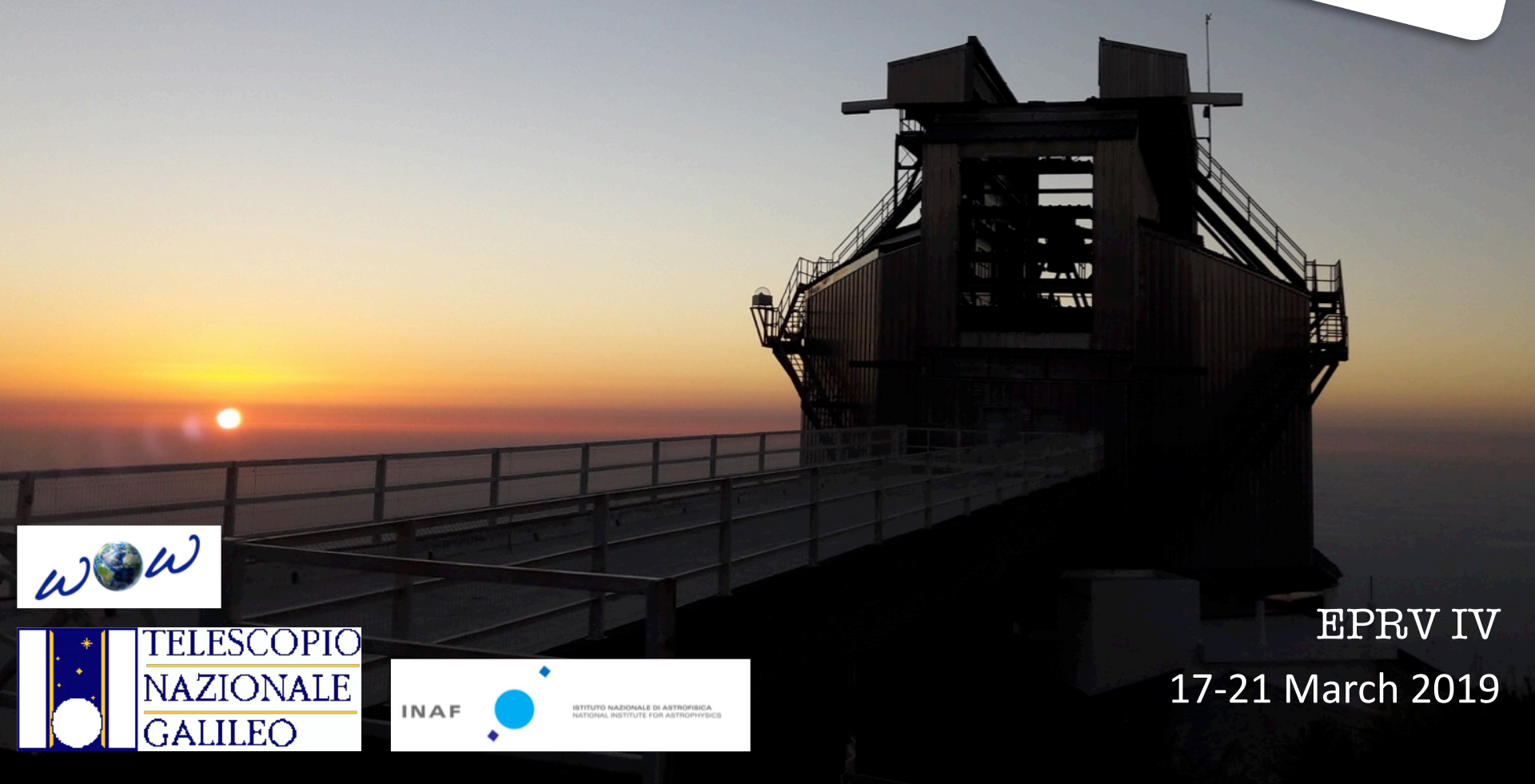


# RVs with

# GIANO-B @ TNG

I. Carleo & the GIARPS Team



EPRV IV

17-21 March 2019



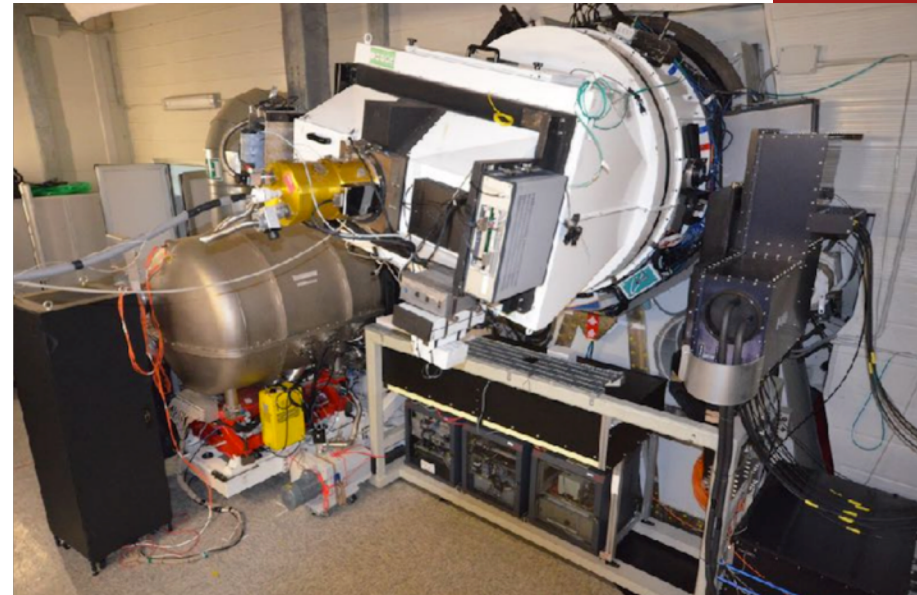
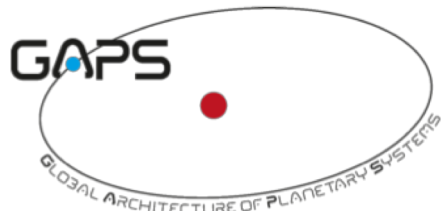
# GIARPS in a nutshell

**Aim:** high resolution VIS-NIR spectra  
+ high precision radial velocities

**Method:** simultaneous use of

- ❑ HARPS-N ( $0.38 \mu\text{m} < \lambda < 0.69 \mu\text{m}$ )
- ❑ GIANO-B ( $0.95 \mu\text{m} < \lambda < 2.45 \mu\text{m}$ )

**Management:** Italian Exo-planetary community through the Premiale WOW, with a particular effort by the GAPS team (funds, manpower, observing time)





# Main actions

- ❑ Transferring GIANO from Nasmyth-A to B
- ❑ Feeding GIANO through a new train of optics rather than fibers, long slit
- ❑ Insertion of a dichroic to split the VIS and NIR light to HARPS-N and GIANO
- ❑ New reduction pipeline for GIANO-B: GOFIO

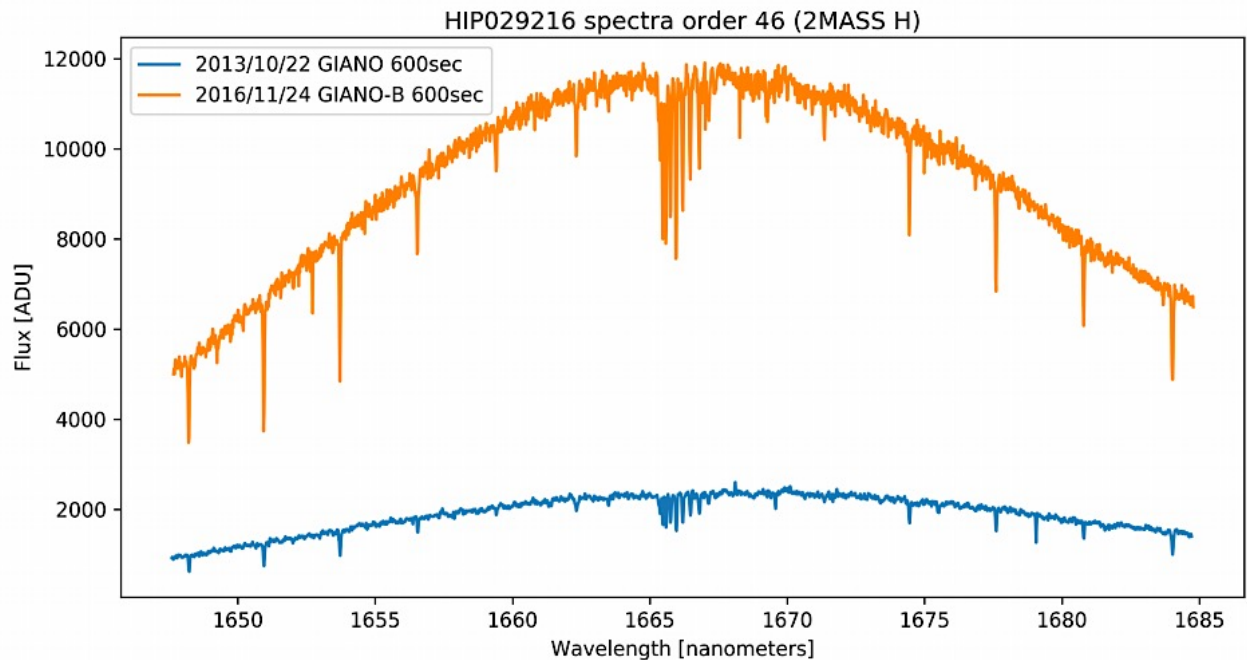


# The removal of GIANO fibers implies:

- Higher efficiency
- Elimination of modal noise
- Stable slit illumination with a closed loop active tip/tilt mirror

Flux ratio between the spectra in the NIR bands

- J: 4.4
- H: 4.7
- K: 5.3

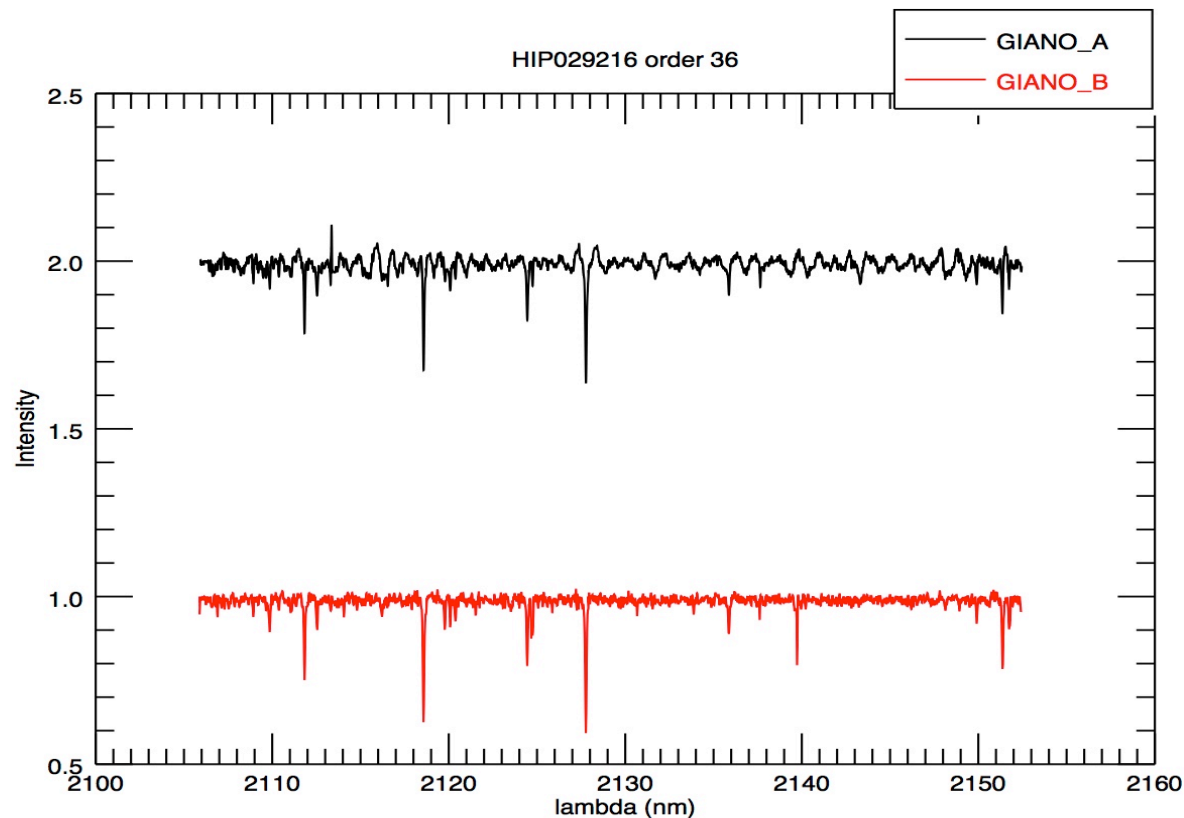


Hmag = 7.4



# The removal of GIANO fibers implies:

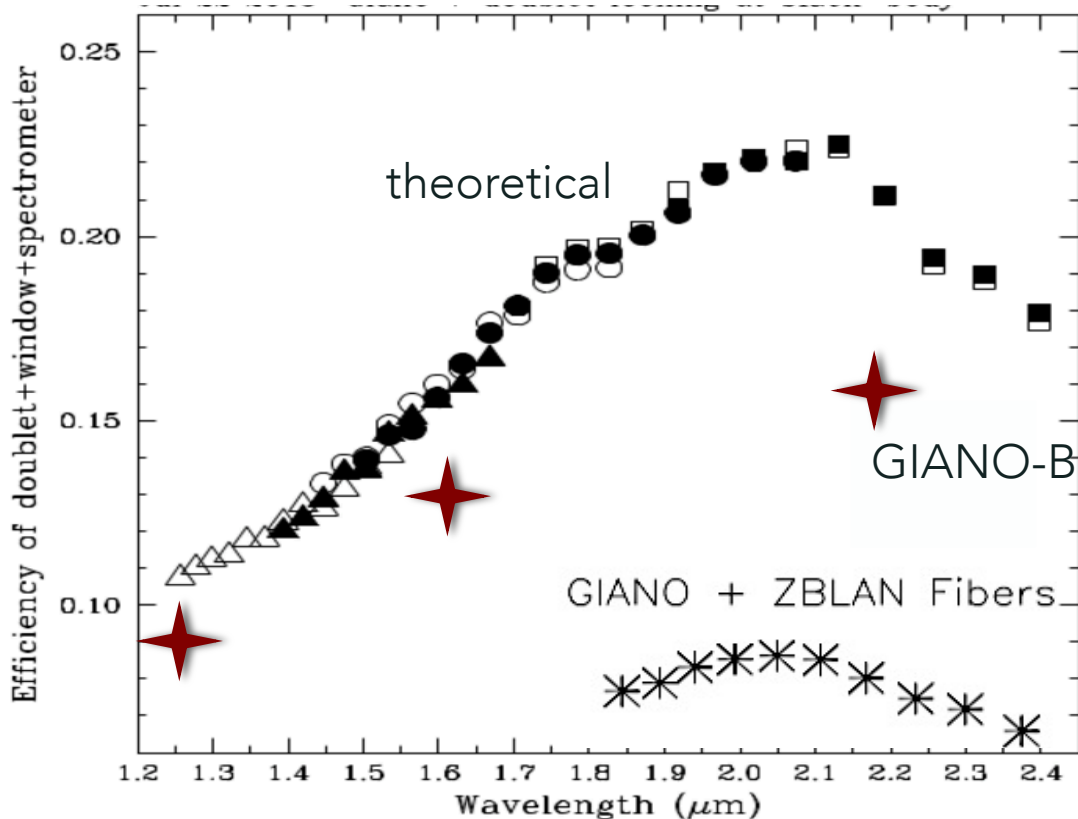
- Higher efficiency
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- Stable slit illumination with a closed loop active tip/tilt mirror





# The removal of GIANO fibers implies:

- Higher efficiency
- Elimination of modal noise
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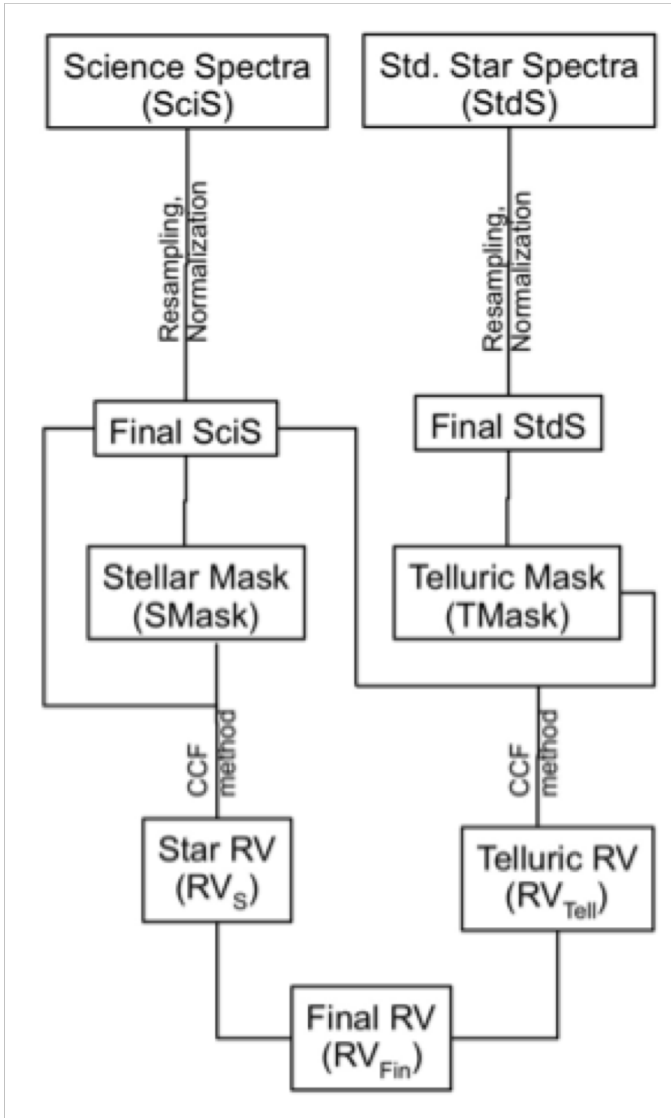


Adapted from Tozzi et al. 2014, SPIE paper

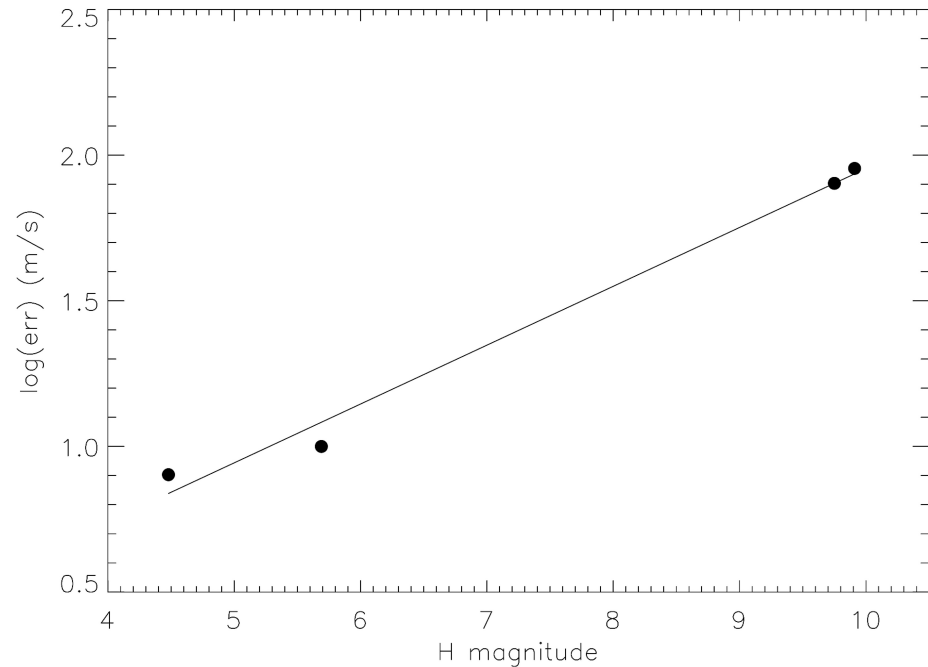


# RVs: telluric method

Carleo et al, 2016



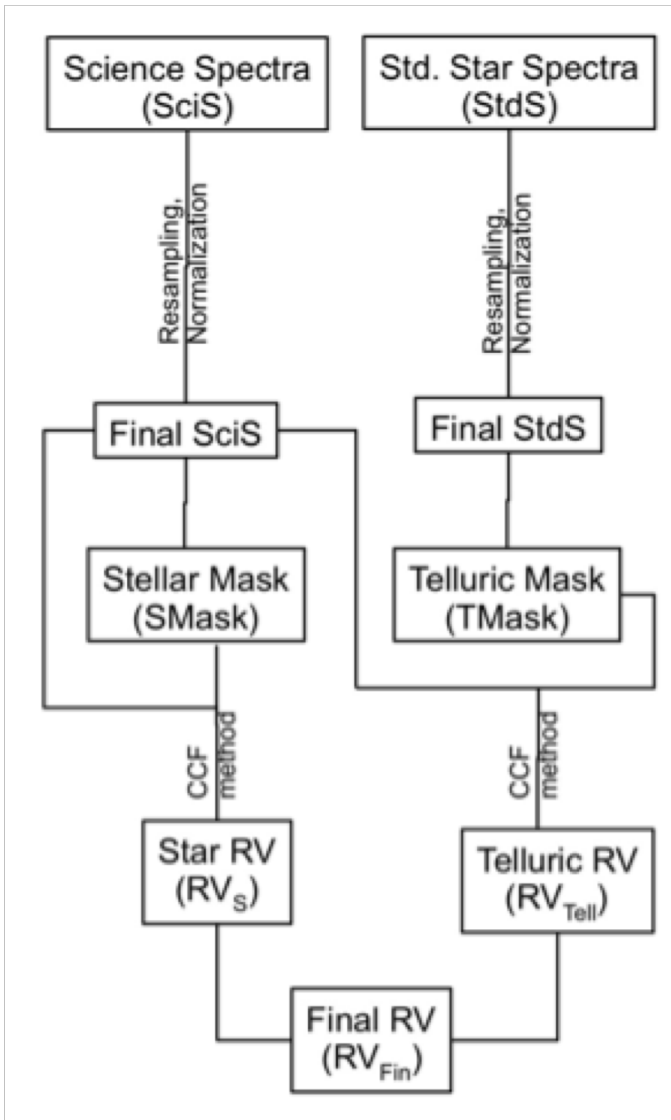
RV accuracy = 8m/s for Hmag = 5





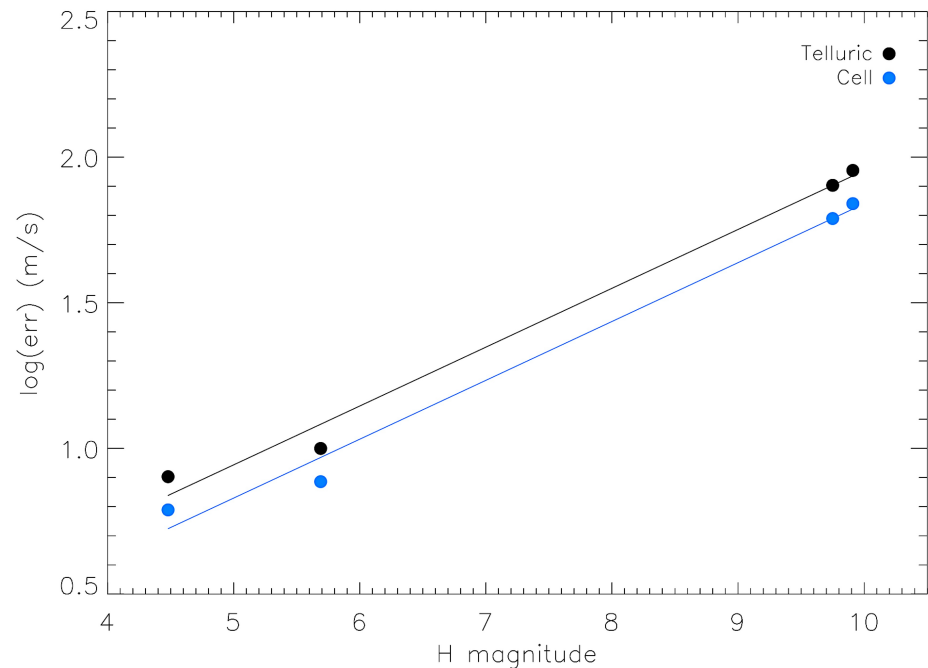
# RVs: telluric method

Carleo et al, 2016



RV accuracy = 8m/s for Hmag = 5

Absorption Cell: 3m/s



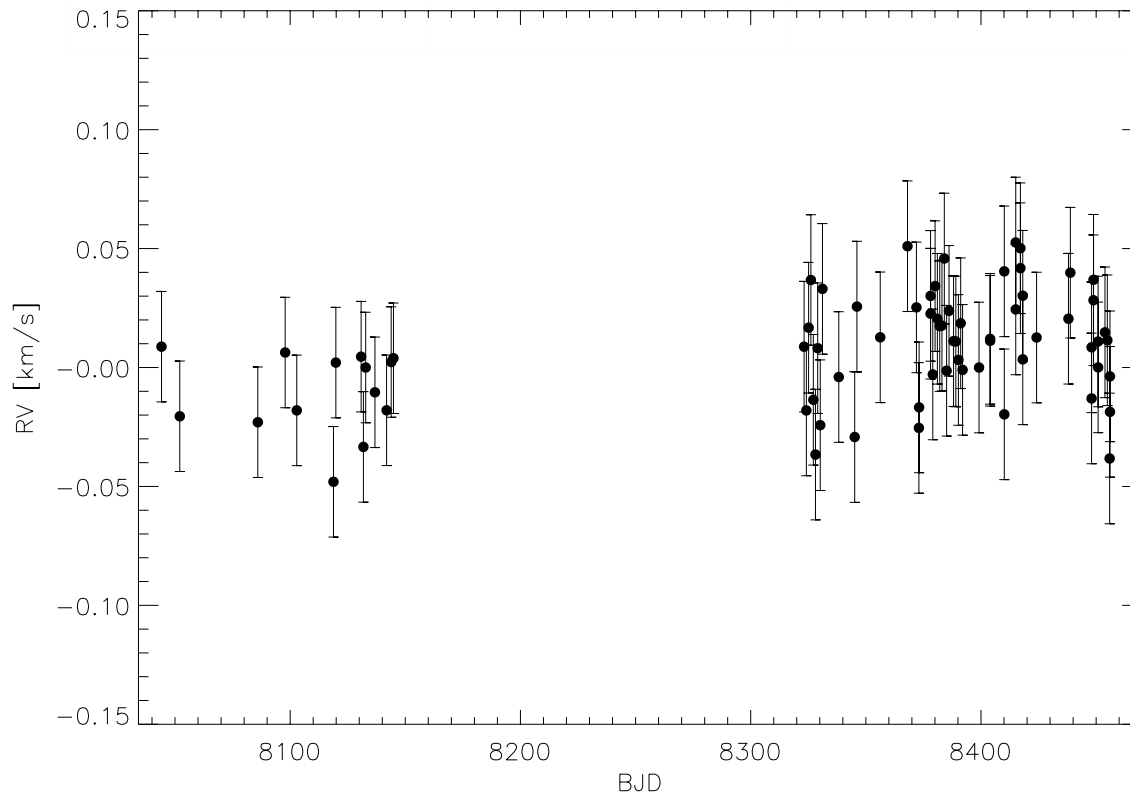




# RV standard star

- HD 3765 - SpT = K2

rms scatter (m/s)	Internal error (m/s)	Typical S/N
23	25	160

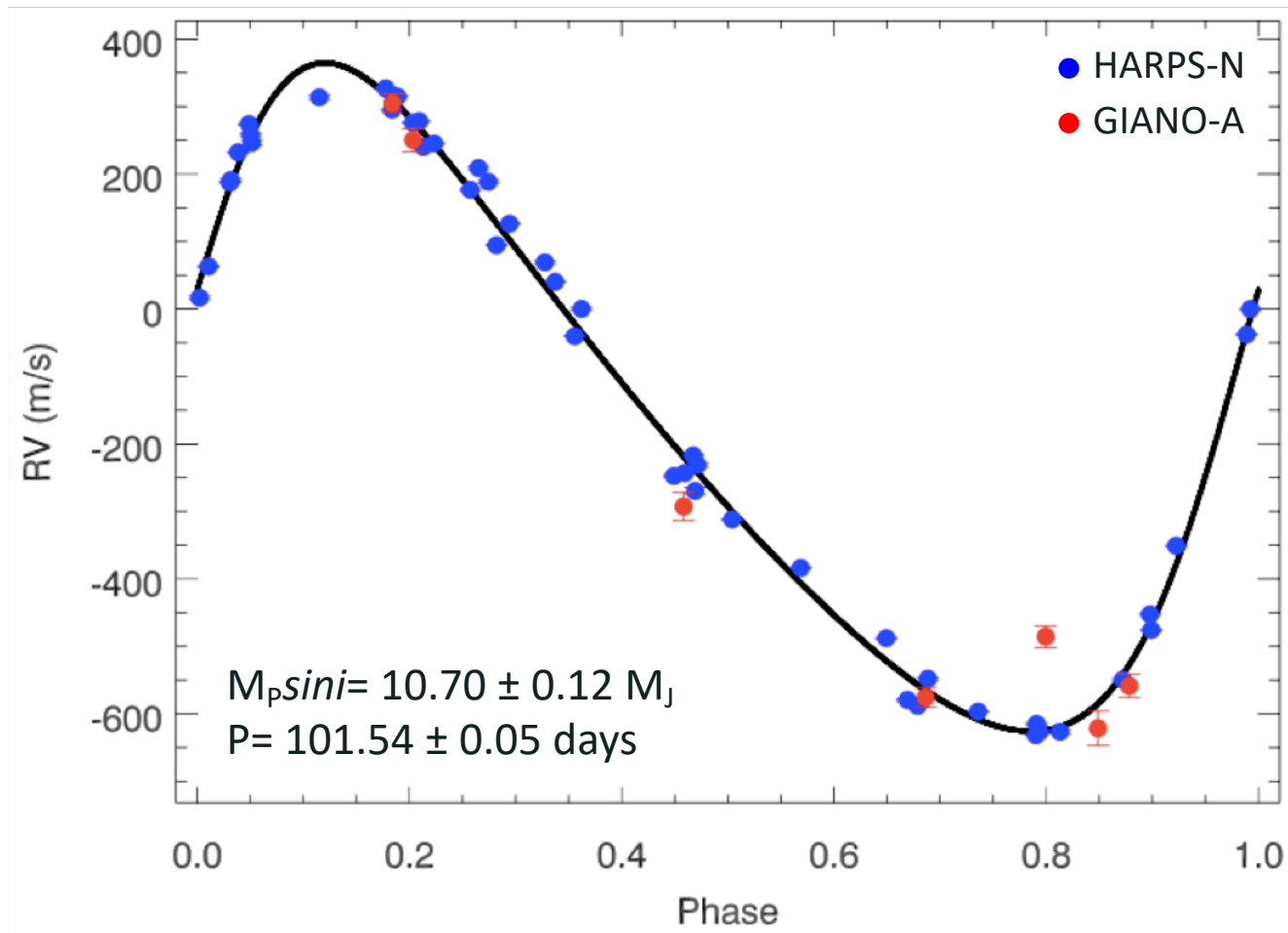




# TYC 4282-605-1

K-giant

González-Alvárez et al., 2017





Available to the scientific community from April 2017

**Semester Apr - Sep 2017:** GAPS makes use of its own observing time to further test GIARPS (science verification)

**Semester Oct - Mar 2017/2018:** GAPS2.0 Pilot Program

**Semester Apr - Sep 2018:** GAPS2.0 Large Program

**Objective: search and characterization of very young Hot Jupiters**

# GIARPS **BD+20 1790**

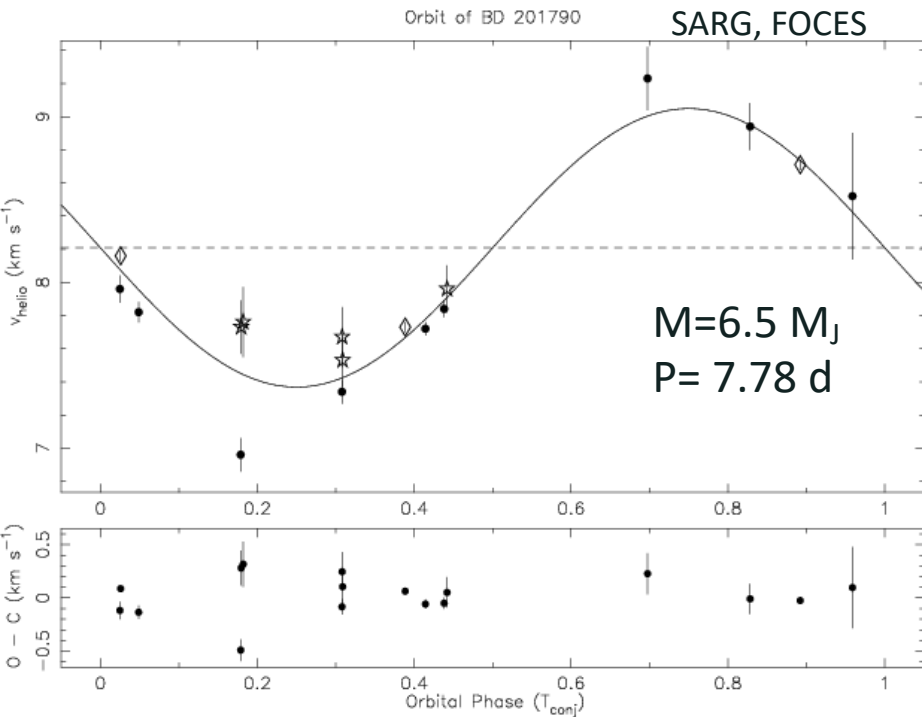
SpT = K5V      vsini= 10 km/s

$P_{\text{rot}} = 2.76 \text{ d}$

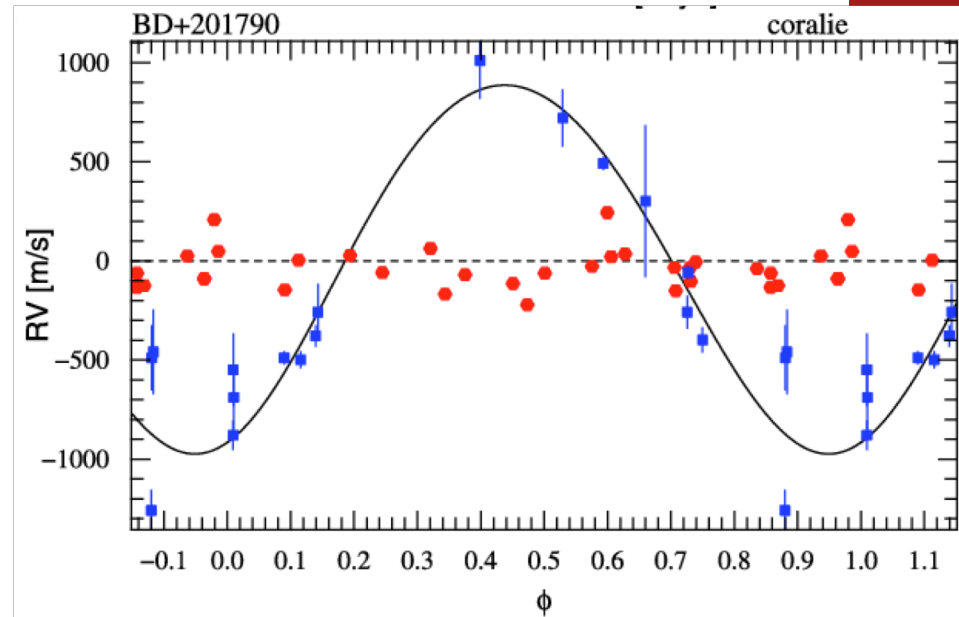
$\log R'_{\text{HK}} = -3.7$  (Hernán-Obispo+2015)

Age  $\approx 150 \text{ Myr}$  (Bell+2015)

Hernán-Obispo et al. 2010



Figueira et al. 2010



# GIARPS **BD+20 1790**

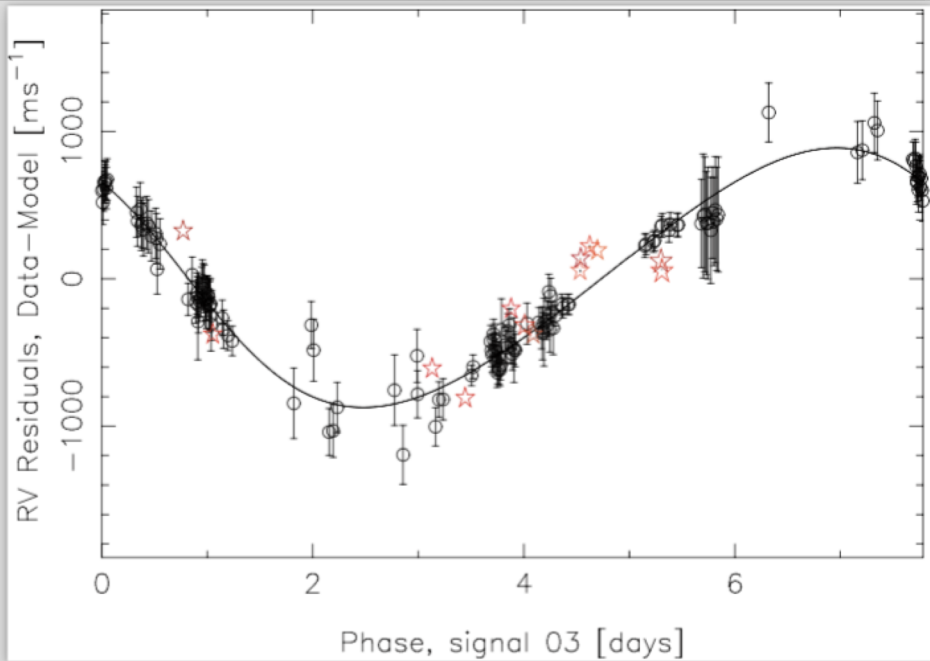
SpT = K5V      vsini= 10 km/s       $P_{\text{rot}} = 2.76 \text{ d}$

$\log R'_{\text{HK}} = -3.7$  (Hernán-Obispo+2015)

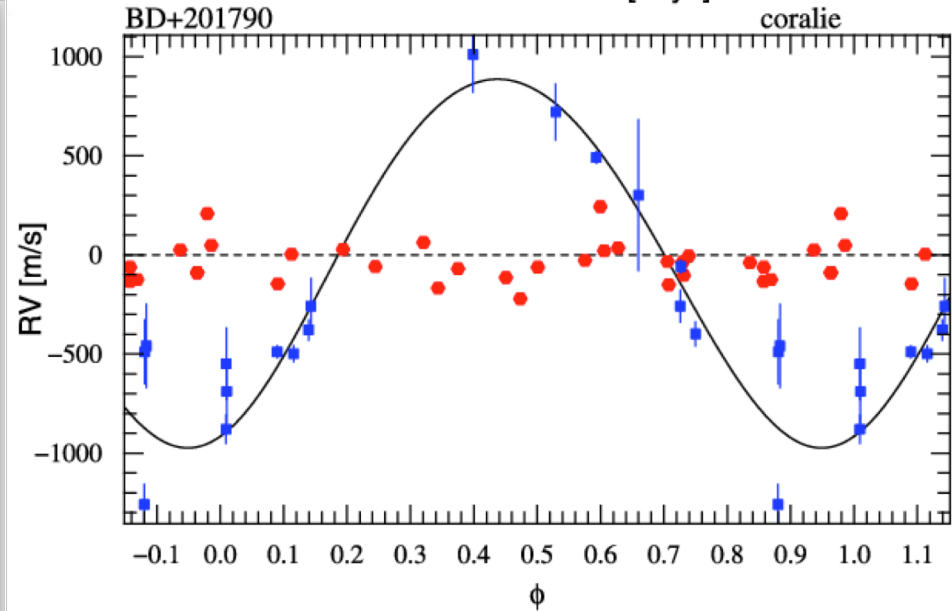
Age  $\approx 150 \text{ Myr}$  (Bell+2015)

Hernán-Obispo et al. 2015

HARPSN, HERMES

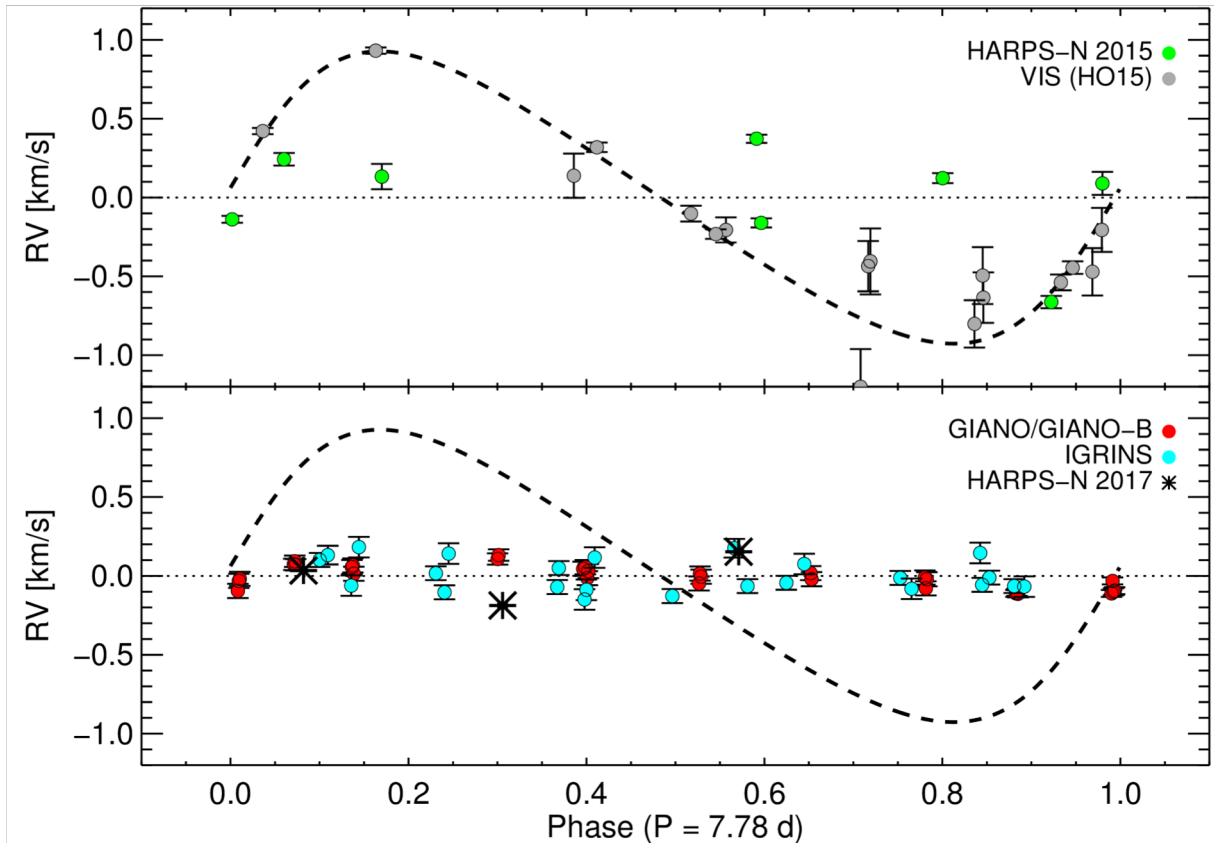


Figueira et al. 2010





# BD+20 1790 Carleo et al, 2018



Only previously known case orbiting a star between 20 to 200 Myr old (Rizzuto et al. 2017; David et al. 2018)



# YO13

Carleo et al, to be subm.

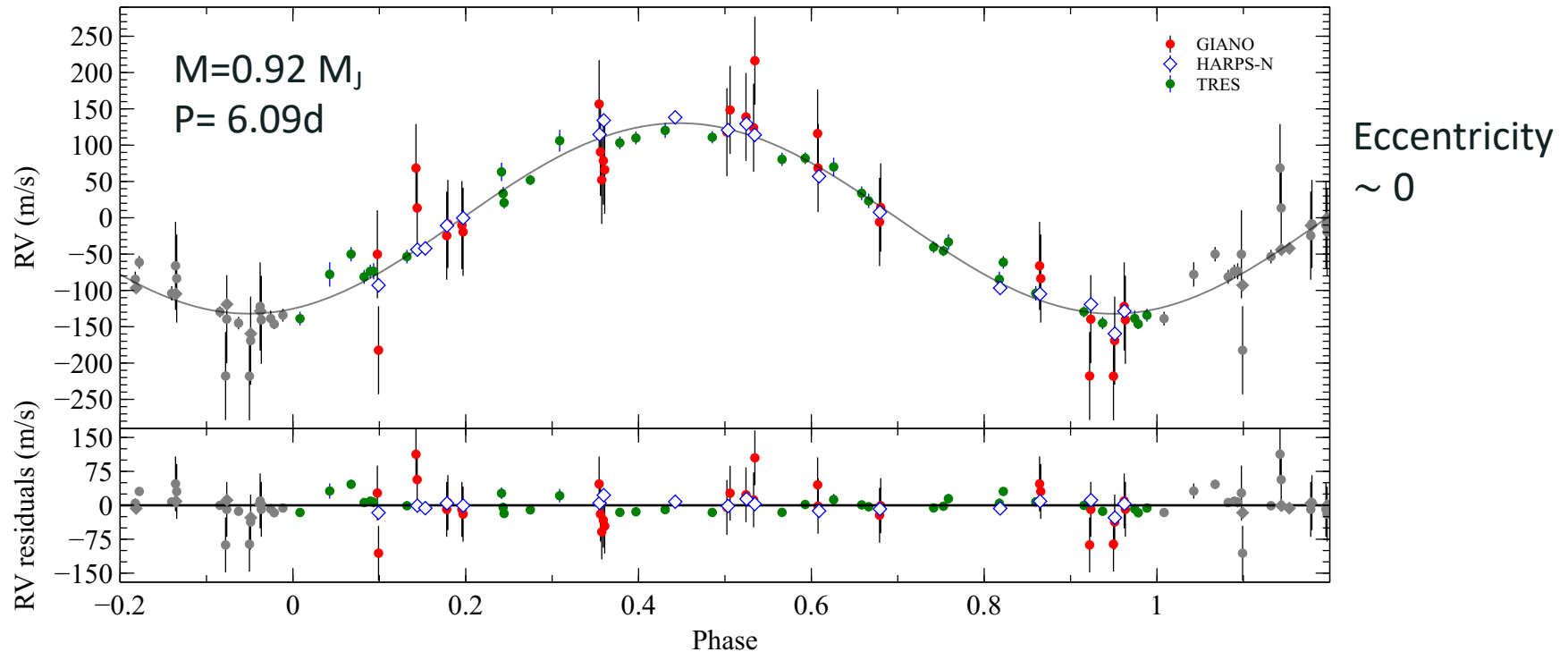
SpT = K5

$v \sin i = 3 \text{ km/s}$

$P_{\text{rot}} = 11.98 \text{ d}$

Age  $\approx 625 \text{ Myr}$

INSTRUMENT	rms scatter (m/s)	Internal error (m/s)	Typical S/N
GIANO - B	119	61	39
HARPS - N	105	3	42





# GIANO-B: RV precisions

SpType	Hmag	SNR	vsini (km/s)	Internal error (m/s)	RV r.m.s. scatter (m/s)
M4.5V	4.8	163	3.0	19	19
G1V	5.1	141	1.0	31	24
K2V	5.3	163	1.7	25	23
K5V	7.0	72	10.0	36	130
K4.5	7.8	39	3.0	61	40
M0	8.6	27	30.5	769	946

