

# Introducing high-school pupils to Modern Cosmology and GR

Dr. Alice Gasparini – Prof. Dr. Andreas Müller  
SwissMAP et Collège Rousseau – Unige

Geneva GIP Day  
Musée d'Ethnographie  
26 janvier 2017



**SwissMAP**

The Mathematics of Physics  
National Centre of Competence in Research



**UNIVERSITÉ  
DE GENÈVE**

Didactique de la physique

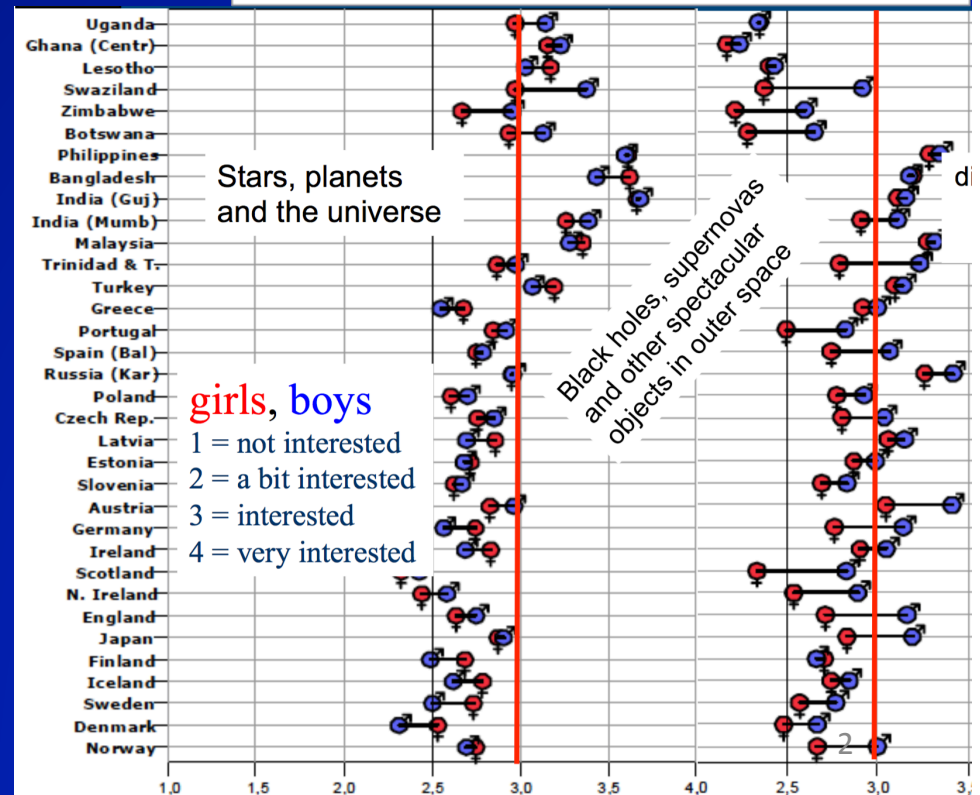
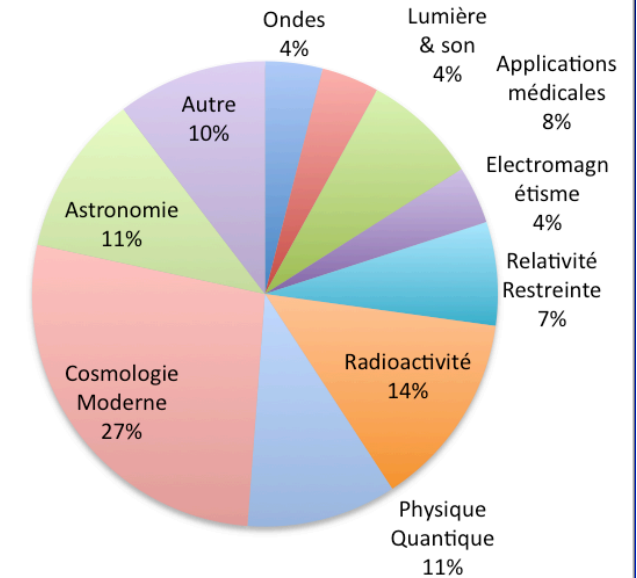
# Motivation & Purposes

1. Boost the pupils' level in maths and in « classical » taught physics by attractive subjects : not replace but complete the "classical" curriculum

+ Research ROSE (Relevance of Science Education)

- Similar results across many countries
- Averages for « ordinary » science subjects are about 2

Préférences élèves 2DF Rousseau - A.Chavanne, Novembre 2016



## 2. Improve the links between high school and actual research:

pupils who learn physics up to the 19<sup>th</sup> century have a distorted idea about the main issues of modern physics

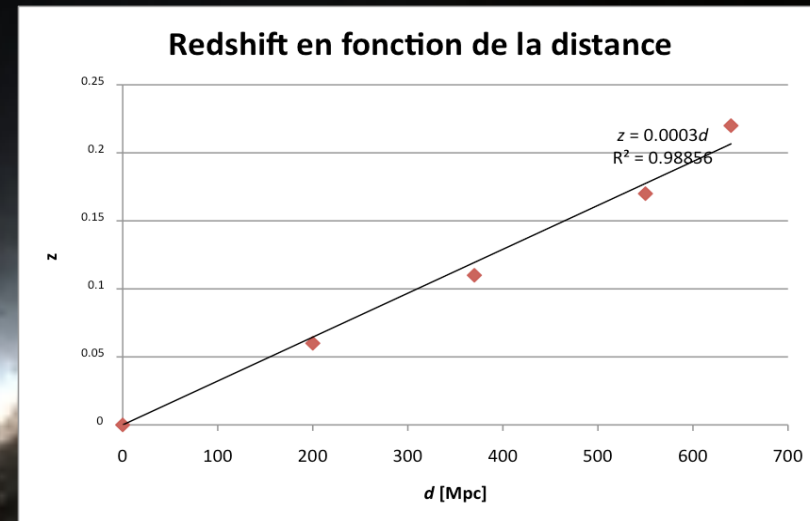
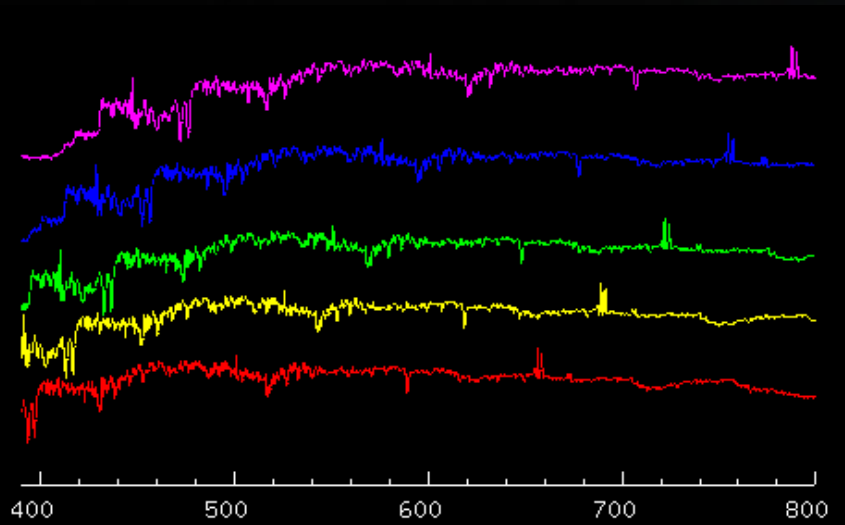
## 3. Medium level of elementarisation:

activities based on high school skills on maths and "classical" taught physics;

=> not for a wide public (zero equations), not academic level.

# Examples of tested activities: Cosmological distances

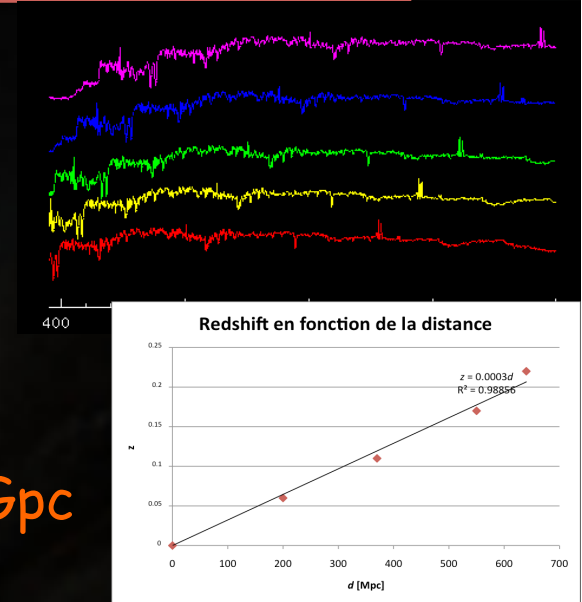
- Finding Hubble's law by comparing some nearby galaxies spectra



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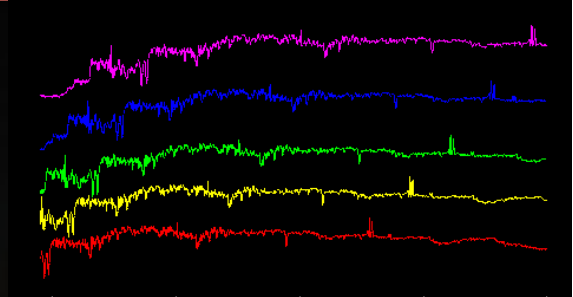
- Finding Hubble's law by comparing some nearby galaxies spectra
- Comparing the OOM of the expansion speed at different scales:

$$H_0 = 70 \text{ km/s / Mpc} = \dots / \text{mm} = \dots / \text{km} = \dots / \text{Gpc}$$



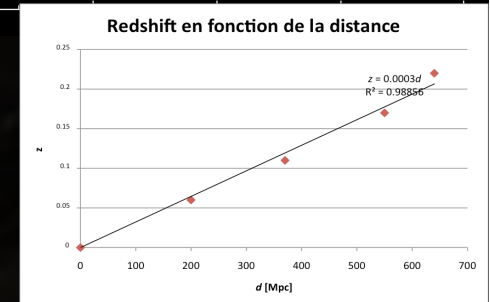
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- Finite value of  $c$  + expansion = the notion of distance split into Proper/Comoving/Time-travel/Angular/Luminosity distances

➤ Deriving the integral formula for each one as a function of  $z$ ,  $\Omega_m$  and  $\Omega_\Lambda$

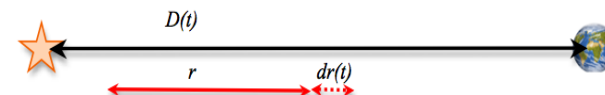
➤ Studying limit cases

➤ Numerical integration + comparing with the Supernova Cosmology Project data

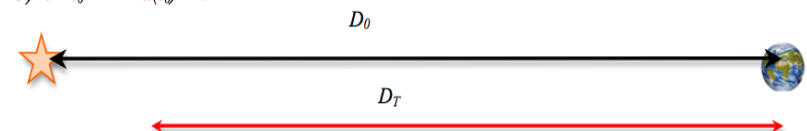
1)  $t \in [t_{em}; t_{em} + dt] \Rightarrow a(t_{em}) = 1/(1+z_{source})$



2)  $t \in [t; t + dt] \Rightarrow a(t_{em}) < a(t) < 1$



3)  $t = t_0 \Rightarrow a(t_0) = 1$



## Examples of tested activities: Strong Lensing

- A simple dimensional analysis to find the deflection angle

$$\alpha = \frac{4GM}{c^2 d}$$



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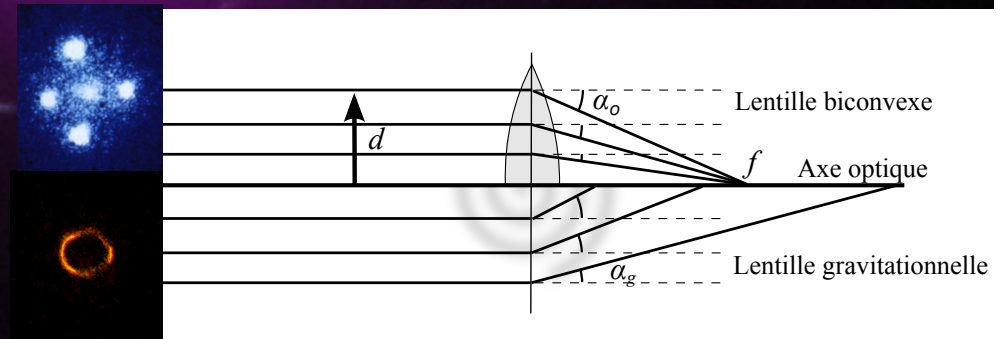


- How the bottom of wineglass can simulate the strong lensing effect?

- Solving a differential equation (simple integration) to find the profile of the optical lens: a curve of kind  $y(x) = y_0 \ln(x)$



- + Manip. Einstein's ring/cross





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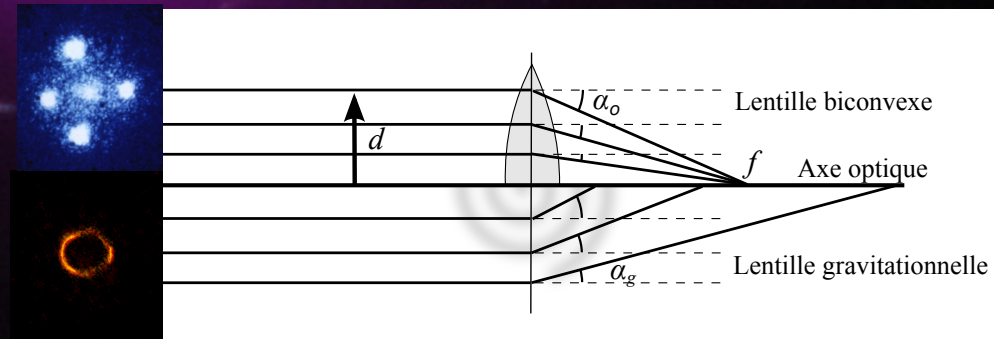


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- Trigonometry + algebra to find  $M$  as a function of the Einstein's radius  $\theta_{Einst}$  and the distances observer - lens - source.

$$\theta_{Einst.} = \sqrt{\frac{4GM D_{SL}}{c^2 D_{LO} D_{SO}}}$$

## Examples of tested activities: Nature of gravity

- Compare gravitational interaction and electromagnetic interaction;
  - For the system  $e^-/p^+$  in the H atom, and for the system Moon/Earth (if we could get all the electrons out of them):  $F_{em}/F_g \approx 10^{39}$
  - For  $F_{em} \approx F_g$  between 2 identical particles, we need their  $m/q \approx 10^{10} \text{ kg/C}$ . Replacing  $q = e$ , we get  $m \approx 10^{27} \text{ eV}$  (unification energy)
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  - Equivalence principle and space-time curvature
- BH: Use conservation of mechanical energy to get the escape speed formula  $v_1^2 = 2GM/R$ 
  - Schwarzschild Radius  $\Rightarrow$  when  $v_1 = c$
  - Compare  $v_1$  and thermal speed of different gases ( $H_2, N_2$ ) to explain the composition of planets' atmosphere

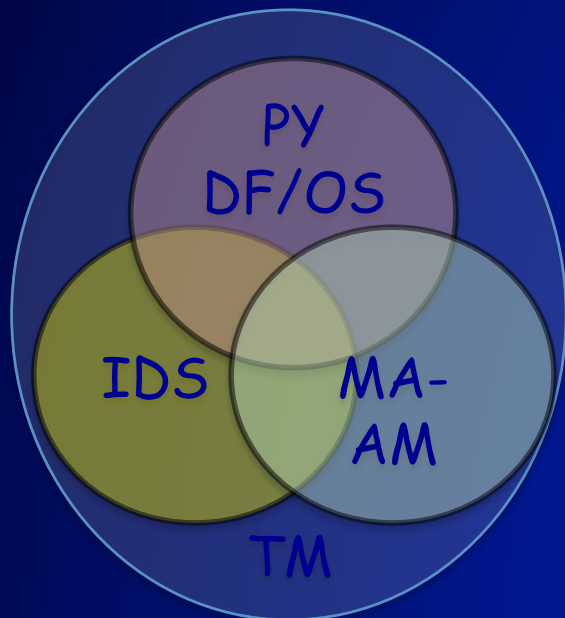
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- BH: Use Wien's law to get its temperature
  - + Estimate the time of evaporation using Stefan-Boltzmann law

# CM & GR: introduction pour élèves du secondaire II

Dr. A. Gasparini (SwissMAP) - Prof. Dr. A. Müller (Dida de la Physique), UniGE

- 9 chapters : theory + activities
- Ideal for a PY OC course (2h/week) planned on 2 semesters
- Toolbox for many others teaching contexts (punctual activities):



## Cosmologie Moderne et Relativité Generale

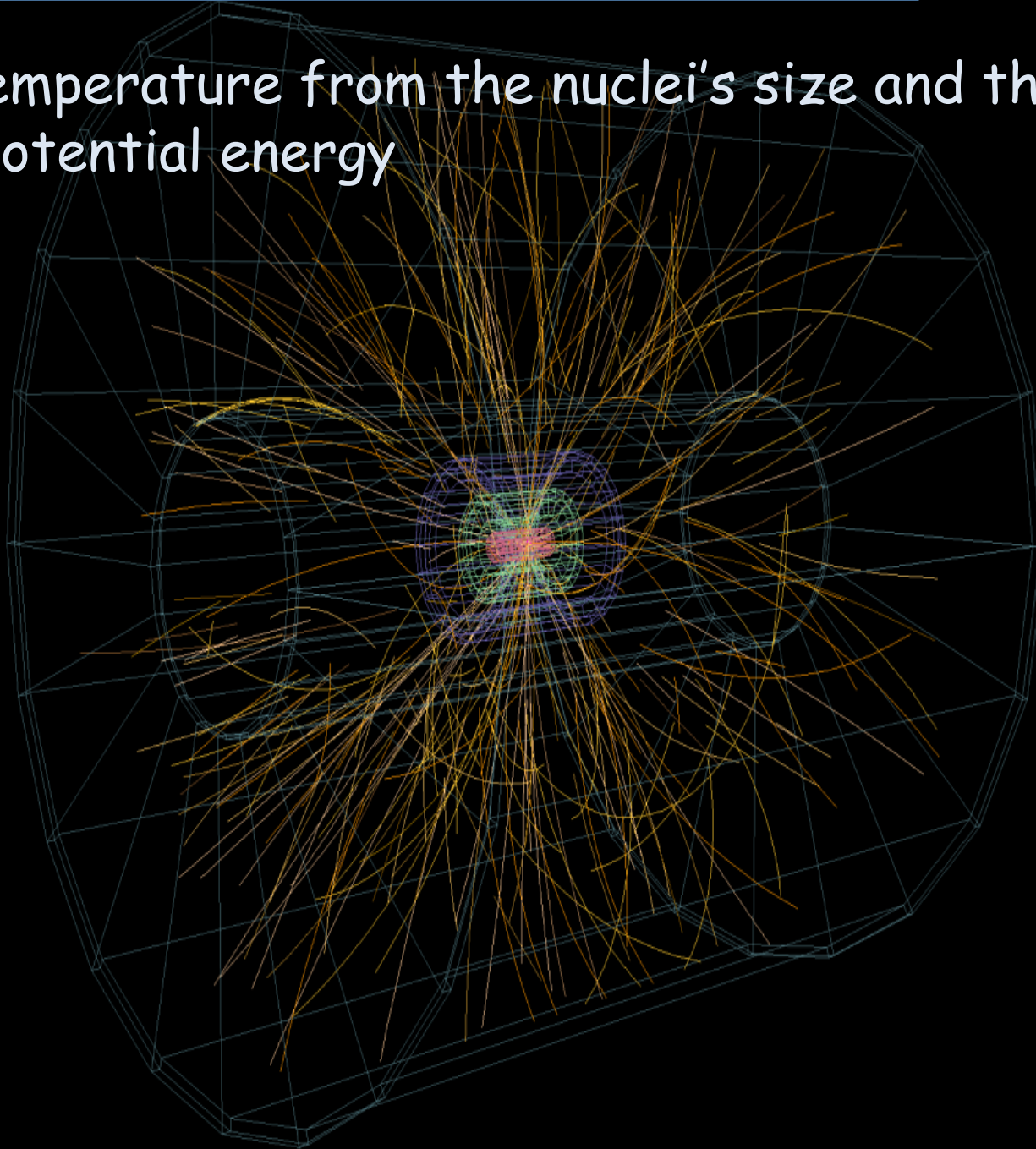
Activités pour les élèves du Secondaire II

Alice Gasparini, Andreas Müller

- Série 1 : Grandeurs
  - Série 2 : Expansion
  - Série 3 : Principe d'équivalence
  - Série 4 : Courbure
  - Série 5 : Lentille gravitationnelle
  - Série 6 : Trous noirs
  - Série 7 : Equations cosmologiques
  - Série 8 : Chronologie du Big Bang
  - Série 9 : Ondes Gravitationnelles
- 
- Activité expérimentale 1 : L'effet Doppler cosmologique
  - Activité expérimentale 2 : La courbure du cône

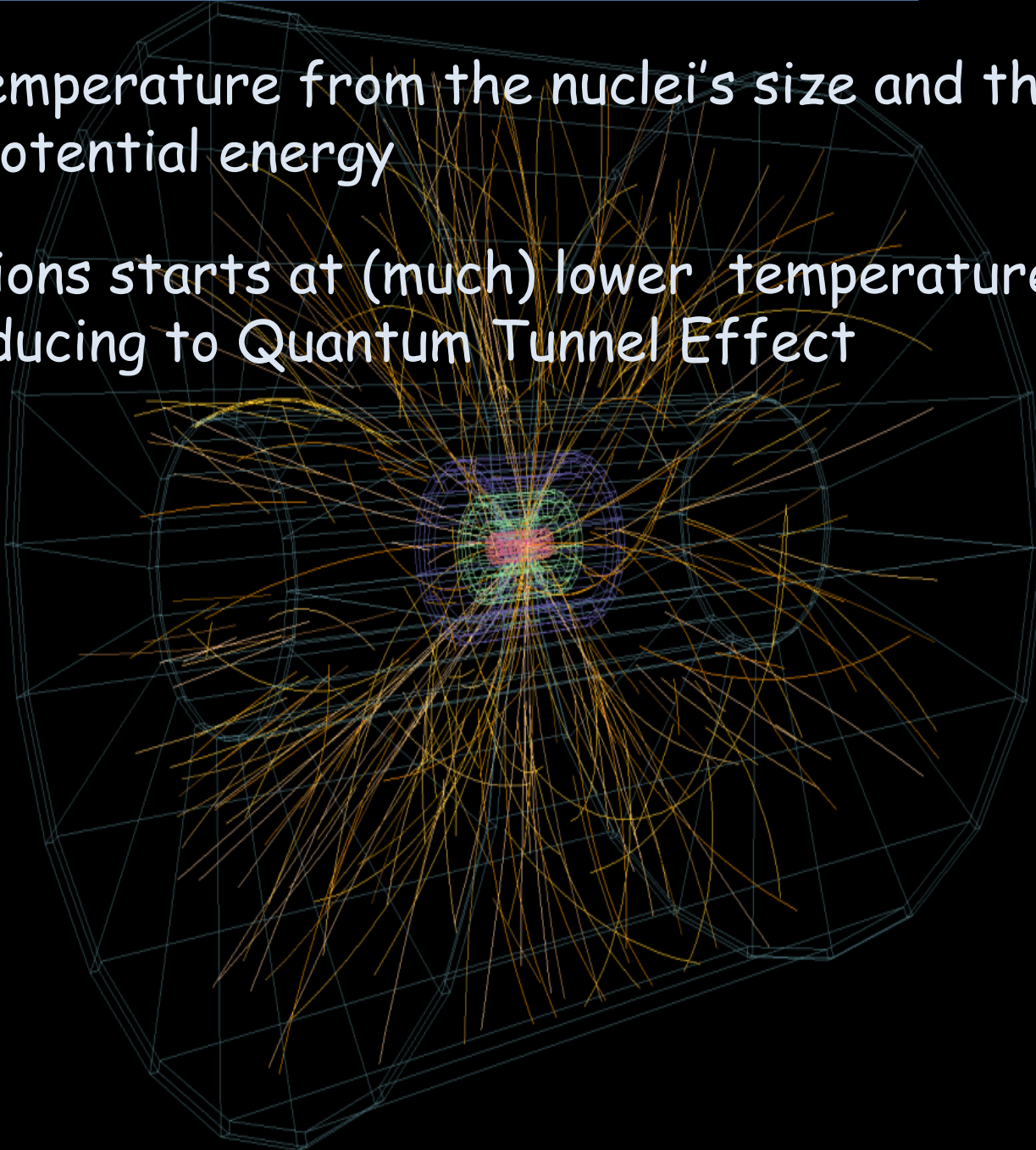
## Outlook: examples of activities to be tested

- Fusion temperature from the nuclei's size and the Coulomb potential energy



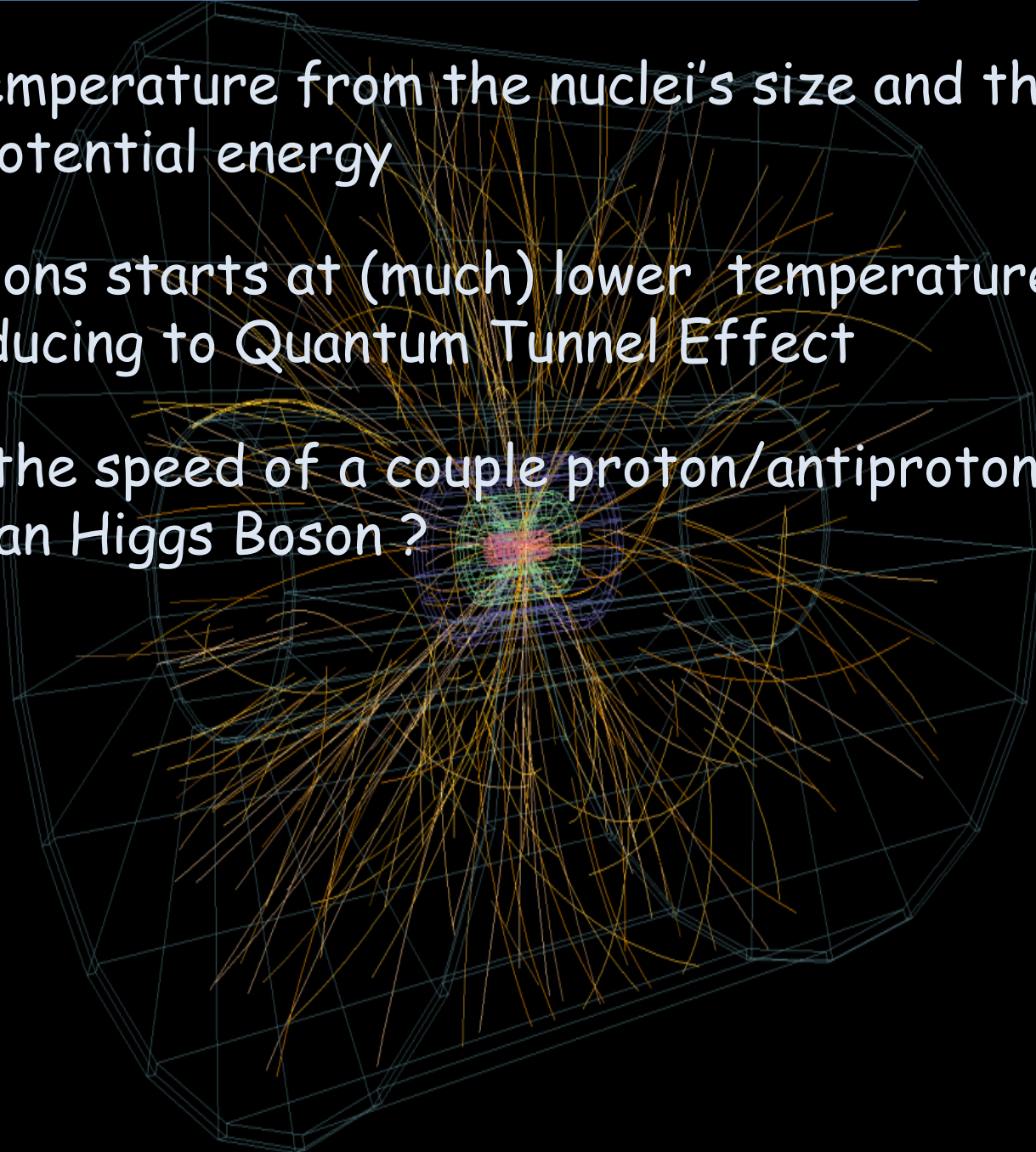
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=> Introducing to Quantum Tunnel Effect
- What is the speed of a couple proton/antiproton to produce an Higgs Boson ?
- Quadrupole formula for the amplitude of a GW:
  - Comparing quadrupole (GW) vs dipole (EMW) emission
  - Why only astrophysical masses can produce such a wave?
  - Why do we need "relativistic sources"? (BH or NS)
  - Why are these waves so important in the nowadays physics?
  - Relation  $M_{\text{source}}$  / frequency / detector's size
  - + all the exercises you can do with waves (DF & OS curriculum)



“

Science is  
competitive,  
aggressive,  
demanding.  
It is also  
imaginative,  
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”

—VERA  
RUBIN

World  
Science  
Festival