

Proxy $SU(3)$ for Nuclear Structure

Nuclear Theory Group

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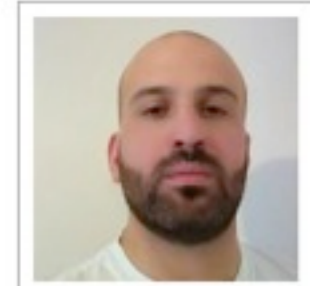


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Who are we?

Elliott SU(3) Symmetry 1957

Collective motion in the nuclear shell model

I. Classification schemes for states of mixed configurations

BY J. P. ELLIOTT

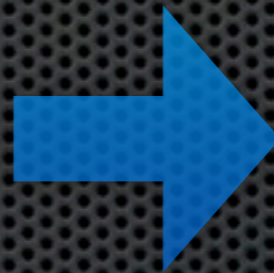
Atomic Energy Research Establishment, Harwell, Didcot, Berkshire

(Communicated by Sir John Cockcroft, F.R.S.—Received 12 December 1957)

The first algebraic Nuclear Model with microscopic justification

Elliott's Breakthrough

Shell Model
Space



All the
collective
observables

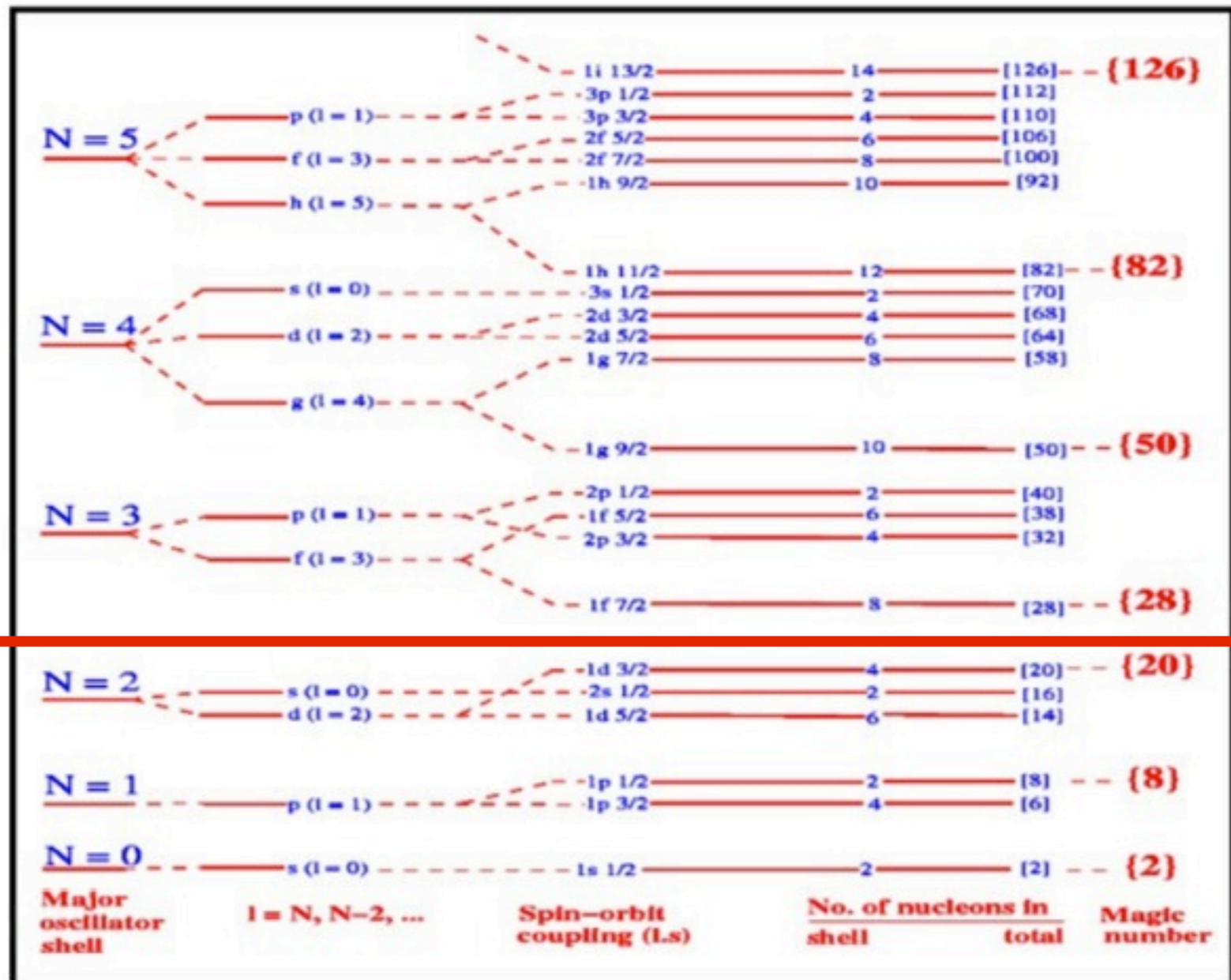
The proxy-SU(3) project

Persons involved: D. Bonatsos, I. E.
Assimakis, N. Minkov, A. Martinou, R. F.
Casten, R. B. Cakirli, K. Blaum

SU(3) Symmetry

Proxy
SU(3)

Elliott
SU(3)



Orbitals differing per 1 quantum in the z axis

Emergent Collectivity in Nuclei and Enhanced Proton-Neutron Interactions

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(Dated: October 30, 2013)

Phys. Rev. C 2013

The replacement of orbitals

Table 2: Nilsson orbitals for the nucleons between 20 and 50. Due to the ls interaction the $1f_{7/2}$ orbital is excluded from the shell under discussion, while the $1g_{9/2}$ orbital is included. Thus the harmonic oscillator shell 20-40, becomes the nuclear shell 28-50. The 28-50 shell consists of some orbitals with $\mathcal{N} = 3$ and some orbitals with $\mathcal{N} = 4$ quanta [14]. Therefore the Elliott SU(3) symmetry is not valid in the 28-50 shell.

Shell model	$K[\mathcal{N}n_z\Lambda]$	$ n_z r s \Sigma\rangle$
$1g_{9/2}$	9/2[404]	040+
	7/2[413]	130+
	5/2[422]	220+
	3/2[431]	310+
	1/2[440]	400+
$2p_{1/2}$	1/2[301]	021-
$1f_{5/2}$	5/2[303]	030-
	3/2[301]	021+
	1/2[310]	111+
$2p_{3/2}$	3/2[312]	120-
	1/2[321]	210-
$1f_{7/2}$	7/2[303]	030+
	5/2[312]	120+
	3/2[321]	210+
	1/2[330]	300+

Edited volume on “Nuclear Structure Physics”, to be published by CRC Press/Taylor and Francis Group, USA.

1st Proxy SU(3) Publication

PHYSICAL REVIEW C 95, 064325 (2017)

Proxy-SU(3) symmetry in heavy deformed nuclei

Dennis Bonatsos,¹ I. E. Assimakis,¹ N. Minkov,² Andriana Martinou,¹ R. B. Cakirli,³ R. F. Casten,^{4,5} and K. Blaum⁶

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(Received 17 January 2017; revised manuscript received 2 June 2017; published 27 June 2017)

Background: Microscopic calculations of heavy nuclei face considerable difficulties due to the sizes of the matrices that need to be solved. Various approximation schemes have been invoked, for example by truncating the spaces, imposing seniority limits, or appealing to various symmetry schemes such as pseudo-SU(3). This paper proposes a new symmetry scheme also based on SU(3). This proxy-SU(3) can be applied to well-deformed nuclei, is simple to use, and can yield analytic predictions.

Purpose: To present the new scheme and its microscopic motivation, and to test it using a Nilsson model calculation with the original shell model orbits and with the new proxy set.

Method: We invoke an approximate, analytic, treatment of the Nilsson model, that allows the above vetting and yet is also transparent in understanding the approximations involved in the new proxy-SU(3).

Results: It is found that the new scheme yields a Nilsson diagram for well-deformed nuclei that is very close to the original Nilsson diagram. The specific levels of approximation in the new scheme are also shown, for each major shell.

Conclusions: The new proxy-SU(3) scheme is a good approximation to the full set of orbits in a major shell. Being able to replace a complex shell model calculation with a symmetry-based description now opens up the possibility to predict many properties of nuclei analytically and often in a parameter-free way. The new scheme works best for heavier nuclei, precisely where full microscopic calculations are most challenged. Some cases in which the new scheme can be used, often analytically, to make specific predictions, are shown in a subsequent paper.

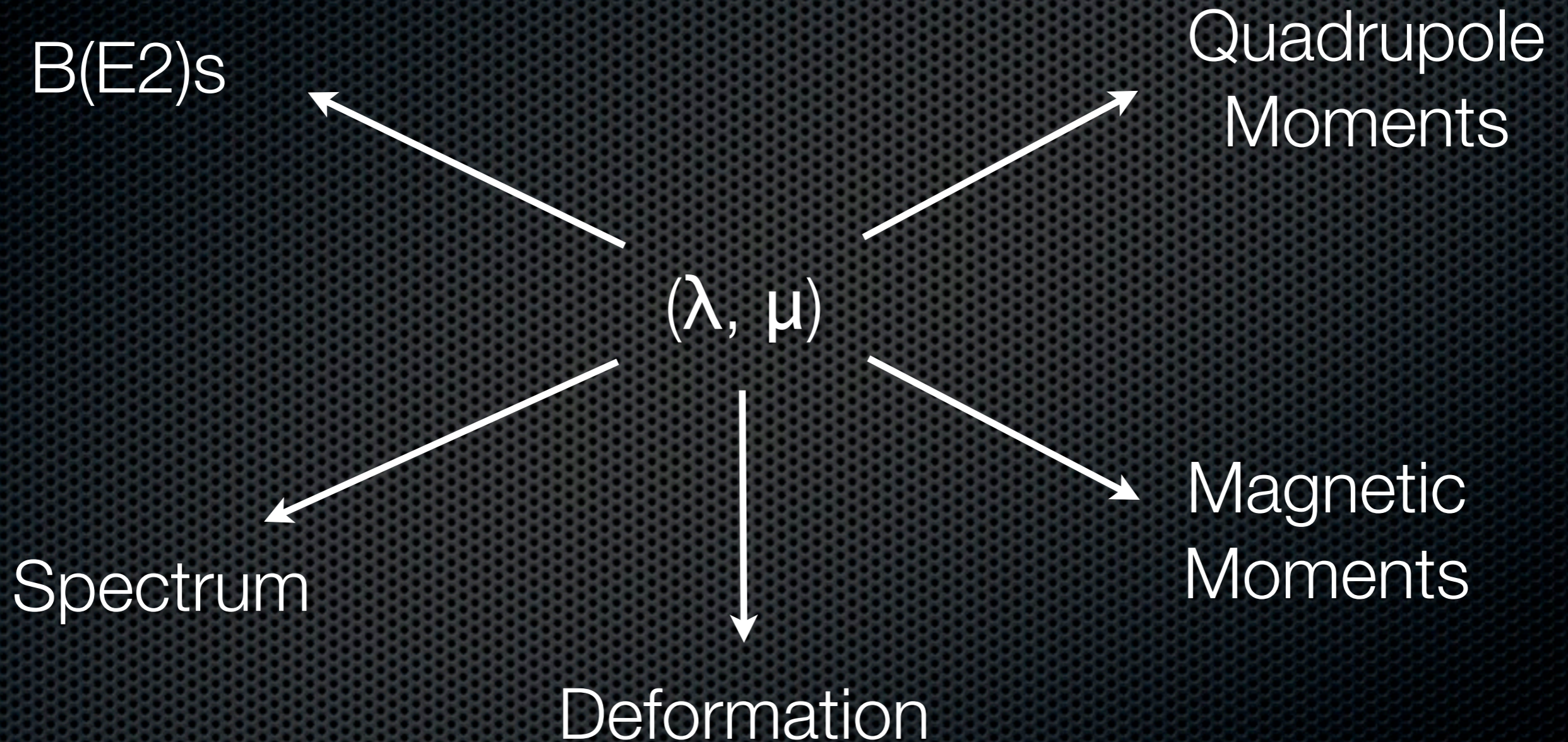
DOI: [10.1103/PhysRevC.95.064325](https://doi.org/10.1103/PhysRevC.95.064325)

The effect on the Hamiltonian

TABLE V. Same as Table III, but for the H matrix elements with $\epsilon = 0.3$ for Nilsson orbitals in the 50–82 proton shell (upper part) and in the sdg proton shell (lower part). Note again that these calculations use the asymptotic wave functions defined in Sec. III. The deformation comes in only through the linear dependence on epsilon of the diagonal matrix elements of H_{osc} . See Sec. IV A for further discussion.

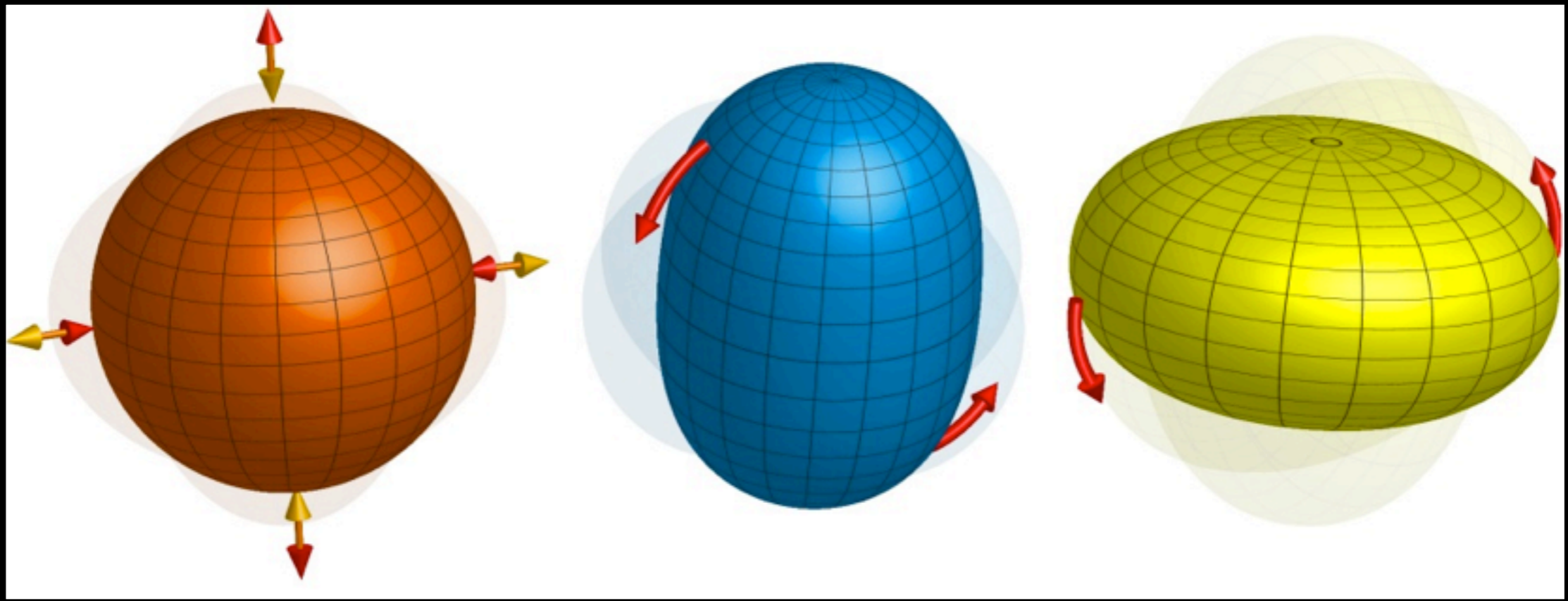
	$\frac{1}{2}[400]$	$\frac{1}{2}[411]$	$\frac{3}{2}[402]$	$\frac{1}{2}[420]$	$\frac{3}{2}[411]$	$\frac{5}{2}[402]$	$\frac{1}{2}[431]$	$\frac{3}{2}[422]$	$\frac{5}{2}[413]$	$\frac{7}{2}[404]$	$\frac{1}{2}[550]$	$\frac{3}{2}[541]$	$\frac{5}{2}[532]$	$\frac{7}{2}[523]$	$\frac{9}{2}[514]$	$\frac{11}{2}[505]$
1/2[400]	6.28	-0.13	0	0.22	0	0	0	0	0	0	0	0	0	0	0	0
1/2[411]		5.74	0	0.18	0	0	0.27	0	0	0	0	0	0	0	0	0
3/2[402]			6.26	0	0.16	0	0	0.19	0	0	0	0	0	0	0	0
1/2[420]				5.30	0	0	-0.16	0	0	0	0	0	0	0	0	0
3/2[411]					5.61	0	0	-0.13	0	0	0	0	0	0	0	0
5/2[402]						6.00	0	0	-0.09	0	0	0	0	0	0	0
1/2[431]							5.06	0	0	0	0	0	0	0	0	0
3/2[422]								5.27	0	0	0	0	0	0	0	0
5/2[413]									5.56	0	0	0	0	0	0	0
7/2[404]										5.93	0	0	0	0	0	0
1/2[550]											5.88	0	0	0	0	0
3/2[541]												5.81	0	0	0	0
5/2[532]													5.82	0	0	0
7/2[523]														5.90	0	0
9/2[514]															6.06	0
11/2[505]																6.30
	$\frac{1}{2}[400]$	$\frac{1}{2}[411]$	$\frac{3}{2}[402]$	$\frac{1}{2}[420]$	$\frac{3}{2}[411]$	$\frac{5}{2}[402]$	$\frac{1}{2}[431]$	$\frac{3}{2}[422]$	$\frac{5}{2}[413]$	$\frac{7}{2}[404]$	$\frac{1}{2}[440]$	$\frac{3}{2}[431]$	$\frac{5}{2}[422]$	$\frac{7}{2}[413]$	$\frac{9}{2}[404]$	
1/2[400]	6.28	-0.13	0	0.22	0	0	0	0	0	0	0	0	0	0	0	
1/2[411]		5.74	0	0.18	0	0	0.27	0	0	0	0	0	0	0	0	
3/2[402]			6.26	0	0.16	0	0	0.19	0	0	0	0	0	0	0	
1/2[420]				5.30	0	0	-0.16	0	0	0	0.27	0	0	0	0	
3/2[411]					5.61	0	0	-0.13	0	0	0	0.27	0	0	0	
5/2[402]						6.00	0	0	-0.09	0	0	0	0.19	0	0	
1/2[431]							5.06	0	0	0	0.18	0	0	0	0	
3/2[422]								5.27	0	0	0	0.22	0	0	0	
5/2[413]									5.56	0	0	0	0.22	0	0	
7/2[404]										5.93	0	0	0	0.18	0	
1/2[440]											5.73	0	0	0	0	
3/2[431]												5.74	0	0	0	
5/2[422]													5.82	0	0	
7/2[413]														5.98	0	
9/2[404]															6.22	

SU(3) Observables



Literature has been enriched by Pseudo-SU(3) publications

Nuclear Shape



spherical

prolate

oblate

Parameter free shape predictions

PHYSICAL REVIEW C 95, 064326 (2017)

Analytic predictions for nuclear shapes, prolate dominance, and the prolate-oblate shape transition in the proxy-SU(3) model

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R. F. Casten,^{4,5} and K. Blaum⁶

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³*Department of Physics, University of Istanbul, 34134 Istanbul, Turkey*

⁴*Wright Laboratory, Yale University, New Haven, Connecticut 06520, USA*

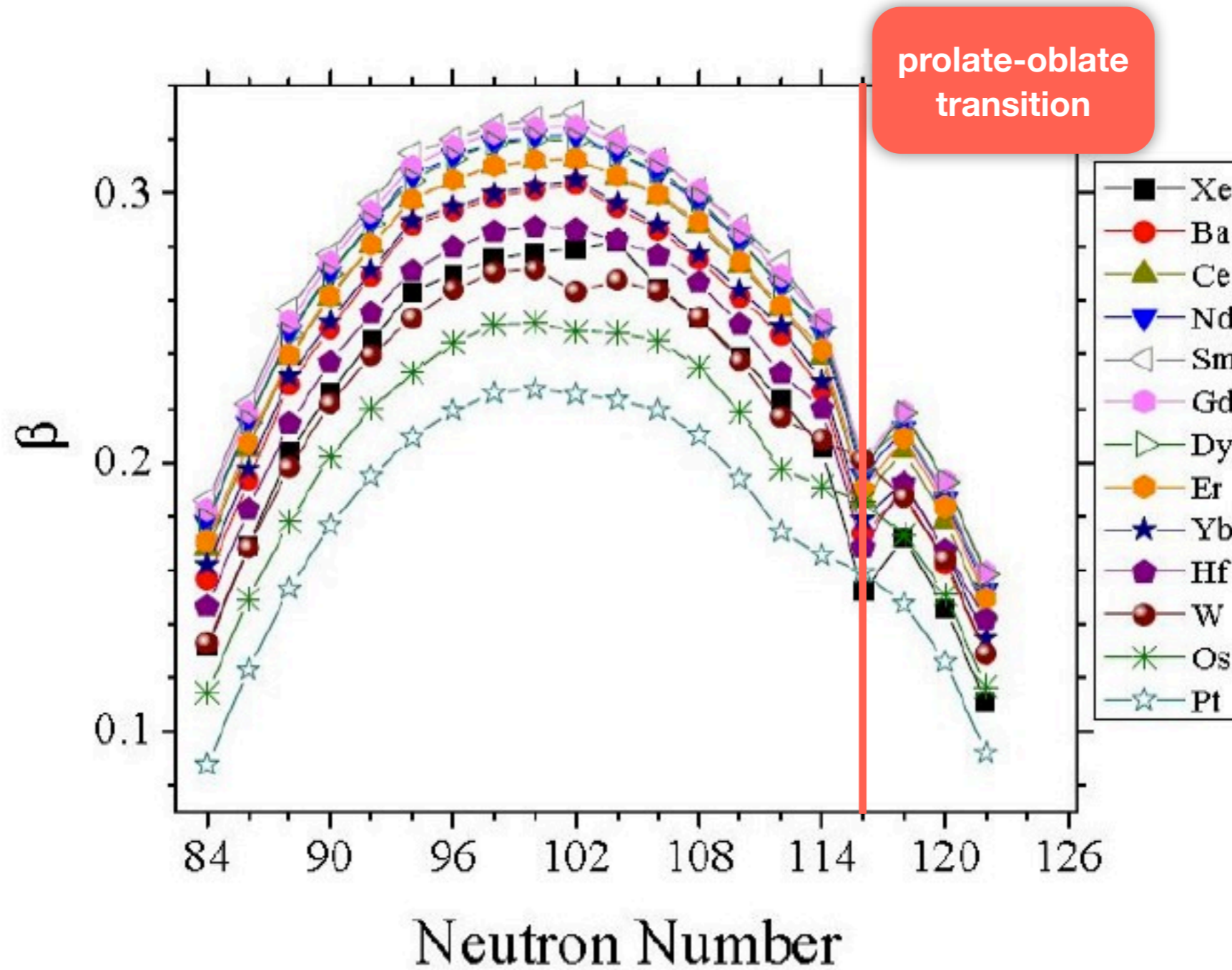
⁵*Facility for Rare Isotope Beams, 640 South Shaw Lane, Michigan State University, East Lansing, Michigan 48824, USA*

⁶*Max-Planck-Institut für Kernphysik, Saupfercheckweg 1, D-69117 Heidelberg, Germany*

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Using a new approximate analytic parameter-free proxy-SU(3) scheme, we make simple predictions of shape observables for deformed nuclei, namely γ and β deformation variables, the global feature of prolate dominance, and the locus of the prolate-oblate shape transition. The predictions are compared with empirical results.

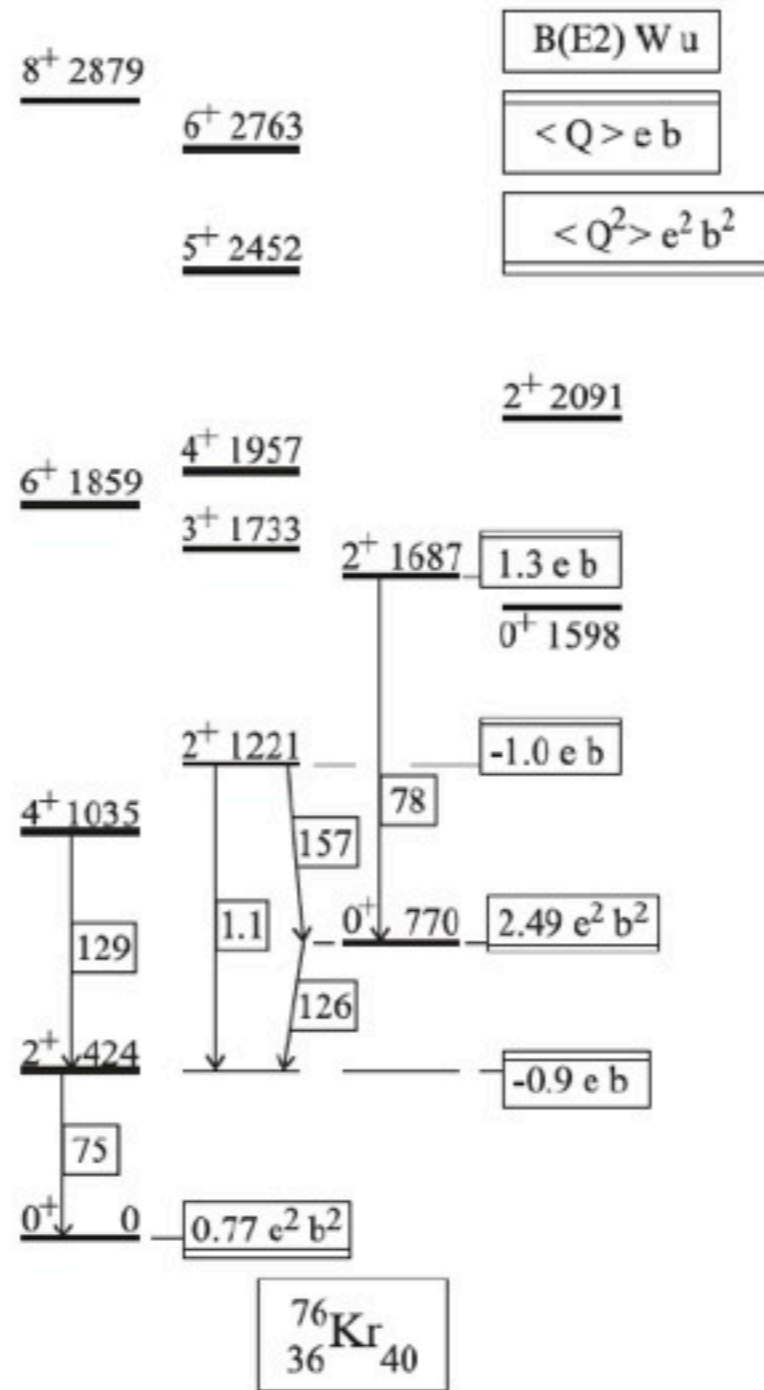
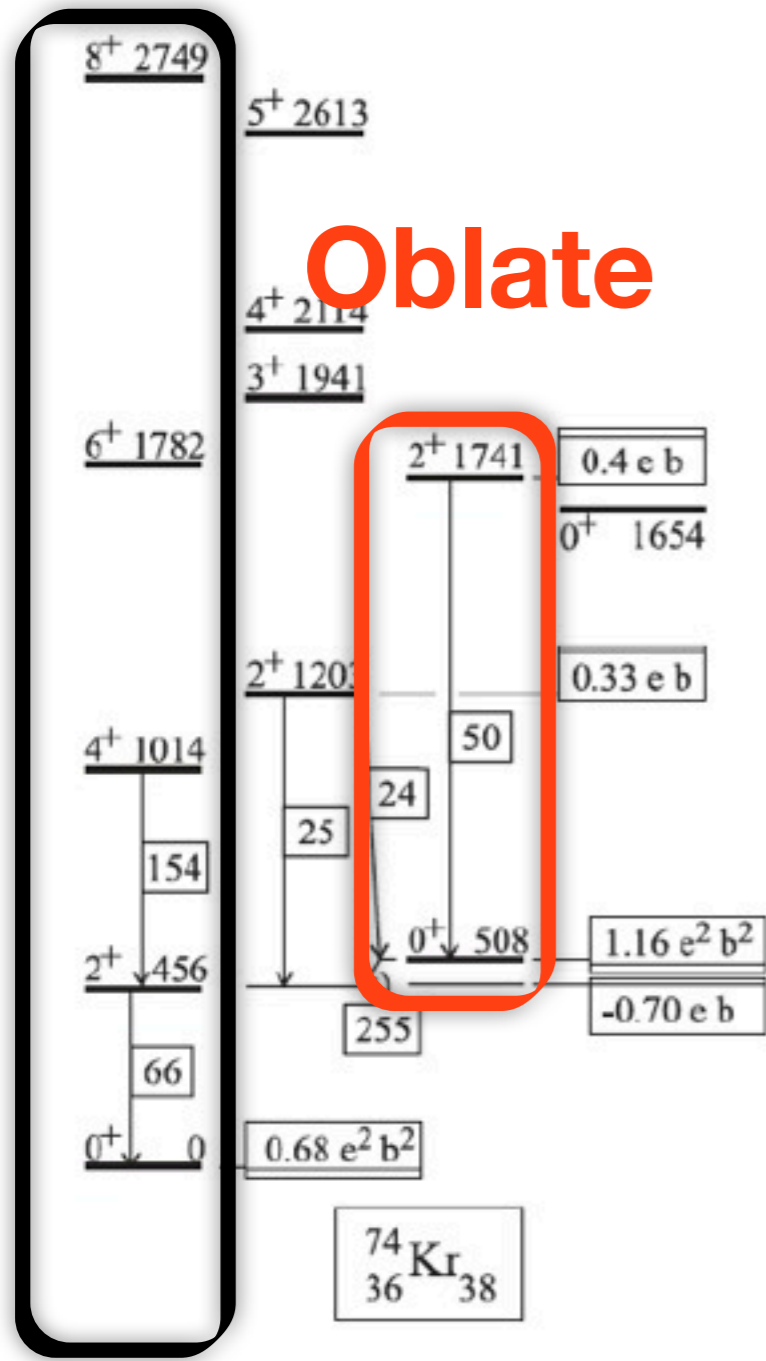
Prolate over oblate dominance



The shape coexistence project

Persons involved: A. Martinou, D. Bonatsos, I.E. Assimakis, S. Peroulis, S. Sarantopoulou, N. Minkov

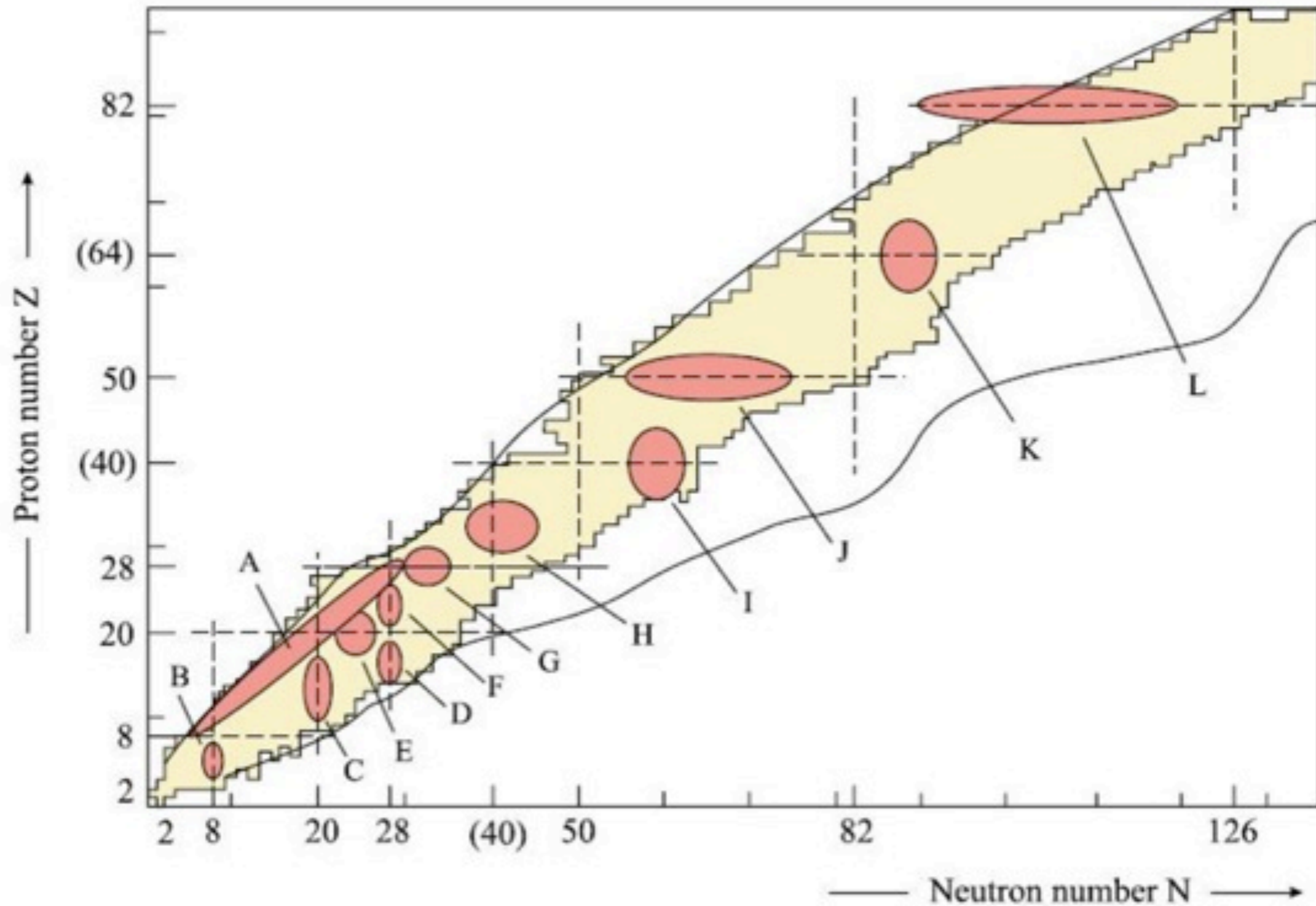
Prolate



Manifestation
of shape
coexistence.

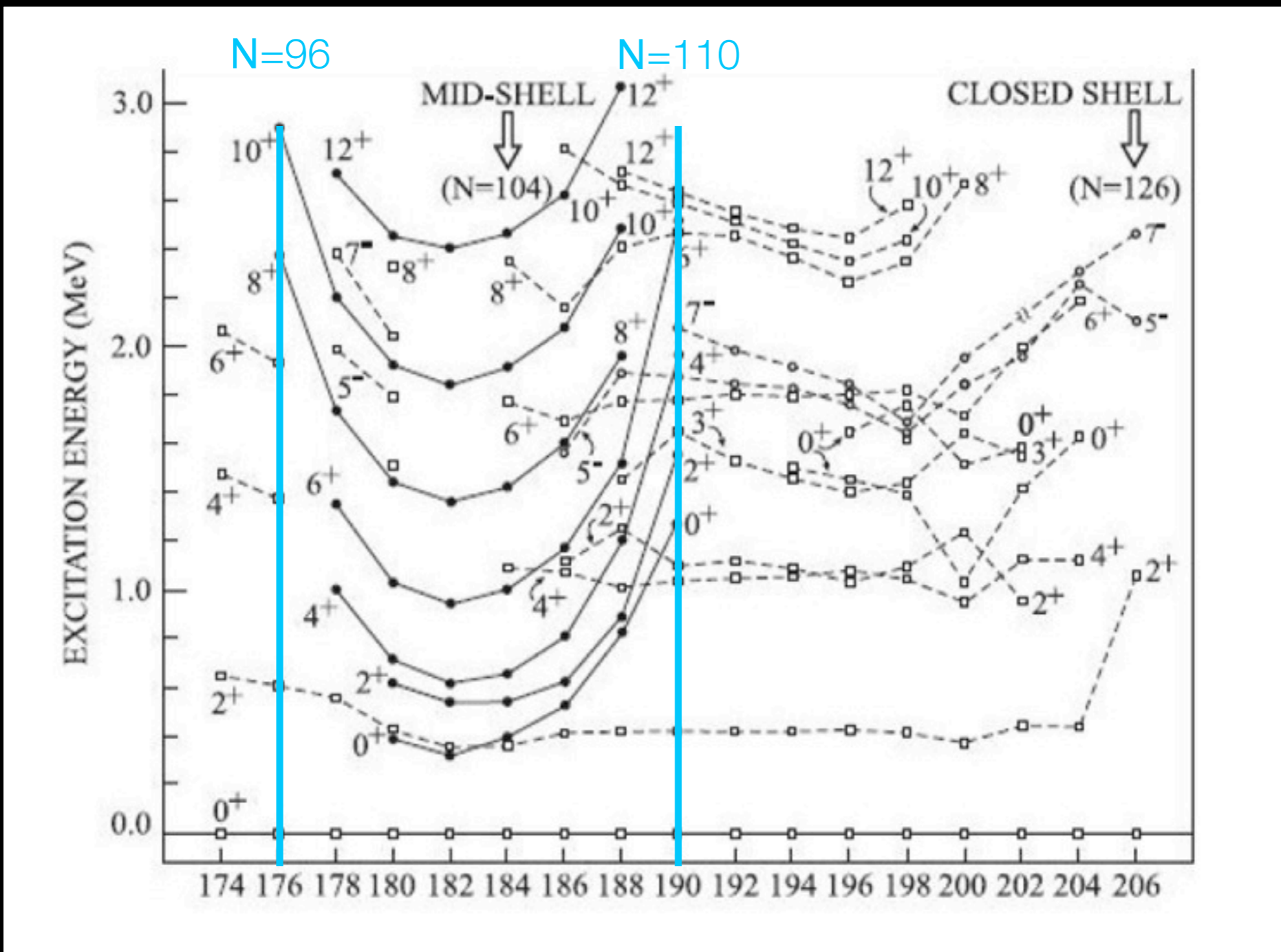
Clement et al. Phys. Rev. C 75, 054313, 2007.

Islands of coexistence



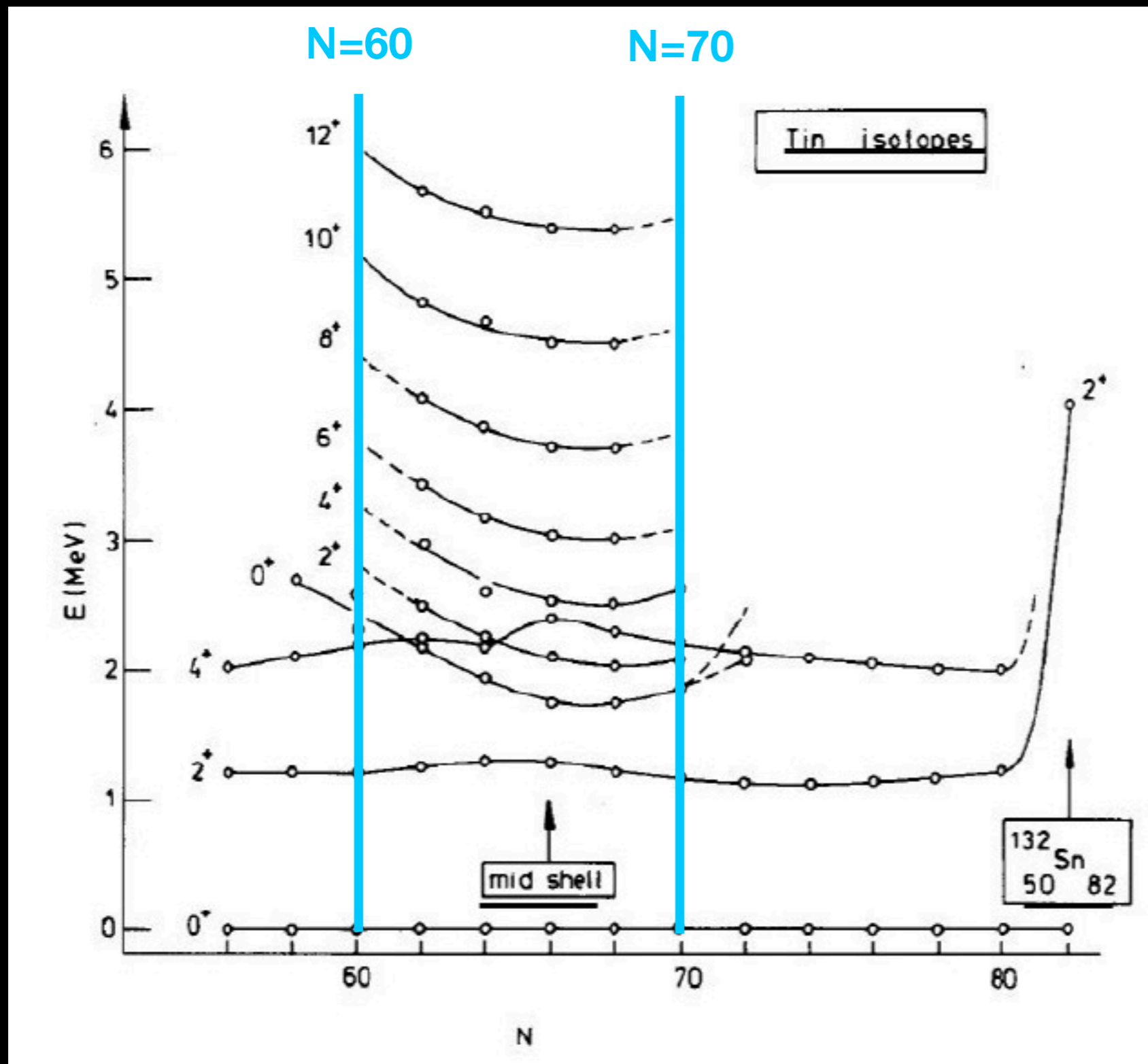
Kris Heyde and J. L. Wood, Rev.
Mod. Phys., Vol 83, 2011

Hg isotopes



Why?

Sn isotopes



Why?

J. L. Wood et al., Phys. Rep. 215, Nos 3 & 4, 101-201, 1992.

The collective quadrupole interaction

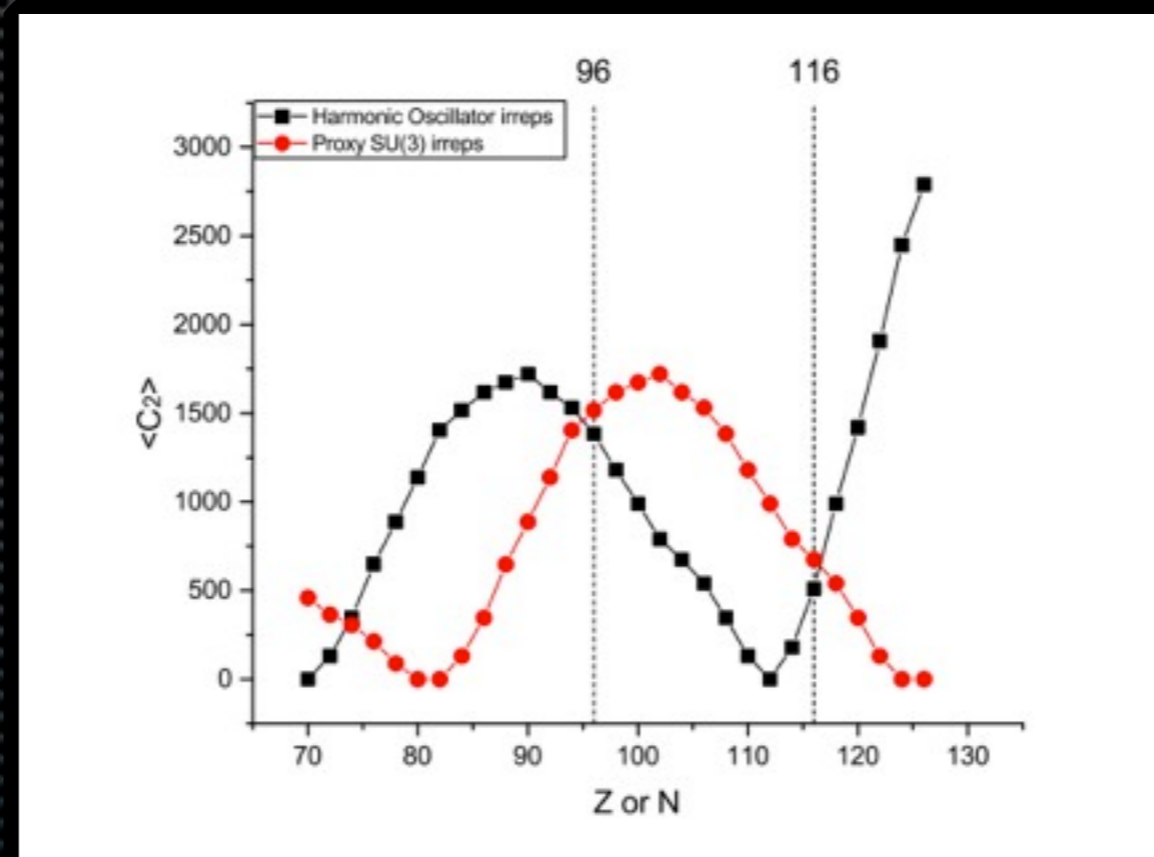
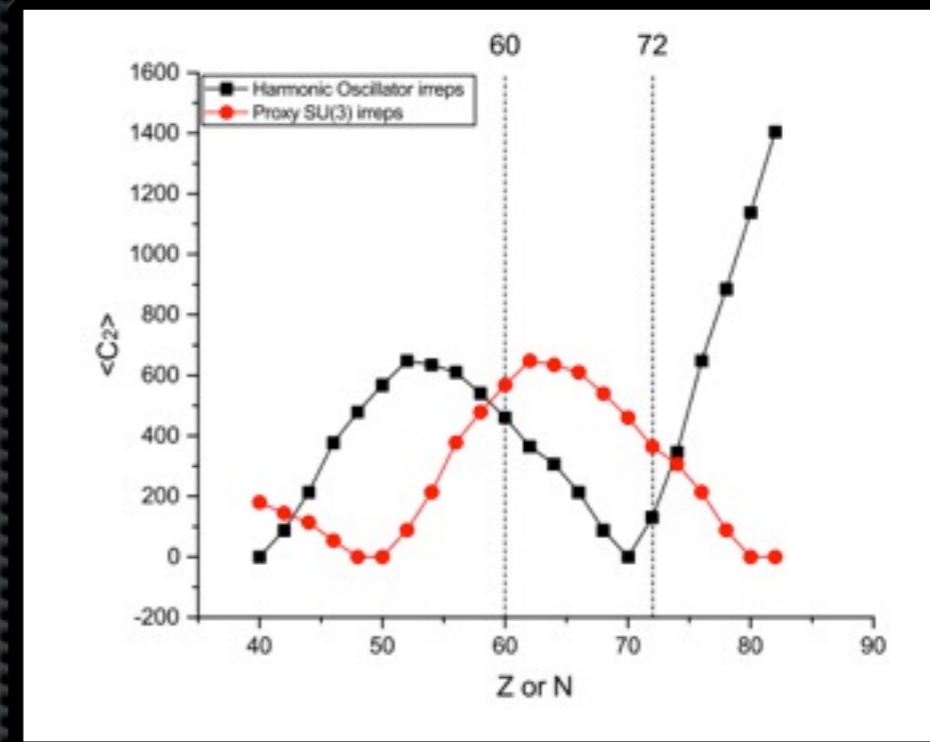
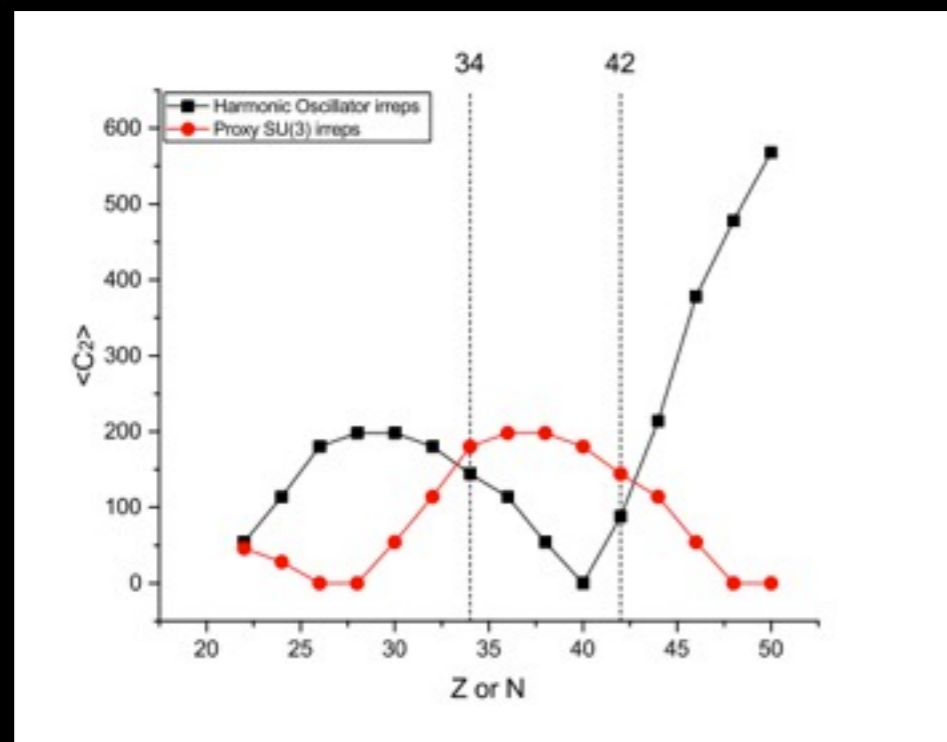
$$Q \cdot Q = 4C_2 - 3L^2,$$
$$C_2 = \lambda^2 + \mu^2 + \lambda\mu + 3(\lambda + \mu),$$
$$\beta^2 = \frac{4\pi}{5(A\bar{r}^2)^2}(C_2 + 3)$$

O. Castanos et al., Z. Phys. A, 329, 33-43, (1988).

$$\gamma = \arctan \left(\frac{\sqrt{3}(\mu + 1)}{2\lambda + \mu + 3} \right)$$

- (λ, μ) are the quantum numbers of SU(3). They depend on the choice of magic numbers. β is the deformation and γ is the angle which distinguishes prolate from oblate.

The islands of coexistence



Invitation for
a Review
Article in
EPJ A

Experiment in ^{180}Hf

IFIN-HH
Proposal for the TANDEM accelerator
experiments 2019

Collectivity and shape coexistence in ^{180}Hf

T.J. Mertzimekis¹, A. Khaliel¹, A. Zyriliou¹, P. Vasiliou¹, K. Bosbotinis¹, S. Pelonis¹,
I. Asimakis², M. Axiotis², D. Bonatsos², A. Lagoyannis², A. Martinou², S. Peroulis^{1,2},
S. Sarantopoulou²,

[...]

C. Nita⁴, N. Mărginean⁴, C. Mihai⁴, D. Bucurescu⁴, Gh. Căta-Danil⁴, I. Căta-Danil⁴, D. Deleanu⁴,
D. Filipescu⁴, I. Gheorghe⁴, D.G. Ghiță⁴, T. Glodariu⁴, R. Lică⁴, R. Mărginean⁴,
A. Negreț⁴, T. Sava⁴, L. Stroe⁴, S. Toma⁴, R. Șuvăilă⁴, N.V. Zamfir⁴, C.A. Ur⁴

¹ *Department of Physics, University of Athens, Greece*

² *Institute of Nuclear and Particle Physics, NCSR "Demokritos", Greece*

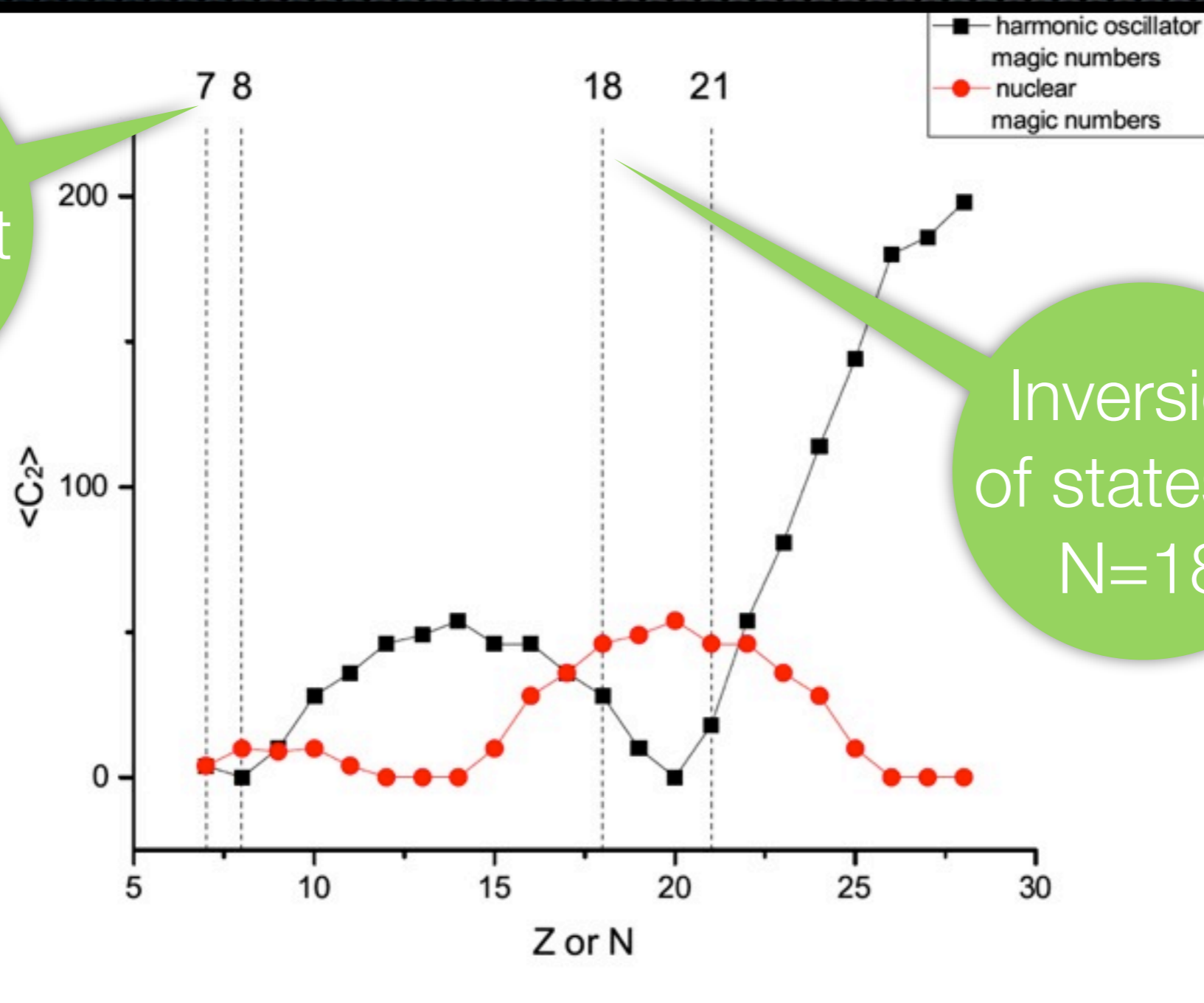
³ *OTHER*

⁴ *National Institute for Physics and Nuclear Engineering, Magurele, Romania*

Spokespersons: T.J. Mertzimekis
Contact Person: N. Mărginean

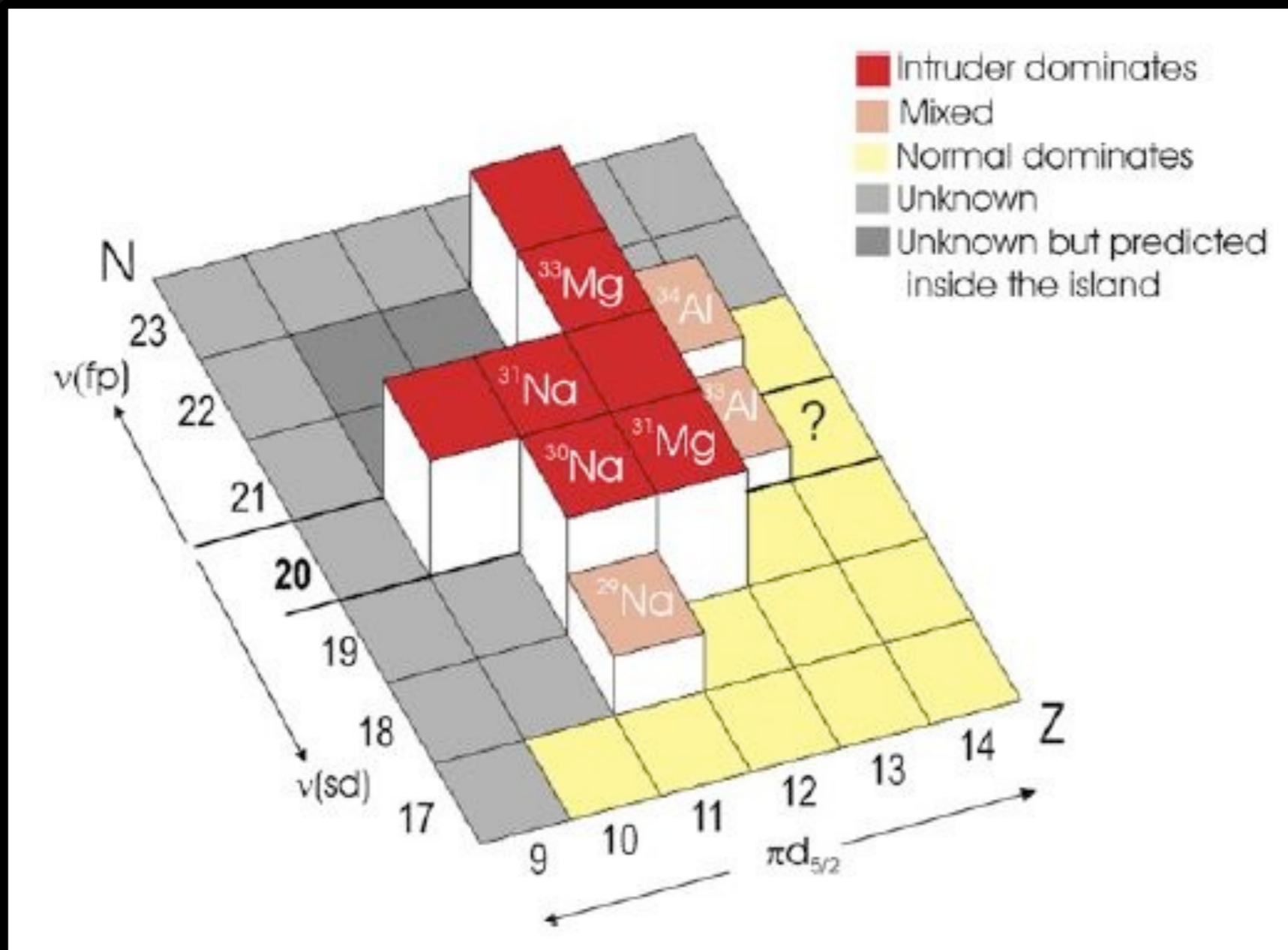
Shape coexistence=portal to the island of inversion

Inversion of states at $N=7$



Inversion of states at $N=18$

The Evidence

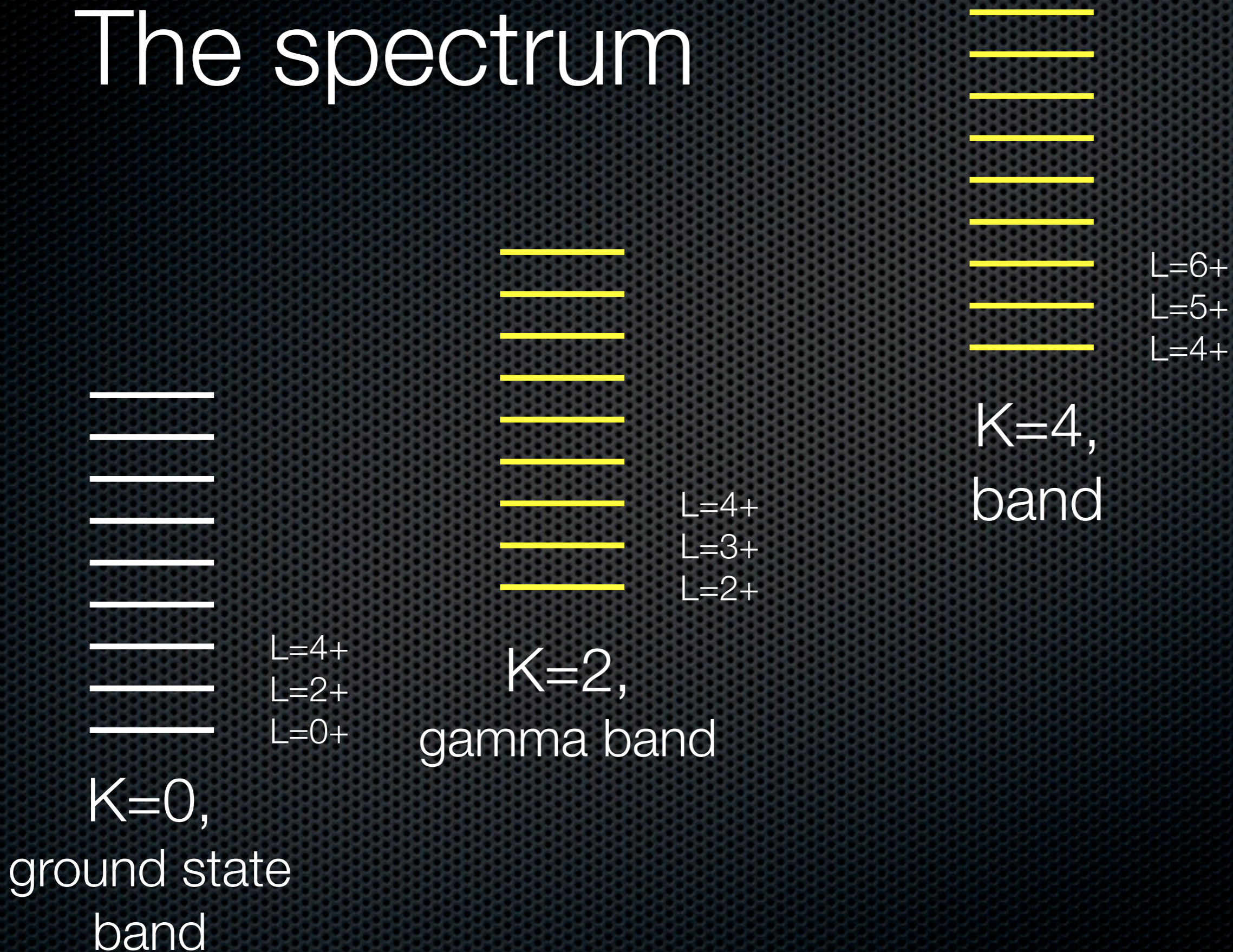


P. Himpe et al., Phys. Lett. B 658, 203-208, 2008.

The nuclear observables project

Persons involved: D. Bonatsos, S. Karambagia, R. F. Casten, N. Minkov, R. B. Cakirli, A. Martinou, I.E. Assimakis, S. Sarantopoulou, S. Peroulis

The spectrum



The SU(3) Hamiltonian

- $H = H_0 - \chi/2 QQ + c_1 \Omega$ or Λ

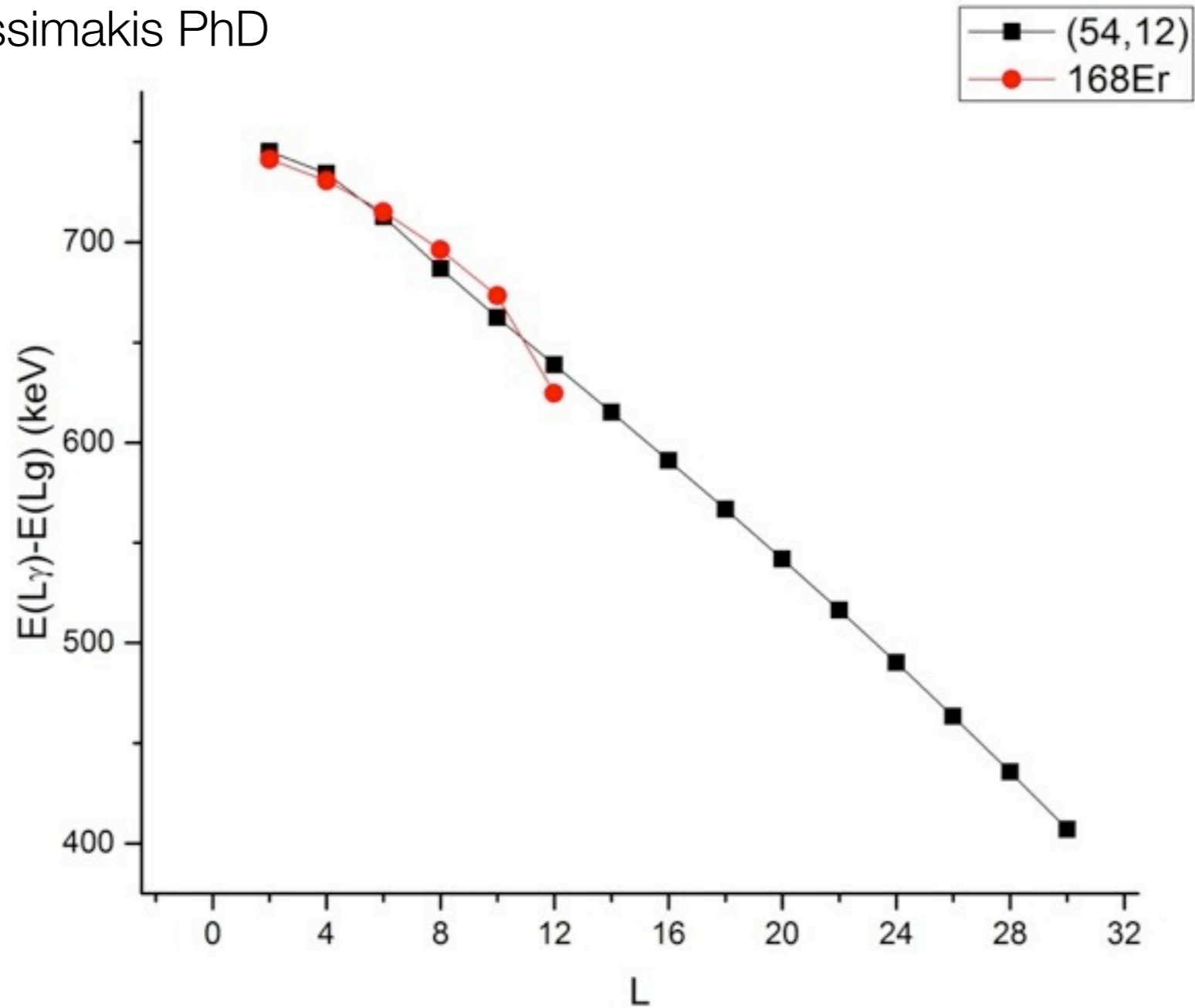
Mean field
potential

Quadrupole-
Quadrupole
interaction

Break the degeneracy
between the K bands

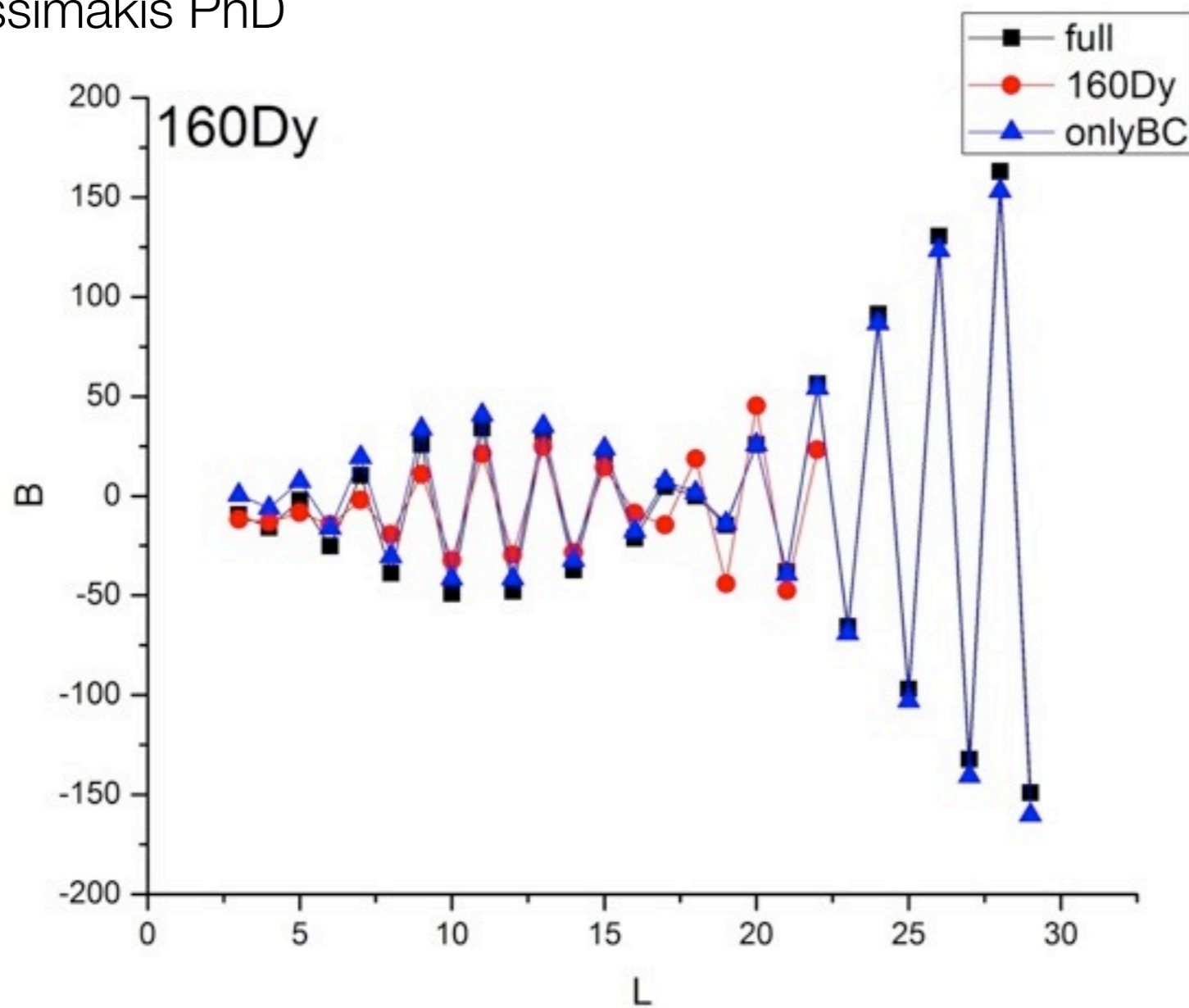
$$E(L)_{\text{gamma}} - E(L)_{\text{ground}}$$

Assimakis PhD

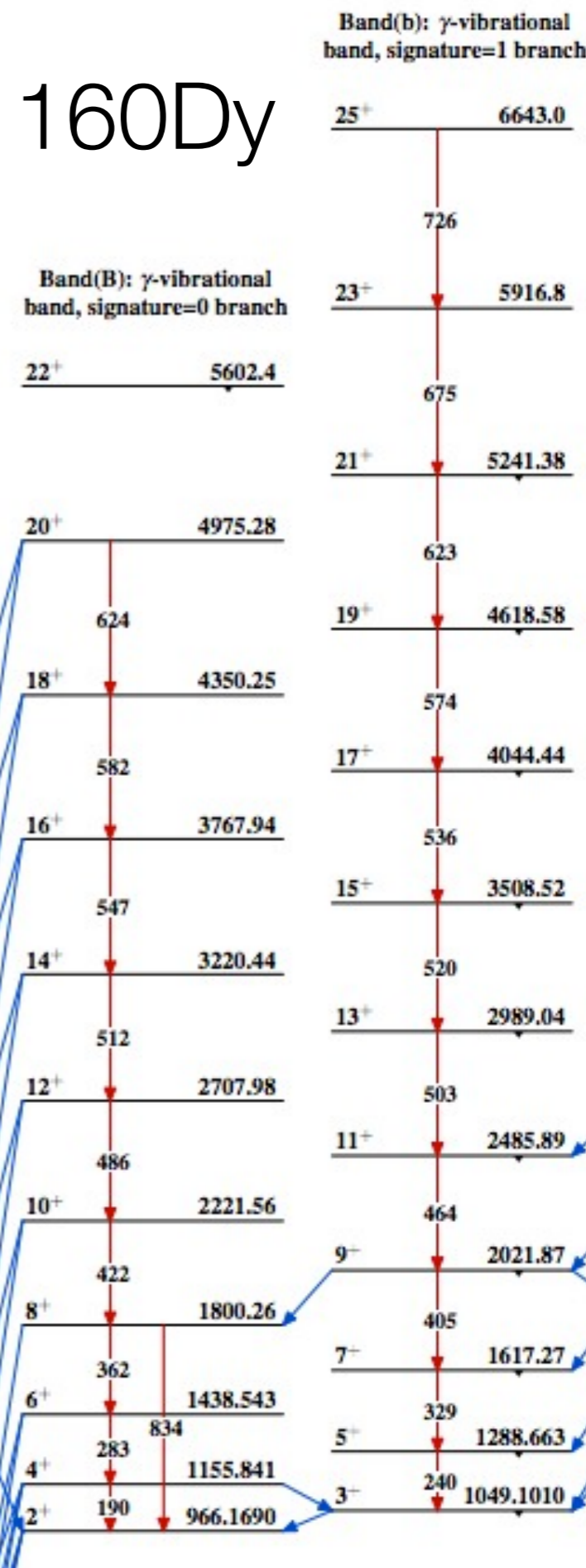


γ band staggering

Assimakis PhD



160Dy



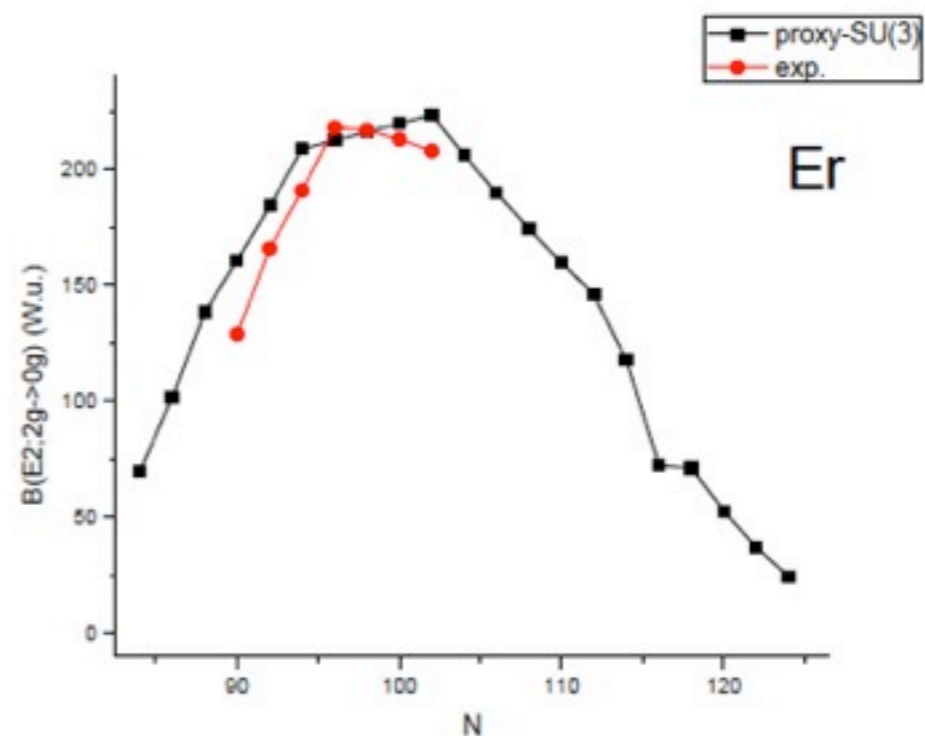
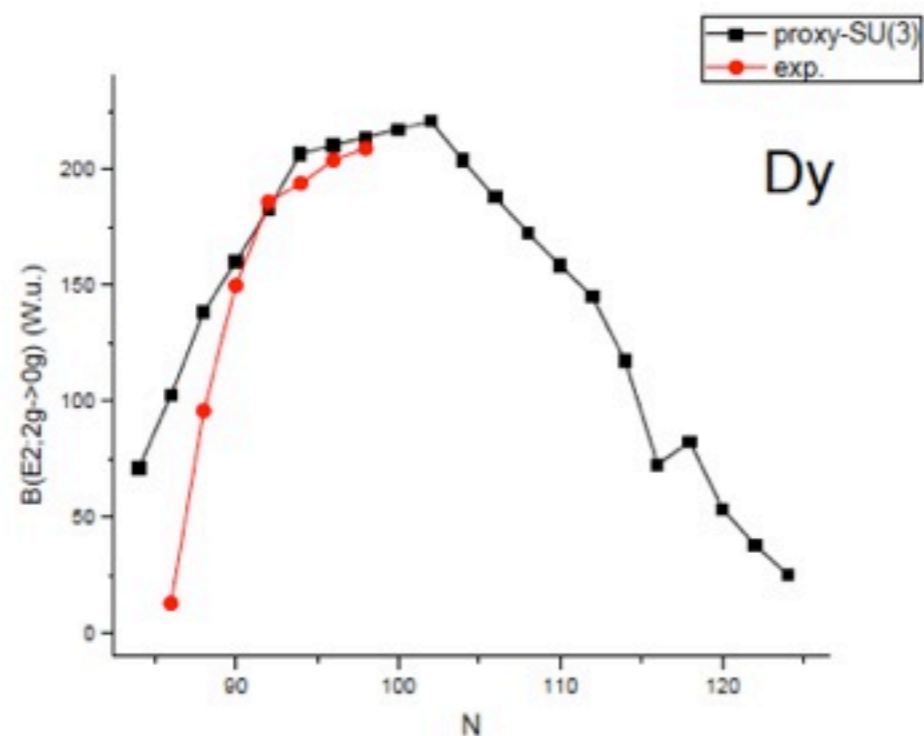
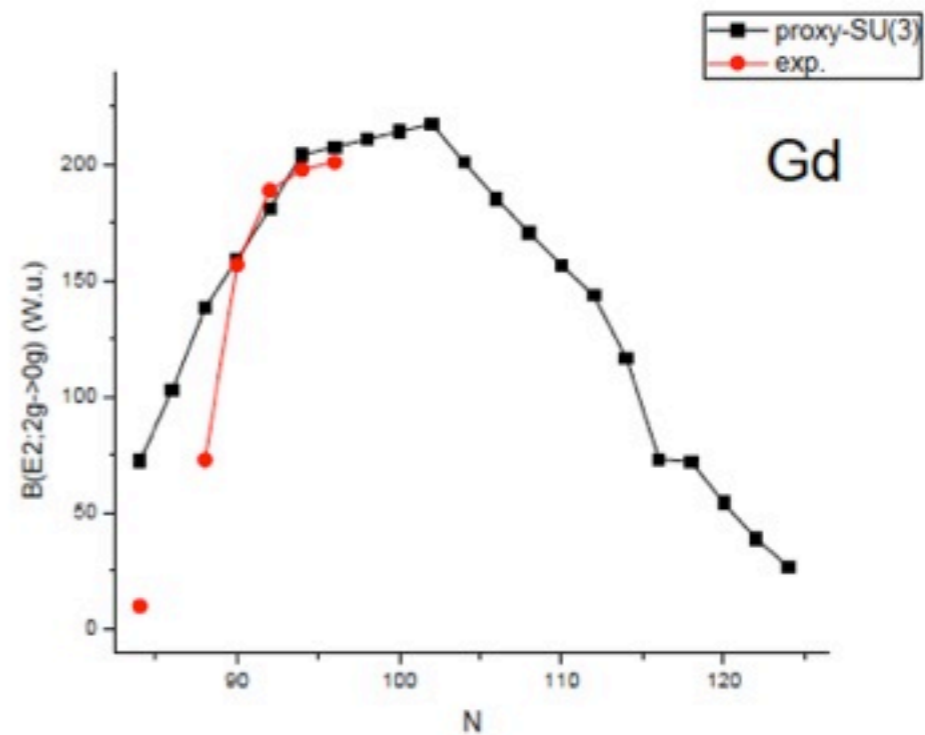
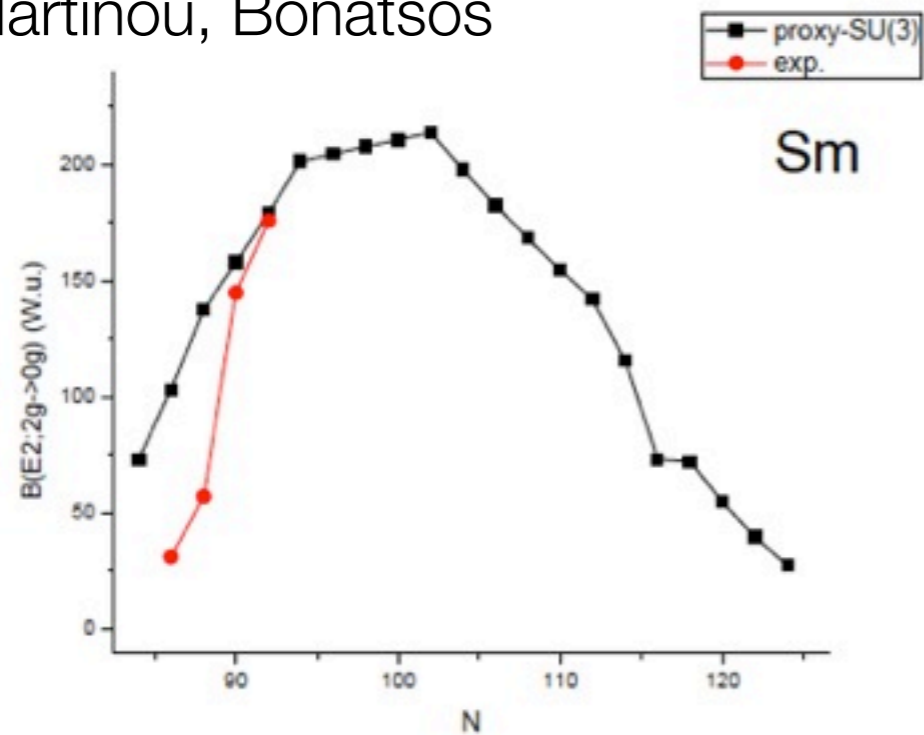
B(E2)s- Transition probabilities

$$B(E2; (\lambda, \mu)KL \rightarrow (\lambda, \mu)K'L') = 4C_2^{(\lambda, \mu)} \frac{2L' + 1}{2L + 1} (\langle (\lambda, \mu)KL; (1, 1)2 || (\lambda, \mu)K'L' \rangle)^2.$$

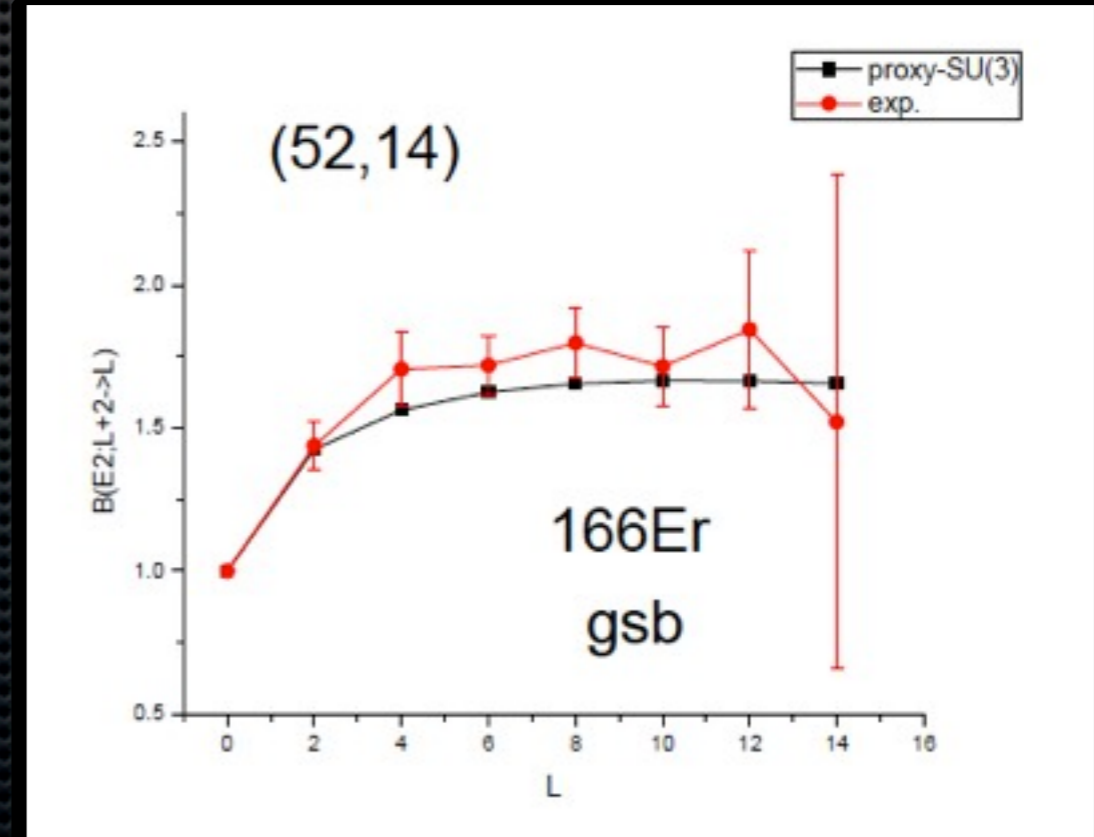
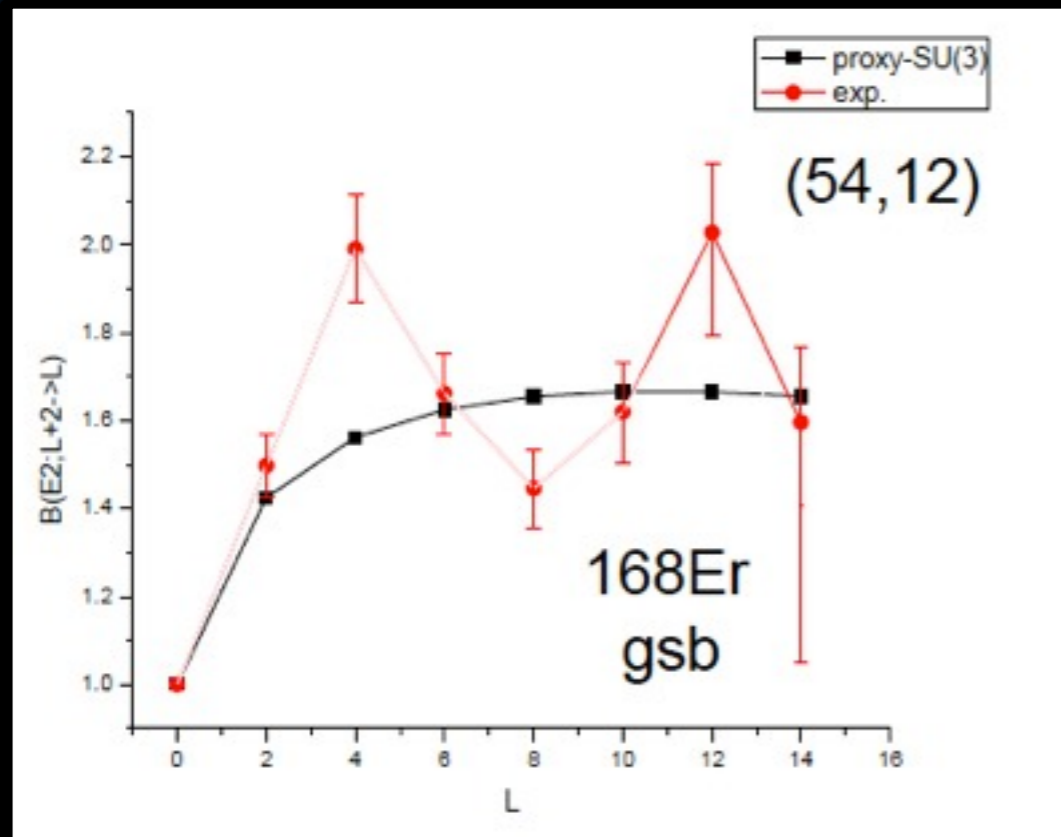
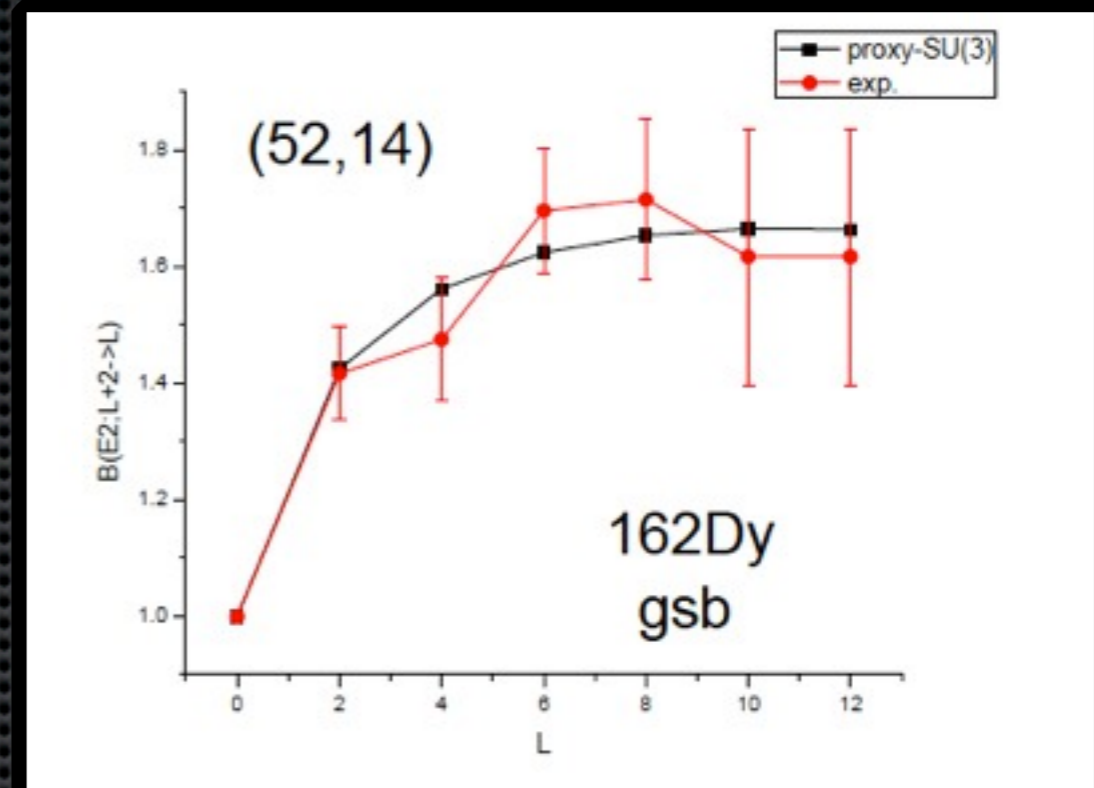
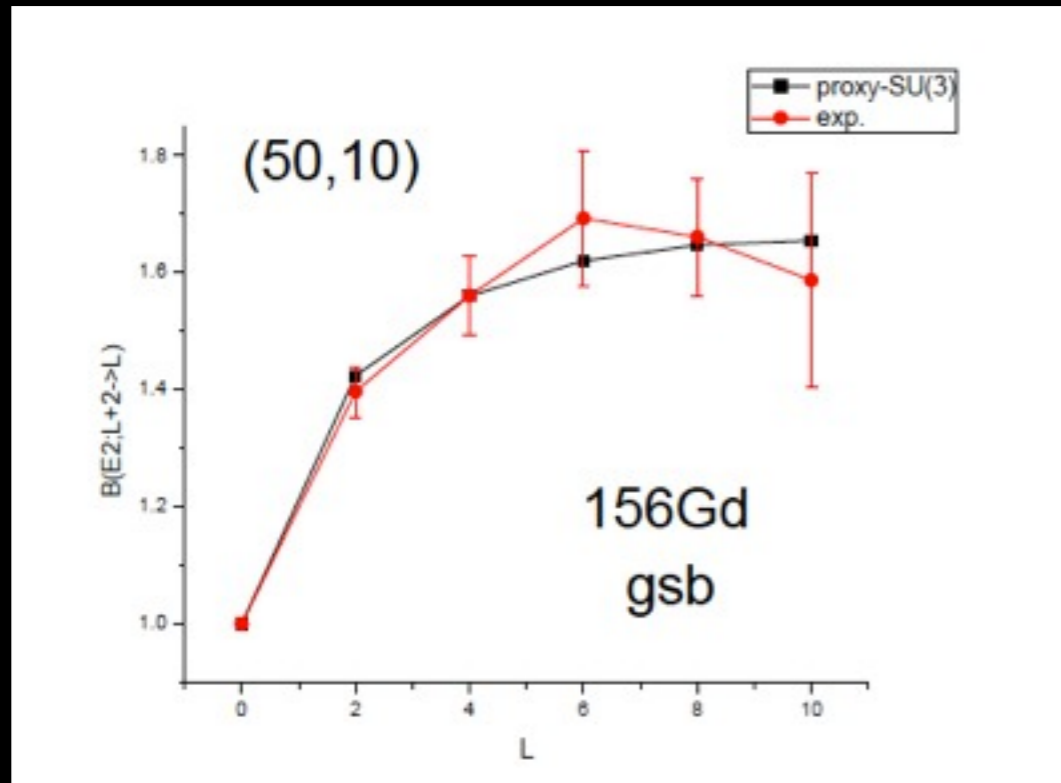
C. Bahri, D. J. Rowe and J.P. Draayer, Comput. Phys. Commun. 159,121 (2004)

Absolute $B(E2)$ s $2g \rightarrow 0g$

Martinou, Bonatsos



B(E2) ratios within g.s.b. normalized to $2g \rightarrow 0g$



Nuclear Lifetimes

$$T_{1/2} = \ln 2 \frac{10^2}{1.23 \Delta E^5 4C_2(\lambda, \mu) \langle (\lambda, \mu)KL; (1, 1)2 || (\lambda, \mu)K'L' \rangle^2} \text{sec}$$

Table 4. Shape-coexisting two-(quasi)particle states in ^{188}Pb .

I^π	$T_{1/2}$	E^* (keV)	Configuration	shape
8^-	$1.2 \mu\text{s}$	2577	$\nu 9/2^+[624], 7/2^-[514]$	High- K prolate
11^-	38 ns	2702	$\pi 9/2^-[505], 13/2^+[606]$	High- K oblate
<u>12^+</u>	<u>136 ns</u>	<u>2710</u>	$\nu i_{13/2}^{-2}$	Spherical

Our value is
134 ns

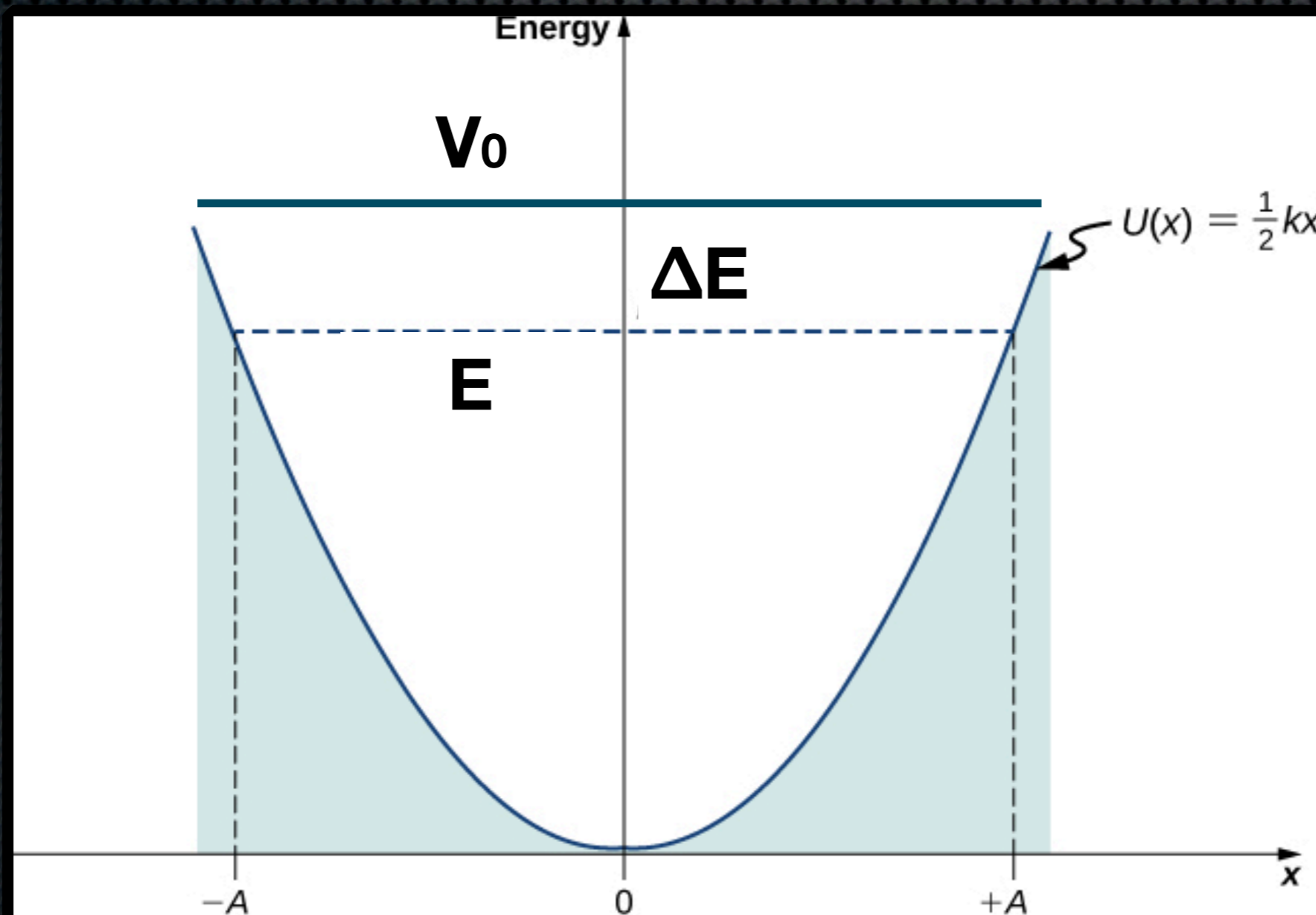
Rep. Prog. Phys. 79
(2016) 076301
Dracoulis et al.

Group Proposal to
Hellenic Foundation
for Research and
Innovation

The nucleon separation energies project

**Persons involved: A. Martinou, S.
Sarantopoulou, D. Bonatsos**

Two neutron separation energy with Proxy SU(3)



$$\Delta E = V_0 - E$$

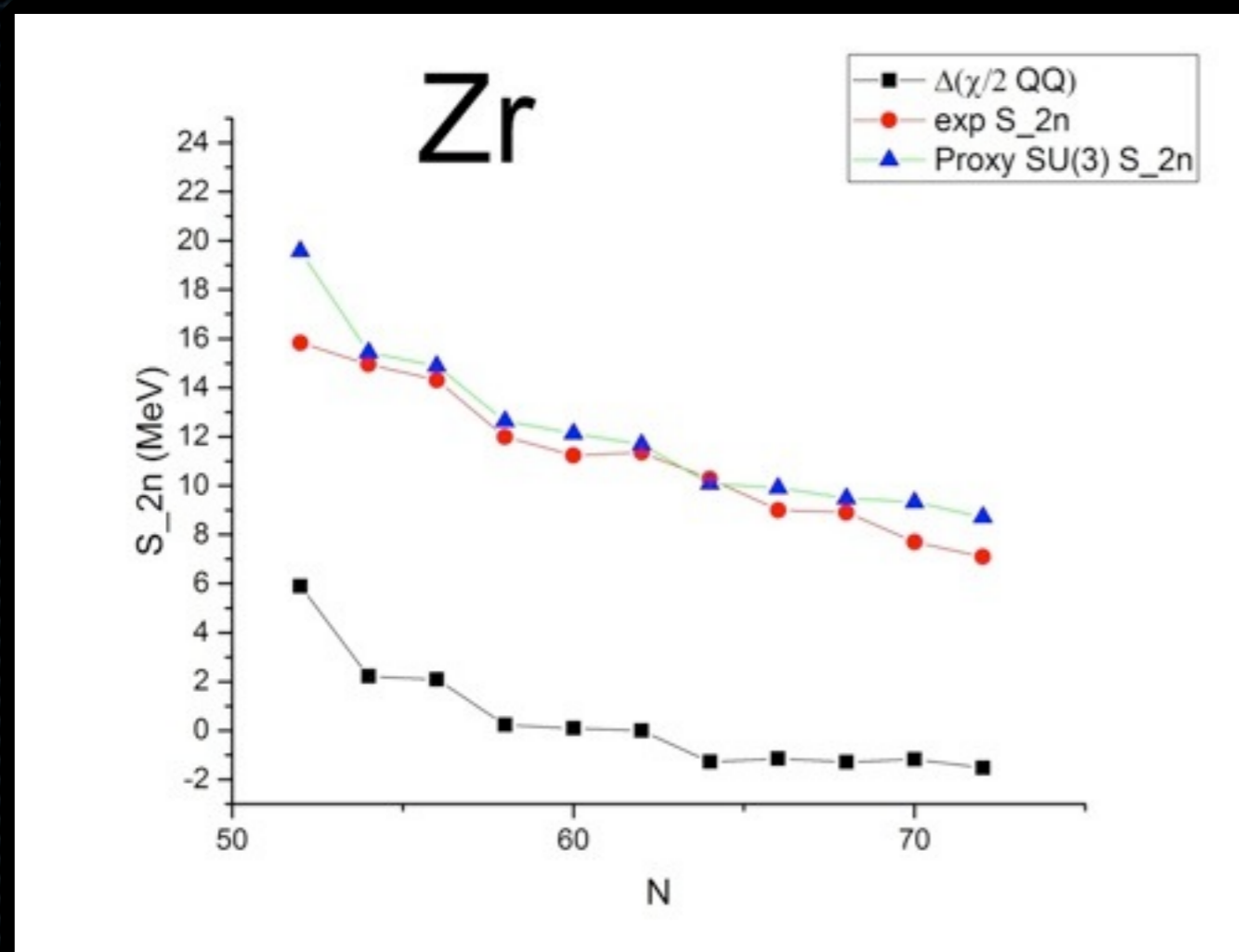
QQ interaction differences

$$\Delta(\chi/2 \text{ QQ}) = (\chi/2 \text{ QQ}(\lambda, \mu))_N - (\chi/2 \text{ QQ}(\lambda, \mu))_{N-2}$$

$$\frac{\chi}{2} = \frac{1}{4N_0}$$

Independent of fitting

Parameter free S_{2n}



$$S_{2n} = \Delta E + \Delta(\chi/2 \text{ QQ})$$

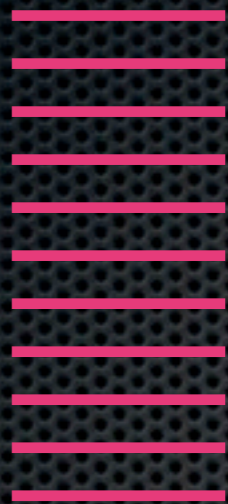
Proposal in Ministry of
Development and Investments
(EΔBM). Grade 95,22%.

The magic numbers project

Persons involved: D. Bonatsos, S. Peroulis,
A. Martinou, I.E. Assimakis

Two options

40



20

harmonic oscillator
magic numbers

50

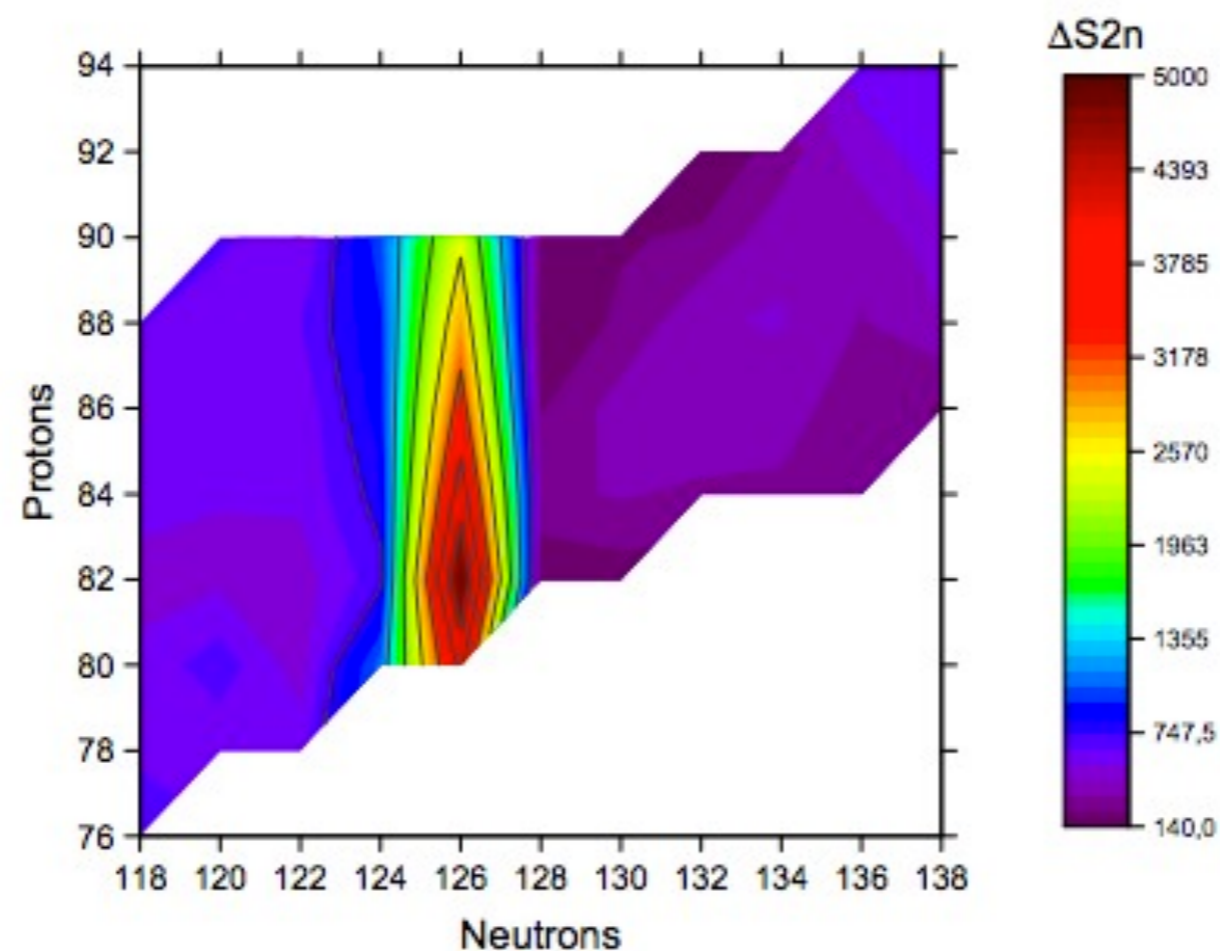


28

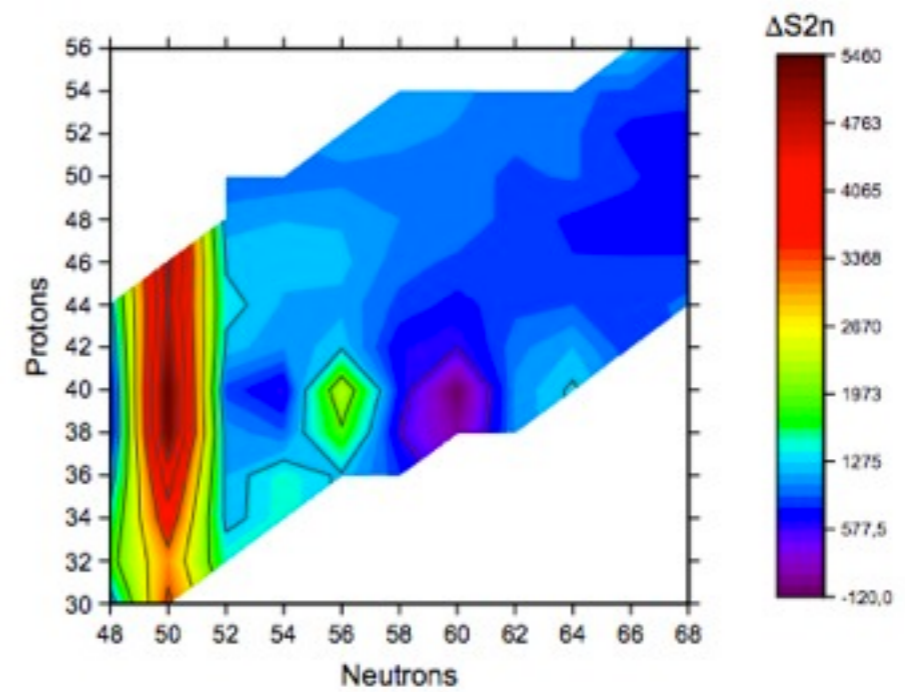
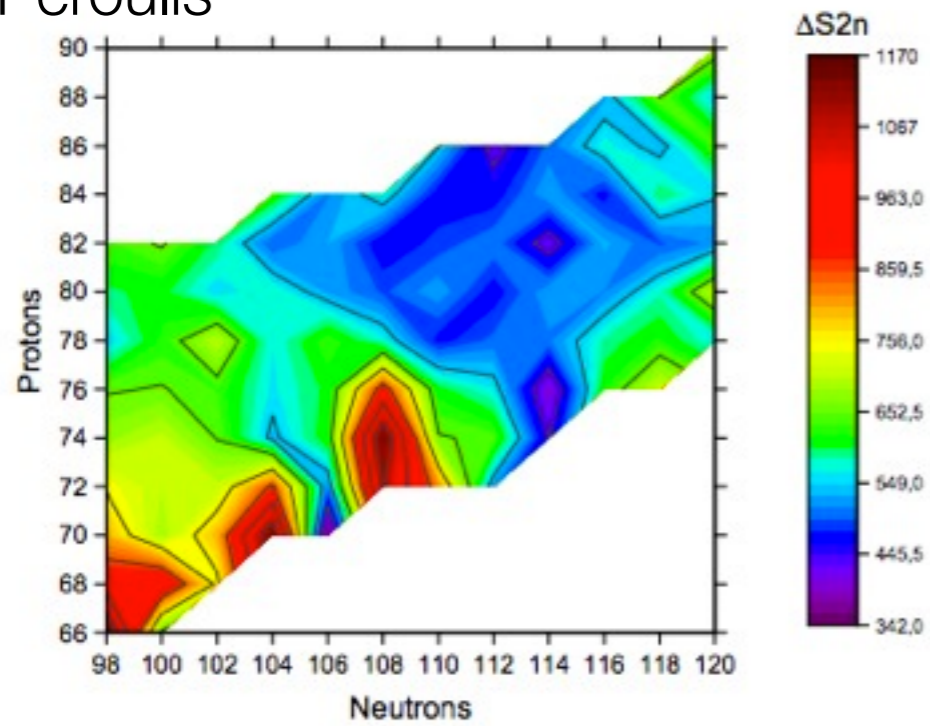
spin-orbit like magic
numbers

Major Magic Numbers

S. Peroulis

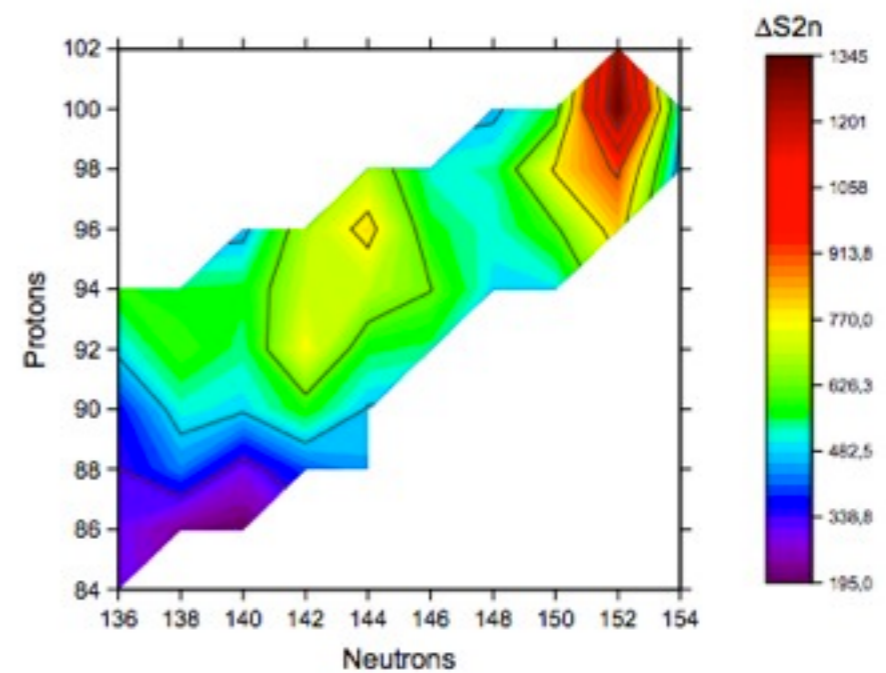


S. Peroulis



(i)

Minor Magic Numbers



Another way to search

Harmonic oscillator
magic numbers

Spin-orbit like magic
numbers

$J_1\pi$

Exotic odd nuclei

$J_2\pi$

Proposal of A.
Martinou to State
Scholarships
Foundation.
Grade 93,75%

μ_1

Compare with data

μ_2



27-28 September 2018

Scientific themes

Nuclear moments
Nuclear structure of exotic nuclei
Single-particle/collective excitations
Neutron-rich nuclei around ^{132}Sn , ^{208}Pb
Proton-rich nuclei around ^{100}Sn

gSPEC workshop on nuclear moments at GSI/FAIR

Organising committee

R. Lozeva (CSNSM/CNRS)
G. Georgiev (CSNSM/ANU)
J. Gerl (GSI/FAIR)
T. Mertzimekis (U. Athens)
A. Stuchbery (ANU)

Site

University of Milan (Milan, Italy)
together with NUSTAR week
24-27 September 2018

<https://indico.in2p3.fr/event/17501/>

Funding

- ✦ **A. Martinou, 7 months Post-Doc position in INPP.**
- ✦ **Proposal of A. Martinou, S. Sarantopoulou in Ministry of Development and Investments (EΔBM), Grade 95,22%, 15 months.**
- ✦ **Proposal of A. Martinou to State Scholarships Foundation, grade 93,75%, 24 months.**
- ✦ **Proposal of Group in Hellenic Foundation for Research and Innovation, grade ???, 3 years.**

Thank you!