

Light (pseudo)scalars lurking undetected? Some suggestions on mass reconstruction

Biswarup Mukhopadhyaya
Regional Centre for Accelerator-based Particle Physics
Harish-Chandra Research Institute
Allahabad, India

May, 2019

- *No clear signal for new physics has emerged so far at the LHC*

- *No clear signal for new physics has emerged so far at the LHC*
- **Possibility 1:** *No new particle/interaction within kinematic reach*

- *No clear signal for new physics has emerged so far at the LHC*
- **Possibility 1:** *No new particle/interaction within kinematic reach*
- **Possibility 2:** *There is something new within reach, but is imperceptible due to insufficient signal rate*

- *No clear signal for new physics has emerged so far at the LHC*
- **Possibility 1:** *No new particle/interaction within kinematic reach*
- **Possibility 2:** *There is something new within reach, but is imperceptible due to insufficient signal rate*
- **Q1:** *Could this be the case in the EWSB sector?*

- *No clear signal for new physics has emerged so far at the LHC*
- **Possibility 1:** *No new particle/interaction within kinematic reach*
- **Possibility 2:** *There is something new within reach, but is imperceptible due to insufficient signal rate*
- **Q1:** *Could this be the case in the EWSB sector?*
- **Q2:** *If so, how well can the high-luminosity run reveal them?*

- *No clear signal for new physics has emerged so far at the LHC*
- **Possibility 1:** *No new particle/interaction within kinematic reach*
- **Possibility 2:** *There is something new within reach, but is imperceptible due to insufficient signal rate*
- **Q1:** *Could this be the case in the EWSB sector?*
- **Q2:** *If so, how well can the high-luminosity run reveal them?*
- *We consider a sample scenario*

The existence of dark matter is indicated/suggested by...

- *Expectation of new physics affecting the EWSB sector arises because of (a) the naturalness issue (b) vacuum stability*

The existence of dark matter is indicated/suggested by...

- *Expectation of new physics affecting the EWSB sector arises because of (a) the naturalness issue (b) vacuum stability*
- *A common new physics option: two Higgs doublet models (2HDM)*

The existence of dark matter is indicated/suggested by...

- *Expectation of new physics affecting the EWSB sector arises because of (a) the naturalness issue (b) vacuum stability*
- *A common new physics option: two Higgs doublet models (2HDM)*
- *Low-energy 2HDM can retain vacuum stability upto the Planck scale*

*N. Chakrabarty, U. K. Dey, BM (2014),
N. Chakrabarty (2016), N. Chakrabarty, BM
(2017)*

$$\Phi_1 = (\phi_1^+, \phi_1^0), \quad \Phi_2 = (\phi_2^+, \phi_2^0)$$

$$\longrightarrow h, H, A, H^\pm$$

Suppression of tree-level flavour violation

$\Rightarrow Z_2$ symmetry hypotheses

\Rightarrow Type 1 / Type 2 / Type X / Type Y.....

Type X: One doublet couples to all quarks, the other to leptons

$$\mathcal{L} = -y_u \bar{Q}_L U_R \tilde{\Phi}_2 + y_d \bar{Q}_L d_R \Phi_2 + y_\ell \bar{L}_L e_R \Phi_1 + h.c.$$

Suppressed quark coupling for H, A, H^\pm close to the 'alignment limit': $(\beta - \alpha) = \pi/2$

$$\tan \beta = v_2/v_1, \quad \alpha = h - H \text{ mixing angle}$$

For large $\tan \beta$

$$\mathcal{L}_{\bar{u}uH} \sim \cot \beta$$

$$\mathcal{L}_{\bar{d}dH} \sim \cot \beta$$

$$\mathcal{L}_{udH^\pm} \sim \cot \beta$$

$$\mathcal{L}_{\bar{\ell}\ell H} \sim \tan \beta$$

$$\mathcal{L}_{\nu\ell H^\pm} \sim \tan \beta$$

$$\mathcal{L}_{\bar{u}uA} \sim \cot \beta$$

$$\mathcal{L}_{\bar{d}dA} \sim \cot \beta$$

$$\mathcal{L}_{\bar{\ell}\ell A} \sim \tan \beta$$

H, H^\pm, A : suppressed interaction with the quark sector

Also, $\mathcal{L}_{VVH} \sim \cos(\beta - \alpha)$

Production suppressed at LHC:

A potential challenge for the high-luminosity run

- *A light A ($m_A \simeq 30 - 60 \text{ GeV}$) is consistent with all data*

- *A light A ($m_A \simeq 30 - 60 \text{ GeV}$) is consistent with all data*
- *Claim: Such a Type- X scenario achieves better consistency with $(g - 2)_\mu$*
A. Broggio et al. (2014), E. J. Chun, J. Kim (2016)

- *A light A ($m_A \simeq 30 - 60 \text{ GeV}$) is consistent with all data*
- *Claim: Such a Type-X scenario achieves better consistency with $(g - 2)_\mu$*
A. Broggio et al. (2014), E. J. Chun, J. Kim (2016)
- *Bottomline: substantial positive contributions may serve to reduce the persistent discrepancy*
A. Keshavarzi et al. (2018)

Contributions to $(g - 2)_\mu$ in 2HDM....

- *One-loop contribution: $\sim m_\mu^4$*

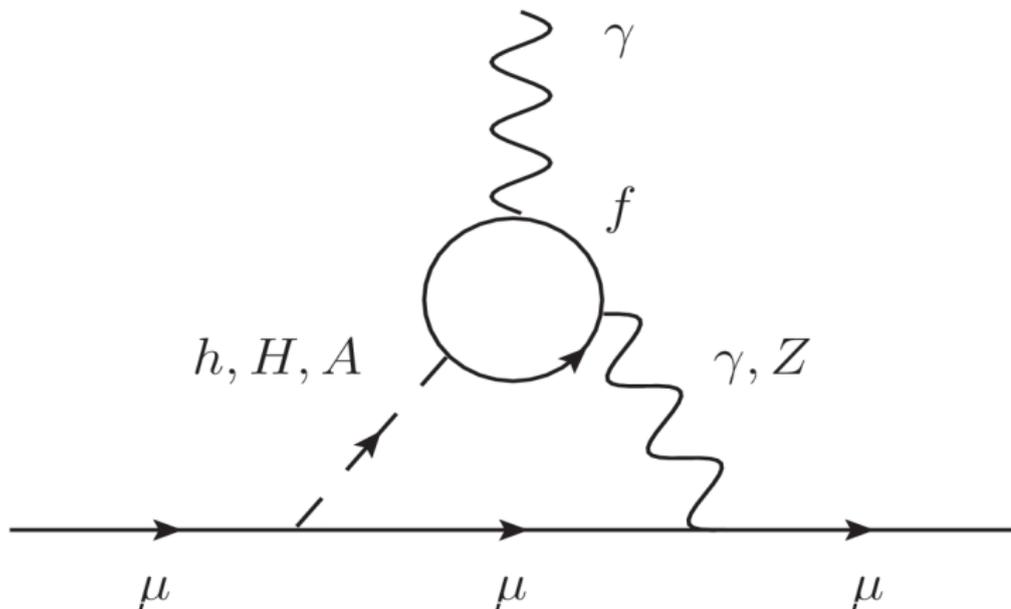
Contributions to $(g - 2)_\mu$ in 2HDM....

- *One-loop contribution: $\sim m_\mu^4$*
- *Two-loops: The Bar-Zee diagrams offer enhancement $\sim \frac{m_f^2}{m_\mu^2} \cdot \frac{\alpha_{em}}{6\pi^2}$
 $f = t, b, \tau$*

Contributions to $(g - 2)_\mu$ in 2HDM....

- *One-loop contribution: $\sim m_\mu^4$*
- *Two-loops: The Bar-Zee diagrams offer enhancement $\sim \frac{m_f^2}{m_\mu^2} \cdot \frac{\alpha_{em}}{6\pi^2}$
 $f = t, b, \tau$*
- *Negative contributions from h, H
Positive contributions from A
D. Chang (2001)*

The Bar-Zee diagrams....



Type-X 2HDM helps for $m_A \lesssim 60\text{GeV}$, $\tan \beta \simeq 40 - 50\text{GeV}$

EW precision data ($m_W, \sin \theta_W, Z\ell\bar{\ell}$ coupling at Z-peak)

$\Rightarrow m_A \simeq 40 - 60 \text{ GeV}$ requires

$$|m_H - m_{H^\pm}| \lesssim 25 \text{ GeV}$$

Vacuum stability:

$$m_A \simeq 40 - 60 \text{ GeV} \Rightarrow m_{H^\pm} \lesssim 200 \text{ GeV}$$

A. Broggio et al. (2014)

Signals at the LHC....

- *Studied mostly for*
 $h \longrightarrow AA, A \longrightarrow \tau\tau$

S. Kanemura *et al.* (2014, 2016), E. J. Chun *et al.* (2015),.....

- *Studied mostly for*
 $h \longrightarrow AA, A \longrightarrow \tau\tau$
S. Kanemura *et al.* (2014, 2016), E. J. Chun *et al.* (2015),.....
- *Difficulty: precise mass reconstruction...*
For $m_A \approx 50$ GeV, the τ is not energetic enough for the collinear approximation to work
Also impossible if the A is pair-produced (more than two particles are sources of MET)

Signals at the LHC....

- *Studied mostly for*
 $h \longrightarrow AA, A \longrightarrow \tau\tau$
S. Kanemura *et al.* (2014, 2016), E. J. Chun *et al.* (2015),.....
- *Difficulty: precise mass reconstruction...*
For $m_A \approx 50$ GeV, the τ is not energetic enough for the collinear approximation to work
Also impossible if the A is pair-produced (more than two particles are sources of MET)
- *A solution: utilise $A \longrightarrow \mu^+ \mu^-$*
E. J. Chun, S. Dwivedi, T. Mondal, BM, PLB 774, 20 (2017)

Signals at the LHC....

In the region of interest,

$$B(A \rightarrow \tau^+\tau^-) \gtrsim 99\%, \quad B(A \rightarrow \mu^+\mu^-) \simeq 0.35\%$$

Preferred channel :

$$pp \rightarrow hX, \quad h \rightarrow AA \rightarrow (\mu^+\mu^-)(\tau^+\tau^-)$$

Constraint on the Type-X 2HDM parameter space:

$$\mu\mu\tau\tau \text{ search,....} \Rightarrow B(\rightarrow AA) \leq 3 - 4\%$$

Higgs coupling studies $\Rightarrow \cos(\beta - \alpha) \lesssim 0.1$ for large $\tan \beta$
(Excepting a narrow strip allowing $\cos(\beta - \alpha) \simeq 0.2 - 0.3$.)

CMS, JHEP11(2018)018 (Nov, 2018),

ATLAS-CONF-2019-005,

D. Taylor, Talk at Moriond, 2019

Signals at the LHC....

Signal to look for:

a $\mu^+\mu^-$ peak + two jets including one tagged τ

Jet- $p_T \geq 20 - 25\text{GeV}$ leads to the jj -and $\mu^+\mu^-$ -peaks close together (within $\lesssim 15\text{ GeV}$)

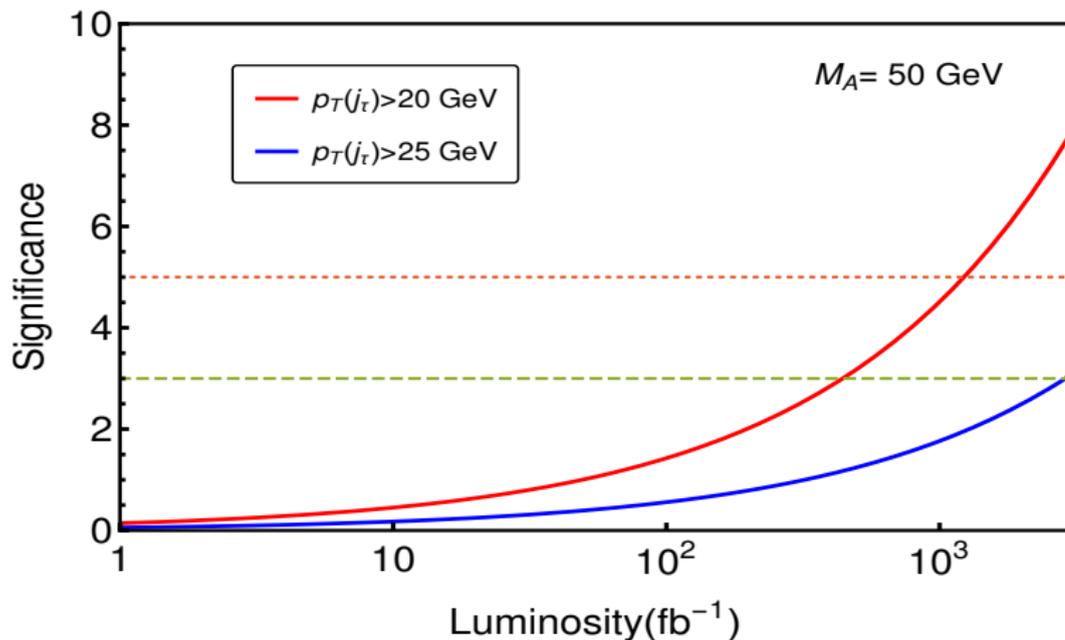
Major SM backgrounds:

- $\mu^+\mu^- + \text{jets}$
- $VV \rightarrow \mu^+\mu^- + \text{jets}$
- $t\bar{t}$

The key feature in background reduction:

Closeness of the $m_{inv}(\mu\mu)$, $m_{inv}(jj)$ peaks

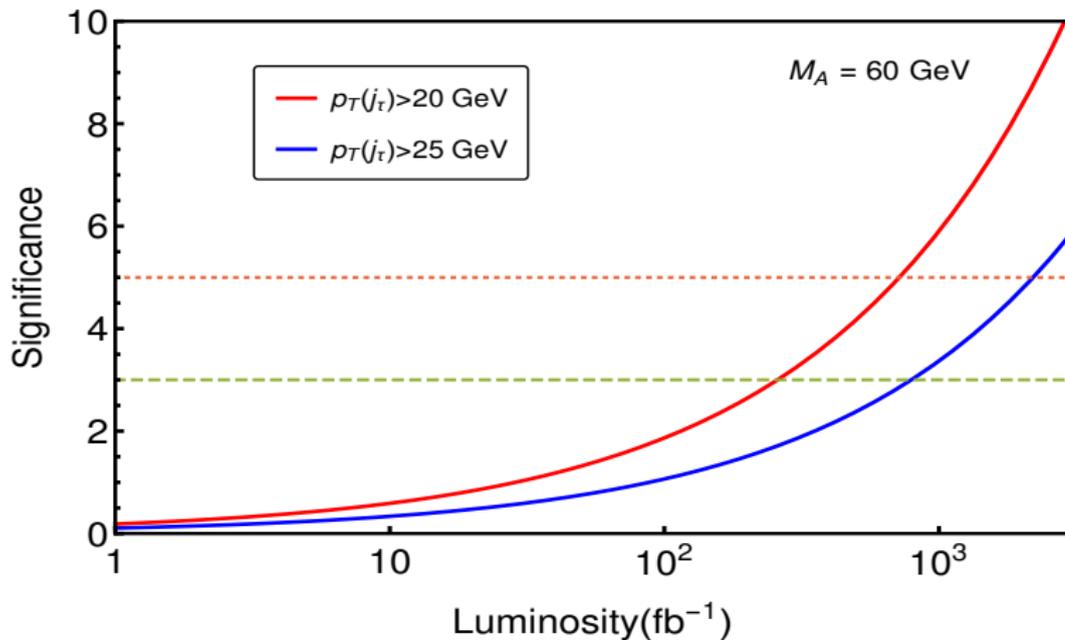
Signals at the LHC....



$m_A = 50 \text{ GeV}$, $B(h \rightarrow AA) = 3\%$

$$\text{Significance} = \sqrt{2[(S + B)\ln(1 + S/B) - S]}$$

Signals at the LHC....



$m_A = 60 \text{ GeV}$, $B(h \rightarrow AA) = 3\%$

Reconstruction of H, H^\pm

E. J. Chun, S. Dwivedi, T. Mondal, BM, S. K Rai, PRD 98, 075008 (2018)

- *The region of interest: $m_H, m_{H^\pm} \approx 200 \text{ GeV}$, with $|m_H - m_{H^\pm}| \lesssim 25 \text{ GeV}$.*

Reconstruction of H, H^\pm, \dots

E. J. Chun, S. Dwivedi, T. Mondal, BM, S. K Rai, PRD 98, 075008 (2018)

- *The region of interest: $m_H, m_{H^\pm} \approx 200 \text{ GeV}$, with $|m_H - m_{H^\pm}| \lesssim 25 \text{ GeV}$.*

- *$pp \longrightarrow H^\pm A, HA$
 $H^\pm \longrightarrow WA, H \longrightarrow ZA$*

$BR \approx 50\%$ for decays

*A-reconstruction information via the μ^\pm mode
 \Rightarrow The heavier scalars reconstructed*

Reconstruction of H, H^\pm, \dots

- $\mu\mu + jj$ with one tagged τ -jet

Various cuts reduce combinatorial backgrounds:

- $m_{\mu\mu} \geq 90$ GeV: to utilise the boost
- Geometry of chosen events $\Rightarrow \Delta\phi_{(\mu\mu),j\tau} \geq 1.6$

Reconstruction of H, H^\pm, \dots

- $\mu\mu + jj$ with one tagged τ -jet

Various cuts reduce combinatorial backgrounds:

- $m_{\mu\mu} \geq 90$ GeV: to utilise the boost
- Geometry of chosen events $\Rightarrow \Delta\phi_{(\mu\mu),j\tau} \geq 1.6$
- $3.0 - 3.5\sigma$ with 3000fb^{-1}

*Excess in a bin covering a broad peak
($\simeq 25 - 30$ GeV)*

Summary and conclusions....

- *A Type-X 2HDM can accommodate a light pseudoscalar (40 - 60 GeV) consistently with all current data, and help in explaining $(g - 2)_\mu$.*

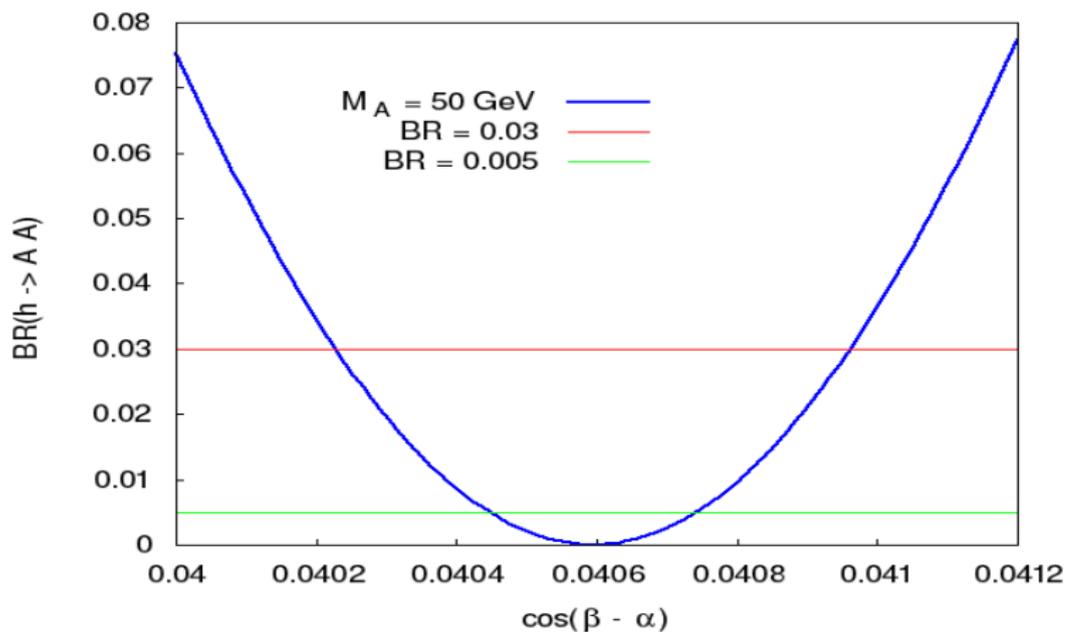
Summary and conclusions....

- *A Type-X 2HDM can accommodate a light pseudoscalar (40 - 60 GeV) consistently with all current data, and help in explaining $(g - 2)_{\mu}$.*
- *The light pseudoscalar mass can be reconstructed precisely using the $\mu^+ \mu^-$ decay modes, to 4-6 σ significance, for a few hundred to 3000 fb^{-1} .*

Summary and conclusions....

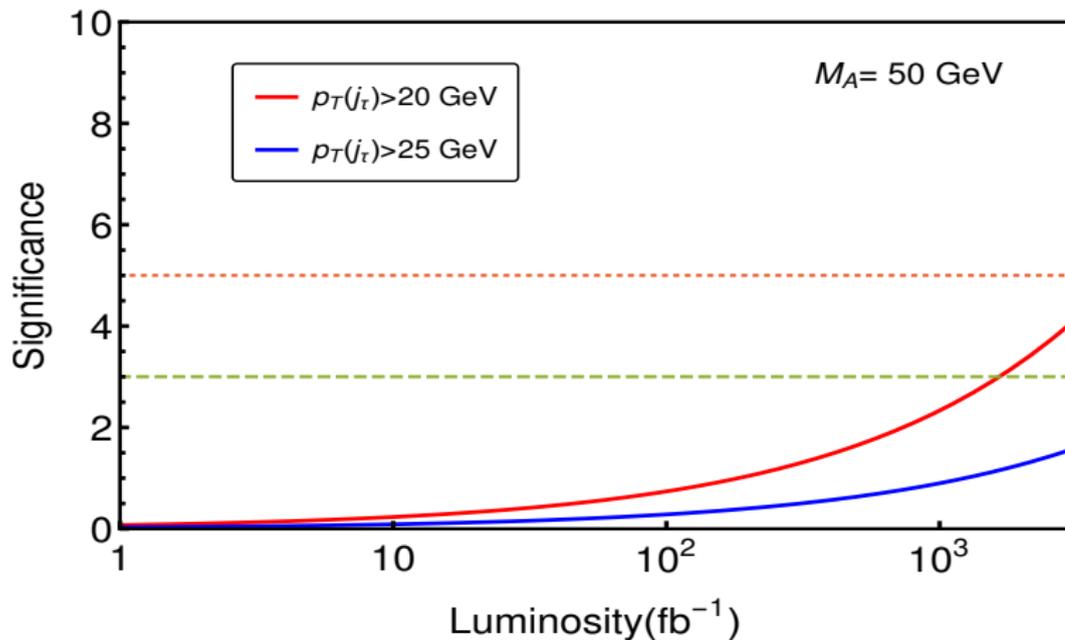
- *A Type-X 2HDM can accommodate a light pseudoscalar (40 - 60 GeV) consistently with all current data, and help in explaining $(g - 2)_\mu$.*
- *The light pseudoscalar mass can be reconstructed precisely using the $\mu^+ \mu^-$ decay modes, to 4-6 σ significance, for a few hundred to 3000 fb^{-1} .*
- *The H and the H^\pm are also reconstructible, and their closeness in mass established, to about 25-30 GeV.*

Signals at the LHC....



Branching ratio against $\cos(\beta - \alpha)$

Signals at the LHC....



$m_A = 50 \text{ GeV}$, $B(h \rightarrow AA) = 1.5\%$