

3D Quantum Bubble Collisions

Jonathan Braden

University College London

TEXAS 2015, Geneva, December 14, 2015

Based on work with Dick Bond and Laura Mersini-Houghton

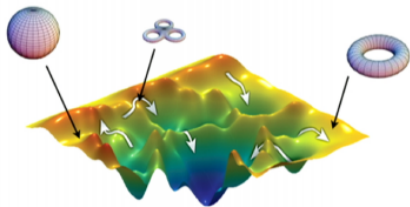
JCAP 1503 (2015) 03, 007 [arXiv:1412.5591]

JCAP 1508 (2015) 08, 048 [arXiv:1505.01857]

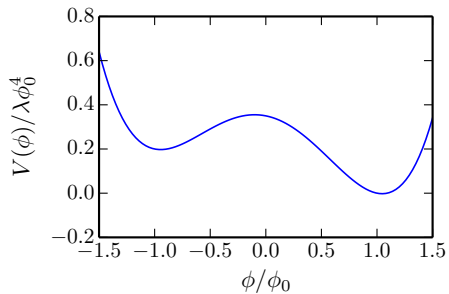
JCAP 1509 (2015) 09, 004 [arXiv:1505.02162]

Videos at www.star.ucl.ac.uk/~jbraden

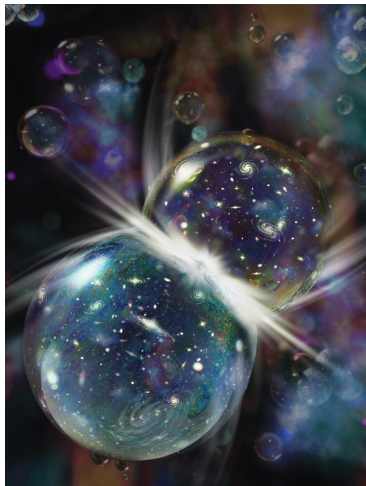
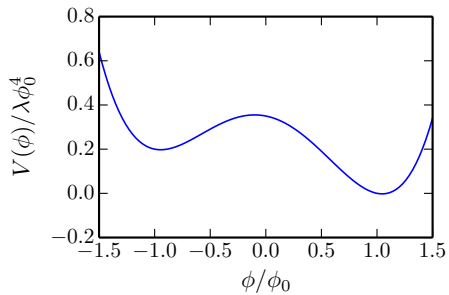
The Bubbly Universe



The Bubbly Universe



The Bubbly Universe



Large Literature Starting in 1982 ...

Single Bubbles

- ▶ Coleman, deLuccia
- ▶ **Hawking, Moss**
- ▶ Turok
- ▶ Sasaki, Linde, Tanaka, Yamamoto
- ▶ Garriga, Vilenkin, Montes, Garcia-Bellido
- ▶ Guth, Guven
- ▶ Freese, Adams
- ▶ Susskind et al
- ▶ ...

Vacuum Bubble Collisions

- ▶ **Hawking, Moss, Stewart**
- ▶ Kosowski, Turner, Watkins, Kamionkowski
- ▶ **Johnson, Aguirre**, Tysanner, Larfors
- ▶ Chang, Kleban, Levy, Sigurdson, Gobbetti
- ▶ Easter, Giblin, Lim, Lau
- ▶ **Johnson**, Lehner, **Peiris**,... (GR)
- ▶ ...

Observations

- ▶ **Johnson, Peiris**, Mortlock, McEwan, Feeney,...
- ▶ Smith, Senatore, Osborne

... All Based on the "Canonical" SO(2,1) Approach

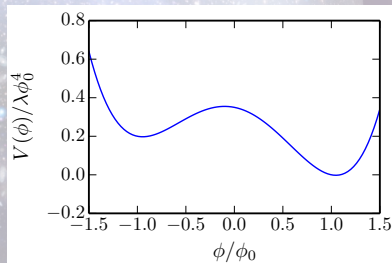
New Coordinates (Minkowski)

$$t = s \cosh \psi$$

$$x = x$$

$$y = s \sinh \psi \cos \theta$$

$$z = s \sinh \psi \sin \theta$$



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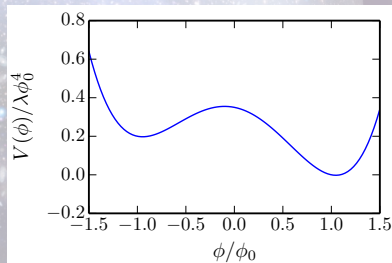
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SO(2,1) Assumption: $\phi(t, x, y, z) = \phi(s, x)$

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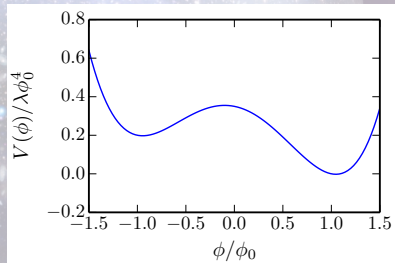
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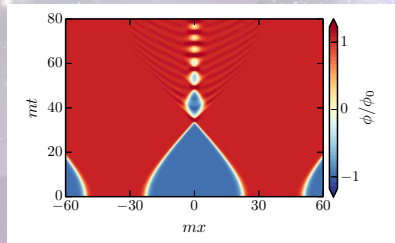
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SO(2,1) Assumption: $\phi(t, x, y, z) = \phi(s, x)$

Individual Bubbles

- ▶ Perfectly Spherical
- ▶ Boost Invariant



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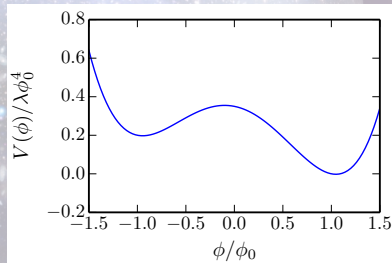
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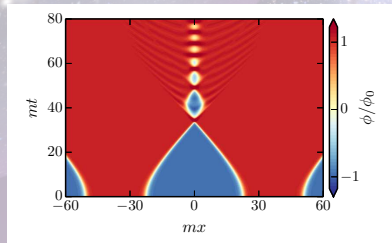
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SO(2,1) Assumption: $\phi(t, x, y, z) = \phi(s, x)$

2nd Bubble Breaks

- ▶ 1 Boost
- ▶ 2 Rotations



... All Based on the "Canonical" SO(2,1) Approach

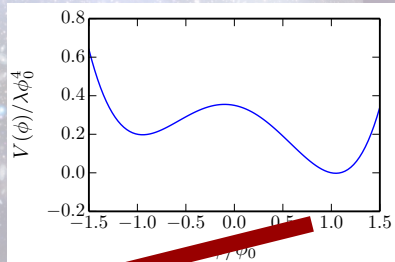
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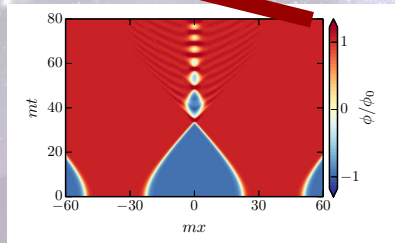
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~~SO(2,1) Assumption: $(t, x, y, z) = \phi(s, x)$~~

2nd Bubble Breaks

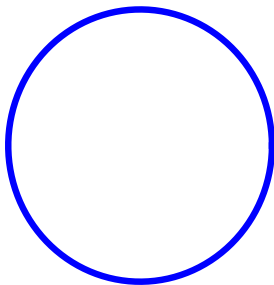
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Limitations of SO(2,1) Formalism

- Quantum (or Stochastic) Fluctuations

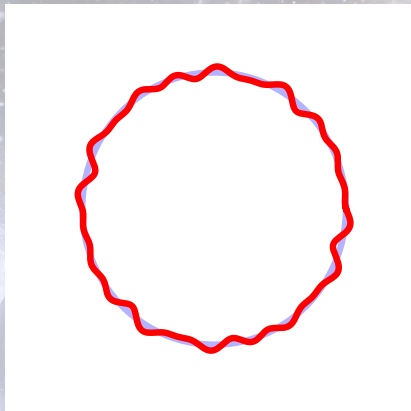
$$\phi(t, x, y, z) = \phi_{bg}(s, x) + \delta\hat{\phi}(t, x, y, z)$$



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Limitations of SO(2,1) Formalism

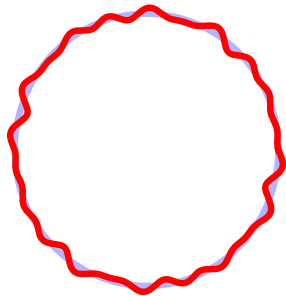
Quantum (or Stochastic) Fluctuations

$$\phi(s, x, \psi, \theta) = \phi_{bg}(s, x) + \delta\hat{\phi}(s, x, \psi, \theta)$$

$\delta\phi$ has dynamics not captured by SO(2,1) formalism

Ignoring $\delta\phi$ Breaks

- ▶ Quantum Mechanics
- ▶ Bubble Nucleation
- ▶ Inflationary perturbations



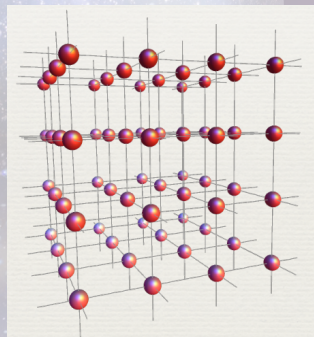
Numerical Approach is Essential [JB, in preparation]

Hybrid MPI/OpenMP Lattice Code

- ▶ Solve field equation (e.g.)

$$\ddot{\phi}_i + 3\frac{\dot{a}}{a}\dot{\phi}_i - \frac{\nabla^2\phi_i}{a^2} + V'(\vec{\phi}) = 0$$

- ▶ 10th order Gauss-Legendre integration (general) or 8th order Yoshida (nonlinear sigma models)
- ▶ Finite-difference (fully parallel) or Pseudospectral (OpenMP)
- ▶ Optional absorbing boundaries
- ▶ Quantum fluctuations \rightarrow realization of random field



- ▶ Energy conservation $\mathcal{O}(10^{-9} - 10^{-14})$

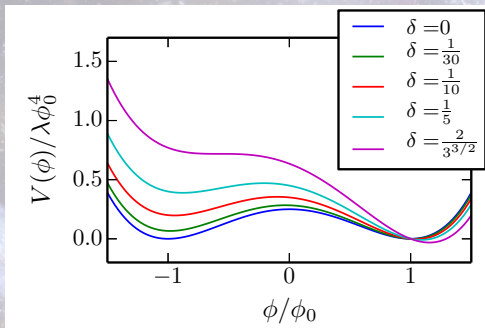
Initial Conditions I: The Bounce Solution [c.f. Coleman]

$$\frac{d^2\phi_b}{dr_E^2} + \frac{3}{r_E} \frac{d\phi_b}{dr_E} - \frac{\partial V}{\partial\phi} = 0$$

$$\phi_b(r_E = \infty) = \phi_{false} \quad \partial_{r_E}\phi_b(r_E = 0) = 0$$

$$r_E^2 \equiv \tau^2 + x^2 + y^2 + z^2 \quad \tau \equiv it$$

Pseudospectral
Approximation



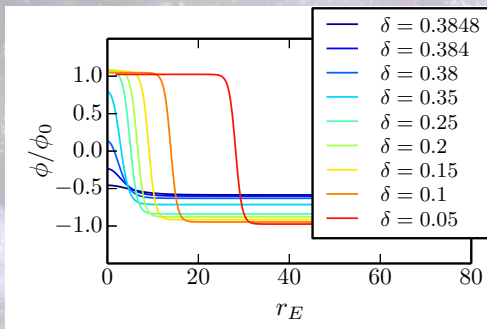
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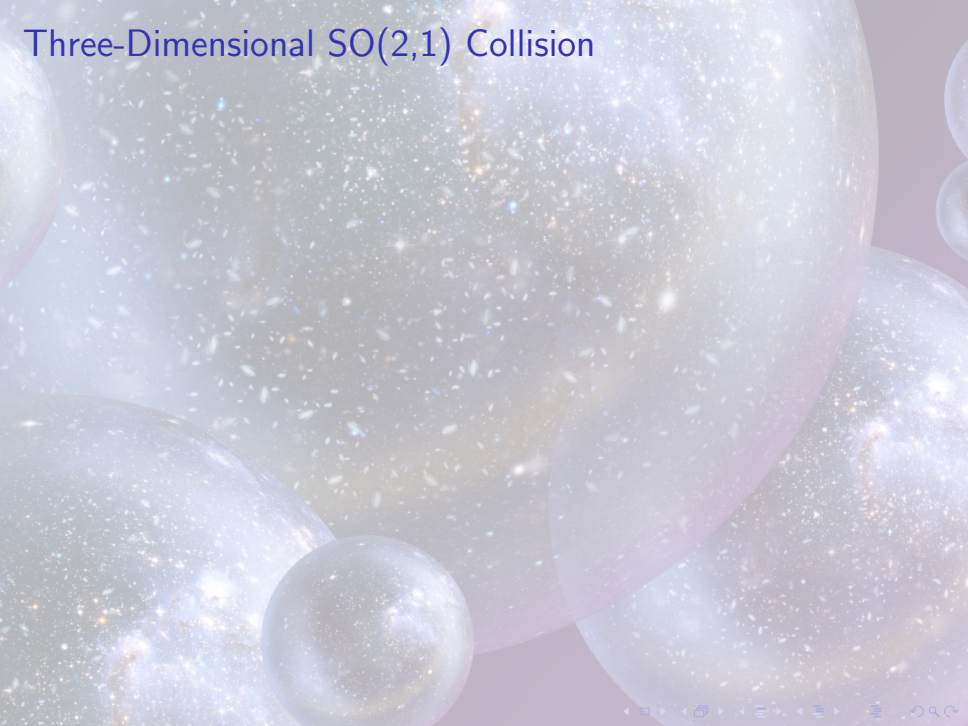
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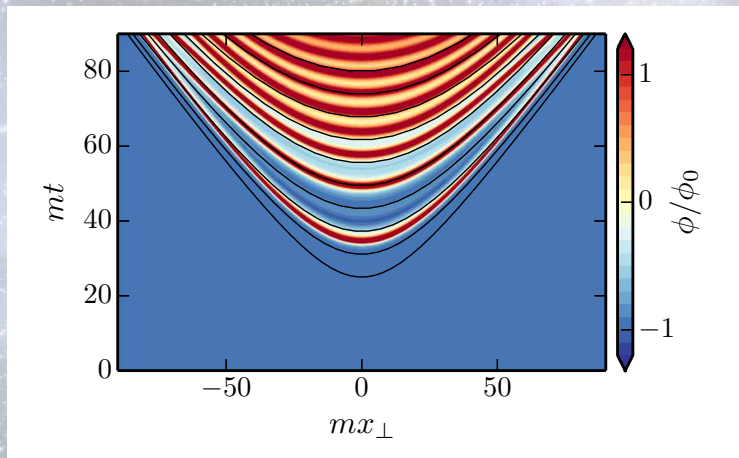
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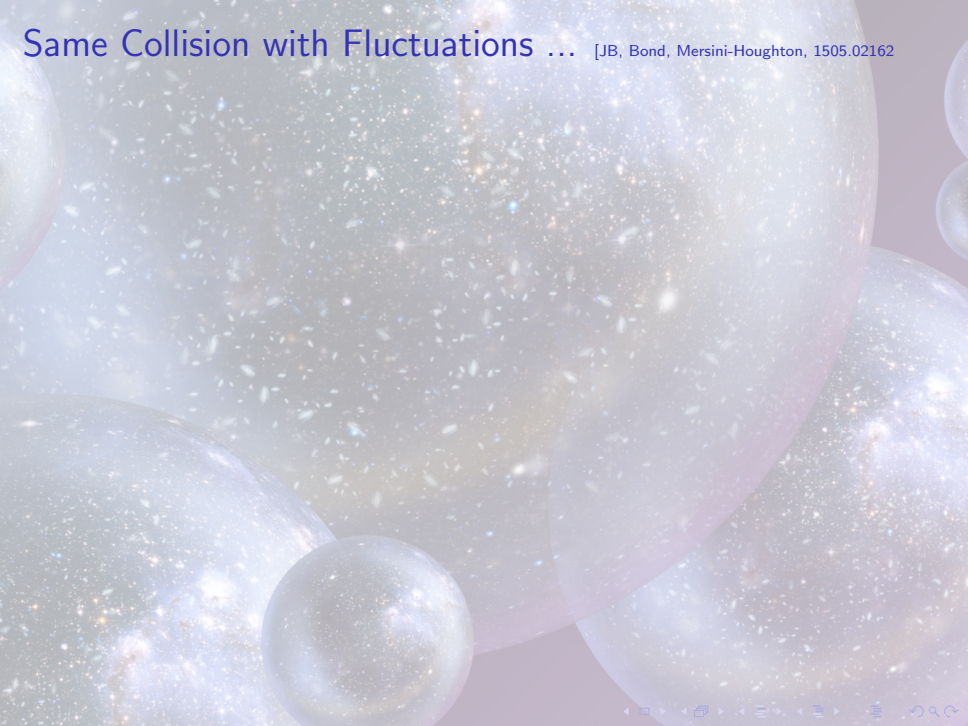
Three-Dimensional $SO(2,1)$ Collision



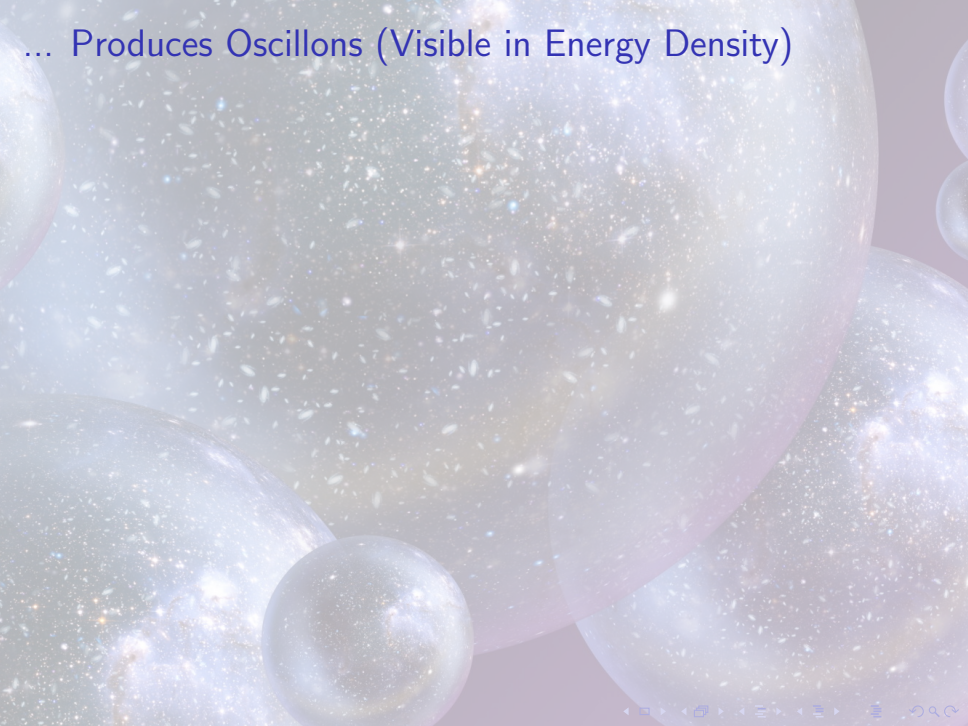
Numerical Preservation of $SO(2,1)$ Symmetry



Same Collision with Fluctuations ... [JB, Bond, Mersini-Houghton, 1505.02162

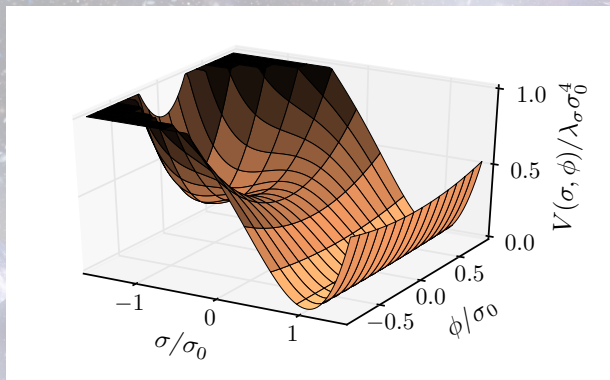


... Produces Oscillons (Visible in Energy Density)



Works for multifield models supporting inflation as well

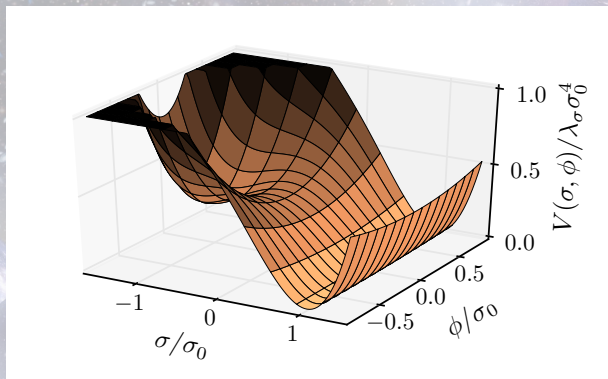
$$V(\sigma, \phi) = \frac{\lambda_\sigma \sigma_0^4}{4} \left[\left(\frac{\sigma^2}{\sigma_0^2} - 1 \right)^2 - \frac{\delta}{\lambda_\sigma} \frac{\sigma}{\sigma_0} \right] + \epsilon \phi + \frac{g^2}{2} \phi^2 \sigma^2$$



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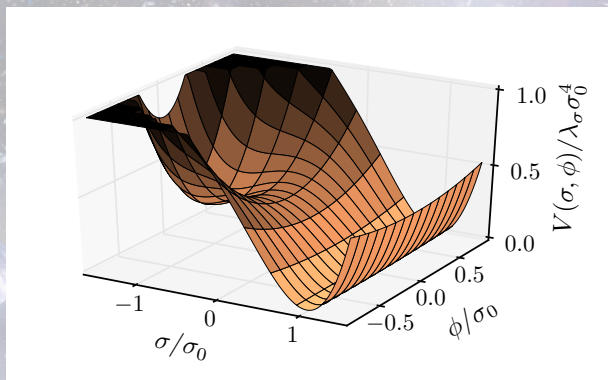
► V_{tunnel}



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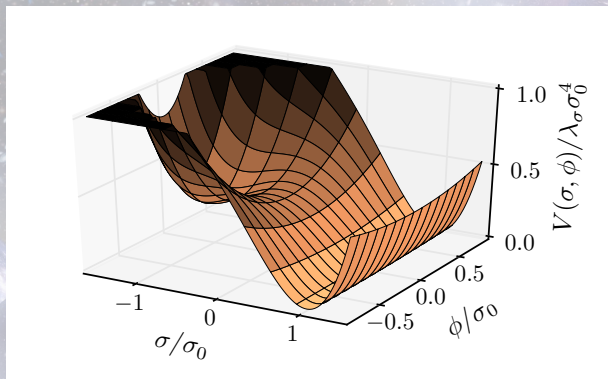
- ▶ V_{tunnel}
- ▶ V_{inflaton}



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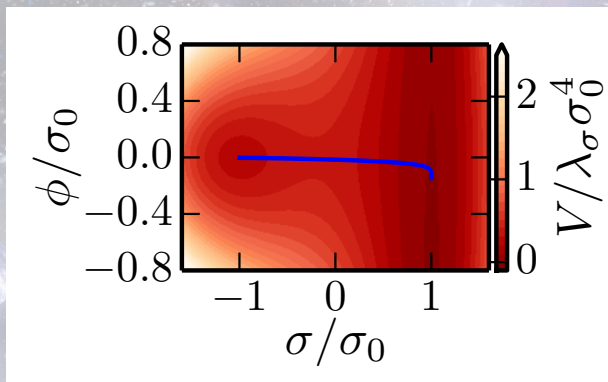
- ▶ V_{tunnel}
- ▶ V_{inflaton}
- ▶ V_{coupling}



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- ▶ V_{tunnel}
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Evolution of σ and ϕ

σ Evolution

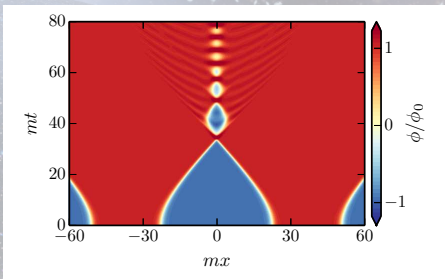
ϕ Evolution

Why Does This Happen?

Linear Parametric Resonance [JB, Bond, Mersini-Houghton, 1412.5591]

Non-SO(2,1) fluctuations evolve in the symmetric background

$$\frac{\partial^2 \phi_{bg}}{\partial s^2} + \frac{2}{s} \frac{\partial \phi_{bg}}{\partial s} - \frac{\partial^2 \phi_{bg}}{\partial x^2} + V'(\phi_{bg}) = 0$$
$$\left[\frac{\partial^2}{\partial s^2} - \frac{\partial^2}{\partial x^2} - \frac{\nabla_{H_2}^2}{s^2} + V''(\phi_{bg}) \right] (s\delta\phi) = 0$$



Floquet Theory

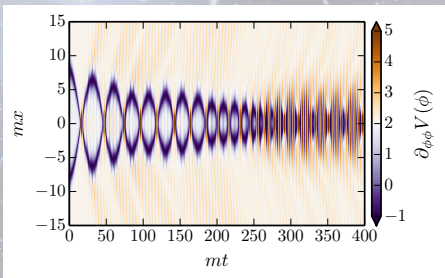
c.f. Preheating [Kofman, Linde, Starobinski '97]

$$\text{Oscillating } V''(\phi_{bg}) \\ \implies \delta\phi \sim e^{\mu t} P(x, t)$$

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$$\frac{\partial^2 \phi_{bg}}{\partial t^2} - \frac{\partial^2 \phi_{bg}}{\partial x^2} + V'(\phi_{planar}) = 0$$
$$\left[\frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial x^2} + (k_{\perp}^2 + V''(\phi_{planar})) \right] \delta \tilde{\phi}_{k_{\perp}} = 0$$



Floquet Theory

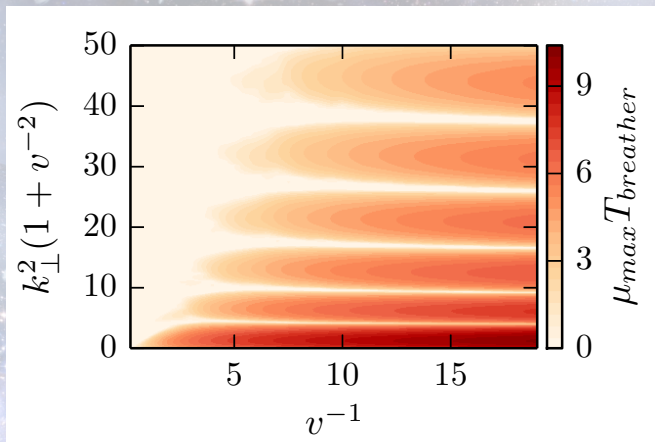
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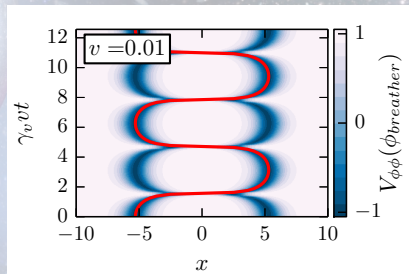
Exponentially Growing Modes Exist

$$V(\phi) = 1 - \cos(\phi)$$

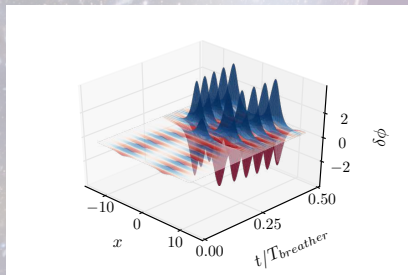
$$\phi_{breather} = 4 \tan^{-1} \left(\frac{\cos(\gamma_v vt)}{v \cosh(\gamma_v x)} \right) \quad \gamma_v \equiv (1 + v^2)^{-1/2}$$



Broad Resonance for Colliding Walls



$$v = 0.01$$

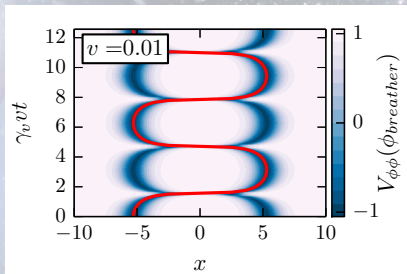


Generic Instability

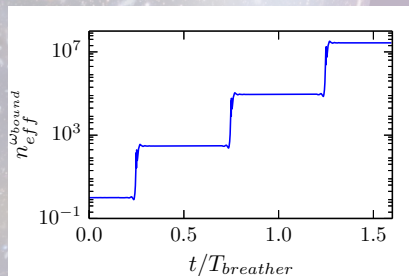
- ▶ Goldstones of Spontaneously Broken Translation Invariance
- ▶ Time-dependent wall “tension”

Broad Resonance for Colliding Walls

$$n_k^{eff} \propto \frac{1}{2k_{\perp}} \int_{-\infty}^{\infty} dx \left(\delta\dot{\phi}_k^2 + k_{\perp}^2 \delta\phi^2 \right)$$



$$v = 0.01$$



Generic Instability

- ▶ Goldstones of Spontaneously Broken Translation Invariance
- ▶ Time-dependent wall “tension”

Implications

SO(2,1) symmetry can be badly broken

Observables don't necessarily have azimuthal symmetry

- ▶ Beam smoothing versus inhomogeneity scale
- ▶ Tensor modes are produced by fracturing of walls
- ▶ Inhomogeneous start to inflation in some regions
- ▶ Sign of $\zeta = \delta \ln(a)$ in one field versus two field model

Qualitative conclusions don't depend on inflationary scenario

- ▶ Oscillons as nonequilibrium environment for baryogenesis?
- ▶ Oscillons dilute as $a^{-3} \rightarrow$ perturbed EOS during phase transition?
- ▶ Application to braneworlds with colliding walls
- ▶ Preheating in unwinding inflation?
- ▶ Bubble baryogenesis

These signals are spatially **intermittent**