

TO WHICH PRECISION CAN WE CORRECT TELLURICS IN THE NIR?

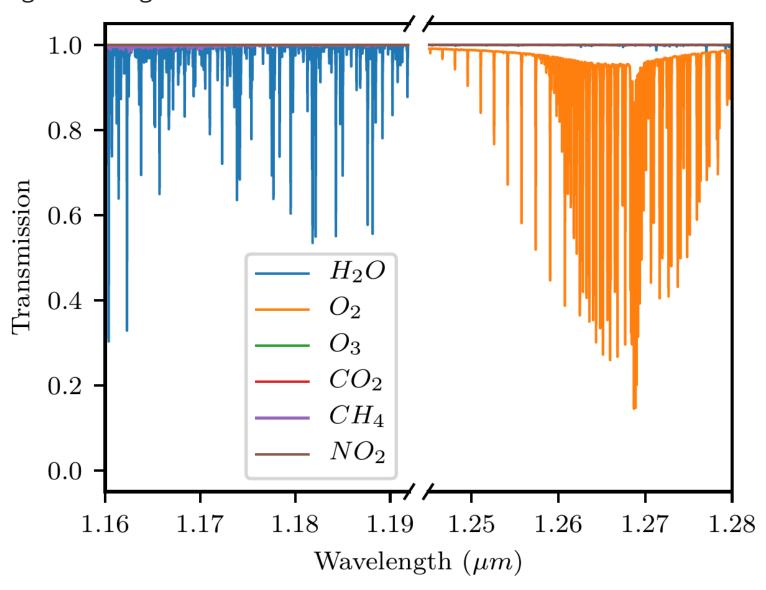


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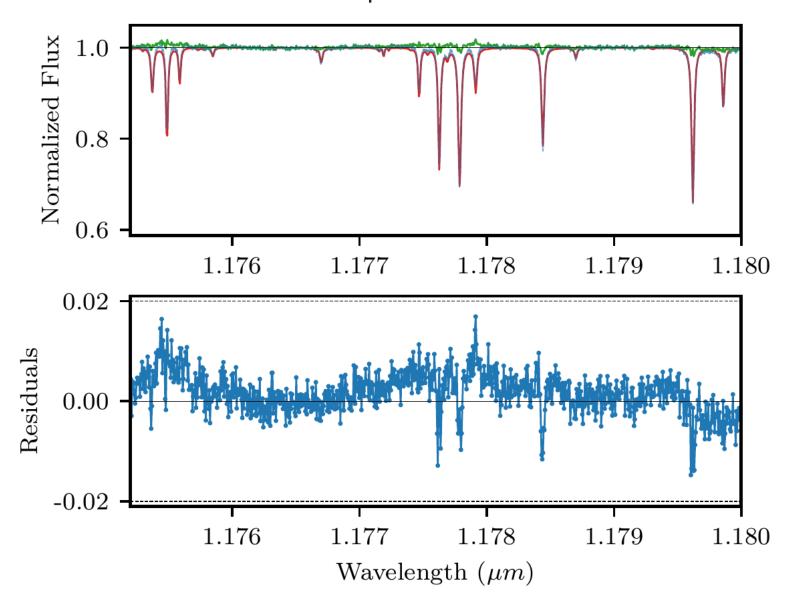
TELLURIC ABSORPTION

Synthetic transmission spectrum of our atmosphere (LBLRTM/TAPAS) Wavelength coverage of our CRIRES data set



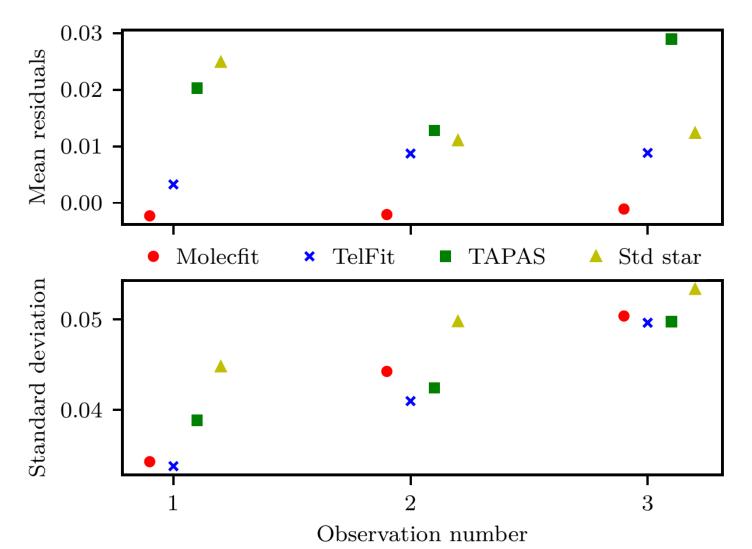
MOLECFIT TELLURIC CORRECTION

CRIRES spectrum of a hot standard star in the J band. Residuals below 2%. Residual patterns could be removed with PCA.

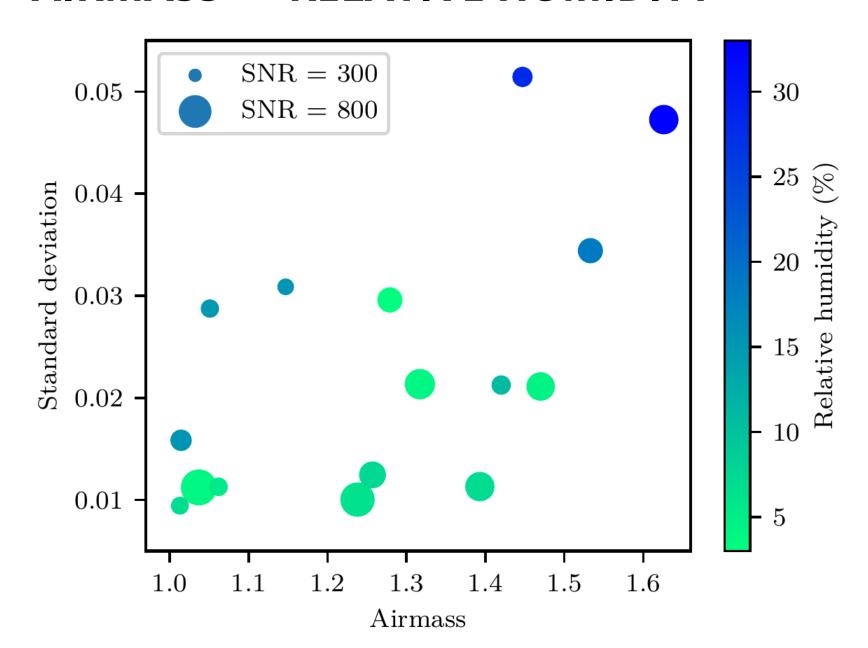


OFFSET AND SCATTER TELLURIC CORRECTION METHODS

Residuals inside the **water telluric lines** for Molecfit, TelFit, TAPAS, and the standard star method around 1.18 microns.

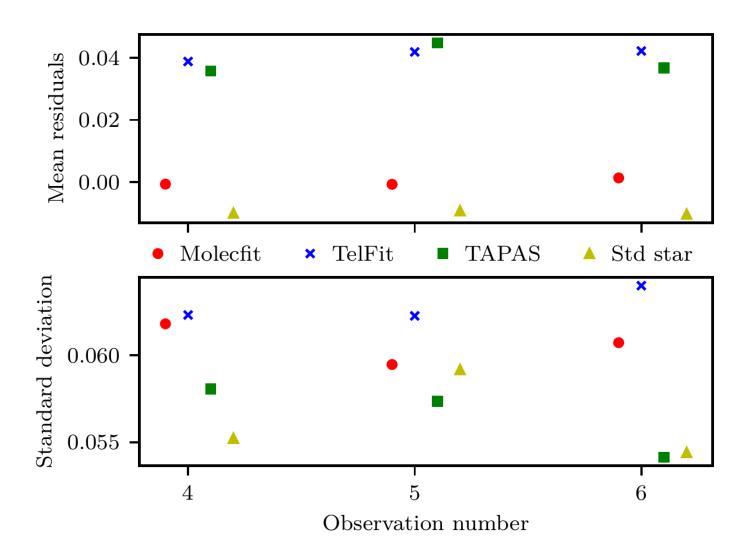


SCATTER EVOLUTION with AIRMASS and RELATIVE HUMIDITY



COMPARISON TELLURIC CORRECTIONS

Residuals inside the telluric lines in the wavelength range dominated by **oxygen absorption**.

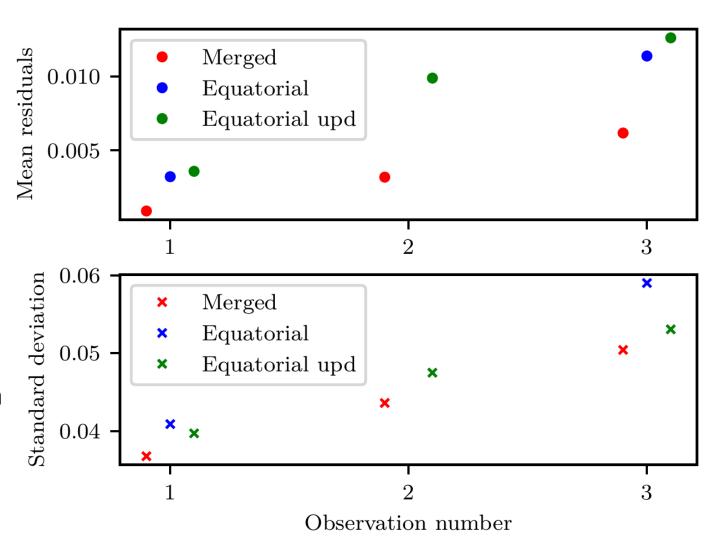


IMPORTANCE ATMOSPHERIC PROFILE

Merged profile: EMM + GDAS + averaged profile

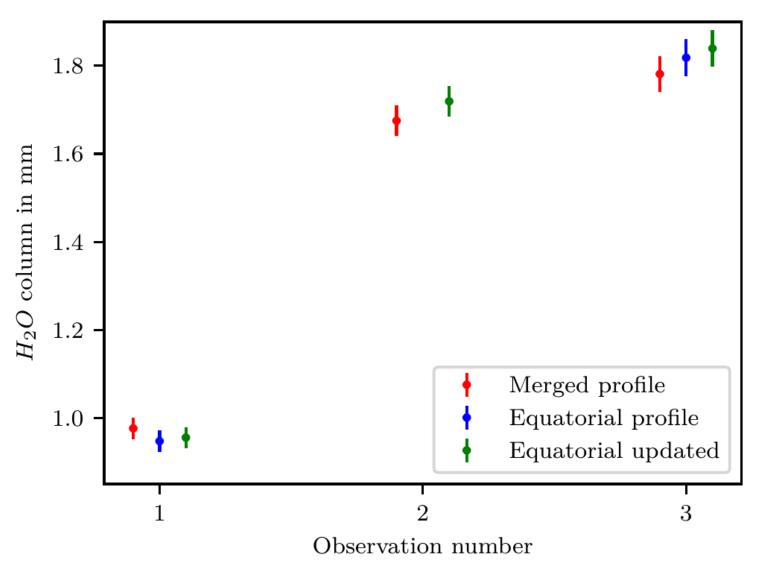
Equatorial profile: yearly averaged atmospheric profile

Equatorial profile updated: profile updated with humidity values from a first telluric correction



WATER COLUMN RESULTS

All atmospheric profiles lead to water column values which are **in agreement** within the error bars.



CONCLUSIONS

- Water lines: synthetic transmission methods correct best scatter is 1.3 times higher with the standard star
- Oxygen lines: standard star method corrects best scatter is 1.2 times smaller with the standard star Molecfit has a comparable level of correction
- Use of tailored atmospheric profile improves the correction level scatter 2 times smaller