

Tidal disruptions rate at hard X-rays

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December 15th, 2015

28th Texas Symposium, Geneva

(based on Hryniewicz & Walter, arXiv:1505.06612)

Introduction to TDEs

Swift BAT and BAT data

Search approach

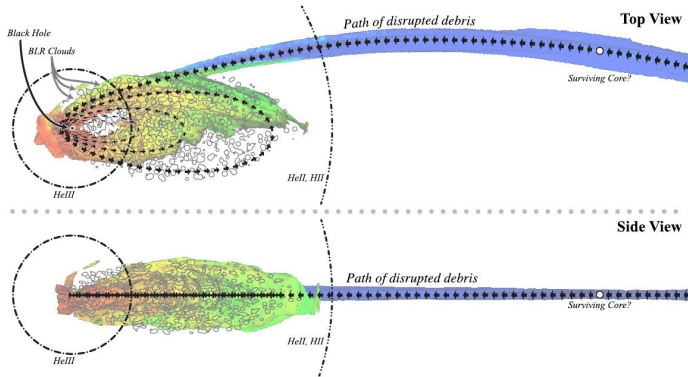
TDE candidates sample

average spectrum

TDE rate

Summary

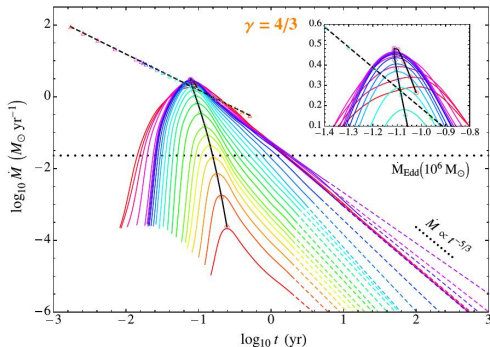
TDEs — simulations



Guillochon et al. (2014)

Star or substellar body disrupted by a compact object (black hole)

Simulations & theory



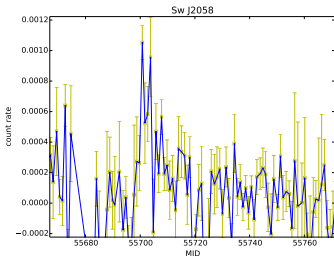
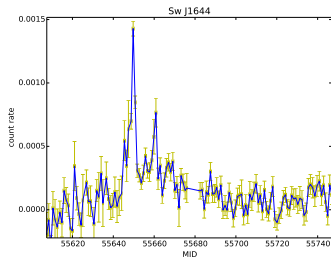
Guillochon & Ramirez-Ruiz (2013)

Accretion rate evolution depending on the impact factor.

Theoretical relation

$$\dot{M}(t) = \frac{2\pi}{3} (GM_h)^{2/3} \frac{dM}{dE} t^{-5/3}$$

Known TDEs



> 46 known TDEs.

Among them:

- ▶ 36 observed in IR/opt/UV,
- ▶ 24 at soft X-rays,
- ▶ 4 at hard X-rays,
- ▶ 2 radio emitters.

SMBH hosts (exception: WD + icy planet).

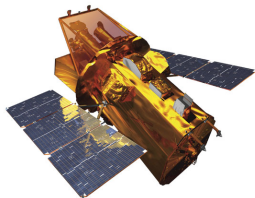
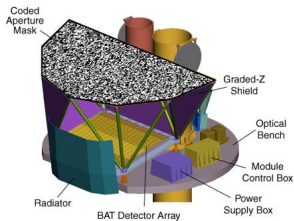
2 hard X-rays flares + radio flares are associated with jets in relativistic TDEs.

2 other are due to disruption of a planet

Swift Burst Alert Telescope

Swift

- ▶ 10 years of continuous operation
- ▶ observing whole sky



BAT

- ▶ 1.4 sr field of view (half coded)
- ▶ 14-195 keV spectral range
- ▶ Gaussian noise distribution

images: NASA

+ Good dust penetration at hard X-rays!

Aim

Estimate TDE rate at hard X-rays.

Approach

- ▶ search in nearby galaxies
- ▶ generating LC and images
- ▶ identification and removal of persistent sources
- ▶ flare detection (continuous positive flux excess in LCs)
- ▶ threshold on flare's significance and duration
- ▶ check for contaminations
- ▶ visual verification

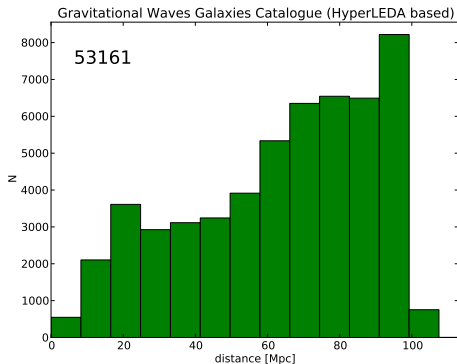
Galaxies

Initial sample

Gravitational Wave Galaxy Catalogue

catalogue (White et al. 2011)

- ▶ 53161 galaxies
- ▶ 150 globular clusters
- ▶ $d \leq 100$ Mpc
- ▶ mostly optical identifications



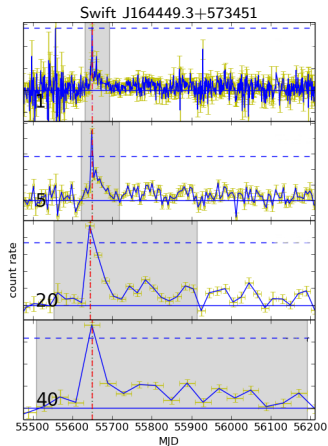
targets overlapping (due to BAT limited
spatial resolution)

BAT LC

We chose 5 days time binning as good option for our purpose.

Steps

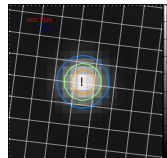
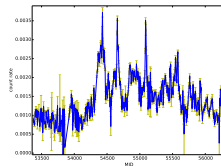
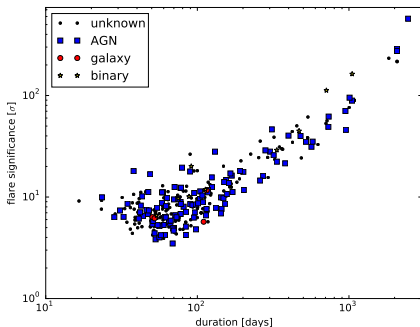
- ▶ LC on the position of every GWGC galaxy
- ▶ the most significant positive flux excess detection
- ▶ image for a flare period
- ▶ PSF fitting and significance estimation



Persistent sources

Persistent sources were identified based on the level of the average flux outside detected flare period.

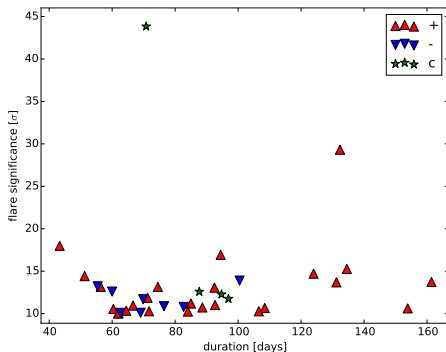
5 new candidates for unknown persistent sources.



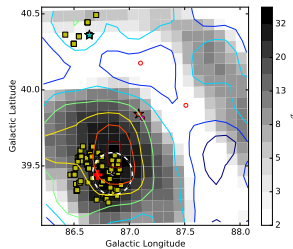
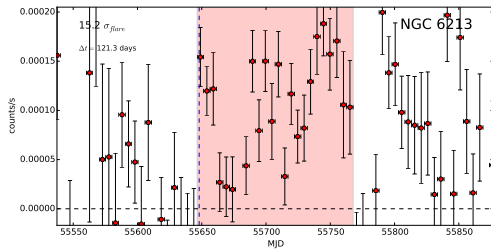
Flares in non-persistent sources

Interesting region

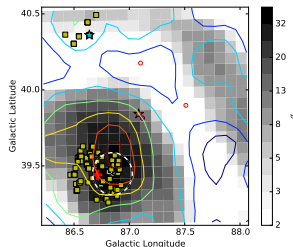
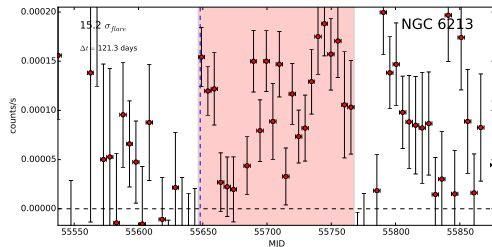
- ▶ high (integrated) significance ($> 10\sigma$)
- ▶ long flares (> 40 days)
- ▶ 27 non-AGNs candidates



Verification

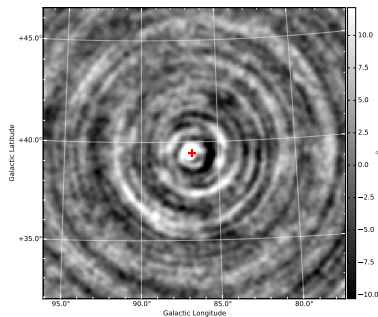


Verification

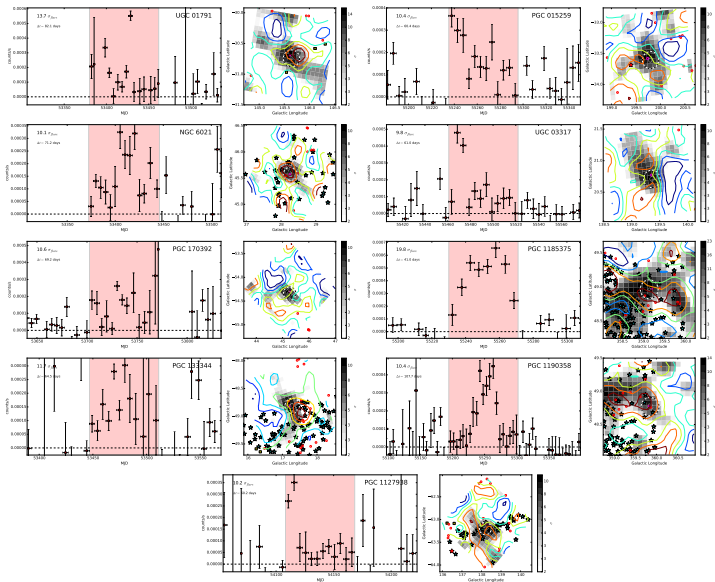


Checks

- ▶ contamination by known bright BAT sources
- ▶ PSF match
- ▶ visual inspection
- ▶ BAT artifacts

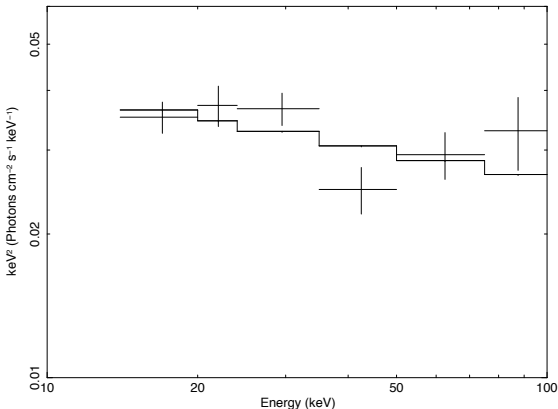


TDE candidates



TDE's average spectrum

The average spectrum in the BAT TDE sample is consistent with power-law AGN spectra. Inverse Compton processes and electrons acceleration in shocks in the accretion flow take place.



~ 4% of the total flux is emitted at hard X-rays? (in AGNs it was suggested by Risaliti & Elvis 2004).

TDE frequency

If we have ~ 9 true candidates from BAT alone:

$$\dot{N}_{TDF} \approx 2.1 \times 10^{-5} \text{ yr}^{-1} \text{ galaxy}^{-1}$$

TDE frequency

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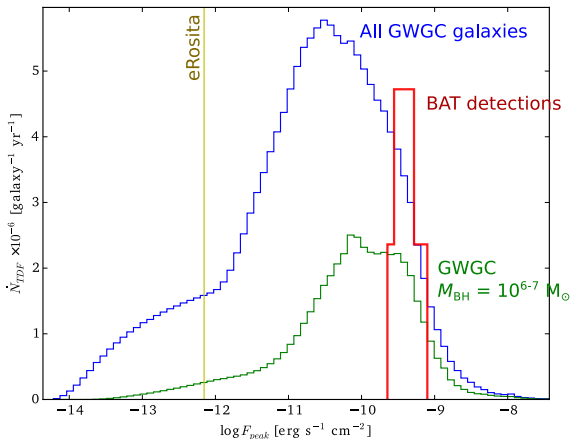
$$\dot{N}_{TDF} \approx 2.1 \times 10^{-5} \text{ yr}^{-1} \text{ galaxy}^{-1}$$

Previous estimations:

- ▶ SDSS galaxies (sensitive to high mass SMBHs; van Velzen et al. 2012)
 $< 3 \times 10^{-4} \text{ yr}^{-1} \text{ galaxy}^{-1}$
- ▶ N-body simulations (Sgr A* like SMBHs; Brockamp et al. 2011)
 $(5.5 \pm 2.7) \times 10^{-5} \text{ yr}^{-1} \text{ galaxy}^{-1}$
- ▶ RASS data (Donley et al. 2002)
 $9.1 \times 10^{-6} \text{ yr}^{-1} \text{ galaxy}^{-1}$ (large X-ray outbursts from inactive galaxies)
 $8.5 \times 10^{-4} \text{ yr}^{-1} \text{ galaxy}^{-1}$ (AGNs)

TDE detectability

Possibility of TDE detection at hard X-rays. Peak flux histogram for GWGC. Assuming stellar MF, estimating GWGC BH masses and geometrical probability of encounter.



Summary

- ▶ Already known TDEs were detected in our search, so procedure works.
- ▶ We propose 9 new non-relativistic TDE candidates.
- ▶ Emission from the TDEs is similar to the AGNs emission.
- ▶ Estimated TDE rate at hard X-rays is consistent with those in lower energy bands.
- ▶ Low sensitivity is compensated by detections of dusty objects?
- ▶ 2 new hard X-ray persistent sources were proposed (edge-on active galaxies).

Thank you