrinos are not expected to traverse Earth

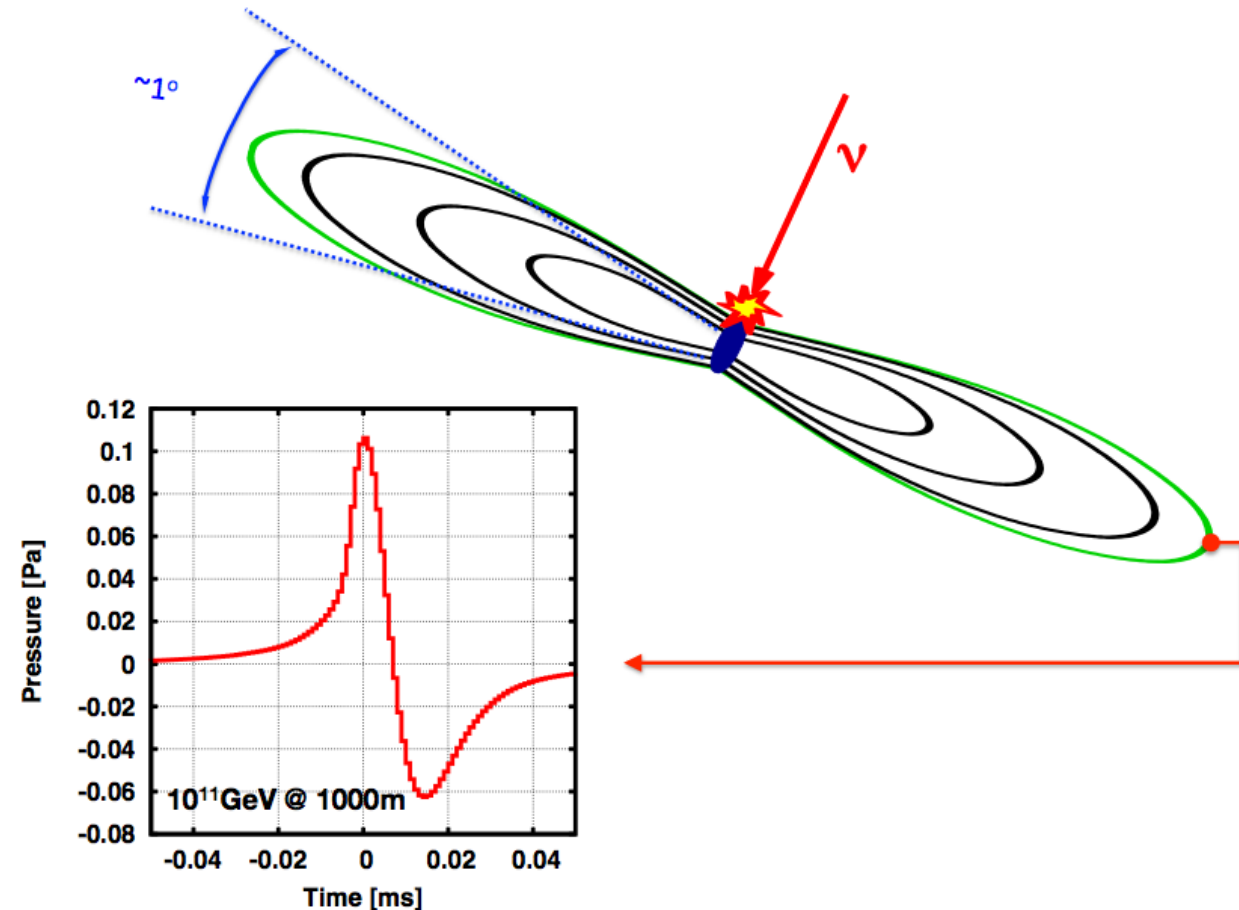


1. The idea: Acoustic detection of EeV neutrinos in bedrock



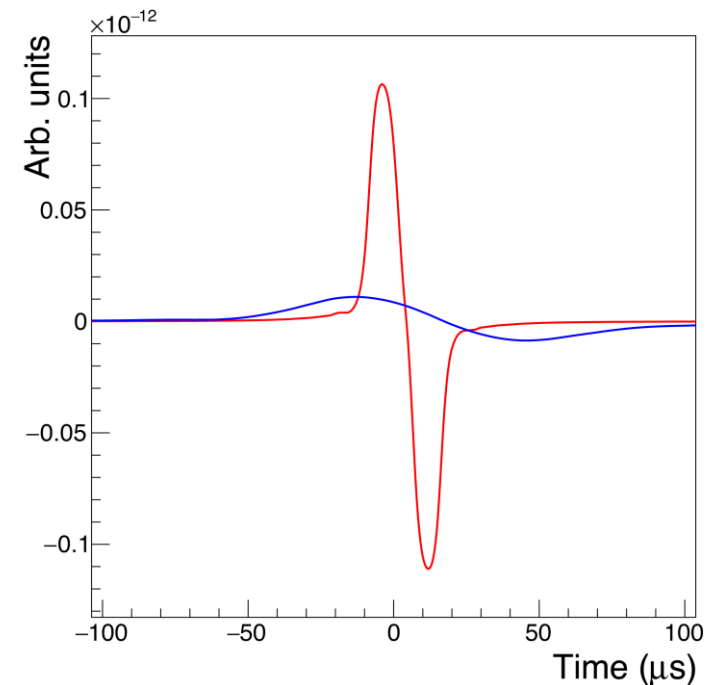
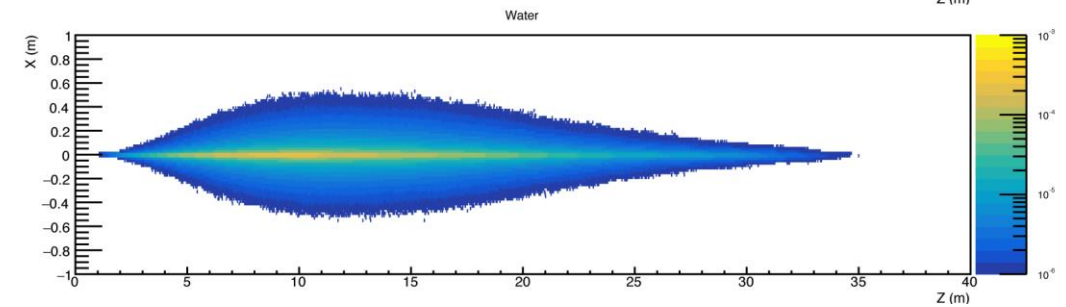
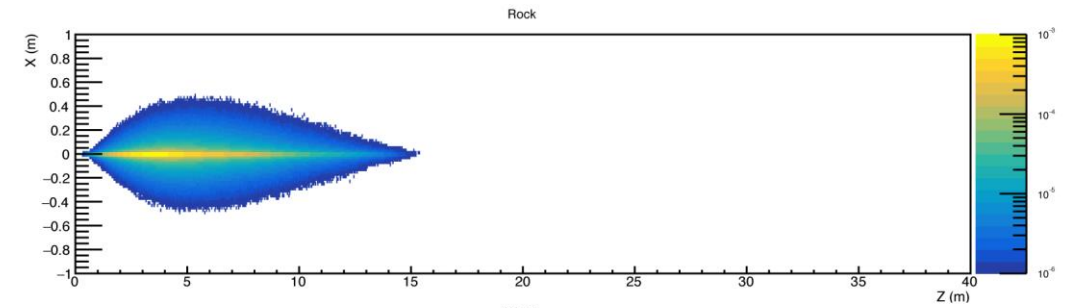
Oral presentation and peer-reviewed proceedings:
EPJ Web of Conferences 216, 04009 (2019)
<https://arxiv.org/abs/1909.00417>

Acoustic and Radio EeV Neutrino detection Activities

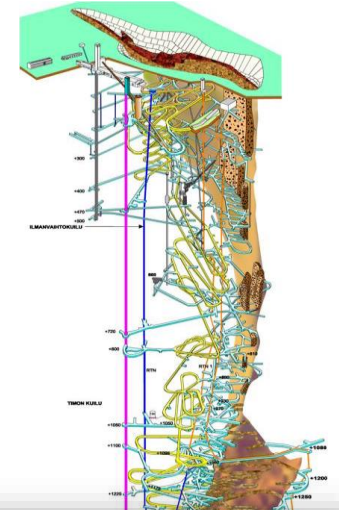
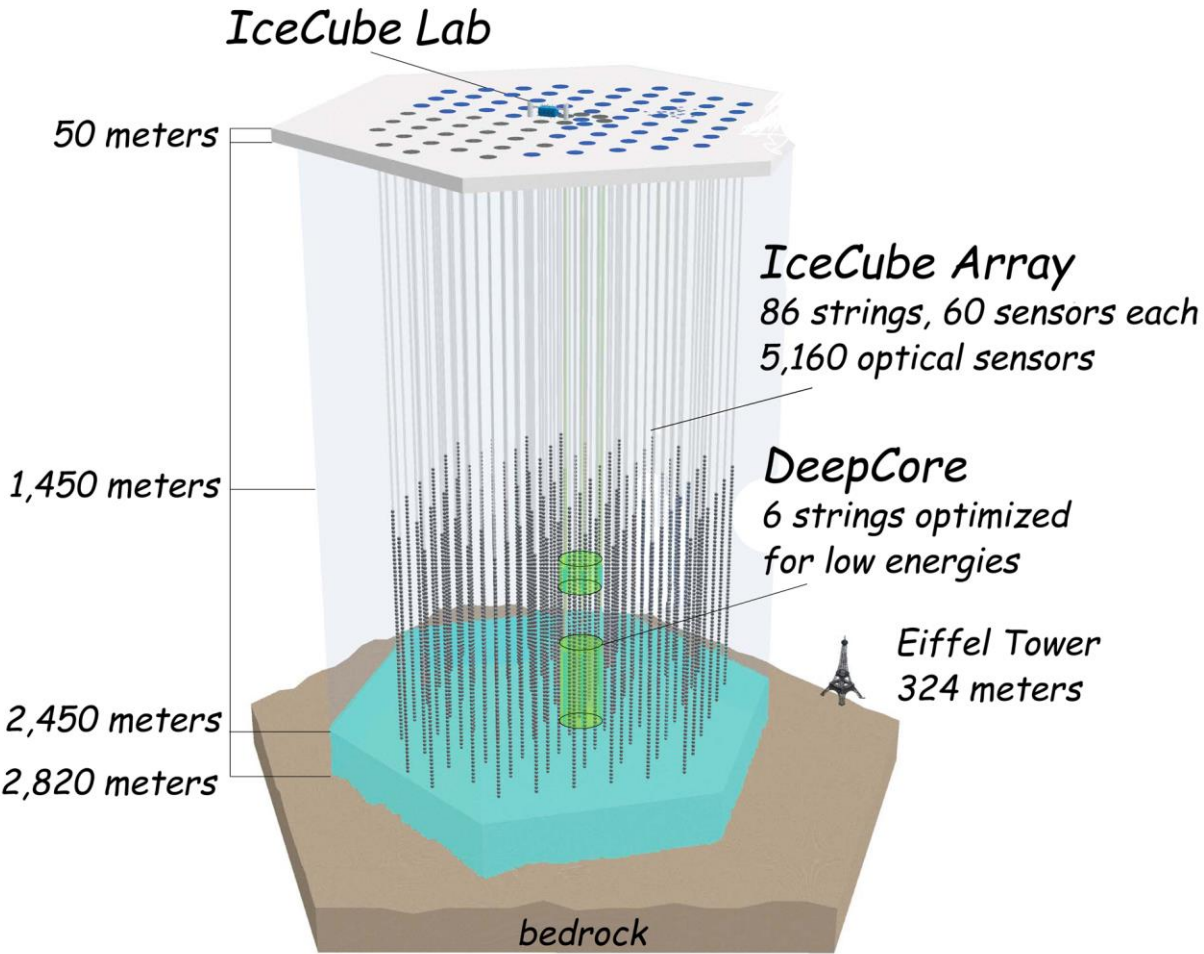


Rock vs. water

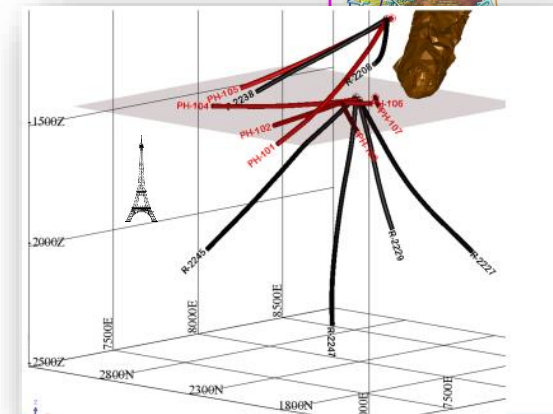
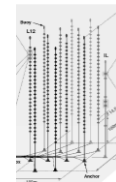
- 3x larger density
- → increased target mass
- → higher event rate for neutrinos
- 4x larger speed of sound
- 10x larger pressure pulse
- Longer attenuation length
- Logistics (mine vs. ocean)



2. Infrastructure

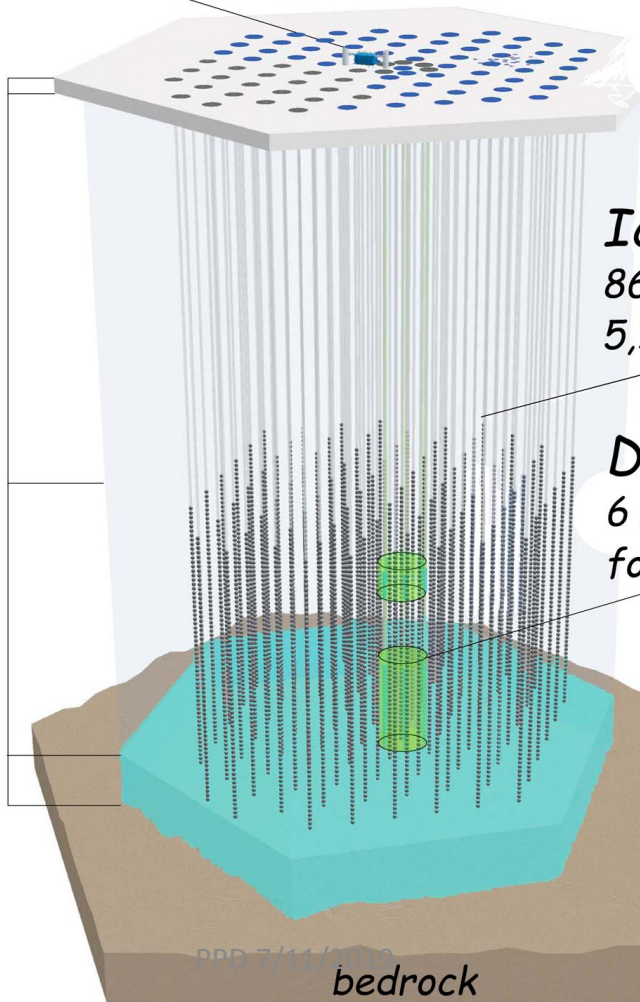


AMADEUS - The Acoustic Neutrino Detection Test System of the ANTARES Deep-Sea Neutrino Telescope



Size corrected
for density

IceCube Lab

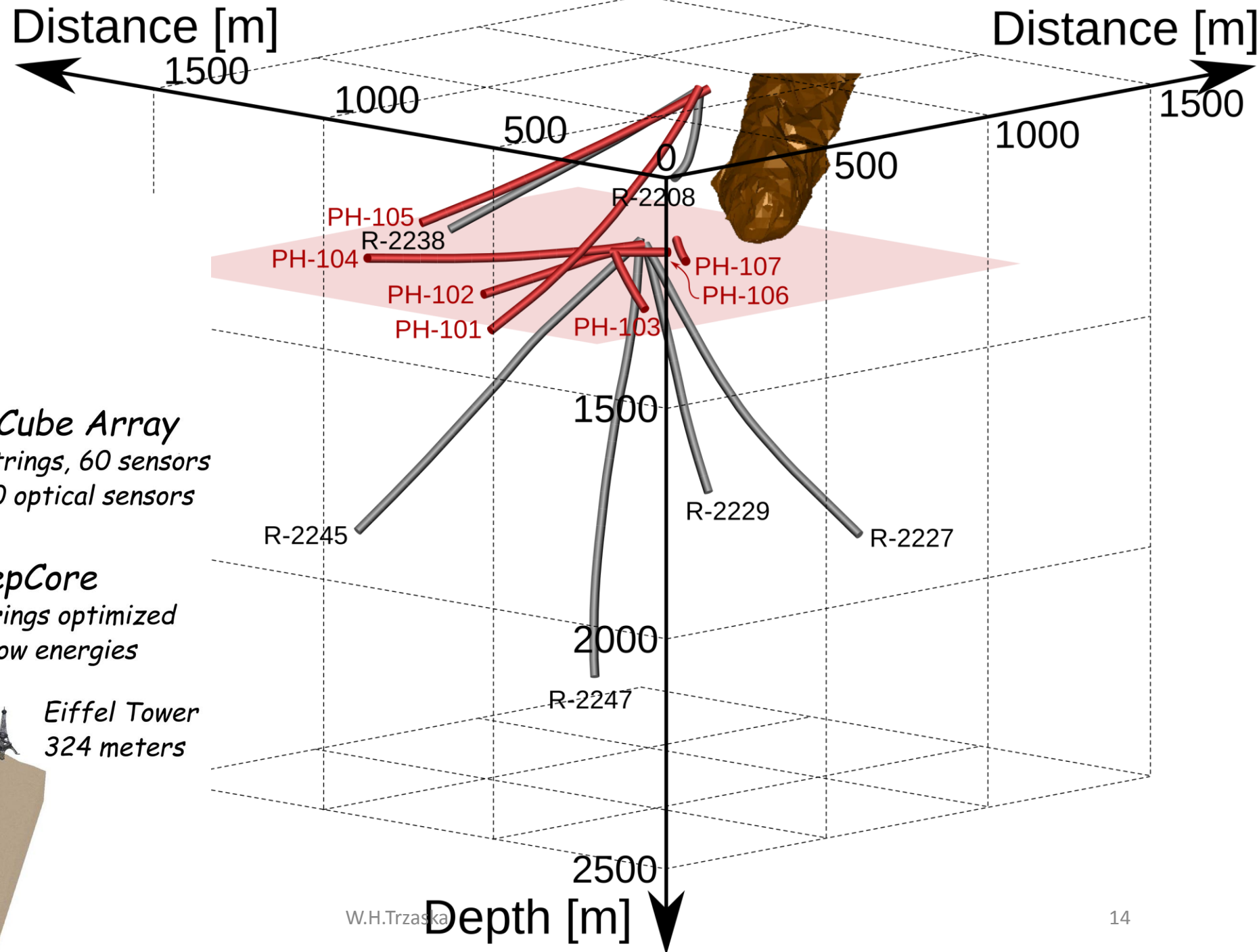


IceCube Array
86 strings, 60 sensors
5,160 optical sensors

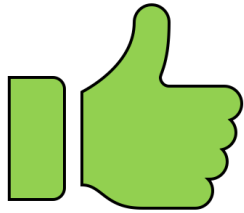
DeepCore
6 strings optimized
for low energies

Eiffel Tower
324 meters

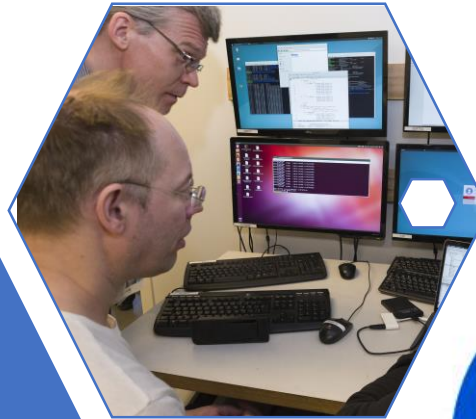
bedrock



3. Know-how network



Geology
HY, OY



Neutrino
community

Applied
Math
LUT



CALLIO LAB

Underground Centre for Science and R & D

THEIA

What is the bottleneck?

Funding for the
on-site team

To make measurements in the mine,
at least two scientist have to be
stationed in Pyhäsalmi

A grant from the Academy
would solve the problem





2015

2002

2002

1995



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Conclusions:

1. Neutrinos are important & relevant
2. Finland has a realistic chance to make a significant contribution

Thank you for your attention!

Courtesy: Wiki, TAUP, NASA, IceCube, JUNO, DUNE, KM3NeT
Artwork by Sandbox Studio, Chicago with Corinne Mucha

