

Testing for inhomogeneities in *real time*

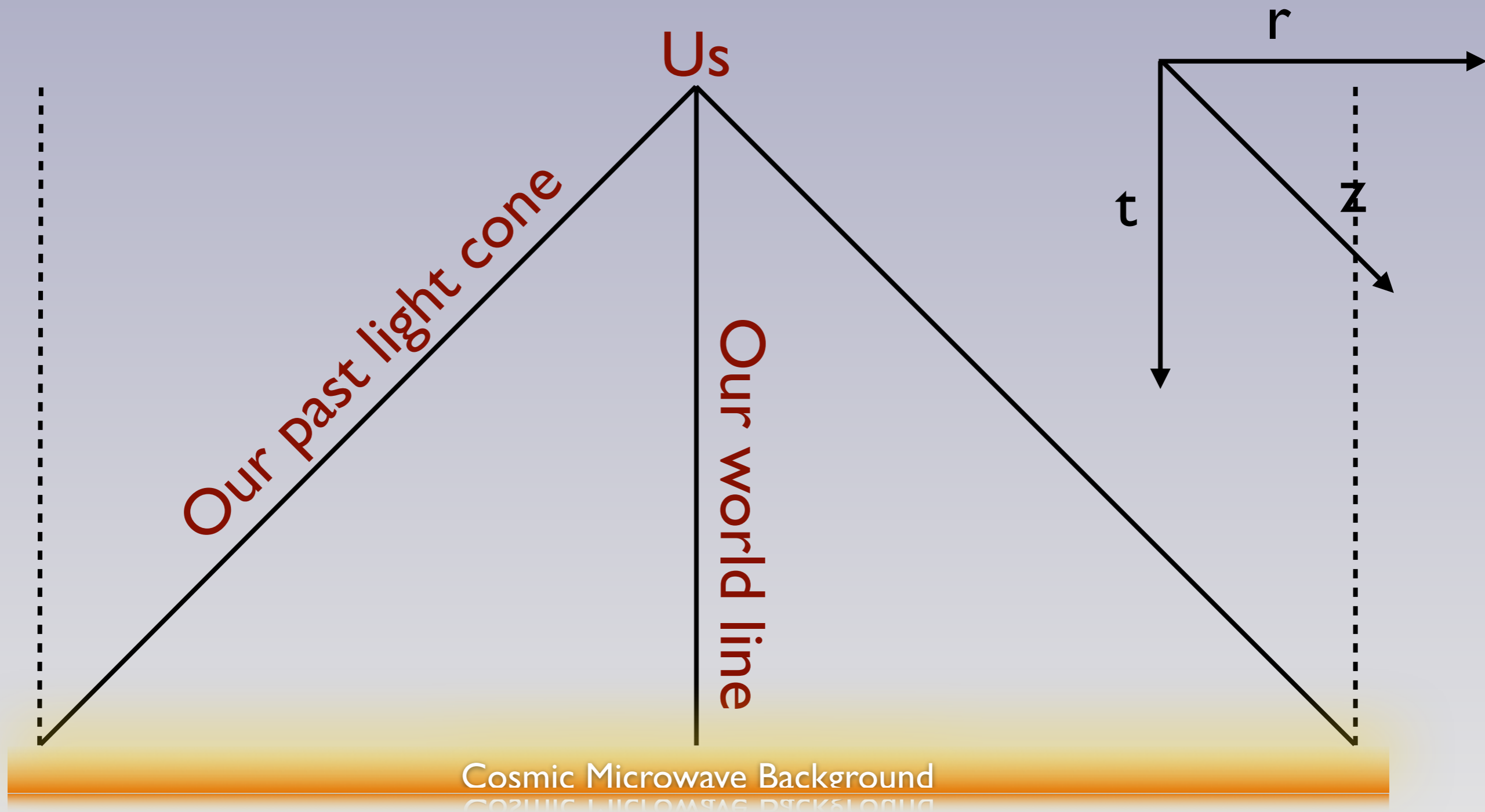
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Based on: Amendola, WV et al., JCAP (2013)

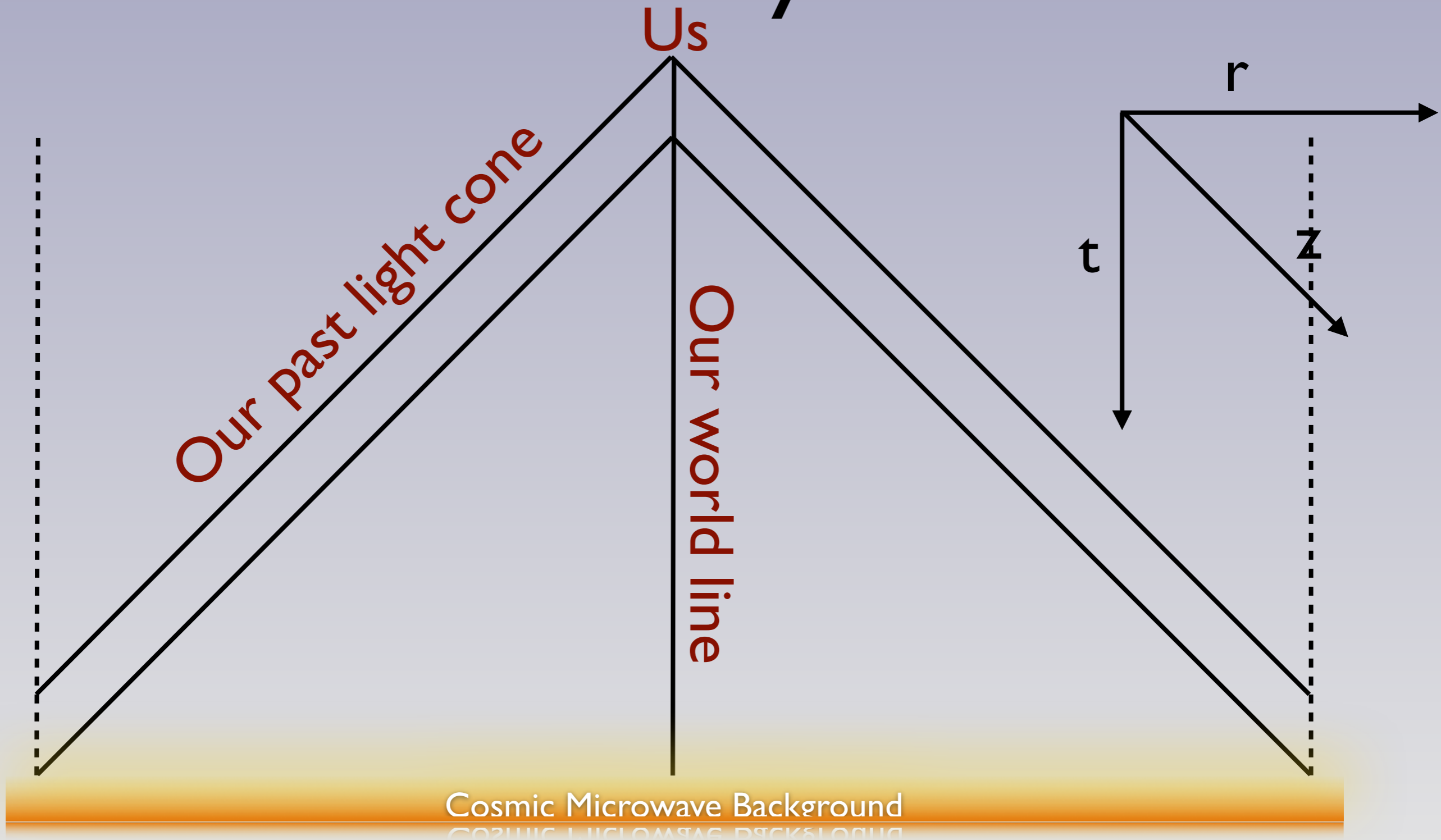
Take home message

- Proof on concept of a tool:
 - Real-time cosmology can be used to observe different inhomogeneous models
- Focus on angular motion using multipole vectors
- This talk: no focus on redshift drift

More insight inside the light cone: realtime cosmology



Realtime cosmology: observe the same 10 yrs later



Real time cosmology

- With e.g. GAIA have enough precision and volume
- only need some more time
- 500.000 quasars
- cross correlate maps from different times

Real time cosmology

The GIQC_5 in a nutshell

[Andrei et al. 2014]

| | |
|---|-----------------------------------|
| Number of sources | 1,248,372 |
| Sources with magnitude | 1,246,512 |
| Sources with redshift | 1,157,285 |
| Astrometry precision | 1 arcsec |
| Magnitude precision | 0.5 |
| Redshift precision | 0.01 |
| Average density | 30.3 sources/deg ² |
| Average neighbor distance | 3.7 arcmin (σ 4.9 arcmin) |
| Maximum distance to neighbor | 5.2 deg |
| Maximum distance to neighbor (average of 100 larger values) | 3.0 deg (σ 0.6 deg) |

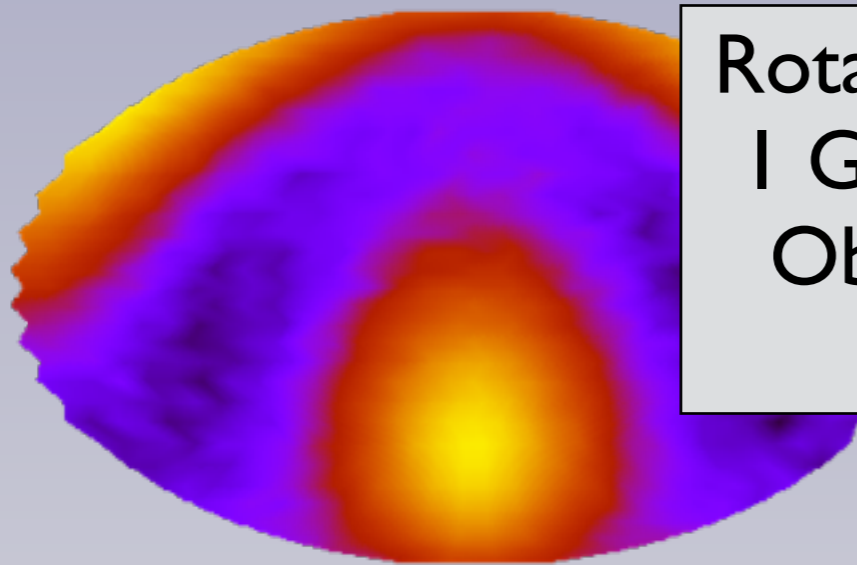
$$F[\hat{n}_1, \hat{n}_2, \hat{n}'_1(\hat{n}_1, \vec{q}_1), \hat{n}'_2(\hat{n}_2, \vec{q}_2)] \equiv \cos \gamma_{12} - \cos \gamma'_{12}$$

$$\langle F \rangle(\theta, \phi) \equiv \int_{r_{\min}}^{r_{\max}} dr n_{\text{obj}}(r) \int_{r_{\min}}^{r_{\max}} dr' n_{\text{obj}}(r') \int \frac{d\Omega'}{4\pi} F,$$

$$\langle F \rangle(\theta, \phi) = \sum_{l=0}^{\infty} \sum_{m=-l}^l a_{lm} Y_{lm}(\theta, \phi).$$

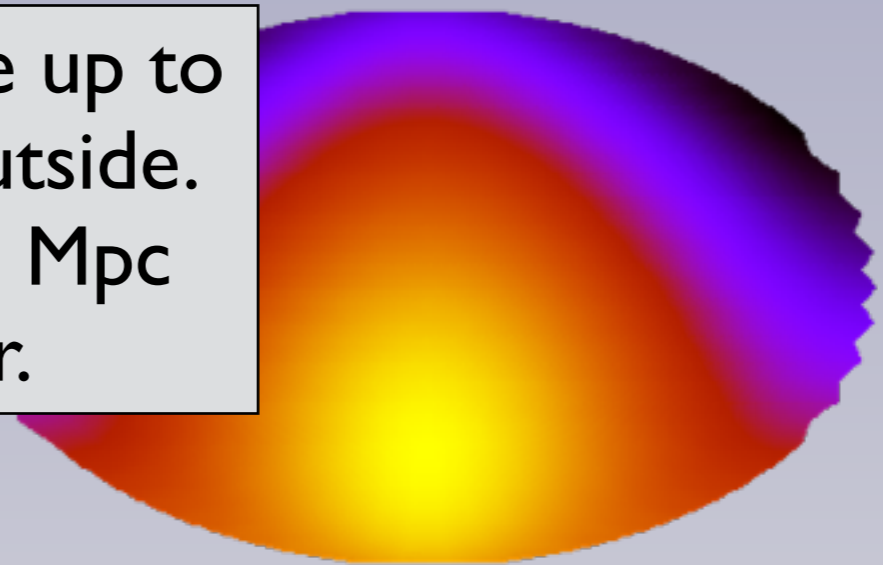
Simulate different anisotropic models

1/r-rotation $\langle F^2 \rangle$

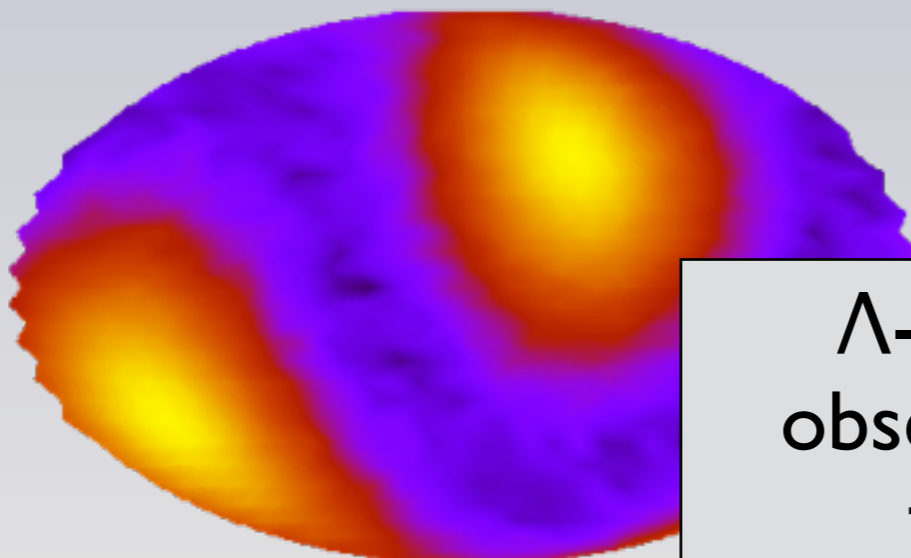


Rotating universe up to 1 Gpc, FLRW outside. Observer at 30 Mpc from center.

1/r-rotation $\langle F \rangle$

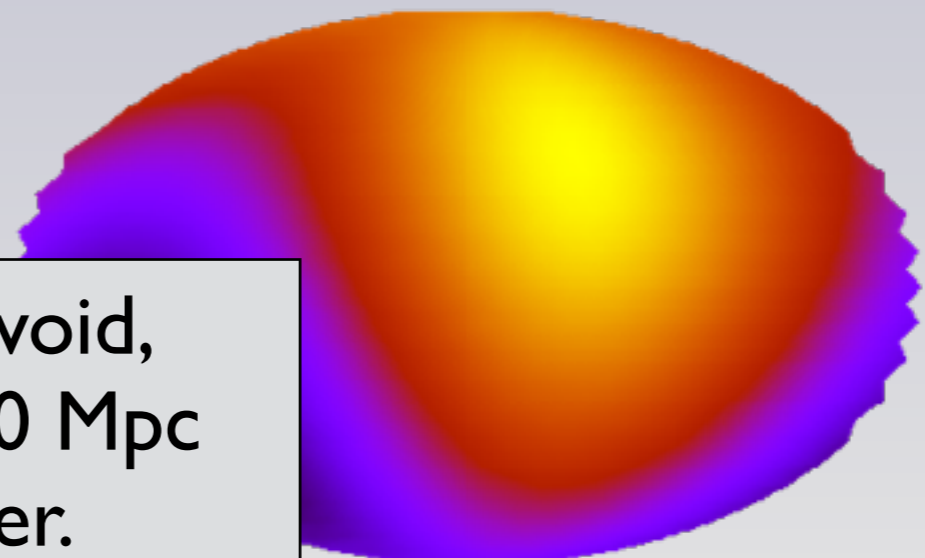


LTB Void $\langle F^2 \rangle$



Λ -replacing void, observer at 30 Mpc from center.

LTB Void $\langle F \rangle$

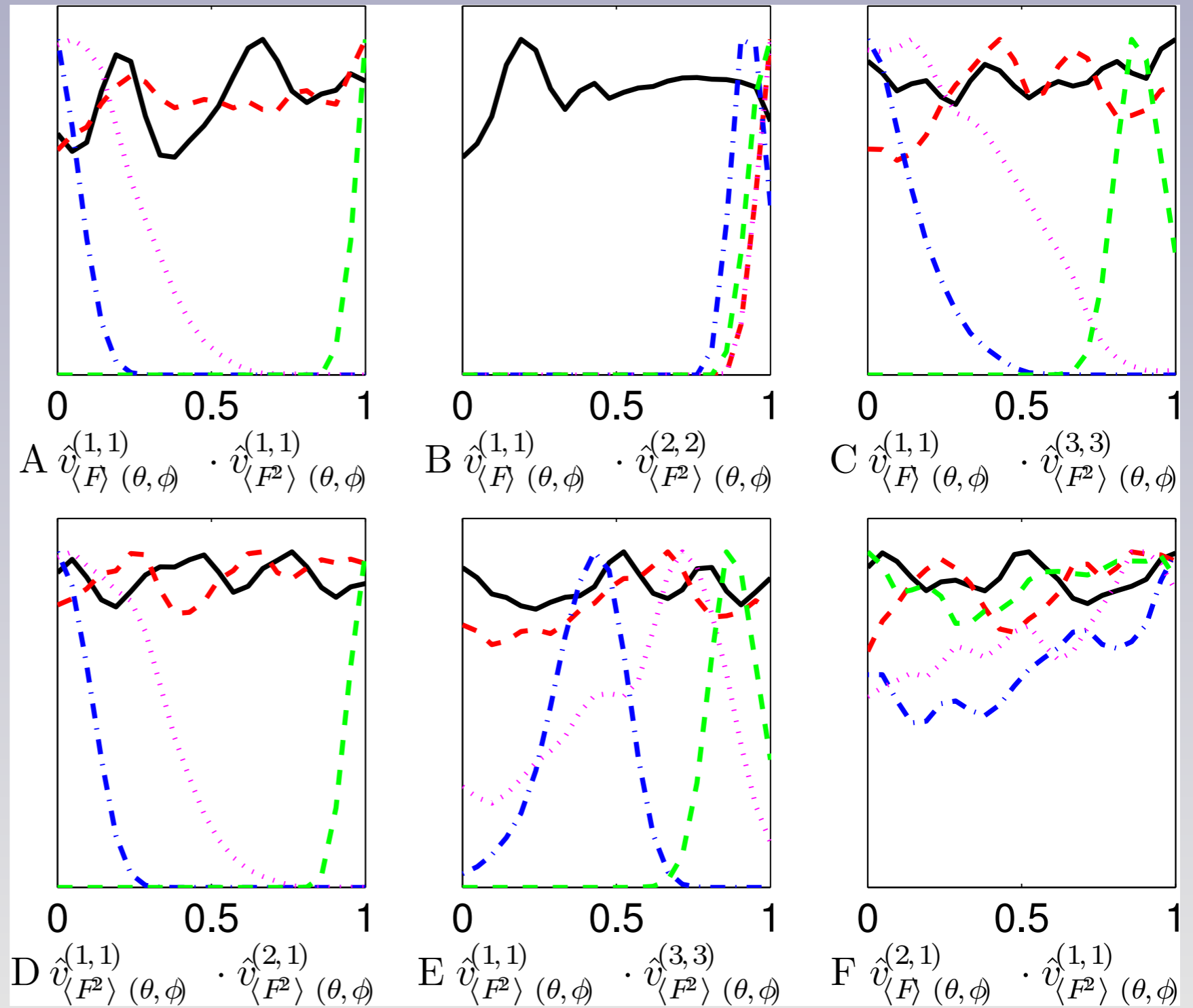


Compare orientations of different multipoles: analysis of axis of symmetry of different models

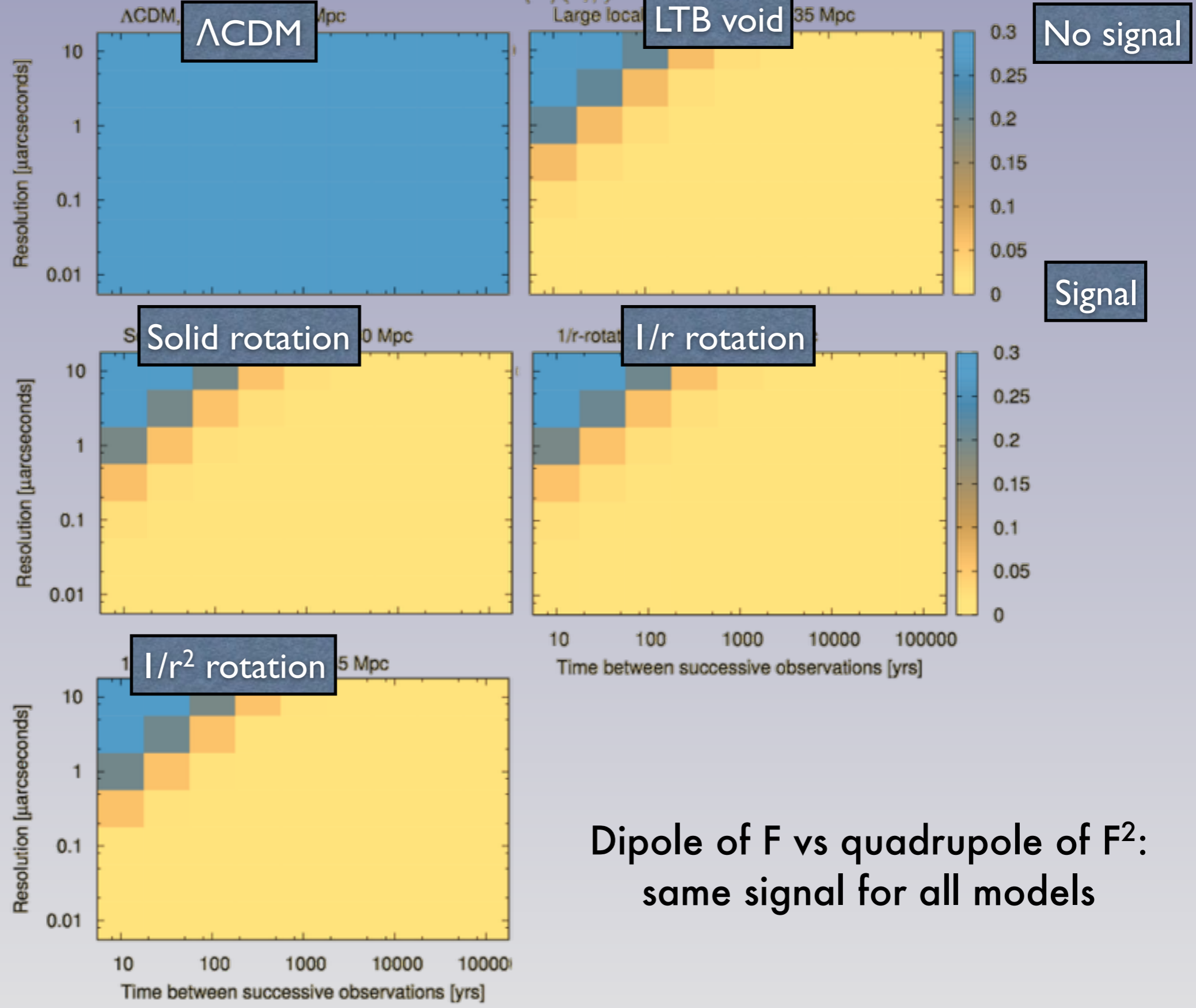
- Decompose $\langle F(\theta, \varphi) \rangle$ maps in multipoles
- Compute multipole vectors to quantify directions of anisotropy Copi et al. (2003)
- Compute inner products of different vectors: coordinate independent handle on directions.
- e.g. dipole orthogonal to quadrupole?
Parallel?

Compare orientations of different multipoles: analysis of axis of symmetry of different models

For an extreme situation: 1000 years of observation time, and high velocities on rotating models.

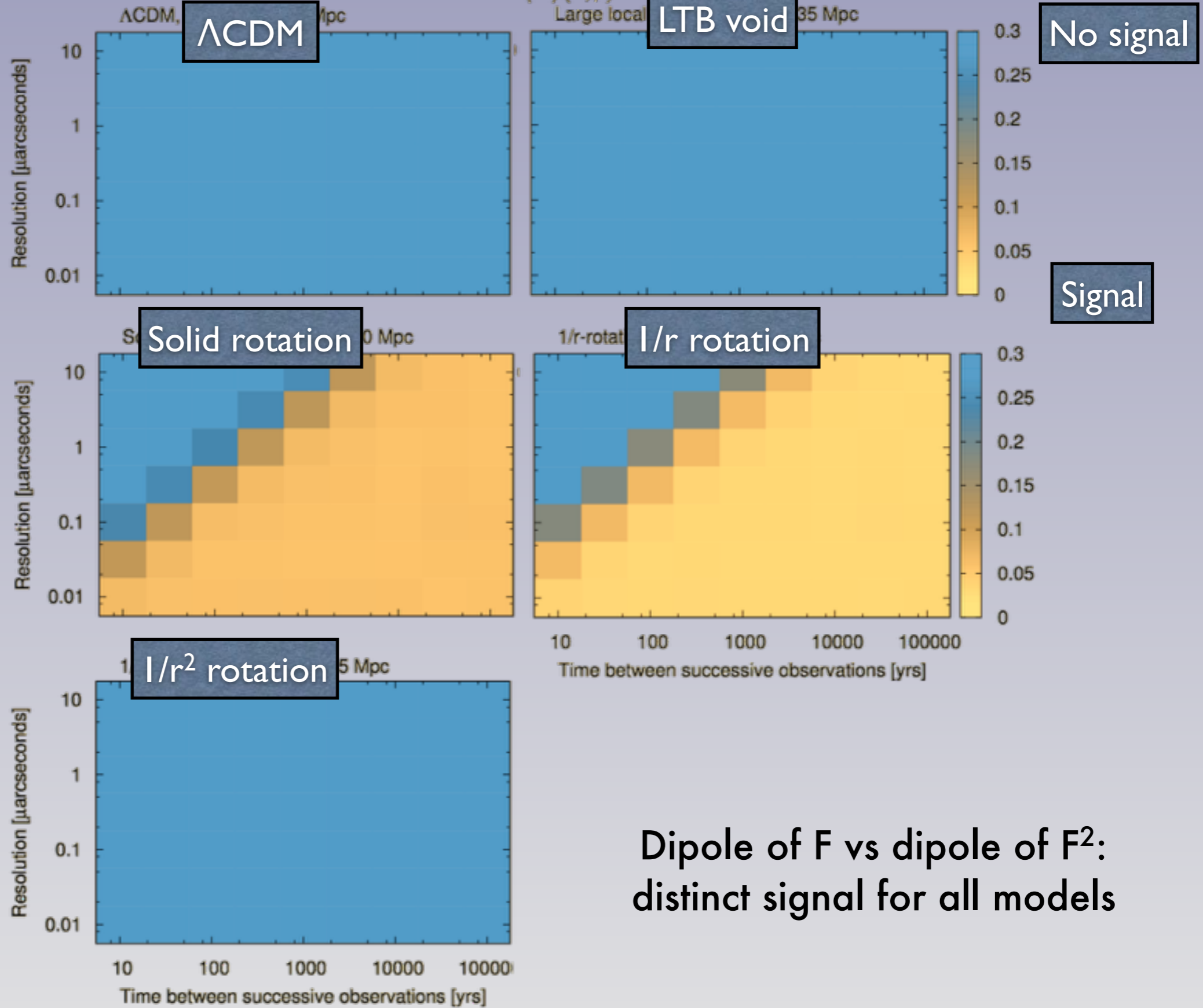


Detectability of $\hat{v}_{(F)}^{(1,1)}(\theta, \phi) \cdot \hat{v}_{(F^2)}^{(2,2)}(\theta, \phi)$



Dipole of F vs quadrupole of F²:
same signal for all models

Detectability of $\hat{v}_{\langle F \rangle}^{(1,1)}(\theta, \phi) \cdot \hat{v}_{\langle F^2 \rangle}^{(1,1)}(\theta, \phi)$



Dipole of F vs dipole of F²:
distinct signal for all models

Conclusion

- Proof on concept of a tool:
- Real-time cosmology can be used to observe different inhomogeneous models
- Using multipole vectors