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CHIPP Roadmap 2020:
Status and Outlook on
Particle Physics in Switzerland

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55 **1 Executive summary**

56 [Main Editor: Rainer] [(1-2 pages) – Summary of the scope, the national and international landscape, the future trends and the
57 major challenges in the field. Identification of the major findings and recommendations.]

58 **2 Findings and recommendations**

59 [Main Editor: Rainer] [(2-4 pages) – More specifically identify here a series of findings and related recommendations. These can
60 be already imbedded in the various sections of the document and listed again here for easier overview and reference. Simplistic
61 example: Finding 1: Infrastructure XYZ is essential for our community in the field ABC, but it will no more meet the international
62 standards in 5 years. Recommendation 1: There is the need for a major upgrade of this facility or the building of a new
63 infrastructure serving the whole Swiss community.]

64 **3 Foreword**

65 [Main Editor: Rainer] [(1-2 pages) – Describe the process that led to this roadmap and its endorsement by the community.
66 Explain how you tried to reach out to the whole community.]

67 **4 List of authors**

68 [Main Editor: Angela] [(1 page) – Provide the list of the authors with affiliations and possibly separating the main editors from
69 other contributors. Possibly add here thanks for support received for the layout work, etc.]

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71 *board members)*

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86 **5 Table of contents**

87 [Main Editor: Angela] [(1-2 pages) – Include a table of contents with page numbers. Possibly with in addition a list of figures
88 and/or of tables, but this is probably only useful for a really long document.]

89 **6 Purpose scope**

90 [Main Editor: Rainer] [(1-3 pages) – Clarify what is very generically meant by the overall field (e.g. biology), because this is
91 not necessarily obvious to the target audience. What are the main scientific questions and challenges in this field? What is the
92 objective of the document and the point of view adopted (very inclusive or more focused)? Clarify the separation with nearby
93 disciplines (i.e. what is or is not covered here). State whether or not there was an attempt to prioritize the needs of specific
94 infrastructures and how this is reflected in the document.]

95 [AS: Intro from Klaus, can be used here if needed]

96 Testing the Standard Model (SM) and searching for physics beyond this theory is done in direct and indirect ways at the high
97 energy frontier at the LHC and future colliders. It can be done in a complementary way also at lower and much lower energies
98 where often much larger intensities of certain particle species allow studying rare processes and small effects at high precision.

99 Two scenarios are usually considered: Either the mass scale of new physics is too high to be directly probed, or the coupling
100 to known particles is so weak that it has escaped detection so far. In both cases, deviations of precision observables from their
101 SM prediction or the occurrence of SM-forbidden effects could provide indirect, but unambiguous, evidence for new physics.
102 Of course, a golden scenario would show deviations in some precision observables and, simultaneously, direct production of
103 new particles in high-energy collisions. Often direct and indirect detection experiments investigate the same underlying physics
104 scenarios and have thus been grouped together into the CHIPP pillar-1. Here we deal with the low-energy, high-intensity and
105 precision frontier.

106 [(1-3 pages) – Clarify what is very generically meant by the overall field (e.g. biology), because this is not necessarily obvious
107 to the target audience. What are the main scientific questions and challenges in this field? What is the objective of the document
108 and the point of view adopted (very inclusive or more focused)? Clarify the separation with nearby disciplines (i.e. what is or
109 is not covered here). State whether or not there was an attempt to prioritize the needs of specific infrastructures and how this is
110 reflected in the document.]

111 [Pillar III scope by Ruth]

112 Particle physics is concerned with the quest for the fundamental laws of nature and its elementary constituents. At the present
113 stage of our knowledge the fundamental interactions are the combination of the Standard Model (SM), described as an $SU(3) \times$
114 $SU(2) \times U(1)$ gauge theory together with the Yukawa interactions with a Higgs field, and of gravity, described by General
115 Relativity (GR). So far no deviations from these laws have been found, but severe open theoretical problems remain. Also some

116 experimental tensions exist, e.g. in the flavor sector and on the generation of neutrino masses. Finally, the nature of dark matter
117 and dark energy, the dominant constituents of our Universe, discovered so far only via gravitational interactions, is completely
118 unknown.

119 These questions are addressed on the one hand by direct, controlled experiments as described in pillars I and II, but also via
120 searches, direct and indirect, for dark matter (DM), dark energy, and for new interactions beyond the standard model at very high
121 energies. This is the approach taken in pillar III.

122 The advantage of indirect searches via astrophysical objects and events is that we can use objects/events in our Universe as
123 high-energy accelerators. Sufficient precision and the combination of data from diverse observations are needed in order to draw
124 solid conclusions. These very challenging endeavours relate pillar III to astrophysics. We want to investigate whether the very
125 high-energy sources and events we observe in the Universe can occur within the standard model of particle physics or whether
126 we have to go beyond. As a goal at its own right, we also want to better understand the Universe, its origin and evolution.

127 In order to understand the physics of the most energetic events in the Universe, observations in all wavelength bands of electro-
128 magnetic radiation, from radio over microwave, optical and X-rays to γ -ray are needed. These are combined with cosmic ray
129 detections as well as neutrino and gravitational wave observations. It has become clear that this multi-messenger approach is
130 required for a full picture. Only the combination of all of these data allows us to understand astrophysical accelerators suffi-
131 ciently well, to determine whether or not we see deviations from the standard model of particle physics at very high energies.
132 The present 'showcase' example is the binary neutron-star coalescence GW170817 discovered first with LIGO/Virgo and later
133 seen by many electromagnetic counter parts in X-ray, optical, IR, UV and radio. These data allow us to address long standing
134 questions about the interior of neutron stars: is there a quark gluon plasma phase, do hyperons form inside neutron stars? But
135 they also provide a measurement of the Hubble constant and show that, as in GR, the speed of gravitons is equal to the speed
136 of light. This fact excludes entire classes of theories modifying GR which were proposed as explanations of dark energy. Other
137 important examples with significant Swiss involvement will be mentioned below.

138 Clearly, pillar III is very complementary to the pillars I and II. Most statements of pillar III are model dependent and request
139 confirmation either by more direct experiments as they are performed in pillars I and II or, if the energy scale is not avail-
140 able at terrestrial accelerators, via independent observations of another messenger. Furthermore, only pillar III is sensitive to
141 gravitational interactions.

142 In the quest for dark matter both direct and indirect searches are part of pillar III. But also here, a signal even in a direct search
143 experiment, requires a significant amount of interpretation and requires confirmation by another, direct or indirect experiment,
144 or at accelerators, before firm conclusions can be drawn. A very recent example is the 2-7keV recoil-electron excess detected by
145 Xenon1t. Is it a hint of axions from the sun or just a contamination of the experiment by Tritium, or something completely else?

146 **[Theory scope by Gino]**

147 The search for a theory able to overcome the difficulties of the SM, extending its validity range, and possibly predicting some
148 of its free parameters, is the ultimate goal of the forefront research in particle physics, both on the experimental and on the
149 theoretical side. We can roughly divide the research directions in theoretical particle physics in four main categories:

150 **I. Precise SM physics.** An essential ingredient to make progress in the field is to develop precise predictions, within the
151 SM, for particle-physics experiments: without precise predictions to compare with, we cannot interpret the experimental

152 results. While the SM is an apparently simple theory in abstract terms, making precise predictions for quantities observed
153 in realistic experiments is often an extremely challenging goal, which requires sophisticated tools, both from the analytical
154 and the numerical side.

155 **II. Model-building and beyond-the-SM (BSM) phenomenology.** A similarly core ingredient of the theoretical research in
156 particle physics is to develop new models of fundamental interactions, able to address some of the shortcomings of the
157 SM, and to understand how these models could possibly be tested in present and future experiments.

158 **III. Cosmology, astroparticle, gravitational physics.** A growing aspects of theoretical particle physics concerns cosmology
159 and, more generally, the connections between particle physics, astrophysics and gravitational physics. The challenge
160 here is understanding particle physics and fundamental interactions through predictions and observations of astrophysical
161 phenomena ranging from cosmic messengers to the large-scale structure of the universe.

162 **IV Formal theory developments.** The last essential ingredient of the theoretical research in fundamental physics deals with
163 more formal aspects. There are still many regimes of QFT that are poorly understood, especially when going beyond a
164 four-dimensional flat geometry, as required by General Relativity. A rich structure of symmetries connecting seemingly
165 different QFT regimes has also emerged in the last few years. A deeper investigation of these aspects, as well as the
166 development of theoretical frameworks able to go beyond QFT, such as String Theory, are the subject of this more formal
167 line of research.

168 **7 The present Swiss landscape**

169 [Main Editor: Rainer and whoever is willing to help] [(5-15 pages) – This section will be very specific to each field. It can
170 be subdivided into research topics and/or per methodology (theoretical versus experimental, laboratory versus field study, and
171 so on). Alternatively, it could be per geographical location, if there are well defined topics covered by different institutions. It
172 shall be as much as possible inclusive of all the community to leave nobody out. It shall also show what are the major topics in
173 Switzerland, which institutes are leaders in specific research areas, and possibly also what is less developed yet (especially if a
174 new infrastructure is foreseen to fill this gap). How strong is the Swiss network: how much do the scientists of different institutes
175 collaborate together? What are the main infrastructures used? Are they accessible to researchers from other institutions?]

176 [The following split in sections is intended as an intermediate step for collection of the necessary material. The main editor of
177 the chapter will merge accordingly after the material is available.]

178 **Input from accelerator research**

179 [Editor: Lenny]

180 **7.1 Accelerator Physics and Technology**

181 **7.1.1 Swiss Landscape**

182 Accelerator science and technology developments in Switzerland are at the heart of several large research infrastructures used
183 in particle physics, but also in a number of fields like chemistry, life and material sciences. Switzerland maintains a strong

184 tradition of accelerator R&D, both at Paul Scherrer Institute and at CERN. Going back to the design and construction of High
185 Intensity Proton Accelerators at PSI in the 1970s with its world highest power proton beam, it resulted in such state of the
186 art synchrotron light sources as Swiss Light Source (SLS) and SwissFEL. The Swiss Accelerator Research and Technology
187 CHART Collaboration between CERN, ETHZ, EPFL, University of Geneva and PSI is maintaining this tradition. The mission
188 of CHART is to support the future oriented accelerator project FCC at CERN and the development of accelerator concepts beyond
189 the existing technology. With extraordinary support by SERI, ETH Board and participating institutions CHART contributes to
190 future accelerator driven research infrastructures benefiting science and society.

191 **Input from Pillar 1**

192 [Editor: Anna]

193 **7.2 High energy**

194 Swiss institutes are founding members and significant contributors to three of the four high energy physics experiments at CERN:
195 ATLAS (U Bern and U Geneva), CMS (UZH, PSI and ETHZ) and LHCb (EPFL and UZH). The activities of the Swiss scientists
196 have been spanning a very broad spectrum of physics pursuits: from precision measurements of the Standard Model (SM) to
197 explorations of new phenomena that could answer the open questions of our universe. Since the discovery of the Higgs boson
198 in 2012, a much deeper characterization of this new particle has been achieved. After the long shutdown from 2013 to 2014
199 to consolidate the LHC magnets to achieve higher bending fields, Run 2 data was collected from 2015 to 2018 at a center-of-
200 mass energy of 13 TeV. This higher energy provided increased cross section of production processes, and along with record
201 performances of instantaneous luminosity, led to enhanced statistics of acquired datasets. These large datasets have allowed for
202 a better understanding of the detectors, new sophisticated developments in reconstruction algorithms, and reduced associated
203 uncertainties. At the same time, a significant leap in the precision of the theoretical predictions has taken place. These are the
204 basic ingredients that have made the Run 2 LHC physics program more compelling than ever before. New results have included
205 the observation of the Higgs boson decay to $\tau^+\tau^-$, evidence for the Higgs decay to bb^- , the observation of electroweak same-
206 sign WW production, and evidence for top-quark pairs produced in association with a Higgs boson. Moreover, ATLAS and CMS
207 results have opened up new phase space for new physics searches and further strengthened bounds on existing models for Beyond
208 the Standard Model (BSM) physics. The large dataset delivered by the LHC so far and the forthcoming runs promise a strong
209 continuous physics output, which will extend the boundaries of knowledge at the energy frontier.

210 [TODO: Should mention somewhere the visibility of Swiss scientists in the large collaborations - convenorships, coordination
211 roles etc, as a global comment maybe rather than providing specific examples?]

212 The most significant discovery in particle physics in the last few decades has been of the Higgs boson, the physical manifestation
213 of the Higgs field, which provides mass to both fermions and bosons, and establishes the mechanism for how the high-energy
214 electroweak interaction is broken into electromagnetism and the weak interaction at low energies.

215 The Higgs boson was discovered, and can currently only be studied, by the ATLAS and CMS experiments at the CERN Large
216 Hadron Collider (LHC). The LHC has been delivering significant rates of proton-proton collisions to ATLAS and CMS since
217 2010. The CDF and D0 experiments in the U.S. had been searching for the Higgs boson using proton-antiproton collisions at
218 the Fermilab Tevatron for the previous decade, but had concluded operation in 2011, reporting evidence of the Higgs boson

219 in its decay to b-quarks by July 2012. At the same time, after collecting data at center-of-mass energies of 7 and 8 TeV, the
220 ATLAS and CMS collaborations each announced definitive observations of a new particle, consistent with the Higgs boson,
221 on July 4th, 2012. The observations were driven by unmistakable, coincident signals of the Higgs boson decaying to photons,
222 which occurs due to the interaction of the Higgs boson with fermions, as well as the decaying to Z bosons, which occurs due
223 to the interaction of the Higgs boson with the electroweak bosons. Both the ATLAS and CMS collaborations observed these
224 two signals, with all four signals appearing at the same mass of approximately $125 \text{ GeV}/c^2$, and with comparable rates that were
225 consistent with the expectation from the standard model. These results were a magnificent success of the LHC and the ATLAS
226 and CMS collaborations, and marked the dawn of a new age of study of the Higgs boson.

227 Of immediate interest was whether the new particle discovered had the correct couplings, spin, parity, and width to be the
228 standard model Higgs boson. All significant Higgs boson production modes and decay modes needed to be established to
229 determine whether any deviations from the standard model rates would reveal themselves. The mass of the Higgs boson also
230 became a high priority to understand, since its value is considered theoretically unnatural, due to large radiative corrections that
231 would tend for it to be at the Planck mass, 10^{16} orders of magnitude larger than the measured value. In the standard model, once
232 the mass of the Higgs boson is known, all branching fractions and production modes can be predicted, although often with large
233 theoretical uncertainties. And while the standard model was constructed with the simplest Higgs mechanism of only one Higgs
234 boson particle, many extensions to the standard model that more completely predict a richer sector of particles with multiple Higgs
235 bosons, as well as new interactions with standard model and BSM particles.

236 Since the discovery of the new particle, ATLAS and CMS have succeeded in measuring the overall production and decay rate of
237 the new particle to within 20% at $\sqrt{s} = 7$ and 8 TeV, and to within 10% at 13 TeV, finding consistent results in all final states
238 with expectations from the standard model Higgs boson. The charge, spin, and parity have been established, and the width and
239 lifetime are also consistent with expectations. The Higgs boson is now an accepted component of the standard model of particle
240 physics.

241 Swiss physicists have made leading contributions towards the discovery and measurements of the Higgs boson, beginning with
242 major contributions to the inner trackers of ATLAS and CMS, which are crucial for identifying the largest decay mode of $H \rightarrow b\bar{b}$,
243 as well as the electromagnetic calorimeter of CMS that provides precise measurements of photons and electrons necessary for
244 establishing the $H \rightarrow \gamma\gamma$, $H \rightarrow ZZ \rightarrow e^+e^- + X$, and $H \rightarrow WW \rightarrow e\nu + X$ channels. In the first years after the discovery, Swiss
245 physicists helped establish the observation and precise measurements of the $H \rightarrow \gamma\gamma$ process, measuring its mass and production
246 rates precisely at 7 and 8 TeV center-of-mass energies. With combined results from $\sqrt{s} = 7, 8$ TeV, Swiss physicists established
247 direct evidence of the Higgs boson decaying to fermions through the $H \rightarrow b\bar{b}$ and $H \rightarrow \tau\tau$ processes. With just a few years
248 of data and analysis, the Higgs boson mass was measured to 0.2% precision in the $H \rightarrow ZZ$ and $H \rightarrow \gamma\gamma$ channels, with Swiss
249 contributions ensuring well-calibrated electromagnetic energy scale and resolution, precise momentum resolution of the trackers,
250 as well as direct contributions to the $H \rightarrow \gamma\gamma$ analysis.

251 Beyond the study of the Higgs properties, the precise measurement of the parameters of the SM is a pillar of the physics program
252 of the LHC. Parameters of the SM, such as the weak mixing angle and the masses of the top quark and the W boson are
253 theoretically predicted within the SM and are also measured with high precision at collider experiments. The comparison of
254 predicted and measured values of these parameters constitute a crucial test of the SM and can uncover new phenomena. Swiss
255 scientists are specifically engaged in the study of the top quark, the heaviest of the elementary particles. It is unique among the
256 SM quarks since it decays before forming hadronic bound states. The top quark has been a natural probe of new physics due

257 to its large mass and strong coupling to the electroweak symmetry breaking sector. Measurements involving top quarks bring
258 key information on fundamental interactions at the electroweak symmetry-breaking scale and also of the strong interactions.
259 The production of SM processes with top quarks is a vital component of the LHC search program as it represents some of the
260 most significant background sources to new physics searches. A better understanding of the properties and the characteristics of
261 top-quark production and decay mechanisms has a direct impact on the constraints on new physics processes.

262 Direct searches for BSM physics have a high priority in the present Swiss landscape and are strongly represented by all Swiss
263 LHC groups. The search program is broad, covering and expanding the whole existing landscape while several centres of exper-
264 tise guarantee a high impact of Swiss contributions. Such a diverse portfolio has many benefits. First of all it puts Switzerland in
265 an excellent position when signs of new phenomena are found. Searches performed by Swiss physicists cover the “classical” sig-
266 natures, such as SUSY or heavy resonances produced by new hypothetical particles, and extend to more unconventional signatures,
267 including rare and forbidden decays from flavour physics phenomena, originating from more exotic, yet viable and interesting
268 models. This approach is a logical consequence of the increase of LHC’s integrated luminosity and the confidence that is gained
269 with operating over many years the LHC detectors. Instead of turning the crank, rapid progress by the Swiss groups is made by
270 innovation, both in covering new ground in the phase space of experimental signatures, and in designing new powerful tools in
271 the areas of trigger, simulation, reconstruction and data analysis, as to optimally exploit the unique LHC data set and guide the
272 way for future experiments. Among the most interesting examples is the use of advanced triggering methods to scout the data
273 for rare signatures in ways that conventional trigger strategies would not allow.

274 In recent years, the exploration of heavy flavour has been dominated by results of the LHCb experiment, which has been designed
275 for precise measurements of CP violation and rare heavy hadron decays, exploiting the large heavy quark production at the LHC.
276 The primary physics goals are to characterize in detail the flavour structure in the quark sector, and look for NP effects in the decay
277 of charm and bottom hadrons. The Swiss groups have played major roles in the LHCb experiment and have carried out a variety
278 of physics analyses, mostly in flavour physics, but also in other areas such as direct searches of long-lived particles. They have
279 pioneered important measurements, among which an angular analysis that showed a yet-to-be-understood significant discrepancy
280 with respect to the Standard Model. They have also performed a number of CP violation measurements exploring a large territory
281 where BSM physics may have appeared. They are looking for hints of lepton flavour violation either by measurements of lepton
282 flavour universality, or direct searches in decays of hadrons.

283 In the HEP research, Switzerland is fully embracing the ongoing transformative period, which was heralded by the Higgs boson
284 discovery in 2012. The enormous amounts of data collected by the LHC experiments allow for the data themselves to move
285 into the limelight, giving way to modern data analysis tools, such as machine learning (ML), which are being pioneered in
286 Switzerland. Such approaches include model-agnostic searches by use of anomaly detection techniques. Use of machine learning
287 is being employed at all levels of event processing, from triggering to reconstruction and to data analysis, creating tools that can
288 enable precision in measurements and reach in searches that was never achieved before.

289 One of the major pursuits in collider high energy physics is the design, construction and commissioning of particle detectors. At
290 the LHC these come with major challenges. The large number of interactions per bunch crossing (pile-up) allows for large event
291 rates, which is essential on the quest for rare phenomena. This also leads to enormous detector occupancy, high trigger rates and
292 severe radiation levels, which drive the requirements on detector granularity, fast electronics and radiation hardness.

293 In the past, Swiss groups have played major roles in the development of the detectors that have been in operation during Run 1

294 and Run 2. Swiss groups have also lead upgrades of or additions to the existing detector. These activities were primarily focused
295 on the tracking detectors (U Geneva and U Bern for ATLAS, UZH, PSI and ETHZ for CMS and EPFL and UZH for LHCb) with
296 design, construction and commissioning of the entire system (sensors, mechanics, read-out electronics, cooling). There have
297 also been major contributions to the calorimeter (ETHZ for CMS) with substantial involvement in construction, commissioning
298 and operation of the ECAL detector. Swiss groups have also been involved in developments in the trigger and data acquisition
299 systems (U Geneva and U Bern for ATLAS). The LHCb collaboration is currently installing a major detector upgrade driven by
300 the need to go to a full readout at 40 MHz and a software-only trigger. This will enable the collection of 5 fb^{-1} per year with
301 a much improved efficiency, especially for heavy-flavour decays without muon in the final state. For this upgrade, the EPFL
302 and UZH groups are involved in the design and construction of the tracker detectors. More specifically, EPFL has proposed and
303 developed the scintillating fibre (SciFi) technology for the replacement of tracking stations downstream of the dipole magnet.

304 The upgraded high-luminosity LHC (HL-LHC) is expected to start operations in 2026 and deliver about 250 fb^{-1} per year
305 until about 2038, which is approximately 5 times the current data size. The challenge associated with data taking under these
306 conditions will be unprecedented and the experiments have developed plans to upgrade their detectors in order to cope with those.
307 ATLAS and CMS will proceed with a the complete replacement of the inner and outer tracker detectors, including an extension
308 of forward tracking to higher pseudorapidity (up to $|\eta| = 4$) with extended pixel detectors and improved track trigger capabilities.
309 In particular CMS will have a hardware-level track trigger. The inner pixel detector specifications are at the forefront of radiation
310 tolerance and rate capabilities for silicon detectors. Both experiments are upgrading the electronics of their calorimeters for faster
311 readout, while the CMS endcap calorimeters are being replaced with high-granularity and radiation-hard silicon detectors. A new
312 timing layer is being proposed to reduce the effects of pileup down to levels similar to current conditions.

313 Overall the goal of the upgrade is to replace sub-detector components as needed to retain a robust, fast and radiation-hard
314 multipurpose-detector using as little material as possible to have the same, or better, performances in HL-LHC conditions as
315 compared to Run 2. In particular pileup rates and occupancy need to be mitigated, while keeping low transverse momentum
316 requirements for the main triggers and guarantee precise measurements up to large rapidity. Switzerland is playing a major
317 role, both in ATLAS (U Geneva and U Bern) and CMS (ETHZ, UZH, PSI), in the design and construction of the inner tracking
318 detector, including detector module and readout chip design, powering and qualification, as well as detector system electronics,
319 mechanics and cooling. In CMS the Swiss groups will also be responsible for the barrel electromagnetic calorimeter electronics,
320 and help build the barrel timing layer. In the ATLAS and CMS experiments, the Swiss groups are contributing to the TDAQ and
321 track trigger upgrade.

322 [TODO: Should add something on generic sensor RnD here??]

323 Equally important to the detector construction is the computing infrastructure, without which the enormous amount of data can
324 not be processed and analysed. The LHC computing in Switzerland has been addressed at two scales. At a small scale, user
325 specific data analyses are performed by each Institute independently adopting what is more convenient for their needs relying
326 on local resources (Tier-3). On the large scale instead each country participating in a LHC experiment provides an agreed
327 amount of computing power and storage to allow the reconstruction of the events collected by the detectors, the analysis of the
328 data and the production of simulated data sets. Switzerland is a part of the Worldwide LHC Computing Grid (WLCG) project,
329 which is a global collaboration of around 170 computing centres in more than 40 countries. WLCG itself is one of the essential
330 partners within the European Grid Infrastructure (EGI) community, which has the role of coordinating the overall operation of
331 the European computing resources based on grid technologies.

332 The load on each LHC participating country is regulated by annual pledges: each country is expected to contribute to the global
333 effort by providing both the hardware and the person power needed to operate it. The Swiss Institutes working at the LHC fulfil
334 their pledges using Tier-2 resources in Bern and at the Swiss National Supercomputing Centre (CSCS) in Lugano. Smaller non
335 LHC experiments, because of their typically lighter computing requirements, use dedicated resources tailored to their needs.

336 *Do we need subsection here?*

337 The lack of direct discovery of new physics by LHC experiments has stimulated people to consider alternatives. One possibility
338 is to extend the SM by adding new light states with feeble couplings to SM particles, creating the so-called Dark Sector (DS),
339 which is becoming an extremely fertile domain of exploration.

340 Swiss researchers are pioneering DS searches in fixed-target beam experiments at CERN. ETHZ is among the original proposers
341 and one of the main drivers of the NA64 experiment searching for DS at the SPS. NA64 is an international collaboration of about
342 50 scientists. It was approved in 2016 and since then has been collecting data.

343 Swiss groups are also involved in future experiments that will shed light in a complementary way to these obscure parts of the
344 new particle landscape. They are involved in the FASER experiment, approved in 2019 and currently under construction, and
345 they are discussing other experiments that will explore the DS at a smaller scale and more targeted ways than the large LHC ones.
346 [TODO: (Following up from EU strategy see if this is a good compromise of a phrasing - everything else on SHiP is removed for
347 now.)].

348 FASER (ForwArd Search ExpeRiment) is a new small that will be placed 480 meters downstream of the ATLAS experiment at
349 the CERN LHC. FASER is designed to capture decays of exotic particles, produced in the very forward region, out of the ATLAS
350 detector acceptance. Beyond searching for new physics, the FASER experiment will also provide capability to measure properties
351 of neutrinos at the highest human-made energies ever recorded. The FASER experiment is being built recycling existing spare
352 detector pieces by other experiments, thus minimising the construction cost, what makes it a low-cost high-gain project. FASER
353 is expected to take data for the whole Run 3, leading to first results that will shed light on currently unexplored phenomena,
354 having the potential to make a revolutionary discovery. Two Swiss universities, U Bern and U Geneva, are significantly involved
355 in the effort since the design of the experiment. Scientist from U Bern have been among the main proponents of the neutrino
356 extension of the experiment. Many leading roles in the collaboration are held by Swiss scientists.

357 [TODO: (FASERnu and other neutrino physics with forward detectors will need to be merged with other neutrino stuff?)]

358 **7.3 Low energy**

359 Switzerland is in a unique position in terms of large-scale facilities for fundamental particle physics. It is a host country of CERN
360 and it operates large-scale infrastructure at the national Paul Scherrer Institute (PSI), in particular the High Intensity Proton
361 Accelerator complex HIPA. HIPA is home to the world's highest power (1.4 MW) proton cyclotron delivering from several target
362 stations the highest intensities of low momentum pion and muon beams, as well as of ultracold neutrons. A substantial fraction
363 of the world-leading research with pions, muons and ultracold neutrons is done at PSI. In order to cope with the requirements of
364 the experiments and to carry forward the leading position in the international context, design and feasibility studies are ongoing

365 on how to further improve beam intensities and quality. CERN houses the antiproton decelerator (AD) facility, the only place
366 on the planet for research with low-energy antiprotons and anti-hydrogen; It is currently being upgraded with the addition of the
367 extreme low energy antiproton (ELENA) ring, which will be fully operational for all experiments in the AD after LS2.

368 All facilities serve an international community and provide the involved Swiss groups with excellent opportunities to initiate,
369 pursue and lead cutting-edge research. It is a considerable advantage of the Swiss groups to have some of the world's best
370 infrastructure within the immediate reach of their scientists, students and technical workforce. In what follows, the activities of
371 Swiss groups are summarised and some strategic considerations and orientation are put forward.

372 Most of the Swiss activities in low-energy particle physics make use of the unique facilities at PSI. PSI's Laboratory for Particle
373 Physics itself has three groups directly involved in the inhouse-physics program. University groups from U Basel, U Bern, U
374 Geneva, ETHZ, and UZH are involved in various international collaborations at PSI. Many more groups from Swiss universities
375 use PSI particle beams of protons, pions, muons, electrons, positrons and neutrons for R&D on detectors and electronics, and
376 for irradiation studies. Smaller Swiss efforts take place at the CERN AD and its new extreme low energy antiproton ring, the
377 neutron sources at ILL and ESS and the positron laboratory at ETHZ,

378 Over the past decade, particle physics at PSI has attracted an increasing number of Swiss groups and individuals, and this trend
379 is likely to continue. On the one hand, this is due to the unique reach of low-energy precision experiments in search for new
380 physics. Some of the tightest constraints on new physics are coming from this field. A growing effort in the global particle
381 theory community is working on the necessary tools to quantitatively evaluate precision experiments and to allow comparisons
382 with bounds obtained from high-energy physics. Swiss particle theory is greatly contributing or even driving the progress of
383 the field. On the other hand, comparatively small collaborations and shorter time scales allow individuals to have an enormous
384 impact on an experiment. PhD students can get a complete experimental physics education from conceiving ideas, via setting
385 up measurements to producing results. PSI is the world-leading center concerning the search for CP violation with the neutron
386 electric dipole moment, for charged lepton flavor violation experiments with muons, and for exotic atom spectroscopy with
387 muons and pions.

388 **Input from Pillar 2**

389 [\[Editor: Michele\]](#)

390 **Input from Pillar 3**

391 [\[Editor: Ruth\]](#) In Switzerland we have strong groups working on different aspects of multi-messenger astrophysics (X-rays, γ -
392 rays, neutrinos, cosmic rays,), we are involved in the forefront of direct dark matter searches and studies of the nature of neutrino
393 masses, and we have strong groups in theoretical cosmology studying the problem of dark energy or modifications of General
394 Relativity. [Also these last groups are involved in observations, but these concern radio (HIRAX, SKA) microwave (Planck, SPT)
395 and optical (Euclid, LSST) wavelengths. These observations are discussed in the CHAPS roadmap.] There are also significant
396 theoretical activities in gravitational wave research in connection to the LIGO/Virgo, LISA and the future Einstein Telescope
397 experiments but so far these are not complemented with experimental work by a Swiss group.

398 In this section we describe the main experiments with Swiss involvement and their goals in some detail.

399 7.4 Cosmic rays, neutrinos, X- and gamma-rays

400 Until the advent of the accelerators in the 60ies, about 20 particles were discovered in cosmic rays, such as the muons, the pions,
401 the kaons and more. Cosmic rays are highly energetic particles, mostly protons, from the Universe with energies beyond those of
402 particles accelerated by the Sun. Their energies reach up to 10^{20} eV and beyond (ultra-high-energy cosmic rays (UHECR)). This
403 energy is orders of magnitudes higher than the ~ 14 TeV achieved by LHC in the centre of mass. What are the mechanisms which
404 accelerate particles in ‘cosmic accelerators’ up to these extreme energies? How do particles interact at these extreme energies?
405 Cosmic accelerators represent impressive laboratories for studying ultra-relativistic particles on the one hand, and accelerated
406 particles themselves are messengers of the astrophysical nature of cosmic accelerators on the other hand. The composition of
407 cosmic rays and their energy spectra help us to unravel the nature of the sources of cosmic rays which remain puzzling in many
408 aspects. Furthermore, some cosmic rays gamma-rays and/or high-energy neutrinos may stem from the decay or annihilation of
409 dark matter. Cosmic rays, can inform us on dark matter accumulated in celestial bodies or in the galactic core and halo.

410 A multi-messenger approach, combining also gravitational waves, has the potential to study these questions, including questions
411 related to gravitational interactions.

412 UNIGE is involved in the long standing very successful AMS/AMS-02, the presently running DAMPE and in the future HERD
413 (DPNC, Wu-group) and EUSO, a space-based detector of UHECR (DA). Switzerland has a long standing interest in high energy
414 γ -ray observations. All groups (about 35 people at UNIGE (DPNC and DA, T. Montaruli, R. Walter, D. della Volpe), ETHZ (A.
415 Biland), UZH (P. Saha) and now also EPFL (E. Charbon)). are currently involved in CTA [?] construction and in the definition of
416 its scientific program, preparing its software and start of data taking of telescopes. The present ‘work horses’ are the first Large
417 Size Telescope (LST-1) of CTA, the two MAGIC telescopes [?], and the small FACT [?] (see Fig. 1). In 2021 the two small size
418 telescopes SST-1M will see first light at Ondrejov Observatory before joining the LHAASO experiment[?]. Several scientists
419 are also participating in LHAASO.



Figure 1: Visible in the photograph, from left to right, the first Large Size Telescope of CTA LST-1 of 24 m mirror-diameter, FACT (4 m) and the MAGIC telescopes (12 m) at Los Roches de Muchachos, La Palma. The LST-1 was inaugurated on Oct. 2018.

420 We briefly describe these experiments with Swiss participation, which study cosmic rays and address the quests indicated above.

421 **From space:**

- 422 ● **AMS-02** : (Alpha Magnetic Spectrometer) is a cosmic ray detector detecting cosmic ray particles and anti-particles, aboard
423 the international space station (ISS). It is taking data since more than 9 years and has e.g. measured in great detail the
424 cosmic ray nuclei, electron, positron and antiproton fluxes in the GeV - TeV range.
- 425 ● **DAMPE** : (DARK Matter Particle Explorer) is a space telescope used for the detection of high energy gamma rays, electrons
426 and cosmic ray ions, to aid in the search for dark matter. It was designed to look for the indirect decay signal of dark matter
427 and for direct cosmic ray measurements in the 1 TeV -100 TeV range. It was launched by the Chinese Space Agency in
428 2015.
- 429 ● **HERD** : (High Energy Cosmic Radiation Detection) facility is a flagship science mission on board Chinas Space Station,
430 planned to be launched around 2025. HERD will extend the direct cosmic ray measurement to the PeV region, allowing
431 the connection to the ground-based observations at the so-called "knee" region. The main science objectives are: detecting
432 dark matter particle, study of cosmic ray flux and composition and high energy gamma-ray observations. HERD will also
433 observe X-rays and gamma-rays which accompany the most energetic events in the Universe and are one of the main in-
434 gredients to understand high-energy accelerators in the Universe. They are very often used in a multi-messenger approach,
435 since being detected with higher statistics than ultra-high energy cosmic rays, neutrinos, and gravitational waves, together
436 with lower energy photons detected from satellites, they drive the searched for cosmic high energy sources.
437 Thanks to the Fiber Tracker (FIT) developed by UNIGE and EPFL, HERD also serves as a sub-GeV gamma-ray observa-
438 tory with unprecedented imaging capability towards the Galactic center.

- 439 • **EUSO:** (Extreme Universe Space Observatory) is a 2.5 meter-apertures wide field of view fluorescence telescope for
440 detection of traces of UHECR in the atmosphere. It is planned for installation at the Russian segment of the ISS around
441 2024. EUSO's goal is to add the ultra-high-energy channel for the multi-messenger astronomy by building the first all-sky
442 high statistics map of arrival directions of UHECR.
- 443 • **POLAR-2:** POLAR-2 is a compact detector of soft gamma-rays with energies below 1 MeV with the goal to measure the
444 polarisation of photons from gamma ray bursts. This is necessary to discriminate between the different physical models
445 which have been put forward to explain the mechanism leading to these most luminous single events in the Universe. It is
446 right now constructed in Geneva (DPNC and DA) to be put on the Chinese space station in 2024.
- 447 • **eXTP:** (enhanced X-ray Timing and Polarimetry mission) is designed to study the state of matter under extreme conditions
448 of density, gravity and magnetism. Primary goals are the determination of the equation of state of matter at supra-nuclear
449 density, the measurement of QED effects in highly magnetized stars, and the study of accretion in the strong-field regime
450 of gravity. Primary targets include isolated and binary neutron stars, strong magnetic field systems like magnetars, and
451 stellar-mass and supermassive black holes. The mission carries a unique and unprecedented suite of scientific instruments
452 enabling for the first time ever simultaneous spectral-timing-polarimetry studies of cosmic sources in the energy range
453 from 0.5-30 keV (and beyond). The mission adoption is expected in 2021 and the planned launch date is 2027.

454 **From the ground:**

- 455 • **IceCube:** Is a 1 km³ instrumented volume of ice between 1.5 to 2.5 km below the South Pole surface detecting high
456 energy neutrinos [?]. At the surface IceTop is detecting the electromagnetic component of cosmic ray showers which
457 inform on the cosmic ray composition in the knee region. The in-ice detector has been completed in 2011. It consists of
458 5'600 photomultipliers attached to 86 strings lowered into the ice which is used as a shield, detecting the Cherenkov light
459 produced by charged particles induced by neutrinos. IceCube is undergoing the Phase 1 upgrade, to increase the size of
460 its dense core detector with additional 7 strings holding 700 new and enhanced optical modules (already financed by the
461 US-NSF). This detector, called PINGU, lowers the energy threshold of IceCube to detect neutrinos down to 1 GeV. It aims
462 at the determination of the more precise neutrino oscillation parameters and of the neutrino ordering. The future upgrade
463 will focus on cosmic neutrinos (IceCube Gen-2) with extending the string to cover an area about a factor of 10 larger and
464 with extended veto capabilities at surface. The photosensor array will be complemented by a radio array exploiting the
465 Askaryan effect to detect signals from high energy neutrinos.
- 466 • **LHAASO:** The Large High Altitude Air Shower Observatory, located at 4'400 a.s.l. on the mountains of the Sichuan
467 province in China, is a new generation Extensive Air Shower (EAS) array for cosmic ray detection in the energy range from
468 10¹¹ to 10¹⁸ eV and for gamma-rays above 1 TeV [?]. At relatively low energies, a water pond of 80'000 tons, equipped
469 with photomultipliers (WCDA), detects gamma-rays and cosmic rays. At higher energies about 20 imaging telescopes (the
470 WFCTA array) can be used in Cherenkov mode and above 10¹⁷ eV in fluorescence mode. The extended array (over an
471 area of 1 km-diameter) of electromagnetic particle and muon detectors measures the lateral and longitudinal distribution
472 of extensive air showers. Detecting shower muons makes LHAASO very powerful for identifying the composition of
473 primary particles in the region of the knee and above [?]. LHAASO is extremely interesting for searches of heavy dark
474 matter, because it may have the opportunity to disentangle the diffuse gamma-ray from cosmic ray interactions in the
475 Galaxy. The observatory is under completion, but 50% is already complete and it is taking data since the end of 2019.

476 Some preliminary physics results will be published after summer 2020. The completion of the installation is expected by
477 the end of 2021. Scientists at UNIGE currently involvement in this experiment and are very excited about the rapid progress
478 and prospects. LHAASO is expected to be the most sensitive project to face the open problems in Galactic cosmic ray
479 physics, with unique ability for detecting the cosmic ray sources and heavy dark matter eventually present in the galactic
480 halo. With large field of view and almost 100% duty cycle it has unique potential to detect the PeVatron in the Galaxy that
481 also contribute to the cosmic neutrino flux detected by IceCube.

- 482 • **CTA:** The Cherenkov Telescope Array [?] is the next generation array of Imaging Atmospheric Cherenkov Telescopes
483 (IACTs), now entering the implementation phase. The CTA Key Science Cases concern three major themes: The under-
484 standing of the origin of the cosmic rays; Probing extreme environments, such as neutron stars, black holes and gamma-ray
485 bursters; Exploring frontiers in physics, such as the nature of dark matter, axions and their interplay with magnetic fields,
486 quantum gravitational effects in photon propagation.

487 The CTAO (CTA Observatory) will be composed of two arrays: one in the Northern hemisphere located at La Palma, and
488 one in the Southern hemispheres at the ESO site of Paranal in Chile, both at about 2000 m. The CTAO also comprises a
489 Science Data Management Centre in Desy-Zeuthen and the Headquarter of the CTAO Project in Bologna. CTA will be an
490 international open access observatory governed by the CTAO ERIC which will become operational in mid-2021. A part of
491 the observation time will be reserved for the CTA Consortium exploitation of the Key Science Cases [?].

492 Its two telescope arrays comprise more than 100 telescopes of three different mirror sizes to cover an energy range from
493 about 20 GeV to 300 TeV. Three different telescopes with mirror sizes of 24m diameter (LSTs - Large Size Telescopes),
494 12m diameter (MSTs - Middle Size Telescopes) and 4 m (SSTs - Small Size Telescopes) will be deployed in order to span
495 the wide energy range indicated above in the Southern Array. The first LST, LST-1 has already been built close to the
496 MAGIC site with significant ETHZ and UNIGE involvement. Subsequently 3 more LSTs will be added in the coming
497 5 years and other MSTs in the Northern CTA site. The preparation of the Southern site is also starting in Chile. The
498 telescope consortia are operative and have mandates to deliver the telescopes as in-kind contribution (IKC) to CTAO that
499 will operate them for 30 years. Swiss groups are involved in the LST Consortium. Some CTA software work packages,
500 such as for array control (ACADA) and data analysis and calibration (DPPS), are defined or being defined.

- 501 • **MAGIC:** (Major Atmospheric Gamma Imaging Cherenkov Telescopes) is a system of two Imaging Atmospheric Cherenkov
502 telescopes situated at the Roque de los Muchachos Observatory on La Palma, at about 2200 m above sea level. MAGIC
503 detects particle showers released by gamma rays, using their Cherenkov radiation. The telescopes have two 17 meter-
504 diameter reflectors, now surpassed in size by the first LST of CTA of 24 m and by the H.E.S.S. II telescope. With H.E.S.S.
505 and VERITAS, MAGIC has opened the gamma-ray field from ground to Big Science, from the observation of pulsation of
506 pulsars above 100 GeV to the extension of the gamma-ray burst spectrum up to above 300 GeV.
- 507 • **FACT:** (First g-App Cherenkov Telescope) is a small 4 m Cherenkov telescope pioneering the usage of SiPM (Silicon
508 Photomultipliers, also called GAPD: Geiger mode avalanche photodiodes) and performing the first unbiased monitoring
509 of variable extragalactic objects at energies > 1 TeV. It is located at La Palma, situated between MAGIC and LST-1.
- 510 • **PORTAL:** Ideas on the post-CTA future are also being considered. One is called 'PORTAL'. Fundamental laws of optics
511 limit the size and field of view of Cherenkov telescopes. Neutrino as well as gravitational wave detectors have intrinsically
512 limited angular resolution, making follow-up observations rather difficult. In his PhD thesis, S.A. Mueller from ETHZ [?]

513 developed the ingenious idea of replacing normal cameras by light-field sensors, allowing to overcome optical restrictions.
514 It seems possible to construct instruments that could reach an energy threshold an order of magnitude lower than possible
515 with CTA array as well as a field of view exceeding 30 degrees, and a preliminary system design named PORTAL was
516 produced together with the civil engineering department of ETHZ. Such a γ -ray telescope would be the ideal companion
517 to work together with future neutrino and gravitational wave telescopes. In addition, at energies below ≈ 10 GeV the
518 geomagnetic field repels charged cosmic ray particles, reducing the background signal by at least a factor 100.

519 7.5 Dark matter, direct detection

520 While there is ample evidence for the existence of dark matter via its gravitational interaction with luminous matter, its nature at
521 the microscopic level remains unknown. The goal to discover non-gravitational interactions of dark matter is a global effort, and
522 Swiss scientists are at the forefront of this search.

523 The Swiss group led by Laura Baudis (UZH) is a founding and leading member of the world's most sensitive direct dark matter
524 detection program based on liquid xenon (LXe) time projection chambers. The XENON1T phase, which used a total of 3.2 t
525 of LXe, acquired data at the Laboratori Nazionali del Gran Sasso (LNGS) of INFN in Italy until December 2018. Currently its
526 successor, XENONnT, with a total of 8.4 t of LXe, is under commissioning at LNGS and will start a first science run by the end
527 of 2020. It will be followed by DARWIN, a 50 t LXe observatory in astroparticle physics, with ultimate sensitivity to particle
528 dark matter over a wide mass range and with a rich program in neutrino physics, and other rare-event searches.

- 529 • **XENON1T:** The XENON1T experiment [?] acquired data at LNGS from 2016-2018 and set the world's best upper limit for
530 the WIMP-nucleon elastic scattering cross-section for DM masses above 85 MeV [?, ?, ?]. as well as on the DM-electron
531 scattering cross section for masses above 30 MeV [?, ?]. XENON1T also sets the most restrictive direct constraints to
532 date on pseudoscalar and vector bosonic dark matter for masses between 1 keV and 210 keV [?]. Furthermore, XENON1T
533 observed for the first time the two neutrino double electron capture of ^{124}Xe , with a half-life of $T_{1/2} = 1.8 \times 10^{22}$ y, the
534 longest half-life ever measured directly [?]. The detector's core - the TPC - with leading contributions from the Swiss
535 group, is back at UZH and will be part of the new Exploratorium on Campus Irchel.
- 536 • **XENONnT:** The XENONnT experiment was installed at LNGS in early 2020, and is under commissioning as of summer
537 2020. With a fiducial LXe mass of 4 t and an exposure of 20 t y, the expected sensitivity to spin-independent interactions
538 will reach a cross-section of $1.4 \times 10^{-48} \text{cm}^2$ for a 50 GeV/c² mass WIMP, a factor of 10 improvement compared to
539 XENON1T. XENONnT will also search for the neutrinoless double beta decay of ^{136}Xe , and will be able to probe the
540 excess of events observed by XENON1T in the (1,7) keV region [?] within a few months of data and distinguish between,
541 e.g., a ^3H component and a solar axion signal. As for XENON1T, the Swiss group has leading contributions to the design
542 and construction of the TPC, to the characterisation of the 3-inch photosensors in LXe [?], as well as the design and
543 construction of their low-background voltage dividers. The group is responsible for the signal transfer between TPC and
544 DAQ on the xenon side, for the low-noise, dual-channel amplification of the signals, for the design and construction of the
545 light calibration system, as well as for the material screening with a dedicated high-purity germanium detector at LNGS [?].
- 546 • **DARWIN:** DARWIN is the ultimate LXe-based DM detector which will explore the full WIMP parameter space (with
547 an exposure of 200 tons \times years) above the so called 'neutrino floor' where neutrinos will start to dominate the signal [?].
548 The Swiss PI (Laura Baudis) is the founder and spokesperson. The project is presently in R&D and design phase, with a

549 CDR planned for 2022 (an invitation was issued by LNGS, after a successful LoI submission and review) and a TDR in
550 2024. The earliest data taking would start in 2026/27. The Swiss group constructs a large prototype (hosting a 2.6 m tall
551 LXe-TPC) at UZH, with funding from an ERC Advanced Grant. DARWIN will have a similar reach to dedicated future
552 neutrinoless double beta decay experiments on the decay of ^{136}Xe [?], and would provide a high-statistics observation
553 of pp neutrinos from the Sun [?]. It would also search for solar axions, galactic ALPs and dark photons, a magnetic
554 moment of the neutrino, and measure coherent elastic neutrino nucleus scatters from ^8B solar neutrinos and eventually
555 from supernovae.

- 556 • **DAMIC, DAMIC-SNOLAB** The DAMIC (Dark Matter in CCDs) experiment has sensitivity to many orders of magnitude
557 in DM mass for various assumptions on how the hidden photon relates to the dark matter and the formation of dark matter
558 in the early universe. The Swiss PI (Ben Kilminster, UZH) is a founding member of DAMIC (Dark Matter in CCDs) since
559 its prototype phase at Fermilab in 2008, which established world-leading search results for weakly interacting, low-mass
560 dark matter. The DAMIC-SNOLAB experiment increased the mass, and decreased backgrounds, and has concluded in
561 2019, after producing several world-leading results, also extending beyond searches for WIMP dark matter, to include
562 hidden-photon DM. Several final publications are soon to be submitted.
- 563 • **DAMIC-M, OSCURA** DAMIC-M is an approved and funded international experiment at Modane Underground laboratory
564 that is set to begin in 2024. It has a mass ten times bigger, a background rate 10 times smaller, and an energy threshold 10
565 times smaller than DAMIC-SNOLAB, this made possible using skipper electronics readout. DAMIC-M has a broad reach
566 to probe 10 orders of magnitude in DM mass over a range of theoretical scenarios. The Swiss group at UZH holds major
567 roles in low-background materials mechanics, electronics, the DCS system. Studies are being done for a new experiment,
568 10 times bigger than DAMIC-M, called OSCURA, with an even lower background rate and energy threshold. Figure 2
569 shows a comparison of the leading CCD experiments searching for dark matter.

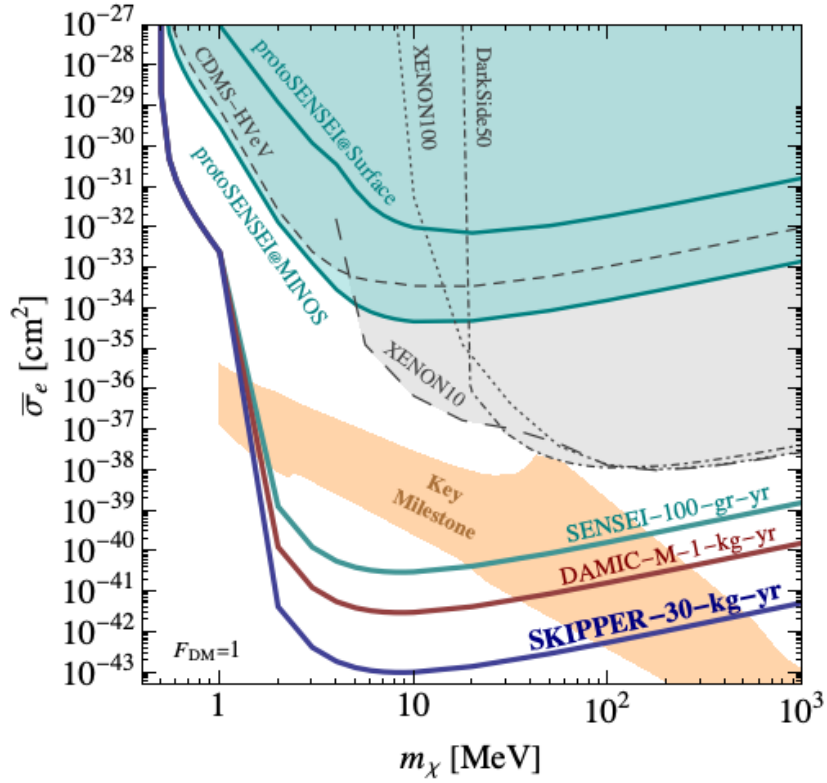


Figure 2: Future sensitivity of CCD experiments for the scattering of dark matter on electrons via a heavy hidden photon mediator. DAMIC-M is shown with its successor, a 30-kg-year proposed skipper CCD experiment named OSCURA. A key milestone is shown indicating the cross-section expected if the entire DM density observed today is due to hidden-sector DM. [?]

570 **7.6 Theoretical physics**

571 Swiss researchers are actively involved in all the main lines of theoretical research mentioned before, namely i) precise SM
 572 physics; ii) BSM phenomenology; iii) Cosmology, astroparticle, gravitational physics; iv) formal theory developments; with
 573 complementary expertise at different universities and research institutes, as qualitatively outlined in Table 1. The Swiss research
 574 in theoretical particle physics is of very high quality in all these different directions. Just to mention a representative figure of
 575 merit: in the last six years Swiss researchers have been awarded four ERC Advanced Grants and four ERC Consolidator Grants
 576 in the area of theoretical particle physics. This is a remarkably high figure given the relatively small community and the highly
 577 competitive nature of these grants.

578 **8 Major successes (2017-2020)**

579 [Main Editor: perhaps Gino] [(1-3 pages) – If relevant, one could identify in this section major recent (within the current ERI
 580 4-year period) Swiss scientific achievements (are there any NCCR, NRP, awards, special EU funding, etc?). It can also be the

Institution	Main research areas
EPFL	(II) High-energy BSM phenomenology, model-building. (III) Cosmology, astroparticle physics, hidden sectors. (IV) Formal aspects of QFT
ETH	(I) High-precision perturbative QCD, collider phenomenology. (III) Physics of gravitational waves. (IV) String theory and formal aspects of QFT.
PSI	(I) Precision low-energy physics, collider phenomenology. (II) BSM phenomenology at low- and high-energies, model building.
University of Basel	(II) Neutrino physics, high-energy BSM phenomenology. (III) Cosmology, astroparticle physics.
University of Bern	(I) Precision low-energy physics, lattice QCD, collider phenomenology. (II) BSM phenomenology at low- and high-energies, model-building. (III) Cosmology, astroparticle physics. (IV) String theory and formal aspects of QFT.
University of Geneva	(II) High-energy BSM phenomenology, model-building. (III) Cosmology, astroparticle physics, physics of GW. (IV) String theory and formal aspects of QFT
University of Zurich	(I) High-precision perturbative QCD, simulation tools for colliders, precision flavour physics. (II) BSM phenomenology at low- and high-energies, model-building. (III) Physics of gravitational waves.

Table 1: Overview of the research activities in theoretical physics in Switzerland.

581 [building of a new infrastructure, the Swiss participation to an international organisation, etc.\]](#)

582 [\[The following split in sections is intended as an intermediate step for collection of the necessary material. The main editor of](#)
583 [the chapter will merge accordingly after the material is available.\]](#)

584 **Input from accelerator research**

585 [\[Editor: Lenny\]](#)

586 **8.1 Major successes 2017-2020**

587 Thanks to the financial support provided by SERI and the matching funds in form of manpower and hardware from the par-
588 ticipating institutes the developments in the CHART program could achieve significant results. The activities of CHART were
589 concentrated on such research topics as high field superconducting magnet developments for FCC, FCC beam dynamic studies
590 and novel methods of laser acceleration. The beam dynamics studies have been concentrated on both future options of the hadron
591 colliders: FCC-hh and HE-LHC. Studies have been performed to define the operational scenario to maximize the luminosity
592 reach of such machines. This is obtained by ensuring particle long-term stability to avoid beam parameters degradation i.e. parti-

593 cle losses or coherent instabilities. Studies have covered both hadron colliders options in the FCC design and have been focused
594 on the top energy setting of 100 TeV and 26 TeV center of mass energy for the FCC-hh and HE-LHC, respectively. Stability is
595 ensured before (called single beam stability) and during collisions (two beams effects). The effect of electron clouds and miti-
596 gation techniques to suppress these collective effects have also been addressed. For both collider options an operational scenario
597 to reach the luminosity goals has been proposed and accepted as the baseline scenario as reported in the final conceptual design
598 reports. Studies have covered several aspects of the design and, where possible, were supported by experimental benchmarking
599 at the present Large Hadron Collider (LHC). The main subjects of studies are listed below with the major achievements:

- 600 • Single beam stability: impedance, Landau damping and electron cloud studies.
- 601 • Two beams dynamics: beam-beam interaction effects and long-term stability.
- 602 • Experimental development of diagnostics for Landau damping studies (i.e. Beam Transfer Function measurements).
- 603 • Feasibility studies of new Landau damping devices: electron lenses and Radio-Frequency quadrupoles.
- 604 • Explorative studies of collider performance optimization using machine learning techniques.
- 605 • Electron cloud studies and mitigation strategies.

606 All of the results have been documented in the FCC Conceptual Design Reports. Several open questions remain to be addressed
607 with the aim to ensure a much higher luminosity reach and to still prove the feasibility of a 100 TeV center of mass energy
608 collider. The main subjects arising are:

- 609 • Alternative crossing schemes (flat optics versus round) with inclusion of relevant effects (mainly synchrotron radiation)
- 610 • Full integration of magnets multiple errors in lattice design using the compensation of beam-beam by octupole magnets.
611 Explore the possibility to use surrogate models of the collider for design optimization with the aim to explore a larger
612 parameters space.
- 613 • Explore the possibility of high precision luminosity measurements in the presence of beam-beam interactions and suggest
614 possible correction methods
- 615 • Continue the study of alternative methods/devices for Landau damping (electron lenses and Radio Frequency Quadrupoles)
616 and keep developing diagnostic devices to benchmark with models expectation in the LHC
- 617 • Different filling schemes (i.e. 5 ns bunch spacing) need to be further explored.

618 CHART Phase-1 engaged in superconducting-magnet R&D at Paul Scherrer Institute (PSI), ETHZ Soft Materials Group (SMG)
619 and EPFL Swiss Plasma Center (SPC). The results included the design of an optimized 16 T Canted Cosine Theta (CCT) dipole
620 magnet, as an option for the FCC hadron collider main magnet, the development (design and prototype) of a high-field dipole
621 magnet with CCT technology (magnet test by the end of this year) and development of reaction resistant splicing techniques for
622 Nb₃Sn based accelerator magnets. The PSI HIPA facility generates a high intensity proton beam with a record beam power of up
623 to 1.4MW for the production of high intensity muon and neutron beams. The acceleration of such high average beam intensities
624 is possible due to a reduction of the unwanted proton beam losses to a relative level of 10E-4. Also the energy efficiency of the
625 facility for conversion of grid power to beam power is outstanding in comparison to other high intensity accelerators and reaches
626 18% [?].

627 Since 2018 the SwissFEL free electron laser is operated at PSI in regular user mode. A key component of the facility is a high
628 brightness 6 GeV electron Linac that utilizes innovative C-band accelerator technology. Normal conducting copper structures in
629 this Linac are realized as brazed stacks of precision manufactured cups. Due to the unprecedented low machining tolerances and
630 the high surface quality of the structures, the usual post-production mechanical tuning is not required and the structures show
631 a good breakdown behavior. Although the primary aim for developing this technology was the generation of high quality and
632 bright electron beams for an FEL, it can also be utilized as injector accelerator for the e^+/e^- version of a future circular collider.
633 Within the CHART project a study on developing an injector concept for FCC-ee has been launched and more details are given
634 in a later section on CHART.

635 **Input from Pillar 1**

636 [Editor: Anna]

637 **8.2 High energy**

638 The ATLAS, CMS and LHCb experiments have achieved major advancements in the understanding of the SM and the exploration
639 of BSM phenomena.

640 [TODO: Here we will need to add links to LHCb, ATLAS and CMS publication pages.]

641 In the recent years, the highest priority pursuit has been the Higgs boson production and decay. Swiss physicists first helped
642 re-establish the observation of the Higgs boson and its expected higher cross section, with primary contributions in the $H \rightarrow \gamma\gamma$
643 channel.

644 Swiss physicists were able to observe both the $H \rightarrow \tau\tau$ and $H \rightarrow b\bar{b}$ processes, establishing definitively the coupling of the
645 Higgs boson to fermions. While previous data periods had established the primary Higgs-boson production modes of gluon
646 fusion, associated production, and vector boson fusion, the 13-TeV data allowed Swiss physicists to establish the $t\bar{t}H$ process,
647 noteworthy due to its direct measurement of the only strong coupling of the Higgs boson, which is to the top quark.

648 The increased luminosity and larger cross-sections have provided the ATLAS and CMS collaborations with a wealth of Higgs
649 bosons to study, allowing them to hone in on classes of events that are either more sensitive to potential BSM physics, or that are
650 not theoretically well-predicted. Swiss physicists measured the differential production of $H \rightarrow \gamma\gamma$ and $H \rightarrow b\bar{b}$, binning events
651 according to momentum of the Higgs boson, and associated jet multiplicity and momenta. Combined Higgs boson measurements,
652 making use of multiple production and decay modes, have moved towards more exclusive topologies, such that the phase space
653 relevant for specific comparisons to theoretical predictions or new-physics models are isolated. For instance, cross-sections are
654 measured separately in events according to jet multiplicity and Higgs momentum, allowing for precise comparisons with both
655 SM and BSM predictions.

656 Beyond the exploration of Higgs physics properties, many other, previously not-well-constrained SM properties, in particular
657 those related to the top quark, are now being measured with unprecedented precision. The experimental focus is on accurately
658 measuring the known interactions and establishing rare processes, while looking for indirect effects in the interactions of known
659 particles. During the Run 2 of the LHC, new rare interactions such as the production of Higgs bosons in association with top
660 quarks, $t\bar{t}H$ and the production of 4-top quarks, $t\bar{t}t\bar{t}$, have been established. One of the main uncertainties in these measurement

661 rises from the modeling of the production of a pair of top quarks in association with jets, a process that also constitutes a common
662 background to NP searches. Swiss scientists have played a leading role in improving the understanding of the $t\bar{t}$ production
663 associated with jets; at ATLAS, they have performed extended measurements of $t\bar{t}$ +jets production, while at CMS, they have
664 studied in detail the case where the extra jets originate from b -quarks, with the goal of reducing the uncertainty in future $t\bar{t}H$
665 measurements and other similar rare processes. Top-quark events are crucial for developing algorithms for identifying b -jets and
666 evaluating the performance of new reconstruction techniques, such as taggers of boosted topologies. Swiss scientists have also
667 led developments in this direction.

668 The steady increase in luminosity and energy in Run 2 is being fully exploited by a large repertoire of well-motivated BSM
669 searches. On the SUSY side this includes searches with missing transverse momentum, and use of powerful variables such as
670 MT_2 , in the all-hadronic, lepton+jets and multilepton final state, third-generation squark searches, as well as electroweakino
671 searches and their combinations. It is complemented by searches for dark-matter candidates and their mediator(s) motivated by
672 several exotic theoretical models. The problems of neutrino masses and matter-antimatter asymmetry are studied by searches for
673 heavy neutral leptons (HNLs) in prompt and displaced leptonic decays of W bosons.

674 The LHCb experiment has achieved major results in physics including the discovery of the very rare $B_s \rightarrow \mu\mu$ decay and the
675 discovery of CP violation in charm decays. For the latter, Tatsuya Nakada from EPFL was awarded the 2019 Enrico Fermi Prize.

676 The successes in physics pursuits would not have been possible without excellent functioning of the detectors, a common success
677 among all LHC experiments. More than 95% of the millions of channels the detectors are made of have been operational at any
678 time during the Run 2 data taking of the LHC. This can be considered a big achievement for the Swiss groups when accounting
679 for the fact that the number of channels is dominated by those present in tracking detectors, whose design, construction and
680 operation have been led by teams within Switzerland. Excellent is also the understanding of the detectors, a fact that has led
681 to novelties in triggering, reconstruction and data analysis techniques. These efforts constitute the continuous focus of Swiss
682 physicists and are documented in the previous chapter. The developments are staged and exploited in major measurements and
683 searches that are described above. Looking in the future, Swiss groups are currently participating to HL-LHC detector and trigger
684 & data acquisition projects that have been approved and are steadily proceeding towards realisation with significant efforts from
685 all Swiss institutes.

686 On the computing side, all LHC experiments are in the process of developing the infrastructure to be able to transparently and
687 interchangeably exploit all available resources in an optimal way. An example is the transfer of CPU based reconstruction
688 algorithms to equivalent parallelised versions to be run on GPUs.

689 In the search for DS and possible dark matter candidates, NA64 set the most stringent limit for light thermal dark matter below
690 0.1 GeV [?, ?]. It also reported the first limits on a new vector boson $X - e^-$ excluding part of the parameter space suggested by
691 the so called X17 anomaly. New bounds could also be set on the mixing strength of photons with dark photons. The latest NA64
692 results set new limits on the scalar/axion like particles (ALP) photon coupling strength [?], in a phase-space that closes the gap
693 in the ALP parameter space between previous fixed target and collider experiments.

694 Extending the pursuit for BSM phenomena beyond what can be done at the LHC, attempting to cover unexplored parts of the
695 parameter phase-space, which cannot be accessed by NA64, a significant recent achievement has been the approval by the CERN

696 Research Board of the FASER experiment. The experiment has been primarily funded by the US Heising and Simons-Heising
697 foundations and SNF supports it with project funding. The construction of the experiment is progressing in a speedy way and the
698 experiment will collect data in Run 3.

699 8.3 Low energy

700 For low energy experiments, major results were obtained with high sensitivity searches for BSM physics as well as for high
701 precision measurements of SM benchmarks and fundamental constants. Four ERC grants were recently (2016-2018) granted,
702 two in neutron EDM searches and two in exotic atom laser spectroscopy with muons, which reflects the considerable progress
703 and impact made over several years now.

704 The nEDM collaboration at PSI has in 2020 released the most stringent limit on the permanent electric dipole moment of the
705 neutron, $d_n < 1.8 \times 10^{-26}$ ecm (90% C.L.) [?]. with direct impact on theories explaining the matter-antimatter asymmetry of
706 the universe. The nEDM data was also analyzed for an oscillating neutron electric dipole moment which could be induced by
707 coupling of ultralight axion-like particles (ALPs) to gluons. Assuming that these ALPs would constitute the dark matter in the
708 universe, first laboratory limits on ALP-gluon coupling for ultralight ALP masses were established [?].

709 The MEG collaboration at PSI established the limit for the lepton flavor violating decay $\mu^+ \rightarrow e^+\gamma$, which is the most stringent
710 upper limit on any branching ratio in physics $\mathcal{B}(\mu^+ \rightarrow e^+\gamma) < 4.2 \times 10^{-13}$ [?]. From their data set, MEG has recently provided
711 the most stringent limits on hypothetical light, neutral particles X in the mass range between 20 and 40 MeV/c² for lifetimes of
712 less than 40 ps and decaying to two photons $\mu^+ \rightarrow e^+X, X \rightarrow \gamma\gamma$ [?].

713 After their successful measurements of the 2S-2P Lambshift in muonic hydrogen and muonic deuterium (2010-16), the CREMA
714 collaboration at PSI has recently measured the 2S-2P Lambshift in the muonic helium isotopes 3 and 4. Besides the extraction of
715 benchmark charge radii in light and calculable systems, sensitive tests of QED and independent determinations of the Rydberg
716 constant become possible [?]. To this aim the Mu-MASS collaboration demonstrated the creation of a muonium 2S metastable
717 beam [?]. The piHe collaboration succeeded in a first time ever laser spectroscopy of a pionic atom [?] further extending the
718 reach of precision optical methods into the realm of particle physics. The muX collaboration succeeded in demonstrating the
719 ability to form heavy muonic atoms ²⁴⁸Cm and ²²⁶Ra using only microgram quantities of target material enabling, e.g., new
720 symmetry tests in heavy nuclear systems with large enhancement effects.

721 Towards the development of a new High intensity Muon Beam (HiMB) first improvements to surface muon production were
722 implemented with a new design of the production target resulting in 40-50% improved muon yield for the same proton beam
723 power, benefiting many muon experiments. The muCool project at PSI succeeded in demonstrating transverse phase space
724 cooling of a positive muon distribution [?]. With the previously demonstrated longitudinal cooling, this confirms the promise
725 of improved phase space quality by ten orders of magnitude at the cost of only three orders of magnitude in muon intensity,
726 translating into improved muon beam brightness by seven orders of magnitude with far-reaching consequences for experiments
727 in fundamental particle physics and beyond.

728 At ETHZ, on a table top experiment, positronium is being used to search for the specific case of massless dark photons which
729 cannot be probed in fixed target or accelerator experiments. Recently the experiment has reached a sensitivity comparable with

730 cosmological bounds [?].

731 **Input from Pillar 2**

732 [Editor: Michele]

733 **Input from Pillar 3**

734 [Editor: Ruth]

735 **8.4 Cosmic rays, high energy neutrinos, X- and γ -rays**

736 Despite the small Swiss groups working in IceCube and MAGIC, some of the relevant scientific outcomes would not have been
737 possible without their contribution. Let us mention some recent highlights:

738 By today about 7.5 yrs of data have been published by IceCube and the significance of the cosmic neutrino flux rose to well above
739 5σ s [?, ?]. The energy spectrum of its muonic component can be fitted by $E^{-2.13}$. The PeV muon-neutrino event, IC-170922A,
740 with energy of about 4.5 PeV has triggered an alert in the astronomer's network[?]. A blazar, TXS 0506+056, located inside the
741 directional uncertainty contour of IC-170922A, was discovered in a flaring state by Fermi-LAT. This strong hint on the discovery
742 of the first cosmic high-energy neutrino source [?] was supported by MAGIC observations revealing also flaring activity between
743 100 – 400 GeV [?]. Enhanced emission was confirmed by VERITAS [?].

744 Lately, MAGIC performed the first detection of very high energy gamma-ray emission from a gamma-ray burst, GRB 190114C,
745 from ground [?, ?, ?].

746 More recently, a new probable neutrino source in a close-by starburst galaxy was discovered by UNIGE. It was published in [?]
747 and awarded the Prix Wurth for the best doctoral thesis of the Science Faculty of 2019.

748 A time-dependent search for coincident IceCube neutrino events with two extremely high-energy neutrino events seen by the
749 ANITA polar balloon flight was recently published [?] and appeared in many journals are possible hint of new physics in extreme
750 energy domain.

751 A multi-messenger group of UHECRs (P. Auger and Telescope Array) and neutrinos (ANTARES, KM3NeT and IceCube), led
752 by UNIGE, looks for common sources of UHECRs and neutrinos [?].

753 Of great importance for the construction of LST telescopes of CTA is the recent detection of the pulsations of the Crab Nebula
754 pulsar at energies of the order of 100 GeV [?].

755 UNIGE and ETHZ were leading institutes for development and the construction of the AMS-02 Silicon Tracker, key to the
756 unprecedented precision of the AMS-02 mission launched in 2011. UNIGE now contributes significantly to AMS-02 data
757 analysis, in particular in tracker charge calibration and most precise flux measurements of heavy nuclei (eg. He, Li, O, Si, Mg)
758 in the GeV - TeV range [?].

759 UNIGE proposed, designed and led the construction of the Silicon-Tungsten Tracker (STK) of the DAMPE mission that is key
760 to the tracking and photon detection capability of DAMPE, launched in 2015. UNIGE contributes significantly to DAMPE
761 operation, data processing data analysis, in particular in tracker calibration and alignment, track reconstruction, MC simulation
762 and the most precise flux measurements of cosmic electron, proton and Helium in the TeV -100 TeV range [?, ?].

763 UNIGE (DPNC and DA) and PSI were leading institutes for development and the construction of POLAR, the first large dedicated

764 Gamma-Ray Burst polarimeter launched in 2016. UNIGE (DPNC and DA) led the POLAR data processing and data analysis,
765 including the high statistic polarization measurement of 5 GRBs, including a time resolved study of one GRB [?].

766 **8.5 Dark matter, direct detection**

767 XENON1T has the leading sensitivity to light dark matter (LDM) in the mass ranges 3-6 GeV for DM-nucleon scattering and
768 above 30 MeV for DM-electron scattering, and in the mass range 0.2-1 keV for the absorption of dark photons and ALPs (axion
769 like particles) [?].

770 XENON1T has improved sensitivity to LDM via electron recoil signals induced by the Migdal effect and bremsstrahlung. It
771 actually has presently the best sensitivity for DM masses between 85 MeV and 2 GeV [?].

772 XENON1T has observed two-neutrino double electron capture events in ^{124}Xe , with $T_{1/2} = 1.8 \times 10^{22}$ y and it has the lowest
773 background ever reached in a direct detection experiment [?].

774 First results of XENON1T on the scalar WIMP-pion coupling have been published [?].

775 DARWIN: a detailed study of the sensitivity to the neutrinoless double beta decay of ^{136}Xe ; competitive to dedicated double beta
776 experiments without additional costs are possible [?].

777 DAMIC at SNOLAB has been collecting data from 2017 to 2019 has produced the worlds best sensitivity for electronic scattering
778 of dark matter and hidden-photon dark matter in some mass ranges [?].

779 **8.6 Theoretical physics**

780 [Editor: Gino] In the following we illustrate in more detail the main motivations the various activities in theoretical physics in
781 Switzerland and briefly outline recent highlights and future prospects.

782 **QCD at colliders.** Theoretical predictions based on the SM are a fundamental ingredient for the interpretation of collider data.
783 The vast majority of experimental analyses make use of perturbative predictions at parton level or in combination with parton
784 showers. Such predictions are relevant to test the SM but they are also a crucial ingredient in the experimental measurements,
785 both for the description of acceptance efficiencies and for the modelling of backgrounds in SM measurements and BSM searches.
786 As a result of the continuously growing precision of the experimental data, and in view of the expected improvements that will
787 be achieved with the HL phase of the LHC, an increasing number of analyses is going to be limited by theoretical uncertainties.

788 Perturbative calculations including NLO corrections in both the QCD and EW couplings are nowadays supported by automated
789 tools, and their systematic application to hadron-collider studies is mandatory in order to reach a precision of $O(10\%)$. To reach
790 percent-level accuracy NNLO corrections in the QCD coupling are required. For selected benchmark processes, even N³LO
791 accuracy may be required.

792 Several NNLO QCD calculations have been completed in recent years for relatively simple hadron-collider processes. The range
793 of processes for which NNLO accuracy can be reached is currently limited by the difficulty in computing two-loop amplitudes
794 in process involving more the four external particles. The precise modelling of complex processes requires to go beyond this
795 limitation. The computation of the relevant loop amplitudes is a formidable challenge from both the algebraic and analytic
796 viewpoint, and the relevant integrals often involve genuinely new classes of functions. At present several directions to break
797 this bottleneck are explored, including the clever use of unitarity, new geometric approaches, and fully numerical techniques.

798 In parallel, an effort is ongoing to improve available methods to handle and cancel infrared singularities and to extend them to
799 more complex processes. This step is also essential to ultimately build flexible tools that are able to fully deploy the achieved
800 theoretical precision into experimental analyses.

801 Members of the particle theory groups at ETH and the University of Zurich work at the forefront of all these activities. Highlights
802 of their recent research activity in this area includes the first ever N^3 LO calculation for a collider process, namely the Higgs
803 production via gluon fusion at the LHC; the completion of fully differential NNLO calculations for a wide class of processes; the
804 development of parton-level Monte Carlo generator for collider processes incorporating NNLO corrections; the development of
805 new techniques for analytical and numerical calculations of multi-loop amplitudes.

806 Related activities are carried out also at PSI and at the University of Bern. At Bern, in particular, the issue of all-order re-
807 summations of soft-and collinear parton emissions, which is particularly relevant for an accurate description jet physics, is
808 investigated via the development of an appropriate effective theory approach.

809 Future efforts in this direction will be targeted towards key processes for the HL-LHC physics program, as well as for future FCC
810 scenarios. The conceptual frontiers to be addressed are high-multiplicity processes where novel approaches to virtual and real
811 radiation corrections are being developed, multi-scale problems combining QCD and electroweak effects or involving top quarks,
812 as well as ultimate precision at N^3 LO for selected benchmark processes.

813 **Precision low-energy physics.** Beside the direct searches for new phenomena performed at high-energy colliders, a comple-
814 mentary way to search for BSM physics is the so-called intensity frontier, namely the search for possible failures in the SM
815 predictions when performing high-precision experiments. Particularly interesting in this respect are high-statistics low-energy
816 experiments testing exact or approximate symmetry properties of SM, such as the absence of flavour-changing processes in the
817 charged-lepton sector, the strong suppression of flavour-changing neutral-current processes in the quark sector, and the approxi-
818 mate matter-antimatter asymmetry in both the quark and the charged-lepton sector.

819 Also in this case an essential ingredient to make progress in the field is the development of accurate SM predictions, a task which
820 is particularly complicated at low-energies due to the phenomenon of quark confinement. A further complication is also the
821 extreme high accuracy required by these experiments, which often involve very different energy scales (e.g. from the 100 GeV
822 of weak interactions down to the 0.5 MeV of the electron mass). A series of effective theory tools have been developed to deal
823 with these problems, as well as methods based on the combination of analytical calculations and experimental data.

824 Swiss researchers at PSI, the University of Bern and the University of Zurich are particularly active on this research line. High-
825 lights of their recent research activity in this area includes precise predictions for the anomalous magnetic moment of the muon;
826 detailed estimates of signals and backgrounds for rare muon experiments at PSI; precise predictions for experiments on rare
827 B meson decays performed at the LHCb experiment at CERN. Future efforts in this direction will be addressed to the new
828 generations of experiments in this field, with special emphasis on planned experiments at PSI, as well as the LHCb upgrades.

829 A special role in this area is played by lattice QCD. The goal here is to overcome the problem of quark confinement through
830 large-scale numerical simulations of strong interactions and determine the properties of hadron physics from first principles.
831 Swiss theoreticians at the University of Bern coordinates the so-called FLAG report, which compiles and critically reviews the
832 results from various lattice collaborations worldwide, and are members of the ETMC (European Twisted Mass Collaboration):
833 one the largest lattice QCD collaborations worldwide.

834 **Model-building and BSM phenomenology.** The phenomenological attempts to build motivated extensions of the SM, with
835 direct implications for current and near-future experimental efforts in particle physics, can be conveniently organised into three
836 distinct but largely complementary research directions which address different aspects of BSM physics.

837 The first direction is focused on understanding the origin of the Fermi scale. The SM Lagrangian contains a single mass pa-
838 rameter, namely the Fermi scale, or the vacuum expectation of the Higgs field. This scale (of the order of 250 GeV) controls
839 the masses of all elementary particles, but is highly unstable with respect to quantum corrections: it would naturally tend to be
840 heavier in presence of heavier degrees of freedom in the theory. Why such scale is much lighter than the fundamental mass scale
841 associated to gravitational interactions (the Plank scale, of the order of 10^{19} GeV) is one of the big open issues in the SM.

842 In the vast majority of proposed BSM extensions, this problem is solved by introducing new degrees of freedom around the TeV
843 scale, whose main purpose is that of screening the Higgs field from its apparent large sensitivity to high energies. On general
844 grounds, this implies new particles in the TeV range. This is why the direct exploration of the TeV energy domain remains a key
845 priority of particle physics.

846 Swiss theory groups at EPFL, PSI, the University of Geneva, and the University of Zurich are working on this front developing
847 explicit models of TeV-scale dynamics and, most important, trying to understand how these models could be detected at present
848 and future high-energy colliders. Highlights of their recent research activity in this area include the development of general
849 effective theories describing SM extensions where the Higgs is a composite particle; the development of experimental techniques
850 to access suppressed BSM effects in high-energy collisions; the development of new theoretical methods to deals with QFT
851 theories in the strong-interaction regime. Future efforts in this direction will be addressed to the HL-LHC physics program but,
852 most important, to evaluate the physics reach of future high-energy facilities.

853 **The flavour puzzle.** The second direction in BSM phenomenology is focused on understanding the origin of fermion masses.
854 Within the SM, the basic constituents of matter are the three families (or three flavours) of quarks and leptons. Each family con-
855 tains four fermions (two quarks and two leptons) with different quantum numbers, which determine completely their properties
856 under strong, weak and electromagnetic interactions. Ordinary matter consists essentially of particles of the first family, while
857 the (unstable) quarks and leptons of the second and third family appear to be identical copies of those in the first family except
858 for their different (heavier) masses. Why we have three almost identical replicas of quarks and leptons, and what the origin of
859 their different masses is, are among the key open issues in the SM. The observed excess of baryons (over anti-baryons) in the
860 Universe, unexplained in the SM and requiring additional sources of CP violation besides that present in the quark mass matrices,
861 is likely to be related to these questions.

862 In many proposed BSM extensions, the flavour puzzle is addressed by a series of new interactions (and new symmetry principles),
863 whose elementary nature manifests itself only at very high energies. The mediators of such new interactions may be too heavy to
864 be directly produced at high-energy colliders. Still, their effect could show up indirectly in deviations from the SM predictions
865 in various rare low-energy processes, such as the decays of the heavy quarks and leptons.

866 Swiss theory groups at PSI and the University of Zurich are particularly active on this front. In the last few years their research ac-
867 tivity has been focused on understanding the interesting phenomenon of the so-called *B*-physics anomalies: a series of deviations
868 from the SM predictions in various rare *B*-meson decays. While the statistical significance of these anomalies is not extremely
869 high yet, the overall picture is quite coherent and might represent a first clue of BSM physics. If confirmed, this BSM physics
870 would certainly have a non-trivial flavour structure. Highlights of the recent research activity in this area include the development

871 of consistent models able to describe these anomalies; the detailed investigation of the implications of these models for future
872 experiments, both at low and at high energies. If the anomalies will persist, future efforts in this direction will further intensified
873 with special emphasis on the LHCb upgrades, the LFV experiments at PSI, and ultimately the flavour-physics program at the
874 FCC.

875 **Dark sectors and neutrino masses.** The SM could be extended not only by the presence of new heavy states, which have
876 not been identified yet because of the energy limitations of existing colliders, but also by new light states, which have not been
877 identified yet because of their weak coupling to ordinary matter, generically denoted as dark sectors. Dark sectors, which are
878 natural candidates to explain the phenomena of dark matter, have received considerable attention in the last few years. While
879 a large fraction of the parameter space of these modes cannot be probed at accelerators, interesting regions of the parameter
880 space give rise to long-lived particles which can be searched for by the existing experiments at hadron colliders, at dedicated
881 fixed target experiments, and also with high-intensity particle beams also at low energies. A partly related issue is the origin of
882 neutrino masses, whose small values naturally point toward the existence of new fundamental scales in the theory and/or new
883 feebly interacting states (such as light quasi-sterile right-handed neutrinos). Moreover, CP violation in the neutrino sector could
884 well be related to the observed matter anti-matter asymmetry in the universe.

885 Swiss theory groups at EPFL and the University of Basel work on both these aspects of BSM physics, building consistent
886 models of neutrino masses, including possible new light exotic states, and analysing their implications for collider experiments
887 taking into account astrophysical observations. Highlights of their recent research activity in this area include the development
888 of consistent modes addressing both the origin of neutrino masses and the problem of dark matter; systematic analyses of the
889 phenomenology of feebly interacting particles at existing and future high-intensity experiments; systematic analyses of the impact
890 of the DUNE experiment in constraining unified models predicting neutrino masses. Future efforts in this direction will continue
891 at the interface of particle-physics experiments and astrophysical observations, with a special emphasis on exploiting the physics
892 reach of the forthcoming long-baseline neutrino experiments, and in planning new dedicated experiments targeting unexplored
893 regions of motivated feebly-interacting exotic sectors.

894 **Formal theory developments.** Many recent developments in formal theory have been concerned with the AdS/CFT corre-
895 spondence that relates superstring theory on a d -dimensional Anti-de Sitter background to the large N limit of a conformal gauge
896 theory living on the $(d - 1)$ -dimensional boundary of AdS. The most promising example is $\mathcal{N} = 4$ super Yang-Mills theory in
897 4 dimensions that is believed to be equivalent to superstring theory on $\text{AdS}_5 \times S^5$. The AdS/CFT correspondence has powerful
898 consequences since the regime where $\mathcal{N} = 4$ SYM is strongly coupled is mapped to the supergravity regime on $\text{AdS}_5 \times S^5$ which
899 is under good quantitative control, thus providing a promising window into strongly coupled QFTs. At the same time, the duality
900 may also give insights into the quantum behaviour of string theory since the regime in which the AdS space is of string size
901 corresponds to (nearly) free super Yang Mills theory.

902 Over the years many precision tests of the duality have been performed, and beautiful confirmation has been found. Swiss theory
903 groups at ETH and Geneva have significantly contributed to these developments. In particular, using techniques of integrability,
904 theorists at ETH have used these ideas to make highly non-trivial predictions about the spectrum of $\mathcal{N} = 4$ super Yang-Mills
905 theory for arbitrary values of the coupling constant; to the extent that they are accessible by perturbation theory, these predictions
906 have been confirmed by explicit computations. More recently, the focus has moved towards identifying correlation functions
907 between the two theories. An independent development carried out at ETH focused on a low-dimensional version of the duality,

908 relating string theory on $\text{AdS}_3 \times \text{S}^3 \times \mathbf{T}^4$ to the $2d$ conformal field theory that is described by the symmetric orbifold of \mathbf{T}^4 . In
909 particular, they managed to prove that the complete spectrum agrees between the two descriptions, and their techniques may be
910 strong enough to lead to a complete proof of the duality in this setting. The AdS/CFT correspondence has also been applied to
911 strongly coupled condensed matter systems.

912 Another important development concerns the so-called modern (or numerical) conformal bootstrap program that was initiated at
913 EPFL. The bootstrap method is based on the insight that one can characterise correlation functions of a (conformal) quantum field
914 theory based on intrinsic consistency conditions, without any direct reference to an underlying action (that may sometimes not
915 even exist). While this general idea had been known for some time, it was only realised recently that, using in particular numerical
916 methods, these constraints are very powerful indeed and allow one to make significant progress for theories that are otherwise
917 out of perturbative control. One highlight has been the high precision determination of various anomalous dimensions for the $3d$
918 Ising model, but these (and related) techniques have now also been successfully applied to $3 + 1$ dimensional QFTs. There has
919 also been interesting progress towards understanding sectors of conformal field theories at large global charge developed at the
920 University of Bern, and the application of resurgence techniques to QFTs carried out at the University of Geneva.

921 **9 The international context**

922 [Main Editor: Guenther] [(2-6 pages) – Explain the main trends and the evolution of research in the field in Europe and in the
923 world. How does Switzerland position itself in this global landscape: are we at the forefront or a small player? Add something
924 on international collaborations: are there many large collaborations or is the research done in smaller groups?]

925 [The following split in sections is intended as an intermediate step for collection of the necessary material. The main editor of
926 the chapter will merge accordingly after the material is available.]

927 **Input from accelerator research**

928 [Editor: Lenny]

929 **Input from Pillar 1**

930 [Editor: Anna]

931 The LHC experiments are composed of international collaborations. The ATLAS collaboration has approximately 5000 members
932 and about 3000 scientific authors affiliated with 182 institutions in 38 countries. CMS has over 4000 particle physicists, engineers,
933 computer scientists, technicians and students from around 200 institutes and universities from more than 40 countries. The
934 LHCb collaboration consists of 1339 members from 83 institutes in 19 countries. The swiss groups in the LHC collaborations
935 work closely with researchers from abroad, both in the context of their physics analysis projects and the detector construction,
936 commissioning and operation. It is interesting to note that within these large collaborations even computing infrastructure is
937 being shared between institutes and countries; for example, Switzerland contributes with standard computing clusters located in
938 Bern for ATLAS and High Performance Computer (HPC) at CSCS for ATLAS, CMS and LHCb.

939 While the energy frontier is currently dominated by the CERN experiments, the intensity frontier in flavour physics is vigorously
940 pursued in Japan, where the energy-asymmetric KEKB electron-positron collider provides beams to the Belle II experiment,

941 which pursues a physics program complementary to the one of LHCb. Even more diverse is the international effort for dark
942 matter searches with dedicated experiments: In addition to NA64 and FASER, other experiments composed of international
943 collaborations, such as MATHUSLA and CODEX-B, have been proposed at CERN.

944 Looking into the low-energy domain, CERN provides the only source for low energy antiprotons and PSI provides the world's
945 highest intensities of low energy pions, muons and ultracold neutrons (UCN). In Europe, other UCN source are located at ILL
946 Grenoble (France) and TRIGA Mainz (Germany). ILL also provides the highest intensity beams of cold neutrons for fundamental
947 physics. Cold neutrons are also available at FRM-2, in Munich (Germany), while at the European Spallation Source (ESS), in
948 Lund (Sweden), at least one fundamental physics beamline should be built. In a global context, more sources for cold and
949 ultracold neutrons with particle physics as part of their program exist, e.g. at LANL (US), SNS (US), NIST (US), TRIUMF
950 (Canada), J-PARC (Japan).

951 Muon beams with different properties than those of PSI are produced at J-PARC (Japan) and FNAL (US). The PSI "continuous
952 wave" muon beams are preferred for coincidence experiments and when high instantaneous rates cause issues. Pulsed beams
953 produced at J-PARC are well-suited for rare event searches with single particle detection, such as $\mu \rightarrow e$ conversion. FNAL
954 produces pulsed muons for dedicated purposes, such as the g-2 experiment. Muons are also available at TRIUMF (Canada)
955 and at RAL (UK), mostly for muon spin spectroscopy and material science, and at lower rates. Some new facilities study the
956 implementation of a muon physics program. The present beams of surface muons at PSI with rates exceeding $10^8/s$ are leading
957 the field. PSI aims to carry forward its leading position in muon beam intensities for the next decades with new high intensity
958 muon beams (HiMB) which could transport on the order of $10^{10}/s$ low energy positive muons to versatile experimental areas.

959 **Input from Pillar 2**

960 [\[Editor: Michele\]](#)

961 **Input from Pillar 3**

962 [\[Editor: Ruth\]](#) All experiments are built within international collaborations. Here we give more details for some of them.
963 DARWIN includes the XENON collaboration with additional groups from Europe and the US. DAMIC-M has collaborators from
964 Europe and the U.S. The future experiment, OSCURA, will unite two international collaborations that search for DM with CCD
965 detectors, DAMIC-M and SENSEI, bringing in additional North and South American institutions. IceCube is a collaboration of
966 about 300 people from 52 institutions in 12 countries. The full Phase 1 upgrade of IceCube has been financed by NSF on
967 June 2019, extending its scientific capabilities to lower energies and thus enabling IceCube to reach neutrino energies that overlap
968 with the energy range of few GeV of smaller existing neutrino detectors worldwide. The seven strings of optical modules of the
969 Phase 1 upgrade are funded by NSF and the second phase of the upgrade foresees the addition of an acoustic array of antennas
970 and further strings (IceCube-Gen2).

971 CTA is a Consortium of about 1500 scientists from more than 200 Institutes in 31 countries all over the world. In order to start
972 construction, CTA defined a baseline asset (e.g. 50 SSTs instead of 70 and no LST at the Southern site instead of 4) and more
973 than 90% of funding are achieved. At La Palma Northern site the first large size telescope (LST-1) is under commissioning and
974 all Swiss groups are members of the LST Consortium. Notice that, as long as LST-1 is not formally approved by CTAO as a CTA
975 element, LST-1 has proprietary access to data and is handled as an individual project. Financial responsibility on its operation is

976 independent from CTAO but relies on the LST Consortium.

977 The MAGIC collaboration encompasses about 150 scientist from 12 countries. While data belong to the collaboration and
978 publications are signed by all collaboration members, proposals for observations can also be submitted by non-members. In
979 addition, non-members can also get access to data as associate scientists. Such non-members have no other duties than keeping
980 information confidential and are included in the author-list if they contribute significantly to a publication. The operation mode
981 of MAGIC beyond 2025 when the CTA array at La Palma will become operational is not yet decided. One likely option is to
982 integrate MAGIC into CTAO operation as an independent sub-array. The small FACT collaboration under leadership of ETHZ
983 consists of scientists from ETHZ, UNIGE and the German universities Dortmund and Wuerzburg (and close association with
984 RWTH Aachen). Its fate beyond 2025 depends on final operation plans of CTAO. One option could be to relocate FACT to the
985 HAWC detector in Mexico to allow hybrid observations.

986 DAMPE is a China-EU collaboration. The EU participation, under the leadership of UNIGE, consisting of 5 institutes from
987 Switzerland and Italy.

988 The POLAR-2 international collaboration, under the leadership of UNIGE, consists of institutes from Switzerland, Germany,
989 Poland and China.

990 The HERD international consortium includes major institutes from China, Italy, Switzerland and Spain.

991 The eXTP international consortium includes major institutions of the Chinese Academy of Sciences and Universities in China,
992 as well as major institutions in several European countries and other International partners.

993 **10 Synergies with other scientific fields**

994 [Main Editor: Mike, Ruth] [(2-6 pages) – Are there synergies with other disciplines (e.g. biology with chemistry, or physics,
995 or medicine, etc.)? Are you benefitting from advances in other fields (e.g. computing, imaging/analysis tools)? Are you using
996 common infrastructures (e.g. SLS at PSI)? Is there transdisciplinary research being pursued?]

997 [The following split in sections is intended as an intermediate step for collection of the necessary material. The main editor of
998 the chapter will merge accordingly after the material is available.]

999 **Input from accelerator research**

1000 [Editor: Lenny]

1001 **10.1 Synergies with other scientific fields**

1002 Cancer treatment using particle beams from an accelerator is an established concept with strong advantages over X-ray treatments
1003 for a range of indications. PSI in Switzerland has pioneered the technology of pencil beam scanning for precision treatments of
1004 deep seated tumors since 1996. In spite of those obvious advantages the technology is expensive, and particularly the treatment
1005 gantries with iron-pole electromagnets are large and heavy, involving cost driving mechanics. Superconducting accelerator
1006 magnet technology, that was developed for particle colliders, provides the potential for reducing the weight of todays gantries
1007 by an order of magnitude. In addition a superconducting final bending section could be realized as an achromat with significant
1008 momentum bandwidth, thereby allowing for fast energy scans. The potential use of high temperature superconductors (HTS)
1009 could further increase the simplicity and attractiveness of such solutions. In summary particle therapy has significant potential

1010 for enhancements of the treatment quality, and with regard to size and weight of facilities by utilizing modern superconducting
1011 technology. Accelerator driven subcritical reactors can be used to reduce the storage time of radioactive waste of nuclear power
1012 stations significantly. Due to the coupling with an accelerator, generating a fraction of the neutrons needed to fission high
1013 level waste incorporated in the sub-critical core, such reactors are passively safe. This application of high intensity proton
1014 accelerators could contribute to the solution of a major problem for the public society in CH. The cyclotron based High Intensity
1015 Proton Accelerator HIPA at PSI generates 1.4 MW proton beam power, and represents a prototype solution for an ADS driver
1016 accelerator. In addition PSI has performed the pioneering MEGAPIE experiment, in which a liquid metal target was operated
1017 with a megawatt class beam. Such target configurations are key elements for any ADS reactor. A collider facility with high
1018 energy reach and luminosity will consume significant electrical energy, of the order of TWh/y. With an increasing fraction of
1019 sustainable energy sources like wind and solar power in the future European energy mix, the production of energy will fluctuate
1020 significantly. One way to mitigate the impact of HEP facilities is to actively manage their energy consumption. The aim should be
1021 to avoid high loads on the grid during low supply conditions, and instead using preferentially excess energy. The possibilities of
1022 energy management using dynamic operation of facilities and energy storage systems are investigated for industrial applications
1023 and potentially synergies with HEP infrastructures could be realised. It is necessary to invest R&D efforts improving the energy
1024 efficiency of HEP facilities through critical technologies. In certain areas such R&D will have an immediate impact on research
1025 facilities operated today, and the savings in energy consumption may be used to co-finance the investments. Certain improved
1026 technologies may also serve the society. The fields of R&D include optimized magnet design, efficient RF power generation,
1027 cryogenics, SRF cavity technology, beam energy recovery, district heating using recovered heat, and energy storage. The use
1028 of permanent magnet material replacing electrical coils for accelerator magnets is a promising technology in this context. The
1029 ongoing design of SLS2.0 at PSI foresees the extensive use of permanent magnets and through the realization of the project the
1030 technology will be refined. Another area is the development of efficient klystron based RF sources by improved beam dynamics
1031 or the use of HTS superconducting focusing coils.

1032 **Input from Pillar 1**

1033 [\[Editor: Anna\]](#)

1034 **10.2 High energy**

1035 In order to discover and measure the Higgs boson, silicon tracking and calorimeter detectors, as well as superconducting-magnet
1036 technologies have been pushed to new limits, with technologies being transferred into the medical sector for biomedical imaging,
1037 as well as molecular and atomic structures. Advances in machine learning for signal processing have also cross-pollinated both
1038 particle physics. Some of the foundational research in solving the mystery of electroweak symmetry breaking in particle physics
1039 was done within the context of superconductivity in condensed-matter physics [?]. A synergy between solid-state processes and
1040 particle-physics processes has cross-pollinated both fields over the years. Also, with the explanation of a scalar Higgs field filling
1041 the universe being confirmed, its impact on cosmology has become a topic of interest, sometimes known as Higgs cosmology.
1042 The effect of the Higgs field on the inflationary period of the beginning of our known universe, as well as cosmological phase
1043 transitions and the stability of the electroweak vacuum have meant that cosmology is discussed in particle physics conferences
1044 and vice versa. Indeed, the future stability of the universe depends on the precise form of the Higgs boson field, and our very
1045 existence depends on whether we are in a stable minimum of the Higgs vacuum potential or not.

1046 For the last several decades modern HEP has successfully relied on human-engineered features, heuristics and algorithms. With
1047 the LHC and its upcoming HL-LHC upgrade, HEP has entered the era of truly Big Data. What is needed is faster MC simulation
1048 of synthetic data, faster data reconstruction algorithms, and to alleviate the data storage bottleneck: a move towards real-time
1049 data analysis. Modern machine learning can provide solutions to these problems. It can also provide a more efficient approach,
1050 given both human and computing resources, to analyzing the LHC data and inferring physics knowledge, e.g. for the identification
1051 of physics objects, event classification, measurements of properties through regression, and a more unified approach to searches
1052 for BSM physics by aid of anomaly detection techniques. In addition LHC's real-world science questions define realistic new
1053 benchmarks, which are of relevance for the ML community as a whole. This approach is complemented by modern engineering
1054 commodity hardware, such as very fast FPGAs, including System-on-Chip (SoC) devices, GPUs and powerful computing farms,
1055 to address the challenge of real-time data analysis.

1056 Synergies flourish in the area of detector development. Collaborations are required between material science and particle physics
1057 for the development of sensors and between electrical engineering and particle physics for the development of fast electronics and
1058 triggering systems. Technologies that are developed for particle physics experiments find applications elsewhere, for example
1059 in medicine, where the applications of detector RnD are numerous and in particular in positron emission tomography (PET)
1060 design. A team from ETHZ is working on a new generation of PET scanners using crystal detector technologies that are based
1061 on developments made for the CMS calorimeter. Teams from U Geneva and U Bern are developing fast silicon sensors that will
1062 constitute the building block of a Time-Of-Flight PET scanner of high granularity for ultimate use in a MRI scanner; this work
1063 is done in close collaboration with the University Hospital in Geneva.

1064 While HEP has very peculiar computing requirements because of the need to process large volumes of data (pushing the use of
1065 fast networks, fast processors and large storage sites), many synergies with other disciplines can be found in the development of
1066 flexible software to allow running on different clusters technologies and sites. [TODO: make the statement more concrete with
1067 examples?]

1068 **10.3 Low energy**

1069 In low-energy particle physics there are three types of synergies that can be outlined: (i) Technology transfer leading to the use
1070 of equipment and know-how developed for particle physics in other applications. (ii) The use of the particles as probes, e.g. in
1071 material science and chemistry, or their application in irradiation, medical physics or isotope production. This is connected to the
1072 application of particle physics techniques to other fields. (iii) Transfer of technology and techniques from other fields leading to
1073 progress in particle physics.

1074 Examples for (i) concern detector technology and electronics. At PSI, technologies for wire chambers, scintillators and light
1075 read-out found their way from particle physics to instrumentation in muon spin rotation and neutron scattering. Chip design
1076 from particle physics (originally coming from the high-energy physics developments at PSI for CMS) found many applications.
1077 Cutting-edge Si pixel detector technology for X-ray detectors in light sources and for medical applications was derived and
1078 commercialized. The DRS4 chip, originally developed for the MEG experiment, is used in many more experiments world-wide
1079 and way beyond particle physics. Space applications have been derived from various chips developed for particle physics and
1080 photon science at PSI. Also certain software, such as the data acquisition system MIDAS and electronic logbook ELOG from
1081 PSI low-energy particle physics, found a very large and versatile user base.

1082 Examples for (ii) are material science and solid state physics research and chemistry with muons and positrons. Spins of positive
1083 muons can be tracked to give information on local magnetic fields. Lifetimes of positrons in material can inform about electron
1084 densities. Detection technology is usually transferred from low energy particle and nuclear physics. Negative muons allow for
1085 non-destructive material analysis techniques with depth information.

1086 Examples for (iii) are found, e.g., from laser physics and radio-chemistry. In the precision spectroscopy of exotic atoms new types
1087 of high-power laser systems are being developed in close cooperation of particle physics and laser science, also with interest for
1088 commercial applications. Radio-chemistry overlaps with low-energy physics in a number of nuclear physics related aspects such
1089 as provision of rare isotopes, preparation of radioactive targets and measurements of physical properties of certain isotopes.

1090 **Input from Pillar 2**

1091 [Editor: Michele]

1092 **Input from Pillar 3**

1093 [Editor: Ruth] To investigate the nature of dark matter, data not only from direct detection, but also from astrophysical ob-
1094 servations, production at colliders, beam-dump experiments and indirect detection are necessary. DM searches are therefore
1095 inherently multi-disciplinary including particle and accelerator physics, solid state physics and astrophysics. DAMIC-M com-
1096 bines solid state physics (device operation of semiconductors), nuclear physics (the major backgrounds are radioactive isotopes),
1097 and particle physics (the main field of research). XENON/DARWIN combine the physics of liquid noble gases with particle
1098 physics, nuclear and atomic physics, as well as detector physics.

1099 Also cosmological observations of the matter distribution in the Universe, e.g. with the Euclid satellite scheduled for launch in
1100 2021 which has a strong Swiss participation, are relevant to unravel the nature of dark matter.

1101 To understand galactic and extragalactic sources, multi-wavelength data are necessary, covering from radio over optical and X-
1102 ray to TeV data. CTA will contribute the γ -ray part to this multi-disciplinary endeavour. In the near future, multi-messenger
1103 astronomy will combine radio and optical observations with high-energy gamma rays, neutrinos, gravitational waves and cosmic
1104 rays. This will allow us to understand the most energetic astrophysical events/objects beyond their electromagnetic emission.
1105 The interpretation of these data requires modelling using most branches of theoretical physics, especially General Relativity,
1106 electrodynamics, plasma physics, quantum field theory and statistics.

1107

1108 The detection of particles of 100 MeV/nucleon to a few GeV/nucleon in deep space are of critical interest for a broad range
1109 of applications in space activities, but they have not yet been measured precisely and monitored long-term in deep space. Cur-
1110 rently UNIGE is leading an international consortium, funded by the EU H2020 FET-OPEN program to develop a demonstrator
1111 (Mini.PAN) in 3 years (2020-2022). The PAN concept, based on a low mass magnetic spectrometer with high precision silicon
1112 strip detector, has been presented to several deep space programs, including the NASA Artemis (Lunar Orbiting Platform-
1113 Gateway.) project, the ESA European Large Logistic Lander (EL3) Call for ideas, and the Jupiters radiation belts studies for the
1114 ESA's Voyage 2050 Call.

1115 **Input from Theory**

1116 **10.4 Cosmology and gravitational waves.**

1117 [Editor: Gino] There is an intimate connection between particle physics and cosmology: the aim is to build a link between the
1118 microscopic laws of physics and the macroscopic observations of the universe as a whole. In the last few years this link has been
1119 extended to gravitational-wave physics, which represent a powerful new probe of fundamental physics on the cosmological scale.
1120 Just to mention an example of such connections, the precise mechanism that implements the electroweak symmetry breaking in
1121 the SM, may have been connected to the inflationary phase of the early universe and may also have led to a phase transition
1122 which could be observed with the LISA gravitational wave observatory in the near future.

1123 Swiss theory groups at EPFL, and at the universities of Basel, Bern, Geneva, and Zurich are heavily involved in such type of
1124 research. Highlights of their recent research activity in this area include important results in the study of CMB anisotropies
1125 and large-scale structure (LSS), providing solid theoretical predictions for the LSS observables and actively participating to
1126 the current experiments in the field, such as Planck and Euclid. More theoretical results include the possible explanation of
1127 the acceleration of the universe and the phenomenon of dark matter as result of primordial black holes, and the corresponding
1128 analysis of gravitational waves (GWs) signatures of interest for present and current GW observers. Other interesting theoretical
1129 results include the development of the so-called Higgs-inflation scenario, where the SM Higgs boson is the field responsible
1130 for inflation, and the detailed analysis of baryogenesis on motivated BSM frameworks. Beside pure theoretical developments,
1131 future efforts in this field will be closely connected to the experimental developments both at the particle-physics and at the
1132 cosmological/gravitational frontier.

1133 **11 Relationship to industry**

1134 [Main Editor: Guenther]

1135 The fields of experimental particle and astroparticle physics have a long-standing tradition of (i) very close collaboration with
1136 (high-tech) industry and (ii) of pushing the technological frontiers, which ultimately results in innovations that are successfully
1137 transferred to the private company sector and industry. These frontiers are typically related to forefront nuclear and particle
1138 physics instrumentation developed for and installed in small- and large-scale detectors, as well as to particle accelerator tech-
1139 nology. In all this, Switzerland is particularly well placed, thanks to (a) its hosting of a considerable number of national and
1140 international high-tech companies, (b) the fertile grounds and resources available for founding spin-off companies and (c) the
1141 substantial support given by the Swiss academic institutions and its national lab (PSI) to those researchers that are interested in
1142 the tech transfer of their ideas, developments and inventions. In the following a not comprehensive list of examples will be given,
1143 intended to provide a glimpse of the rich spectrum of tech transfer activities, spin-off companies and other relations to industry
1144 that exist in the Swiss particle physics landscape.

1145 **11.1 Examples of spin-off companies**

1146 Spin-off companies founded by Swiss researchers in the last years typically have their origin in novel particle detector techniques
1147 and its related data acquisition systems, such as silicon-based pixel detectors, scintillating crystals or other materials, photosen-

1148 sors (most notably, silicon photo multipliers in recent times), or dedicated Application-Specific Integrated Circuits (ASICs) as
1149 front-end readout elements of such detectors. In addition, know-how on the usage and control of particle beams from accelerators
1150 turns out to be highly valuable in a number of applications. A particularly targeted sector is the "med-tech" field with focus on
1151 biomedical imaging, but also the fields of (homeland) security or the handling of nuclear waste have been addressed. A few
1152 examples of such spin-off companies are shortly described in the alphabetically ordered list below:

- 1153 • *Advanced Accelerator Technologies AG (AAT)* is a joint venture of leading global industrial suppliers for research and high-
1154 tech enterprise equipment and a commercialising and licensing partner to PSI. Its main mission is the commercialization of
1155 PSI-IP in accelerator technologies and applications, to create value beyond the shareholders' individual expertise, such as
1156 accelerator component & system design and realization, proton therapy instrumentation and services, compact accelerators
1157 such as synchrotron sources spanning various energies, neutron instrumentation, as well as services and consulting.
- 1158 • *Arktis Radiation Detectors Ltd*, co-founded by ETH professor A. Rubbia and former ETH PhD students R. Chandra and
1159 G. Davatz, was built on expertise related to the detection of neutrons and high-energy photons ("gammas") and addresses
1160 the issue of detecting radiological and nuclear materials that pose a threat to customer's safety and security. Arktis de-
1161 velops next generation systems that categorize, prevent, and intercept radiological and nuclear materials in addition to
1162 contaminated cargo.
- 1163 • *Dectris* has been established by former students of Prof. R. Horisberger (PSI) in 2006 thanks to their expertise in silicon
1164 pixel detector technology and has grown from an initial 4 to 130 employees by now. Its main products are 1D and 2D
1165 hybrid photon detectors for scientific, industrial and medical applications, such as the Pilatus pixel detector, the Mythen
1166 strip detectors, a new fast pixel detector with 3kHz frame rate (Eiger), or a high-Z detector for higher energy X-rays.
1167 While initially the main application was X-ray imaging at synchrotron facilities, e.g. protein crystallography, with time the
1168 product range has spread to other imaging applications in the industry and medicine (Human CT and Mammography).
- 1169 • *Positrigo AG* has been co-founded by Prof. G. Dissertori and Dr. W. Lustermann (ETH Zurich), together with former ETH
1170 PhD students (Dr. M. Ahnen and Dr. J. Fischer) and colleagues from the University of Zurich (Prof. B. Weber) and the Uni-
1171 versity Hospital Zurich (Prof. A. Buck). Building on their expertise on scintillating crystals and silicon-photomultipliers,
1172 that are key components for PET scanners (PET=Positron-Emission-Tomography, a biomedical imaging modality), and
1173 on previous experience with the development of a pre-clinical PET scanner, the company aims at the development of a
1174 cost-effective and versatile brain PET scanner for the early diagnosis and treatment follow-up of Alzheimer's disease. In
1175 addition to an ETH pioneer fellowship and a donation through the ETH foundation, setting up of this effort was supported
1176 by a dedicated Innosuisse grant.
- 1177 • *RADEC GmbH* was founded in 2017 and performs tests of existing electronic components and materials for their radiation
1178 hardness using particles generated at accelerators located at PSI or other facilities. It also offers advice and assistance in the
1179 development and construction of radiation-hard components and systems. RADEC GmbH collaborates with companies
1180 involved in creating technologies for space (e.g., ARC POWER GmbH, Kramert GmbH, Teledyne (E2V)), as well as
1181 companies developing terrestrial technology where failure due to radiation must be eliminated.
- 1182 • *SE2S GmbH* - Space Environment Systems and Service, founded in 2020 by PSI researchers and based on over 30 years
1183 of experience in radiation qualification, particle detection and data analysis, offers next generation services and products
1184 in radiation effects and qualification (e.g. radiation modelling, assessment of radiation effects), detection of particles and

1185 radiation (e.g. novel detector technologies, equipment, software), space weather services (e.g. space weather and radiation
1186 modelling, analysis of space weather data, impact assessment and risk prediction).

- 1187 • *TransMutex SA* is a Swiss company founded in 2019 by a team of present and former scientists mostly linked to CERN, with
1188 Prof.em. M. Bourquin (Geneva University) being a member of its Scientific Board. The company is developing the concept
1189 of accelerator-driven systems (ADS) invented by Prof. C. Rubbia to solve the issue of long-lived nuclear waste, based on
1190 key experiments performed at CERN and at PSI that have validated his idea. Furthermore, very significant advances in
1191 particle accelerators (e.g. high-power cyclotron technology) and computing power (e.g. simulation tools) have helped in
1192 building confidence that the ADS concept is ready for industrial development. Interestingly, in the recent *Energiebericht*
1193 of SERI a report by Prof. Bourquin on thorium ADS has been included. As a consequence, SERI now recommends that
1194 Switzerland should engage in that research.

1195 **11.2 Contacts and collaborations with industry**

1196 Besides the obvious contacts of Swiss researchers with spin-off companies, that are based on their own research and/or have
1197 been co-founded by them, our field profits from a very extensive portfolio of contacts and collaborations with small-, medium-
1198 and large-scale national or international companies, typically working in the high-tech industry sector. The usual trigger for
1199 such collaborations arises from the need of pushing technology beyond its current boundaries, when new detectors or accelerator
1200 components are under development for specific research applications - the detectors of the experiments at CERN's LHC being a
1201 prime example, but not only. Again, the scope of such collaborations is multi-faceted, ranging from simple orders of equipment
1202 based on the researcher's own in-house developments, to joint prototyping and/or large-scale production efforts, as well as
1203 joint ventures towards transferring and/or licensing and future commercialisation of intellectual property by industry. A typical
1204 approach is the development of tooling and testing equipment by researchers in order to enable the companies to improve and
1205 determine their production tolerances. As an example, in the context of the upgrade of the CMS pixel detector (in particular, for
1206 its CO₂ cooling system), researchers from UZH, together with Swiss industry, have developed tooling for aligning components to
1207 be welded; furthermore, an UZH scientist developed for the company a testing setup to pressure-test welded stainless steel pipes
1208 to reach the standards required by the experiment. Due to such relationships maintained with these companies, the scientists have
1209 been able to convince them to do further R&D work with titanium pipes to see if a reliable and lighter-weight system can be
1210 produced. As a consequence, the companies have been interested in exploring this R&D for other future contracts with industry,
1211 as well as other interested partners at CERN. This (not isolated) example shows how industry not only profits commercially from
1212 orders received from our field, but more importantly how the companies' internal expertise, its quality management and/or its
1213 product portfolio can be enhanced thanks to the close interactions with demanding customers, namely particle physicists. A few
1214 further examples of such collaborations, by far non exhaustive, are given in the list below.

- 1215 • *ESPROS photonics corporation - EPC*: Researchers at ETH Zurich, UZH and PSI, who lead the initial construction and
1216 upgrade of the CMS pixel detector, collaborate with this high-tech company located in Sargans (SG). The medium-scale
1217 company with 50 employees specializes in Integrated Circuits design and production. In particular, the Swiss researchers
1218 collaborate with EPC on the design of Monolithic Active Pixel (MAP) pixel sensors for future applications, using the
1219 company's special CMOS technology.
- 1220 • In the context of the aforementioned upgrade of the CMS pixel detector, especially related to the development and con-
1221 struction of the so-called supply tube and the cooling system, UZH scientists have collaborated with a long list of Swiss

1222 companies (such as MEDELEC SA in Puidoux-Gare, Createch AG in Langenthal, Spalinger Präzisionsmechanik GmbH
1223 in Marthalen or Bolleter Composites AG in Arbon, just to mention a few), for the production of thin-wall precision tubes
1224 made of stainless steel and titanium, the bending and precision cutting of such tubes and related laser welding, the produc-
1225 tion of carbon-fibre or foam core based support structures, as well as complex plastic parts.

- 1226 • Scientists from the University of Geneva, involved in the ATLAS experiment, collaborate with Intel, towards the devel-
1227 opment of firmware using Intel's high-end Field Programmable Gate Array (FPGA) that will be used for ATLAS' future
1228 trigger system. This joint venture results in interesting experience gained on both ends, since, e.g., ATLAS' low latency
1229 applications do not fall into the typical use case spectrum of Intel's FPGAs.
- 1230 • Researchers at PSI, lead by Dr. S. Ritt, have developed the so-called DRS4 readout chip for the MEG experiment, which
1231 they sell through PSI's technology transfer program. So far, this has resulted in already more than 200 international
1232 companies and institutes as customers. As an example, the Italian company CAEN SpA (specialized in High/Low Voltage
1233 Power Supply systems and Front-End/Data Acquisition modules) has a product with the DRS4 chip. The PSI group is
1234 currently in discussion with a start-up company (RADEC) to outsource the chip distribution.
- 1235 • CAEN SpA, mentioned above, is also a partner of choice for many other particle physics experiments, not only at the LHC,
1236 resulting in numerous commercial contacts and joint developments with Swiss scientists.
- 1237 • In the context of the aforementioned PET scanner developments at ETH Zurich, close contacts for the production of silicon-
1238 photomultipliers (photosensors) have been established with the Japanese company Hamamatsu, and with the Chinese
1239 company Sichuan Tiangle Photonics Co. for the delivery of scintillating crystals made of LYSO. Hamamatsu will also
1240 be the main supplier of silicon strip and pixel detector modules for the major upgrades of the ATLAS and CMS tracking
1241 detectors.
- 1242 • In the past, during the first construction of the CMS experiment, important collaborations and major industrial contracts
1243 with Swiss industry had been established, in particular for the construction of the CMS superconducting magnet cables
1244 (Kabelwerke Brugg AG and Nexans in Cortaillod) and the large-scale manufacturing of printed circuit boards (ASCOM
1245 Systec AG).
- 1246 • In general, accelerator-driven large research facilities, such as PSI or CERN, give rise to a substantial number of close
1247 collaborations and joint ventures with industry, combined with tech transfer. Here we mention only a few examples,
1248 related to (a) the (co-)development of components (Daetwyler Industries, Cosylab, SCS-Super Computing Systems, Fer-
1249 rovac GmbH, VDL etc.), (b) imaging and analytics (Anaxam, ABB, Roche, GE-General Electric, Novartis, Nestle, BASF,
1250 etc.), (c) medtech such as proton therapy and medical imaging (Varian, Schär Engineering AG, etc.), and (d) business
1251 development (SwissNeutronics, InterAx, GratXRay, Eulitha, etc., or Dectris and AAT as mentioned earlier).

1252 In conclusion, research in particle physics instrumentation and accelerators has provided very fertile grounds, and will continue
1253 to do so in the near and far future, for win-win collaborations between academia and industry, with Swiss companies playing a
1254 particularly relevant role.

1255 [TODO: add links (as hyperlinks) to all company names? for a printer version also as dedicated references or footnotes?

1256 QUESTION: even more on accelerator technologies? is there anything worth mentioning related to computing and/or software,
1257 eg. deep learning?]

1258 [Ruth Durrer]

1259 From the hardware developed for CTA prototypes, silicon photomultipliers have been adopted which are now used in a project
1260 for the development of a β probe to drive surgeons in the ATTRACT program. The FET-OPEN EXCHANGE project SENSE
1261 has developed a roadmap for such low light level sensors.

1262 Direct dark matter detection experiments with liquid Noble gases have a wide range of possible industrial applications: Devel-
1263 opment of materials with extremely low levels of radioactivity and low radon emission.

1264 Development of low-noise, VUV sensitive SiPMs (and other photosensors) for operation in liquid xenon detectors.

1265 Development of low-noise, low-radioactivity electronics that works at cryogenic temperatures.

1266

1267 DAMIC-M is pushing the way for extremely low-energy threshold detectors, which may find a use commercially measuring
1268 extremely small interaction processes. Such CCD detectors function as very small detectors that can detect low-rate nuclear
1269 processes.

1270 **12 Impact on education and society**

1271 [Main Editor: Katharina] [(2-6 pages) – How is your research positively impacting on education and society? What are the
1272 benefits of the pursued research? Is your field offering a service to society (e.g. health, meteorology, agriculture, environment,
1273 energy, hazard warning, etc.)? Is there a link with the politics, do you provide advice for political decisions? Do you have links
1274 with museums and scientific collections? Do you support the promotion of young talents in your field? Do you have outreach
1275 activities? Is there something to say in relation with the sustainable developments goals (SDGs) or with possible citizen science?]

1276 **12.1 Education**

1277 The purpose of this section is to briefly discuss the structure of particle physics education in Switzerland offered at the universities
1278 with a focus on students studying physics as a major. The success of particle physics research in Switzerland largely results from
1279 the high-qualified and innovative scientific and technical teams within Swiss institutes. To maintain that quality, the best highly
1280 motivated students must be attracted to the field. For this to achieve education in particle physics in all undergraduate physics
1281 curricula is mandatory.

1282 **12.1.1 Bachelor and master**

1283 At Swiss universities general courses in nuclear and particle physics are commonly included in the final year of the Bachelor
1284 programs. Some counter example exist, however. There are Bachelor programs in physics for which an option with strong
1285 emphasis on nanoscience and technology or an option with an extended minor can be chosen. These curricula include nuclear
1286 and particle physics only as a elective or core elective modules, allowing the possibility of a student being awarded a Bachelor
1287 degree in physics without having followed these courses.

1288 All Swiss universities offer master programs with a strong focus on particle or astroparticle physics. Some of the master programs
1289 are clearly structured, targeting towards a specialisation in the chosen topic areas, while others encourage breadth but allow
1290 specialisation, if the student so wishes. ETHZ offers a unique and very attractive joint Masters degree in High Energy Physics

1291 together with École Polytechnique (Institut Polytechnique de Paris (IP Paris)) preparing excellent students for a future research
1292 career in High Energy Physics. The two-year Master program is set up symmetrically between the two universities: students
1293 spend one year in Zurich, and one year in Paris.

1294 **12.1.2 PhD**

1295 CHIPP initiated a specialised education program in particle physics open to PhD students all over Switzerland with the CHIPP
1296 Winter School and the Zuoz Summer School organised bi-annually. The purpose of the schools is to learn about recent advances
1297 in elementary-particle physics from local and world-leading researchers and our PhD students are expected to participate at
1298 least once during der PhD studies. The program includes lectures on accelerator and non-accelerator particle physics from an
1299 experimental and phenomenological perspective based on the activities of the swiss institutes involved in particle and astro-
1300 particle physics. Further education of the students is guaranteed and supported by the institutes through specialised schools
1301 offered by CERN, Fermilab, DESY or other institutions.

1302 **12.2 Outreach activities in Switzerland**

1303 Any new large-scale project to be proposed in particle physics will need concerted, global education, outreach, and commu-
1304 nication efforts, with a strong and committed dialogue with the public and stakeholders, and adequately educating pupils and
1305 students at all ages. Scientific outreach fulfils important and necessary obligations to society. The activities involve direct partic-
1306 ipation of scientists active in current research of particle and astro-particle physics to improve public understanding of our field,
1307 appreciation of the benefits of fundamental research, to rise interest and enthusiasm among young people, and to strengthen the
1308 integration of science in society.

1309 Current efforts in particle physics outreach in Switzerland raise awareness, appreciation and understanding of the field and its
1310 current state of research. The different outreach activities address diverse audiences and different venues ranging from traditional
1311 ones such as schools, science festivals or museums to YouTube videos, science slams, bars or music festivals. Audiences include
1312 primary and secondary school pupils, teachers, journalists and communicators, key stakeholders and policy makers, as well as
1313 the general public. Outreach activities of all institutes and universities are thus developing broad, long-term impact, making use
1314 of current research to raise and maintain the interest of the audience, but taking also the time to address the underlying nature of
1315 the scientific process, the strength of fundamental research and its key role in society.

1316 Outreach activities within CHIPP are intended to inform the political platform and the general public but also to target specifically
1317 potential young physicist and high-school students in general. In the view of the Swiss particle physics community, the primary
1318 aim is to convey to young secondary school students by conveying the importance, excitement and fantasy of basic physics and
1319 in particular recent particle physics and related cosmology developments. In this process the importance of a sound mathematical
1320 background is transferred as physics is by definition a mathematical description of fundamental phenomena. By convincing the
1321 audience of the importance of fundamental research in general and particle physics in particular these outreach activities serve to
1322 the benefit of all STEM related subjects and fundamental research in general.

1323 In the following we outline a few of the key activities for the general public and high-school students of the past years

- 1324 • Visits to CERN: CERN as the centre of high energy research is extremely attractive for visits which are organised regularly
1325 by CHIPP members. In recent years about 50 visits a year were organised for university students in physics and other

1326 disciplines, high-school students, alumni, politicians, members of societies, media, and the general public at large.

- 1327 ● Talks to the general public, the industry and high schools: many CHIPP members, involving all CHIPP institutes are
1328 actively participating giving talks to the general public in addition to the regular public talks organised by the institutes.
- 1329 ● Teacher education: we collaborate with secondary-school teachers in the development of innovative and interesting physics
1330 demonstrations, sometimes using particle physics data. Education of secondary-school teachers is done by providing
1331 teaching material, via the CERN Teacher program (<https://teacher-programmes.web.cern.ch/>) and specific topological
1332 workshops as well as open days for teacher at our institutes.
- 1333 ● YouTube video 'How particle-physics works: hope and worries on the B-physics anomalies': This short movie illustrates
1334 how experimental and theoretical physicists at UZH that work together to understand recent puzzling results in B-physics
1335 reported by the LHCb experiment (<https://www.youtube.com/watch?v=9dLyTS0Xscw>).
- 1336 ● Exhibits: A multidisciplinary Art& Science exhibition at the Espace Ballon in Château-d'Oex presenting on the discovery
1337 of cosmic rays, protagonating the Swiss physicist Albert Gockel from Fribourg who established first hints in his balloon
1338 flights over a century ago. Experiments carried out at high altitudes, in balloons, airplanes and in high mountain stations,
1339 such as the Jungfrauoch and Gornergrat research stations, allowed researchers to detect radioactivity in the atmosphere
1340 and to conclude on the existence of cosmic radiation (<https://www.chateau-doex.ch/de/P395/ballonraum-espace-ballon>).
- 1341 ● Scientifica – the Zurich Science Days: this bi-annual event attracts typically more than 25'000 visitors. Particle and
1342 astro-particle physicists of ETHZ and University of Zurich contribute regularly with topical talks and booths. In 2019 for
1343 example the general topic of the Scientifica 'Science fiction - Science facts' was perfectly suited to discuss antimatter and
1344 Dark Matter with the general public. In 2017 the discovery of the Higgs boson was discussed in the overall context of
1345 'What data reveals' (<https://www.scientifica.ch>).
- 1346 ● Dark Matter Day: Since a few years, the world celebrates end of October the hunt for the unseensomething that scientists
1347 refer to as dark matter. Swiss institutes regularly contribute with local events and highlight the experiments that could
1348 deepen our understanding of the mystery of dark Matter.

1349 Specialised school labs as well as lectures and workshops for school classes play a key role in attracting young students to
1350 study STEM related subjects. There are several dedicated laboratories at our institutes that offer special courses in cosmology
1351 as well as particle, astroparticle and neutrino physics for school classes targeting different ages of young students. With
1352 hands-on experiments, visits to the labs and by meeting bachelor and master students they get in contact with state of the art
1353 research and passionate researchers (<https://www.psi.ch/ilab/>, <http://www.sciencelab.uzh.ch>,<https://dqmp.unige.ch/physics-for-all/physiscope/>).

1355 **12.2.1 International outreach network**

1356 The International Particle Physics Outreach Group (IPPOG) (<http://ippog.org/>) has the mission to maximise the impact of edu-
1357 cation and outreach efforts related to particle physics and is an excellent example how outreach is done in a collaborative effort.
1358 Since 2016 IPPOG is an international scientific collaboration of scientists with experience in research, education & outreach from
1359 26 countries, six experiments and two international laboratories. IPPOG provides a network of scientists, science educators and
1360 communication specialists working across the globe sharing knowledge and providing tools for outreach in particle physics and

1361 related topics such as astro-particle, neutrino physics, radiation treatment and gravitational waves. Hans Peter Beck (University
1362 of Berne and Fribourg) served as IPPOG co-chair from 2016 – 2019.

1363 The European Particle Physics Communication Network (EPPCN) (<https://espace.cern.ch/EPPCN-site>) is a network established
1364 by the CERN Council in 2005 following the approval of the European strategy for particle physics. It is a network of professional
1365 communication officers, with Angela Benelli for Switzerland, from each member and associate states with the mandate to support
1366 and strengthen communication between CERN and the member states.

1367 **12.3 Support of young talents**

1368 Particle physics is a field that equally fascinates and attracts high-school and university students; the research field is therefore
1369 very well suited to attract interested, talented high-school students to study physics and later to motivate excellent students to
1370 follow a career as a researcher. Several activities for different ages are already in place others will be developed in the next years.

1371 In the following the different activities to attract young talents to our field of research are summarised:

- 1372 ● International Particle Physics Masterclass program for high-school students (<https://physicsmasterclasses.org/>): This in-
1373 ternational one-day program is targeting high school students that are very interested in physics and particle physics in
1374 particular. After an introduction in the concepts of the Standard Model and the measurement techniques the students learn
1375 through hands-on experiments and interaction with physicists at CERN how to perform a simple analysis. In Switzerland
1376 the various masterclass events typically attract about 200 students each year;
- 1377 ● High-school internship at CERN (<https://hssip.web.cern.ch/>): Two weeks internship at CERN for 24 high-school students
1378 from Switzerland which is offered in 2021 for the first time. More than 60 excellent applications of extremely motivated
1379 students were received which made the selection of the candidates a challenging task;
- 1380 ● Individual coaching for high-school students, eg. an internship or support with the matura thesis;
- 1381 ● Internship for students in physics or related fields in our research groups;
- 1382 ● CERN summer student program for students pursuing bachelors or masters degrees in physics, computing, engineering
1383 or mathematics (<https://home.cern/summer-student-programme>). Students attend lectures and perform their own research
1384 projects at CERN during eight to thirteen weeks. As this is an international project with more than 3000 applications for
1385 the 340 places, competition is high and only well prepared applications of extremely motivated students succeed;
- 1386 ● Mentoring of PhD students and Early Postdocs in the LHC collaborations ALICE, ATLAS, CMS and LHCb. All LHC
1387 collaborations have installed early career offices that organise trainings for newcomers and provide help and advice in the
1388 early career of young scientists. Senior scientists from the Swiss universities are actively supporting these early career
1389 efforts.

1390 Our efforts will be further enhanced in the coming years to strongly motivate and support young talents in order to strengthen the
1391 next generation of young scientists that choose to engage in the particle physics community.

1392 **12.4 Service to society**

1393 Traditionally, particle physics has been a driving force for new technical developments in medicine, such as positron-emission
1394 tomography or cancer therapy with proton or heavy ion beams. In addition detectors, that have been developed for measurements
1395 of charged particles in high-energy experiments are nowadays used for precise position in medical imaging. Many tools and soft-
1396 ware applications are in fact very similar in particle physics and medical applications. With the need of even preciser and faster
1397 detectors, radiation harder detector material and readout electronics, R&D is continuously ongoing, our understanding deepens
1398 and it can thus be expected that further developments in the detector material and types, electronics and analysis algorithm will
1399 have a significant impact in other fields such as medicine, material sciences or space sciences.

1400 **12.5 Summary**

1401 As the number of scientists engaged in outreach increases, so do the variety and ingenuity of their efforts reflected in the wide se-
1402 lection of activities, ranging from Open Days to Public Lectures, from lab tours to special workshops for high-schools. Travelling
1403 and standing exhibitions attract broad audiences and various events are organised at schools, universities, museums, and science
1404 cafes to raise interest and engage the audience. Many institutes offer projects primarily aimed toward high school students and
1405 teachers. Often, as a result of these efforts, young students might become more inclined to choose a STEM-related subjects for
1406 their studies. They might even go on to join the next generation of particle physicists. What is most important, however, is
1407 the fact that they will be more educated and appreciative of the importance of research, and thus more suited to make informed
1408 decisions about science and scientific questions in their future.

1409 **13 Vision for the future**

1410 [Main Editor: XXX] [(6-12pages) – Explain how the landscape is foreseen to evolve until 2025-2028. What are the future trends
1411 and the development opportunities. What fields of research are getting more momentum and what is rather to stay constant or
1412 get less interest in the future? Are there game-changing new technological possibilities to be expected (e.g. Big Data, artificial
1413 intelligence, new imaging/analysis capabilities, etc.)? Are there new infrastructures already being built in the years to come?
1414 Are there new international collaborations foreseen? Where shall Switzerland reinforce its position, follow-up new international
1415 trends, etc.?)

1416 [The following split in sections is intended as an intermediate step for collection of the necessary material. The main editor of
1417 the chapter will merge accordingly after the material is available.]

1418 **Input from Theory**

1419 [Editor: Gino]

1420 **Input from accelerator research**

1421 [Editor: Lenny]

1422 **13.1 Vision for the future**

1423 The Swiss Accelerator Research and Technology CHART projects contribute to one of the highest priorities of the European
1424 Strategy for Particle Physics to focus on advanced accelerator technologies, in particular the high-field superconducting magnets,
1425 including high-temperature superconductors. Swiss scientists in close collaboration with international partners are investigating
1426 the technical and financial feasibility of a future hadron collider at CERN with a centre-of-mass energy of at least 100 TeV
1427 and with an electron-positron Higgs and electroweak factory as a possible first stage [?]. Work on full understanding of the
1428 subsurface geology that will be crossed by both the tunnel and the access shafts is being carried out as one of the CHART
1429 projects at the University of Geneva's Geology Department. The innovative Linac technology developed for SwissFEL can be
1430 used advantageously for the realization of an injector concept for FCC-ee. Established low tolerance manufacturing methods
1431 allow for a cost effective mass production and good performance at the same time. A CHART project aims at developing an
1432 injector concept including an efficient positron source. As compared to an FEL the collider facility requires significantly higher
1433 bunch charges for maximum luminosity and positron production by the Linac beam. Controlling the collective effects triggered
1434 by the high intensity beam is one of the challenges for this project. Another challenge is the realization of a positron source
1435 that delivers the desired beam intensity for a collider. In particular the conversion efficiency from electrons to positrons must be
1436 maximized, while keeping the thermal and thermomechanical requirements for the conversion target realistic. It is planned to test
1437 a prototype of this newly developed positron source in the SwissFEL facility at 6 GeV. Any future collider facility serving particle
1438 physics research will represent a large accelerator based research infrastructure (RI) with significant investment and operating
1439 cost, as well as electrical power consumption which is not only a cost factor. For the realization of such projects sustainability
1440 aspects as energy efficiency and other factors impacting the environment are becoming increasingly important. A proposal for a
1441 competitive European RI must be optimized in view of many aspects, foremost the physics reach and the cost, and also in view of
1442 these sustainability aspects. The particle and accelerator physics community should work towards developing technological and
1443 conceptual advancements in multiple fields that contribute to an overall optimization of the concept. In the field of accelerator
1444 R&D the CHART program focuses on important developments in this context, such as the development of high field s.c. magnets
1445 aimed at maximizing the energy reach of a circular collider facility for a given size. CHART is addressing the most pressing
1446 problems in the design and construction of superconducting accelerator magnets:

- 1447 ● establishing a magnet laboratory at PSI capable of the design and construction of superconducting accelerator magnets in
1448 Nb₃Sn and HTS (REBCO tape) technologies,
- 1449 ● providing generic enabling-technology R&D and apply the results to improve the performance of Nb₃Sn canted-cosine-
1450 theta (CCT) magnets, as demonstrated by the delivery of a CCT technology demonstrator
- 1451 ● integrating with CERN's HTS technology-coil program, delivering a number of technology-coil assemblies that are to be
1452 tested in PSI's upgraded cryogen-free test station, and thereby introducing the full chain of HTS magnet design, construction
1453 and testing at PSI
- 1454 ● Investigation of the superconducting wire at the University of Geneva in order to increase its performance under transverse
1455 stress.
- 1456 ● Investigation of novel epoxy systems by the collaboration with the ETHZ Soft-Materials-Group, in order to provide optimal
1457 mechanical support to the superconducting wire in the coil matrix.

1458 The luminosity production per grid power is maximized through beam dynamics studies and advanced collision schemes. Alter-
1459 native collider scenarios with potential performance and efficiency related advantages like the muon collider may be studied in
1460 parallel.

- 1461 • perhaps comment on high power target expertise for muon production at PSI

1462 Another topic is the exploration of high gradient acceleration schemes utilizing micron scale accelerator structures and high
1463 power lasers. With regard to proton beam power and the intensity of generated low energy muon beams, the PSI HIPA facility
1464 provides a very competitive performance. While a significant further increase of the primary proton beam intensity is difficult to
1465 achieve, the conversion efficiency to muons has a good potential for further improvements. The improved target configuration
1466 and capture optics will not contribute to a higher power consumption but rather enhance the energy efficiency of the facility. The
1467 High Intensity Muon Beam (HIMB) project at PSI has the potential to enhance the intensity by more than an order of magnitude.

1468 **Input from Pillar 1**

1469 [\[Editor: Anna\]](#)

1470 **13.2 High energy**

1471 **Physics pursuits with the HL-LHC ATLAS and CMS experiments**

1472 The major motivation for the HL-LHC program, being installed from 2025 to 2027, and running from approximately 2027-2036,
1473 is to measure with high precision the least known Higgs boson properties, as well as to probe in depth the weak scale, using
1474 a dataset approximately 10 times larger than the previously existing dataset. With this dataset, improved ATLAS and CMS
1475 detectors for mitigating the pile-up due to higher instantaneous luminosity, and improvements on theoretical uncertainties, the
1476 HL-LHC is expected to deliver measurements of Higgs couplings with uncertainties reduced by a factor of two. The study of
1477 differential (and double differential) cross-section measurements, which are currently statistically limited, will also provide more
1478 opportunities for the discovery of new physics.

1479 One of the major goals of the HL-LHC will be to find evidence for the self-coupling of the Higgs boson. This effect leads to
1480 SM double Higgs production, HH , will not be observed or constrained strongly during the LHC running period. Both CMS and
1481 ATLAS have endeavored to estimate their sensitivity to this process, which requires two Higgs bosons to be identified in a single
1482 event. The best signal significance for this process is expected to be in the combination of one high-rate, high-background Higgs
1483 boson decay, with one of low-rate and low-background, leading to the golden channel of $HH \rightarrow b\bar{b}\gamma\gamma$. Swiss physicists have
1484 been active in $H \rightarrow b\bar{b}$ and $H \rightarrow \gamma\gamma$, are now leading the current $HH \rightarrow b\bar{b}\gamma\gamma$ analyses, and are continuing to develop detectors
1485 and triggering systems that are specialized for measuring these processes.

1486 The coupling of the Higgs boson to fermions in the first and second generation has not yet been observed. An observation of
1487 the Higgs boson coupling to muons is expected during run 3 of the LHC, however, since the branching ratio $H \rightarrow c\bar{c}$ is 20 times
1488 lower than that of $H \rightarrow b\bar{b}$, and c jets are identified with efficiencies 10 times lower than b jets, a measurement of the SM $H \rightarrow c\bar{c}$
1489 process is not expected at the HL-LHC. There is, however, an opportunity for discovery of new physics in the rare decays of
1490 the Higgs boson to various second-generation vector mesons and photons which have a SM branching ratio of the order of 10^{-6}
1491 and, while sensitivity to SM rates is not expected, BSM contributions can greatly enhance these rates. Swiss physicists will be
1492 investigating such rare Higgs-boson decays, as well as flavor-violating interactions of the Higgs boson such as $H \rightarrow \mu\tau$.

1493 Searches for new physics will carry outstanding importance in the HL-LHC program, with the large datasets giving the opportu-
1494 nity to probe rare phenomena where we would not have had access previously. The top quark, being the heaviest of the particles
1495 in the SM, will carry a central role in the future searches for NP due to its potentially increased sensitivity to BSM effects. In
1496 order to extend the discovery reach of the LHC, the use of indirect approaches such as automatized calculations, commonly done
1497 in the context of effective field theory (EFT) to analyse possible deviations with the SM is expected to take centre stage in the
1498 near future. Only recently have experimental measurements started to test directly the coupling of the top quark to Z, W, and
1499 Higgs bosons. The current and future ATLAS and CMS datasets will provide an intriguing opportunity to study these processes
1500 in more detail. The resonance search program will be extended to challenging areas of low signal rate, large signal width, in-
1501 cluding tails of distributions, as well as hard-to-trigger low mass region. The di-boson resonance program will be extended to
1502 non-standard boson polarisations. The Higgs physics program will be further expanded to various exotic Higgs scenarios. The
1503 SUSY physics program will further probe feeble cross-sections, such as those associated with electroweak production; it will
1504 explore R-parity-violating models; and it will be expanded towards compressed mass spectra and smaller couplings, resulting in
1505 soft and displaced objects in the final state. The search for HNLs in leptonic decays of W bosons will be extended to searches in
1506 B decays, taking full advantage of improved triggers strategies. This vast increase in statistics of B decays will also benefit other
1507 indirect searches for new physics in the context of lepton flavor violation.

1508 Important to achieving these research goals are improvements in the ATLAS and CMS detectors. In particular, the new timing
1509 layer upgrade at CMS, being built with Swiss participation, will improve object identification efficiency amid pileup, and will
1510 improve identification and energy reconstruction of photons in the central detector region to maintain high-quality $H \rightarrow \gamma\gamma$
1511 measurements. The new inner trackers of both ATLAS and CMS, being built with major participation from Swiss institutions
1512 will greatly improve measurements of $H \rightarrow b\bar{b}$ and $H \rightarrow \tau\tau$ measurements, as well as reduce the effects of pileup in all analyses.
1513 The introduction of tracking reconstruction early on in the triggering stages will equally be paramount for maximizing the
1514 acceptance to rare phenomena that are typically swamped in large rates of SM processes; this is a driving motivation behind the
1515 ATLAS and CMS trigger architectures.

1516 Optimized detector design will be followed by resource efficiency in the aforementioned areas of triggering, reconstruction and
1517 simulation. These translate directly to improved precision of SM measurements and increased sensitivity to NP given higher
1518 trigger efficiencies, improved reconstruction algorithms and higher statistics of simulated data to optimise the analysis strategies.
1519 Areas of particular interest to the Swiss research teams are searches for new physics objects leading to unconventional signatures
1520 in the tracking volume, or to anomalous jet substructure, as well as the combination of both phenomena. Modern tools based on
1521 machine learning provide cutting-edge technology that can be used to take full advantage of the unique LHC data set and at the
1522 same time to revolutionise the way we do science far beyond High-Energy Physics.

1523 **Flavour physics with LHCb at the HL-LHC**

1524 Flavour physics plays a unique role in the search for BSM physics, allowing the exploration of a region of mass and coupling
1525 inaccessible to current and planned direct detection experiments that could pave the way to NP discovery. Flavour physics is
1526 strongly linked to theoretical QCD computations on a lattice since some measurements require knowledge of the hadronic system
1527 to be interpreted. The correlations between the different measurements is a powerful weapon in flavour physics to disentangle
1528 NP from hadronic effects and can be used to advance theoretical knowledge of low-energy QCD. Since most key measurements
1529 in heavy flavour are statistically limited, it is of paramount importance to have a flavour physics experiment in the HL-LHC era.
1530 Multi-purpose flavour experiments at colliders, such as LHCb, are those offering the highest yields of hadrons containing bottom

1531 and charm quarks, as well as of tau leptons, and the widest spectrum of interesting measurements. An expression of interest has
1532 been submitted in February 2017 to the LHC committee for a second upgrade (Upgrade 2) after Run 4 (in ~ 2030). The idea is
1533 to operate at a luminosity of $2 \times 10^{34} \text{cm}^{-2} \text{s}^{-1}$, i.e. ten times that of the first upgraded detector, and improve the performance of
1534 the detector in key areas. With an accumulated sample of at least 300fb^{-1} , LHCb would then take full advantage of the flavour
1535 physics opportunities at HL-LHC. Switzerland intends to play a crucial role in this endeavour thanks to the experience of the
1536 EPFL and UZH groups in the current LHCb experiment and its upgrade.

1537 **Detector and computing**

1538 The Swiss particle physics community masters a wide range of detector technology: tracking detectors, calorimetry, triggering
1539 and DAQ. Due to the diverse expertise present in all institutions, the Swiss community is well poised to develop/adapt any
1540 hardware technology that would be needed for future facilities. Hardware expertise is therefore not perceived as a limiting factor
1541 to pursue future directions in the field. In the close future and beyond 2025, the focus of Swiss scientists is expected to be
1542 three-fold: the commissioning and operation of the HL-LHC detectors, detector and trigger upgrades within HL-LHC, and R&D
1543 for future facilities, in line with the European strategy recommendations.

1544 While the initial HL-LHC detector upgrades for the LHC Run 4 are well underway, discussions are now starting within the LHC
1545 experiments on detector upgrades for Run5. These upgrades will accommodate flexibility and challenges that are expected not to
1546 be fully addressed beforehand. They will also allow the experiments to respond to potential change in the physics landscape, in
1547 the case of an observed anomaly in data. As an example, the ATLAS collaboration is envisaging the replacement of the innermost
1548 tracking layers to account for radiation damage; at the same time, it considers an upgrade in the read-out electronics, which will
1549 in turns allow for an evolution in the TDAQ architecture of the experiment.

1550 The HL-LHC will require an increase in computing resources by a factor of order 50. A combination of scaling of the present
1551 resources and increase of processors performance by Moore's law will most probably not be enough. The present solution
1552 pursued by the HEP community is instead to enhance the parallelism of the algorithms and use more heterogeneous computing
1553 architecture including GPUs and FPGAs to run them. Machine learning will play a definite role in shaping those reconstruction
1554 algorithms (e.g. tracking and clustering running on GPUs), boosting the speed of simulations and in general in increasing
1555 the efficiency in extracting information from data. The investment in the hardware facilities will have to be paralleled by an
1556 investment in developing the software needed to accomplish these goals. To facilitate the cooperation within the HEP community
1557 towards the development of software and computing infrastructures several fora have been created, among which are the HEP
1558 Software Foundation (HSF) and the CERN "Scientific Computing Forum".[TODO: Add a reference to / quote from European
1559 strategy, chapter 4d.]

1560 **Probing particle physics further**

1561 As indicated in the European Strategy, new experiments beyond the ones belonging to the general purpose collider ones and
1562 which are exploring the dark sector have a rich future. The NA64 experiment is currently being upgraded and will resume data
1563 taking after LS2. The goal is to probe most of the remaining parameter space motivated by light thermal dark matter models and
1564 to completely cover the X17 anomaly parameter space. Moreover, a pilot run using the unique 150 GeV muon beamline at the
1565 SPS was approved to search for a new dark boson Z_μ with a mass in sub-GeV range, which is coupled predominantly to the second
1566 and third lepton generations. The existence of Z_μ would provide an explanation of the muon $g-2$ anomaly and is complementary
1567 to NA64 in electron mode to search for DS at higher masses [?]. The FASER collaboration is exploring ways to increase the

1568 detector precision and acceptance in what will become the FASER2 experiment, rendering it sensitive to a variety of additional
1569 physics channels that are currently inaccessible. Such a FASER2 detector would start design after FASER is commissioned in
1570 2022, aiming at being installed during LS3 for data taking at the HL-LHC.

1571 Beyond the HL-LHC upgrades, the high energy physics community views with enthusiasm the European strategy outcome, which
1572 supports R&D for a large Future Circular Collider (FCC), opening up enormous potential in the comprehension of our world.
1573 Exploring the properties of the Higgs boson continues to be one of the most pertinent tasks of the field, both in understanding
1574 electroweak symmetry breaking, the mechanism by which particles acquire mass, as well as searching for new clues to answer
1575 deep questions in the understanding of the universe. In the coming years, the community is asked to produce design reports
1576 for future detectors to be hosted in the prospective FCC, which is expected to motivate the Swiss scientists and the younger
1577 generations alike.

1578 **13.3 Low energy**

1579 A goal for the future, of course, is the discovery of new physics in low energy precision observables and/or forbidden decays.
1580 Ideally, this would come together with the observation of clear direct signals from high-energy collisions. The chances are good
1581 and some of the most promising and sensitive discovery channels are searches for violation of the symmetry between matter and
1582 antimatter (CP) and between leptons from different families (lepton flavor LF, here: muons and electrons). As such discoveries
1583 cannot be planned, measurements of SM parameters at the highest precision are also important, provide crucial input, confirm
1584 theoretical understanding in detail and exclude BSM theories.

1585 After LS2, the ELENA ring at the CERN AD will provide an unprecedented flux of low energy antiprotons. This will open a
1586 new era for precision tests with antimatter. Among those the measurement in GBAR of the gravitational acceleration \bar{g} imparted
1587 to freely falling anti-hydrogen atoms which will allow for a direct experimental test of the Weak Equivalence Principle with
1588 anti-matter [?, ?] and a stringent test of the CPT theorem [?].

1589 PSI is offering world-leading beams of low momentum pions, muons and ultracold neutrons used by a large and growing com-
1590 munity with strong Swiss participation and leadership. There is a unique opportunity to maintain the leadership in this attractive
1591 field and to substantially upgrade these facilities in terms of beam intensity and quality. This will translate into a significantly
1592 enhanced reach of the experiments and their physics potential, and pave the way for completely new experiments and research
1593 directions.

1594 On the one hand, this concerns the intensity of the source of ultracold neutrons (UCN) at PSI at which the search for the neutron
1595 electric dipole moment will also in 5-10 years still be statistically limited. On the other hand, this concerns the intensity of
1596 PSI's secondary muon beams which could be boosted by almost two orders of magnitude by the High Intensity Muon Beam
1597 project HiMB. In a similar direction, many experiments would benefit from improved muon beam quality, where the muCool
1598 project promises seven orders of magnitude improvement for the brilliance of slow positive muon beams with a plethora of
1599 applications in fundamental particle physics and in applied sciences. Obviously, the combination of muCool and HiMB will be
1600 highly attractive. With an additional project for cooling of slow, negative muons many more applications would show up, directly
1601 for muonic atom research and material surface studies, but it might impact future muon collider options as well.

1602 While important installations at other international facilities, such as at the CERN AD, at ILL and ESS with their existing or
1603 envisaged fundamental neutron physics programs, will be driven by the international community, partially with strong Swiss par-

1604 ticipation, the installations at PSI will be driven by the Swiss community (with strong international participation in experiments
1605 and applications).

1606 The single most important facility project of the next 5-10 years, with exploitation over the next more than 20 years will be the
1607 realization of the HiMB project. One very strong science driver on the particle physics side is the search for charged lepton
1608 flavor violation (cLFV), as ongoing with the MEG II and Mu3e experiments. The international Mu3e collaboration with leading
1609 contributions by groups from PSI, U Geneva, UZH and ETHZ has laid out a phased approach which ultimately needs HiMB to
1610 push the limits of cLFV searches with muons. HiMB at HIPA at PSI is of great interest for the Swiss particle physics community
1611 and beyond. Besides Mu3e, many particle physics experiments with muons can be tailored to benefit from a HiMB, and with the
1612 installation of two such beamlines a second one could serve material science applications with unprecedented statistical power.

1613 **Input from Pillar 2**

1614 [\[Editor: Michele\]](#)

1615 **Input from Pillar 3**

1616 [\[Editor: Ruth\]](#)

1617 DARWIN will probe WIMP dark matter down to the neutrino floor, and continue to broaden the DM reach by using the ionisation
1618 signal only, the Migdal effect and bremsstrahlung, as well as DM-electron scattering. It will search for the neutrinoless double
1619 beta decay of ^{136}Xe with sensitivity of 2×10^{27} y (in the baseline scenario) and several other double beta processes (see e.g., e-
1620 Print: 2002.04239), measure the solar pp-neutrino flux (via neutrino-electron scattering) with $\pm 1\%$ precision and the weak mixing
1621 angle at low energies. It will search for solar axions, DM ALPs and dark photons, nucleon decay and many other processes (a
1622 global white paper A Next-Generation Liquid Xenon Observatory for Dark Matter and Neutrino Physics is under preparation).

1623 LEGEND-200 will achieve a discovery potential of the neutrinoless double beta decay of ^{76}Ge of 1×10^{27} y with 1000 kg
1624 y exposure, while LEGEND-1000 will extend the sensitivity to 1×10^{28} y. This will allow us to cover the so-called inverted
1625 neutrino mass ordering scenario, probing effective Majorana neutrino masses in the range (10-20) meV (see also the Double Beta
1626 Decay APPEC Report, e-Print: 1910.04688).

1627 XENONnT will improve the sensitivity to WIMP dark matter by one order of magnitude. It will also probe LDM via DM-electron
1628 scattering, as well as ALPs and dark photons via absorption in liquid xenon. It will search for solar axions with unprecedented
1629 sensitivity, and detect solar neutrinos (8B) via coherent neutrino-nucleus scattering.

1630 DAMIC-M is set to begin in 2024, and will probe several theoretically viable models for low-energy interactions between DM
1631 and matter. It is positioned to be the world-leading experiment in studying hidden-photon DM, hidden photons mediating the
1632 interactions of dark matter, and electron scattering of DM at low energy scales. It has a broad reach to probe 10 orders of
1633 magnitude in DM mass over a range of theoretical scenarios. OSCURA, whose feasibility studies are supported by a U.S. DOE
1634 grant, has a timeline after DAMIC-M, and will have thousands of CCD detector modules, making its production similar in scale
1635 to detector production for the LHC experiments. The Swiss institutions plan to continue a leading role in this international and
1636 large-scale project.

1637 CTA has very high potential for the exploration of the universe of the most violent processes forming compact objects and

1638 accelerating particles to extreme energies. It has also high potential for the exploration of dark matter, whether it is made up
1639 of axions or WIMPs. CTA will be an extremely important observer to drive multi-messenger observations, which combines
1640 information from high-energy gamma, neutrinos, gravitation waves, charged cosmic rays. It will improve energy coverage and
1641 sensitivity of current ground based gamma-ray observatories by about an order of magnitude. Taking into account those being
1642 background dominated, this corresponds to a factor 100 in observation time.

1643 IceCube will extend its reach to cover neutrino oscillations and detect order of 100 and more cosmic events per year. Running
1644 with CTA and Advanced gravitational wave detectors, it will enhance the reach of the multi-messenger astrophysics.

1645 In the next decades CHIPP institutes will also have a very rich research program at the forefront of space astroparticle physics.
1646 There are 3 main themes: 1) High-energy astroparticle physics with direct particle detection in space from GeV to PeV (AMS-
1647 02, DAMPE, HERD); 2) Multimessenger astrophysics with X-ray and gamma-ray missions (POLAR-2, HERD, eXTP); 3)
1648 Multidisciplinary particle detection instrument development for deep space (PAN).

1649 **14 Development of national infrastructures (2025-2028)**

1650 [Main Editor: XXX] [(2-8 pages) – On the basis of the previous sections, what are possible developments (in some cases even
1651 essential needs) in terms of infrastructures to maintain or strengthen the Swiss scientific expertise in the field. Are there infras-
1652 tructures at national level, i.e. beyond what can be afforded by single institutes, that would be essential? Focus on the scientific
1653 benefits and the breadth of the community of users, whilst keep it very general on the size, the costs, possible geographical
1654 location, management structure, etc. These points will be defined in a second step involving the ETH-Board, swissuniversities
1655 and the institutions. They will be described in any specific proposal for a given infrastructure to be submitted by end 2021 (or
1656 early 2022) in view of an international evaluation conducted by the SNSF.]

1657 [Ruth Durrer]

1658 A large liquid xenon demonstrator (2.6 m tall TPC, in a 3.5 m cryostat) is in construction at UZH; this infrastructure (financed
1659 by an ERC Adv. Grant) will be available to other institutions in CH and Europe.

1660 **15 Swiss participation to international organisations (2025-2028)**

1661 [Main Editor: XXX] [(1-5 pages) – Is there a need for Switzerland to join an international organization to get access to one or
1662 several international facilities? Explain the benefits of this. Which specific community would benefit from this? Is it of strategic
1663 importance for Switzerland? Are there also positive implications to be expected for the industry or society?]

1664 [Ruth Durrer]

1665 CTA ERIC will be a worldwide organization on governmental level with solid Swiss participation (ETHZ, UNIGE). It is not yet
1666 clear if Switzerland will be a founding member or will have another status, e.g. a strategic partnership. The final decision about
1667 Switzerland joining the ERIC will require the Swiss Parliament approval.

1668 DARWIN is on the APPEC roadmap, the Swiss SERI roadmap and on the roadmap of several European funding agencies.

1669 LEGEND-1000 is one of the three double beta experiments under consideration by DoE in the US.

1670 The next big international DAMIC experiment is OSCURA, planned for 2026. OSCURA will push the limits of the CCD tech-
1671 nology to 10 kg of silicon, single-electron ionization threshold, and a detector with a background of only 0.01 events/kg/keV/day.
1672 Switzerland will continue its leading role in this experiment, which is currently a DoE BRN-funded R&D experiment.

1673 **16 Conclusion**

1674 [Main Editor: Rainer] [(1-2 pages) – In this section or imbedded in various recommendations in the text and listed in Sect. 2,
1675 there should be some consideration about the prioritization of investments in the field. It is a difficult topic to agree upon by
1676 the community, but leaving this completely open to decision-makers is not always the best alternative. As the funds are not
1677 infinite it would be good that the community gives basic recommendations on how to serve them optimally in case the list of
1678 possible investments clearly exceeds the available means. Some simplistic examples could be: focus on research infrastructures
1679 serving the widest community of users; avoid prestige infrastructures in areas not yet having a strong scientific community in
1680 Switzerland; rather consider joining a European facility than building something smaller in Switzerland, etc.]

1681 **17 Appendix**

1682 [(1-6 pages) – An appendix could be a list of people involved in the sub-groups formed in the preparation of this document.
1683 Another annex can give a list of acronyms used in the text. Concerning acronyms, try to refrain using them too widely to ease
1684 the reading by somebody not directly in the field. The same applies to references to scientific publications. Some key references
1685 can be given in appendix, if useful, but the roadmap shall not be a scientific paper with many references. Finally, the credit for
1686 figures and images shall also be included somewhere, either in the figure caption, or in an appendix, or an inside cover page.]

¹⁶⁸⁷ **test section**

¹⁶⁸⁸ This is some test text. See how to use a reference, e.g. of a great recent public result [?].