Can Loop Quantum Cosmology describe an unified and consistent scenario since bounce until the end of inflation?

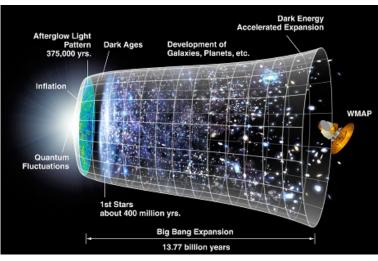
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Introduction

Very Early Universe

- Planck scale;
- Standard Cosmology (from Big Bang until the end of inflation);
- Inflationary Universe;



Credit: NASA/WMAP

Open Problems

- Big Bang singularity;
- Inflaton nature;
- Quantum theory of gravity;

Loop Quantum Cosmology

- Loop Quantum Gravity (LQG);
- Quantization of cosmological space-times through techniques and variables from LQG;
- The Big Bang initial singularity is replaced by a Big Bounce:

$$H^2 = \frac{8\pi G}{3} \rho \left(1 - \frac{\rho}{\rho_c} \right); \tag{1}$$

- It allows the existence of a previous version of our universe;
- Mimetic gravity description of LQC;

Higgs Inflation

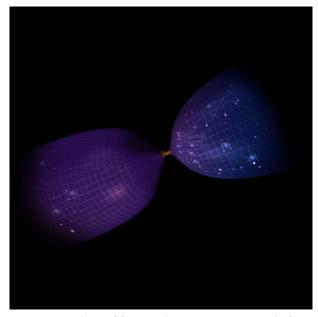
- The Higgs field is the only scalar field currently observed;
- Primordial version of Higgs field;
- Non-minimal coupling between gravity and matter;

How it can be implemented?

Mimetic Gravity description of LQC + like-Higgs Inflation;

Evolutionary Phases of Primordial Universe

- Contracting;
- Bounce;
- Expanding;



Source: http://www.physics.princeton.edu/

Effective Friedmann Equation

• Reinterpreting the effective Friedmann equation obtained in Langlois *et al.*(2017)

$$H^{2} = \frac{1}{3M_{Pl}^{2}} \left[\rho + 3M_{Pl}^{2} \left(V_{k}(h) - \frac{k}{a^{2}} \right) \right] \times \left\{ 1 - \frac{1}{\rho_{c}} \left[\rho + 3M_{Pl}^{2} \left(V_{k}(h) - \frac{k}{a^{2}} \right) \right] \right\}; \tag{2}$$

Effective energy density

$$\rho_{\text{eff}} = \rho + 3M_{Pl}^2 \left[V_k(h) - \frac{k}{a^2} \right] \text{ where } \rho = \frac{\pi_h^2}{2a^6}; \qquad (3)$$

Recovering the effective Friedmann equation

$$H^2 = \frac{\rho_{\text{eff}}}{3M_{Pl}^2} \left(1 - \frac{\rho_{\text{eff}}}{\rho_c} \right); \tag{4}$$

Bounce

$$H^2 = \frac{8\pi G}{3} \rho \left(1 - \frac{\rho}{\rho_c} \right); \tag{5}$$

Outset of Inflation

$$H^{2} = \left[V_{k}(h) - \frac{k}{a^{2}}\right] \left\{1 - \frac{3M_{Pl}^{2}}{\rho_{c}} \left[V_{k}(h) - \frac{k}{a^{2}}\right]\right\}; \qquad (6)$$

End of inflation

$$H^2 \approx \frac{1}{3M_{Pl}^2} \left[\rho + 3M_{Pl}^2 V_k(h) \right];$$
 (7)

Final remarks

- Origin of current universe;
- Higgs field;
- Regarding curvature;
- Pair Higgs field and gravity;
- Recovering of both standard models;

Thank you!

Questions?