

Overview of left-right symmetric extensions of the SM

Particle Physics Day / Katri Huitu, 7.11.2019



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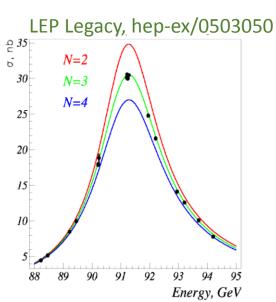
MOTIVATION

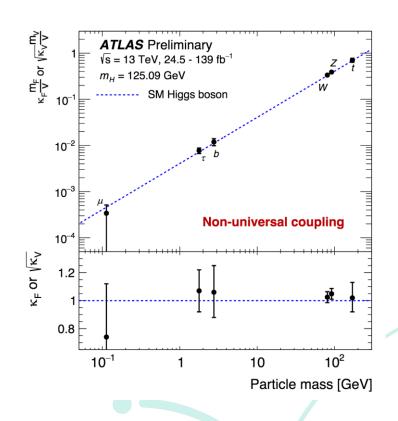
Problems in the Standard Model:

- Neutrino mass
- Dark matter
- Higgs mass
- Strong CP violation
- Weak CP violation
- No quantum gravity
- •

Some mundane questions:
Why three generations?
Why large mass differences?
Why left- and right-handed for

Why left- and right-handed fermions are fundamentally different?









Panico in EPS-HEP conference, 2019



A large number of models, including all kinds of combinations of earlier attempts.

Working close to experiments:

- Choose a model (personal preference is some well motivated model)
- Check the experimentally allowed parameter space (note: not a trivial task)
- Find testable predictions
- Discard the model/parameter region, if not viable



Left-right symmetric extensions of the Standard Model

Electromagnetic, strong and gravitational interactions conserve parity.

Maybe conserved at high energies also by weak interactions and dynamically broken

Pati, Salam (1974); Mohapatra, Pati (1975); Senjanovic, Mohapatra (1975)

If weak interactions were parity conserving at a high scale, every left-handed particle would need a right-handed counterpart



right-handed neutrinos needed

After neutrinos were deemed massive, neutrino mass has also become a motivation.

Smallest gauge group where left-right symmetry can be implemented is $SU(3)_c \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$ Davidson (1979), Mohapatra, Marshak (1980) Gauge symmetry could originate from SO(10) or E_6 unified model



Simple Left-Right Model with $SU(2)_L \times SU(2)_R \times U(1)_{B-L}$

Fermions:

$$(Q_L)^i = \begin{pmatrix} u_L^i \\ d_L^i \end{pmatrix} = (\mathbf{3}, \mathbf{2}, \mathbf{1}, \frac{1}{3}), (Q_R)^i = \begin{pmatrix} u_R^i \\ d_R^i \end{pmatrix} = (\mathbf{3}, \mathbf{1}, \mathbf{2}, \frac{1}{3}),$$

$$(L_L)^i = \begin{pmatrix} v_L^i \\ \ell_L^i \end{pmatrix} = (\mathbf{1}, \mathbf{2}, \mathbf{1}, -1), (L_R)^i = \begin{pmatrix} v_R^i \\ \ell_R^i \end{pmatrix} = (\mathbf{1}, \mathbf{1}, \mathbf{2}, -1)$$

with L,R the left- and right-chiral components, $P_{L,R} = (1 \pm \gamma_5)$

$$Q = I_{3L} + I_{3R} + \frac{B-L}{2}$$

Higgs sector: need to break $SU(2)_L \times SU(2)_R \times U(1)_{B-L} \rightarrow SU(2)_L \times U(1)_Y$

Break $SU(2)_R \times U(1)_{B-L} \rightarrow U(1)_Y$ by a triplet:

$$\Delta_R = \begin{pmatrix} \Delta_R^+ / \sqrt{2} & \Delta_R^{++} \\ \Delta_R^0 & -\Delta_R^+ / \sqrt{2} \end{pmatrix} = (\mathbf{1}, \mathbf{1}, \mathbf{3}, 2); \ \langle \Delta_R^0 \rangle = v_R$$

For simplicity, assume that Δ_L is decoupled Chang, Mohapatra, Parida, PRL (1984)

 $SU(2)_L \times U(1)_Y \rightarrow U(1)_{em}$ by a bidoublet:

$$\Phi = \begin{pmatrix} \phi_1^0 & \phi_2^+ \\ \phi_1^- & \phi_2^0 \end{pmatrix} = (\mathbf{1}, \mathbf{2}, \mathbf{2}, 0); \ \langle \Phi \rangle = \begin{pmatrix} \kappa & 0 \\ 0 & \kappa' e^{i\alpha} \end{pmatrix}$$

$$v_R \gg \kappa \gg \kappa'$$
; $\kappa^2 + \kappa'^2 = v_{SM}^2$

 $K^0 - \overline{K^0}$ mixing constrains $\approx \kappa \kappa' e^{i\alpha}$

Beall, Bander, Soni (1982)



Masses can be found from:

$$\mathcal{L}_{Y} = h_{q,ij}^{a} \overline{Q}_{L,i} \Phi_{a} Q_{R,j} + \tilde{h}_{q,ij}^{a} \widetilde{Q}_{L,i} \widetilde{\Phi}_{a} Q_{R,j} + h_{\ell,ij}^{a} \overline{L}_{L,i} \Phi_{a} L_{R,j} + \tilde{h}_{\ell,ij}^{a} \overline{L}_{L,i} \widetilde{\Phi}_{a} L_{R,j}$$
$$+ f_{ij} L_{R,i}^{T} Ci \tau_{2} \Delta_{R} L_{R,j} + h.c.$$

With
$$\widetilde{\Phi} = i\sigma_2 \Phi^* \sigma_2$$

Quark masses:

$$M_u = h_u \kappa + \tilde{h}_u \, e^{-i\alpha} \kappa'; \\ M_d = h_d e^{i\alpha} \kappa' + \tilde{h}_d \, \kappa$$

Charged lepton masses
$$M_\ell = h_\ell e^{ilpha} \kappa' + \tilde{h}_\ell \; \kappa$$



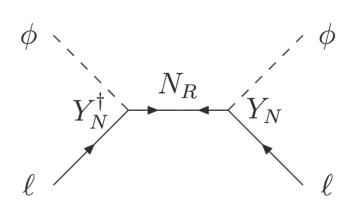
Neutrino masses
$$M_{\nu}=\begin{pmatrix} 0 & m_D \\ m_D^T & M_N \end{pmatrix}$$
; m_D = $h_{\nu}\kappa+\tilde{h}_{\nu}~e^{-i\alpha}\kappa'$; M_N = fv_R

Type-I seesaw, no neutrino Majorana mass: $m_{\nu} = -m_D M_N^{-1} m_D^T$

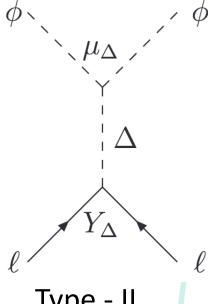
Note: if Δ_L included, one gets also Majorana mass for the light neutrinos, and

thus type-II seesaw

Gell-Mann, Ramond, Slansky (1979); Yanagida (1979); Mohapatra, Senjanovic (1980)



Type - I





Gauge bosons, $SU(2)_L \times SU(2)_R \times U(1)_{B-L}$

New gauge bosons connected to enlarged gauge group W_R , W, Z_R , Z, γ

$$m_W^2 = \frac{g_L^2}{2} (\kappa^2 + \kappa'^2);$$
 $m_{W_R}^2 = g_R^2 v_R^2$

$$m_Z^2 = \frac{g_L^2}{2 \cos^2 \theta_W} (\kappa^2 + \kappa'^2); \quad m_{Z_R}^2 = 2(g_R^2 + g_{BL}^2) v_R^2$$

Theoretically $\frac{g_R}{g_L} \gtrsim 0.55$; the smaller the ratio is, the heavier Z_R is

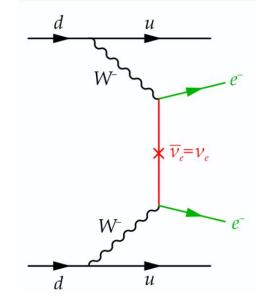
Mixings
$$\approx \frac{g_R}{g_L} \frac{m_V}{m_{V_R}}$$
, $V = W, Z$



Tests of left-right model

Lepton number violating processes:

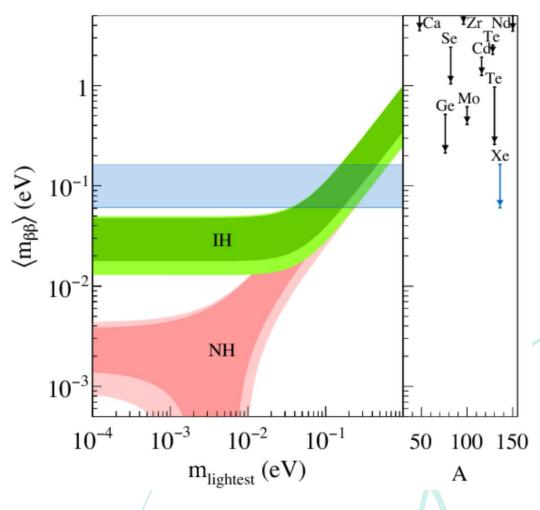
Neutrinoless double beta decay, $0\nu\beta\beta$



Lepton number violation in

Meson decays, $\tau \to \mu$, $e\gamma$, $\mu \to e\gamma$, $\mu \leftrightarrow e$ –conversion,...

J. Engel et al, Rept.Prog.Phys (2017)

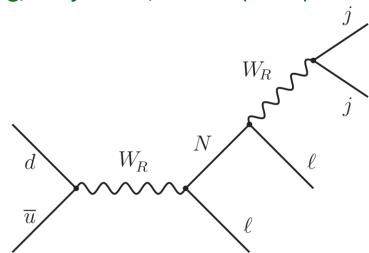


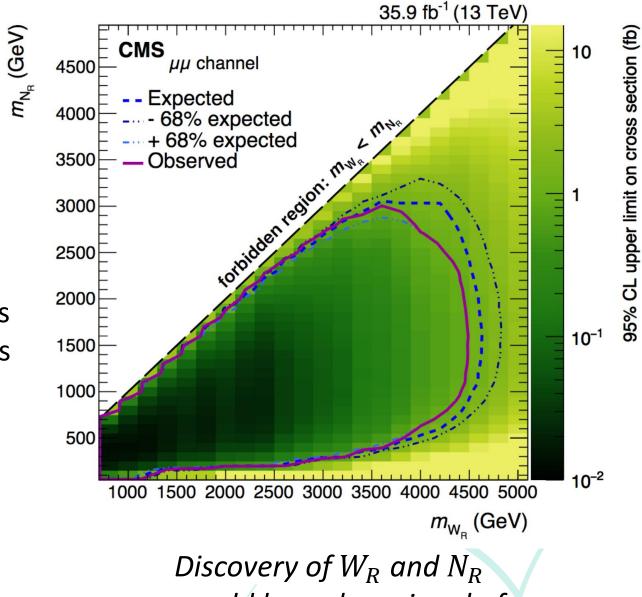


CMS, JHEP 05 (2018), arXiv: 1803.11116

$$W_R \to N_R \ell \to \ell \ell j j$$

N_R Majorana similar amounts of same sign and opposite sign dileptons Keung, Senjanovic, PRL 50 (1983) 1427





Discovery of W_R and N_R would be a clear signal of left-right symmetry

$$\Phi = \begin{pmatrix} \phi_1^0 & \phi_2^+ \\ \phi_1^- & \phi_2^0 \end{pmatrix}, \Delta_R = \begin{pmatrix} \Delta_R^+/\sqrt{2} & \Delta_R^{++} \\ \Delta_R^0 & -\Delta_R^+/\sqrt{2} \end{pmatrix}$$

Physical scalars:

 Φ : 4 singly charged d.o.f.; 2 scalar and 2 pseudoscalar d.o.f.

 Δ : 2 doubly charged d.o.f.; 2 singly charged, 1 scalar, 1 pseudoscalar d.o.f.

 W_R, W, Z_R, Z, γ : need 4 charged and 2 neutral longitudinal d.o.f.



$$H^{\pm\pm}, H_1^{\pm}, A, H_{1,2}^0, h$$

Of these, H_1^{\pm} , A, H_1^0 are nearly degenerate in mass and heavy from FCNC requirements; maybe not reachable at the LHC

 $H^{\pm\pm}$, H_2^0 do not (almost) couple to hadrons, and limits for their masses are < 1 TeV

Supersymmetric left-right model (SUSYLR)

 $R_{parity} = (-1)^{3(B-L)+2s}$ is an exact symmetry as B-L is a gauge symmetry;

if $SU(2)_R x U(1)_{B-L}$ is broken by $L=\pm 2$ triplet, remains unbroken also after symmetry breaking LSP a dark matter candidate

Mohapatra, PRD (1986); Font, Ibanez, Quevedo, PLB (1989); Martin PRD (1992)



There is no SUSY CP problem because of parity invariance

Kuchimanchi, PRL (1996), Mohapatra, Rasin, PRL (1996)

Also strong CP problem may be solved

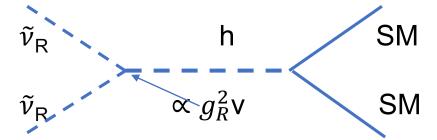
Beg, Tsao (1978); Mohapatra, Rasin, (1996); Babu, Dutta, Mohapatra (2002)



Dark matter in SUSYLR

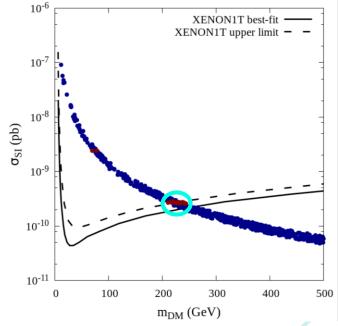
Frank, Fuks, KH, Rai, Waltari (2017); Chatterjee, Frank, Fuks, KH, Mondal, Rai, Waltari (2019)

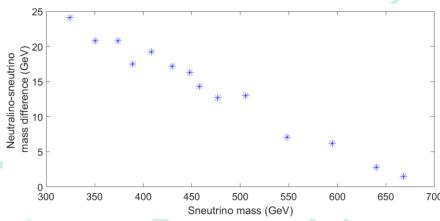
1) partners of right-handed neutrinos can produce DM



Right-sneutrino mass is the only free parameter \rightarrow relic density determines $m_{\widetilde{\nu}_R} \approx 250 \text{ GeV}$

Right sneutrino LSP with coannihilations with another sneutrino or neutralinos: $m_{\widetilde{\nu}_R} \approx 700 \text{ GeV}$





2) SUSY partners of gauge bosons or Higgses can be the dark matter

Chatterjee, Frank, Fuks, KH, Mondal, Rai, Waltari (2019)

Bidoublet higgsinos form a nearly degenerate set of four neutralinos and two charginos

Coannihilations cannot be avoided, when the lightest higgsino is the LSP

The Planck-value for relic density is achieved with 750 GeV LSP higgsino (could be silghtly decreased with further coannihilations, with e.g. sneutrino)

Note that in this case the spectrum is rather heavy and compressed difficult to detect.

3) Right-handed neutrinos can contribute to dark matter Bhattacharya, Ma, Wegman (2014)

In the minimal form need to take into account the one-loop corrections to scalar potential to stabilize the model and $v_R \lesssim 10$ TeV

In addition to h, A_1 , H_1 , H_1^{\pm} could be within LHC reach: almost degenerate with mass $\gtrsim 700$ GeV

One $H^{\pm\pm}$ light, not much heavier than 1 TeV possibility to exclude the model

 $H^{\pm\pm} \to \ell^{\pm} \ell'^{\pm}$; $m_{H^{\pm\pm}} \gtrsim 700$ GeV, if decays to electrons or muons

...KH, Maalampi, PLB (1995), KH, Maalampi, Raidal, (1994); KH, Maalampi, Pietilä, Raidal (1997); Barenboim, KH, Maalampi, Raidal (1997); Dutta, Mohapatra (1998); Chacko, Mohapatra (1998); KH, Pandita, Puolamäki (1998); Babu, Mohapatra (2008); Frank, Korutlu (2011); Chun, Sharma, PLB (2013); Basso, Fuks, Krauss, Porod (2015); Babu, Patra (2016); Frank, Ghosh, KH, Rai, Saha, Waltari (2014); Frank, Fuks, KH, Rai, Waltari, (2017); Chatterjee, Frank, Fuks, KH, Mondal, Rai, Waltari, (2019)...



Conclusions

Need some experimental guidance

Meanwhile:

consider relevant possibilities check all possible implications indirect experiments can provide the scale

Experiments need to know the new type of signals to analyze and

theorists need to apply experimental results to their favourite theory

For the field it is essential to fully use the LHC potential