

Search for dark photons in heavy-ion collisions

Elena Bratkovskaya

(GSI, Darmstadt & Uni. Frankfurt)



Kazimir Malevich, "Black Square" (1915)

10th International Workshop on Astronomy
and Relativistic Astrophysics
IWARA-2022
5-9 September 2022, Antigua, Guatemala

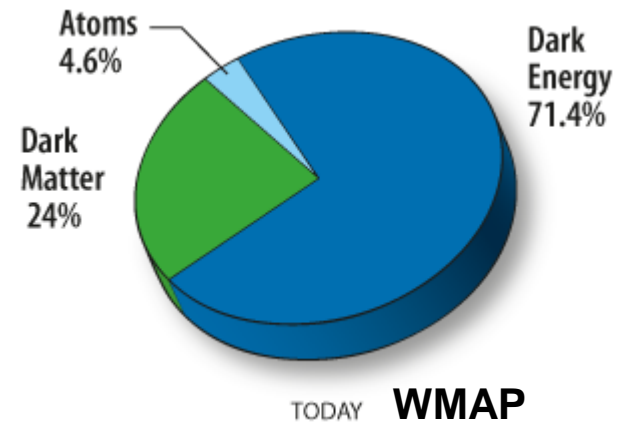
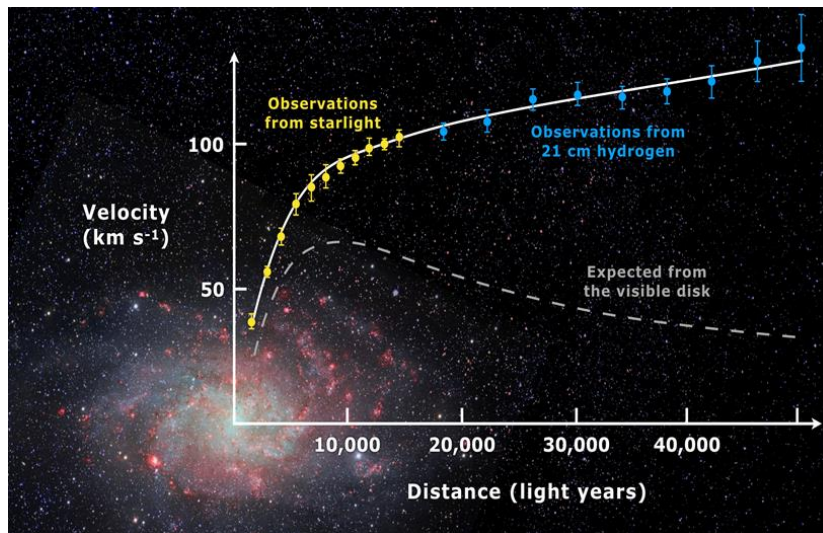


Structure of Universe

- ❑ Dark matter (DM) ~24%
- ❑ DM detected by astrophysical observations based on **gravitational** effects:

1933: F. Zwicky: observation of galaxy clusters

1970: V. Rubin: rotation anomalies in galaxies



Hubble: gravitational lensing



Dark matter portals

Search for **non-gravitational** dark matter (DM) **interactions with normal matter**, i.e. with standard model (SM) particles

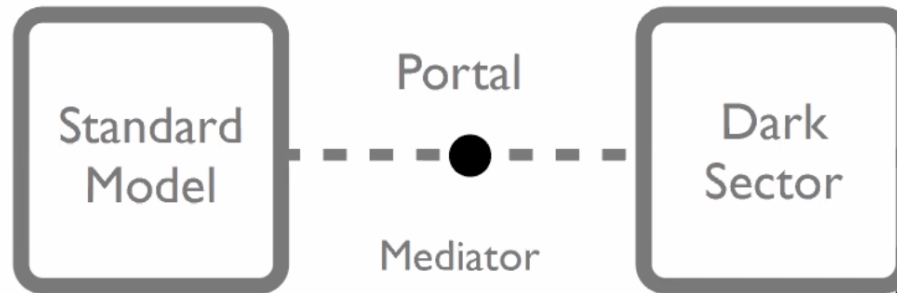
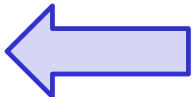


Figure from Brian Battel

$$\mathcal{L} \supset \begin{cases} -\frac{\epsilon}{2 \cos \theta_W} B_{\mu\nu} F'^{\mu\nu}, & \text{vector portal} \\ (\mu\phi + \lambda\phi^2) H^\dagger H, & \text{Higgs portal} \\ y_n L H N, & \text{neutrino portal} \\ \frac{a}{f_a} F_{\mu\nu} \tilde{F}^{\mu\nu}, & \text{axion portal.} \end{cases}$$


Dark matter detection

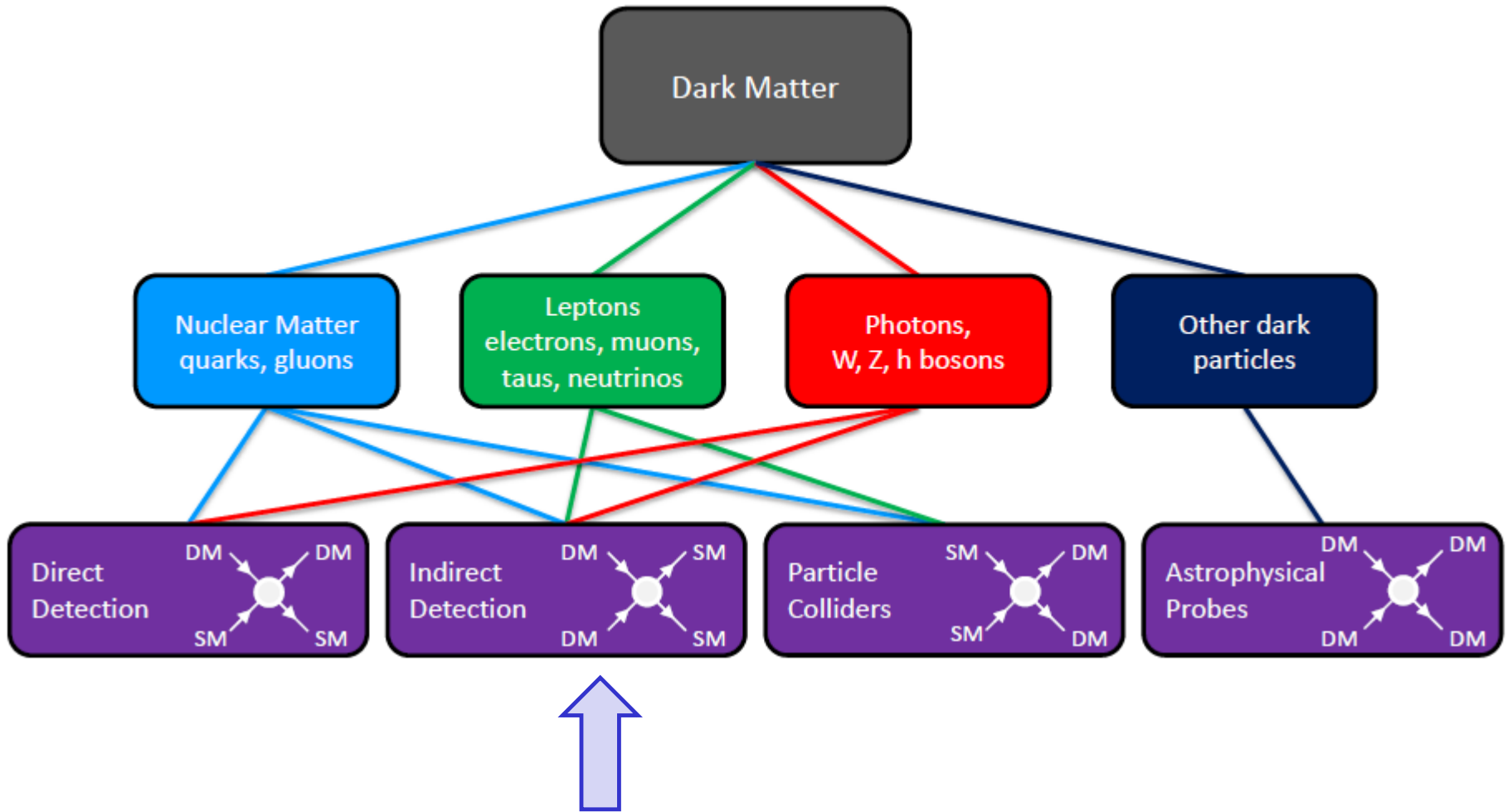


Figure from D. Bauer et al. 2015, Physics of the Dark Universe, 7, 16

Vector portal

The '**vector portal**' assumes the mixing of SM and DM via a **U(1)-U(1)'** gauge symmetry group mixing

L.B. Okun, Sov. Phys. 56 JETP (1982);
B. Holdom, Phys. Lett. B 166, 196 (1986)

$$\mathcal{L}_{A'} = -\frac{1}{4}F'^{\mu\nu}F'_{\mu\nu} + \boxed{\frac{1}{2} \frac{\epsilon}{\cos \theta_W} B^{\mu\nu} F'_{\mu\nu}} - \frac{1}{2}m_{A'}^2 A'^{\mu} A'_{\mu}$$

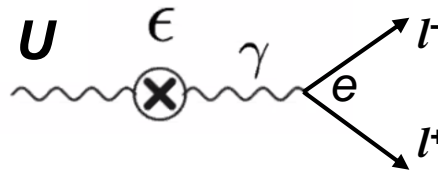
Dark photon field strength:

$$F'_{\mu\nu} \equiv \partial_{\mu}A'_{\nu} - \partial_{\nu}A'_{\mu}$$

SM hypercharge field strength:

$$B_{\mu\nu} \equiv \partial_{\mu}B_{\nu} - \partial_{\nu}B_{\mu}$$

ϵ - kinetic mixing parameter:



$$\epsilon^2 = \alpha'/\alpha$$

Due to the kinetic mixing the dark photon (U-boson) couples to the electromagnetic current with strength **ϵe**

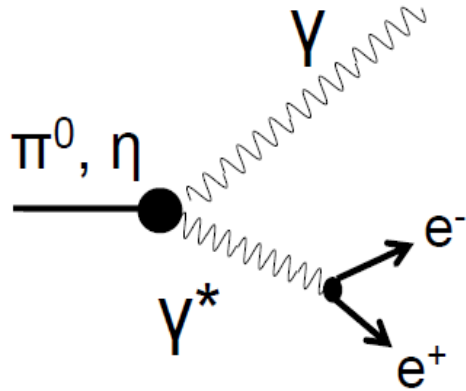
Unknown: kinetic mixing parameter **ϵ** and mass **M_U**

* Notation in literature for the 'dark photon': A' , V , U -boson

Dalitz decay of the dark photon to dileptons

- Dalitz decays of pseudoscalar mesons π^0, η and Δ -resonances to dileptons via the U-boson mediator

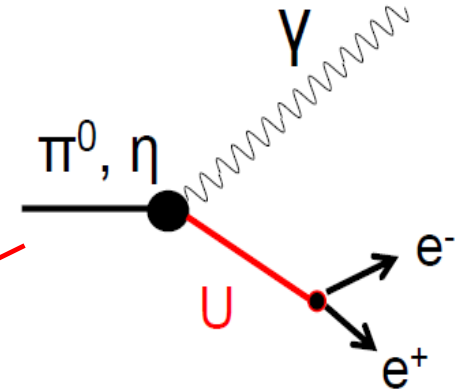
Standard model



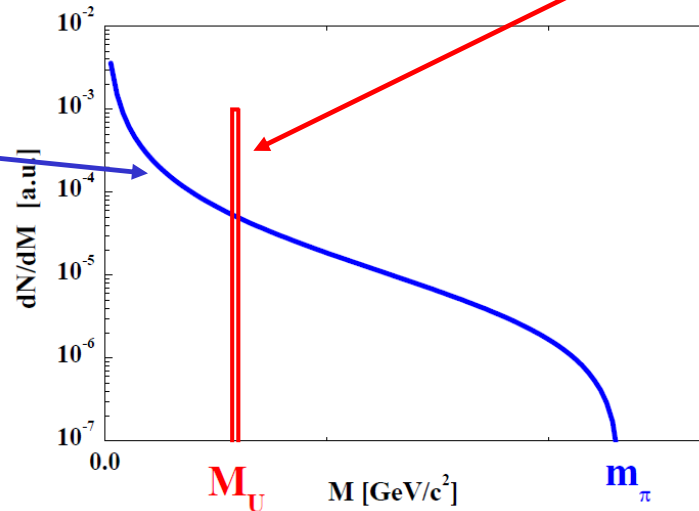
$$\begin{aligned} \pi^0 &\rightarrow \gamma + \gamma^*, \\ \eta &\rightarrow \gamma + \gamma^*, \quad \gamma^* \rightarrow e^+ e^-, \\ \Delta &\rightarrow N + \gamma^*, \end{aligned}$$

U(1)-U(1)
kinetic mixing

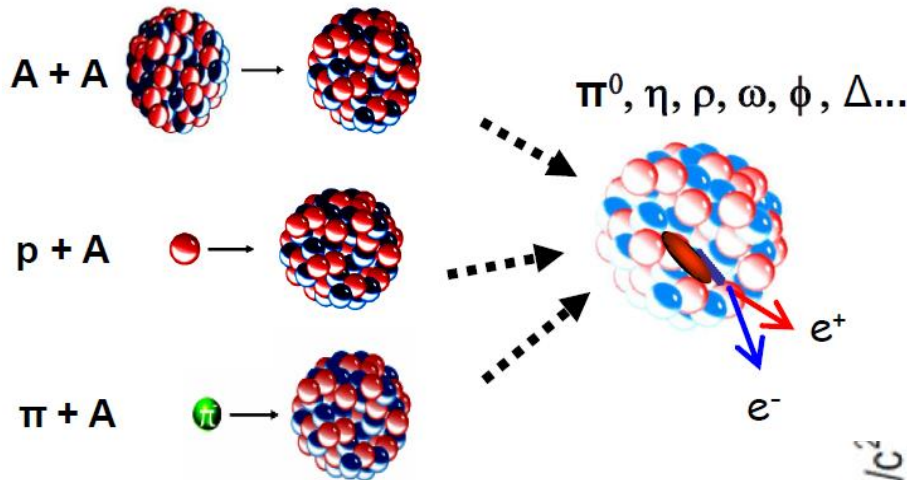
Beyond SM: DM scenario



$$\begin{aligned} \pi^0 &\rightarrow \gamma + U, \\ \eta &\rightarrow \gamma + U, \quad U \rightarrow e^+ e^-, \\ \Delta &\rightarrow N + U, \end{aligned}$$



Possible dark photon observation by dilepton experiments



□ Dilepton spectra from SM sources are well studied by dilepton experiments from SIS to LHC energies

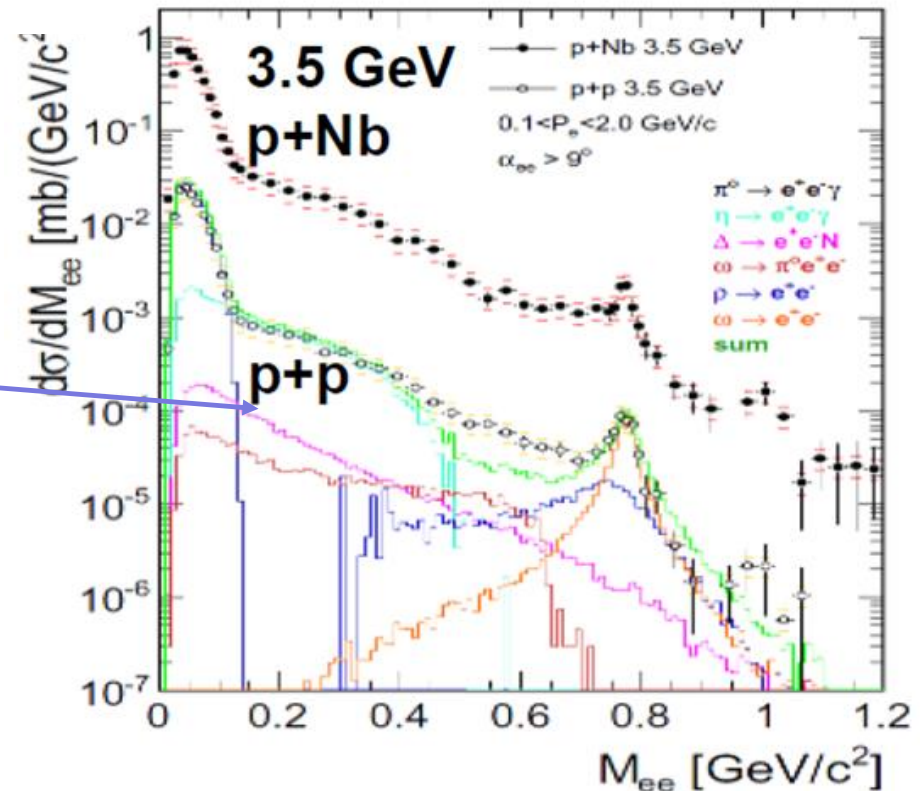
□ Hadron production by p+p, p+A, A+A

□ Hadron decay to dileptons

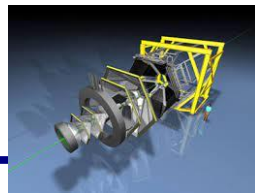
□ Dalitz π^0, η and Δ decays are the dominant dilepton sources at low M

→ Possibility for an experimental observation of dark photons by electromagnetic decays $U \rightarrow e^+e^-$ in heavy-ion experiments

Dilepton spectra at low M ('cocktail')



Search for dark photons with HADES (GSI)

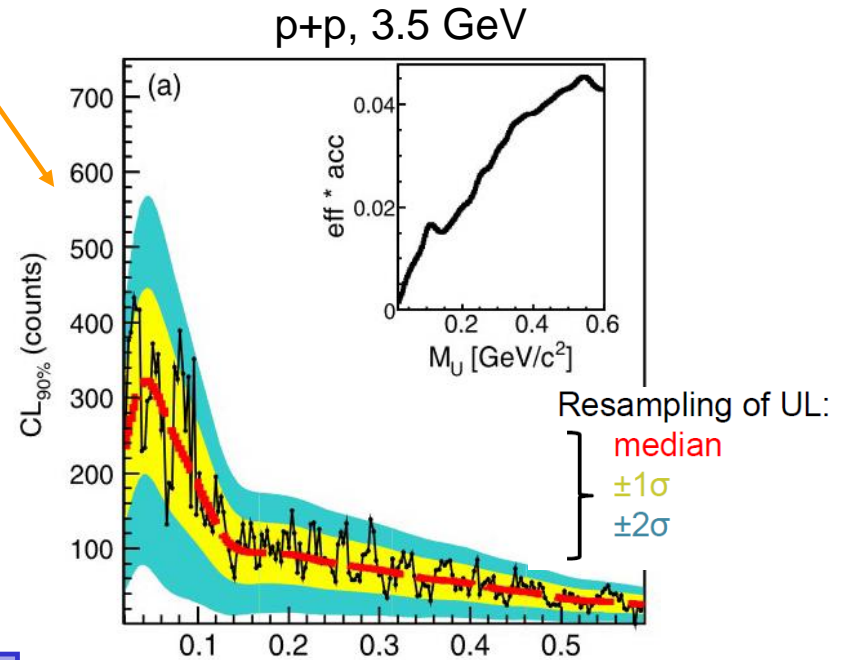
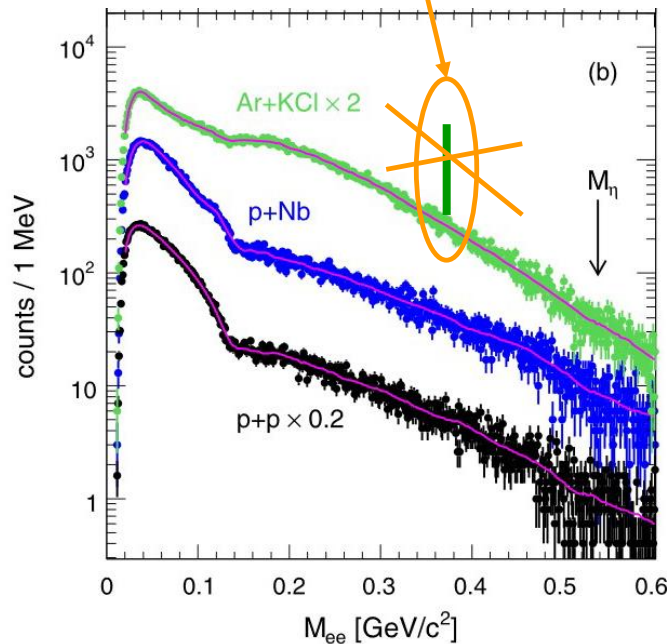


HADES data:

p+p, p+Nb at 3.5 GeV, Ar+KCl at 1.76 A GeV

G. Agakishiev et al. (HADES), Phys. Lett. B 731, 265 (2014)

- 1) **Search for a peak structure** in the raw dN/dM spectrum taking into account mass resolution: fit with 5th-order polynomial + Gauss peak for each fixed M bin
- 2) If no peak found, get an **UL (upper limit)** on peak



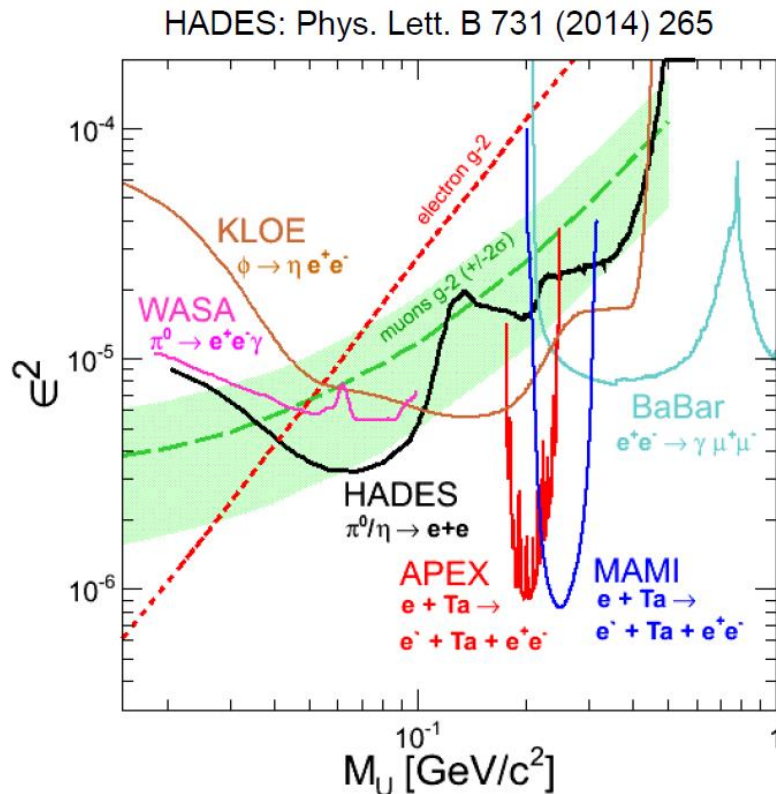
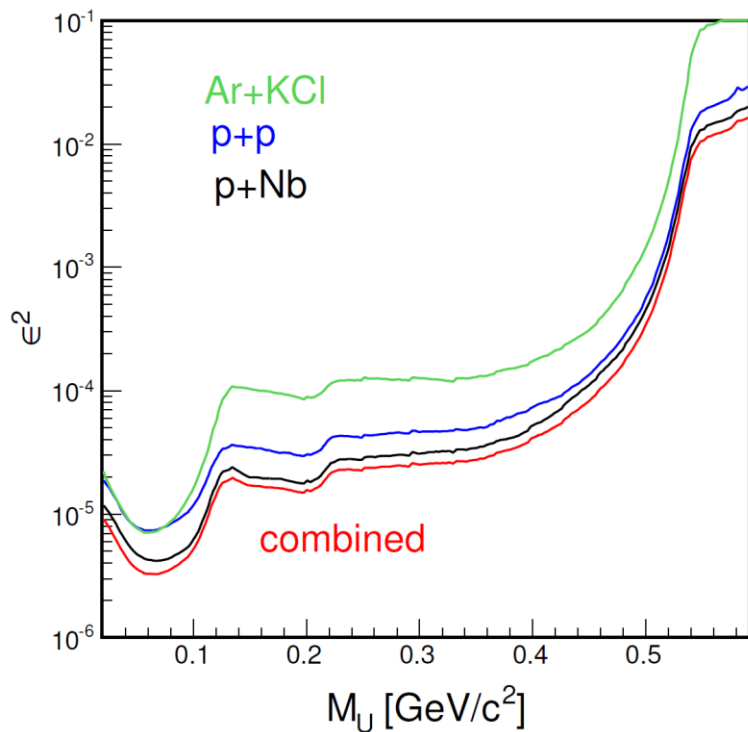
- 3) Transform this UL into an UL on the **mixing parameter ε^2** based on the **modelling of the U-boson production rate** (B. Batell, M. Pospelov, and A. Ritz, PRD 80,095024 (2009))

Search for dark photons with HADES



G. Agakishiev et al. (HADES), Phys. Lett. B 731, 265 (2014)

Upper limit on the mixing parameter ϵ^2



HADES coverage : $0.02 < M_U < 0.6$ GeV

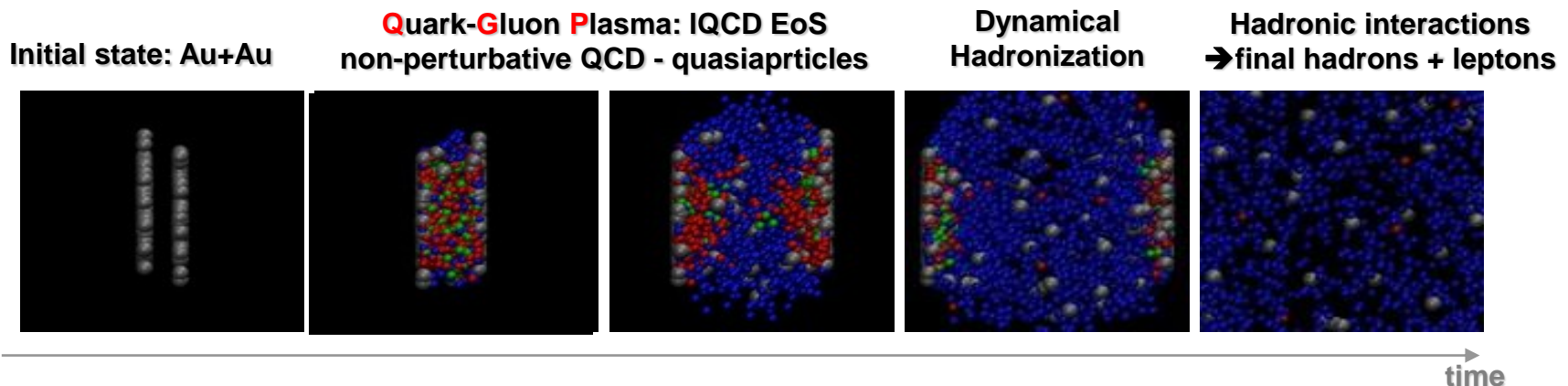
Dark photons are not observed so far!

Theoretical modeling of U-boson production

Goal: estimate the upper limit for the kinetic mixing parameter $\varepsilon^2(M_U)$ of the U-boson **from the theoretically calculated dilepton spectra** using the microscopic **PHSD** transport approach

Parton-Hadron-String Dynamics (PHSD) is a **non-equilibrium microscopic transport approach** for the description of strongly-interacting hadronic and partonic matter created in heavy-ion collisions

Dynamics: based on the solution of generalized off-shell transport equations derived from Kadanoff-Baym many-body theory



→ **PHSD** provides a good description of ‘bulk’ hadronic observables as well as **dilepton spectra** from SIS to LHC energies



Production of hadron \rightarrow decay to U \rightarrow **dilepton yield from U-boson decay of mass M_U** :

$$\begin{aligned} \pi^0 &\rightarrow \gamma + \mathbf{U}, \\ \eta &\rightarrow \gamma + \mathbf{U}, \quad \mathbf{U} \rightarrow e^+e^- \\ \Delta &\rightarrow \mathbf{N} + \mathbf{U}, \end{aligned}$$

$$\begin{aligned} N^{U \rightarrow e^+e^-} &= N_{\pi^0 \rightarrow e^+e^-}^U + N_{\eta \rightarrow e^+e^-}^U + N_{\Delta \rightarrow e^+e^-}^U \\ &= \boxed{Br^{U \rightarrow e^+e^-}} \boxed{(N_{\pi^0 \rightarrow \gamma U} + N_{\eta \rightarrow \gamma U} + N_{\Delta \rightarrow NU})}, \end{aligned}$$

□ **Dalitz decay of π^0, η mesons and Δ -resonances to U-bosons and real photons or N**

Based on the model:
(used in the HADES analysis)

B. Batell, M. Pospelov, and A. Ritz, Phys. Rev. D 79, 115008 (2009);
Phys. Rev. D 80, 095024 (2009)

- Ratio of the partial widths $\pi^0(\eta) \rightarrow \gamma + \mathbf{U}$ and $\pi^0(\eta) \rightarrow \gamma + \gamma$:

$$\frac{\Gamma_{i \rightarrow \gamma U}}{\Gamma_{i \rightarrow \gamma \gamma}} = 2 \underline{\epsilon^2} |F_i(q^2 = M_U^2)| \frac{\lambda^{3/2}(m_i^2, m_\gamma^2, M_U^2)}{\lambda^{3/2}(m_i^2, m_\gamma^2, m_\gamma^2)} \quad i = \pi^0, \eta$$

Formfactors:

$$|F_{\pi^0}(q^2)| = 1 + 0.032 \frac{q^2}{m_{\pi^0}^2}$$

$$|F_\eta(q^2)| = \left(1 - \frac{q^2}{\Lambda^2}\right)^{-1}$$

$\Lambda = 0.72 \text{ GeV}$

$$|F_\Delta(M_U^2)| = 1$$

- Ratio of the partial widths $\Delta \rightarrow \mathbf{N} + \mathbf{U}$ and $\Delta \rightarrow \gamma + \mathbf{N}$:

$$\frac{\Gamma_{\Delta \rightarrow NU}}{\Gamma_{\Delta \rightarrow N\gamma}} = \underline{\epsilon^2} \int A(M_\Delta) |F_\Delta(M_U^2)| \frac{\lambda^{3/2}(M_\Delta^2, m_N^2, M_U^2)}{\lambda^{3/2}(M_\Delta^2, m_N^2, m_\gamma^2)} dM_\Delta$$

Spectral function of Δ -resonance:
$$A_\Delta(M_\Delta) = C_1 \cdot \frac{2}{\pi} \frac{M_\Delta^2 \Gamma_\Delta^{tot}(M_\Delta)}{(M_\Delta^2 - M_{\Delta 0}^2)^2 + (M_\Delta \Gamma_\Delta^{tot}(M_\Delta))^2}$$

Dark photon production in PHSD

I.) $(N_{\pi^0 \rightarrow \gamma U} + N_{\eta \rightarrow \gamma U} + N_{\Delta \rightarrow NU})$

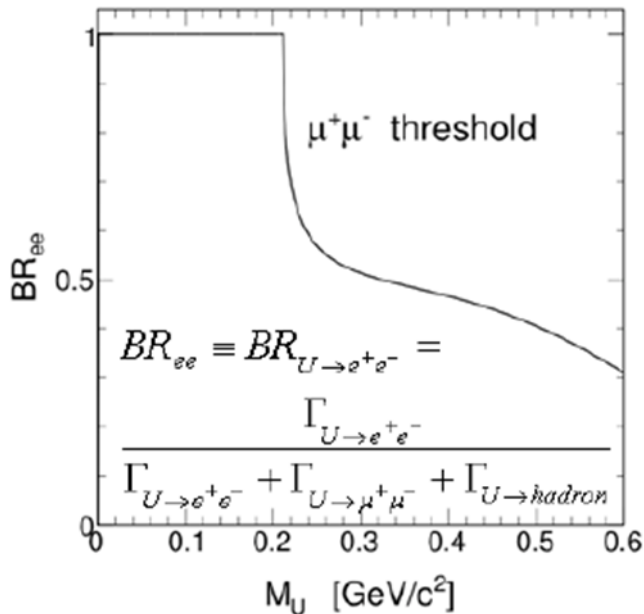
The **yield of U-bosons** of mass M_U :

$$N_{i \rightarrow \gamma U} = N_i Br_{i \rightarrow \gamma \gamma} \cdot \frac{\Gamma_{i \rightarrow \gamma U}}{\Gamma_{i \rightarrow \gamma \gamma}}, \quad i = \pi^0, \eta$$

$$N_{\Delta \rightarrow NU} = N_{\Delta} Br_{\Delta \rightarrow N \gamma} \cdot \frac{\Gamma_{\Delta \rightarrow NU}}{\Gamma_{\Delta \rightarrow N \gamma}}$$

II.) $Br^{U \rightarrow e^+ e^-}$

Branching ratio for the decay of U-bosons to e^+e^- : $Br_{ee} = Br^{U \rightarrow ee} = \frac{\Gamma_{U \rightarrow e^+ e^-}}{\Gamma_{tot}^U}$



$$\Gamma_{tot}^U = \Gamma_{U \rightarrow hadr} + \Gamma_{U \rightarrow e^+ e^-} + \Gamma_{U \rightarrow \mu^+ \mu^-}$$

$$\Gamma_{U \rightarrow hadr} = R(\sqrt{s} = M_U) \Gamma_{U \rightarrow \mu^+ \mu^-}$$

$$R(\sqrt{s}) = \sigma_{e^+ e^- \rightarrow hadrons} / \sigma_{e^+ e^- \rightarrow \mu^+ \mu^-}$$

$$Br^{U \rightarrow ee} = \frac{\Gamma_{U \rightarrow e^+ e^-}}{\Gamma_{tot}^U} = \frac{1}{1 + \sqrt{1 - \frac{4m_{\mu}^2}{M_U^2}} \left(1 + \frac{2m_{\mu}^2}{M_U}\right) (1 + R(M_U))}$$



How to obtain theoretical constraints on $\varepsilon^2(M_U)$?

Basic ideas:

- 1) There is **NO evidence for an experimental observation of dark photons** in dilepton experiments **so far**
- 2) Theoretical model (**PHSD**) provides a good description of exp. data on SM dileptons (from SIS to LHC energies)
- 3) Dark photon yield is proportional to $\varepsilon^2(M_U)$
- 4) **Use the theoretical spectra** – instead of exp. data – **to estimate an upper limit for $\varepsilon^2(M_U)$** from the **constraint** that an additional yield from dark photons **can not exceed the total yield** from standard sources by more than a small factor (for each M_U), which we can vary in our calculations
- 5) Variation of this **'surplus' factor** can provide the range of possible $\varepsilon^2(M_U)$ and can be related to **experimental accuracy** e.g. via error bars, mass resolution, acceptance etc.

Procedure to obtain constraints on $\epsilon^2(M_U)$

1) For each bin $[M_U, M_U + dM]$ calculate the **sum of all $U \rightarrow e^+e^-$ contributions** (kinematically possible in this mass bin)

$$\frac{dN^{sumU}}{dM} = \frac{dN_{\pi^0}^{U \rightarrow e^+e^-}}{dM} + \frac{dN_{\eta}^{U \rightarrow e^+e^-}}{dM} + \frac{dN_{\Delta}^{U \rightarrow e^+e^-}}{dM}$$

Can be presented as $\frac{dN^{sumU}}{dM} = \epsilon^2 \frac{dN_{\epsilon=1}^{sumU}}{dM}$

2) Calculate the **sum of all SM contributions** and ‘dark matter’ (DM) contributions :

$$\frac{dN^{total}}{dM} = \frac{dN^{sumSM}}{dM} + \frac{dN^{sumU}}{dM} = \frac{dN^{sumSM}}{dM} + \epsilon^2 \frac{dN_{\epsilon=1}^{sumU}}{dM}$$

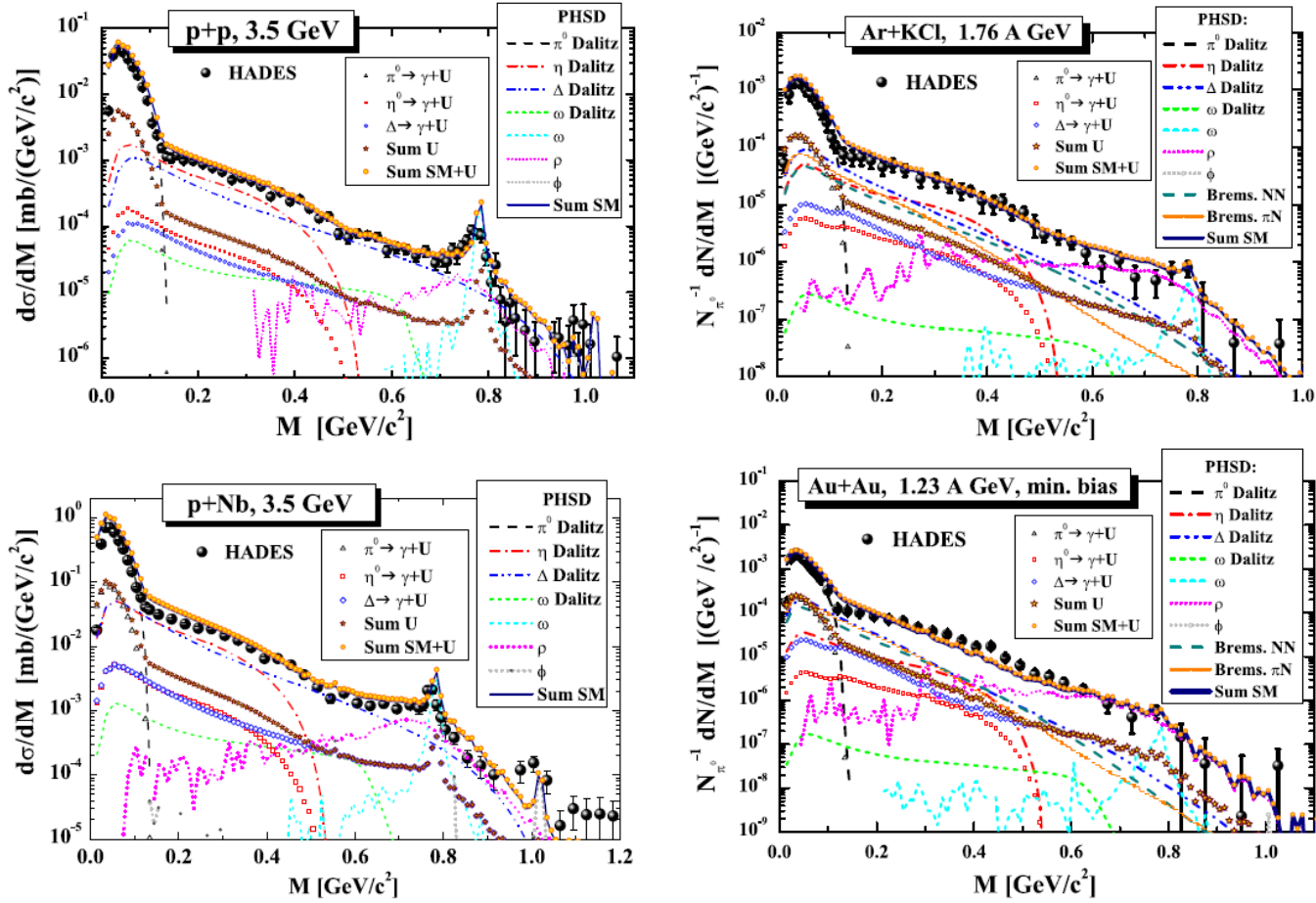
3) Obtain **constraints** by requesting that dN^{total}/dM cannot **exceed** the **sum of SM channels** (i.e. exp. data!) by more than a factor C_U in each bin dM , i.e.

$$\frac{dN^{total}}{dM} = (1 + C_U) \frac{dN^{sumSM}}{dM} \quad \rightarrow \quad C_U \text{ controls the allowed "surplus" dilepton yield resulting from dark photons on top of the total SM yield}$$

4) Calculate $\epsilon^2(M_U)$ by assuming C_U : e.g. $C_U = 0.1 \rightarrow 10\%$ DM extra yield to the SM yield

$$\epsilon^2(M_U) = C_U \cdot \left(\frac{dN^{sumSM}}{dM} \right) / \left(\frac{dN_{\epsilon=1}^{sumU}}{dM} \right)$$

Dileptons yields including dark photons vs. HADES data

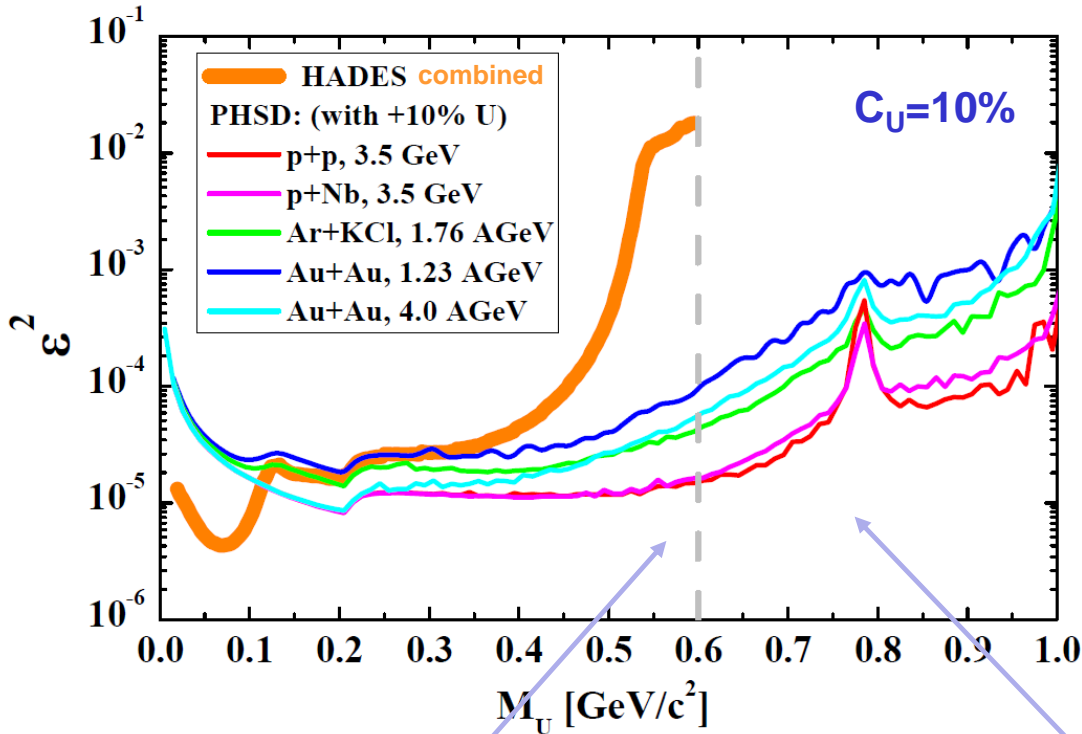


- The HADES data, i.e. **SM contributions** (including exp. acceptance) are well described by the PHSD
- The contributions from **$U \rightarrow e+e^-$** are added with **$C_U=10\%$** allowed surplus of the total SM yield \rightarrow the **total sum** is still in a good agreement with exp. data

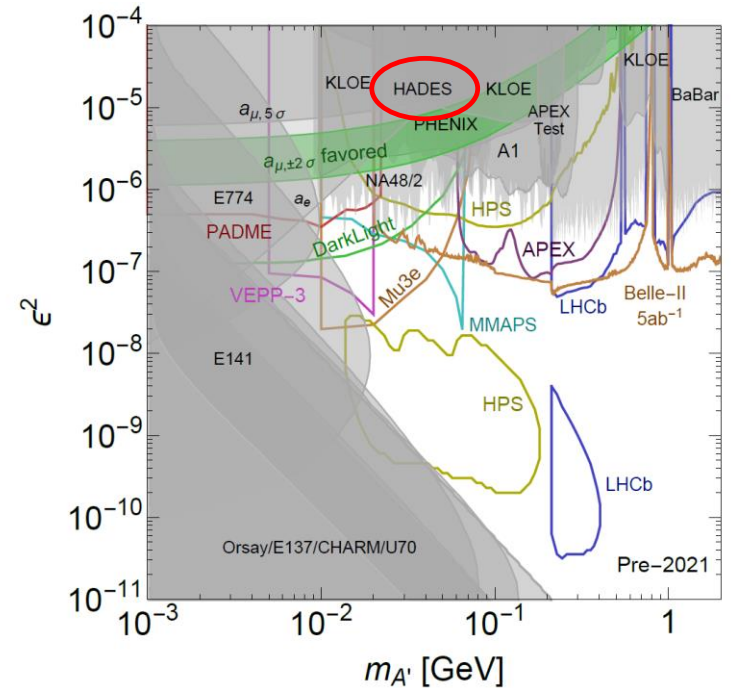


Mixing parameter $\epsilon^2(M_U)$

The **upper limit for the kinetic mixing parameter $\epsilon^2(M_U)$** of light dark photons extracted from the PHSD dilepton spectra - with **10% allowed surplus** of the total SM yield



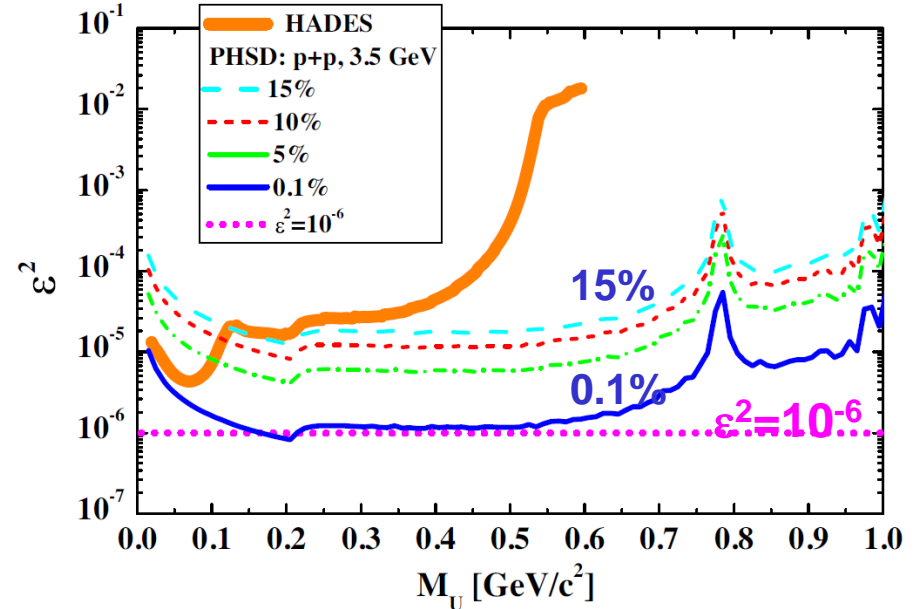
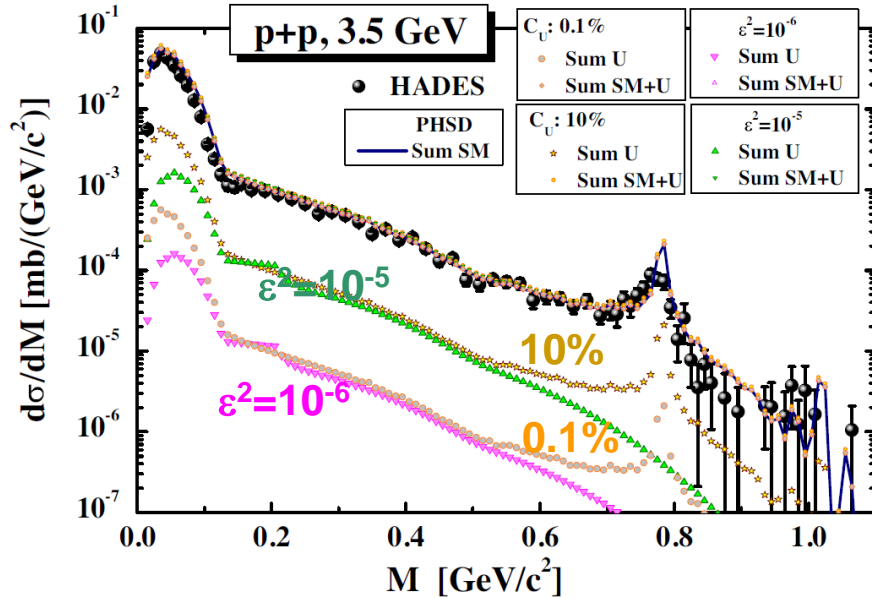
Validity range of extracted kinetic mixing parameter: **$0 < M_U < 0.6$ GeV** based on low energy dilepton spectra



Possible contribution from other dark photon channels

Limits for the mixing parameter $\varepsilon^2(M_U)$

- The PHSD predictions for $\varepsilon^2(M_U)$ with 0.1%, 5%, 10%, and 15% allowed surplus of the U-boson contributions over the total SM yield



The **theoretically** extracted upper limit of the kinetic mixing parameter $\varepsilon^2(M_U)$ of light dark photons from Dalitz decays of π^0, η mesons and Δ -resonances:

- strongly reduces by reducing the allowed 'surplus'
- exp. data of high precision is needed to reduce the upper limit for $\varepsilon^2(M_U)$



Summary

- ❑ We presented **first microscopic transport calculations**, based on the PHSD approach, for the **dilepton yield from the decay of hypothetical dark photons** (or U-bosons), $U \rightarrow e^+e^-$ from $p + p$, $p + A$ and heavy-ion collisions at SIS18 energies
- ❑ For that we incorporated in the PHSD the **production of U-bosons** by the Dalitz decay $\pi^0 \rightarrow \gamma + U$, $\eta \rightarrow \gamma + U$, $\Delta \rightarrow N + U$ with further dilepton decays $U \rightarrow e^+e^-$ based on the theoretical model by Batell, Pospelov and Ritz, which describes the interaction of DM and SM particles by the **U(1)-U(1)' mixing**
- ❑ We **introduced a procedure to define theoretical constraints on the upper limit of the kinetic mixing parameter $\varepsilon^2(M_U)$** :
Since dark photons are not observed in dilepton experiments so far, we can require that their contribution **can not exceed some limit** which would make them visible in experimental data
- ❑ We found that the **extracted upper limit of $\varepsilon^2(M_U)$ is consistent with the experimental results of the HADES experiment** for $0.15 < M_U < 0.4$ GeV, as well as with the world-wide experimental compilation
- ➔ **Proposed theoretical procedure allows:**
 - to check any theoretical ideas on the $\varepsilon^2(M_U)$ independent on exp. data
 - to study the influence of exp. acceptance, system and centrality selection
 - to perform the simulation for testing experimental set-ups for the search of U-bosons