

Higgs inflation: connecting electroweak physics to inflation

(work with Vera-Maria Enckell, Kari Enqvist,
Sami Raatikainen, Eemeli Tomberg and Lumi-Pyry Wahlman)

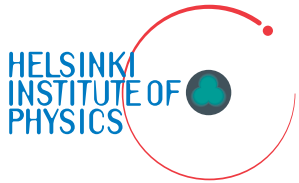
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Using what you have



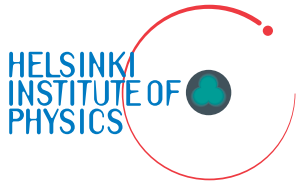
$$S = \int d^4x \sqrt{-g} \left[\frac{M_{\text{Pl}}^2 + \xi h^2}{2} R - \frac{1}{2} g^{\alpha\beta} \partial_\alpha h \partial_\beta h - V(h) \right]$$

$$V(h) = \frac{\lambda}{4} (h^2 - v^2)^2 \simeq \frac{\lambda}{4} h^4$$

- Non-minimal coupling $\xi h^2 R$ is the only new dimension 4 term for the combined Einstein-Hilbert + SM action.
- The coupling ξ is generated by renormalisation, even if it is classically zero.
- Non-minimal coupling enables Higgs inflation, which uses the only known scalar field that may be elementary.
(Bezrukov and Shaposhnikov: 0710.3755)



When the action is not enough



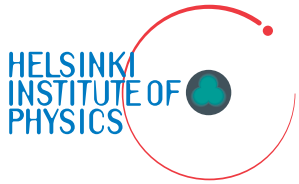
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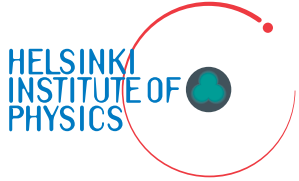
- Complication: classical low-energy action is not enough to specify the theory.
- Two sources of ambiguity.
 - Quantum theory: how to calculate loop corrections?
 - General relativity: what are the gravitational degrees of freedom?



Two aspects of spacetime

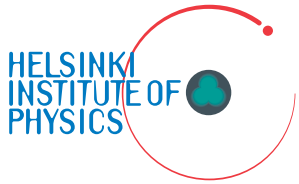


- There are different formulations of general relativity (GR).
- They are not modified theories, but alternative theories.
- In general, metric and connection are independent degrees of freedom.
- Metric describes distances in spacetime and inner products of vectors.
- Connection describes straight lines in spacetime and derivatives of vectors.



Metric and Palatini

- One formulation of GR is the metric formulation (1915), where the connection is defined in terms of the metric.
- Another is the Palatini formulation (1925), where the connection is determined by the field equations, instead of being constrained a priori.
- For the Einstein-Hilbert action R , the formulations are equivalent.

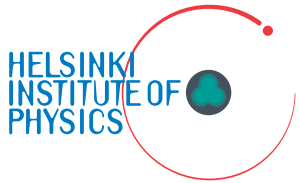


Higgs breaks the equivalence

- Different formulations of GR are inequivalent if the gravity action is more complicated than R or matter couples to the connection.
- Higgs exists, so nature is described by a scalar-tensor theory.
- Higgs couples to the connection, breaking the equivalence between the formulations. (Bauer and Demir: 0803.2664)



Two faces of Einstein gravity



$$S = \int d^4x \sqrt{-g} \left[\frac{1 + \xi h^2}{2} g^{\alpha\beta} R_{\alpha\beta}(\Gamma, \partial\Gamma) - \frac{1}{2} g^{\alpha\beta} \partial_\alpha h \partial_\beta h - V(h) \right]$$

- The Einstein frame is reached with the conformal transformation
 $g_{\alpha\beta} \rightarrow (1 + \xi h^2)^{-1} g_{\alpha\beta}$
- In the Palatini case, the conformal transformation does not affect the Ricci tensor.
- To recover canonical kinetic term, define new field χ :

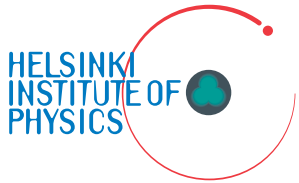
- metric: $\frac{d\chi}{dh} = \sqrt{\frac{1 + \xi h^2 + 6\xi^2 h^2}{(1 + \xi h^2)^2}} \simeq \frac{\sqrt{6}}{h} \Rightarrow h \propto e^{\chi/\sqrt{6}}$

- Palatini: $\frac{d\chi}{dh} = \sqrt{\frac{1 + \xi h^2}{(1 + \xi h^2)^2}} \simeq \frac{1}{\sqrt{\xi} h} \Rightarrow h \propto e^{\sqrt{\xi} \chi}$

- Polynomials are transformed into sums of exponentials.



Two faces of Einstein gravity



$$S = \int d^4x \sqrt{-g} \left[\frac{1}{2} g^{\alpha\beta} R_{\alpha\beta}(\Gamma, \partial\Gamma) - \frac{1}{2} g^{\alpha\beta} \frac{\partial_\alpha h \partial_\beta h}{1 + \xi h^2} - \frac{V(h)}{(1 + \xi h^2)^2} \right]$$

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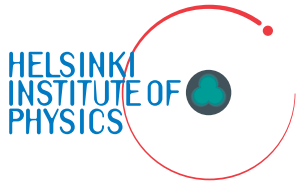
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Higgs potential in Metric vs Palatini



$$S = \int d^4x \sqrt{-g} \left[\frac{1}{2} R - \frac{1}{2} g^{\alpha\beta} \partial_\alpha \chi \partial_\beta \chi - U(\chi) \right]$$

- We get a different Einstein frame potential depending on the formulation of GR:

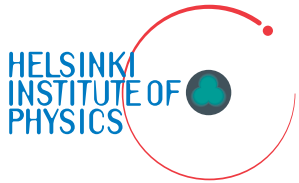
- metric: $U(\chi) \equiv \frac{V[h(\chi)]}{[1 + \xi h(\chi)^2]^2} \simeq \frac{\lambda}{4\xi^2} (1 - 2e^{-\frac{2}{\sqrt{6}}\chi})$

- Palatini: $U(\chi) \simeq \frac{\lambda}{4\xi^2} (1 - 8e^{-2\sqrt{\xi}\chi})$

- The potential is exponentially flat.



Predictions of Higgs inflation on the plateau

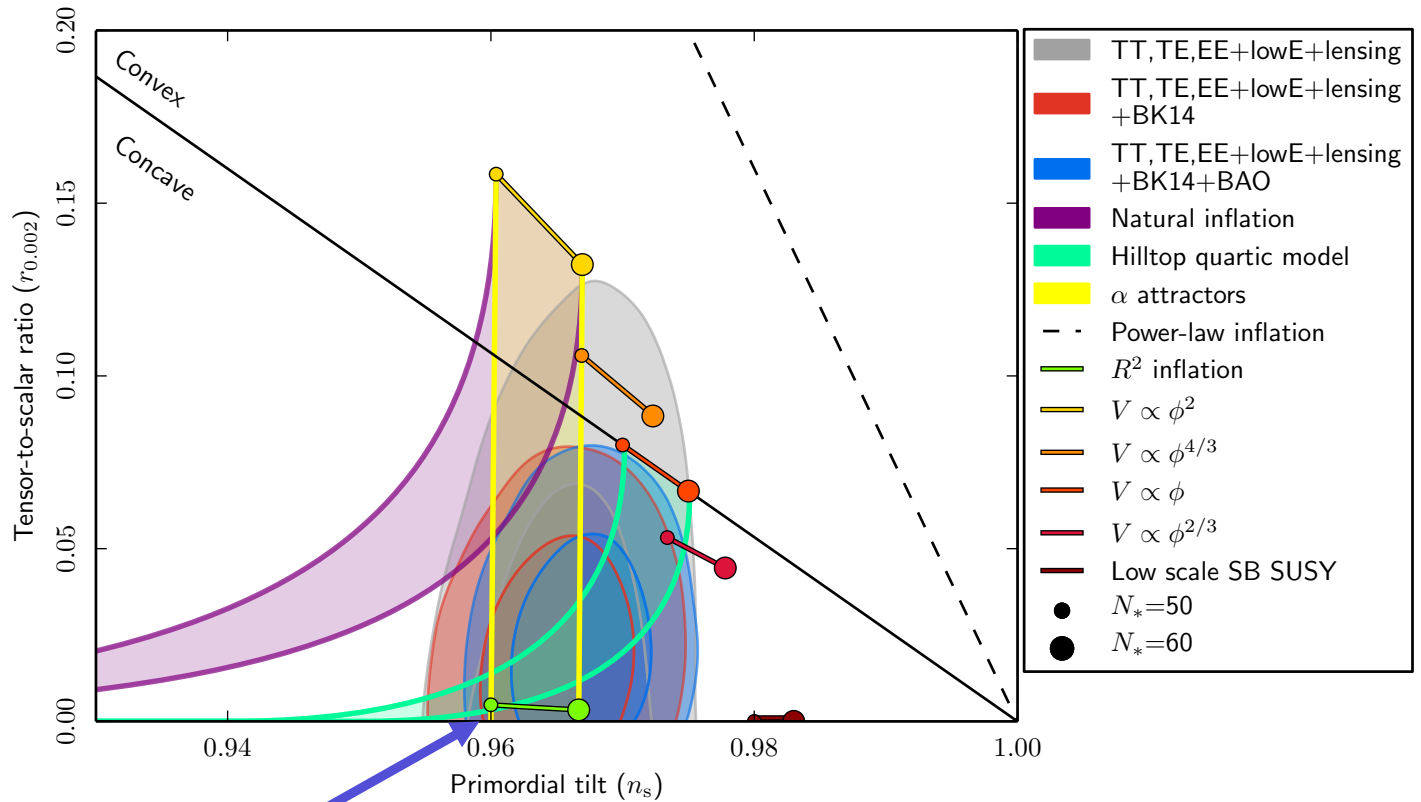
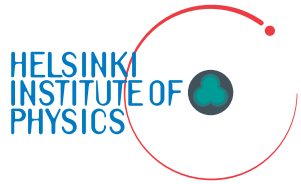


- Predictions for inflation depend on reheating.
- This is the SM, so reheating is known. (Garcia-Bellido, Figueroa and Rubio: 0812.4624, Rubio and Tomberg: 1902.10148)

- metric: $n_s = 0.96$, $r = 5 \times 10^{-3}$, $\xi = 4 \times 10^4 \sqrt{\lambda}$
- Palatini: $n_s = 0.96$, $r = \frac{8 \times 10^{-4}}{\xi} = \frac{8 \times 10^{-14}}{\lambda}$, $\xi = 10^{10} \lambda$

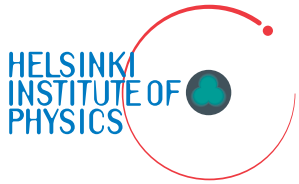


The data likes Higgs inflation





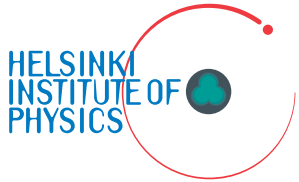
Renormalising the non-renormalisable



- Dimension 4 is key for flat potential: $U(\chi) = \frac{V[h(\chi)]}{[1 + \xi h(\chi)^2]^2} \simeq \frac{\frac{\lambda}{4} h^4}{\xi^2 h^4}$
- Due to loop corrections, inflationary predictions in principle depend on Higgs and top mass.
 - Consistency condition between cosmology and colliders?
- Due to gravity, the theory is not renormalisable.
 - Loop corrections boil down to prescriptions.
- Perturbative unitarity is violated at scales below M_{Pl} :
 $1/\xi$ (metric) or $1/\sqrt{\xi}$ (Palatini) $\sim 10^{14}$ GeV
- Loop corrections can also open up new inflation regimes.



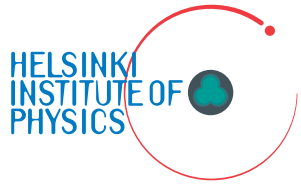
Higgs window to gravity



- Higgs inflation is a minimal bottom-up scenario for particle physics and gravity, using only known degrees of freedom.
- Different formulations of GR become inequivalent as Higgs couples to the connection.
- Predictions of Higgs inflation depend on the choice of theory of gravity and loop corrections.
 - Observations can tell what are the gravitational degrees of freedom.
- Different results for teleparallel formulation, loop quantum gravity action, ...

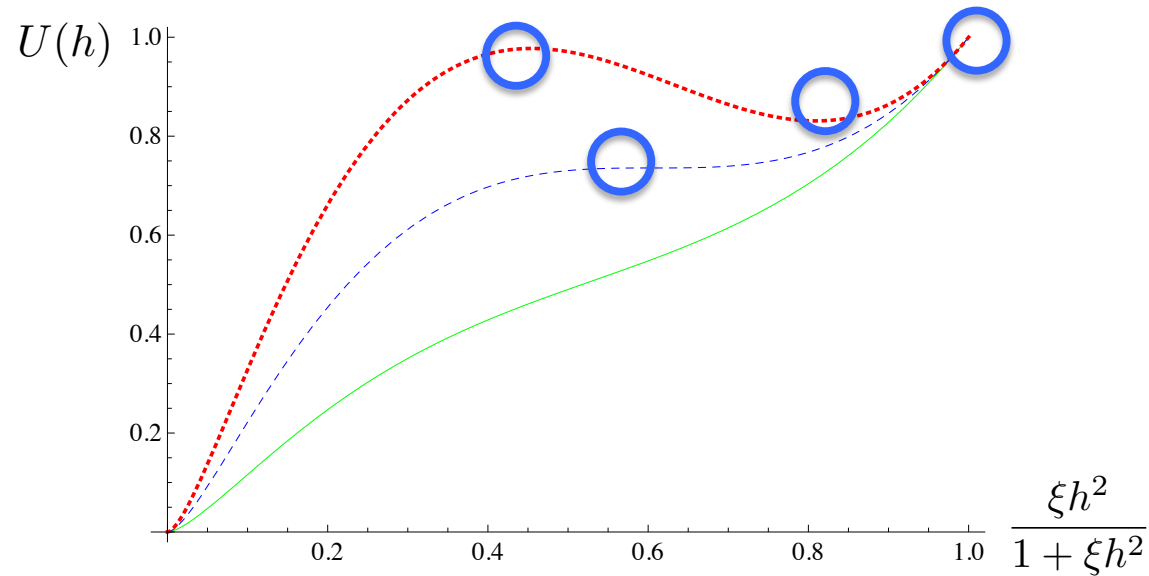
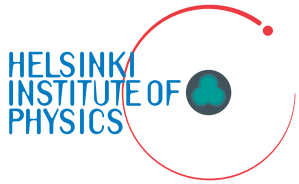


Backup slides





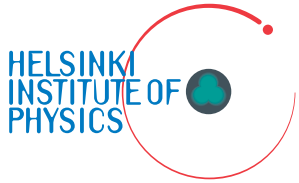
Loop-corrected potential



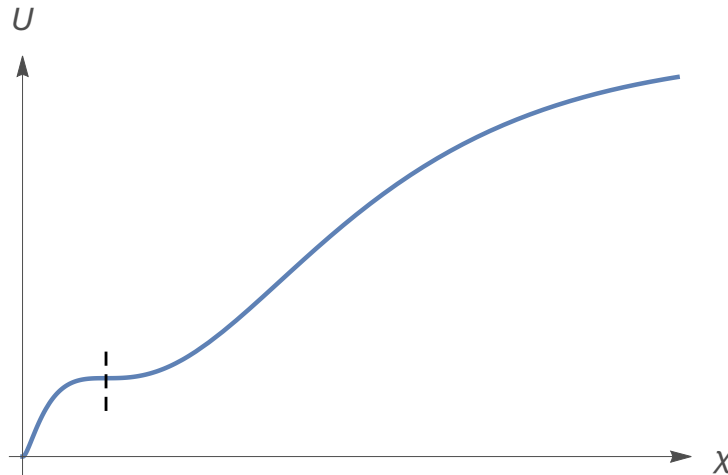
- Different inflationary possibilities:
 - Plateau: apparently not spoiled by loops.
 - Inflection point: can give $r \sim 0.1$.
 - False vacuum: new physics needed for graceful exit.
 - Hilltop: can give even smaller r than plateau.



A feature below the plateau

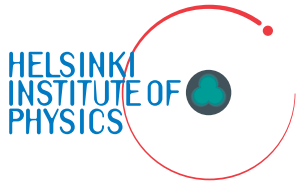


- We want to produce CMB perturbations on the plateau and have a slow-down feature at smaller field values.
- This can be generated by Standard Model loop corrections.





Black hole dark matter



- Loop corrections can be used to produce dark matter from the Higgs. (SR and Tomberg: 1810.12608)
- This requires a feature in the potential that slows down the field, enhancing perturbations.
- Once the perturbations re-enter Hubble radius after inflation, they collapse into black holes (BHs).
- These then evaporate down to Planck-scale relics.
- The relic BHs can be all of the dark matter.