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# New tools for better-characterizing astroparticle physics sites and detectors for the Latin American Giant Observatory

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**IWARA**

From Quarks to Cosmos

**IWARA 2022**

11th. International Workshop on Astronomy  
and Relativistic Astrophysics.  
Antigua Guatemala, Guatemala.  
5 - 9 September, 2022



Horizon 2020 RI project 857647

# LAGO



LAGO is a giant network of astroparticle detectors at global scale, currently operating in 11 countries

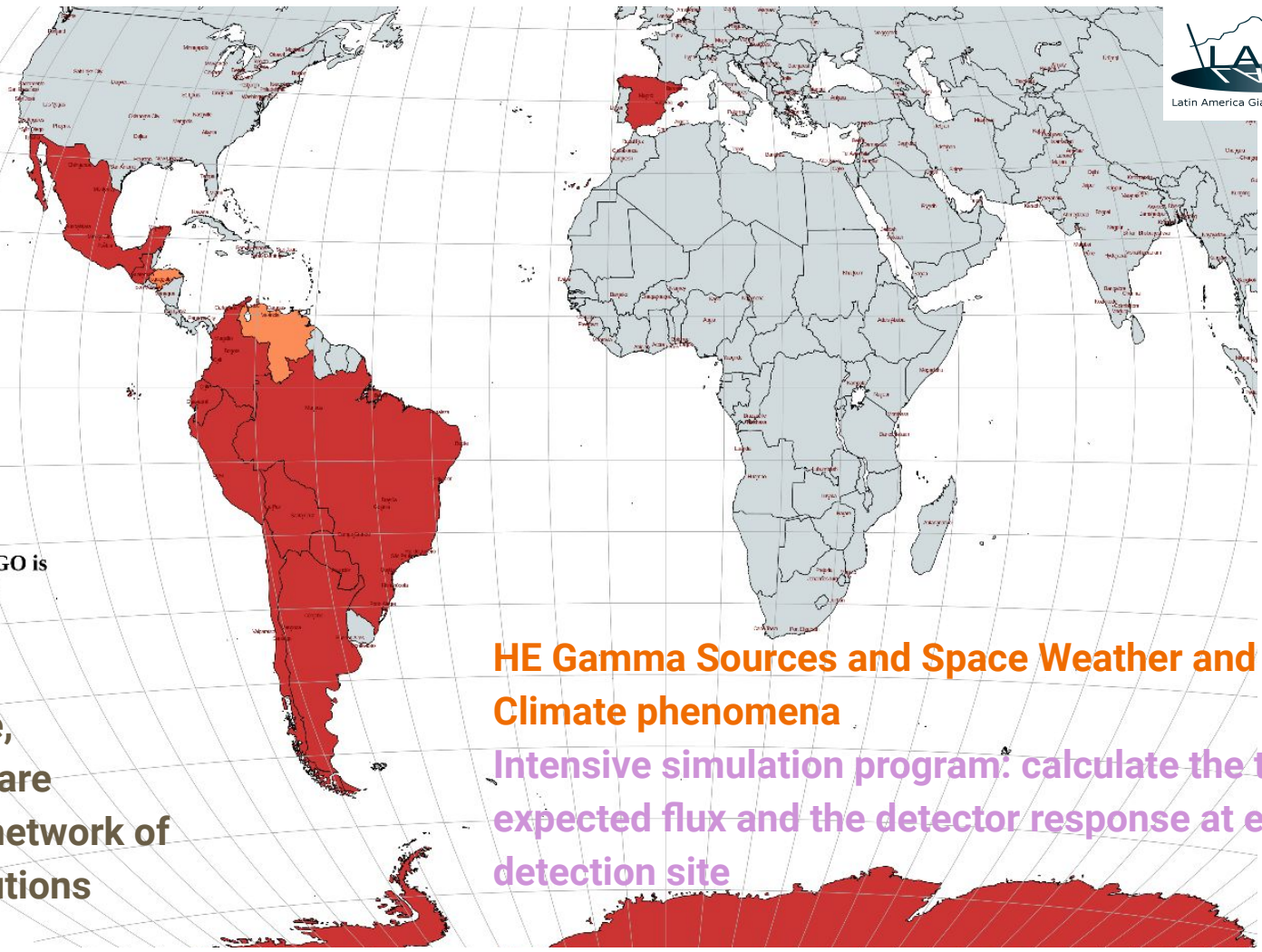
Countries where LAGO is operating

- Active
- Development

Hardware, software, expertise and data are shared across the network of participating institutions

HE Gamma Sources and Space Weather and Climate phenomena

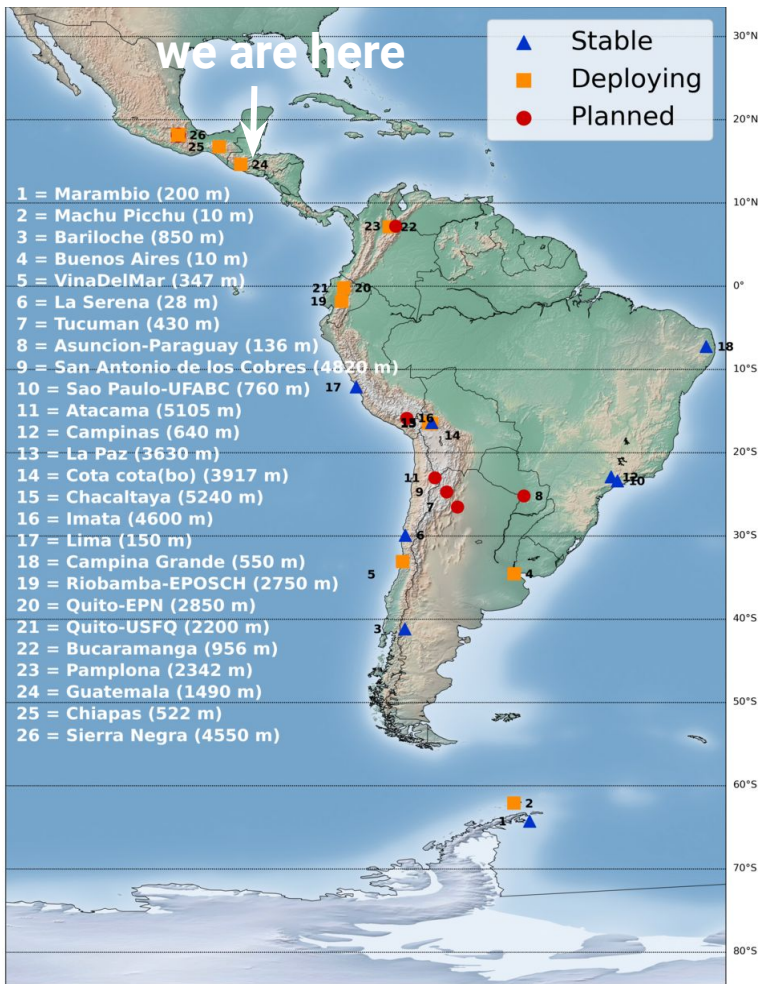
Intensive simulation program: calculate the total expected flux and the detector response at each detection site



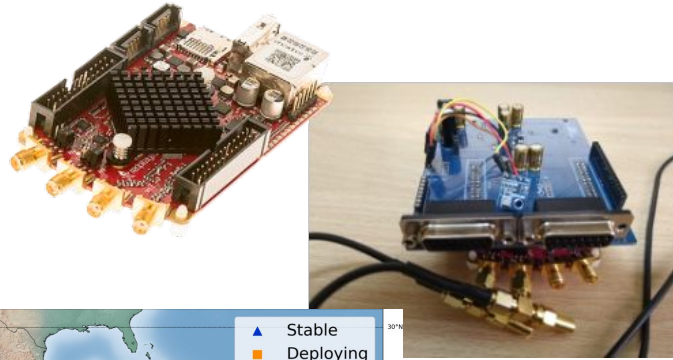
# The Latin American Giant Observatory

LAGO is an extended astroparticle observatory at continental scale: from Mexico to Antarctica

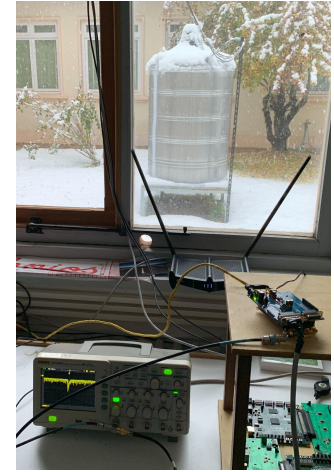
- **Scientific goals**
  - Astroparticle physics to study the extreme universe
  - Transient and long term space weather phenomena trough Solar modulation of Cosmic Rays
  - Measurement of background radiation at ground level
- **Academic goals**
  - We are a HEP and astro-ph seedbed for Latin American students
  - Build a Latin American network of Astroparticle and Cosmic Rays researchers



# The Latin American Giant Observatory

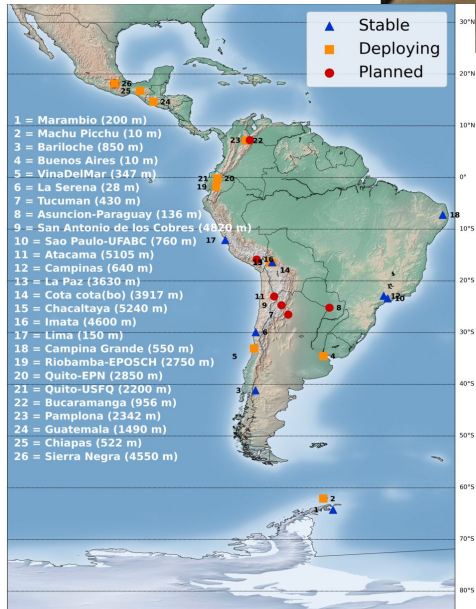


Own designed electronics based on RedPitaya SteamLab. Provides services for other experiments



LAGO WCDs have a single PMT and are based on SBC and COTS. They are autonomous, reliable, simple, smart, and inexpensive.

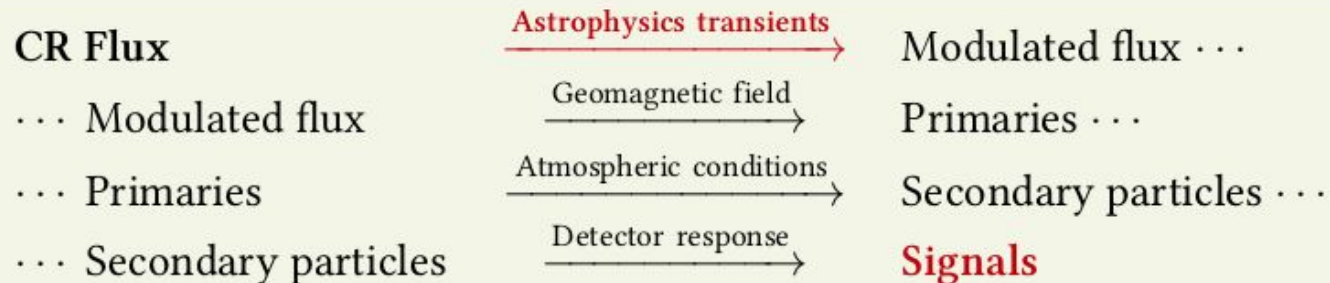
1-10 m<sup>3</sup> WCD deployed at different altitudes and geomagnetic coordinates



## LAGO Capabilities: Multi-spectral analysis

- Simultaneous measurements of secondaries at ground level
- Intensive simulation and data analysis frameworks

## Connections



## Synergy

Flux variation of signals at detector level  $\Leftrightarrow$  Transients

**Extensive Air Showers (EAS):**  
each cosmic ray interacts with  
the atmosphere producing a  
cascade of up to  $>10^{10}$   
secondary particles

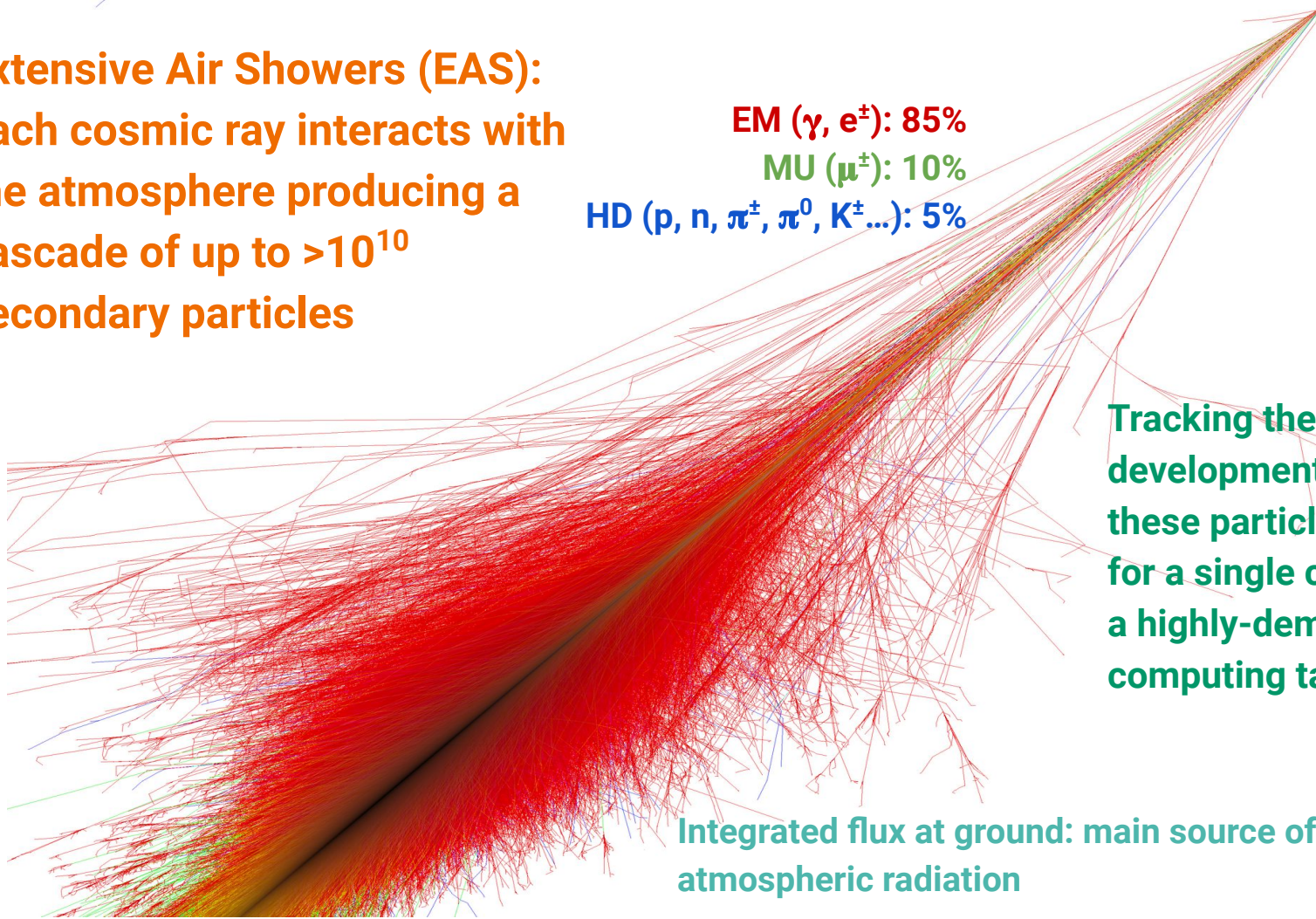
EM ( $\gamma$ ,  $e^\pm$ ): 85%

MU ( $\mu^\pm$ ): 10%

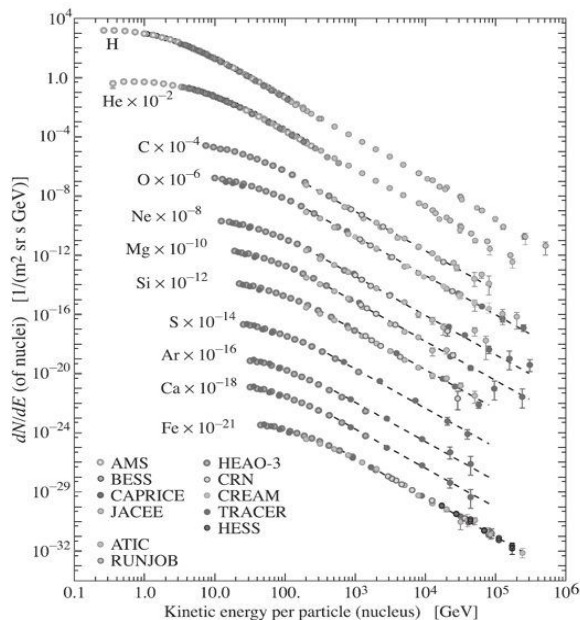
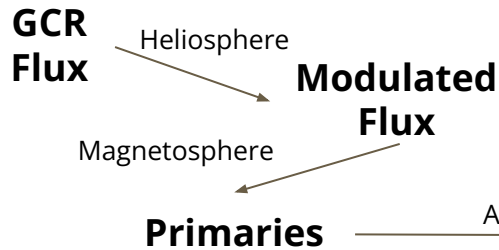
HD (p, n,  $\pi^\pm$ ,  $\pi^0$ ,  $K^\pm$ ...): 5%

Tracking the  
development of all of  
these particles, even  
for a single cascade, is  
a highly-demanding  
computing task

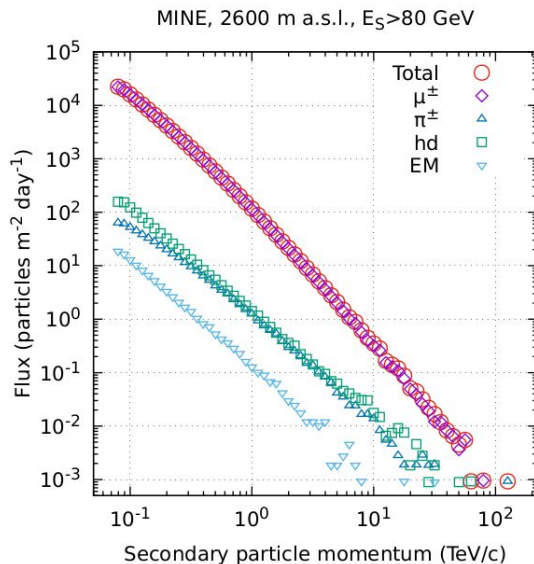
Integrated flux at ground: main source of  
atmospheric radiation



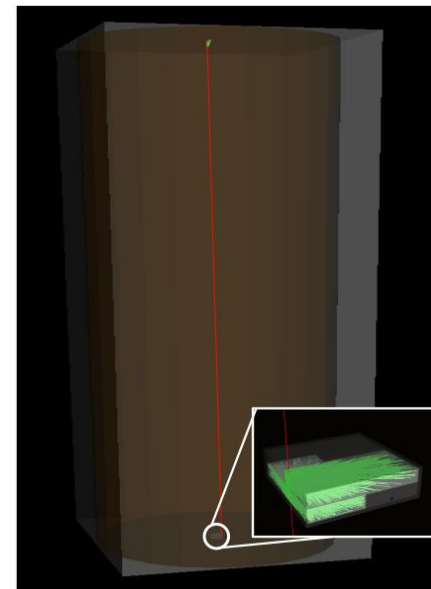
# ARTI: The LAGO Simulation framework



**ARTI + CORSIKA**

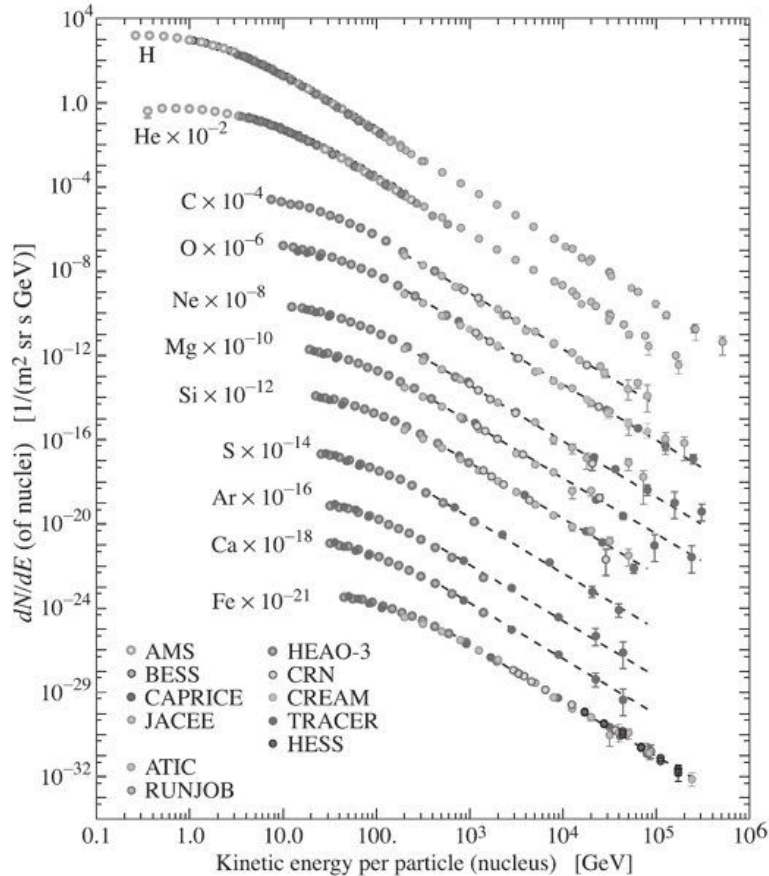


**ARTI**



**ARTI + MEIGA (Geant4)**

# Primary flux integration



For each primary, we need to integrate its spectrum to get the expected number of primaries at the top of the atmosphere

$$N_{t,S} = \int_t \int_S \int_{\Omega} \int_{E_p} j_0(E_p, Z_p)^{\alpha(E_p, Z_p)} dt dS d\Omega dE$$

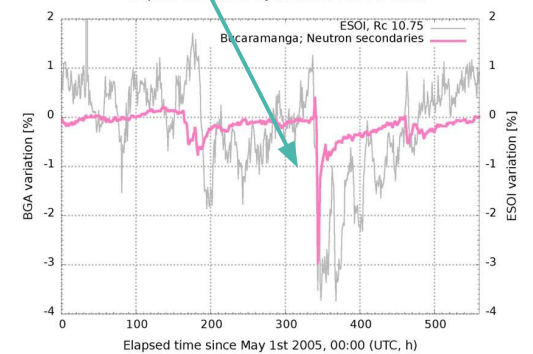
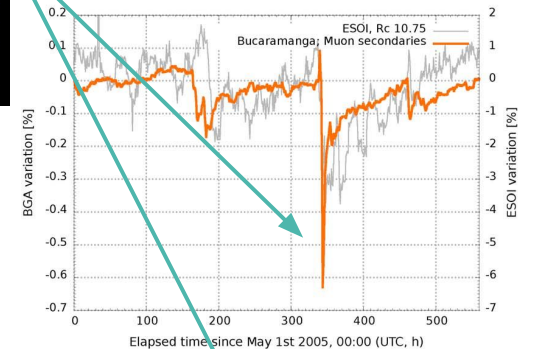
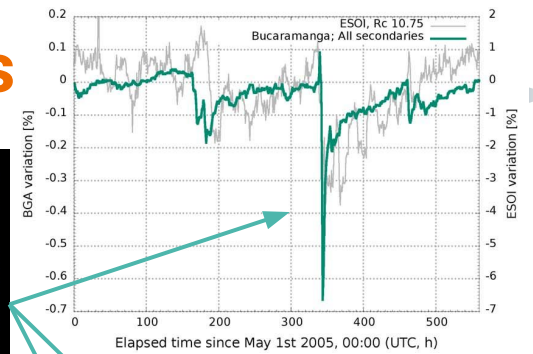
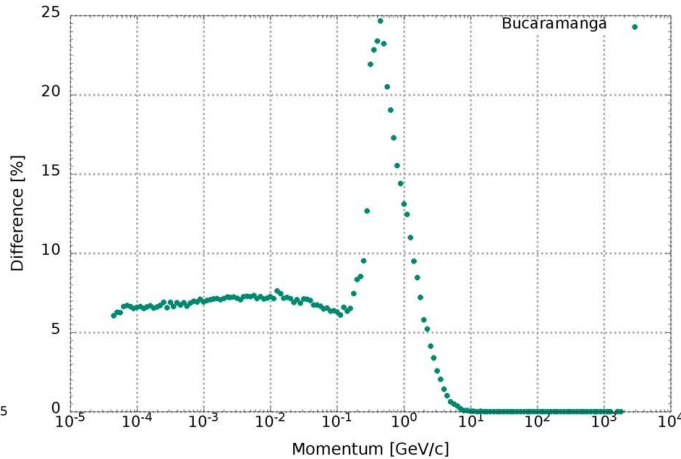
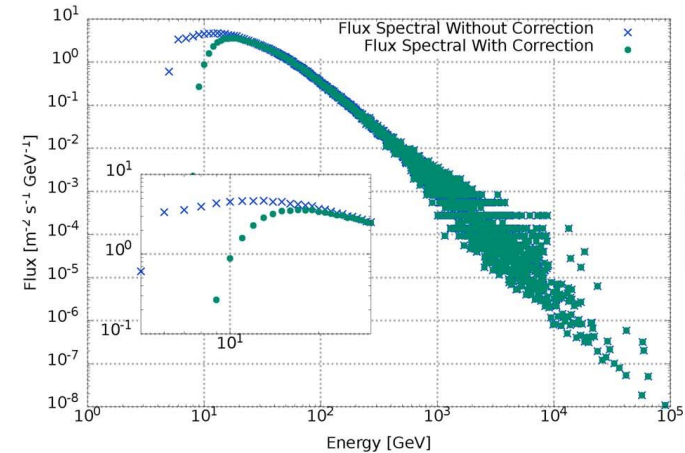
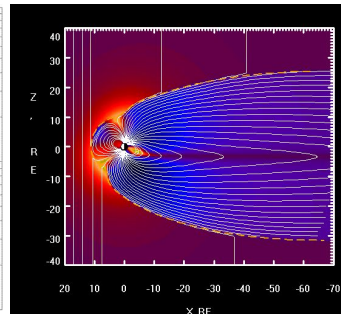
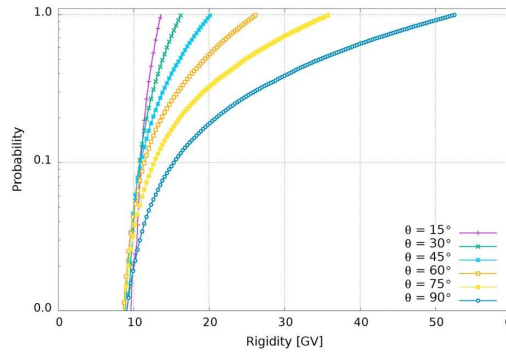
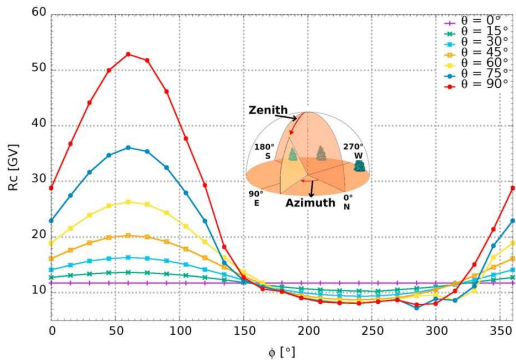
We integrate:

- full spectra,  $1 < Z < 26$
- hemisphere,  $0 \leq \theta \leq \pi/2, -\pi \leq \phi \leq \pi$
- energy range,  $(R_C \times Z_p) < E/GeV < E_{max}$

$R_C$  is the local, time-dependent, geomagnetic rigidity cut-off  
 $E_{max}$  depending on application

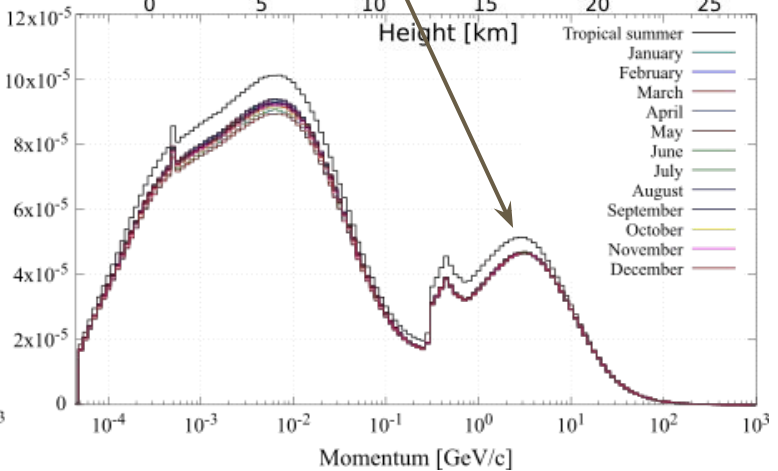
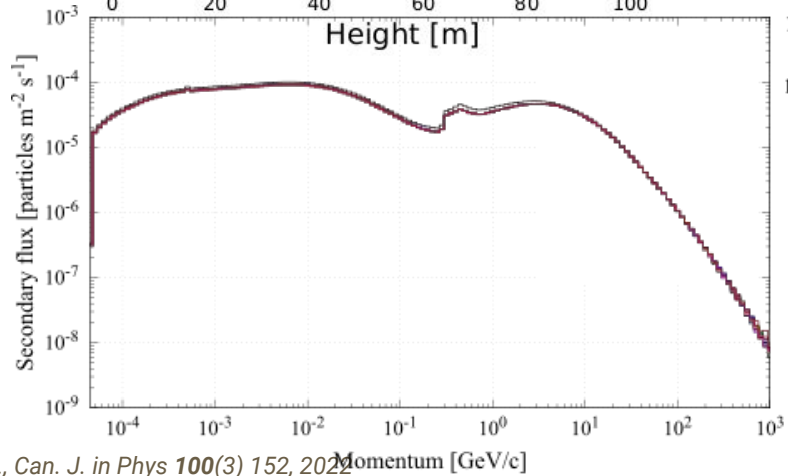
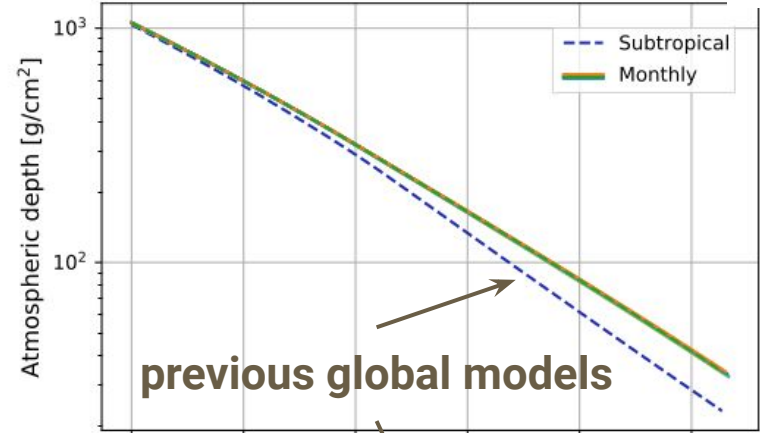
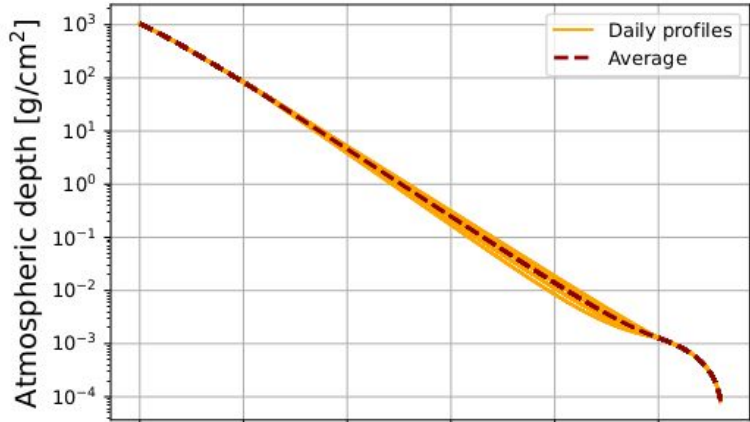


# Time-dependent local geomagnetic effects

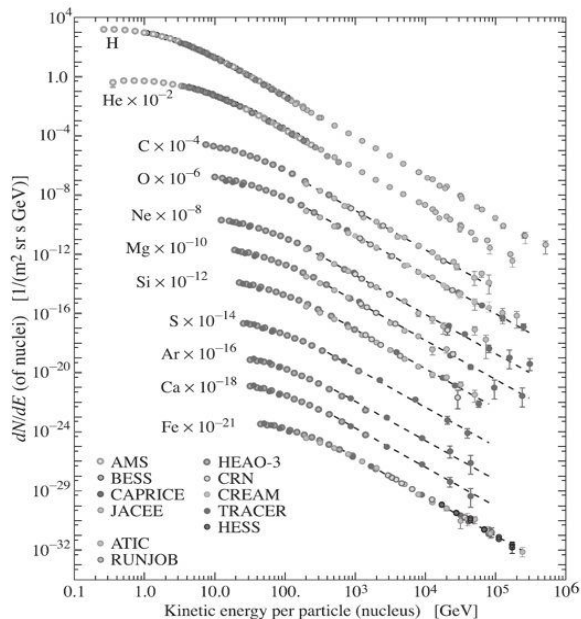
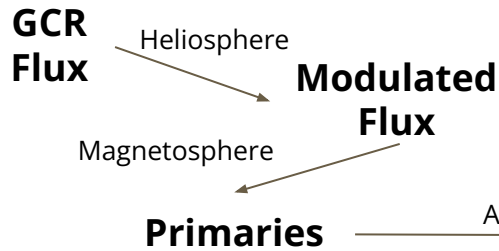


# Local atmospheric effects

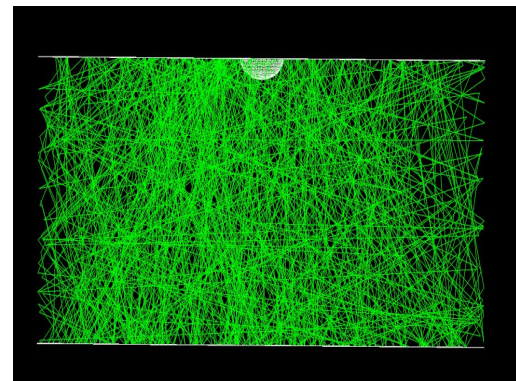
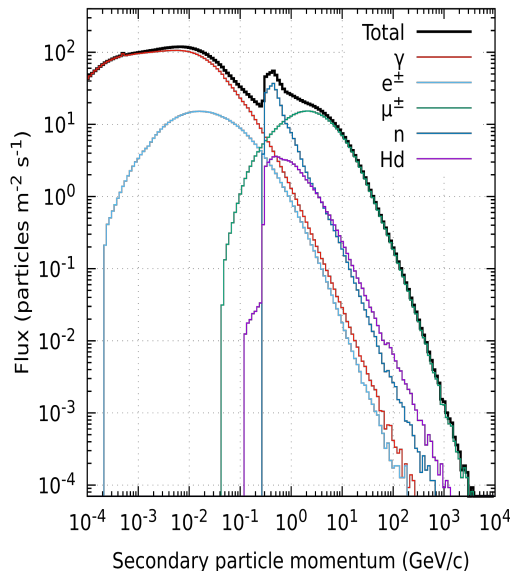
# Monthly-averaged or instantaneous local atm profiles from GDAS



# ARTI: The LAGO Simulation framework



IMA, 4 days, 4600 m a.s.l.



**ARTI (CORSIKA)**

**ARTI**

**ARTI+MEIGA (Geant4)**

# Challenge: profit from the effort of every collaborator

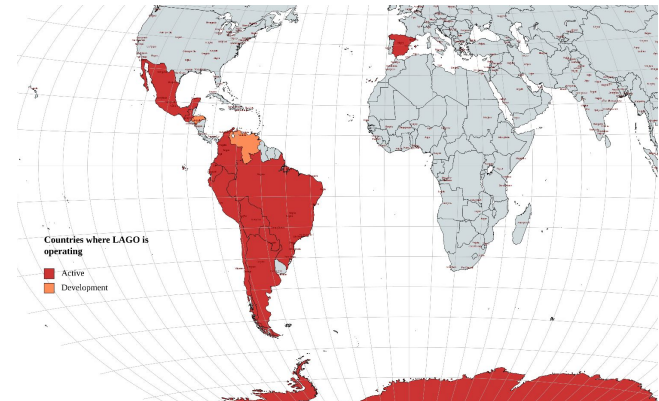
- Non-centralized alliance (30 institutions, 11 countries)
- >100 researchers grouped into autonomous research units
- Externals & students are join the project for short periods of time
- Computation is supported by the heterogeneous resources of each institution (private HPC facilities)



Every result (simulations, measurements, code or designs) is of interest to other members and external actors (industrial, government...)

Need of effective mechanisms for:

- curation and data publishing
- reproducibility
- coordination of the development
- sharing computational resources



# The FAIR paradigm

## Findability



discoverable,  
identifiable and  
locatable by  
means of a  
standard  
identification  
mechanism

PiD, Metadata

## Accessibility



always available  
and obtainable;  
even if the data  
is restricted, the  
metadata is  
open

Cloud storage

## Interoperability



syntactically  
parseable and  
semantically  
understandable,  
allowing data  
exchange and  
reuse

Vocabulary

## Reusability



described and  
shared with the  
least restrictive  
licences,  
allowing the  
widest reuse  
possible

Harvester,  
copyleft

# Our Approach: “FAIRification” of every result

**Data, software, materials** should be considered as **Linked Open Data**:

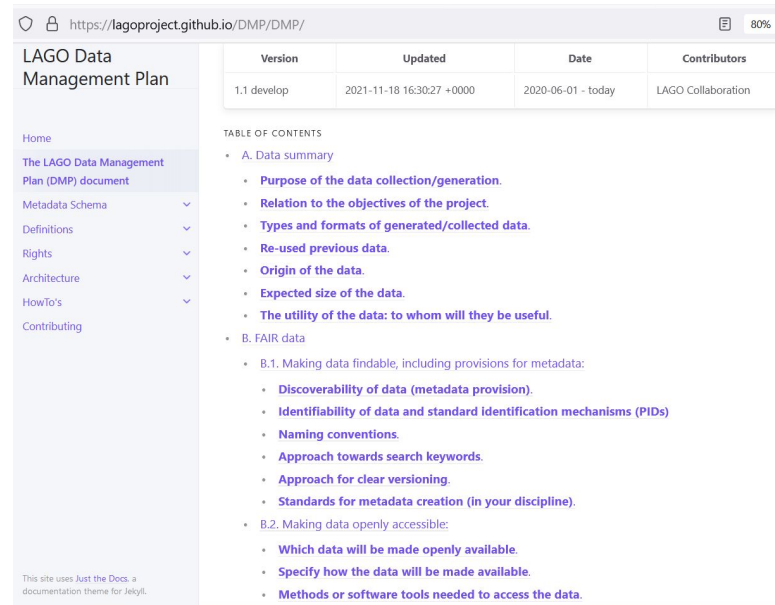
- Univocally referenced with **PiDs and/or IRIs**
- **Metadata** in current standard formats for linked-data
- **Provenance (IRIs in metadata)**
- Version control (CVS commits)
- Open policies and licences
- **Structured** and open-formatted (if possible)

All generated and available on **Open platforms**:

- **Computing, storage, and provisioning tools**
- **Virtual Organisation management**
- CD/CI, quality assurance tools

Whole described in an extended **DMP document**:

- not only a “Data” Management Plan
- comprises whole docs: how-to’s, team rules, **workflows**
- is a living document
- is another result (with PiD/IRI, version control, public)



https://lagoproject.github.io/DMP/DMP/ 80%

| Version     | Updated                   | Date               | Contributors       |
|-------------|---------------------------|--------------------|--------------------|
| 1.1 develop | 2021-11-18 16:30:27 +0000 | 2020-06-01 - today | LAGO Collaboration |

LAGO Data Management Plan

Home

The LAGO Data Management Plan (DMP) document

Metadata Schema

Definitions

Rights

Architecture

HowTo's

Contributing

TABLE OF CONTENTS

- A. Data summary
  - **Purpose of the data collection/generation.**
  - **Relation to the objectives of the project.**
  - **Types and formats of generated/collected data.**
  - **Re-used previous data.**
  - **Origin of the data.**
  - **Expected size of the data.**
  - **The utility of the data: to whom will they be useful.**
- B. FAIR data
  - B.1. Making data findable, including provisions for metadata:
    - **Discoverability of data (metadata provision).**
    - **Identifiability of data and standard identification mechanisms (PIDs)**
    - **Naming conventions.**
    - **Approach towards search keywords.**
    - **Approach for clear versioning.**
    - **Standards for metadata creation (in your discipline).**
  - B.2. Making data openly accessible:
    - **Which data will be made openly available.**
    - **Specify how the data will be made available.**
    - **Methods or software tools needed to access the data.**

This site uses Just the Docs, a documentation theme for Jekyll.



**Expanding the capacity and capabilities of European Open Science Cloud by leveraging the experience, effort and resources of national publicly-funded digital infrastructures**



**EUROPEAN OPEN SCIENCE CLOUD**



INDIGO - DataCloud  
Better Software for Better Science

**EOSC-Synergy Horizon 2020 RI project 857647**

Objectives aligned to Open Science

+

Coordination with other EU Projects

=

Open Platforms & Tools for:

- Computing & Storage & Publishing
- Community Management (IdP...)
- CI/CD & quality assurance
- Expertise

# Standardizing: data-sets & encapsulates pipeline steps

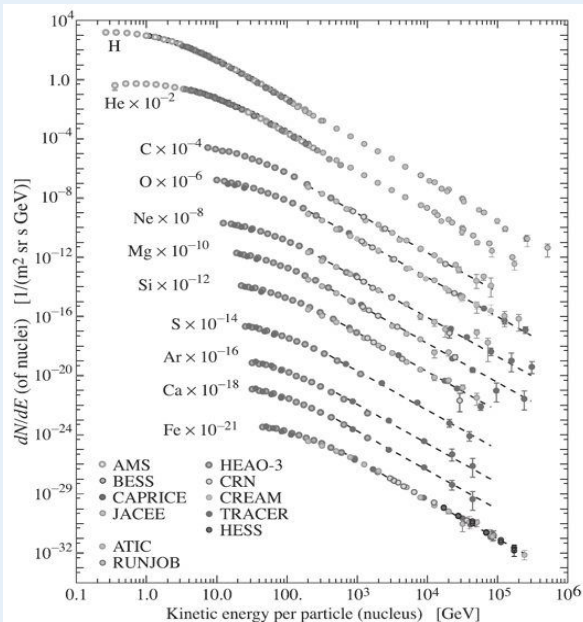
GCR Flux → Heliosphere → Modulated Flux  
 Modulated Flux → Magnetosphere → Primaries (S0)

Atmosphere →

Secondaries (S1)

Detector →

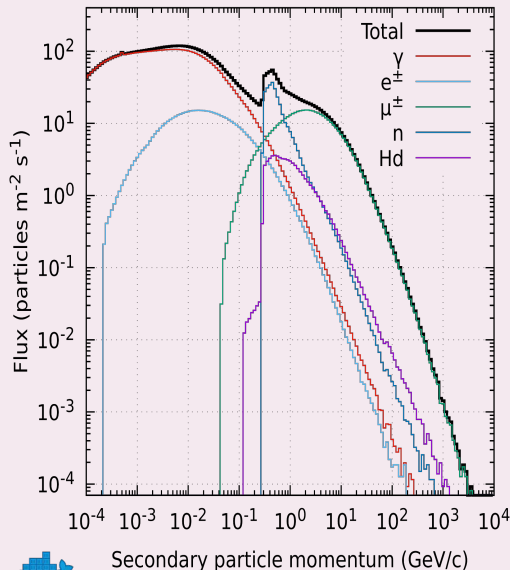
Signals or Doses (S2)



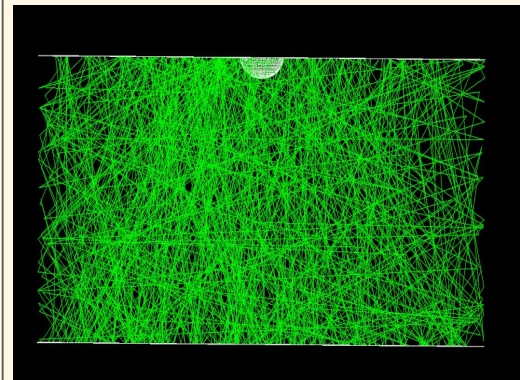
ondataSim-S0



ondataSim-S1

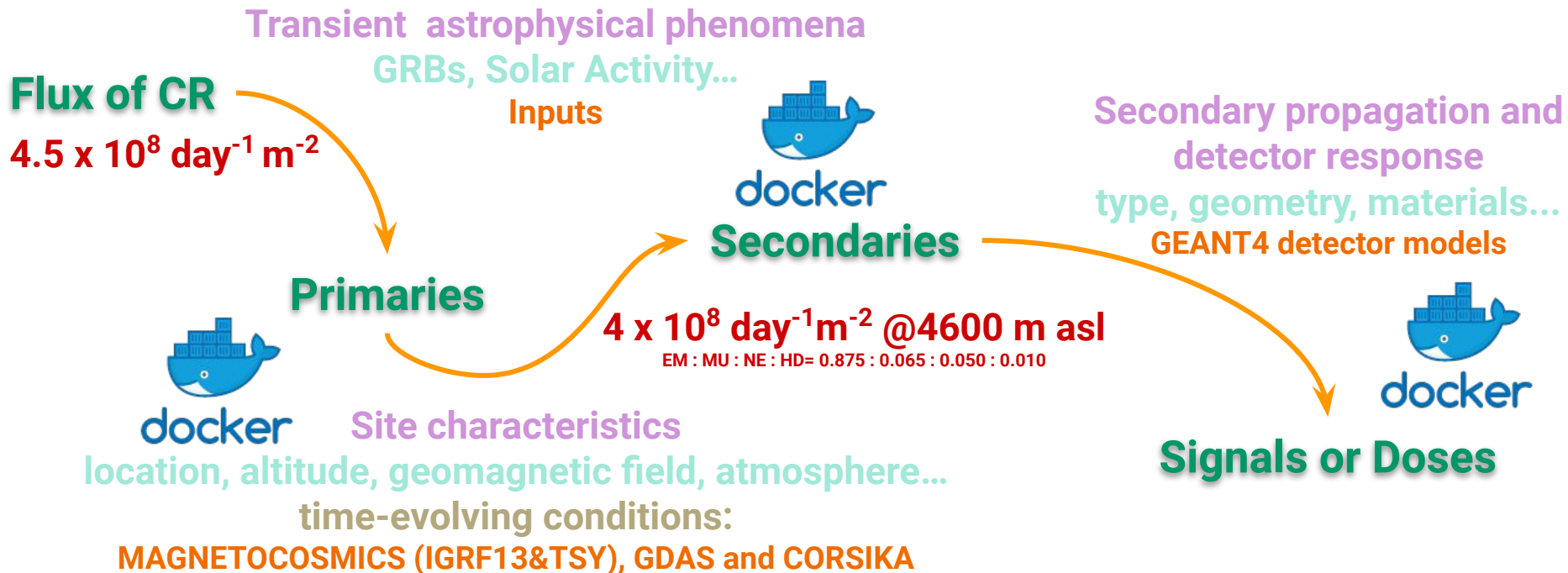


ondataSim-S2





# Standardizing: workflow and some figures



# Standardizing: workflow and some figures

Transient astrophysical phenomena

GRBs, Solar Activity, ...

Flux of CR

4.5 x 10<sup>17</sup>

Too many parameters and files: must ensure the provenance and reproducibility defining the data and metadata structures

location, altitude, geomagnetic field, atmosphere...

time-evolving conditions:

MAGNETOCOSMICS (IGRF13&TSY), GDAS and CORSIKA

Signals of Doses



# How does it work?

HPC provider assigns cloud resources to EOSC:  
 $n$  Nodes,  $r$  GB/TB of RAM,  $d$  TB local storage



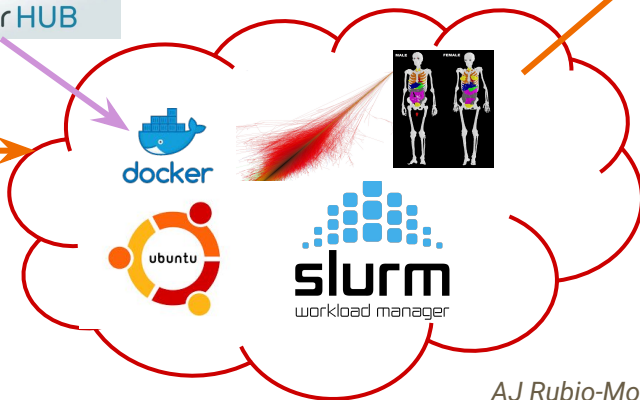
**Infrastructure Manager**

web-based service with templates for distributing to available resources in virtual clusters. Installs OS, in our case, ubuntu 20.04 + slurm manager

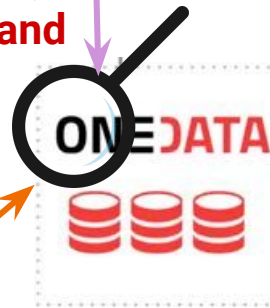


docker containers with our codes are deployed from docker HUB in the virtual cluster

**virtual cluster**



**FAIR: catalogs are findable, accessible, interoperable, and reusable**



Results are stored at cloud-based storage services. Access through personal tokens



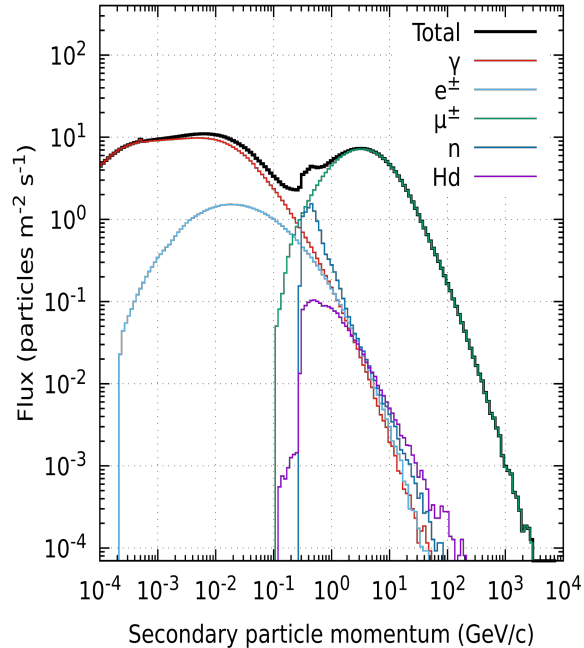
PID (Persistent Identifiers) are assigned for each data catalog



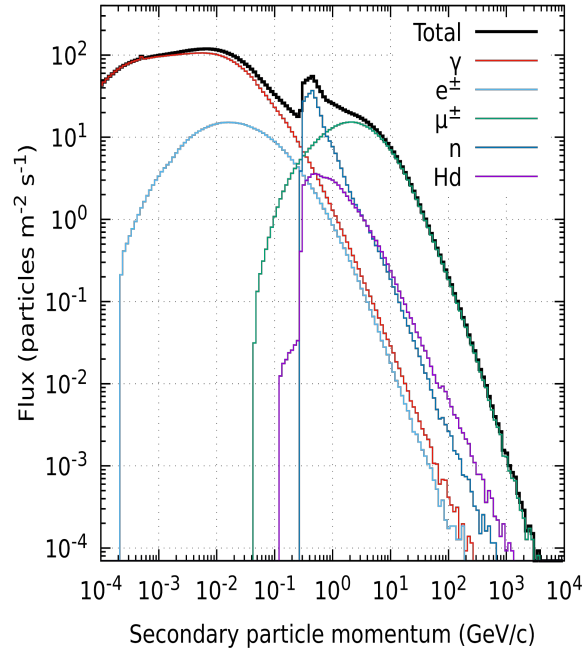
# Expected secondary flux

Detailed flux of secondary particles at any location around the World.

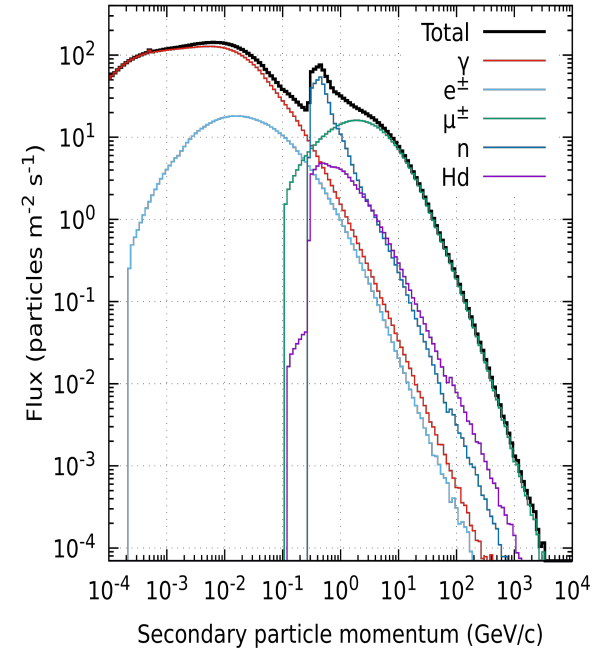
LSC, 3 days, 28 m a.s.l.



IMA, 4 days, 4600 m a.s.l.

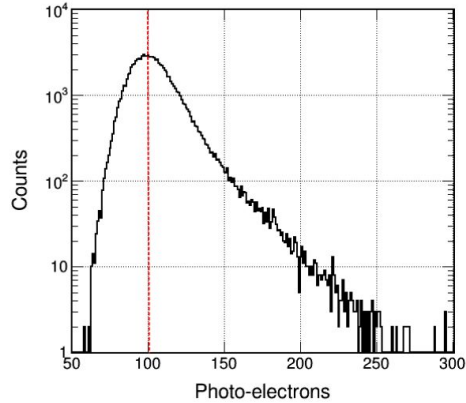
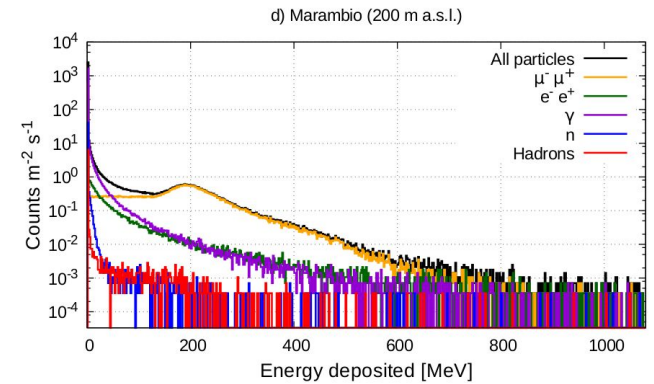
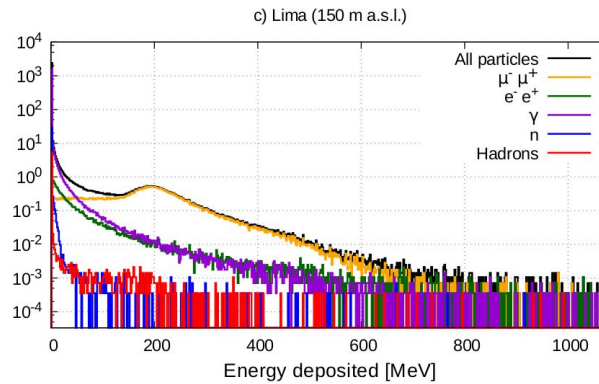
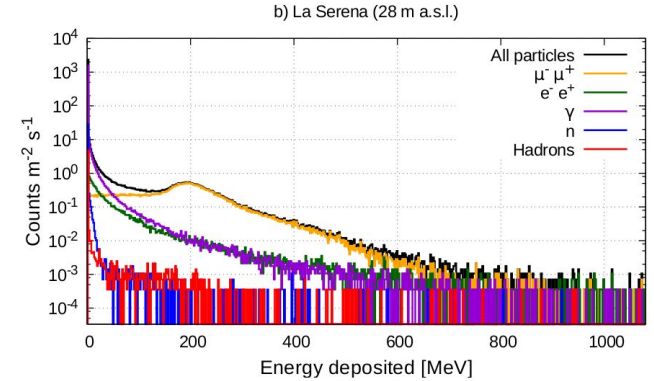
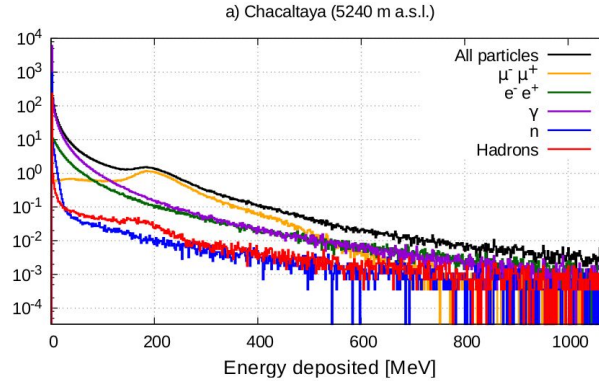
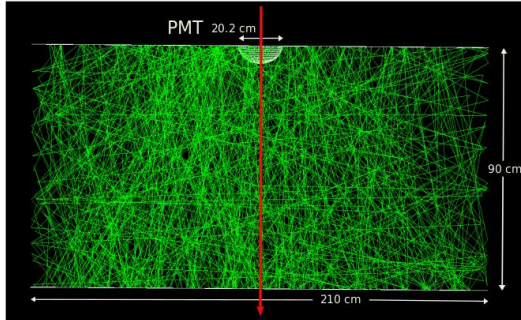


SAC, 7 days, 4500 m a.s.l.



Secondary flux in 4 LAGO sites: La Serena (CL), Imata (PE) and San Antonio de los Cobres (AR)

# Signals simulations: deposited charge histograms



Vertical Equivalent Muon (VEM)

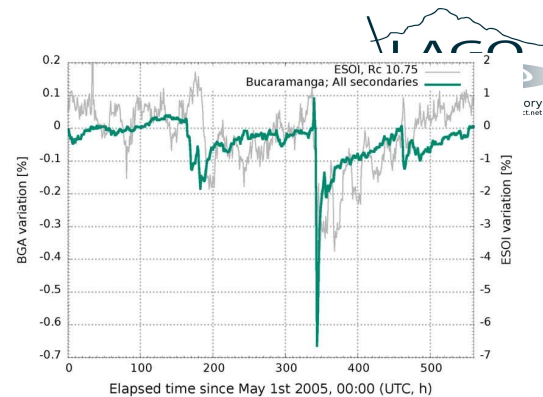
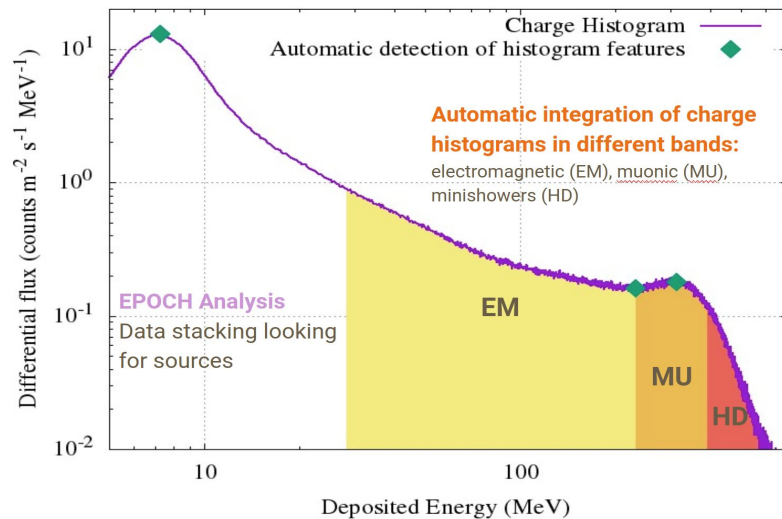
Expected histograms at Chacaltaya (BO), La Serena (CL), Lima (PE) and Marambio (ANT)

# LAGO Space Weather

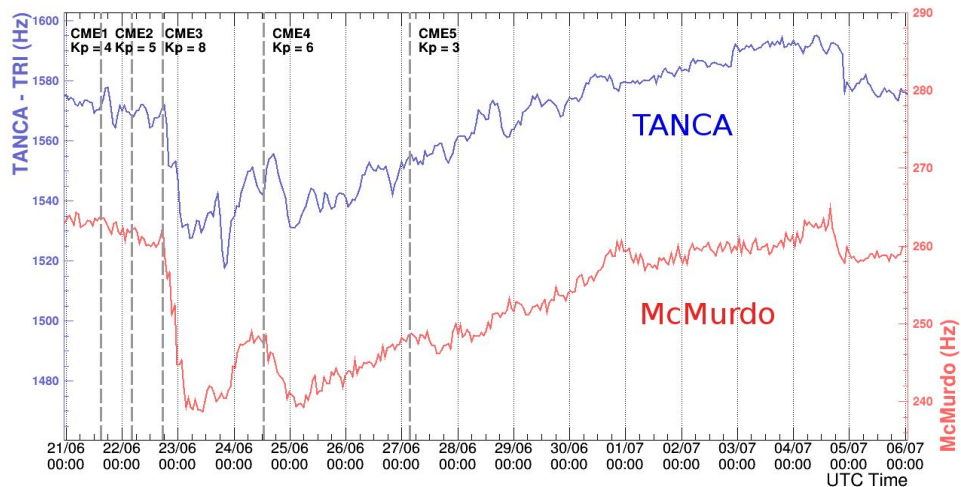
LAGO studies Earth-Sun connection by measuring the time-evolving secondary signals from ground level.

Atmospheric conditions are continuously monitored

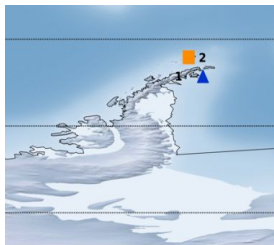
## Multispectral analysis of the charge histogram at the WCD



## Forbush Decrease Observation at Campina (BR)



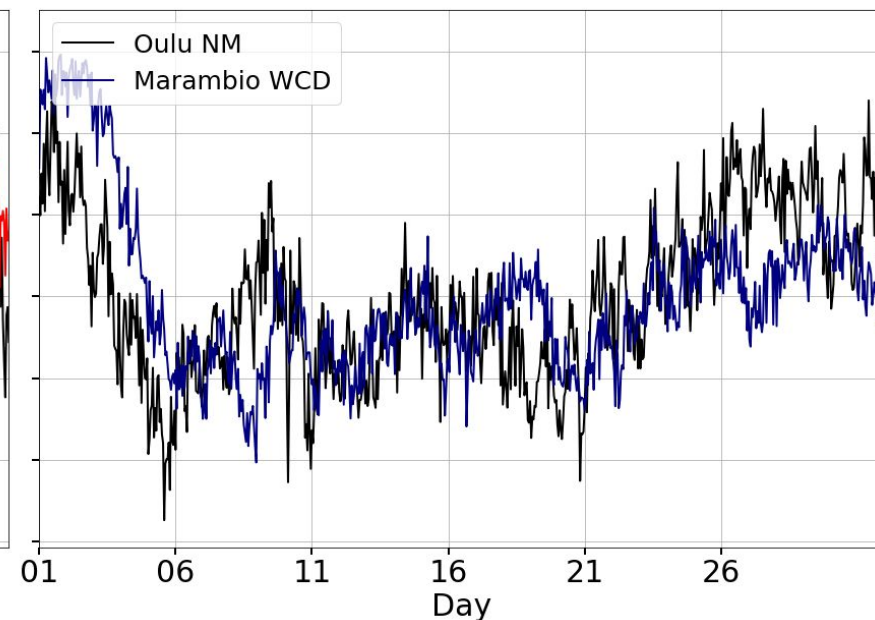
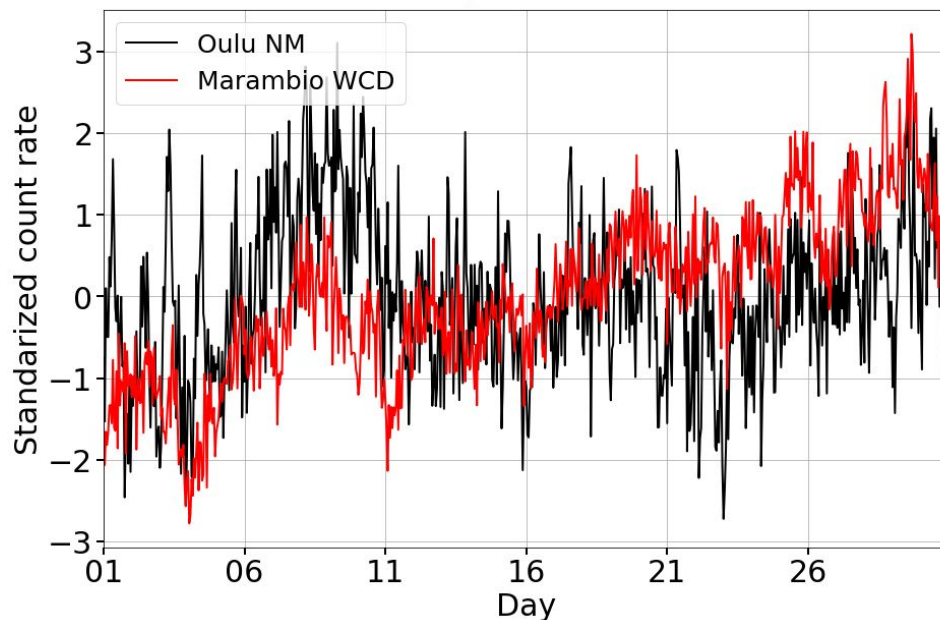
# LAGO Space Weather



Long term analysis for space climate at low rigidity sites: LAGO observations at the Antarctica Peninsula

April 2019

December 2020



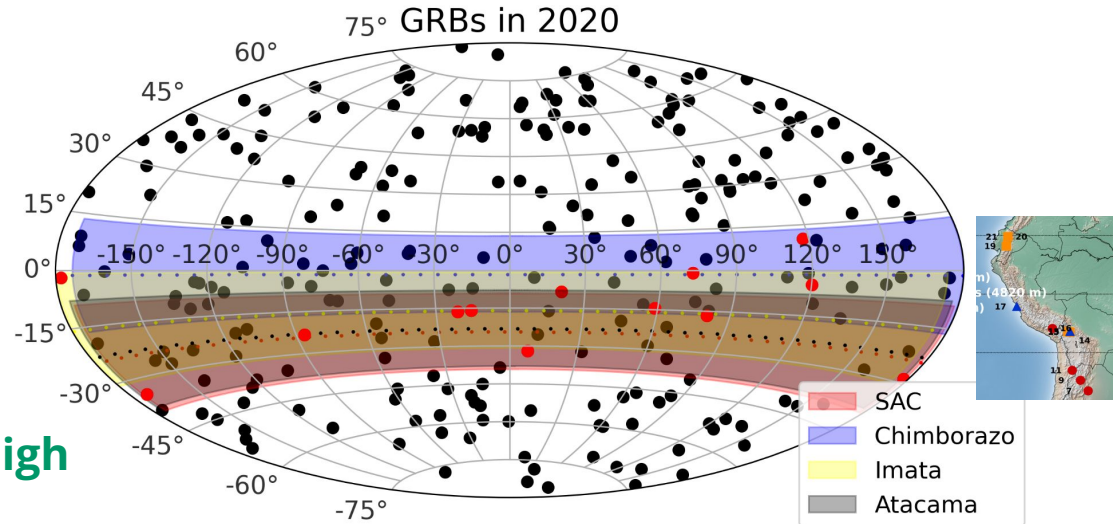
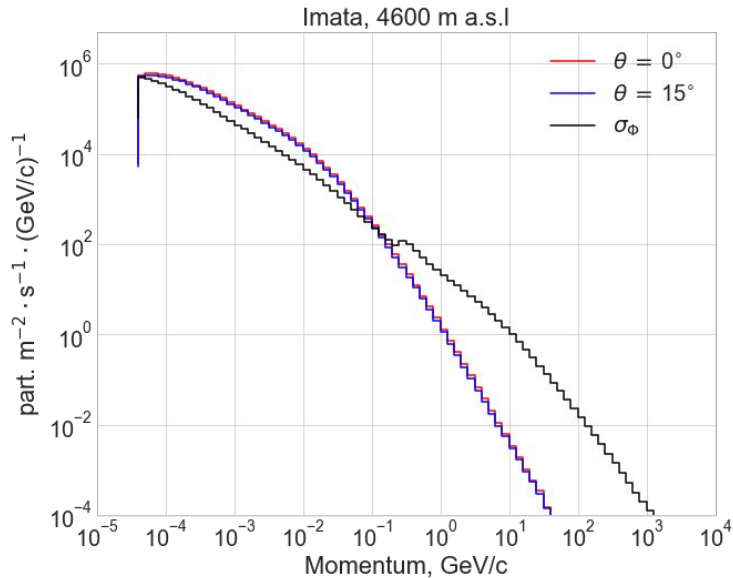


# LAGO HE program: GRBs

Small arrays of WCD at very high altitude sites (<4500 m asl)

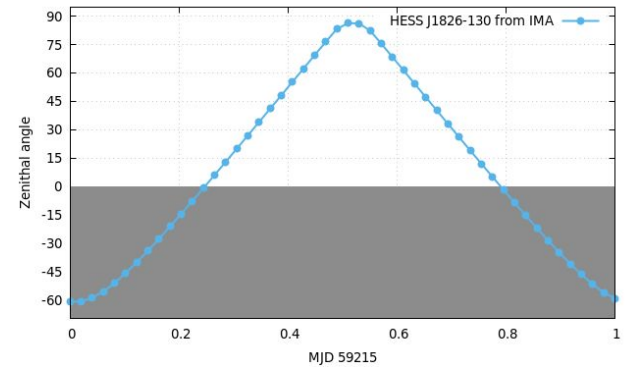
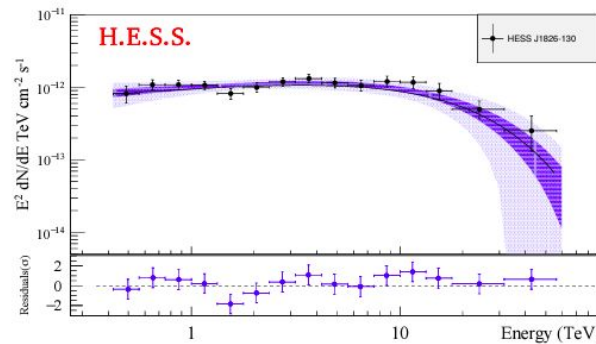
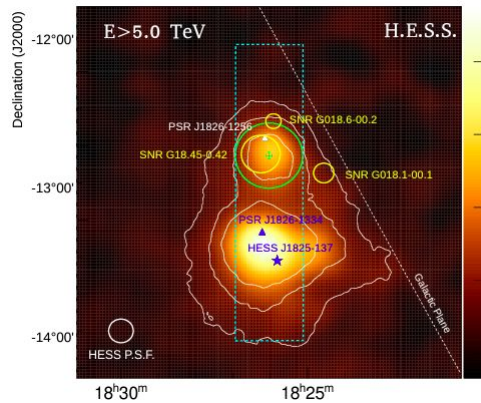
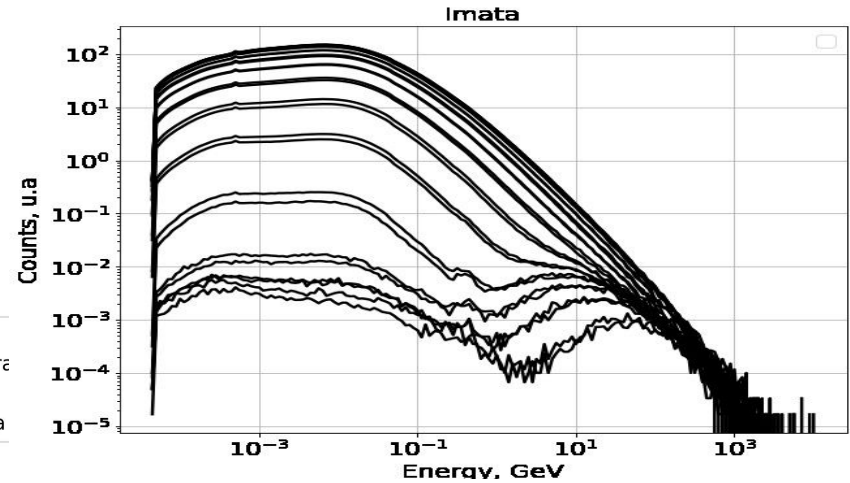
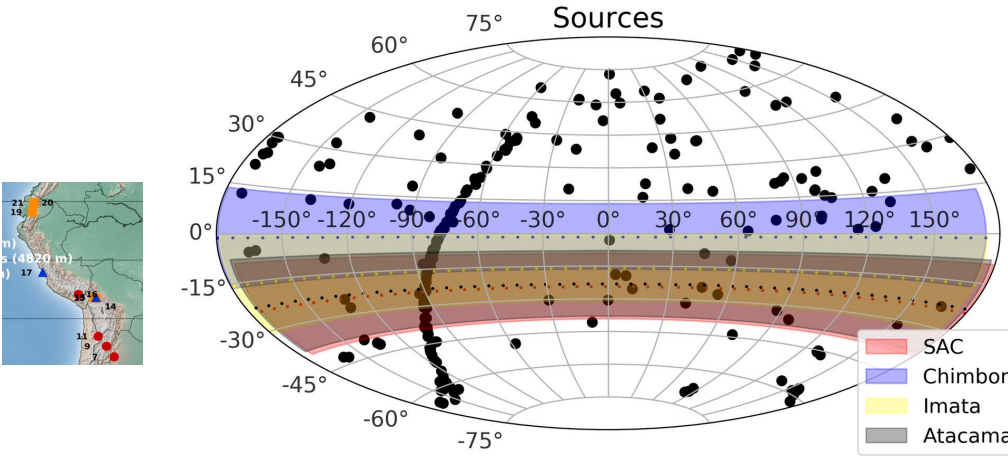
FOV overlapped by design for simultaneous measurements

muon decay for signal-to-noise ratio improvement

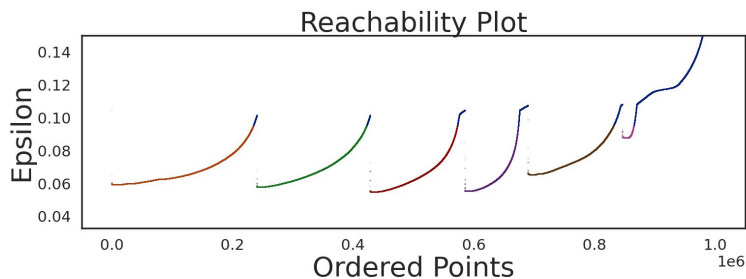
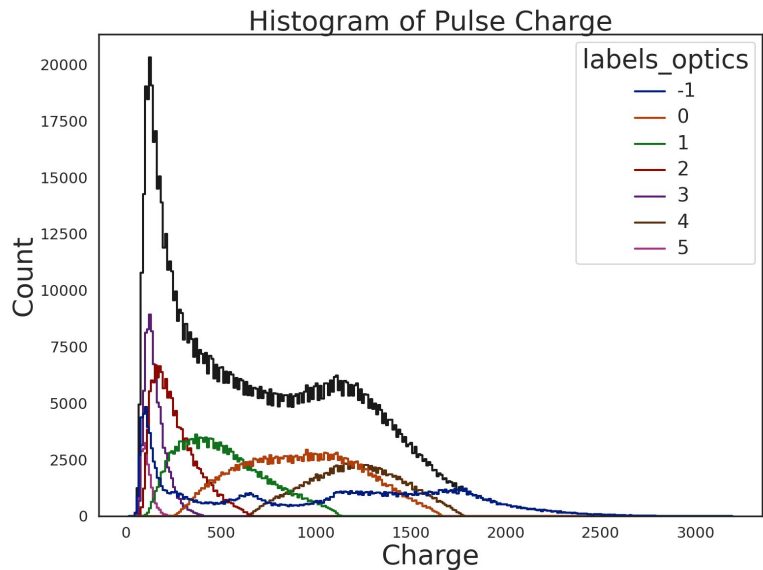


LAGO Capabilities for detecting high energy component of GRBs and Gamma Sources

# LAGO HE program: Steady $\gamma$ sources: HESS J1826-130



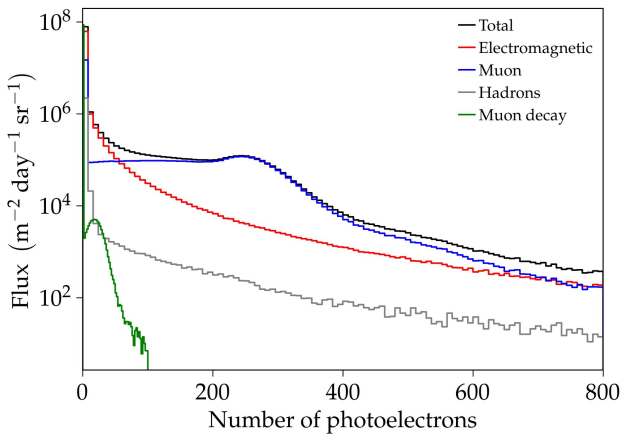
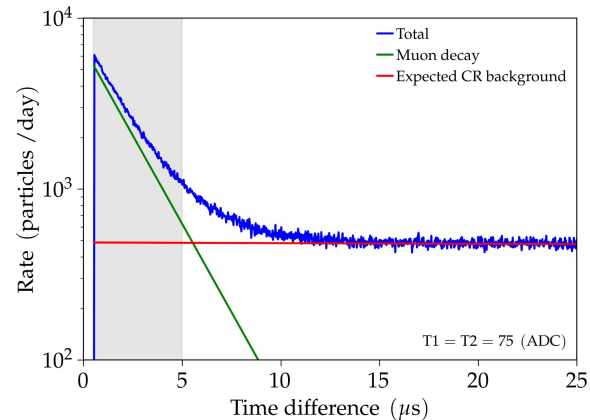
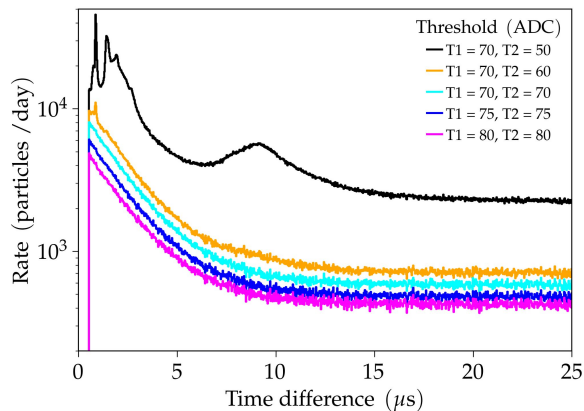
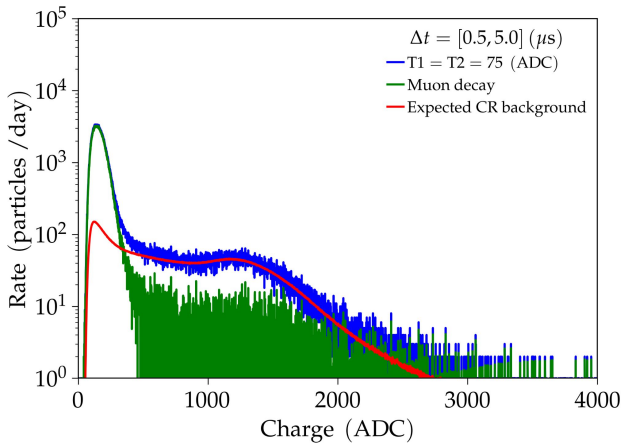
# Particle classification using ML algorithms



## What kind of particles produce the observed features in the charge histogram?

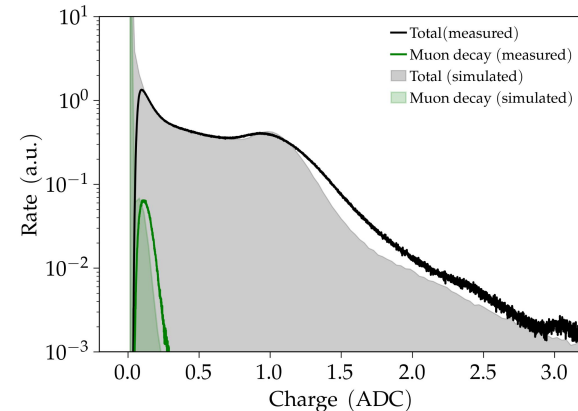
- OPTICS clustering algorithms
- Expected features observe
  - EM (photons and electrons), muon and multi-particle contributions are clearly observed
- **Multi-spectral analysis allowed**
  - Gamma/hadron separation
  - Space Weather spectroscopy
  - Extensive Air Showers analysis: LAGO sensitivity up to the knee

# Measurement of the Muon Lifetime & Michel Spectrum



## Muon decay measurements

- Low-energy muons decay within the water volume
- Observed in all our WCD
- **Large improvement in the signal-to-noise ratio and energy calibration**
- **An educational tool for associated Universities**

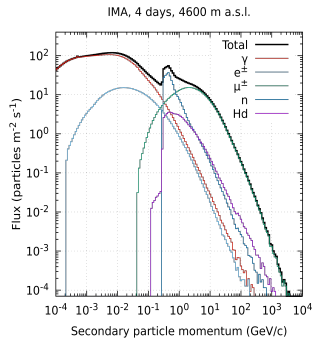


# Conclusions: ARTI/OneDataSim

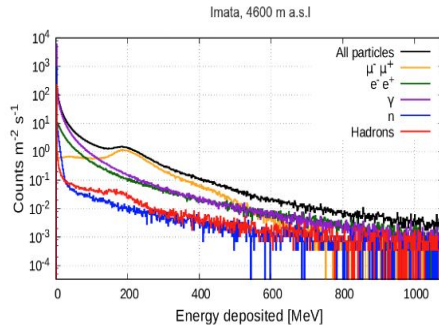
>10<sup>13</sup> simulated EAS in 1.6



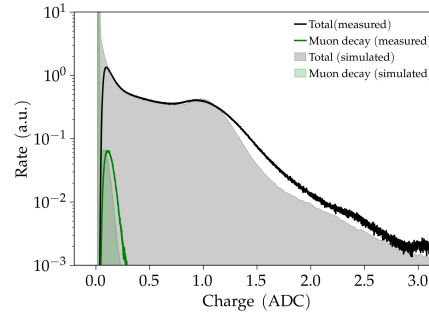
Functional, fully-integrated, and publicly-available cloud-based FAIR-compliant implementation



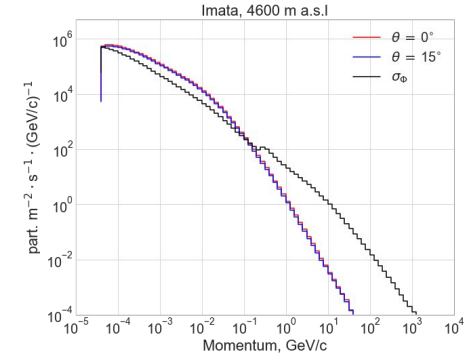
Secondaries at ground



WCD charges histograms



Muon decay and Michel spectrum



Site characterization for Gama sources

Expected flux of secondaries and the corresponding signals at any particle detector installed in any place in the World and under realistic atmospheric and geomagnetic conditions.

ARTI/OneDataSim in astroparticle physics: sites characterization, new detector designs, signal-to-background ratios...

ARTI/OneDataSim in other fields: volcano muography, mining prospecting, underground labs, errors at HPC, doses in flights, personal and safeguard protection, space exploration... (see R. Mayo-García, this conference)

ARTI/OneDataSim is publicly available: LAGO [GitHub repository](#) and in the [EOSC Marketplace](#)

[hernan.asorey@iteda.cnea.gov.ar](mailto:hernan.asorey@iteda.cnea.gov.ar)

**Thank you for your attention**





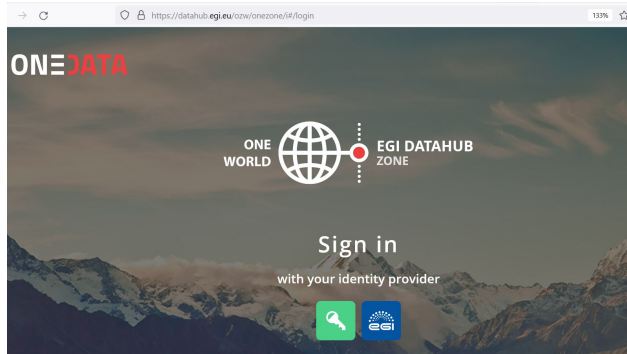
# Where are the meta & data stored?

**ONE DATA** is an object-based distributed FS focused on cloud and HPC:

- offers unified data access and roles across globally distributed environments;
- manages metadata and expose them to harvesters.

EGI/EOSC supports the “DataHub” instance that:

- allows extending the capacity by adding providers around the World;
- asks for (Handle.net) PiDs;
- enable pan-European IdPs and translate Virtual Organisation roles.



<https://datahub.egi.eu>

