First hundred years of GR: successes, status and prospects

T.Padmanabhan IUCAA, Pune, India

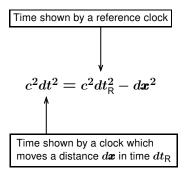
"It ... represents probably most beautiful of all existing physical theories."

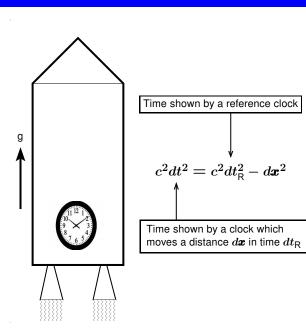
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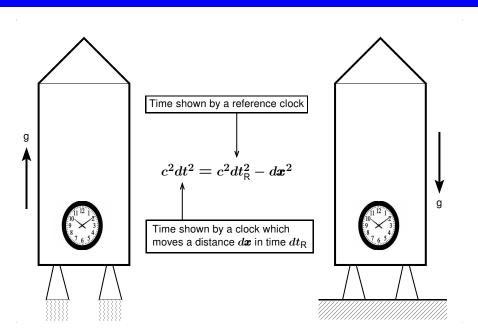
Landau and Liftshitz Volume II of the Course of Theoretical Physics

Principle of Equivalence + Special Relativity

Gravity = Curvature of Spacetime







$$V^2 = v^2 + 2gx = v^2 - 2\phi$$

$$c^2 dt^2 = c^2 dt_R^2 + 2\phi dt_R^2 - v^2 dt_R^2$$

$$c^2dt^2=c^2\left(1+rac{2\phi}{c^2}
ight)dt_R^2-doldsymbol{x}^2$$

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$$\phi(t, {m x}) o g_{ab}(t, {m x})$$

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ight)dt_R^2-doldsymbol{x}^2$$

This metric represents a curved spacetime

$$\phi(t, {m x}) o g_{ab}(t, {m x})$$

"This fusion of ... metric and gravitation must be considered as the most beautiful achievement of GR." – W. Pauli

$$abla^2 \phi \propto
ho = T_{ab} \ u^a \ u^b = \left(egin{array}{l} ext{Energy density} \ ext{measured by} \ ext{that observer} \end{array}
ight)$$

$$\left(egin{array}{c} {\sf Radius \ of \ curvature} \\ {\sf of \ space \ orthogonal} \\ {\sf to \ an \ observer} \end{array}
ight)^{-2} \propto \left(egin{array}{c} {\sf Energy \ density} \\ {\sf measured \ by} \\ {\sf that \ observer} \end{array}
ight)$$

$$\left(\begin{array}{c} \text{Radius of curvature} \\ \text{of space orthogonal} \\ \text{to an observer} \end{array} \right)^{-2} \propto \left(\begin{array}{c} \text{Energy density} \\ \text{measured by} \\ \text{that observer} \end{array} \right)$$

$$egin{aligned} \mathcal{R}_{ijkl} &\equiv R_{abcd} \ u^a_i \ u^b_j \ u^c_k \ u^d_l \ R^{-2}_{ ext{curv}} &\equiv \mathcal{R}^{ij}_{ij} = 2 G^a_b u^b u_a \end{aligned}$$

$$ho = T_{ab} \, u^a \, u^b$$

$$\left(\begin{array}{c} \text{Radius of curvature} \\ \text{of space orthogonal} \\ \text{to an observer} \end{array} \right)^{-2} \propto \left(\begin{array}{c} \text{Energy density} \\ \text{measured by} \\ \text{that observer} \end{array} \right)$$

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ho = T_{ab} \ u^a \ u^b$$

$$R_{
m curv}^{-2}=16\pi G
ho$$

PRECESSION OF MERCURY

BENDING OF LIGHT

GRAVITATIONAL WAVES

Observed precession $\sim 574''/cent$ Planetary perturbations $\sim 532''/cent$ Difference $\sim 42''/cent$

Einstein used this as a bench mark!

First attempt was with wrong field equation and with approximate metric!

Discovered within few months; understood after few decades!

thin four months.

Discovered within few months; understood after few decades!

$$ds^2 = -dt^2 + rac{dx^2}{
ho^4(x)} +
ho^2 \left(rac{d\psi^2}{\sin^2 heta} + \sin^2 heta d\phi^2
ight)$$

$$x \equiv \rho^3/3; \quad \psi \equiv -\cos\theta$$

Discovered within few months; understood after few decades!

$$ds^2 = -f_0 dt^2 + f_1 dx^2 + f_2 \left(rac{d\psi^2}{\sin^2 heta} + \sin^2 heta d\phi^2
ight)$$

$$f_0 f_1 f_2^2 = -1$$

The First Exact Solution

understood after few decades!

 $ds^2=-f_0dt^2+f_1dx^2+f_2\left(rac{d\psi^2}{\sin^2 heta}+\sin^2 heta d\phi^2
ight)$

 $f_0 f_1 f_2^2 = -1$

Looked non-singular at origin, causing considerable confusion!

$$ds^2 = -(1{+}2\Phi)dt^2{+}rac{dr^2}{(1+2\Phi)}{+}r^2(d heta^2{+}\sin^2 heta d\phi^2)$$

$$T_0^0=T_r^r=
ho(r); \quad T_{ heta}^ heta=T_{\phi}^\phi=\mu(r)$$

Einstein's equations become linear in $\Phi!$

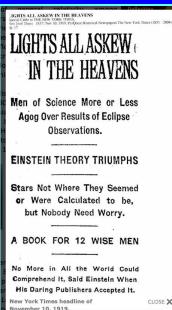
$$\Phi = -rac{lpha}{r} + rac{G}{r} \int 4\pi r^2
ho(r) dr; \quad \mu =
ho + rac{1}{2} r
ho'(r)$$

Bending of Light

A small effect in 1919 ...



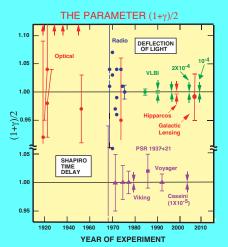
Eddington's telegram to Einstein announcing the observation of the bending of light

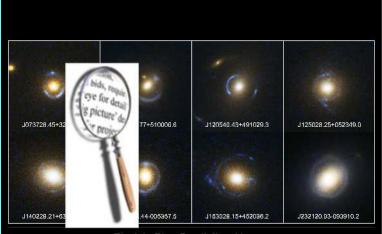


Bending of Light

...a large effect today!

$$heta=1.7505$$
 arc-seconds $\dfrac{1}{2}(1+\gamma)\dfrac{1+\cos\Phi}{2}$





Einstein Ring Gravitational Lenses
Hubble Space Telescope • Advanced Camera for Surveys

NASA, ESA, A. Bollon (Harvard-Smithson an CIA), and the SLACS Team

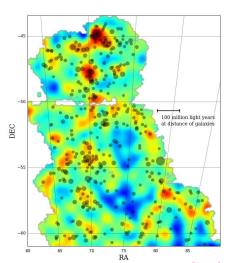
STScI-PRC05-32

Einstein (1936): Lensing by stars:

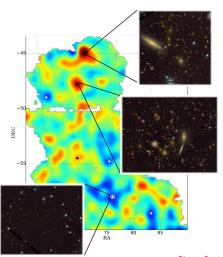
"It is of little value."

Zwicky (1937): Lensing by galaxies:

Visionary 1-page paper: testing relativity,
magnifying faint objects, measuring
masses



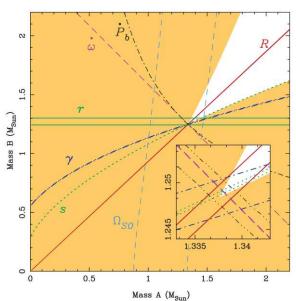
Chang C. et al., 2015; Vikram V. et al., 2015



Chang C. et al., 2015; Vikram V. et al., 2015

Gravitational Wave Emission





Source: M. Kramer

Gravitational Wave Emission

$$m_1 = 1.4398 \pm 0.0002$$

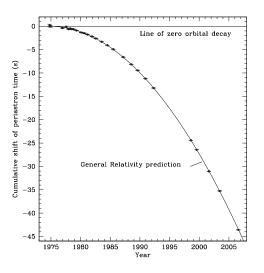
$$m_2 = 1.3886 \pm 0.0002$$

$$\dot{P} = (-2.402531 \pm 0.000014) \times 10^{-12}$$

$$P_{obs}/P_{GR} = 0.997 \pm 0.002$$

Gravitational Wave Emission

B1913 + 16



Two Major Areas Of Impact

BLACK HOLES IN ASTROPHYSICS

THE EXPANDING UNIVERSE

Black Holes in Astrophysics

"...the relativists...are suddenly experts in a field they hardly knew existed;..."

Thomas Gold 1st Texas Astrophysics

Black Holes in Astrophysics

Powers the AGN: $M pprox (10^6-10^{10}) M_{\odot}$

X-ray Binaries: $M \approx (5-20) M_{\odot}$

Represents a rotating black hole with an angular momentum parameter a=J/M

$$\begin{split} ds^2 &= -\left(1 - \frac{2\mu r}{\rho^2}\right)dt^2 - \frac{4\mu ar\sin^2\theta}{\rho^2}dtd\phi + \frac{\rho^2}{\Delta}dr^2 + \rho^2d\theta^2 \\ &+ \left(r^2 + a^2 + \frac{2\mu ra^2\sin^2\theta}{\rho^2}\right)\sin^2\theta d\phi^2 \end{split}$$

$$\rho^2 \equiv r^2 + a^2 \cos^2 \theta \; , \qquad \Delta \equiv r^2 - 2\mu r + a^2 \label{eq:rho2}$$

Kerr Metric in the Sky

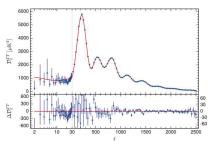
BH Masses and Spins

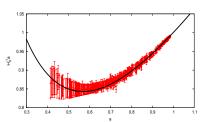
Source Name	BH Mass (M ₀)	BH Spin (a,)
A0620-00	6.3–6.9	0.12 ± 0.19
LMC X-3	5.9-9.2	//_0.25 //
XTE J1550-564	8.5-9.7	0.34±0.24
GRO J1655-40	6.0-6.6	0.70 ± 0.05
4U1543-47	8.4—10.4	0.80 ± 0.05
M33 X-7	14.2–17.1	0.84 ± 0.05
LMC X-1	9.4—12.4	0.92/±/0.06
Cyg X-1	13.8—15.8	//> 0.97 //
GRS 1915+105	10-18	///>0.98 //

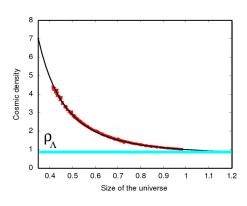
Shafee et al. (2006); McClintock et al. (2006); Davis et al. (2006); Liu et al. (2007,2009); Gou et al. (2009,2010, 2011); Steiner et al. (2010)

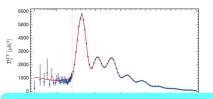
A missed opportunity by Einstein!

Tremendous progress in recent years



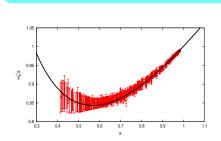


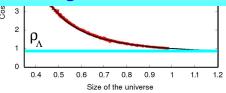


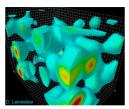




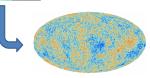
Observations are way ahead of theoretical understanding!





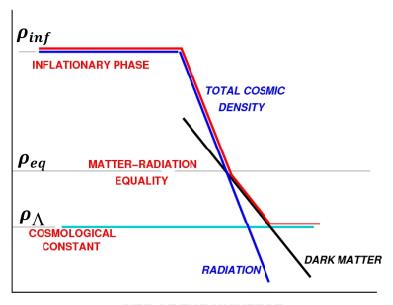


VACUUM FLUCTUATIONS IN THE VERY EARLY UNIVERSE



COSMIC STRUCTURES
SEEN TODAY





SIZE OF THE UNIVERSE

The Signature Of The Universe

$$\left(rac{\dot{a}}{a}
ight)^2 = \left\{ egin{array}{l} rac{8\pi G}{3}
ho_{
m inf} \ & \ rac{8\pi G}{3} \left[
ho_{\Lambda} +
ho_{
m eq} \left(\left(rac{a_{
m eq}}{a}
ight)^3 + \left(rac{a_{
m eq}}{a}
ight)^4
ight)
ight] \end{array}
ight.$$

The Signature Of The Universe

$$ho_{
m inf} < (1.94 imes 10^{16}~{
m GeV})^4$$

$$ho_{
m eq} = rac{
ho_m^4}{
ho_R^3} = [(0.86 \pm 0.09) \ {
m eV}]^4$$

$$ho_{\Lambda} = [(2.26 \pm 0.05) imes 10^{-3} ext{eV}]^4$$

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Three Major Challenges

SINGULARITIES: BLACK HOLES, UNIVERSE

COSMOLOGICAL CONSTANT

THE THERMODYNAMIC CONNECTION

"The existence of spacetime singularities represents an end to ... the predictability gained by science. How could physics lead to ... no physics?"

– John Wheeler

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QUANTUM EFFECTS AT $A_{Planck} = (G\hbar/c^3)$?

The Cosmological Constant

Gravity breaks a symmetry H o H + C of the matter sector

The Cosmological Constant

Gravity breaks a symmetry H o H + C of the matter sector

But then:

- It seems unaffected by changes in the zero-level of the energy
- ▶ It couples to a small cosmological constant

$$\Lambda\left(rac{G\hbar}{c^3}
ight)pprox 10^{-123}$$

Spacetimes, Like Matter, can be Hot

The most beautiful result in the interface of quantum theory and gravity

Spacetimes, Like Matter, can be Hot

The most beautiful result in the interface of quantum theory and gravity

OBSERVERS WHO PERCEIVE A HORIZON ATTRIBUTE A TEMPERATURE TO SPACETIME

$$k_BT=rac{\hbar}{c}\left(rac{g}{2\pi}
ight)$$

[Davies (1975), Unruh (1976)]

Spacetimes, Like Matter, can be Hot

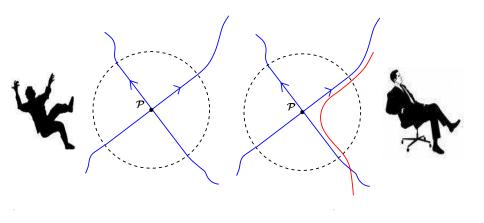
The most beautiful result in the interface of quantum theory and gravity

OBSERVERS WHO PERCEIVE A HORIZON ATTRIBUTE A TEMPERATURE TO SPACETIME

$$k_BT=rac{\hbar}{c}\left(rac{g}{2\pi}
ight)$$

[Davies (1975), Unruh (1976)]

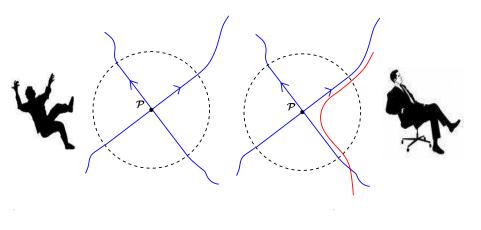
This allows you to associate a heat density $\mathcal{H} = Ts$ with every event of spacetime!



Vacuum fluctuations



Thermal fluctuations



Vacuum fluctuations



Thermal fluctuations

Why does spacetime exhibit thermal properties?!

The three challenges involve $\hbar!$

$$A_{Planck} = rac{G\hbar}{c^3}; \quad \Lambda\left(rac{G\hbar}{c^3}
ight) pprox 10^{-123}; \quad k_BT = rac{\hbar}{c}\left(rac{g}{2\pi}
ight)$$

HOW DO WE PUT TOGETHER THE PRINCIPLES OF QUANTUM THEORY AND GRAVITY?

Everybody Wants To Quantize Gravity!

Everybody Wants To Quantize Gravity!





... But Nobody Has Succeeded!

► The perturbative approach does not work

 Virtually every interesting question about gravity is non-perturbative by nature

No guiding principle; metric is assumed to be a quantum variable

GR: The Next 100 Years

Needs another paradigm shift!

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If gravity is immune to zero level of energy it must have a thermodynamic interpretation!

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Connects two features usually thought to be completely separate!

[TP, arxiv:1508.06286]

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Needs another paradigm shift!

If gravity is immune to zero level of energy it must have a thermodynamic interpretation!

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[TP, arxiv:1508.06286]

Classical Gravity has the same conceptual status as elasticity/hydrodynamics

If gravity is immune to zero level of energy it must have a thermodynamic interpretation!

Connects two features usually thought to be completely separate!

[TP, arxiv:1508.06286]

Study spacetime dynamics the way physicists studied fluids before knowing the atomic structure of matter

Atoms Of Spacetime

TP, arxiv: 1003.5665, 1508.06286

Atoms Of Spacetime

TP, arxiv: 1003.5665, 1508.06286

▶ To store energy ΔE at temperature T, you need $\Delta n = \Delta E/(1/2)k_BT$ degrees of freedom.

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- You can heat up spacetime!
- Count the microscopic spacetime degrees of freedom in terms of distribution function for atoms of spacetime!

- ▶ To store energy ΔE at temperature T, you need $\Delta n = \Delta E/(1/2)k_BT$ degrees of freedom.
- You can heat up spacetime!
- Count the microscopic spacetime degrees of freedom in terms of distribution function for atoms of spacetime!
- The distribution function determines the entropy density of spacetime.

Nature Of Gravity

$$\mathcal{H} = rac{Q}{V} = rac{TS}{V} = rac{1}{V}(E-F) = Ts$$

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$$Q=\int d\lambda\, d^2x\, \sqrt{\gamma}\, \left({\cal H}_g+{\cal H}_m
ight)$$

$$\mathcal{H} = rac{Q}{V} = rac{TS}{V} = rac{1}{V}(E - F) = Ts$$

$$Q = \int d\lambda \, d^2 x \, \sqrt{\gamma} \, \left({\cal H}_g + {\cal H}_m
ight) \, .$$

Works for a wide class gravitational theories; entropy decides the theory.

Gravity responds to heat density $(Ts = p + \rho)$ — not energy density!

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Cosmological constant arises as an integration constant.

Gravity responds to heat density (Ts = p +
ho) — not energy density!

Cosmological constant arises as an integration constant.

Its value is determined by a new conserved quantity for the universe!

What Makes Spacetime Evolve?

T.P., arXiv:1312.3253

$$\underbrace{\int rac{d\Sigma_a}{8\pi L_P^2} \left[q^{\ell m}\partial\; p_{\ell m}^a
ight]} = -rac{1}{2}k_B T_{
m av}\; \underbrace{\left(N_{
m sur}-N_{
m bulk}
ight)}$$

time evolution of spacetime

= heating of spacetime

deviation from holographic equipartition

$$\underbrace{\int rac{d\Sigma_a}{8\pi L_P^2} \left[q^{\ell m}\partial\; p_{\ell m}^a
ight]} = -rac{1}{2}k_B T_{
m av}\; \underbrace{\left(N_{
m sur}-N_{
m bulk}
ight)}$$

time evolution of spacetime

= heating of spacetime

deviation from holographic equipartition

This replaces the field equation for gravity

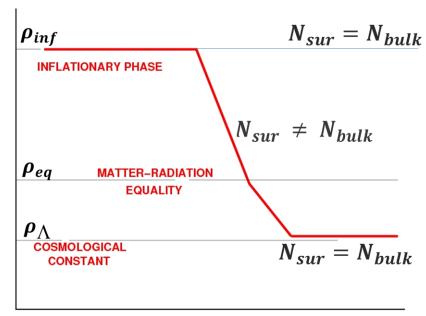
$$\underbrace{\int rac{d\Sigma_a}{8\pi L_P^2} [q^{\ell m}\partial\; p_{\ell m}^a]}_{} = -rac{1}{2} k_B T_{
m av} \; \underbrace{(N_{
m sur}-N_{
m bulk})}_{}$$

time evolution of spacetime

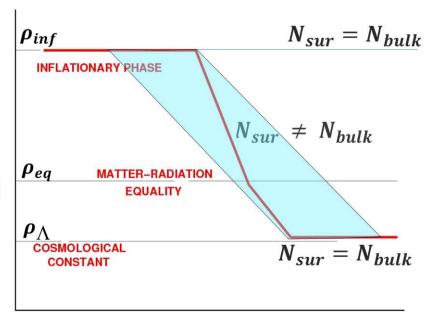
= heating of spacetime

deviation from holographic equipartition

$$rac{d\mathcal{R}}{dt} = 1 - rac{N_{
m bulk}}{N_{
m sur}}$$



SIZE OF THE UNIVERSE



SIZE OF THE UNIVERSE

$$ho_{\Lambda} = rac{4}{27} \, rac{
ho_{inf}^{3/2}}{
ho_{eq}^{1/2}} \, \exp(-36\pi^2)$$

Makes a falsifiable prediction!

$$\rho_{inf} = (1-6) \times 10^{15} GeV$$