

# Transit Timing Variations with NGTS



**Daniel Bayliss, University of Warwick**

On behalf of the NGTS Consortium



**UNIVERSITY  
OF WARWICK**



Science and  
Technology  
Facilities Council

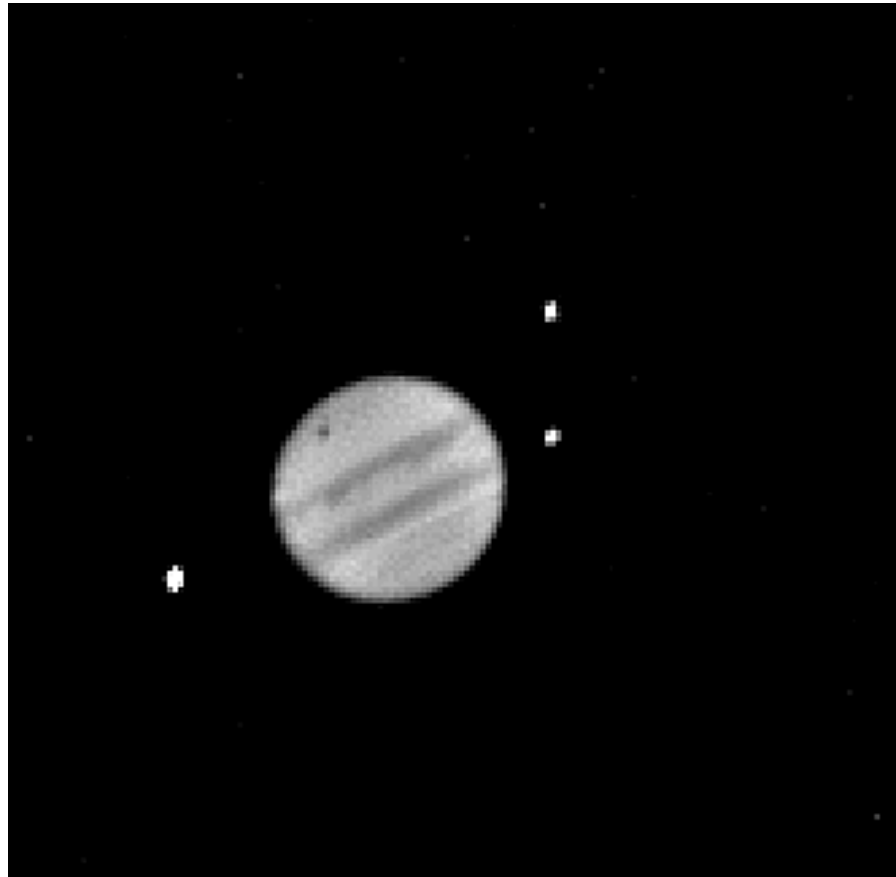
**NGTS**  
NEXT-GENERATION TRANSIT SURVEY



# 1554 Lopo Homem world map

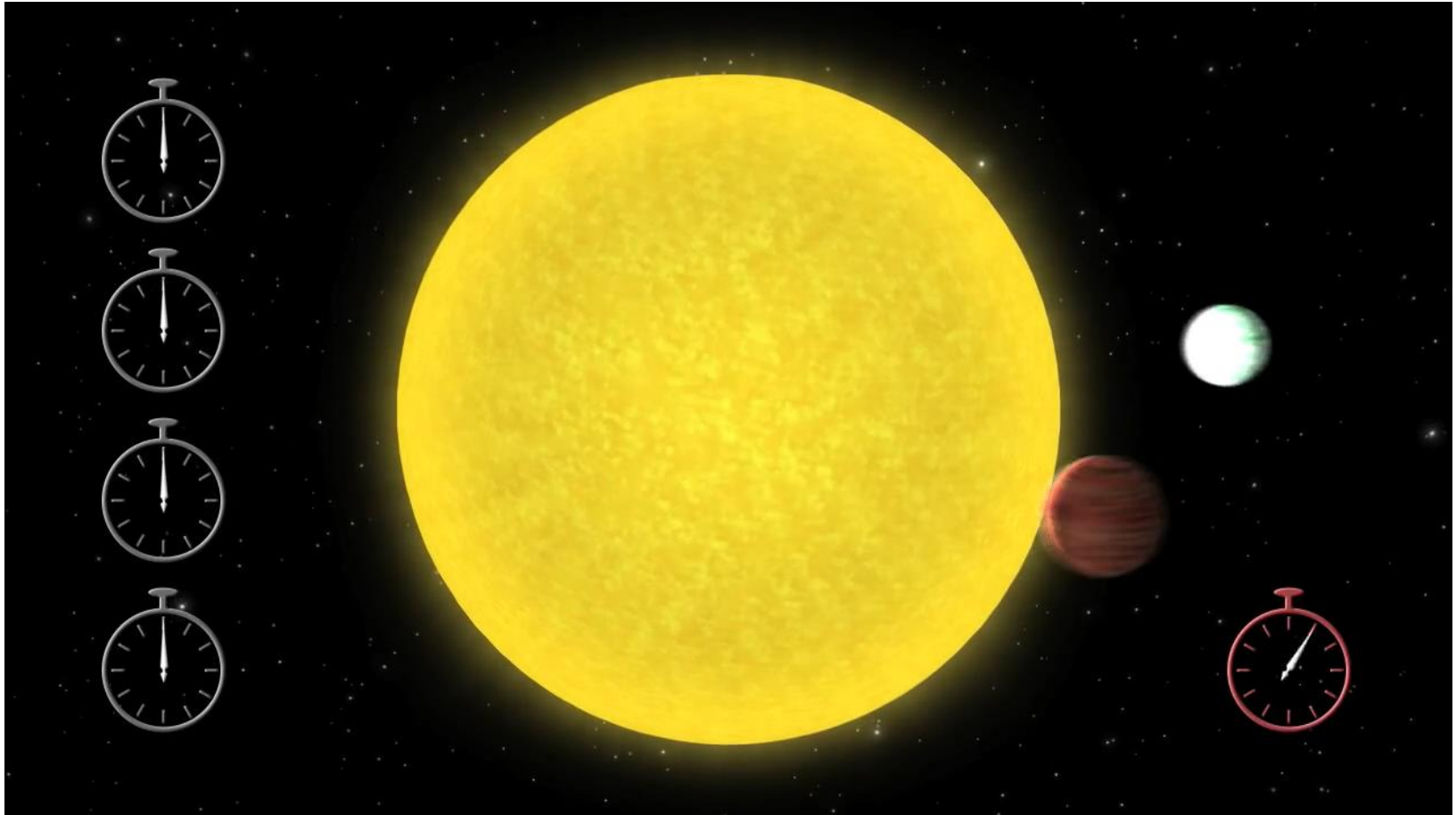


# 17<sup>th</sup> Century - Transit timing to map the new worlds



2h observation of Jupiter (Bruce Gray)

# 21<sup>st</sup> Century - Transit timing to map the new worlds



Credit: NASA Ames/Kepler mission

# NGTS Facility – 12 x 20cm telescopes at Paranal Observatory

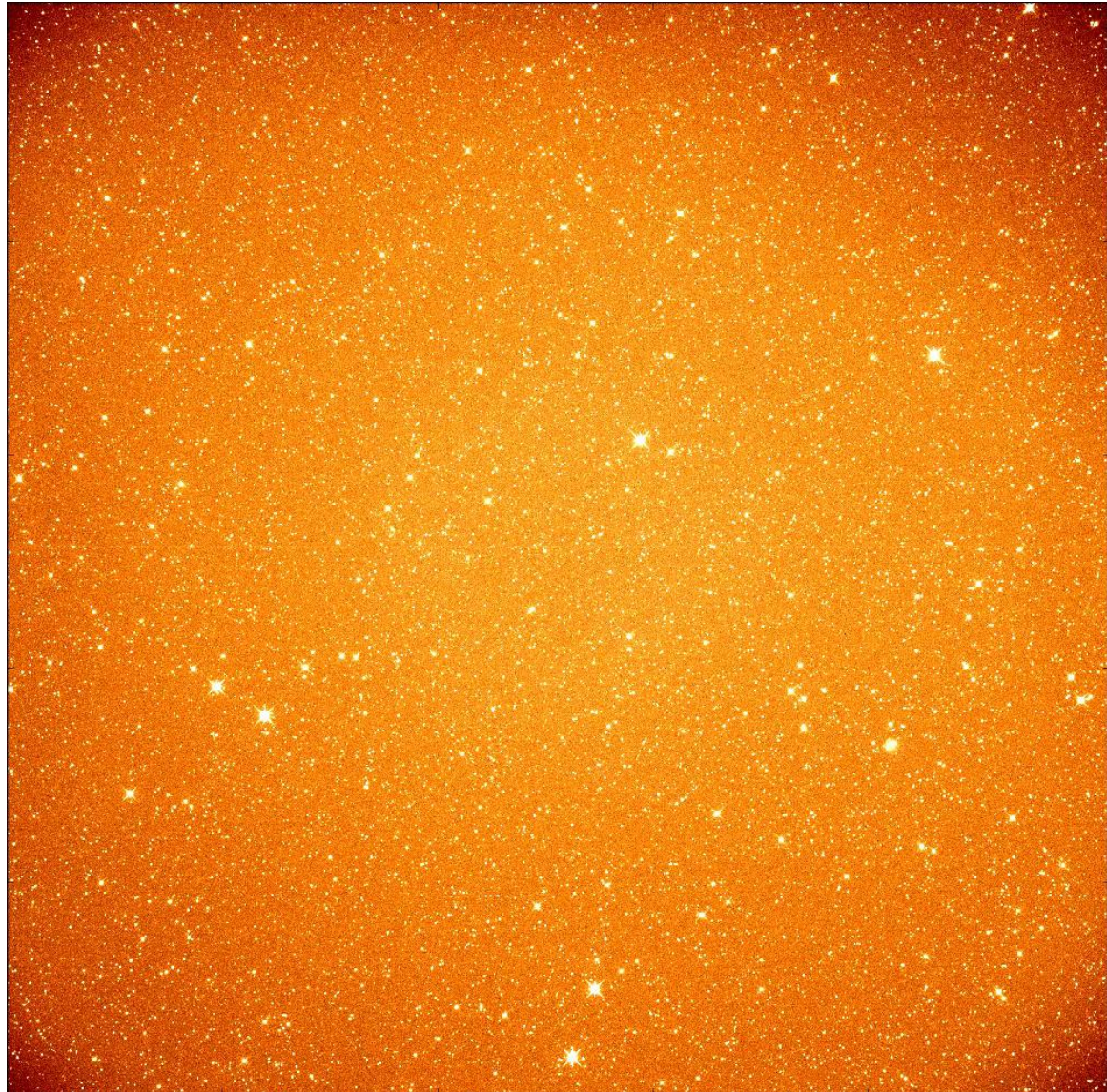


# NGTS Facility – Fully automated and robotic



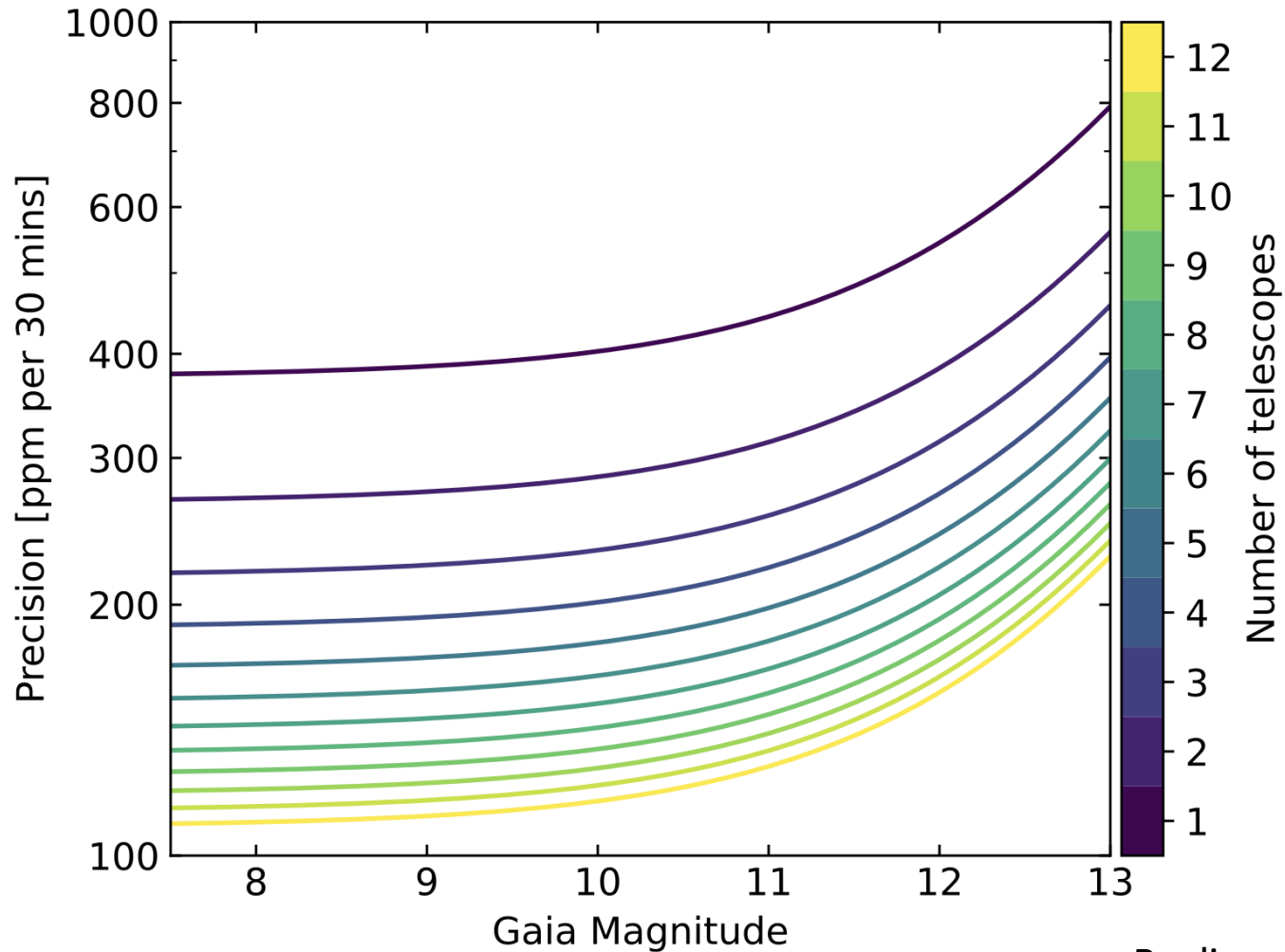
Wheatley et al., 2018

# 13s cadence in wide (R+I) band



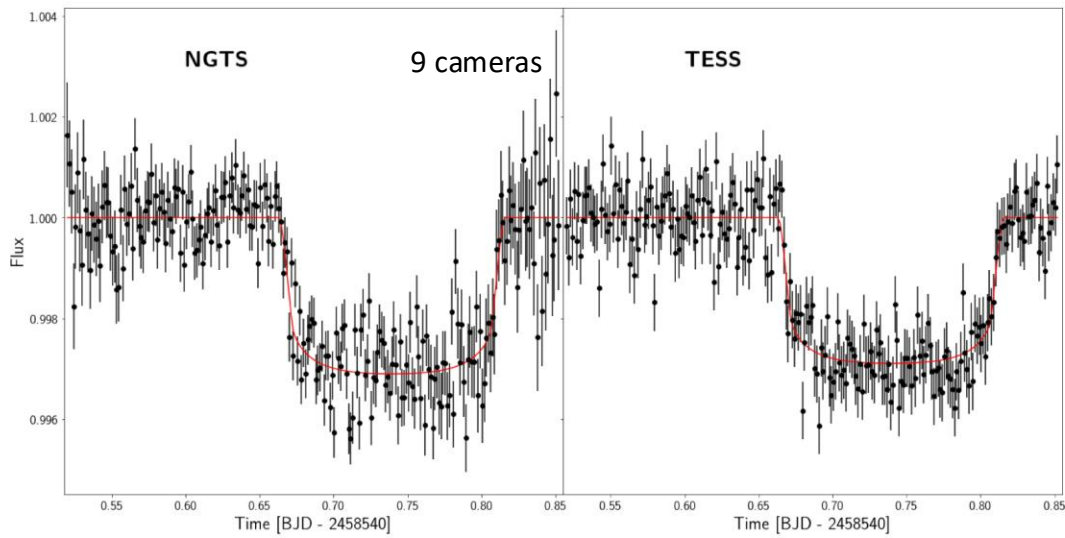
5'' pixels  
2.8x2.8 deg FOV

# Precision for 1 to 12 telescopes

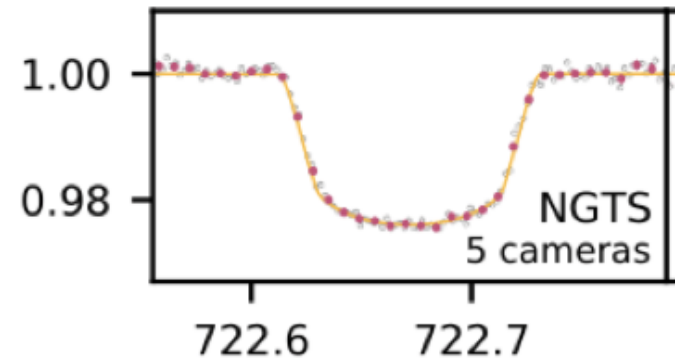
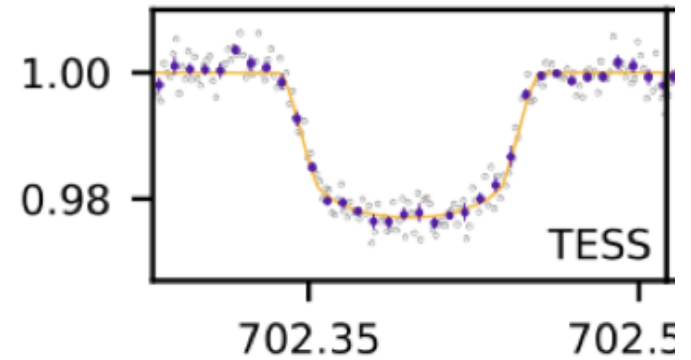


Bayliss et al., 2022.

# Photometric precision similar to TESS



Simultaneous transits of WASP-166b on 25 February 2019 , Bryant et al., 2020.



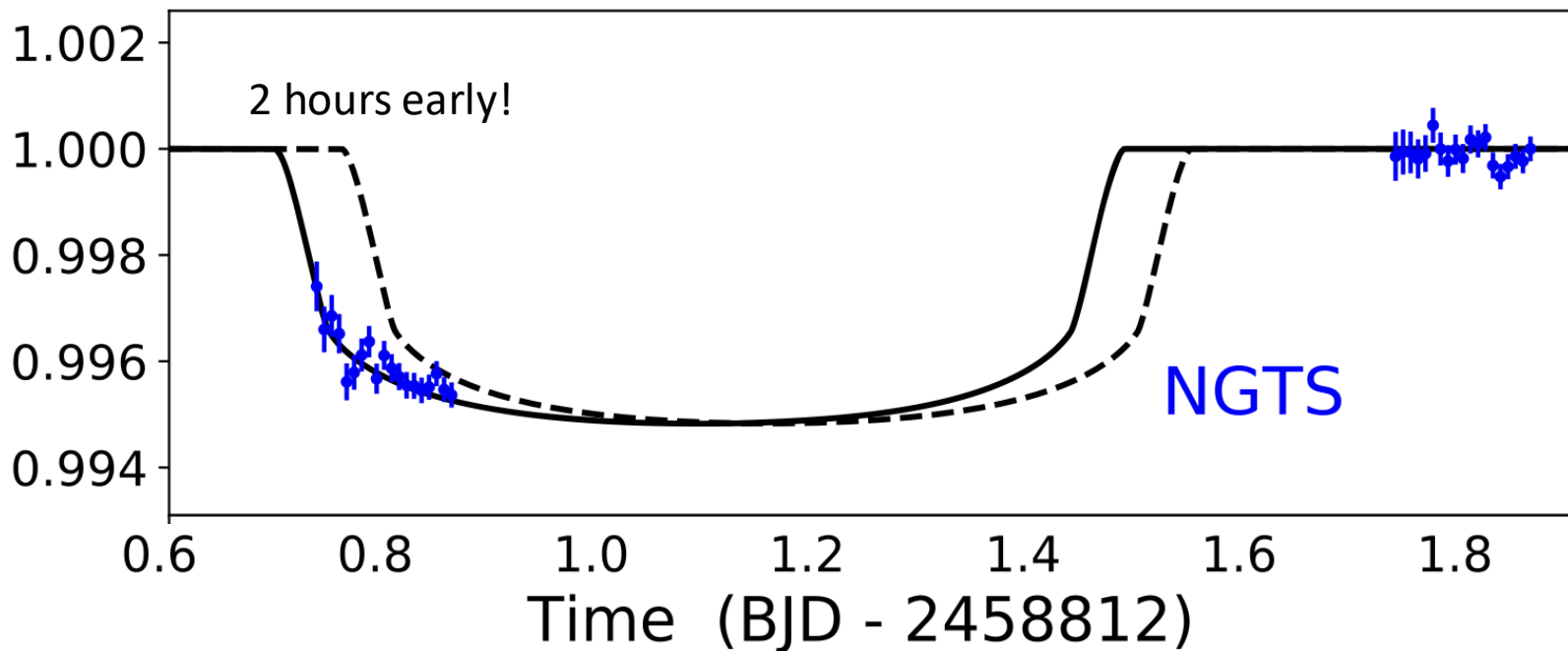
Transits of WASP-39b, Ahrer et al., 2023.

# 1. NGTS Orbital Decay Program

- Sample of 12 very hot Jupiters well suited to measuring TTVs to detect orbital decay.
- Began in 2019. Transits measured several times per year. Aim to build up a decade-long homogeneous set of transits.

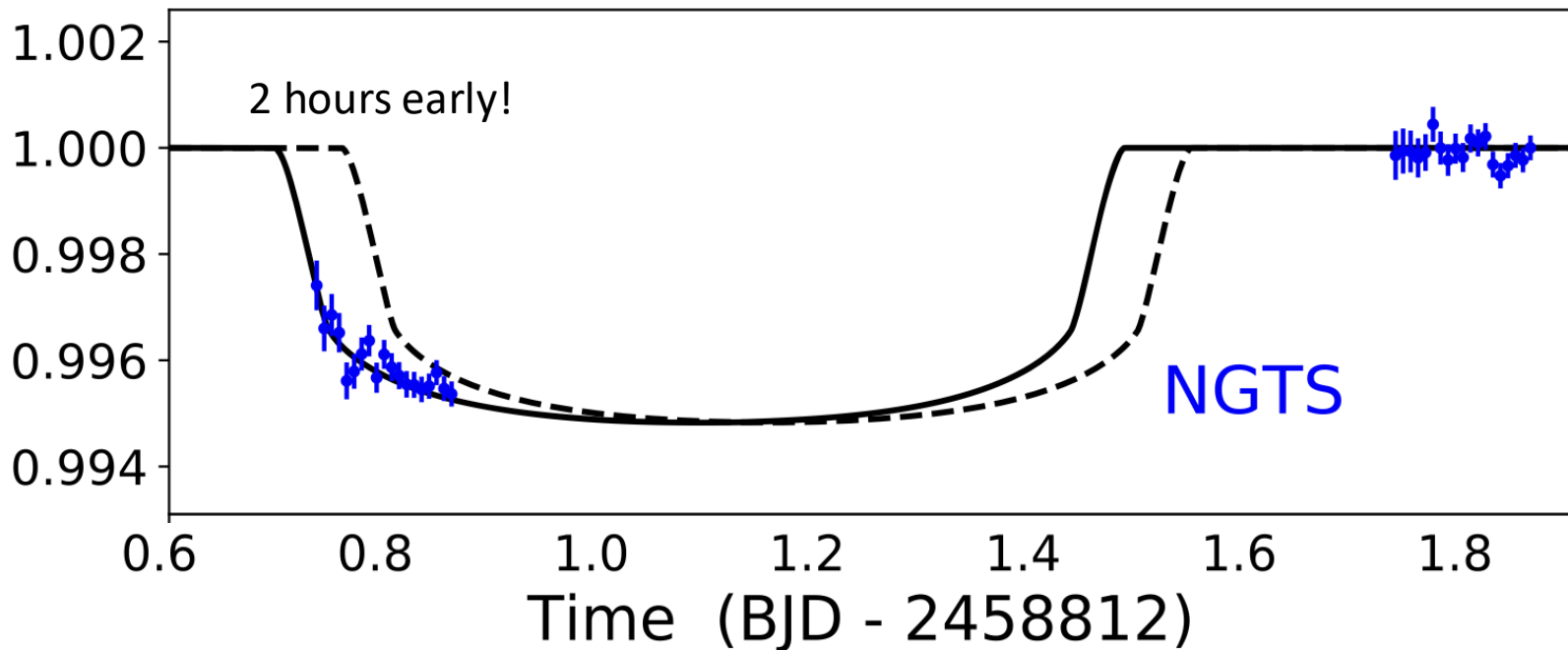
## 2. Multi-planet TTV program

- Sample of  $\sim 10$  bright star multi-planet systems from TESS.
- Example: HIP-41378f (Bryant et al., 2021)

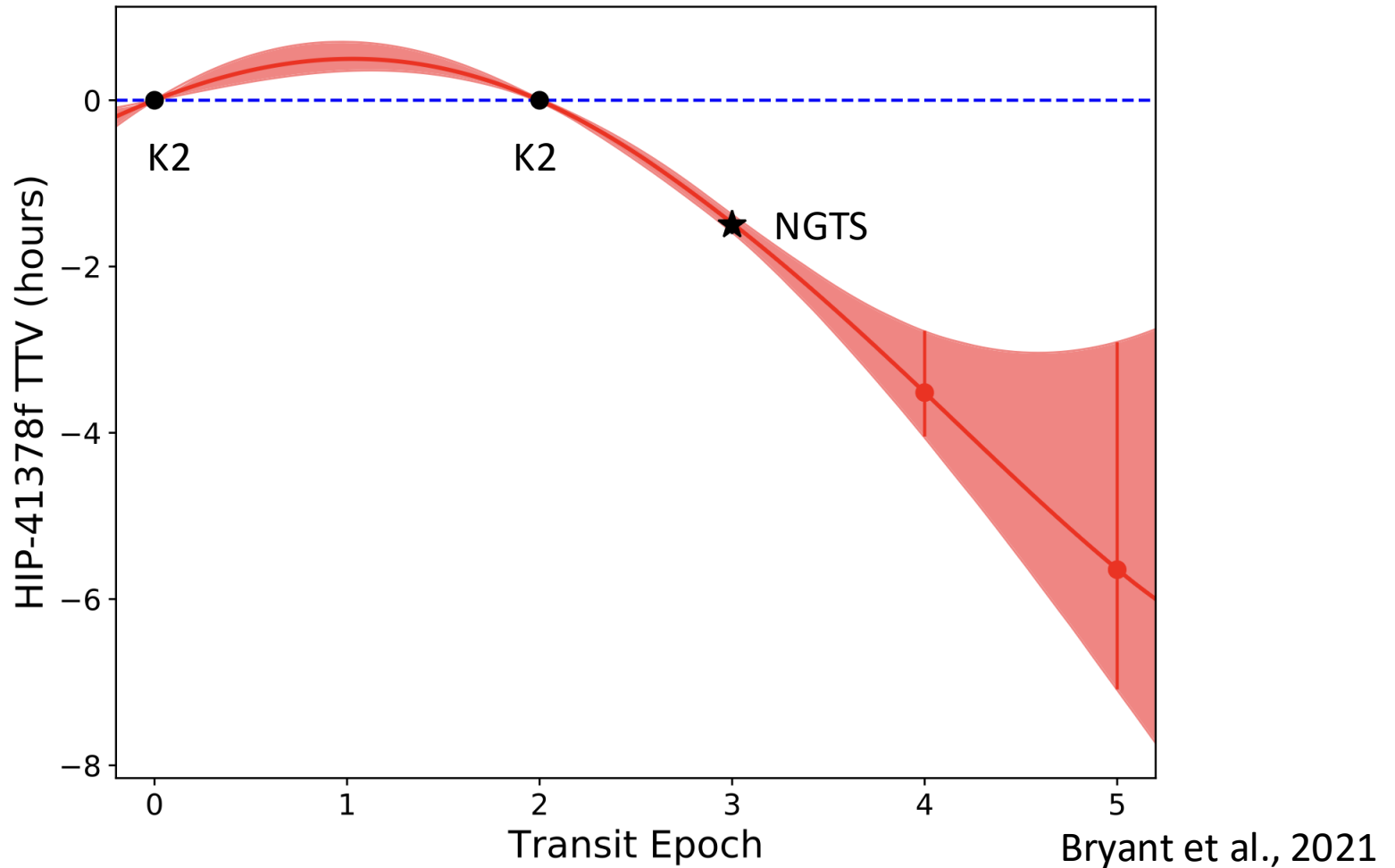


## 2. Multi-planet TTV program

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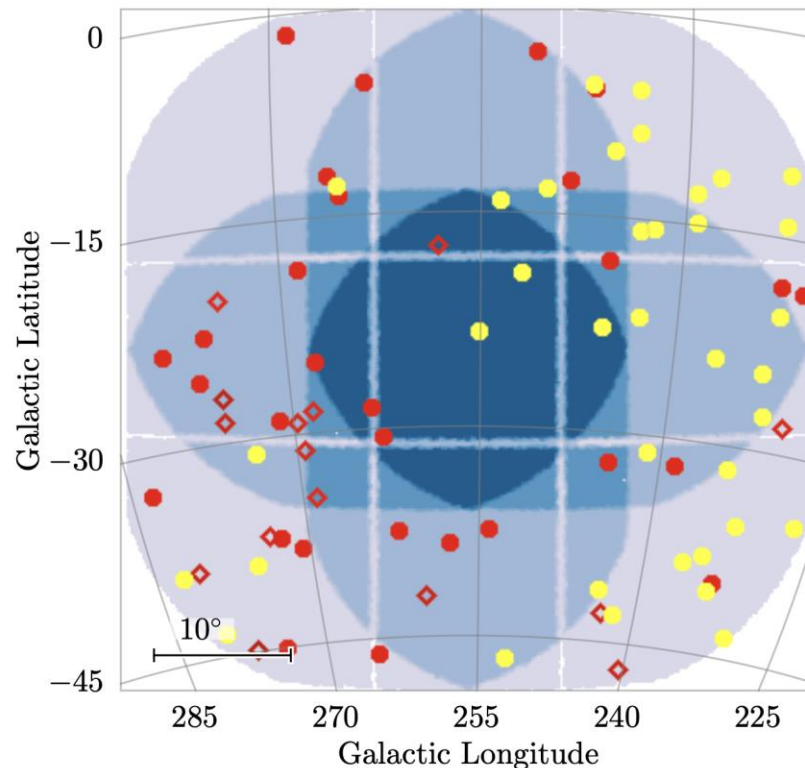


# Crucial TTV discovery for HIP-41378f



# 3. PLATO LOPS2 TTV program

- Sample “long period” ( $P > 20d$ ) transiting planets, focusing on the PLATO LOPS2 FOV .



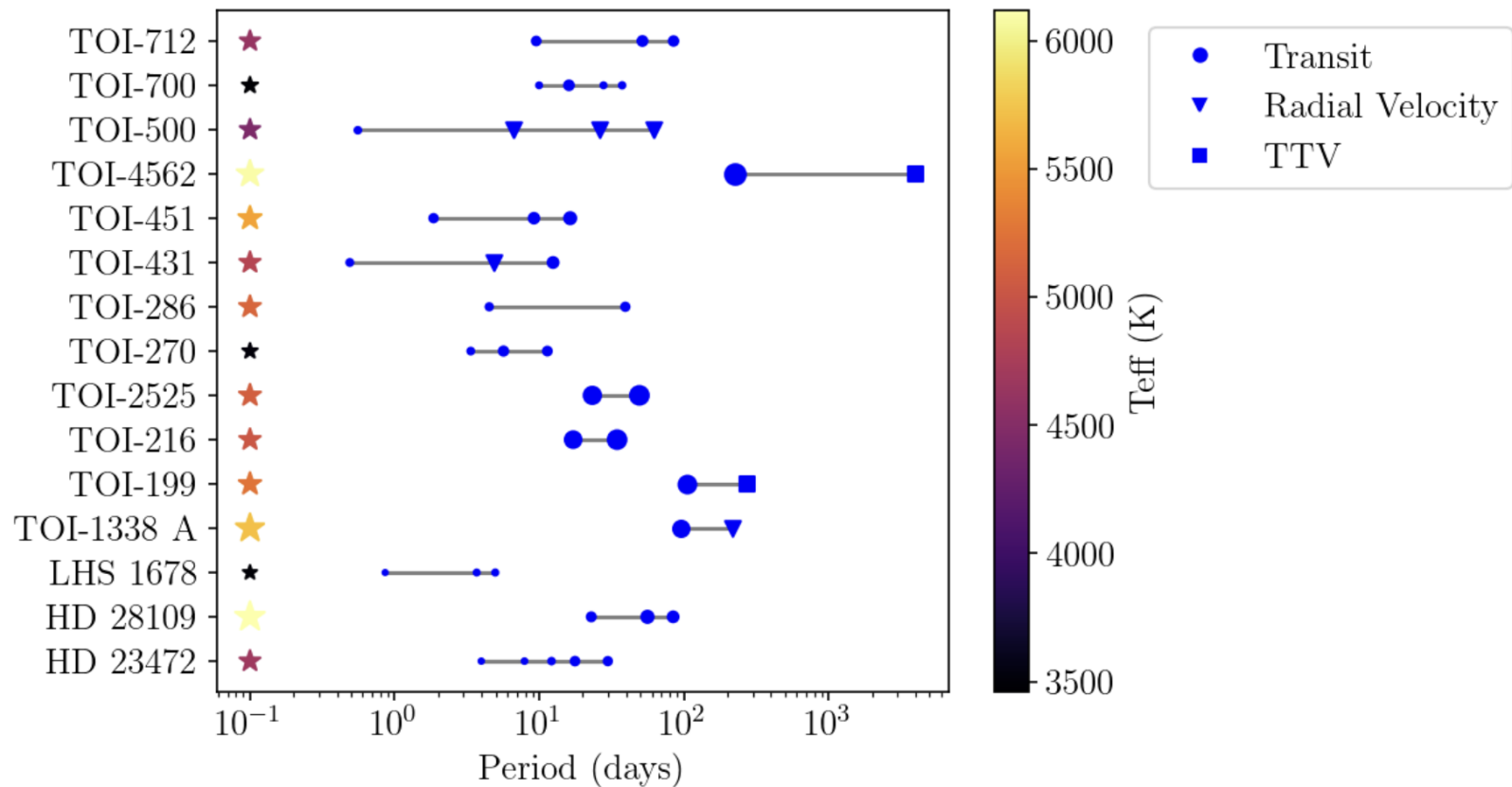
Yellow – Known transiting planets

Red – TESS discovered transiting planets

Diamonds – Multi-planet systems

Eschen et al., 2024

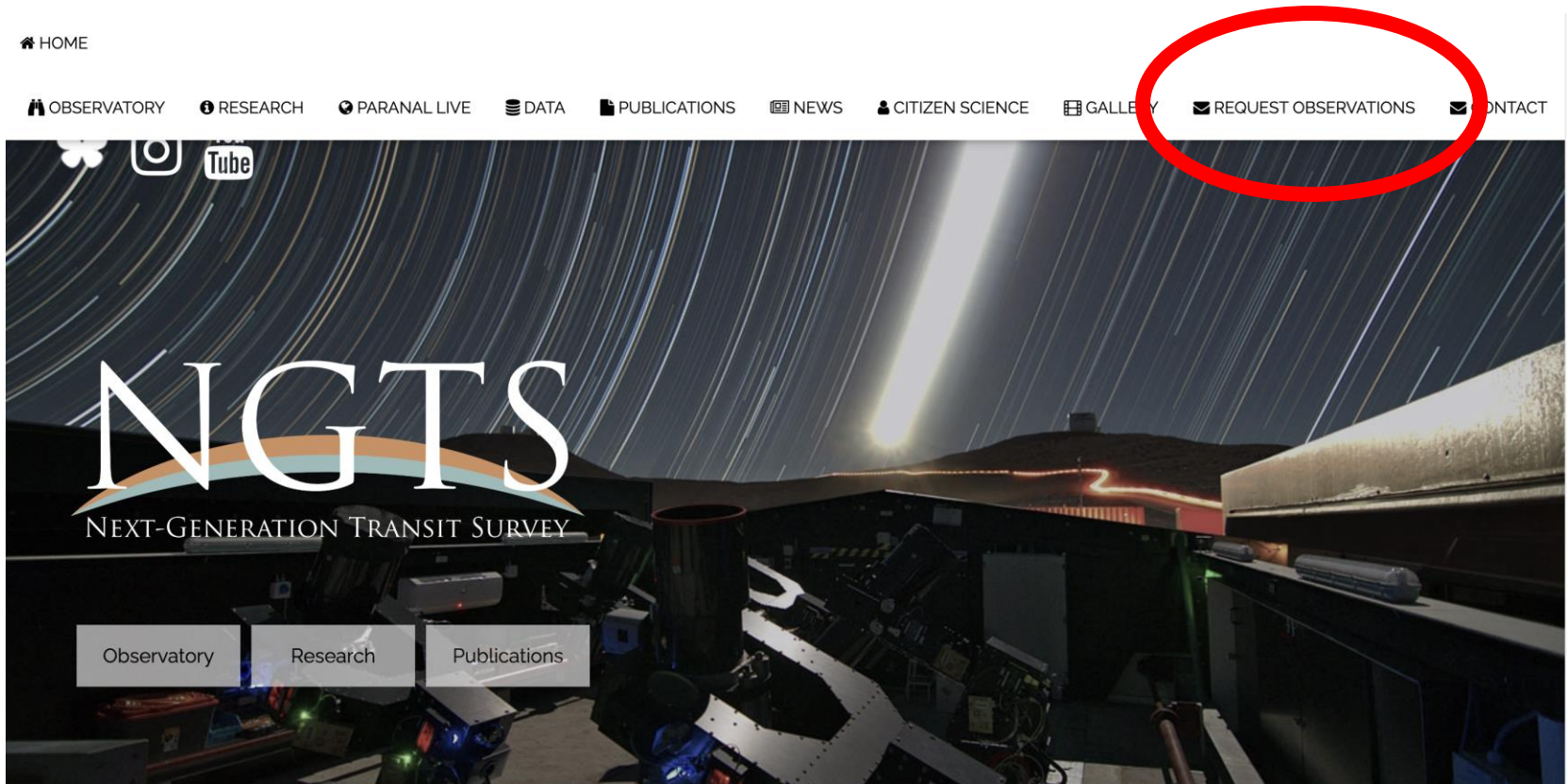
# Multi-planet Systems in PLATO LOPS2 FOV



Eschen et al., 2024

# 4. [Insert your target(s) here]

[ngtransits.org](https://ngtransits.org)



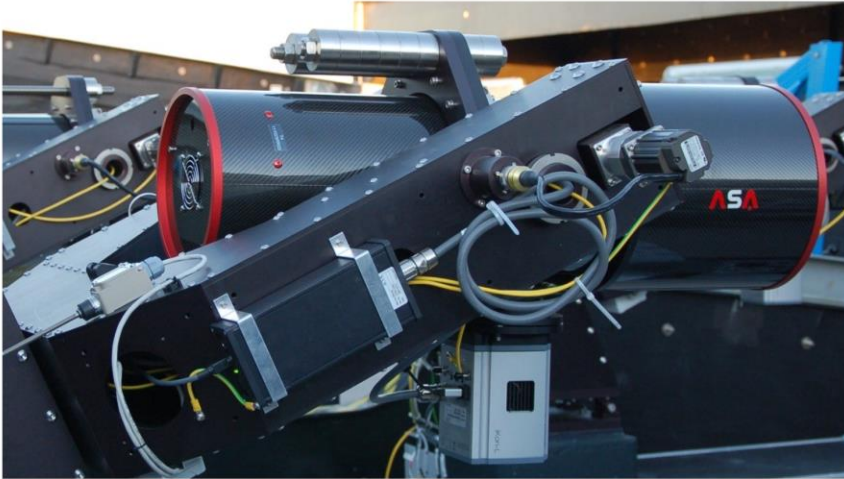
# Obrigado!



# Bonus Slides



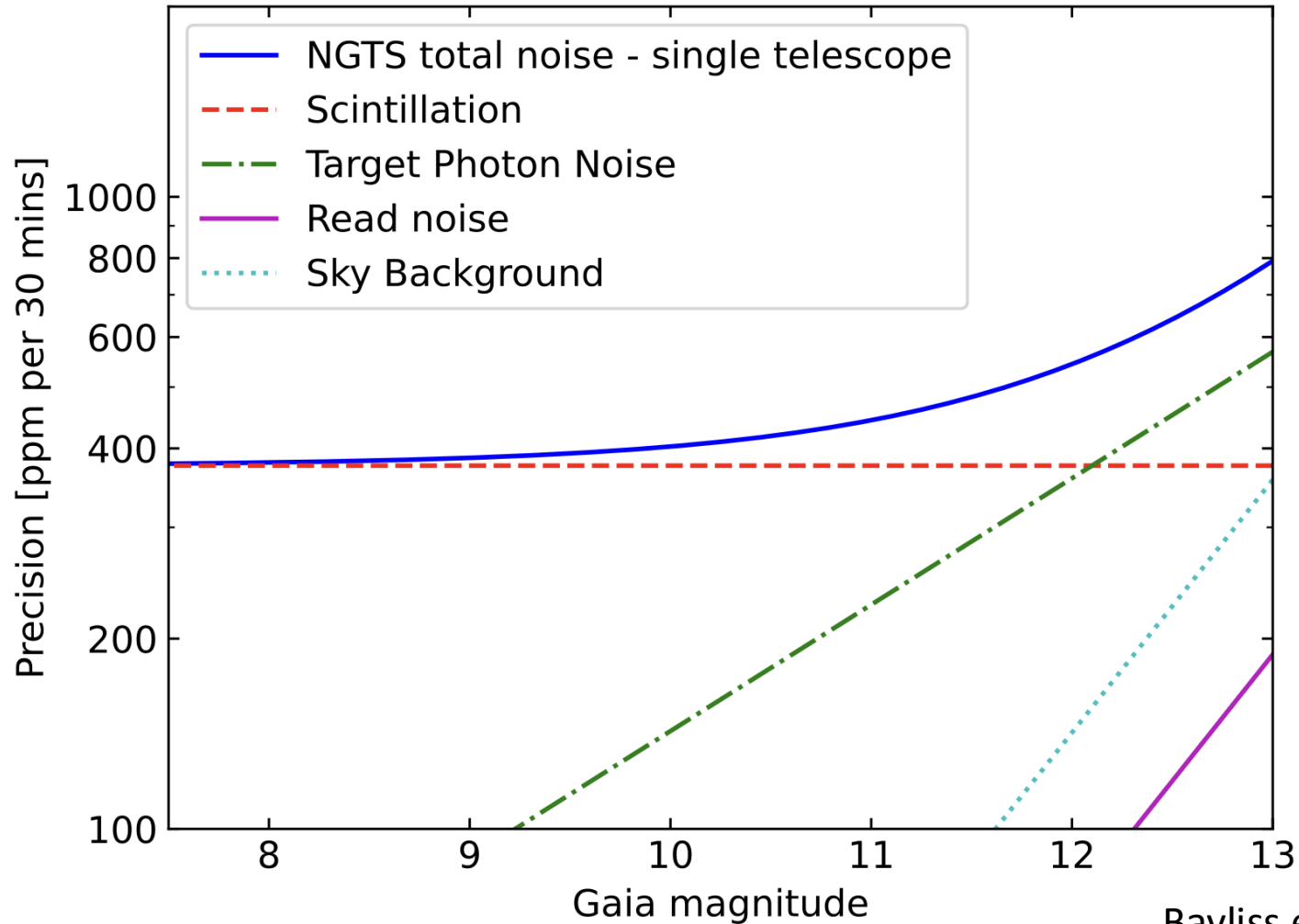
# NGTS CCD Camera Details



Description	iKon-L 936 (CCD)
Active pixels	2048 × 2048
Detector size	27.6×27.6 mm
Pixel size	13.5×13.5 μm
Read Noise	12 e <sup>-</sup> in fast readout mode (1.5 s)
Dark Current	0.006 e <sup>-</sup> /pix/s at -80°C
Sensor Type	e2v BR-DD
QE	peak QE > 95% at 800nm
Full Frame Readout Time	1.5 s
Operating temperature	-70°C
Shutter	45 mm programable shutter

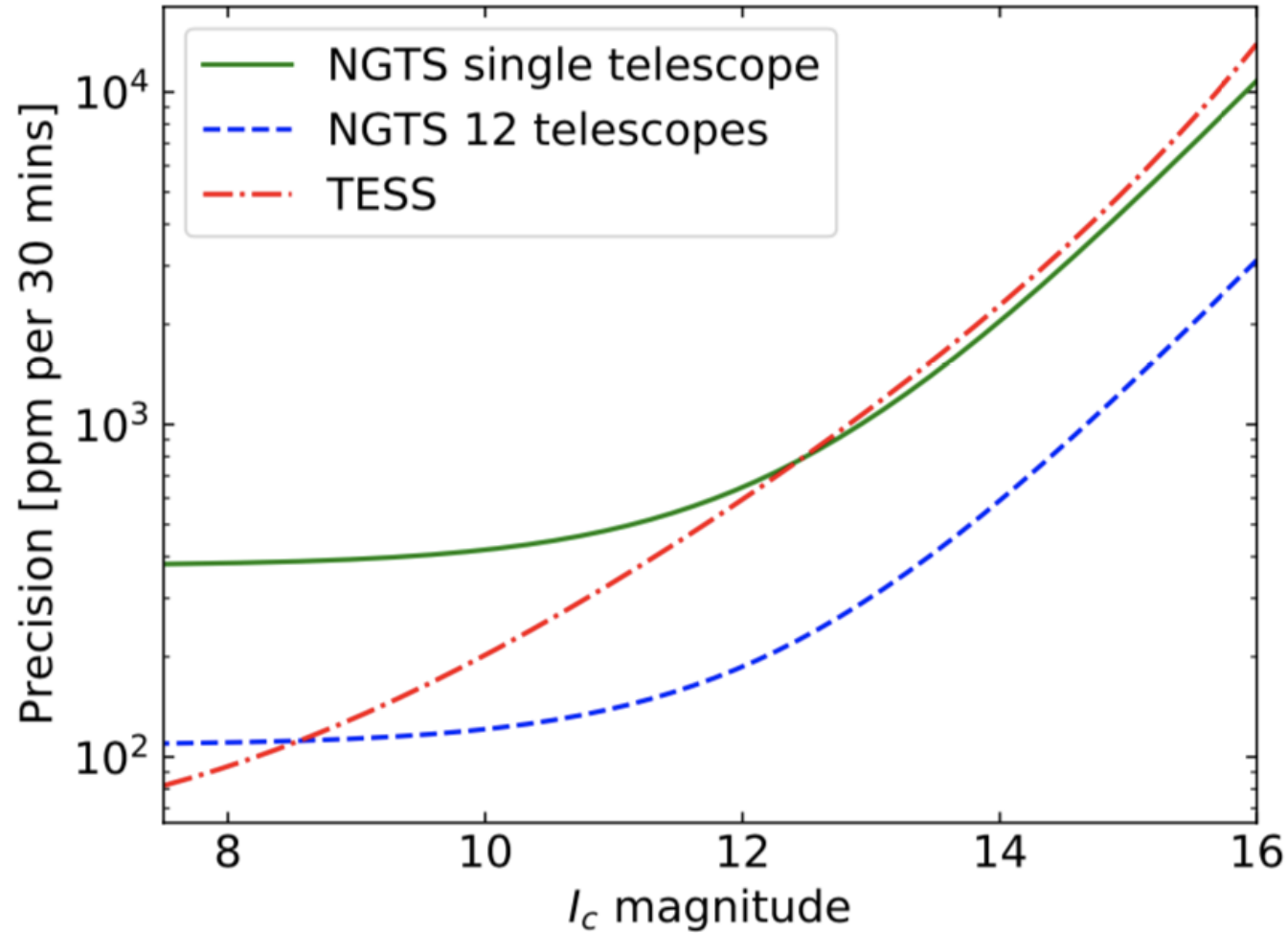
Bayliss et al., 2022

# Photometric Noise



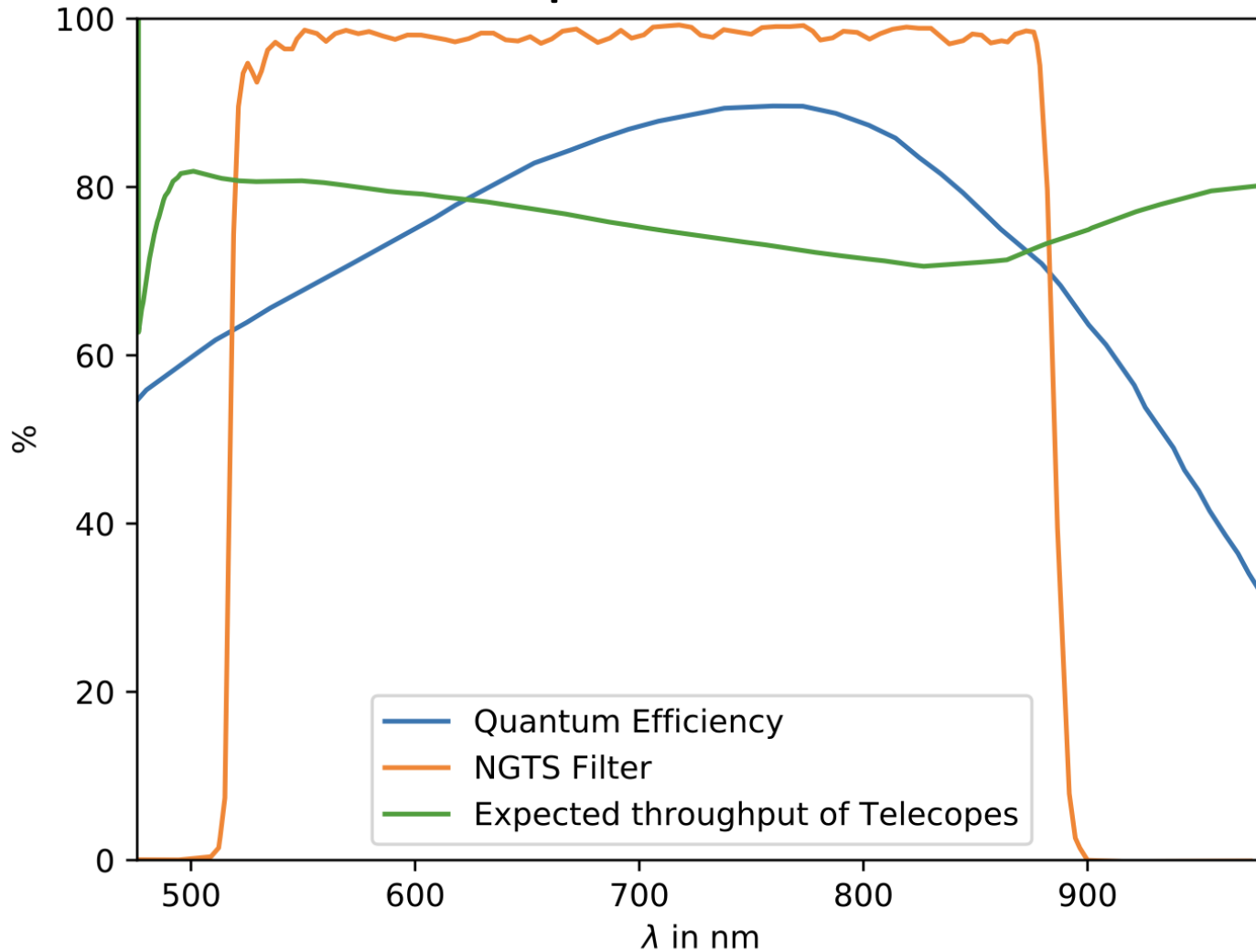
Bayliss et al., 2022

# Precision: NGTS v TESS



O'Brien et al., 2022

# NGTS passband

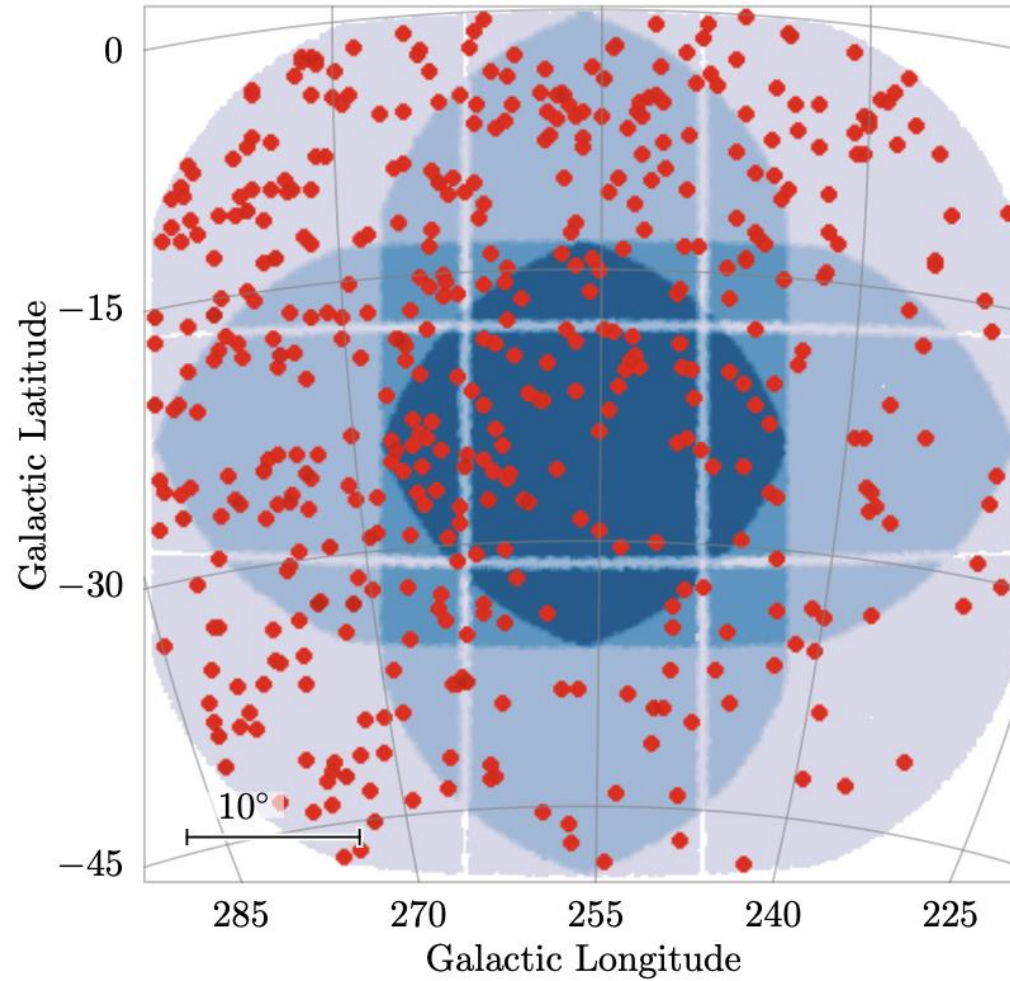


Wheatley et al., 2018

# NGTS Details

- Wheatley et al., 2018, Bayliss et al., 2022, Bryant et al., 2020, O'Brien et al., 2022.
- NGTS multi-telescope suited to bright stars ( $T < 10$ ) that are difficult on 1m telescopes.
- Single telescope can get to 1000 ppm per  $\frac{1}{2}$  hour.
- Multi-telescope can get to 150 ppm per  $\frac{1}{2}$  hour.
- Plate scale is 5 arcsec/pix.

# TOI Candidates in PLATO LOPS2



Eschen et al., 2024