PHAROS Conference 2020: The multi-messenger physics and astrophysics of neutron stars



Contribution ID: 35

Type: Oral Presentation

Vortex dynamics in neutron stars

Thursday 2 April 2020 15:15 (15 minutes)

In the superfluid interior of a neutron star the presence of quantized vortex lines defines an intermediate scale (in between the microscopic fermi-scale and the centimeter-scale) ranging from the radius of a vortex core to the typical separation between vortices. This complicates the hydrodynamic description of a neutron star interior. A classical treatment of a vortex moving through the lattice of nuclear impurities in the crust can be achieved by means of the vortex-filament model. Understanding the complex dynamics of vortices by means of vortex-filament simulations can deepen our understanding of superfluidity-related phenomena in neutron stars, like pulsar glitches.

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Track Classification: Pulsar glitches and superfluidity