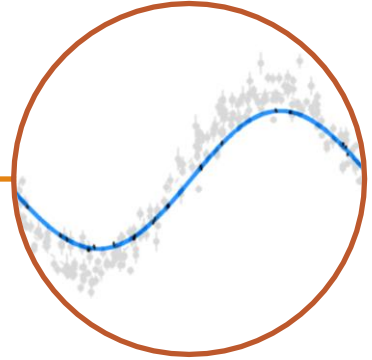
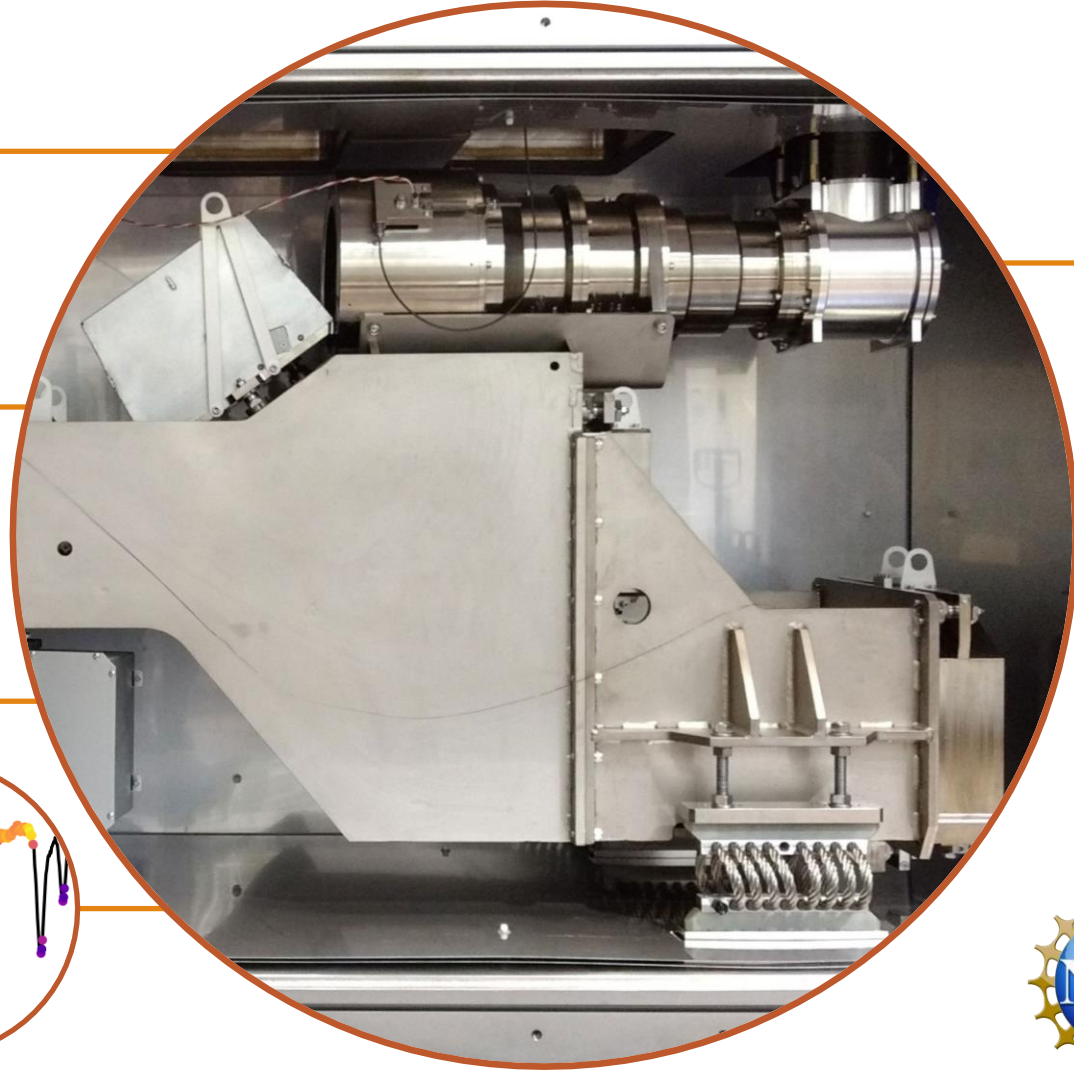
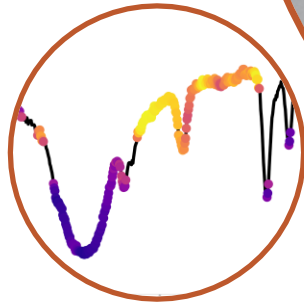
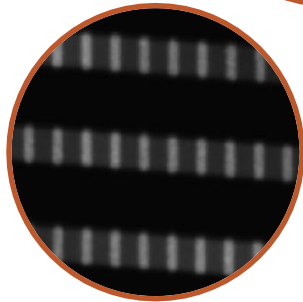
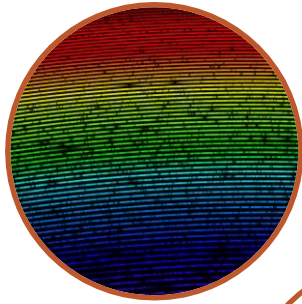


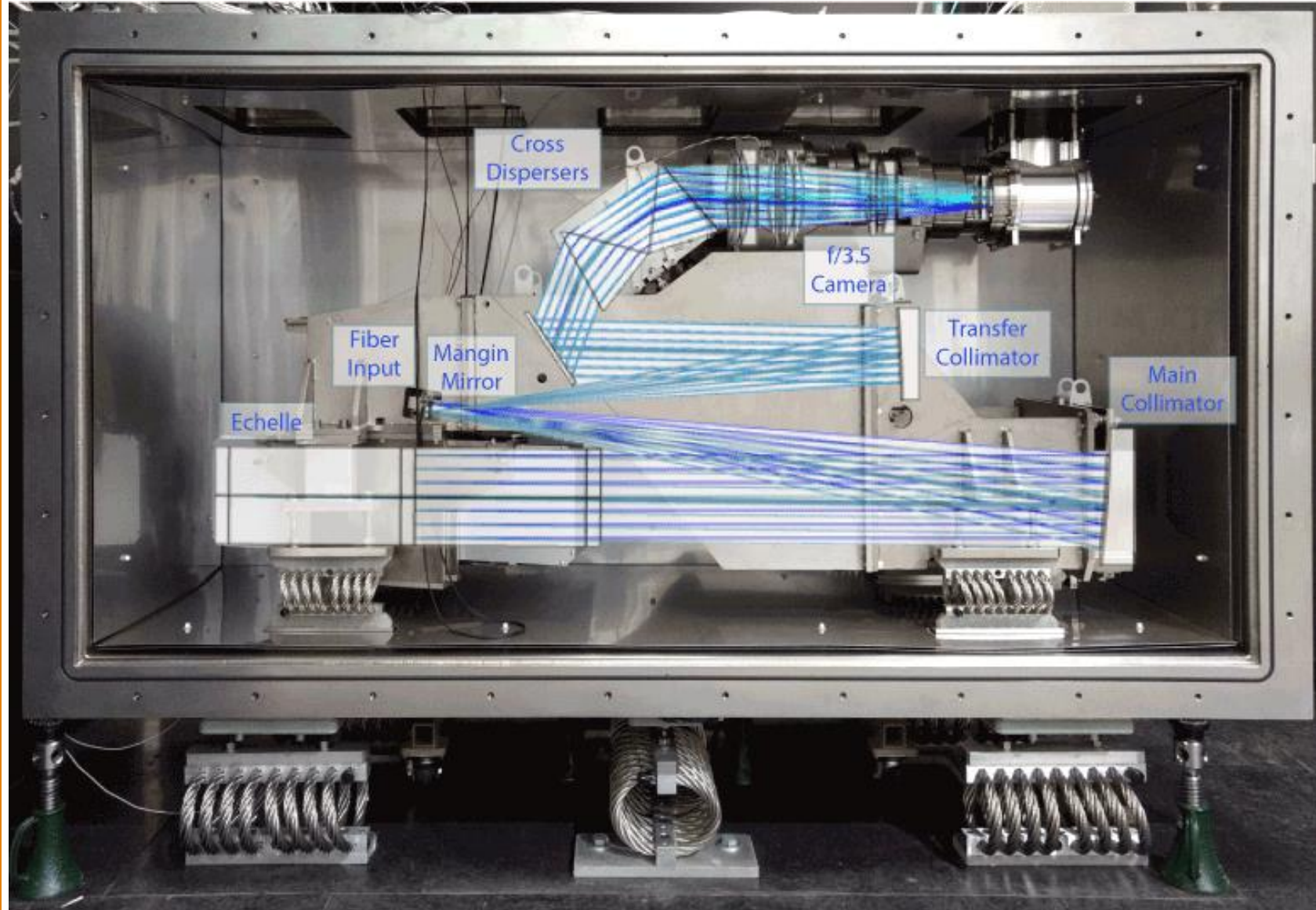
EXPRES | THE EXTREME PRECISION SPECTROGRAPH

Speaker: Lily Zhao
Debra Fischer
Ryan Blackman
John Brewer
Lars Buchhave
Jessi Cisewski
Allen Davis
Parker Holzer
Colby Jurgenson
Tyler McCracken
Bo Ning
Joel Ong
Ryan Petersburg
Dave Sawyer
Andrew Szymkowiak
René Tronsgaard
Xin Xu

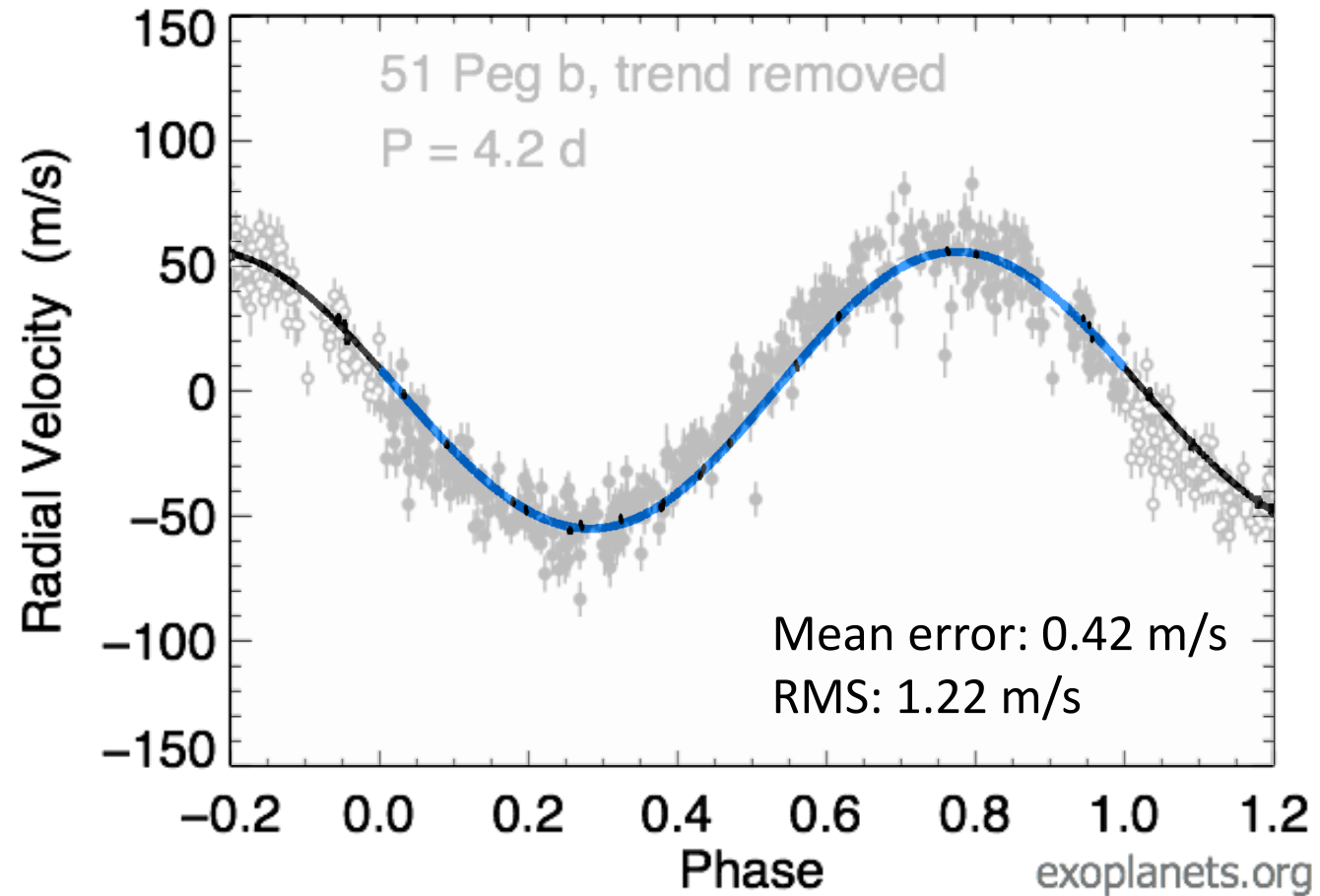


EXPRES by the Numbers

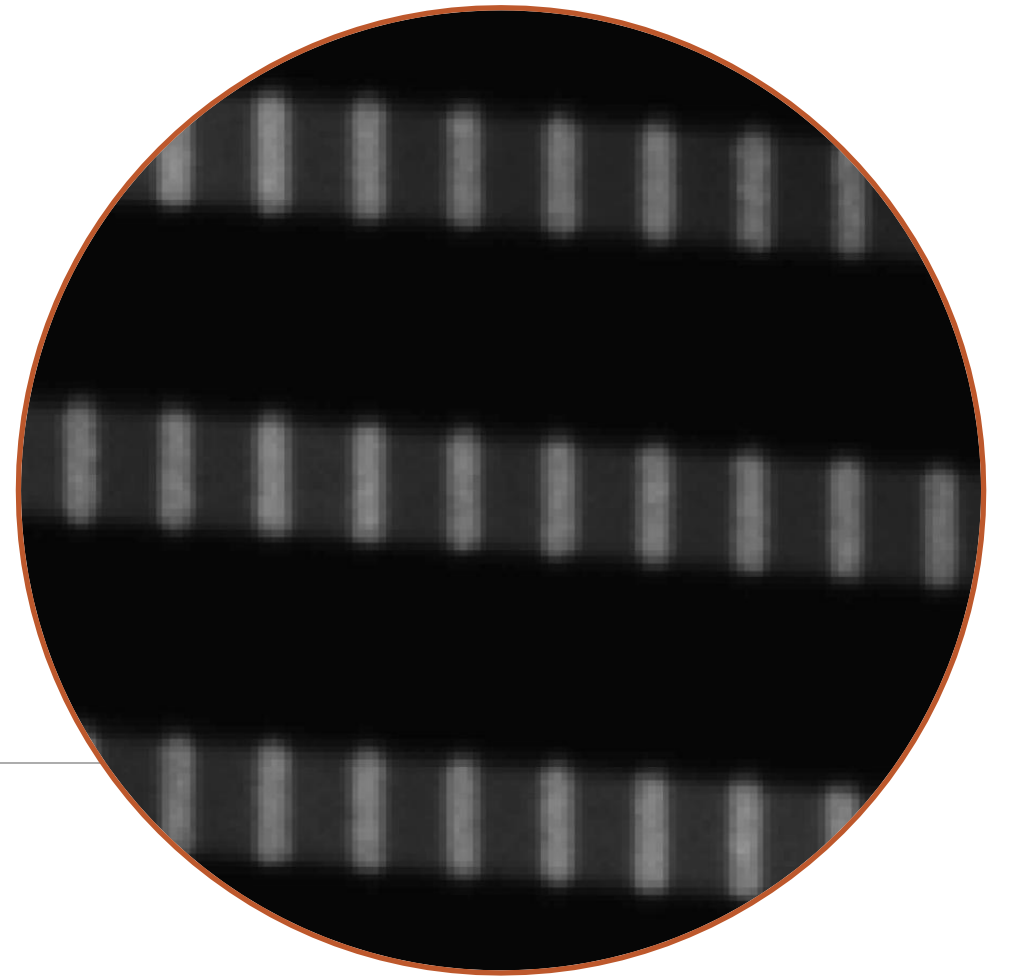
- Environmentally stabilized
- $R \sim 150,000$
- 390-700 nm
- DCT (4.3 m) in Flagstaff, AZ
- Up to 280 partial nights/year
- Comissioned in 2018



Observations of 51 Peg

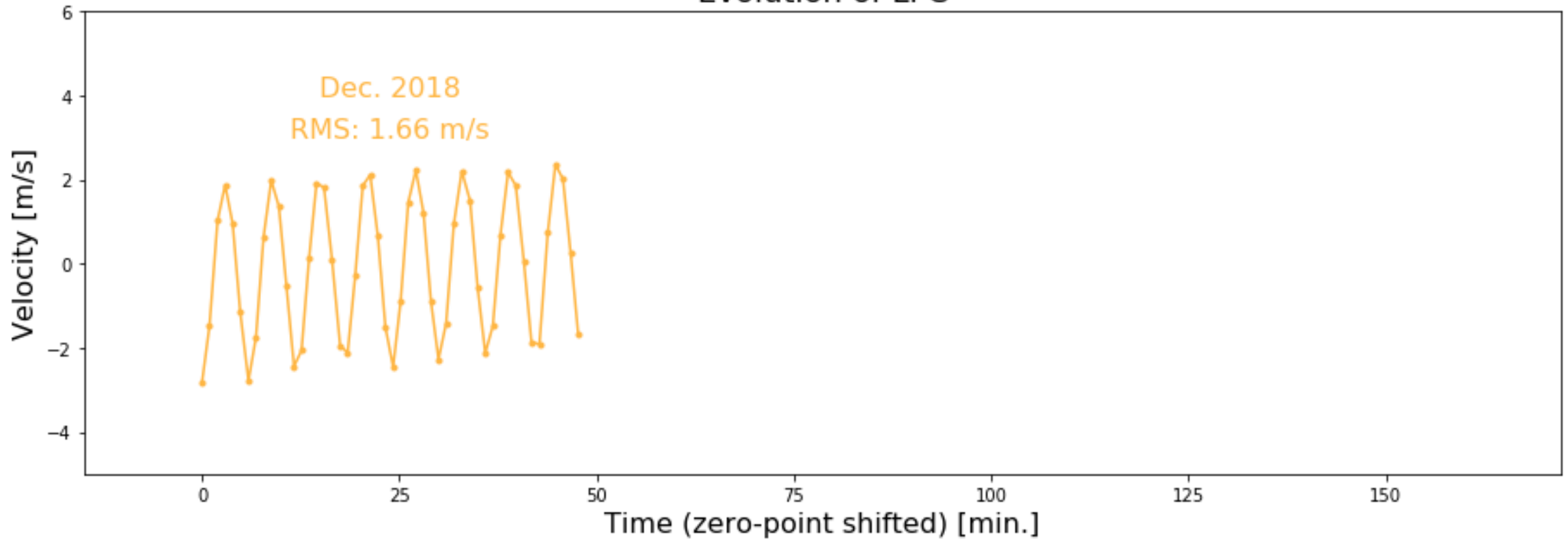


LFC Variability



Consecutive LFCs consistently showed a sinusoidal variation with a period of ~ 6 min.

Evolution of LFC



The ~6 min variation was due to a beat frequency.

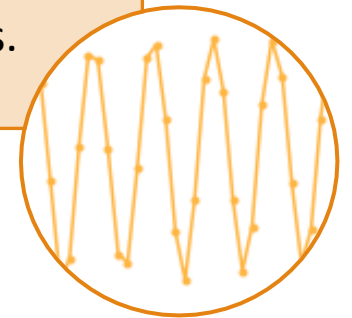
FREQUENCY ONE

The CCD's CryoTel cooler's active vibration correction system

FREQUENCY TWO

The LFC's Spatial Light Modulator (SLM) using a pulse width modulation technique
Frequency: ~120 Hz

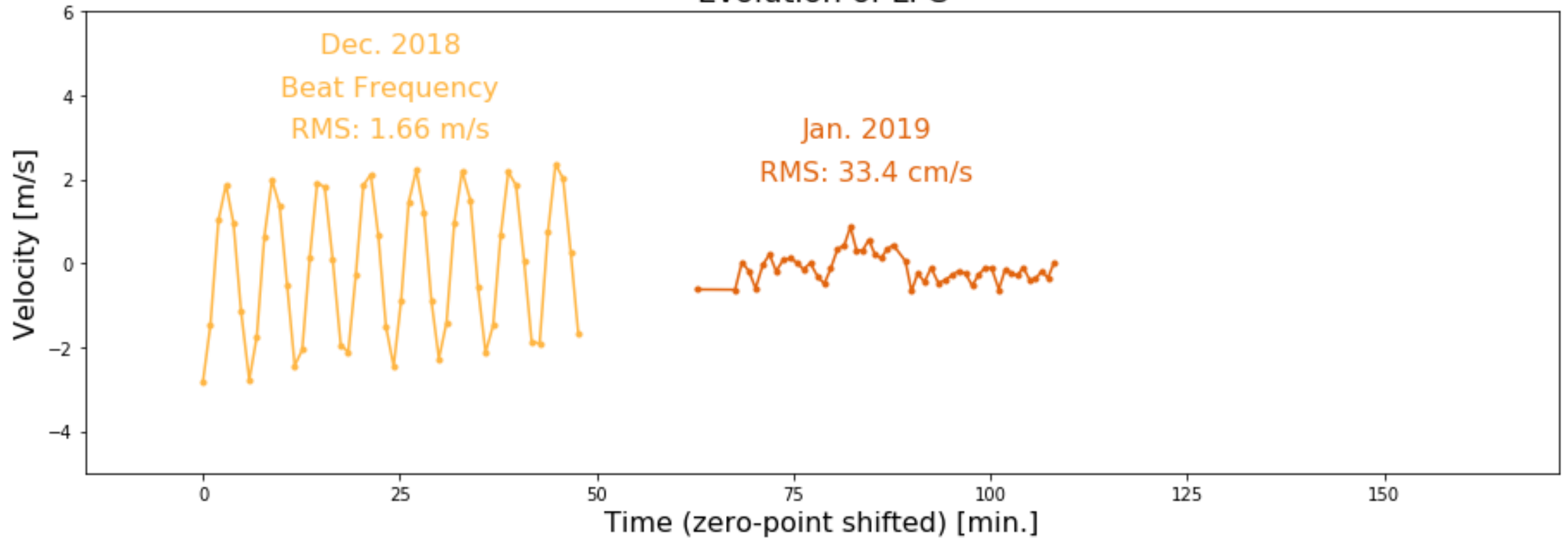
Beat frequency with a period of ~6 min.
and an amplitude of ~1.2 m/s.



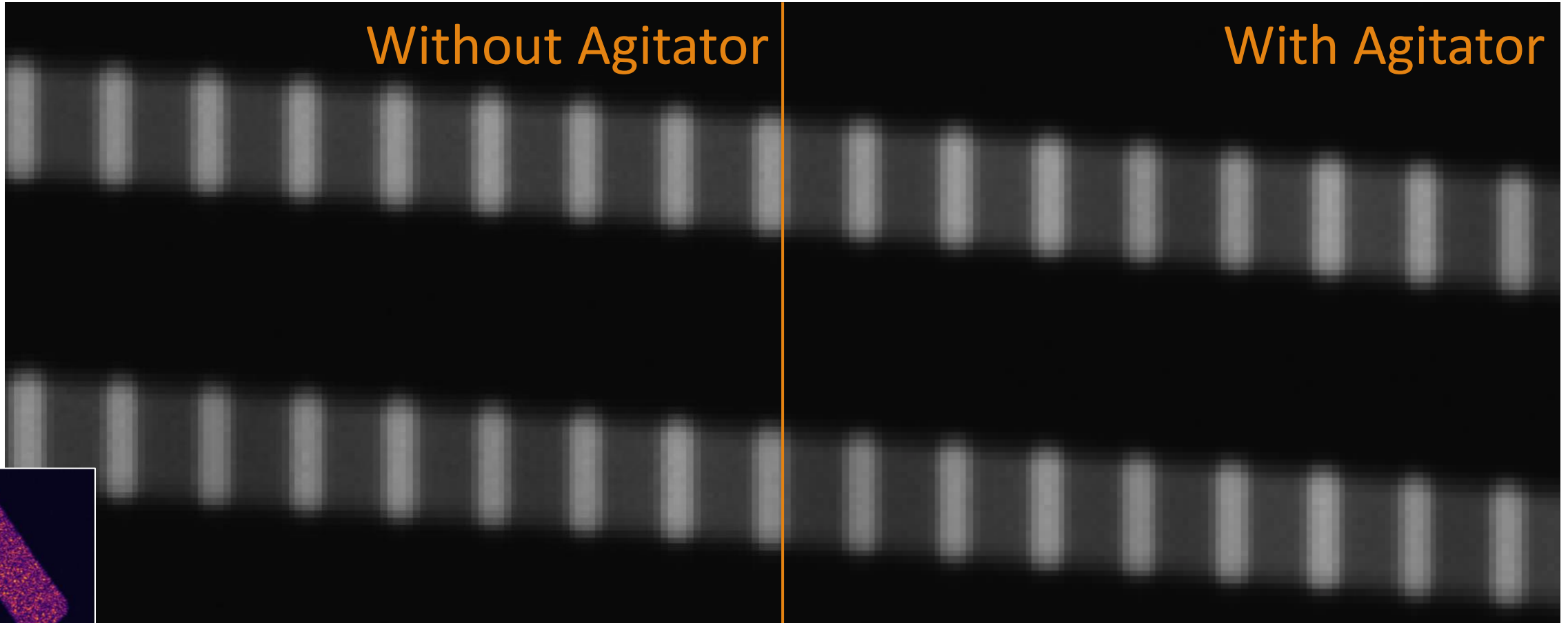
Changed the LFC's flicker to ~110 Hz

A ~ 30 cm/s variation persisted.

Evolution of LFC

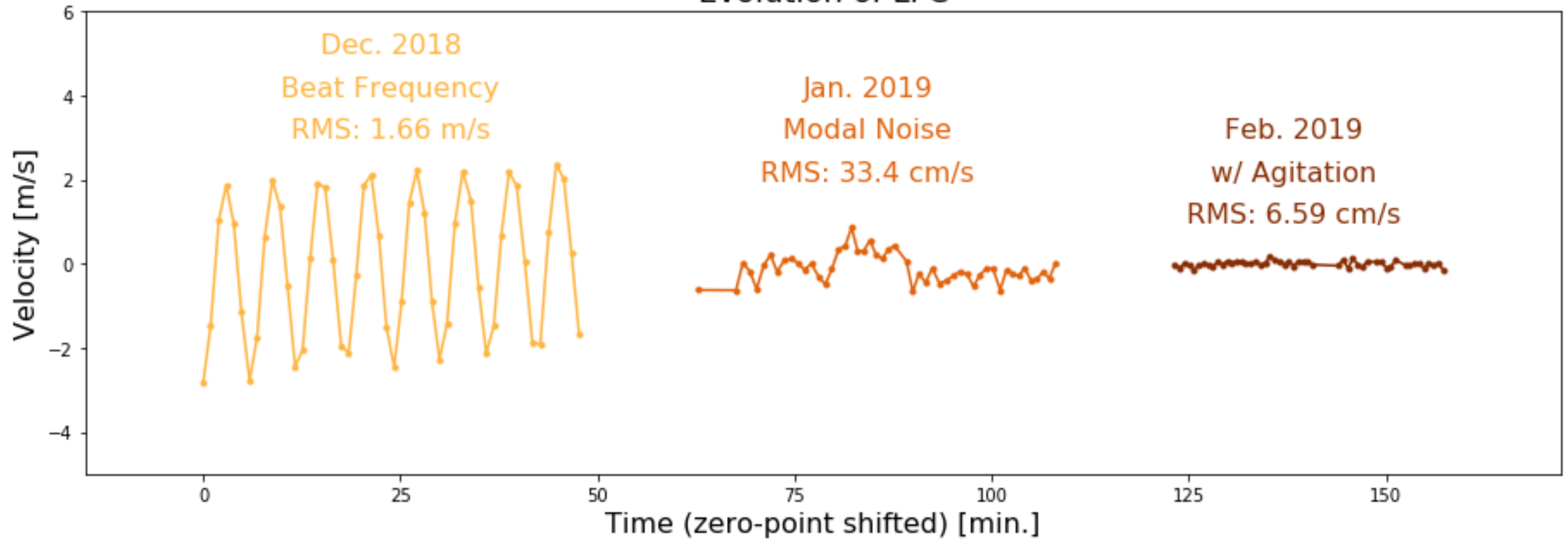


Modal noise was actually visible in the LFC lines.
Our agitator had stopped working.

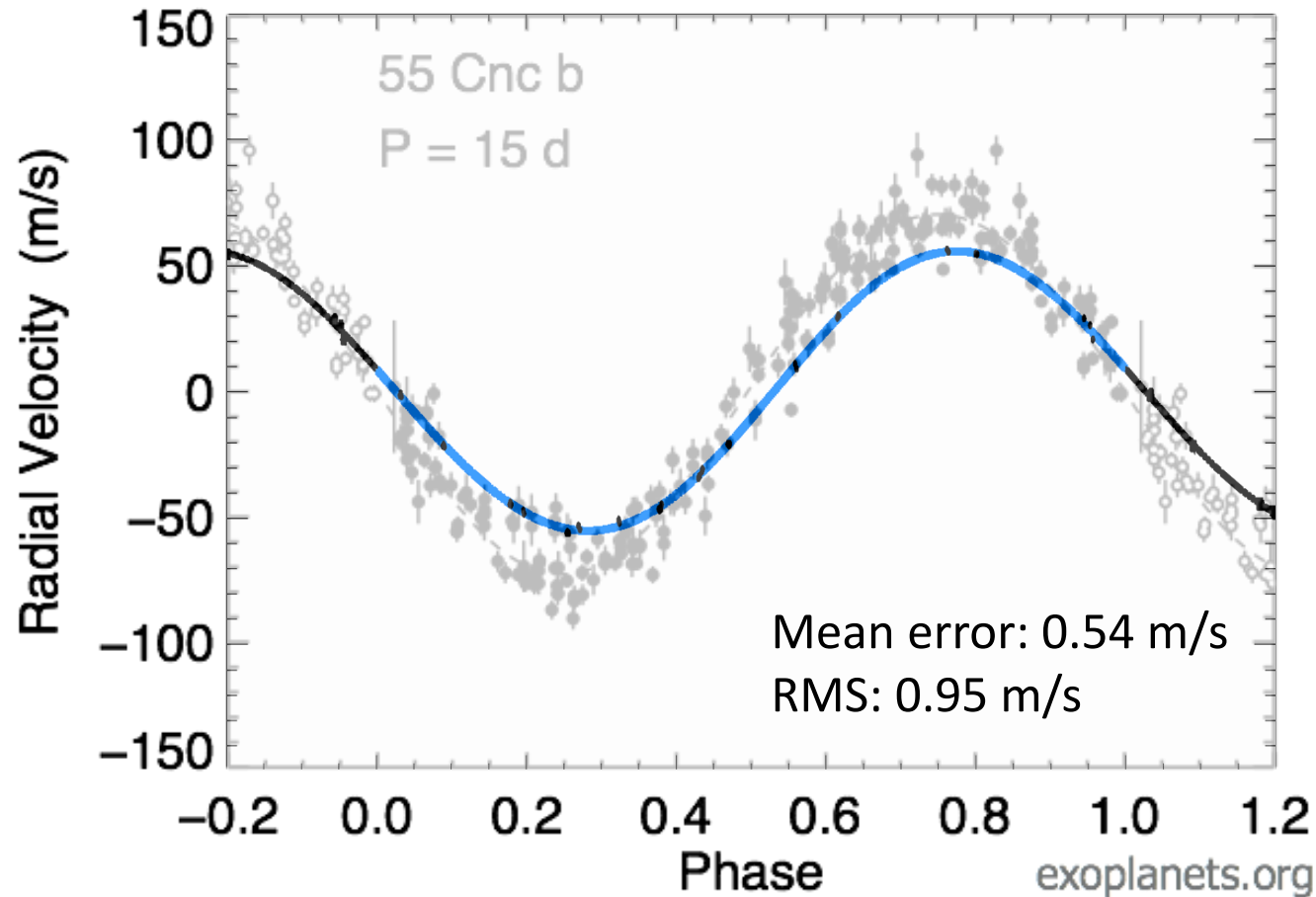


LFC measurements now show an instrument stability of 5-7 cm/s.

Evolution of LFC



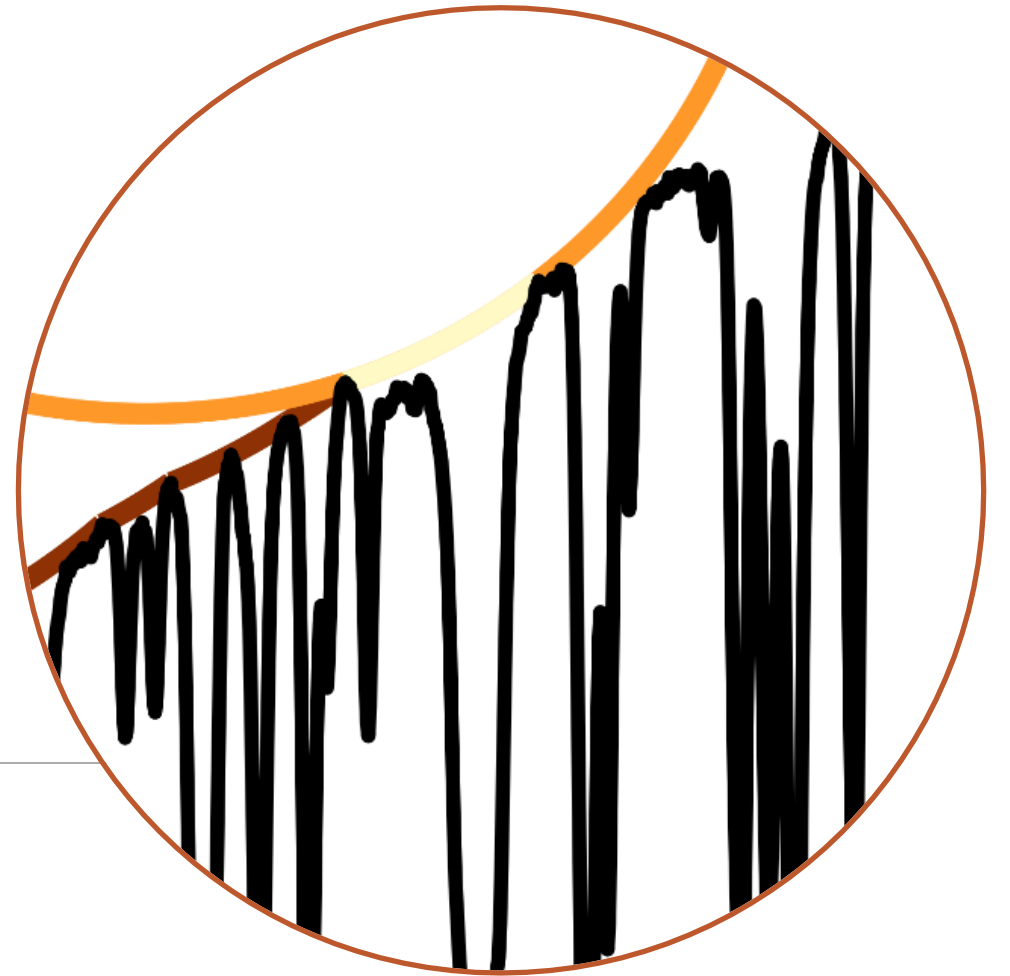
Observations of 55 Cnc



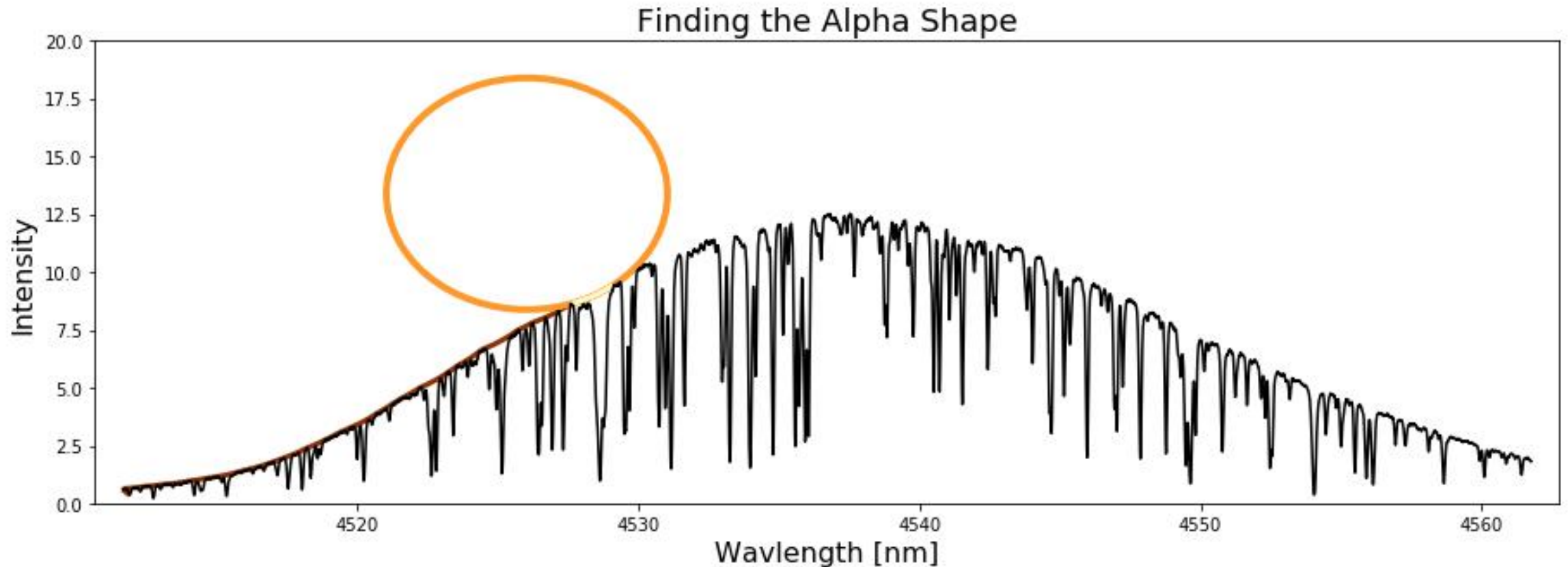
Results still include stellar noise and are affected by transit data.

Note: We calculate a lower mass of 232 Earth masses (vs. 253) though the exact same period.

Convex Hull Blaze Removal

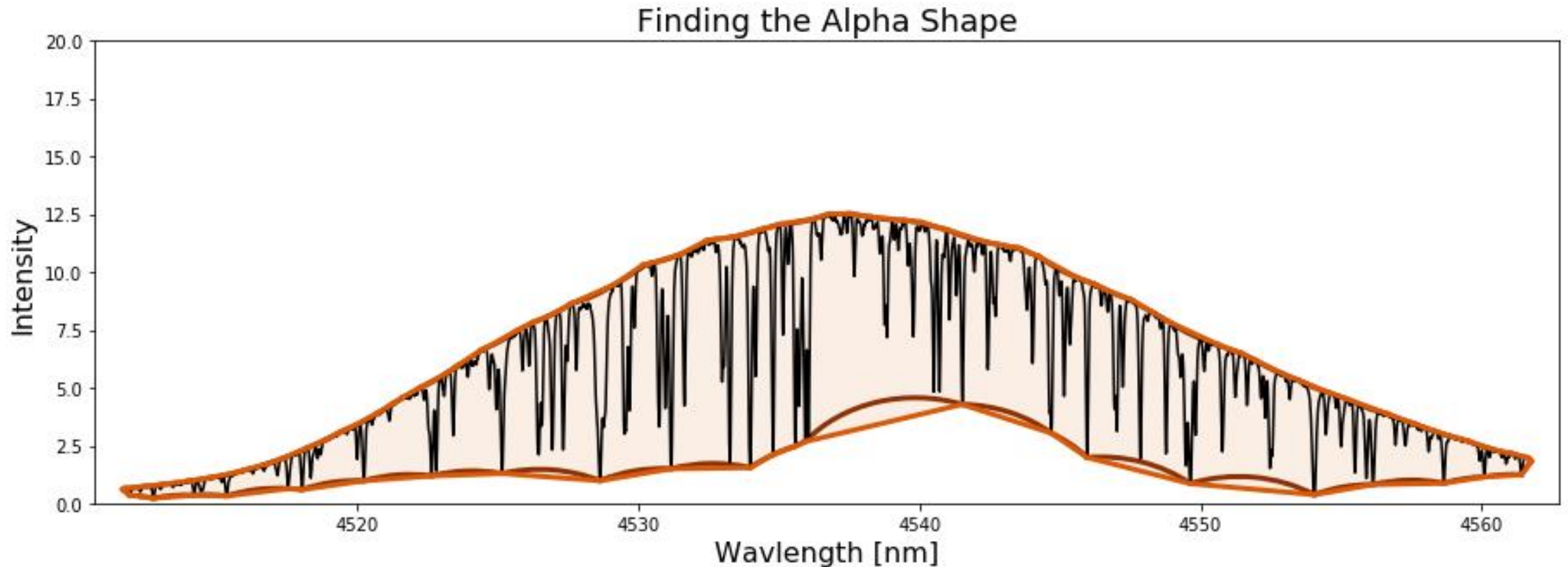


The continuum/blaze function can be well fit by constructing an α shape of each order.



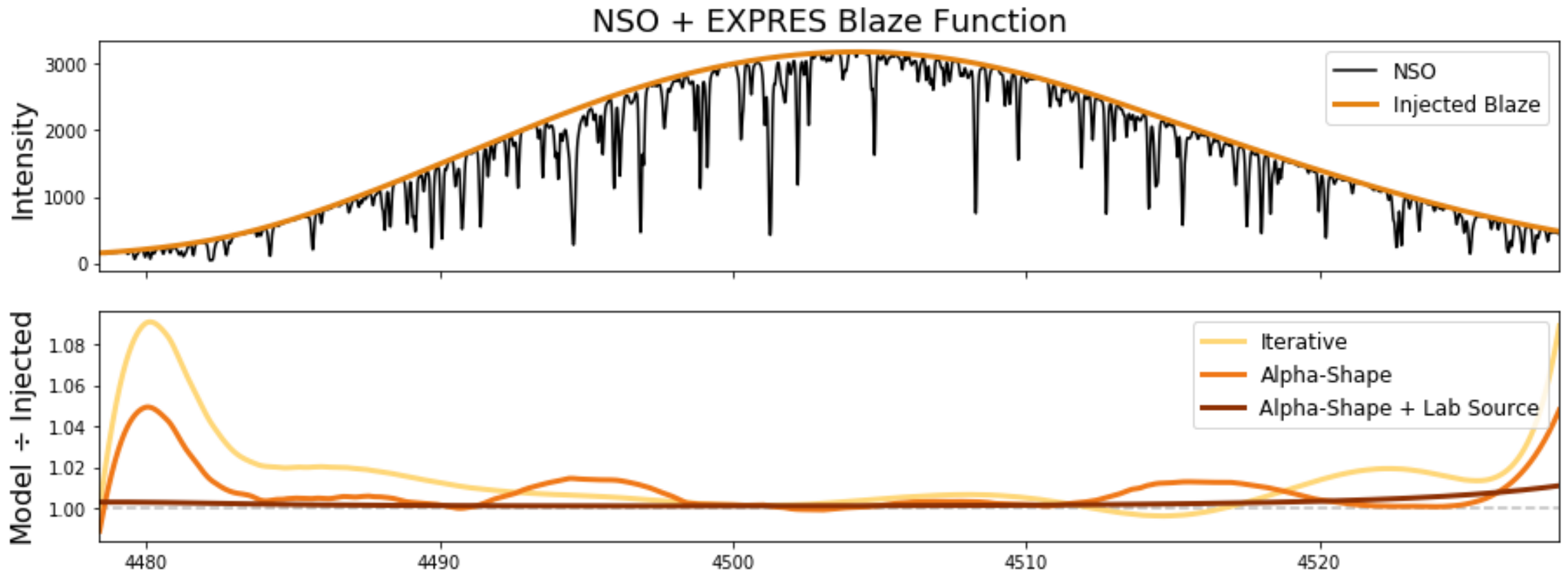
Xu et al. (submitted)

The continuum/blaze function can be well fit by constructing an α shape of each order.



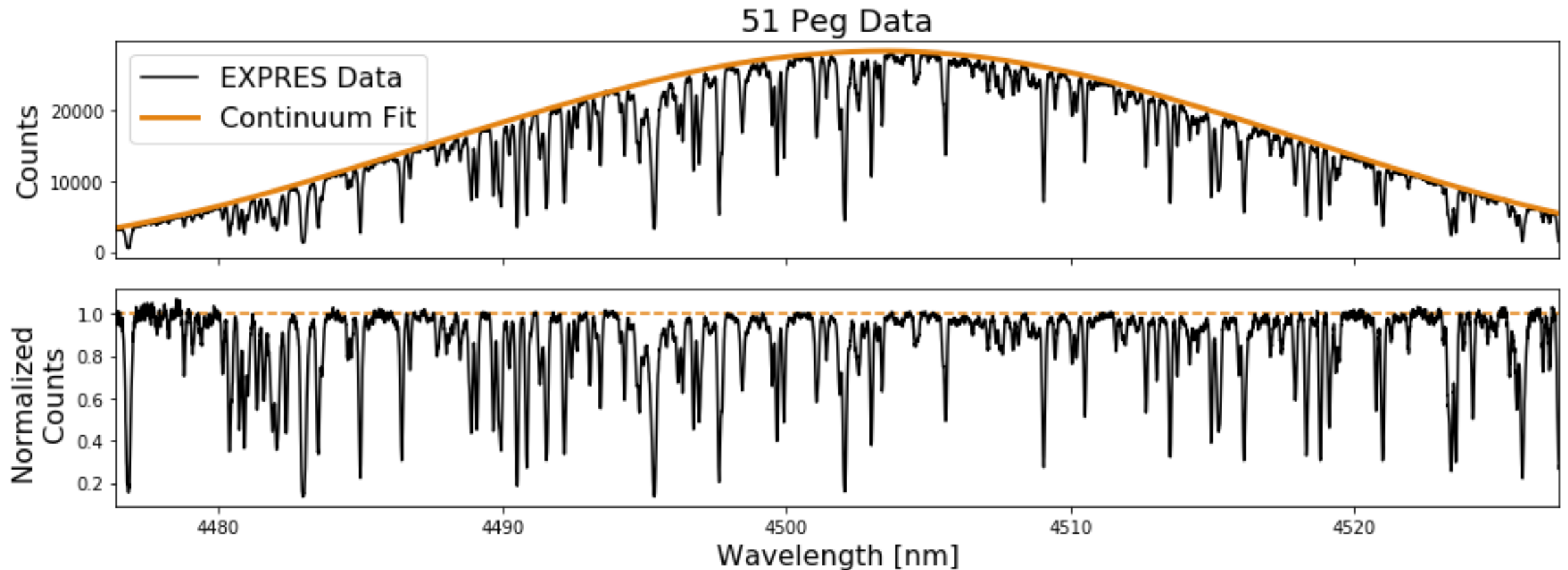
Xu et al. (submitted)

Using the α shape returns flatter results than an iterative, continuum fitting method.



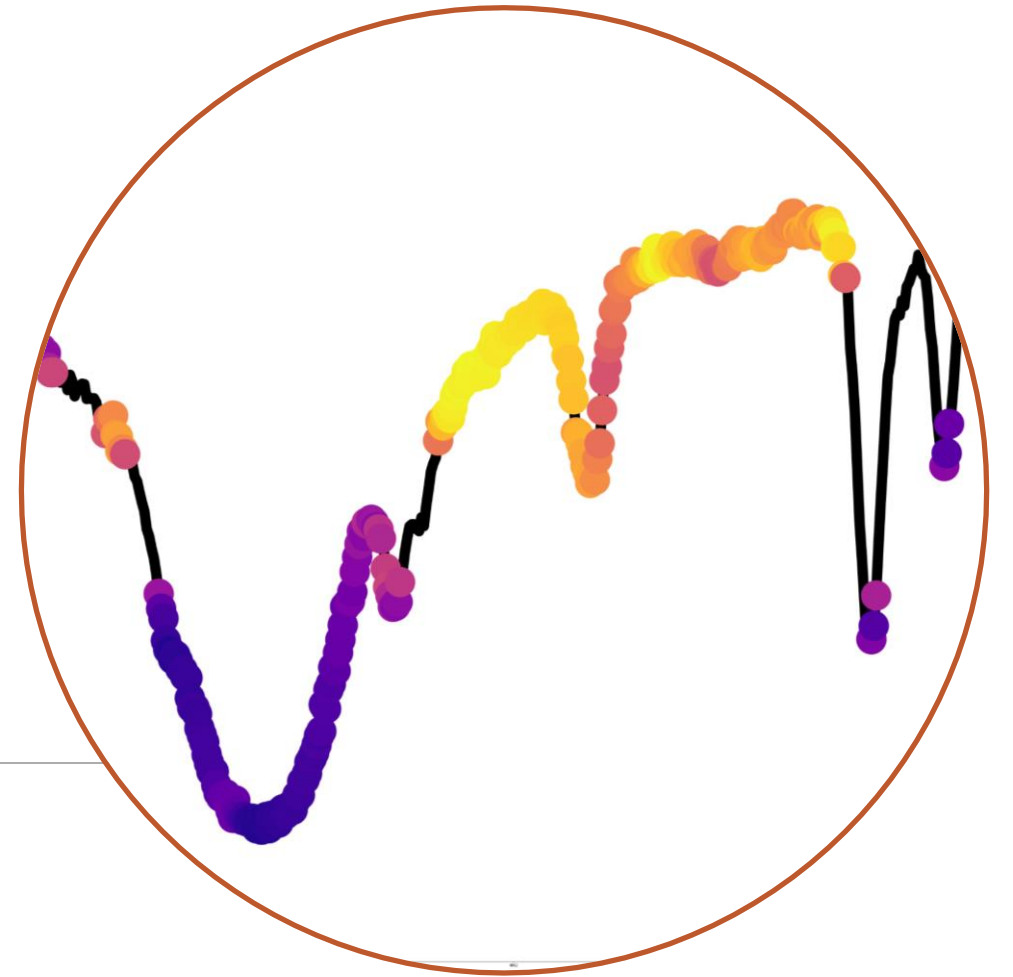
Xu et al. (submitted)

The α shape method is shown to work well on actual EXPRES data.

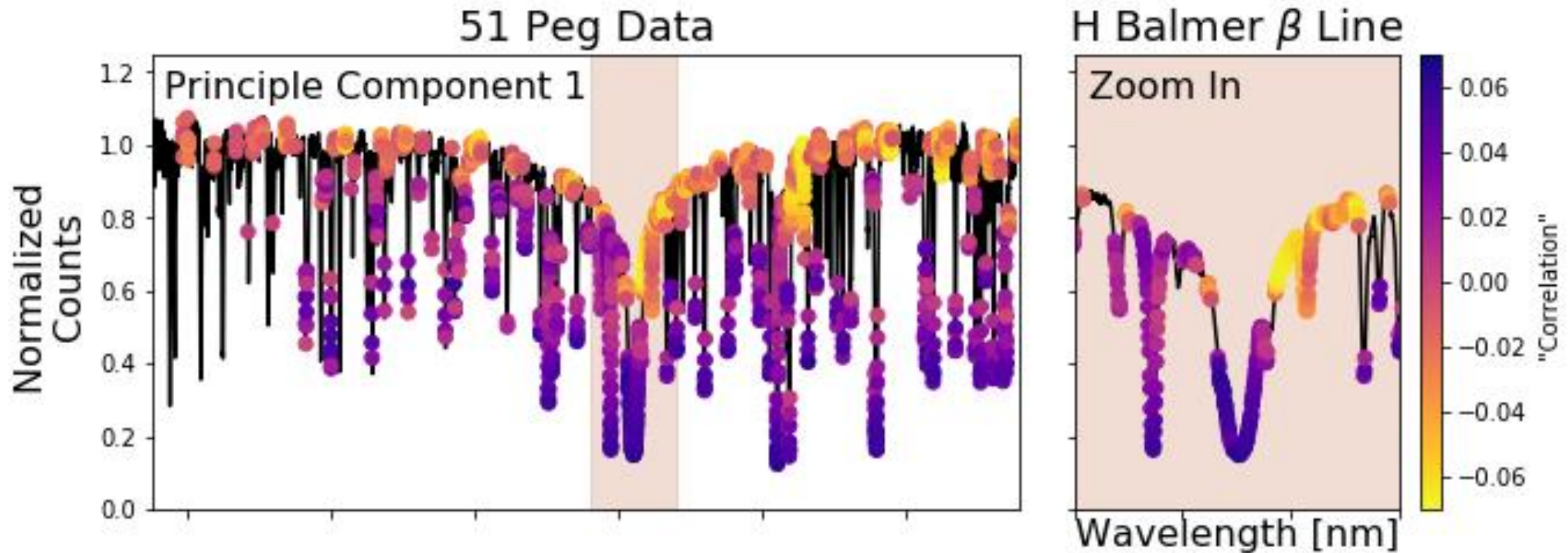


Xu et al. (submitted)

Line Variability w/ Sparse PCA



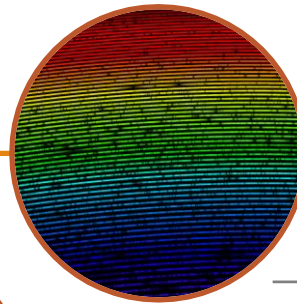
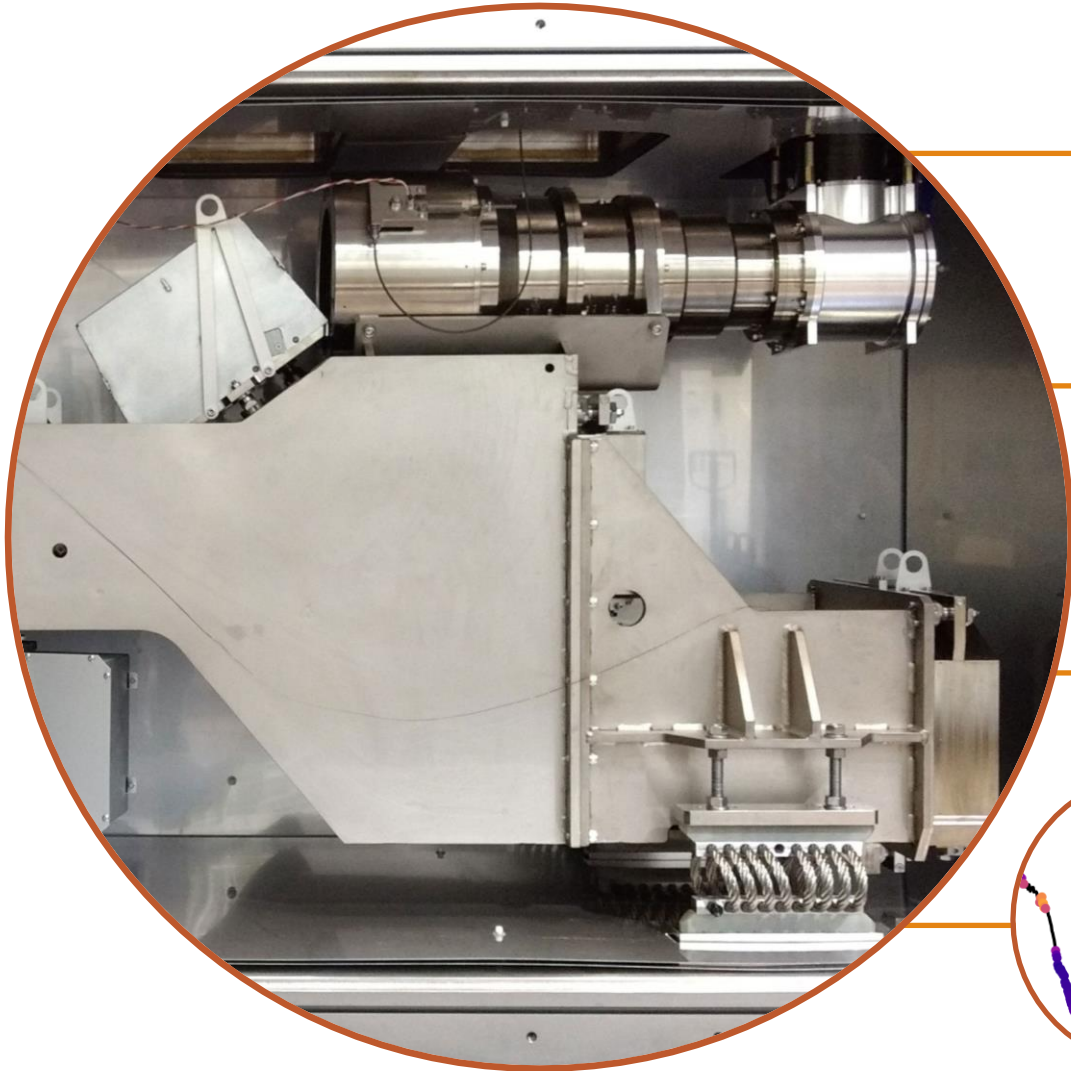
We are using sparse PSA to identify lines with variability that may be indicative of stellar activity.



Ning et al. (in prep.)

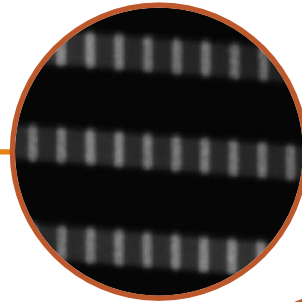
EXPRES | THE EXTREME PRECISION SPECTROGRAPH

SUMMARY



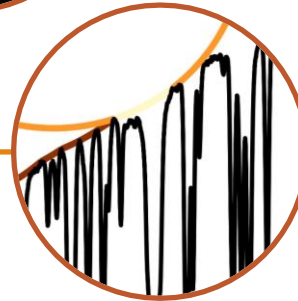
EXPRES is running smoothly and returning sub-meter precision data.

(Ong +)



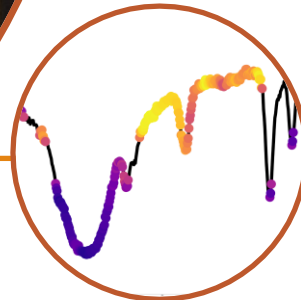
A ~ 6 min. beat frequency arose between the LFC SLM and the CCD CryoTel, creating a ~ 1 m/s variability.

(Symkowiak +)



Using the α shape of each order allows for better continuum fitting than an iterative model.

(Xu +)



Sparse PCA is promising for identifying lines with greater variability, likely due to stellar activity.

(Ning +)