

# Outline of the talk

## ① Introduction

- ▶ Summary **escape model** & its approach

## ② A recent nearby SN?

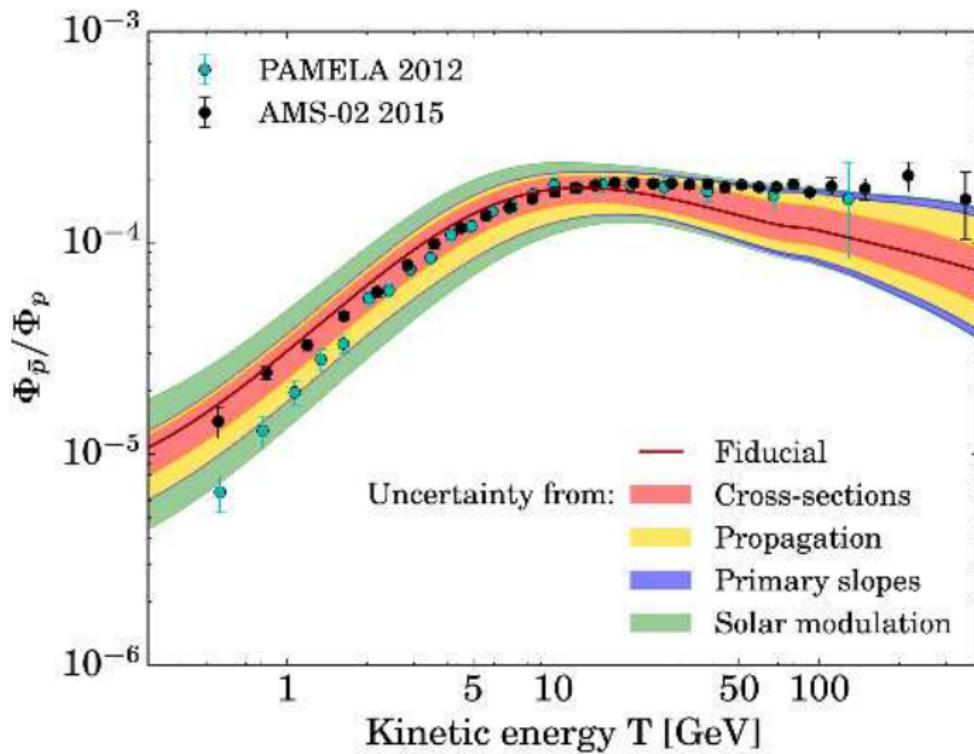
- ▶ Anisotropy [V.Savchenko, MK, D.Semikoz' 15 ]
- ▶ Antimatter fluxes [MK, A.Nernov, D.Semikoz' 15 ]

## ③ Antiproton production in hadronic collisions [MK, I.Moskalenko, S.Ostapchenko '15 ]

## ④ Conclusions

## Uncertainties in $\bar{p}$ flux prediction

[Giesen et al. [1504.04276] ]



# CR propagation in the escape model

[Giacinti, MK, Semikoz '14, '15]

- Standard approach: **diffusion** as effective theory
  - ▶ **effective parameters**: not well constrained, often degenerated ( $D/h$ )
  - ▶ fast, but too **simplistic**: geometry uniform box,  $D_{ij} = D\delta_{ij}$ , etc,
- our approach: **trajectories** of individual CRs in **GMF model**
  - + adds info from RM + polarised synchrotron:  $D = \text{const} \rightarrow \mathbf{B}(\mathbf{x})$
  - + easy to include  $n_{\text{gas}}(\mathbf{x})$  and  $n_{\text{src}}(\mathbf{x}, t)$  (like spiral structure)
  - computationally expansive

⇒ model reproduces all CR data from  $\sim 200 \text{ GeV}$  up  $10^{18} \text{ eV}$

[talk in VHE session ]

# Anisotropy of a single source

- if **only turbulent field**:

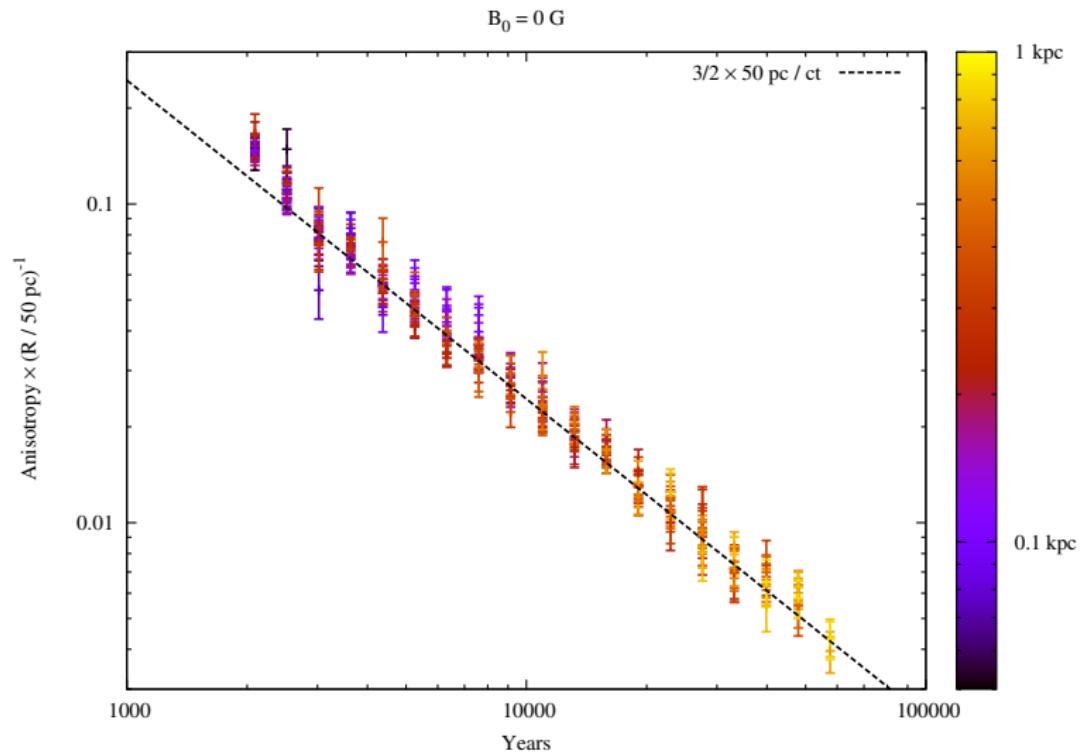
diffusion = random walk = free quantum particle

- number density is **Gaussian** with  $\sigma^2 = 4DT$

$$\delta = \frac{3D}{c} \frac{\nabla n}{n} = \frac{3R}{2cT}$$

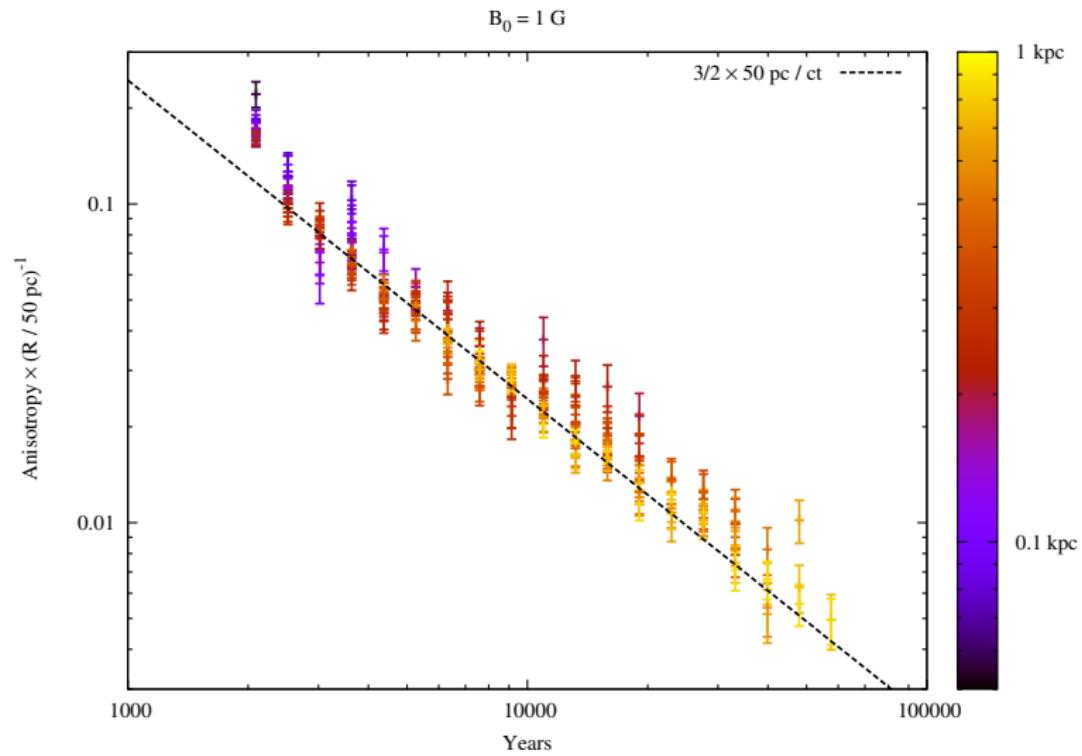
- what happens for general fields?

# Anisotropy of a single source: only turbulent field

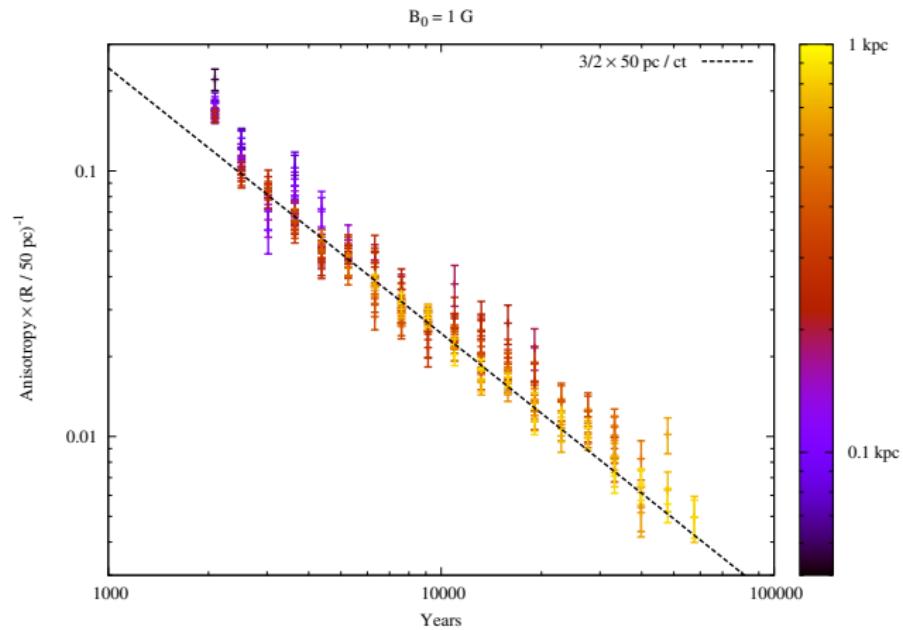


[Savchenko, MK, Semikoz '15]

# Anisotropy of a single source: plus regular

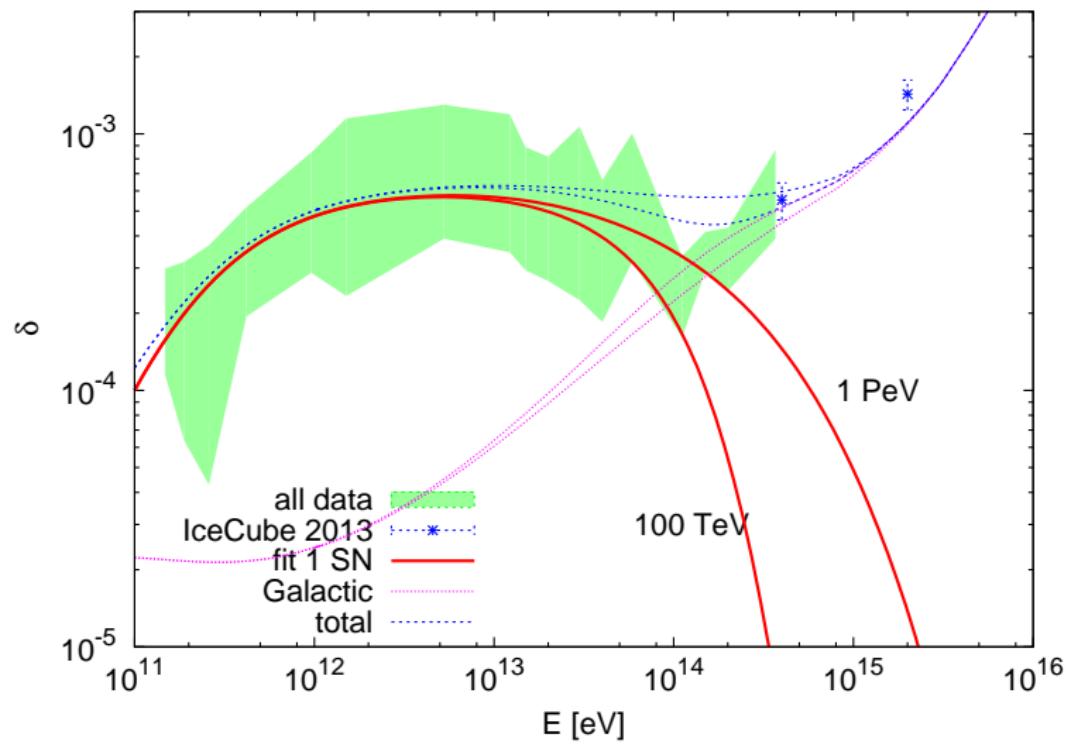


# Anisotropy of a single source:



- regular field changes  $n(x)$ , but keeps it Gaussian  
⇒ no change in  $\delta$

# Anisotropy of a single source:



[Savchenko, MK, Semikoz '15]

# Single source: other signatures

- 2 Myr SN explains anomalous  $^{60}\text{Fe}$  sediments

[Ellis+ '96 ]

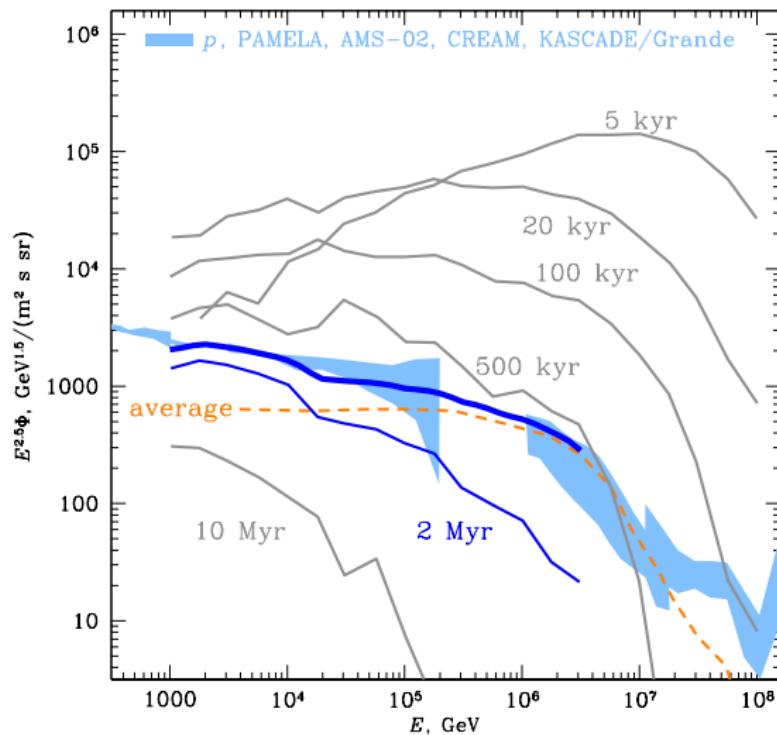
## Single source: other signatures

- 2 Myr SN explains anomalous  $^{60}\text{Fe}$  sediments [Ellis+ '96]
- secondaries:
  - ▶  $\bar{p}$  diffuse as  $p \Rightarrow$  leads to **constant  $\bar{p}/p$  ratio**
  - ▶  $\bar{p}/p$  ratio fixed by source age and  $\delta \Rightarrow \bar{p}$  flux is predicted
  - ▶  $e^+$  flux is predicted
  - ▶ relative ratio of  $\bar{p}$  and  $e^+$  depends only on their  $Z$  factors

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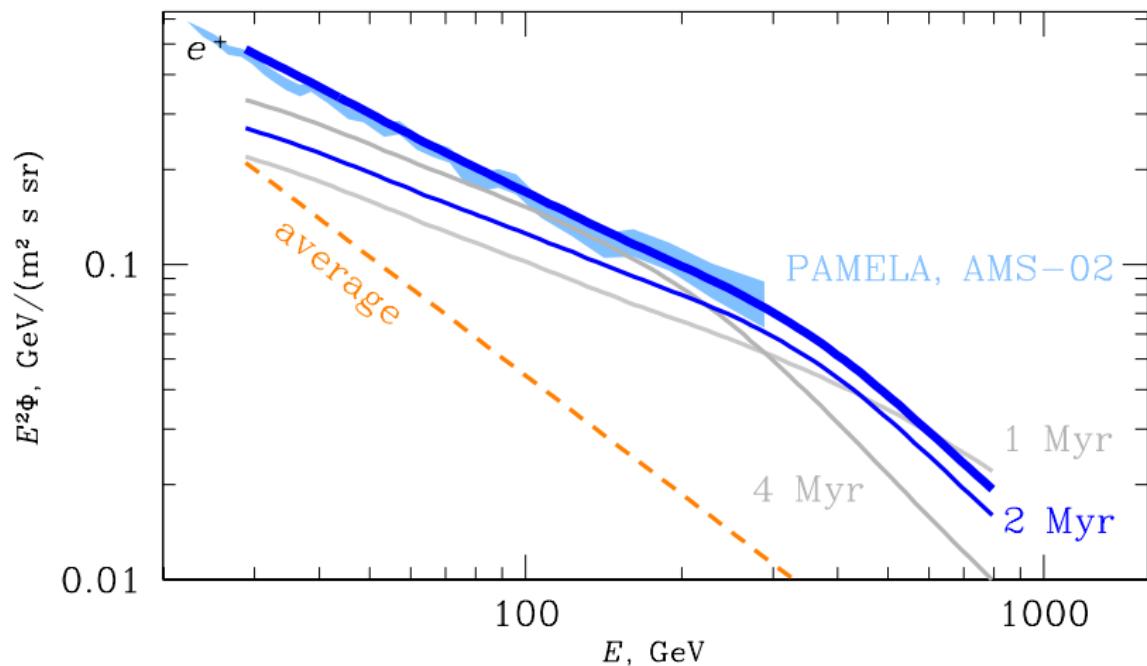
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- may responsible for different slopes of local  $p$  and nuclei fluxes

# Single source: proton flux



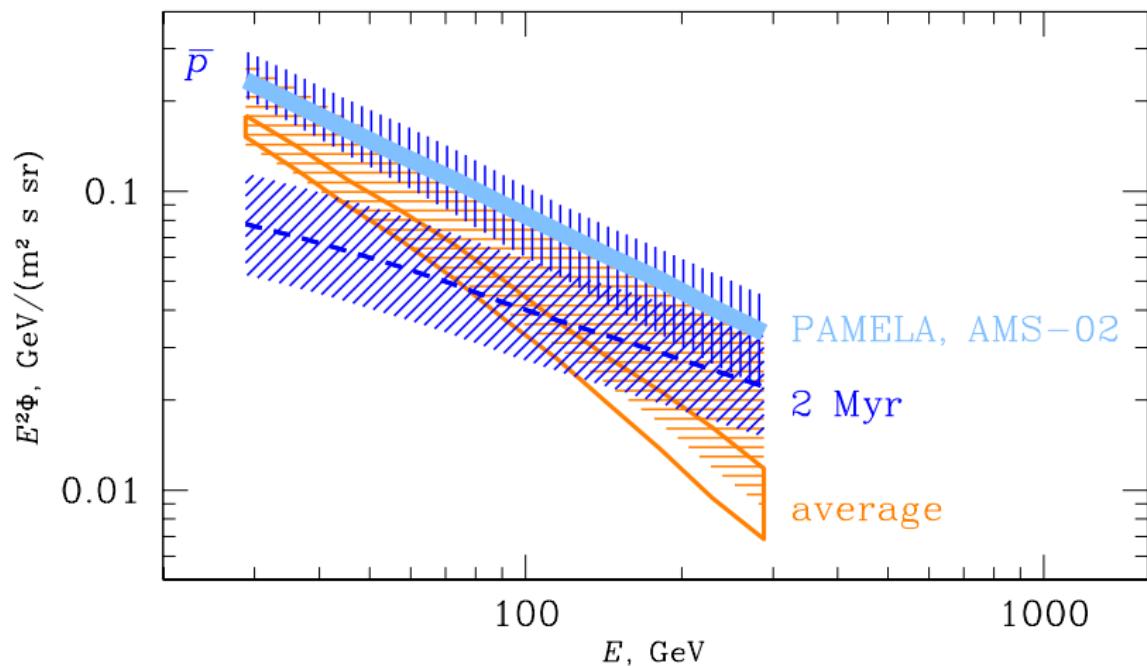
[MK, Neronov, Semikoz '15]

## Single source: positrons



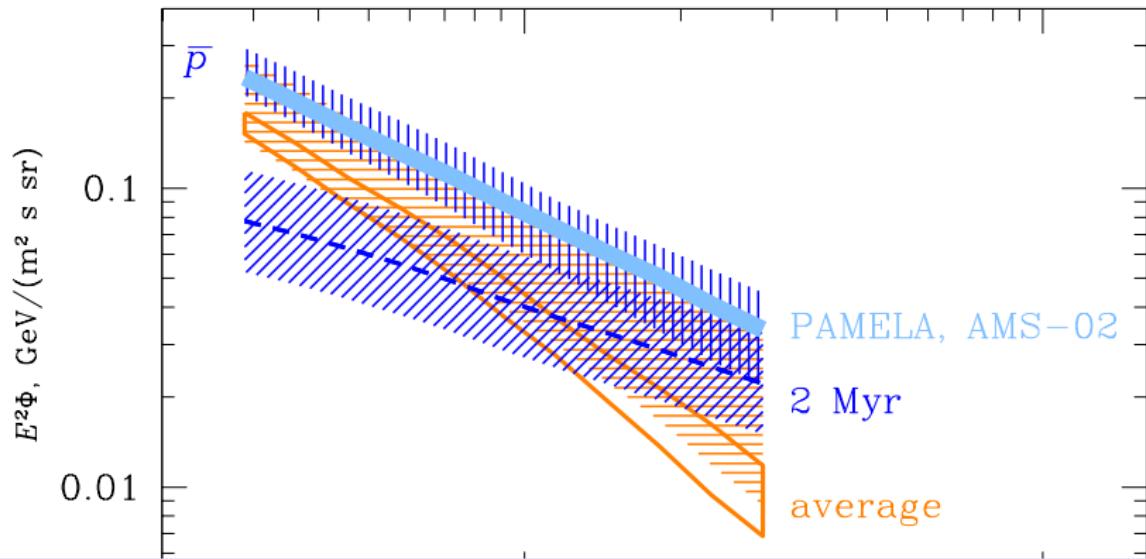
[MK, Neronov, Semikoz '15]

# Single source: antiprotons



[MK, Neronov, Semikoz '15]

# Single source: antiprotons



$\bar{p}$  flux = average + loc. source + DM

- ▶ DM is not needed, but how to constrain it best?

[MK, Neronov, Semikoz '15]

# Fitting $\bar{p}$ production vs. modelling

- common problems:

- ▶ low energies  $\cong$  old exp.  $\cong$  large, badly documented syst. errors
- ▶ Ex.: some “pp” measurements are rescaled  $pA$  data
- ▶ small  $\Omega$  coverage, typically fixed angle
- ▶ low  $E$  not covered

- fitting:

- ▶ required extrapolation depends on quality of fit function
- ▶ based on obsolet Ng&Tang parametrisations

- Simulations:

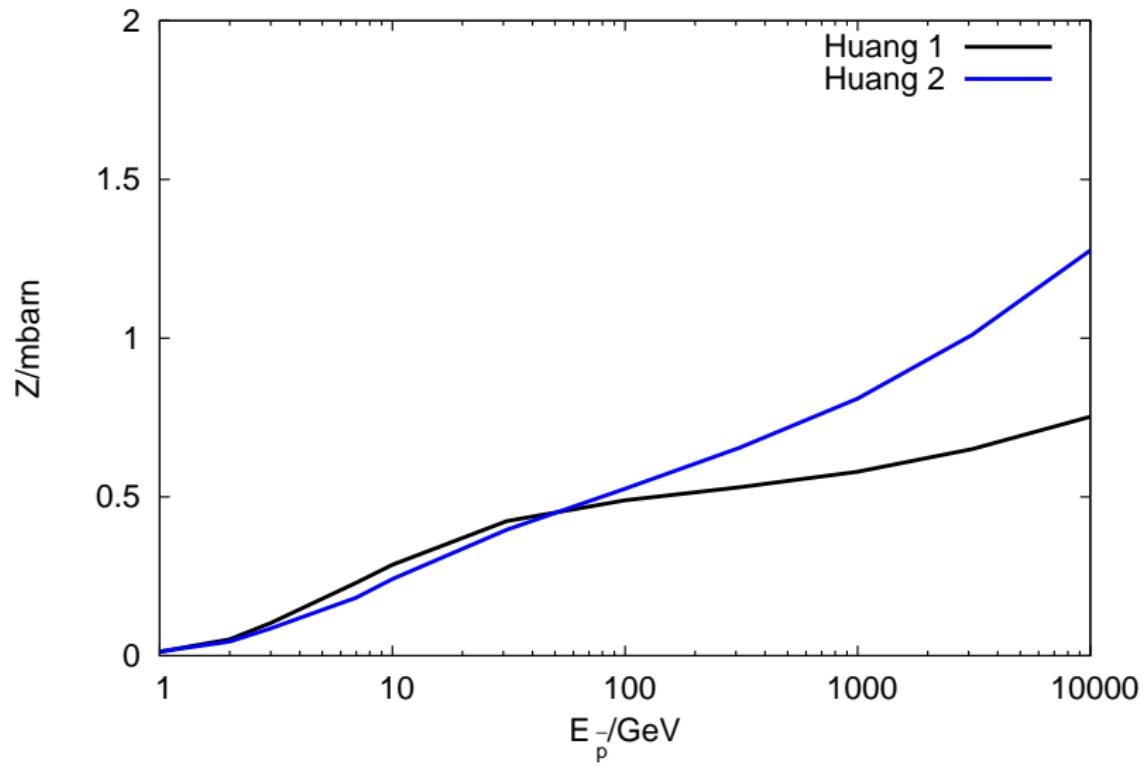
- ▶ require models for soft interactions and hadronisation
- ▶ models like QGSJet or EPOS calibrated on large data sets (SPS, Tevatron, LHC, Na49, ..., CR)
- ▶ consistent framework for pA, Ap and AA collisions

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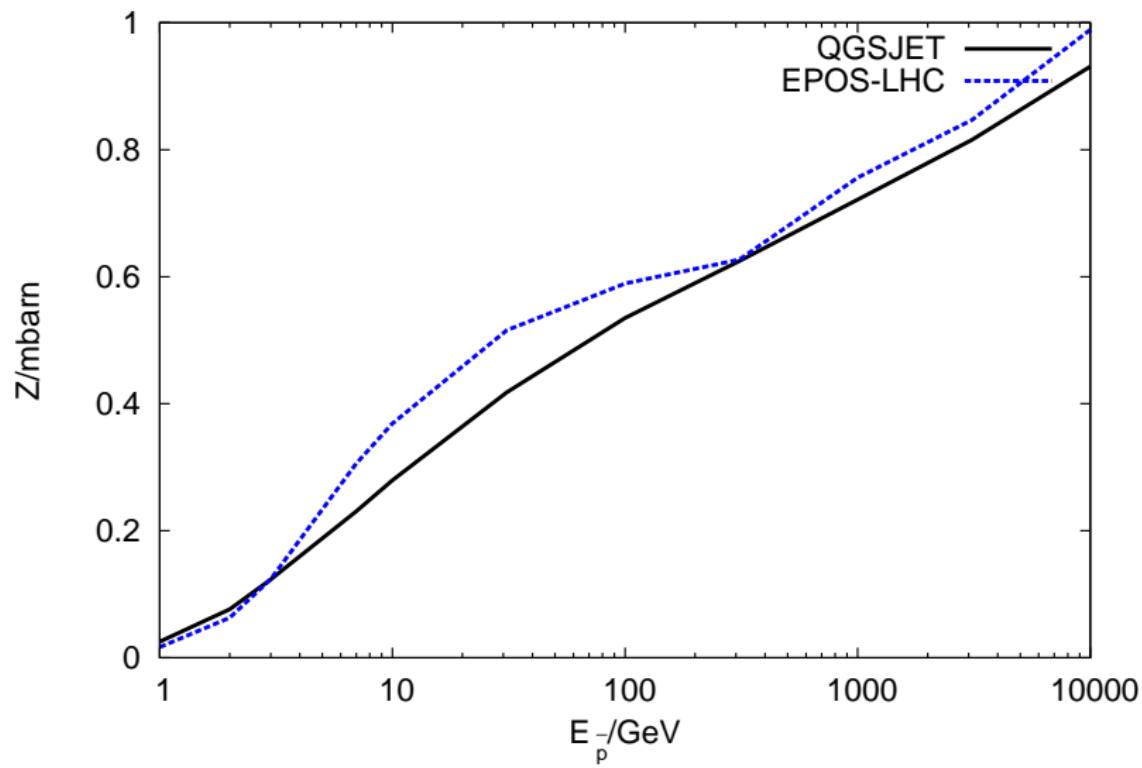
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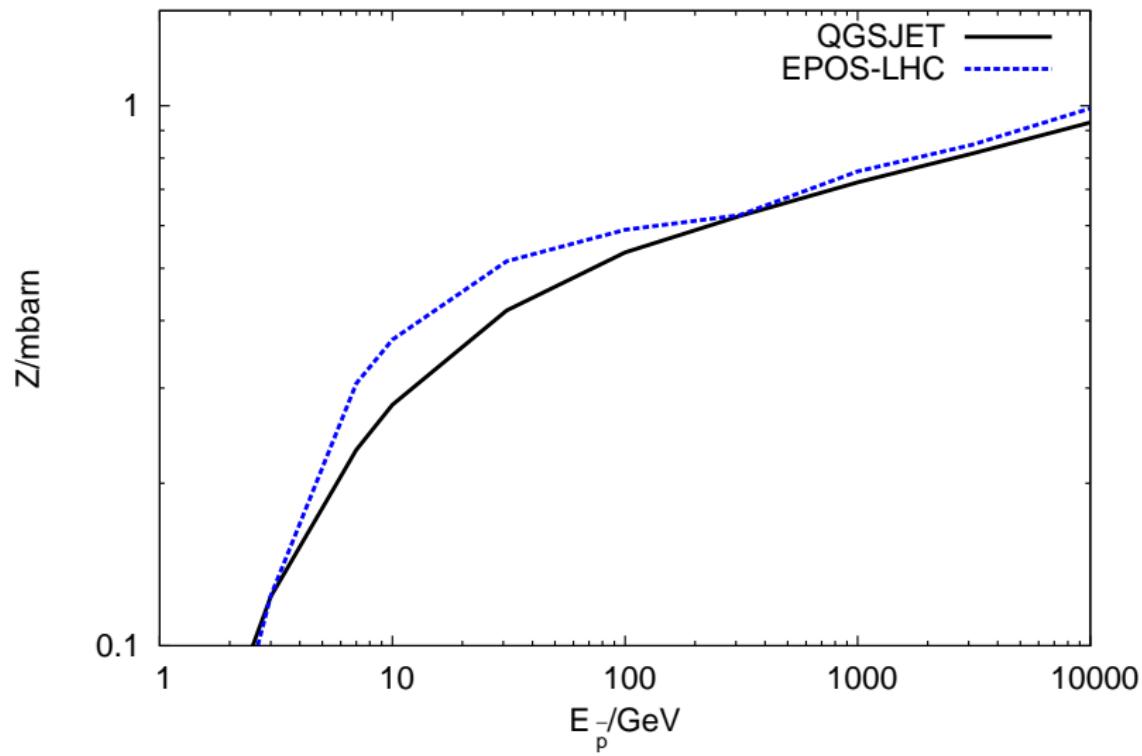
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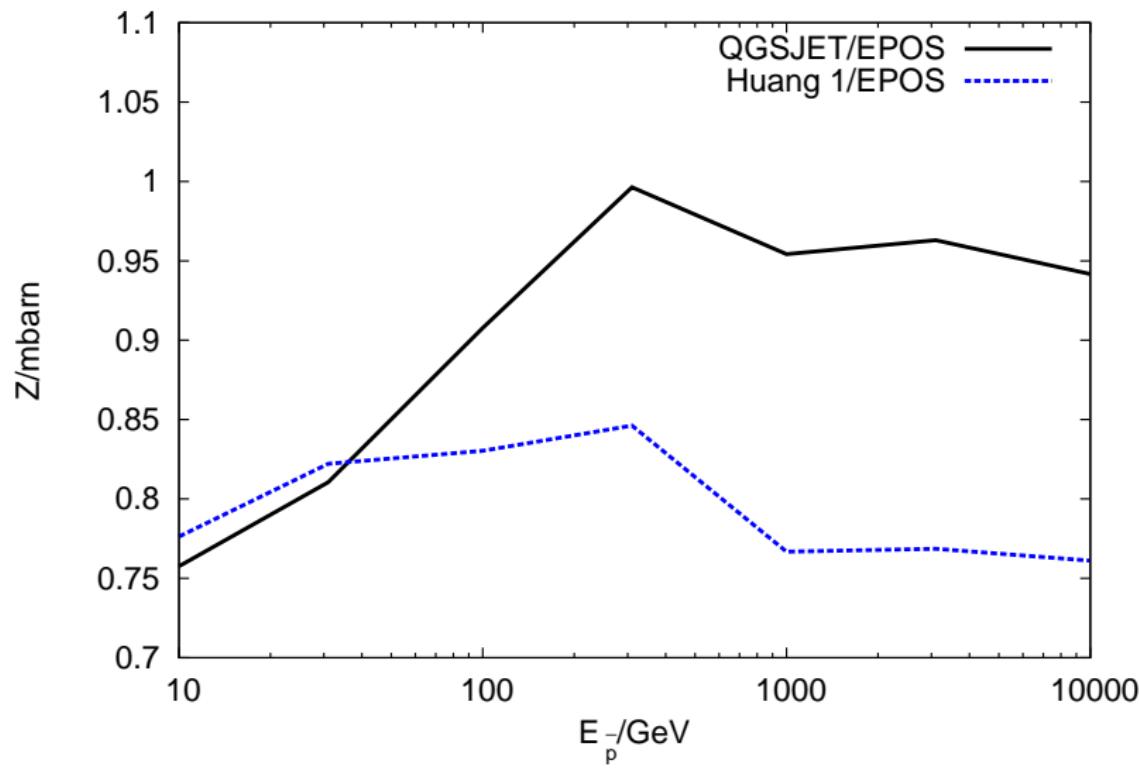
Fitting: Variation of Huang models ( $\alpha = 2$ )

# Comparison of QCD models ( $\alpha = 2$ )

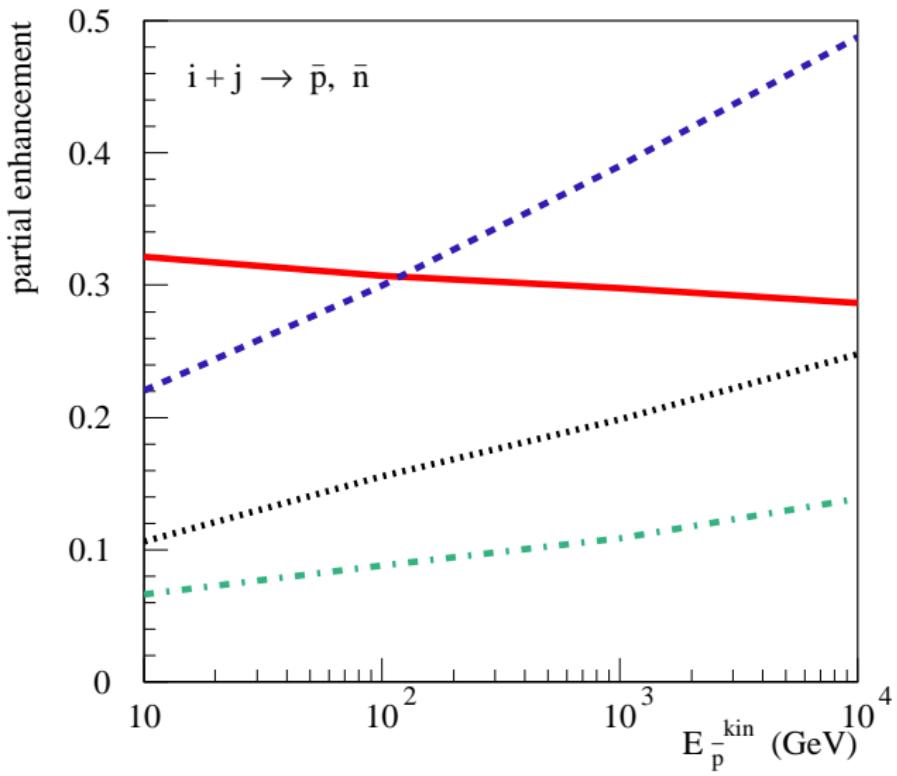


# Comparison of QCD models ( $\alpha = 2$ )



Ratios of  $Z$ -factors ( $\alpha = 2$ )

# Nuclear enhancement: adding He ( $\alpha = 2$ )



# Conclusions

## ① Single source: anisotropy

- ▶ dipole formula  $\delta = 3R/2T$  holds universally in quasi-gaussian regime
- ▶ plateau of  $\delta$  points to dominance of single source

## ② Single source: antimatter

- ▶ consistent explanation of  $p$ ,  $\bar{p}$  and  $e^+$  fluxes
- ▶ explains  $\gamma$  flux  $\propto E^{-2.5}$
- ▶ consistent with  $^{60}\text{Fe}$  and  $\delta$

## ③ Uncertainty in $\sigma(pp \rightarrow \bar{p})$ :

- ▶  $E_{\bar{p}} \gtrsim 100 \text{ GeV}$ : models agree within 15%
- ▶ below: no improvement without new exp. data
- ▶ parametrisations:  $\varepsilon_{\text{nuc}}$  adds additional uncertainty