Summary and Future plan for Phenomenology study

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NOTING

Signature

### Prepared the initial facility parameters

веат	ESSNUSB total running time	10			yrs	1.7.10' s per year assumed (200 days)
Beam	ESSnuSB Neutrino/Antineutrino runs	5/5	1/9	9/1	yrs	To be optimized
Far Detector	Expected performance for signal and background components	yes				Migration matrices and efficiencies from ESSnuSB project
Far Detector	Overall normalization systematic uncertainty for signal	5 %	2 %	10 %		
Far Detector	Overall normalization systematic uncertainty for background	10 %	5 %	15 %		
Far Detector	Overall shape systematic uncertainty for signal	2 %	0 %	5 %		
Far Detector	Overall shape systematic uncertainty for background	2 %	0 %	5 %		
Near Detectors	Neutrino energy resolution	15 %	5%	25%		
Near Detectors	Signal selection efficiency	50 %	40 %	70 %		
Near Detectors	Background rejection efficiency	99 %	90 %	99 %		
Near Detectors	Overall normalization systematic uncertainty for signal	5 %	5 %	10 %		
Near Detectors	Overall normalization systematic uncertainty for background	10 %	10 %	20 %		
Near Detectors	Overall shape systematic uncertainty for signal	2 %	0 %	10 %		
Near Detectors	Overall shape systematic uncertainty for background	2 %	0 %	10 %		
Oscillation Parameters	θ <sub>12</sub>	33.41	31.31	35.75	degrees	NuFit 5.2 (http://www.nu-fit.org)
Oscillation Parameters	θ <sub>13</sub>	8.58	8.23	8.94	degrees	NuFit 5.2 (http://www.nu-fit.org)
Oscillation Parameters	θ <sub>23</sub>	42.2	39.0	52.0	degrees	NuFit 5.2 (http://www.nu-fit.org)
Oscillation Parameters	Δm <sup>2</sup> <sub>21</sub>	7.41	6.82	8.03	10^-5 eV^2	NuFit 5.2 (http://www.nu-fit.org)
Oscillation Parameters	Δm² <sub>31</sub>	2.507	2.406	2.590	10^-3 eV^2	NuFit 5.2 (http://www.nu-fit.org)
Oscillation Parameters	$\delta_{CP}$	232	0	360	degrees	NuFit 5.2 (http://www.nu-fit.org)

### Statistical Model

- Discussed an improved satitiscal analysis method Discussed by Spyros and summarized by George F and Spyros
- Will proceed to implement this method in the phenomenology studies
- Till then will continue with the current method i.e.,

$$\chi 2 = 2 \sum_{i=1}^{n} \left[ N_i^{test} - N_i^{true} - N_i^{true} \ln \left( \frac{N_i^{test}}{N_i^{true}} \right) \right]$$

# Supernova Study at FD

- Will wait for the migration matrices
- In the meantime will proceed with a Gaussian resolution
- Estimate event rates
- Study physics scenarios

#### Atmospheric Study at FD

- Will continue to develop a code
- Study physics scenarios

Solar Study at FD

• Will look at the possibilities

## Study at NDs

- Will wait for the fluxes for LEnuSTORM and LEMNB
- Will wait for the migration matrices for LEMON-D and END (Can work with Gaussian energy resolution)
- Study physics scenarios
- Sterile: Performed a preliminary studies considering production of neutrinos at different points in the race-track/decay tunnel Alessio

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Thanks everyone for very fruiyful meeting