

Hamiltonian Monte Carlo & eigenvector continuation for *ab initio* nuclear physics

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UNIVERSITY OF TECHNOLOGY

ISNET-9, 22-26 May 2023, Washington University in St. Louis



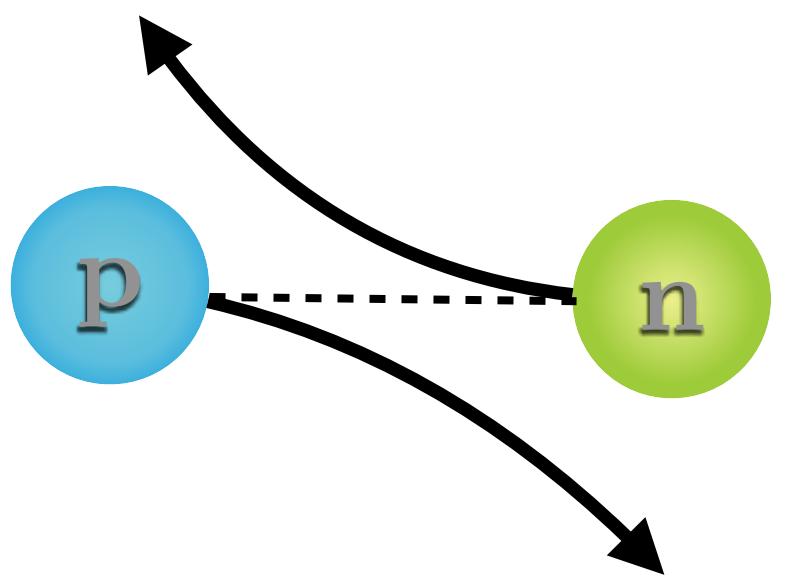
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Contents

Ab initio modelling of nuclear systems

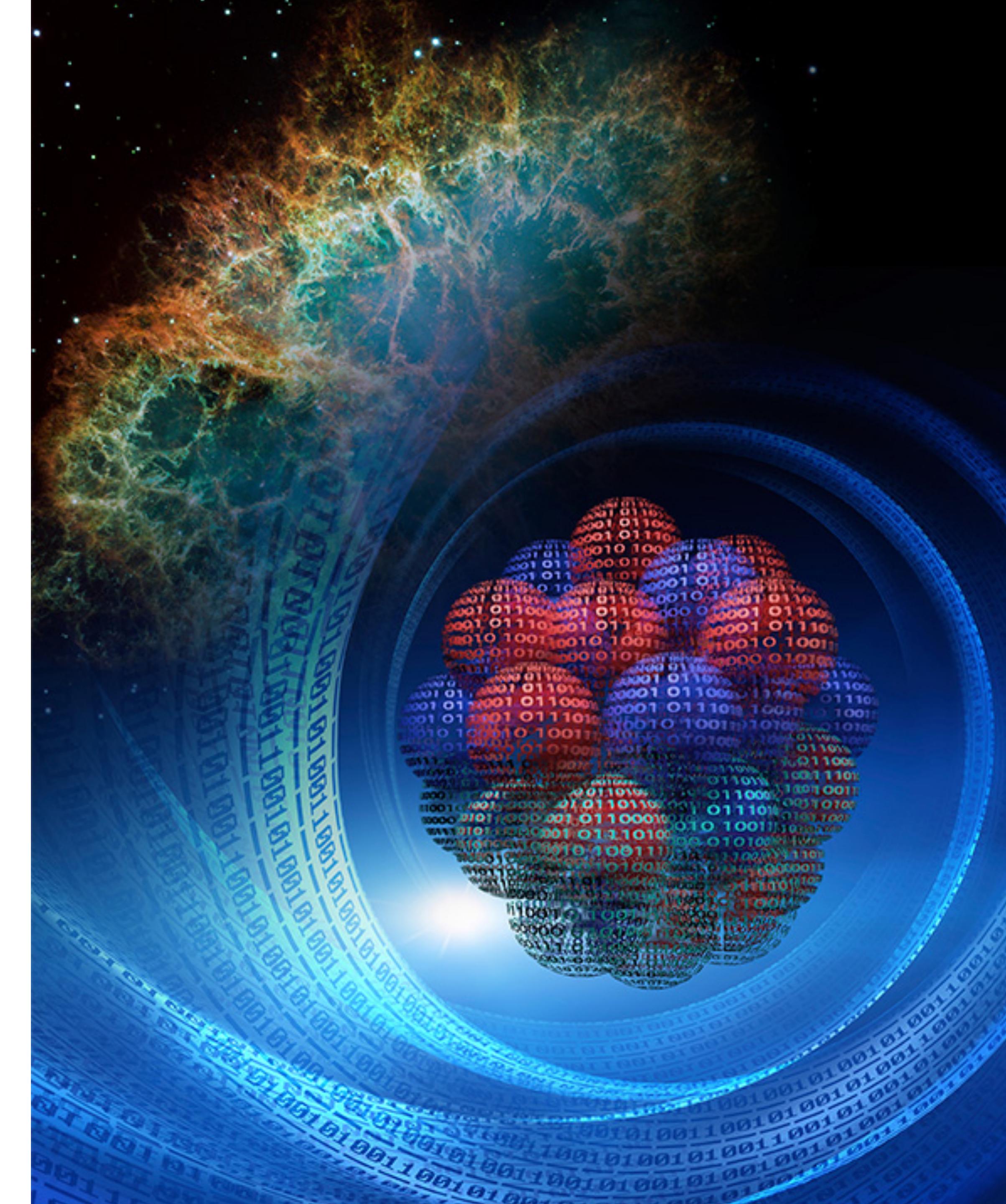
Bayesian inference of neutron-proton scattering



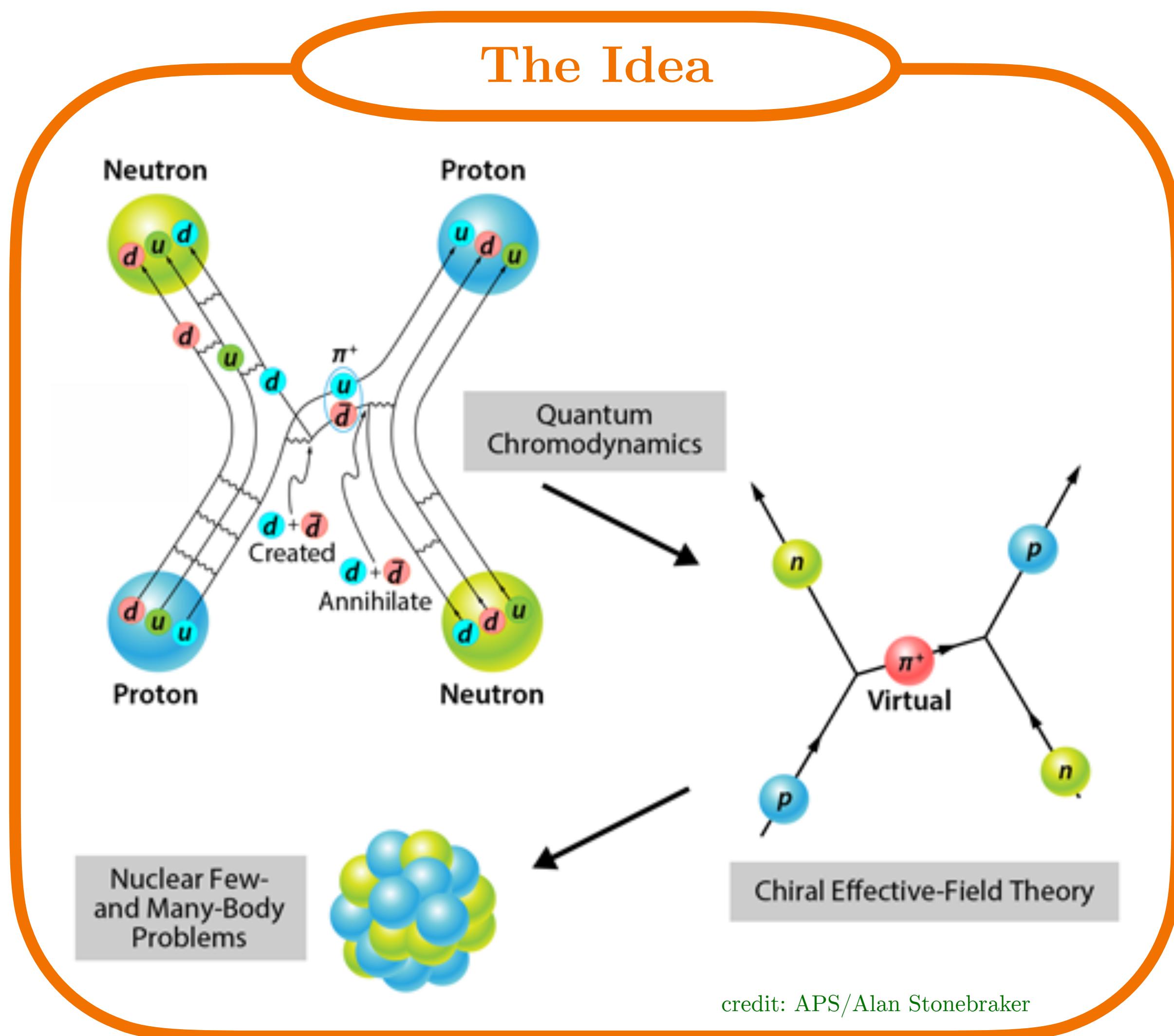
a prototypical system for studying the
strong nuclear interaction

For details, see our recent preprint

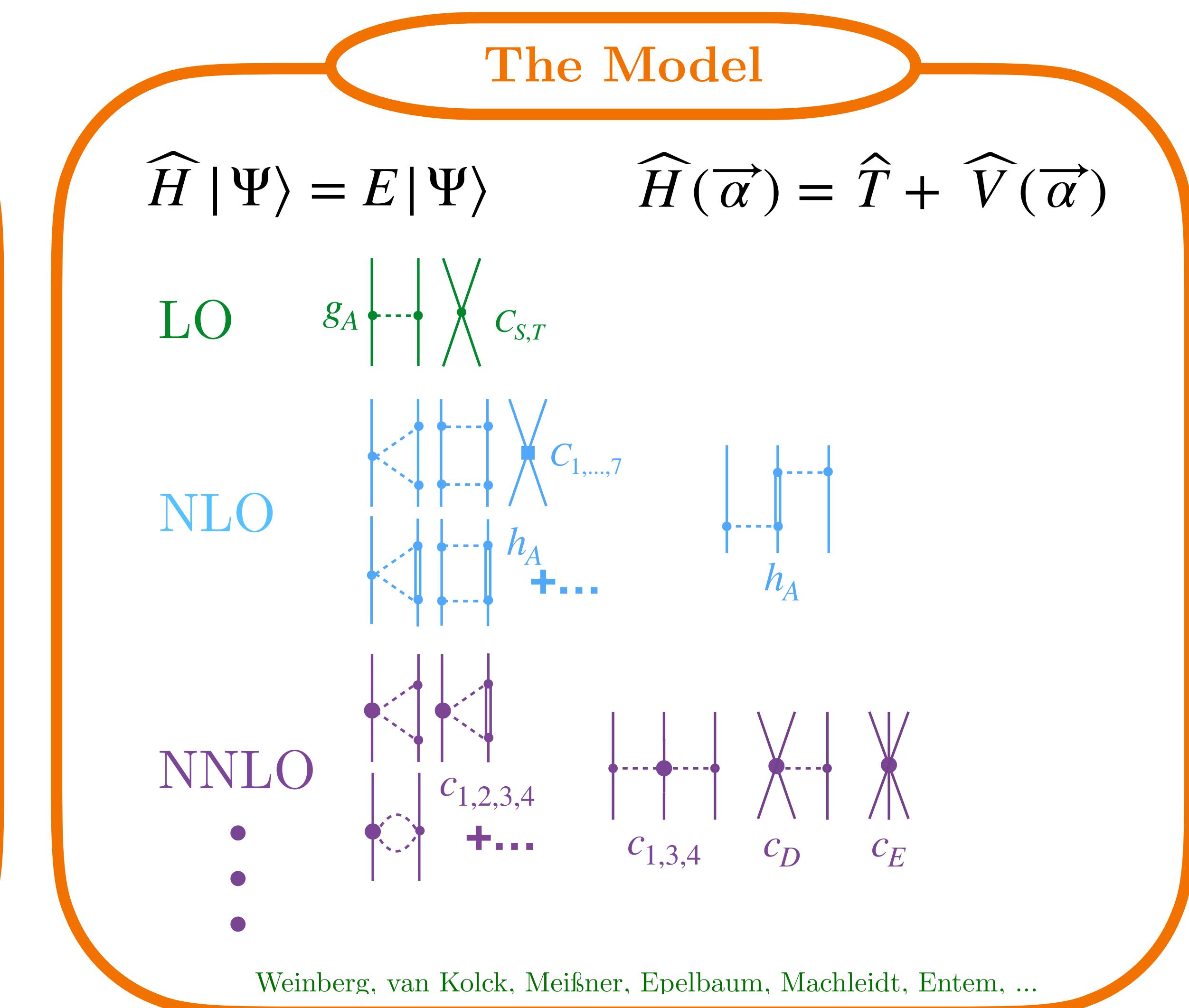
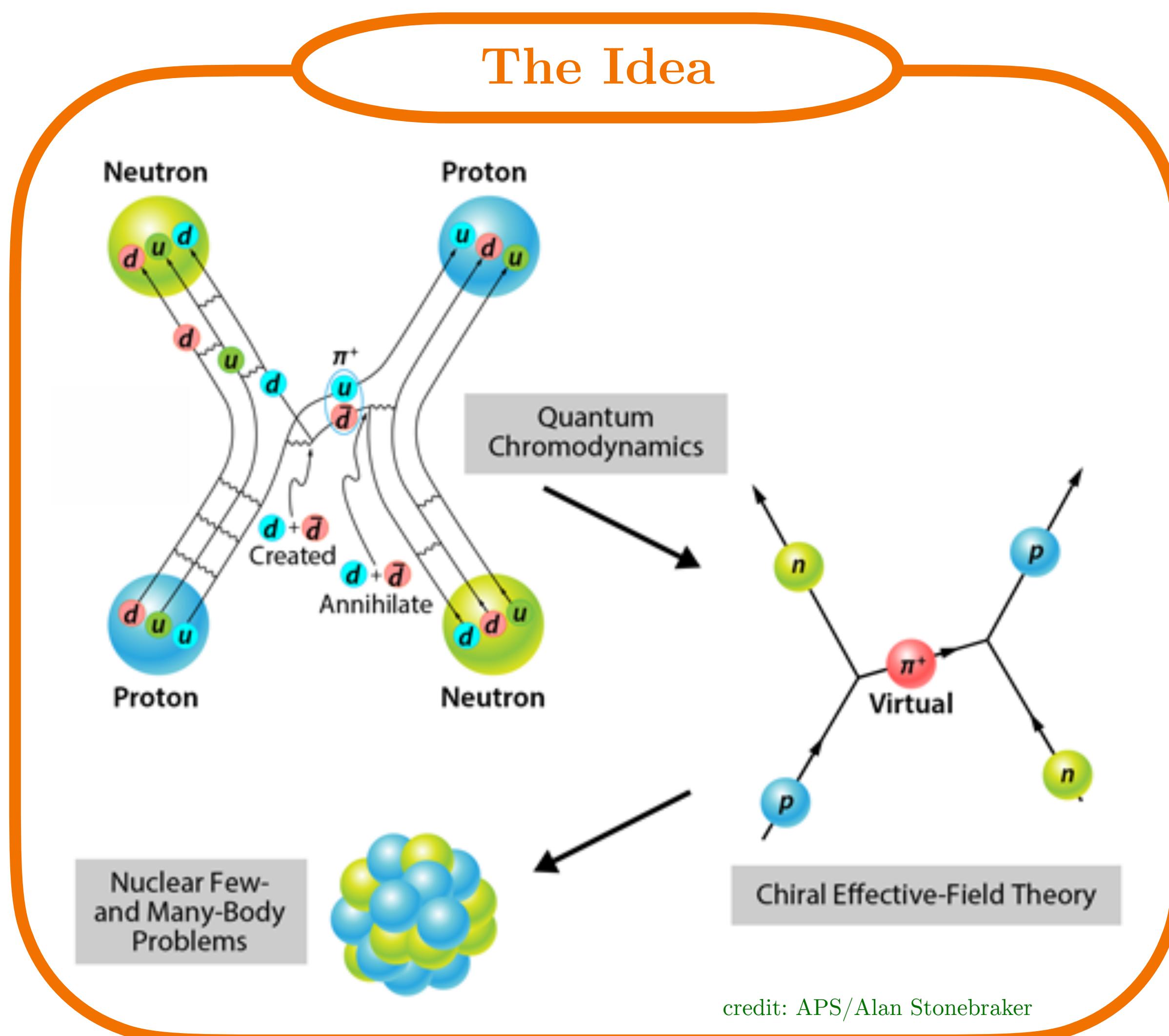
Isak Svensson, Andreas Ekström, Christian Forssén arXiv:2304.02004 [nucl-th]



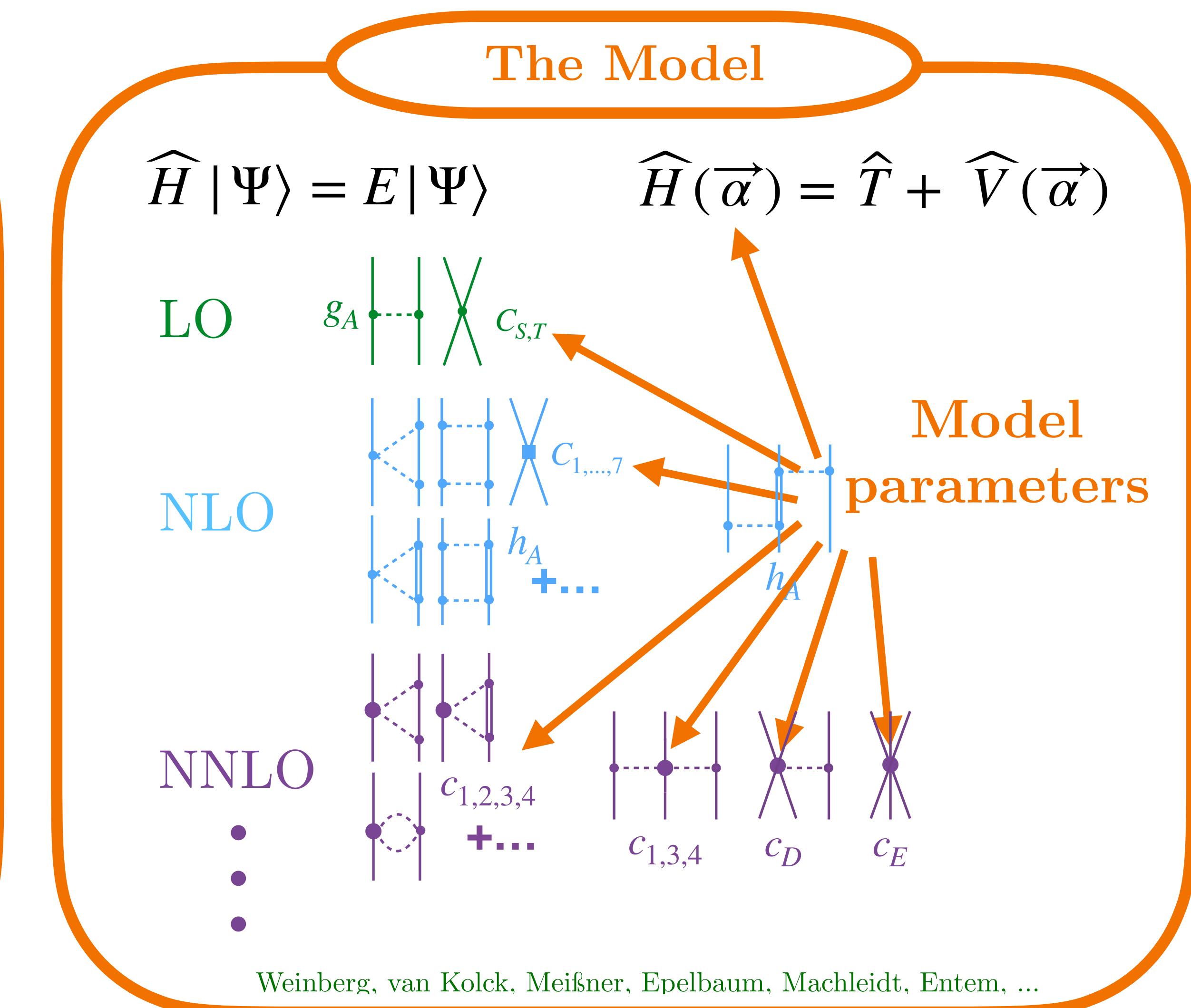
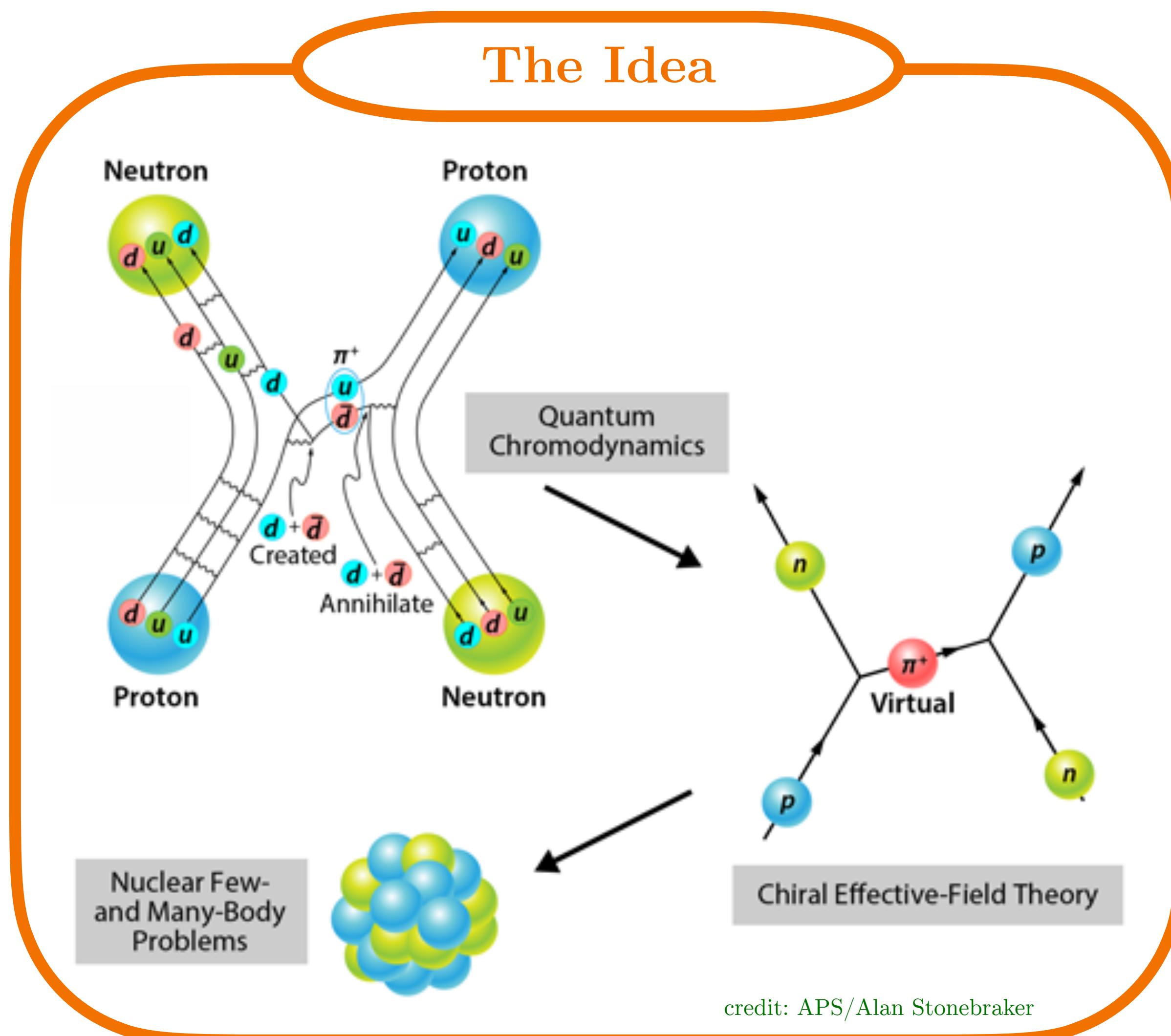
Ab initio modelling of nuclear systems



Ab initio modelling of nuclear systems



Ab initio modelling of nuclear systems



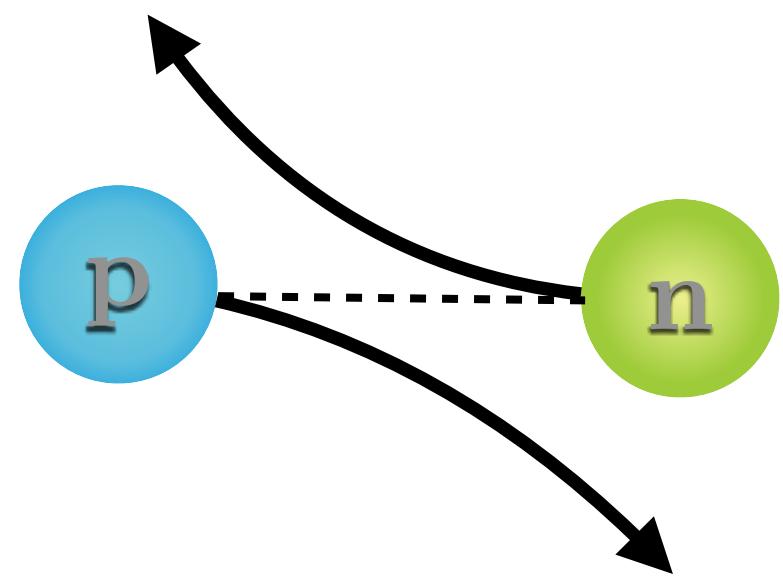
Why *ab initio* nuclear theory?

$$D_i = M_i(\vec{\alpha}) + \delta D_i + \delta M_i$$

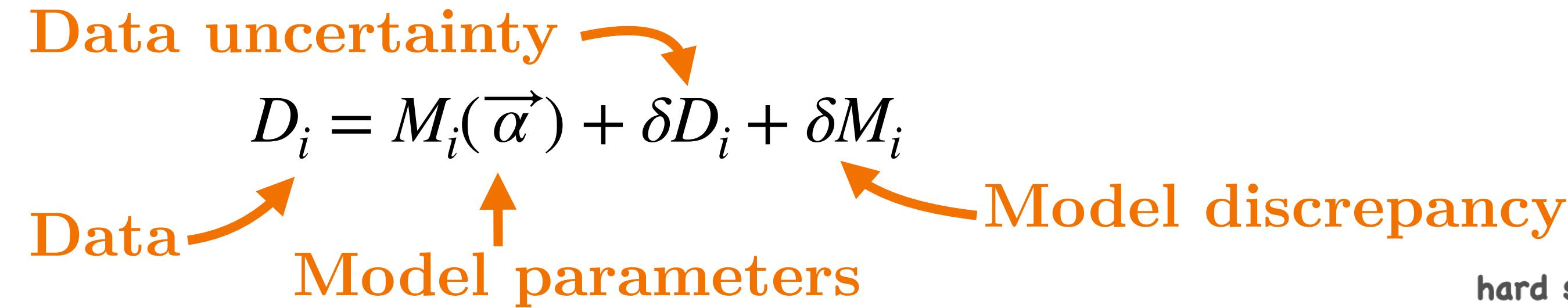
Diagram illustrating the components of data uncertainty:

- Data**: Represented by an orange arrow pointing to the term $M_i(\vec{\alpha})$.
- Model parameters**: Represented by an orange arrow pointing to the term δM_i .
- Data uncertainty**: Represented by an orange arrow pointing to the term δD_i .
- Model discrepancy**: Represented by an orange arrow pointing to the term δM_i .

Why *ab initio* nuclear theory?



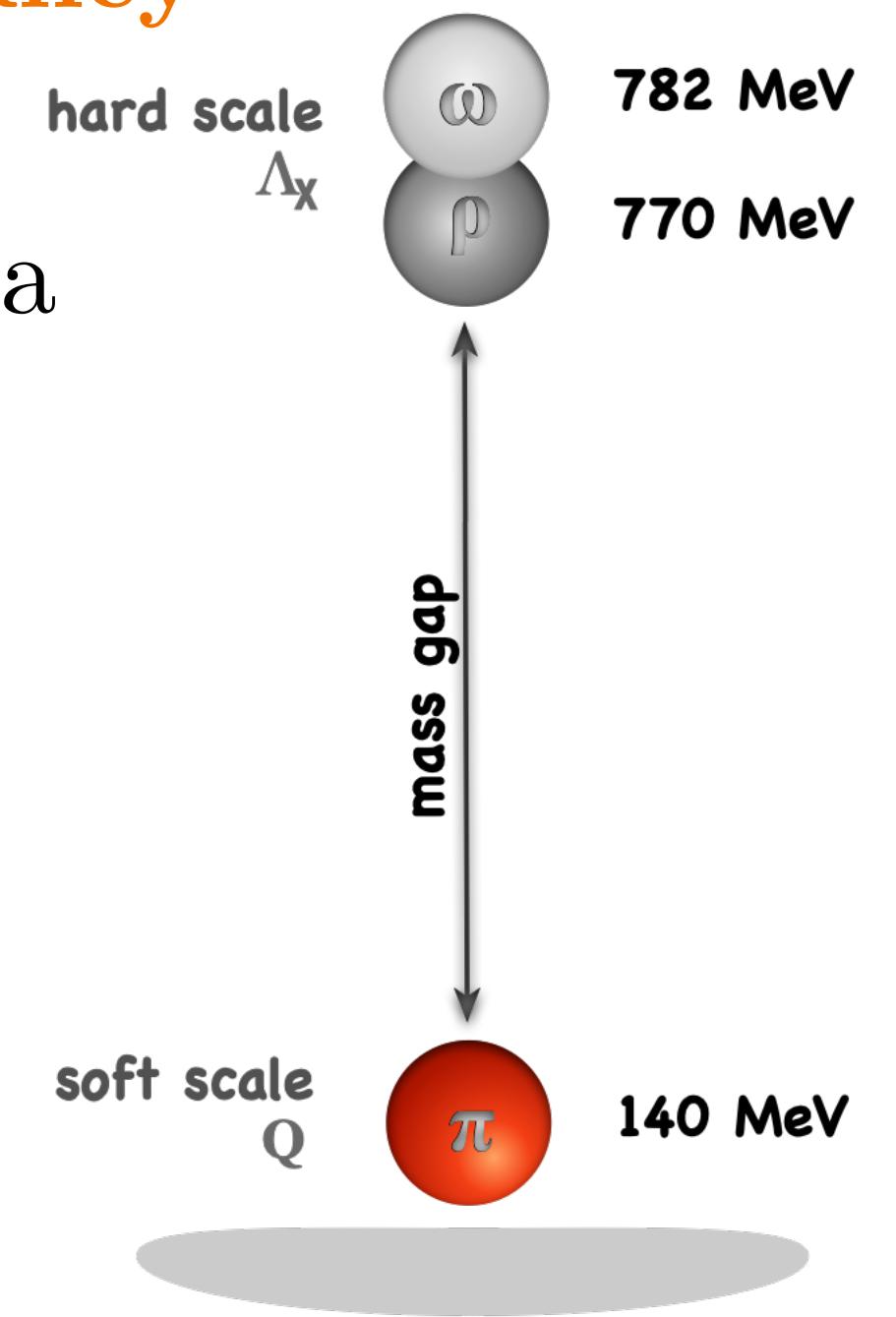
nucleon-nucleon
scattering data
(~2000 of relevance)



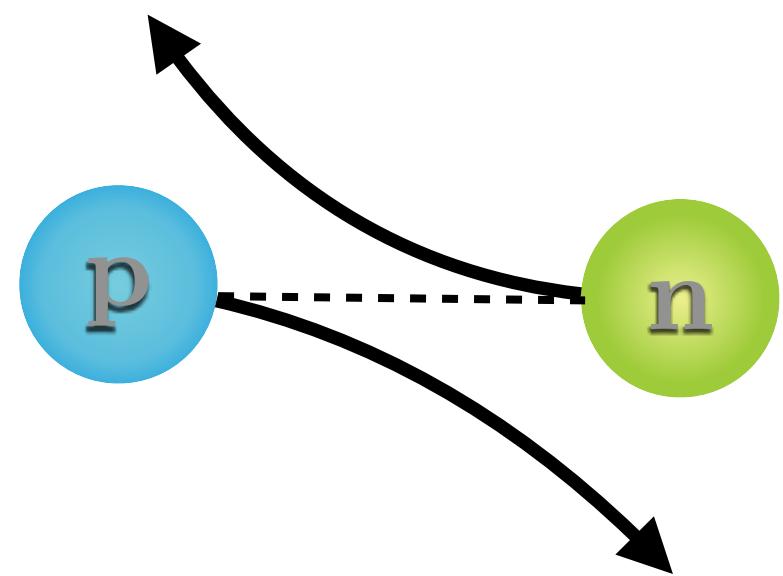
We have a model (an effective field theory) for describing data

$$M_i(\vec{\alpha}) = y^{(k)}(\vec{\alpha}; \vec{x}_i) = y_{\text{ref}} \left[\gamma_0 \left(\frac{Q}{\Lambda_\chi} \right)^0 + \gamma_1 \left(\frac{Q}{\Lambda_\chi} \right)^1 + \gamma_2 \left(\frac{Q}{\Lambda_\chi} \right)^2 + \dots + \gamma_k \left(\frac{Q}{\Lambda_\chi} \right)^k \right]$$

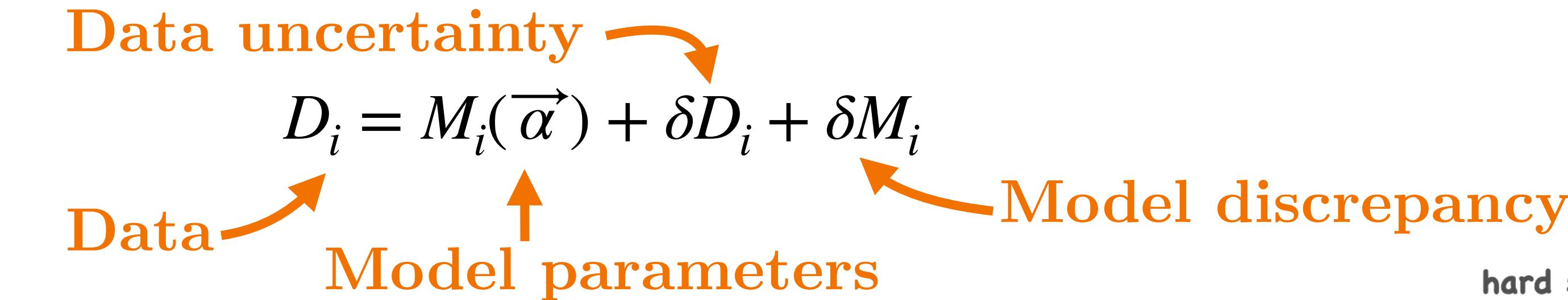
LO NLO N2LO N_kLO



Why *ab initio* nuclear theory?



nucleon-nucleon
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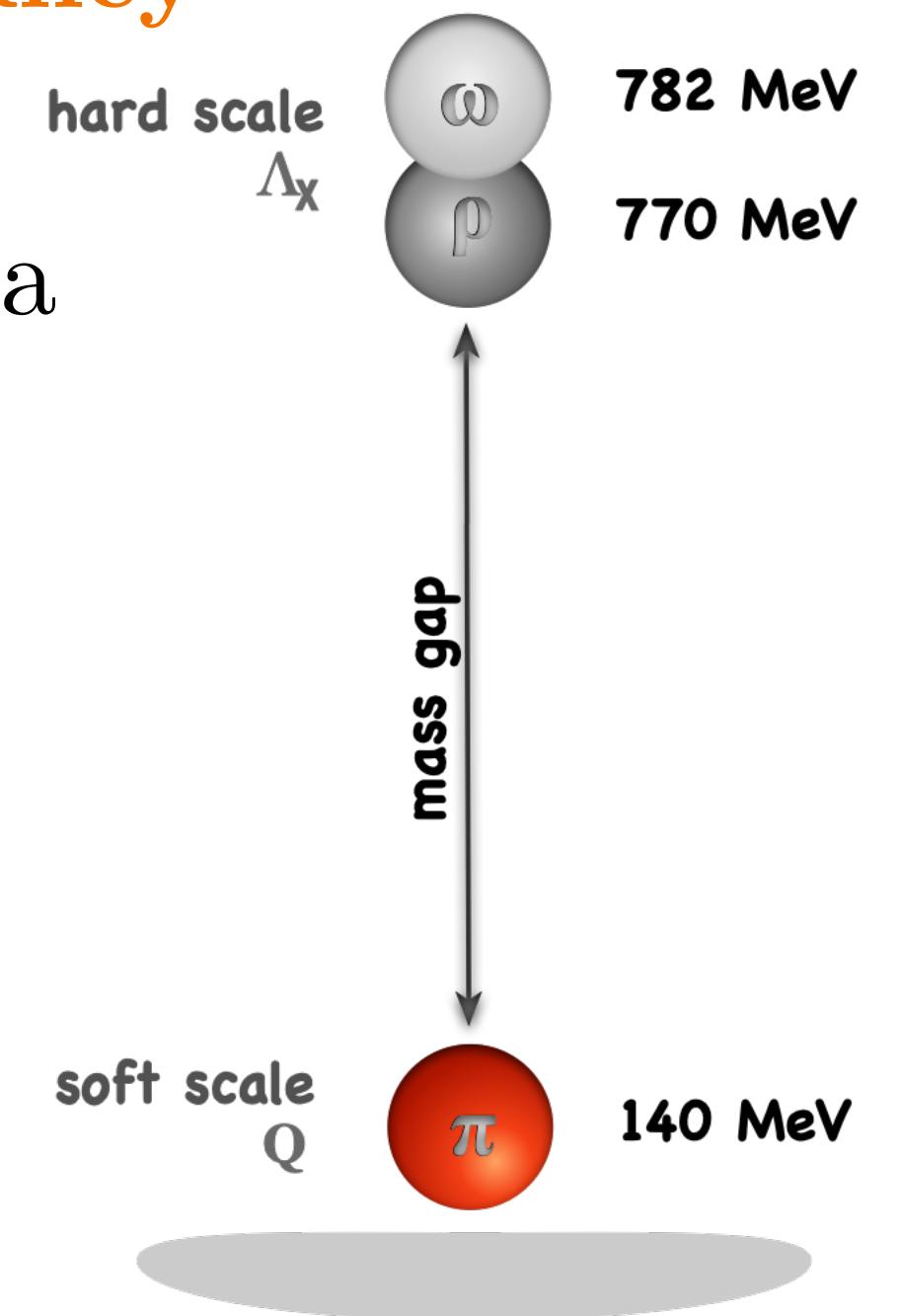
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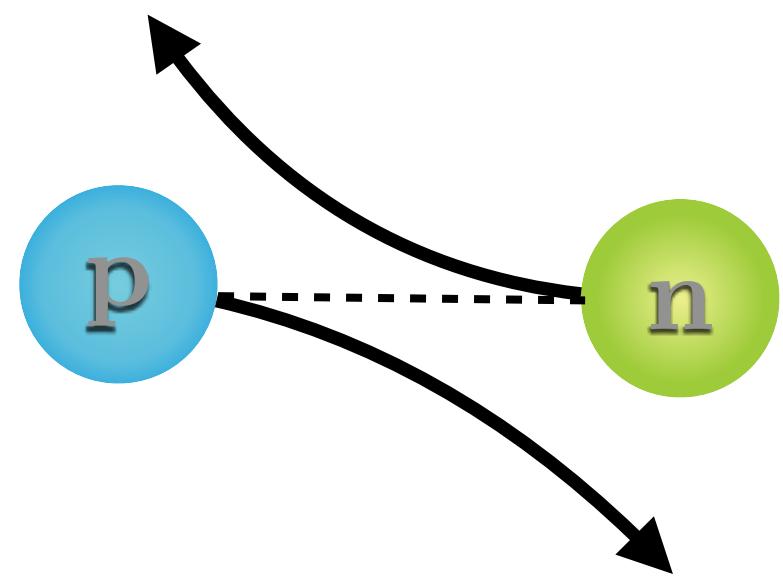
LO NLO N2LO NkLO

The systematicity of the *ab initio* method creates an **inferential advantage** since we claim to know something about the model discrepancy.

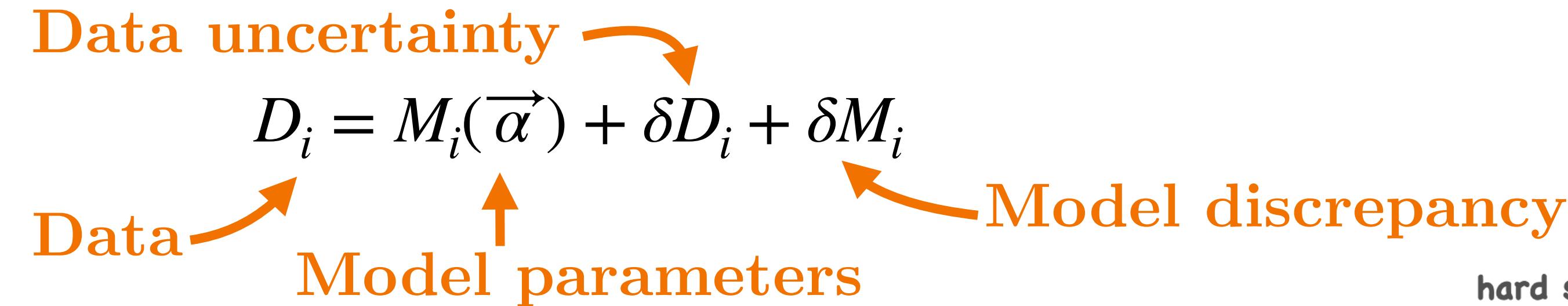
$$\delta M_i = y_{\text{ref}} \sum_{i=k+1}^{\infty} \gamma_i \left(\frac{\mathcal{Q}}{\Lambda_\chi} \right)^i$$



Why *ab initio* nuclear theory?



nucleon-nucleon
scattering data
(~2000 of relevance)



We have a model (an effective field theory) for describing data

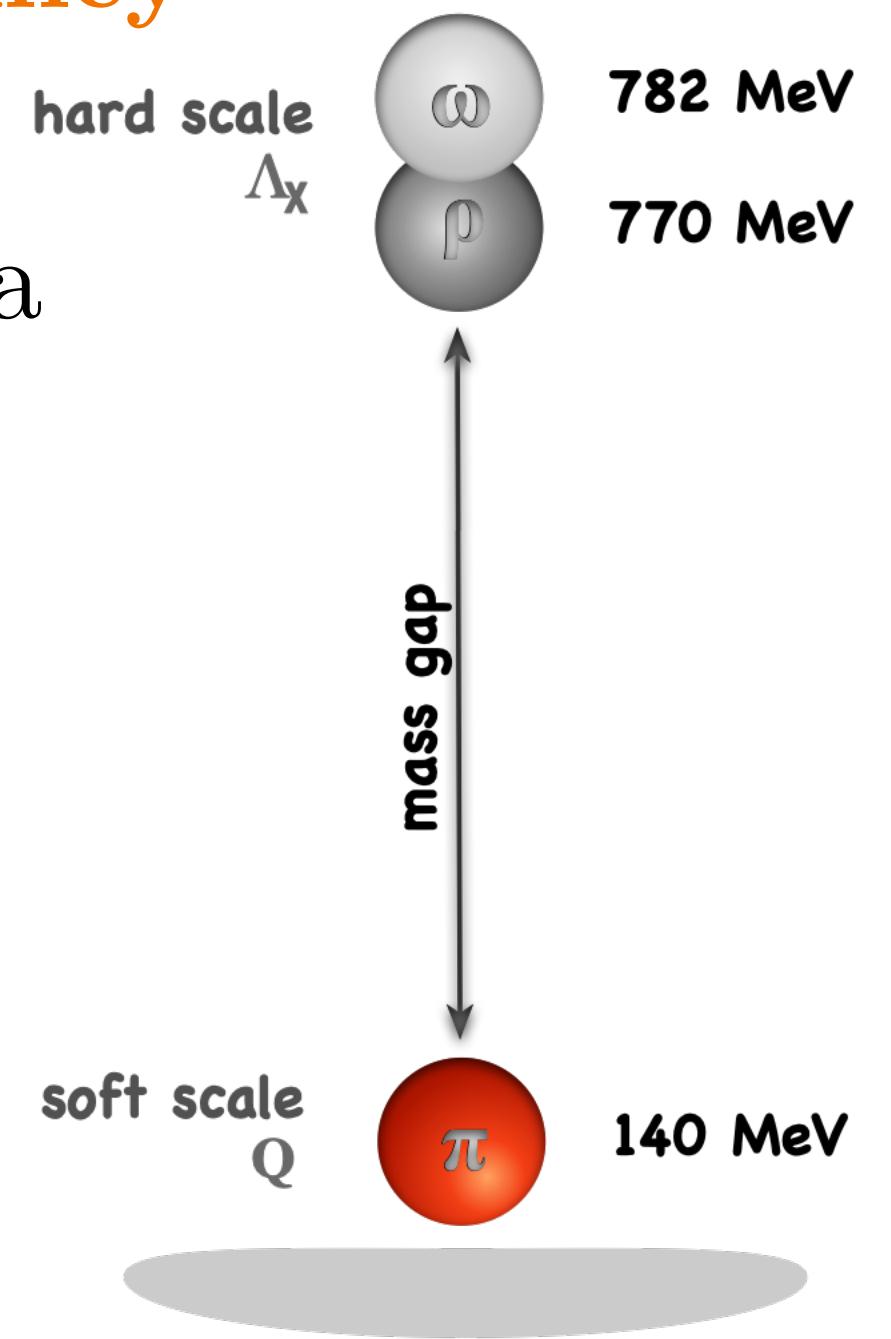
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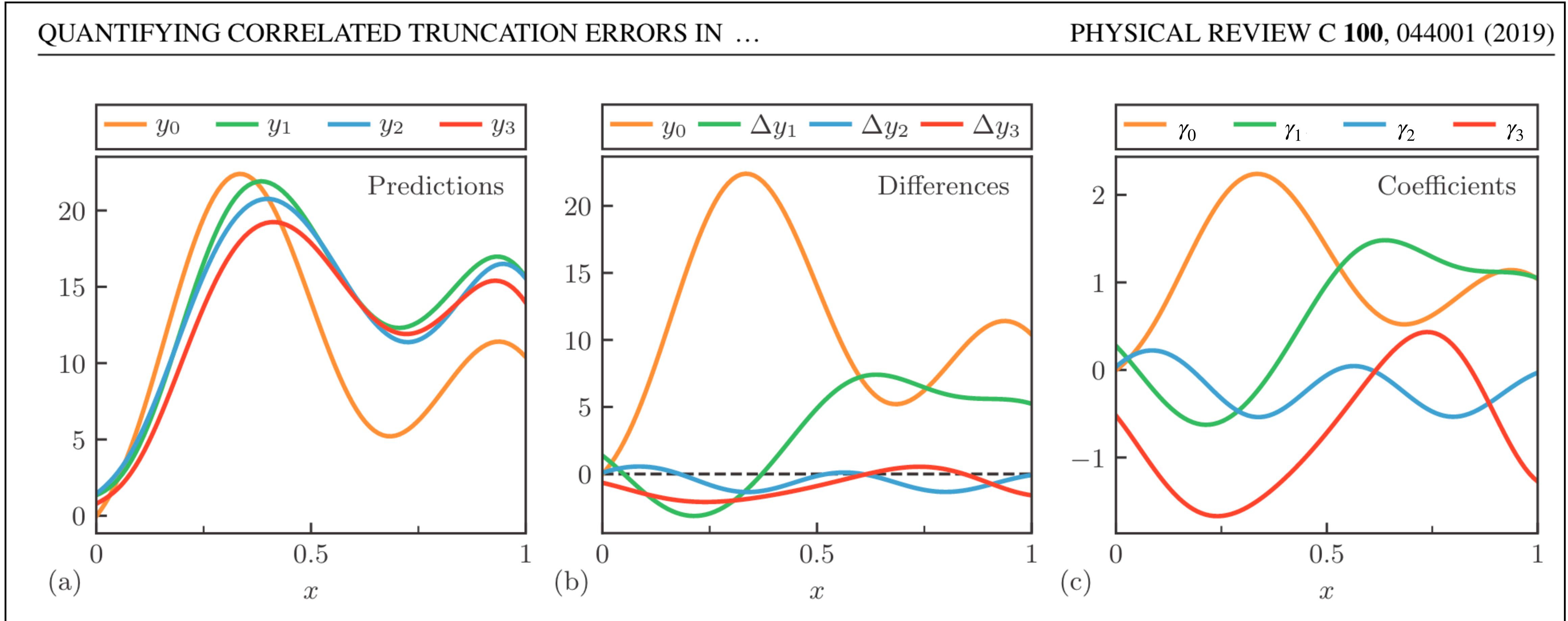
$$\delta M_i = y_{\text{ref}} \sum_{i=k+1}^{\infty} \gamma_i \left(\frac{\mathcal{Q}}{\Lambda_\chi} \right)^i$$

We are uncertain about the expansion coefficients



Predictions $y^{(k)}$ correlated across scattering energy & angle

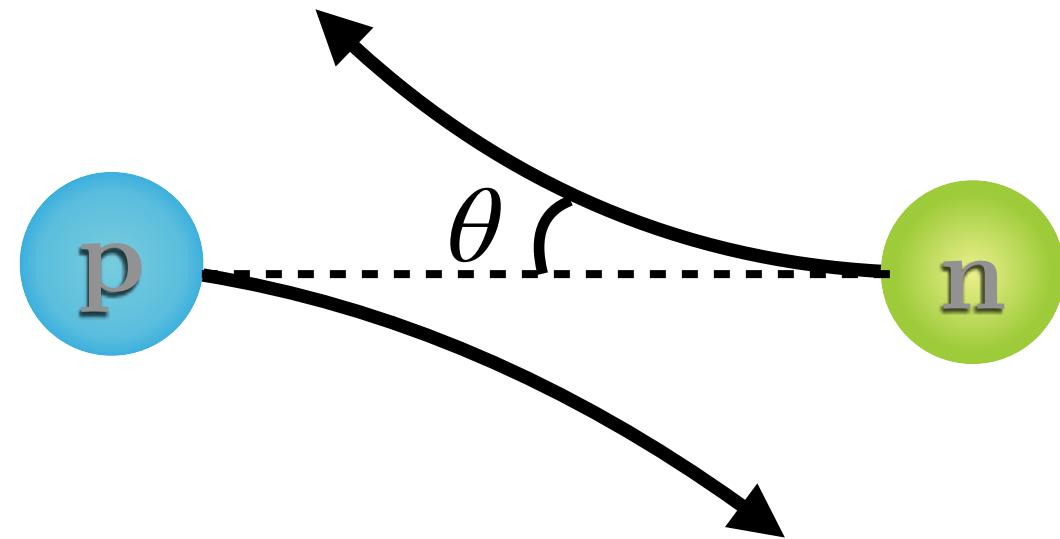
adapted from J. A. Melendez, R. J. Furnstahl, D. R. Phillips, M. T. Pratola, S. Wesolowski Phys. Rev. C 100 (4), 044001 (2019)



we use $y_{\text{ref}} = 0.5$ (Idaho-N3LO) for polarization (σ_{tot} & $\sigma(\theta)$)

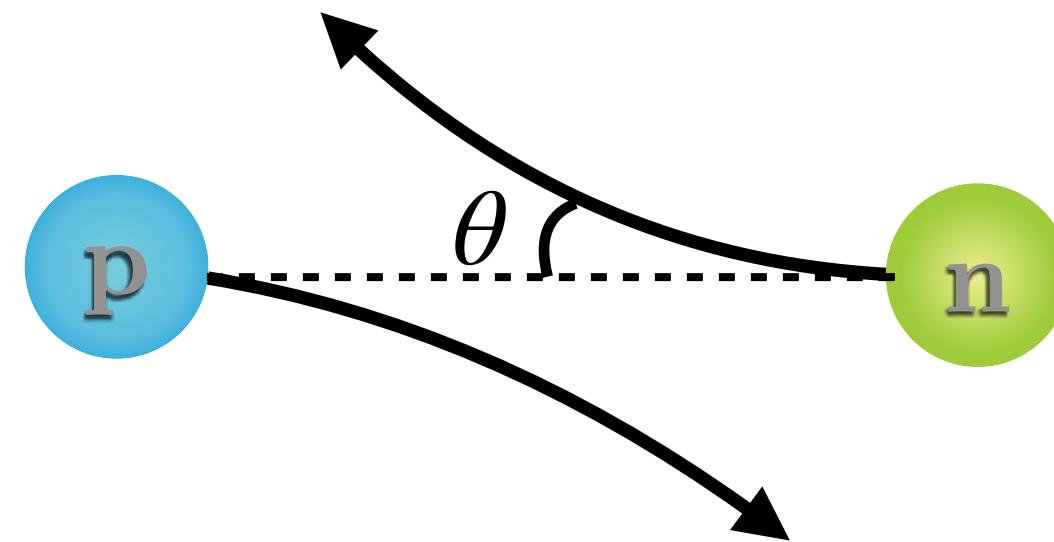
$$\gamma^{(k)}(\vec{x}_j) = \frac{y_{\text{th}}^{(k)}(\vec{\alpha}^{\star}; \vec{x}_j) - y_{\text{th}}^{(k-1)}(\vec{\alpha}^{\star}; \vec{x}_j)}{y_{\text{ref}}(Q/\Lambda_{\chi})^k}$$

A Gaussian-process model for correlated δM



Quantify expansion coefficients by evaluating the model at consecutive chiral orders for a 2D grid of \vec{x}_j values (energies T_{lab} and angles θ).
n.b. we do this for a sensible choice of $\vec{\alpha} = \vec{\alpha}^$*

A Gaussian-process model for correlated δM

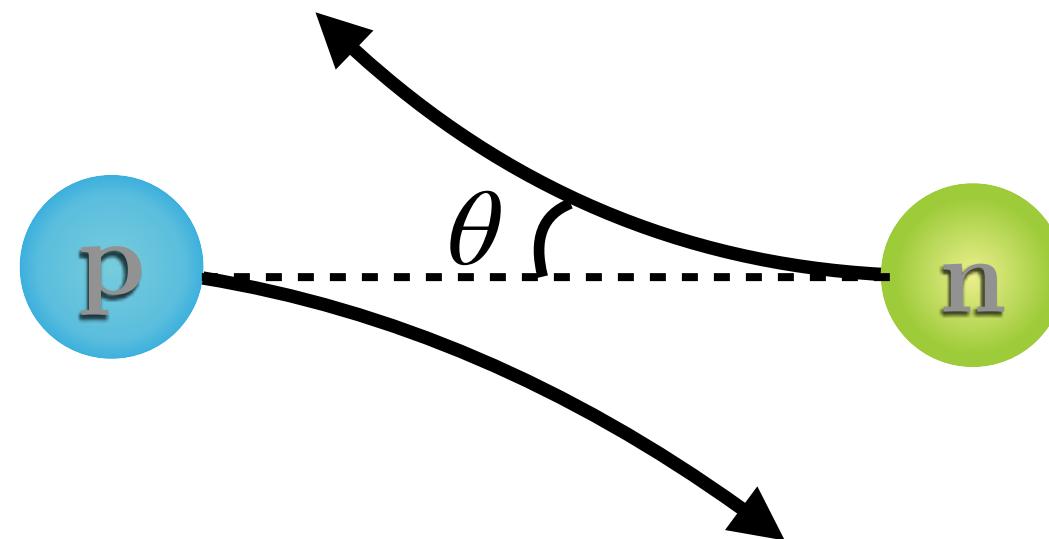


Quantify expansion coefficients by evaluating the model at consecutive chiral orders for a 2D grid of \vec{x}_j values (energies T_{lab} and angles θ).
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We model the correlated expansion parameters as a two-feature Gaussian process using a SQE kernel

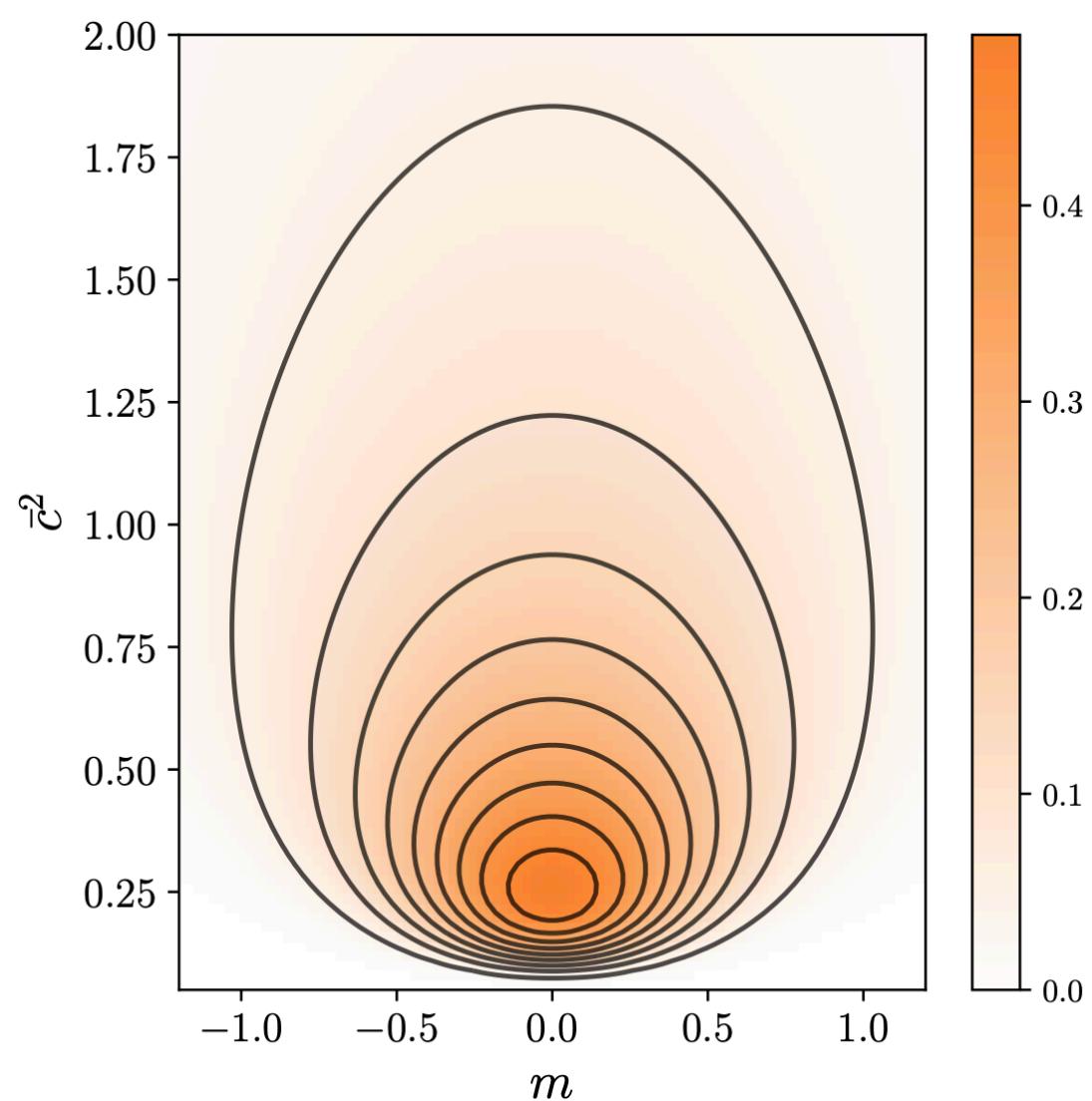
$$\gamma(\mathbf{x}) | m, \vec{\ell}, \bar{c}^2 \sim \mathcal{GP}[m, \bar{c}^2 k(\mathbf{x}', \mathbf{x}; \vec{\ell})]$$

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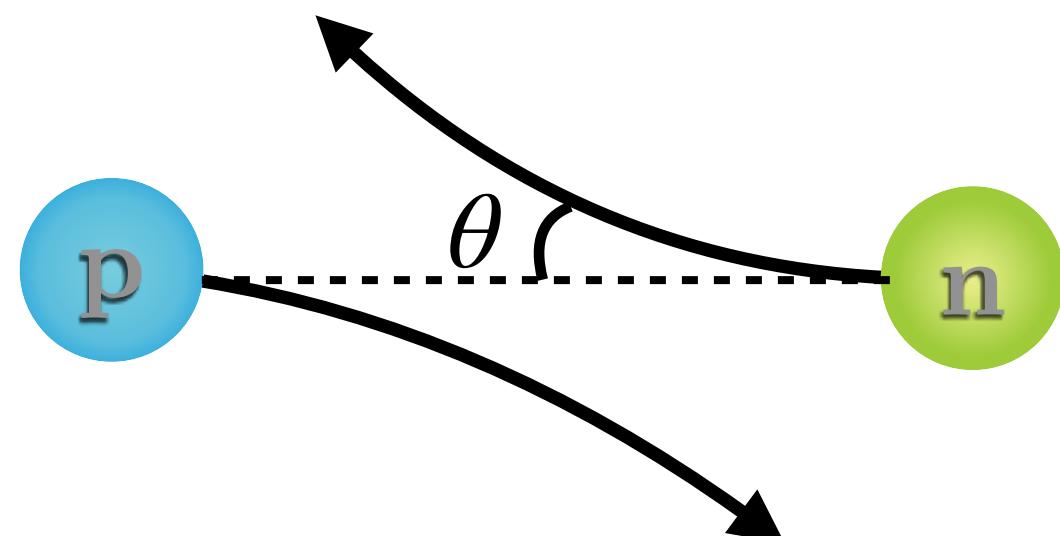


$$\gamma(\mathbf{x}) | m, \vec{\ell}, \bar{c}^2 \sim \mathcal{GP}[m, \bar{c}^2 k(\mathbf{x}', \mathbf{x}; \vec{\ell})]$$

Our prior for the GP hyperparameters

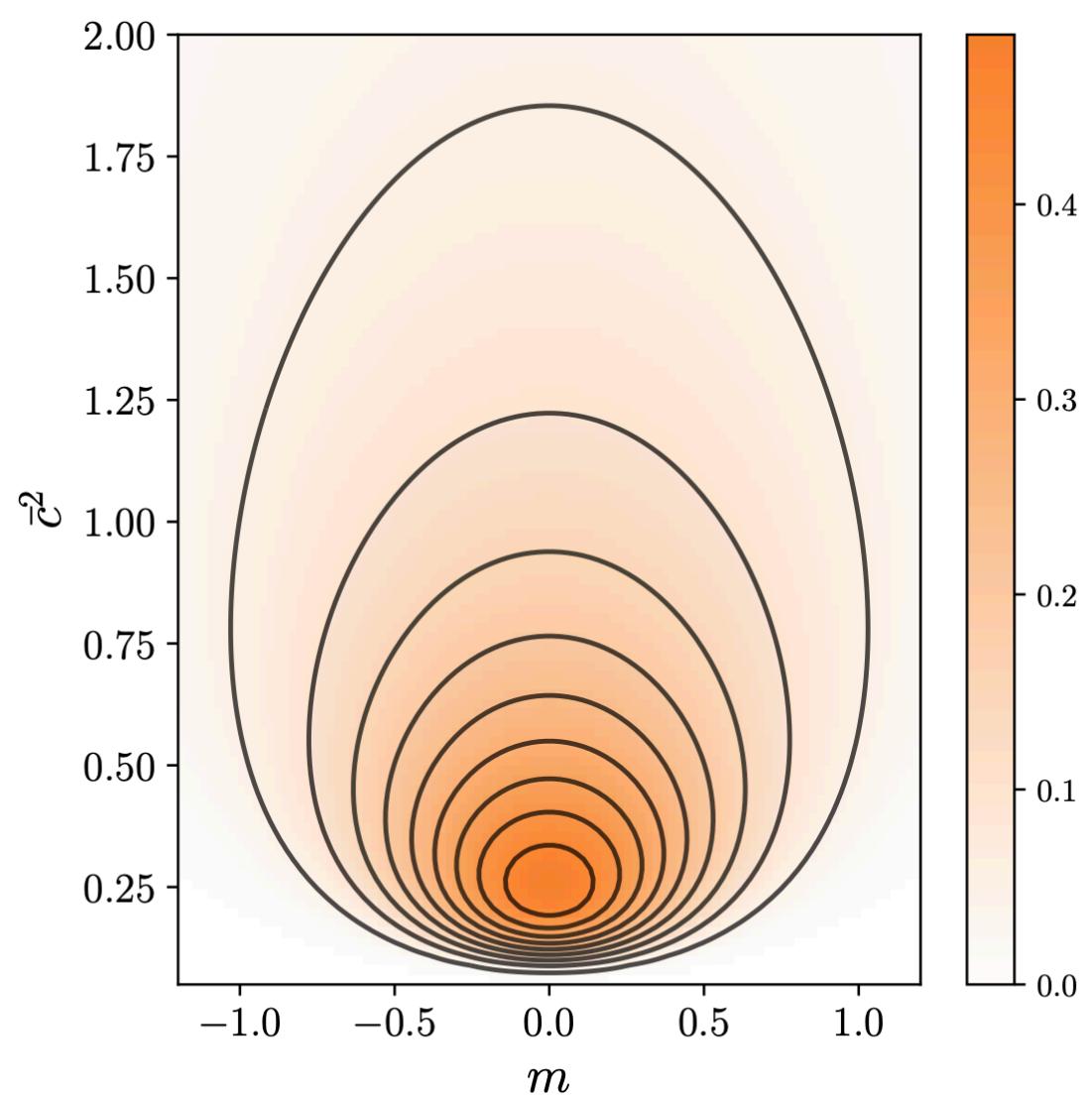
uniform distribution
for the lengthscales $\vec{\ell}$

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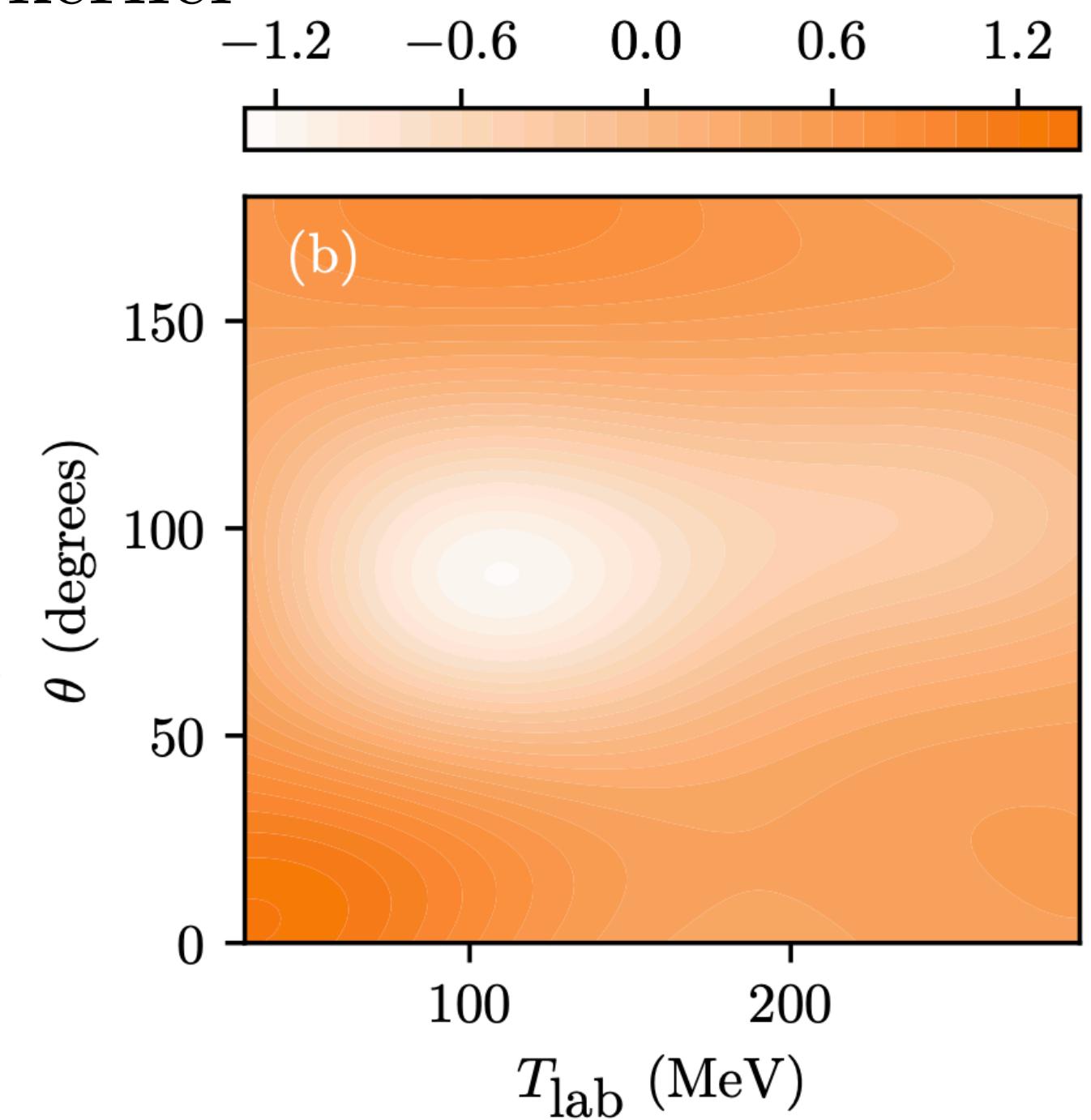
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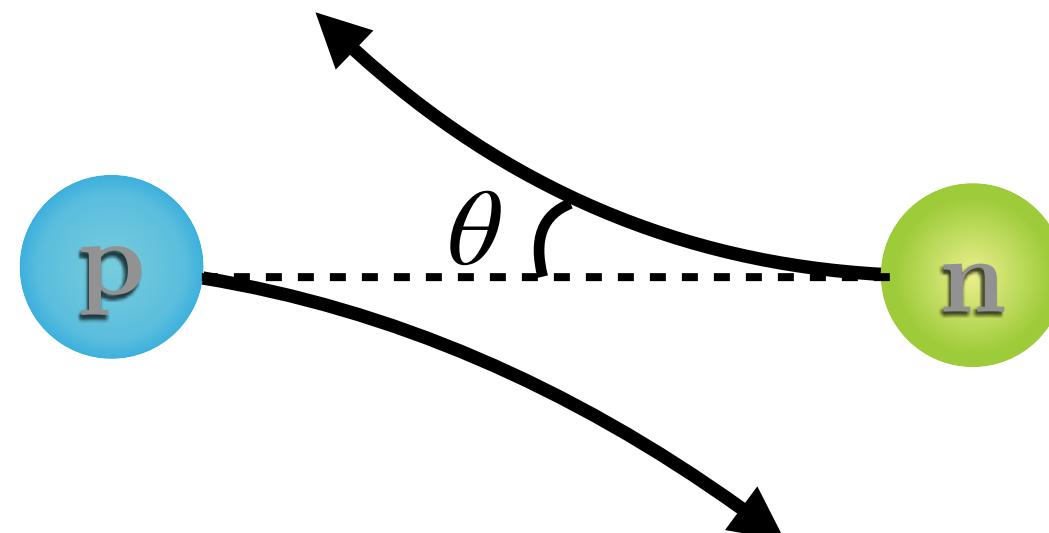
Our prior for the GP hyperparameters

Our posterior for the EFT expansion coefficients



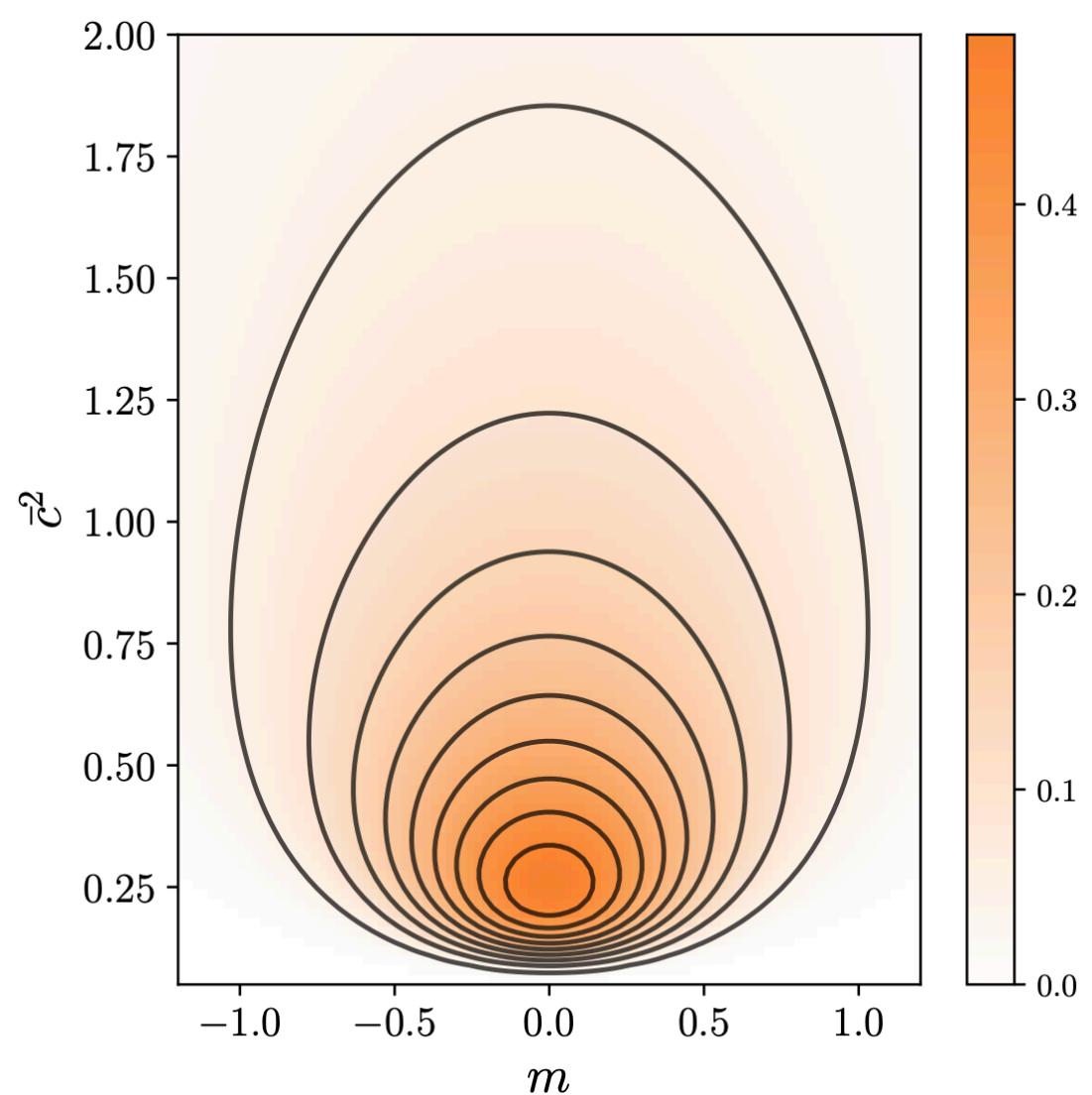
uniform distribution \rightarrow
for the lengthscales $\vec{\ell}$

A Gaussian-process model for correlated δM



Quantify expansion coefficients by evaluating the model at consecutive chiral orders for a 2D grid of \vec{x}_j values (energies T_{lab} and angles θ).
n.b. we do this for a sensible choice of $\vec{\alpha} = \vec{\alpha}^$*

We model the correlated expansion parameters as a two-feature Gaussian process using a SQE kernel



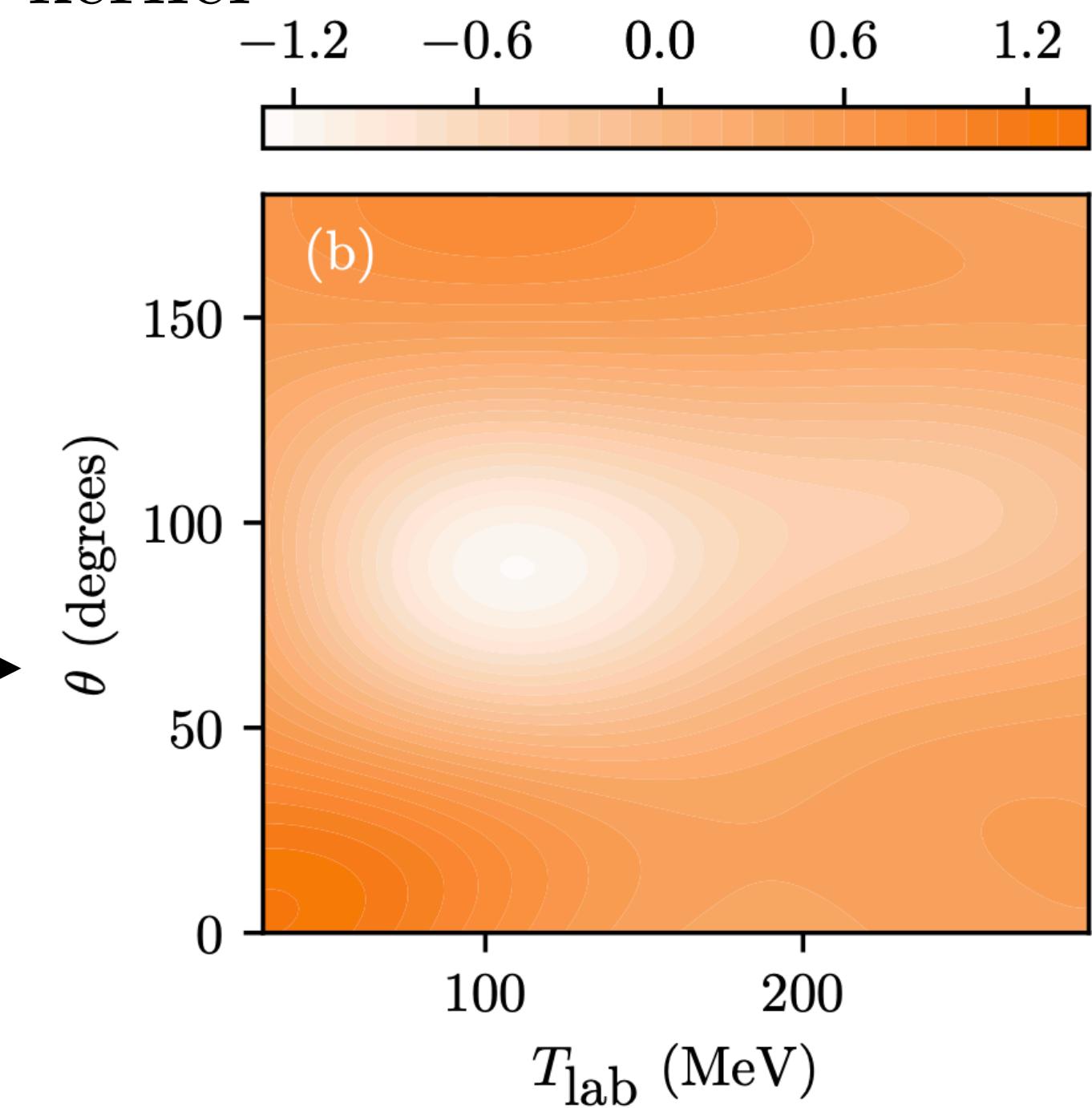
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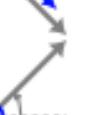
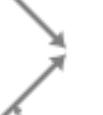
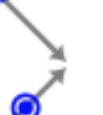
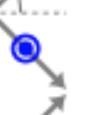
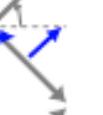
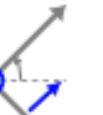
Our prior for the GP hyperparameters

Our posterior for the EFT expansion coefficients

$$\delta M \sim \mathcal{GP}[M, K]$$

uniform distribution
for the lengthscales $\vec{\ell}$



Notation	Definition	Acronym	$N_{d,y}$	$N_{T_{\text{lab}},y}$	n_{eff}	$\widehat{\ell}_{T_{\text{lab}}} \text{ (MeV)}$	\widehat{c}^2		
σ_{tot}	total cross section	SGT	119	113	26.7	61	0.61^2		
σ_T	$\sigma_{\text{tot}}(\uparrow\downarrow) - \sigma_{\text{tot}}(\uparrow\uparrow)$	SGTT	3	3	—	—	—		
σ_L	$\sigma_{\text{tot}}(\leftarrow\rightarrow) - \sigma_{\text{tot}}(\rightarrow\leftarrow)$	SGTL	4	4	3.7	59	2.28^2		
Notation	Tensor	Illustration	$N_{d,y}$	$N_{T_{\text{lab}},y}$	n_{eff}	$\widehat{\ell}_{T_{\text{lab}}} \text{ (MeV)}$	$\widehat{\ell}_\theta \text{ (deg)}$	\widehat{c}^2	
$\sigma(\theta)$	I_{0000}		DSG	1207	68	383.8	73	39	0.56^2
$A(\theta)$	D_{s0k0}		A	5	1	5.0	70	37	0.66^2
$A_t(\theta)$	K_{0ks0}		AT	30	2	23.0	60	36	0.61^2
$A_{yy}(\theta)$	A_{00nn}		AYY	58	4	19.4	60	33	0.97^2
$A_{zz}(\theta)$	A_{00kk}		AZZ	45	2	12.1	120	37	1.45^2
$D(\theta)$	D_{n0n0}		D	13	1	4.7	68	26	0.6^2
$D_t(\theta)$	K_{0nn0}		DT	36	5	33.6	41	45	0.52^2
$D_x^z(\theta)$	D_{0s0k}		D0SK	8	1	3.1	59	29	0.52^2
$A_y(\theta)$	P_{n000}		P	503	28	269.6	65	32	0.35^2
$N_{zz}^y(\theta)$	N_{0nk0}		NNKK	8	1	8.0	45	25	0.22^2
$N_{zy}^x(\theta)$	N_{0skn}		NSKN	12	1	11.0	83	36	0.85^2
$N_{xy}^x(\theta)$	N_{0ssn}		NSSN	4	1	4.0	78	31	0.62^2
$R(\theta)$	D_{s0s0}		R	5	1	5.0	74	32	0.67^2
$R_t(\theta)$	K_{0ss0}		RT	29	3	24.2	87	28	0.59^2
$R'_t(\theta)$	K_{0sk0}		RPT	1	1	1.0	62	35	0.54^2



We construct a model discrepancy for each observable type (y) we condition our inference on.

correlation lengths for np scattering energies and angles in the ranges **40–120 MeV and 25–45 degrees.**

small marginal variances

The likelihood

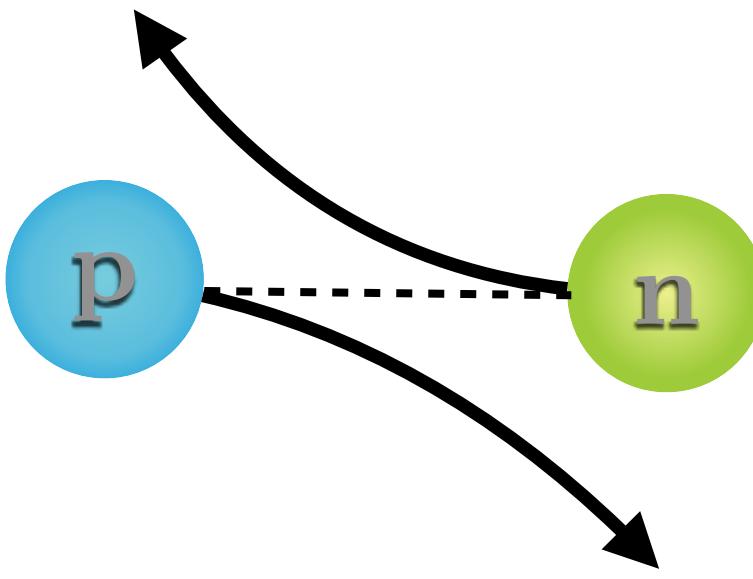
$$\boxed{\text{pr}(\mathcal{D}|\vec{\alpha}, I) = \prod_{y=1}^{N_y} \text{pr}(\mathcal{D}_y|\vec{\alpha}, I_y),}$$

$$\boxed{\text{pr}(\mathcal{D}_y|\vec{\alpha}, I_y) \propto \exp\left[-\frac{\mathbf{r}_y^T(\vec{\alpha})(\boldsymbol{\Sigma}_{\text{exp},y} + \boldsymbol{\Sigma}_{\text{th},y})^{-1}\mathbf{r}_y(\vec{\alpha})}{2}\right]}$$

$$\boxed{r_{y,j}(\vec{\alpha}) = [y_{\text{exp}}(\vec{x}_j) - y_{\text{th}}^{(k)}(\vec{\alpha}; \vec{x}_j)]}$$

$$\boxed{\boldsymbol{\Sigma}_{\text{th}} = \begin{bmatrix} \boldsymbol{\Sigma}_{\text{th},1} & \mathbf{0} & \cdots & \mathbf{0} \\ \mathbf{0} & \boldsymbol{\Sigma}_{\text{th},2} & \cdots & \mathbf{0} \\ \vdots & \vdots & \ddots & \vdots \\ \mathbf{0} & \mathbf{0} & \cdots & \boldsymbol{\Sigma}_{\text{th},N_y} \end{bmatrix}}$$

$$\boxed{(\boldsymbol{\Sigma}_{\text{th},y})_{mn} = \text{cov}[\delta y_{\text{th}}^{(k)}(\vec{x}_m), \delta y_{\text{th}}^{(k)}(\vec{x}_n)]}$$



The likelihood

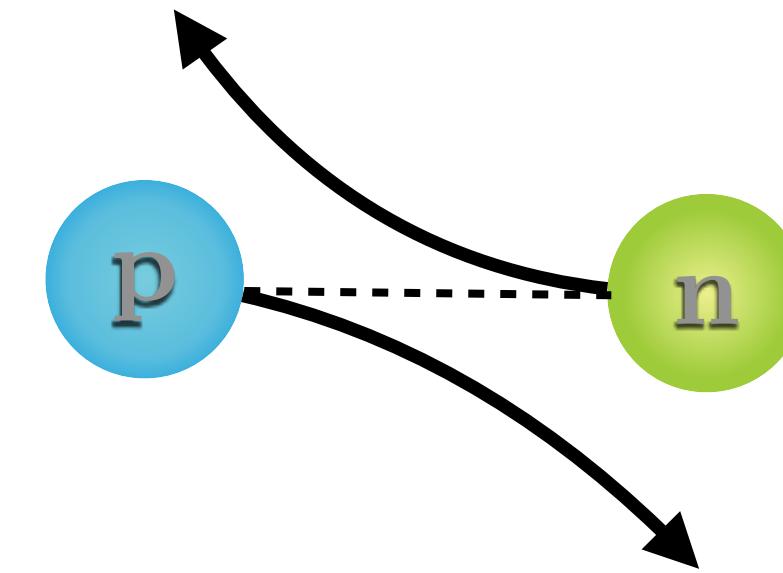
$$\text{pr}(\mathcal{D}|\vec{\alpha}, I) = \prod_{y=1}^{N_y} \text{pr}(\mathcal{D}_y|\vec{\alpha}, I_y),$$

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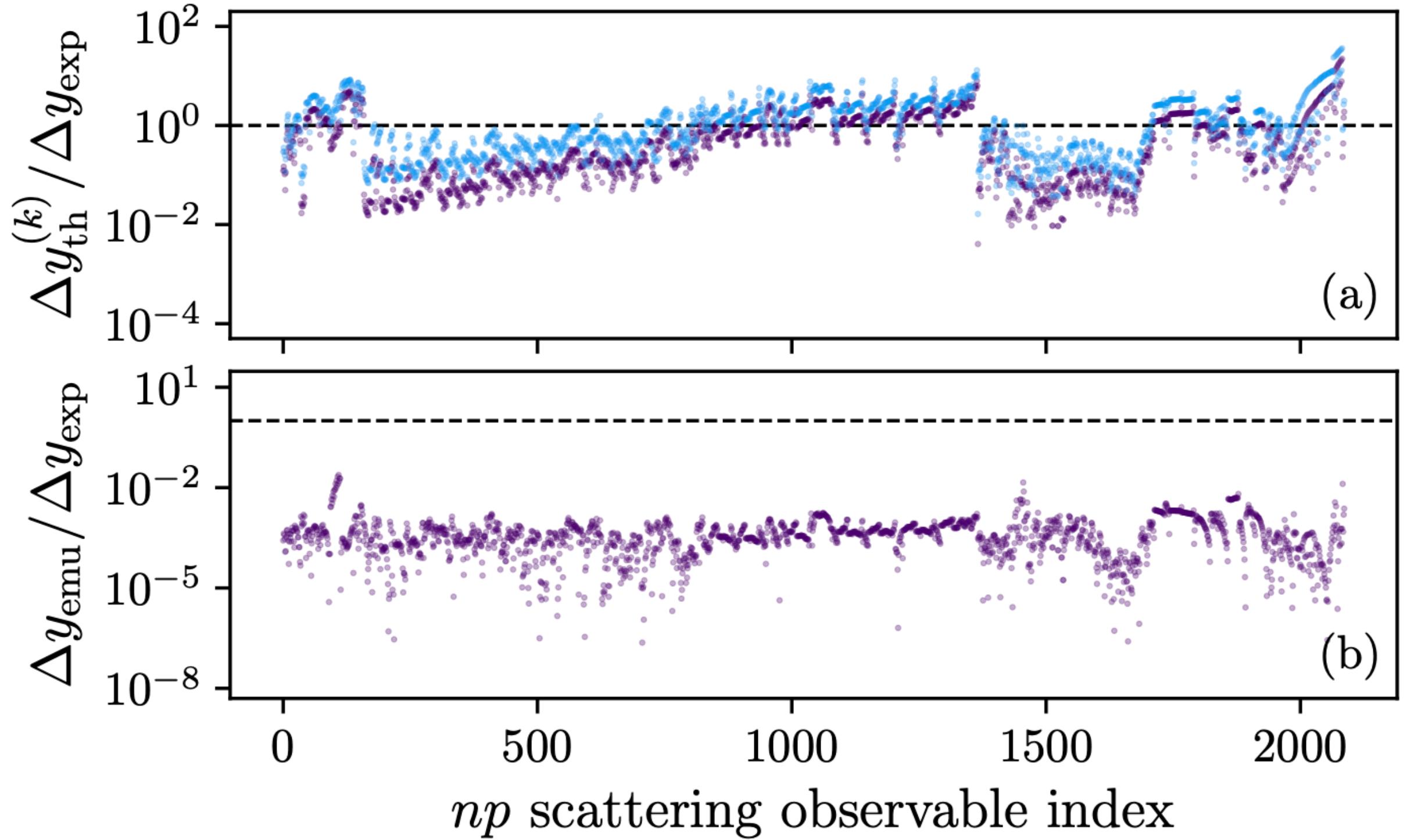
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$$(\Sigma_{\text{th},y})_{mn} = \text{cov}[\delta y_{\text{th}}^{(k)}(\vec{x}_m), \delta y_{\text{th}}^{(k)}(\vec{x}_n)]$$

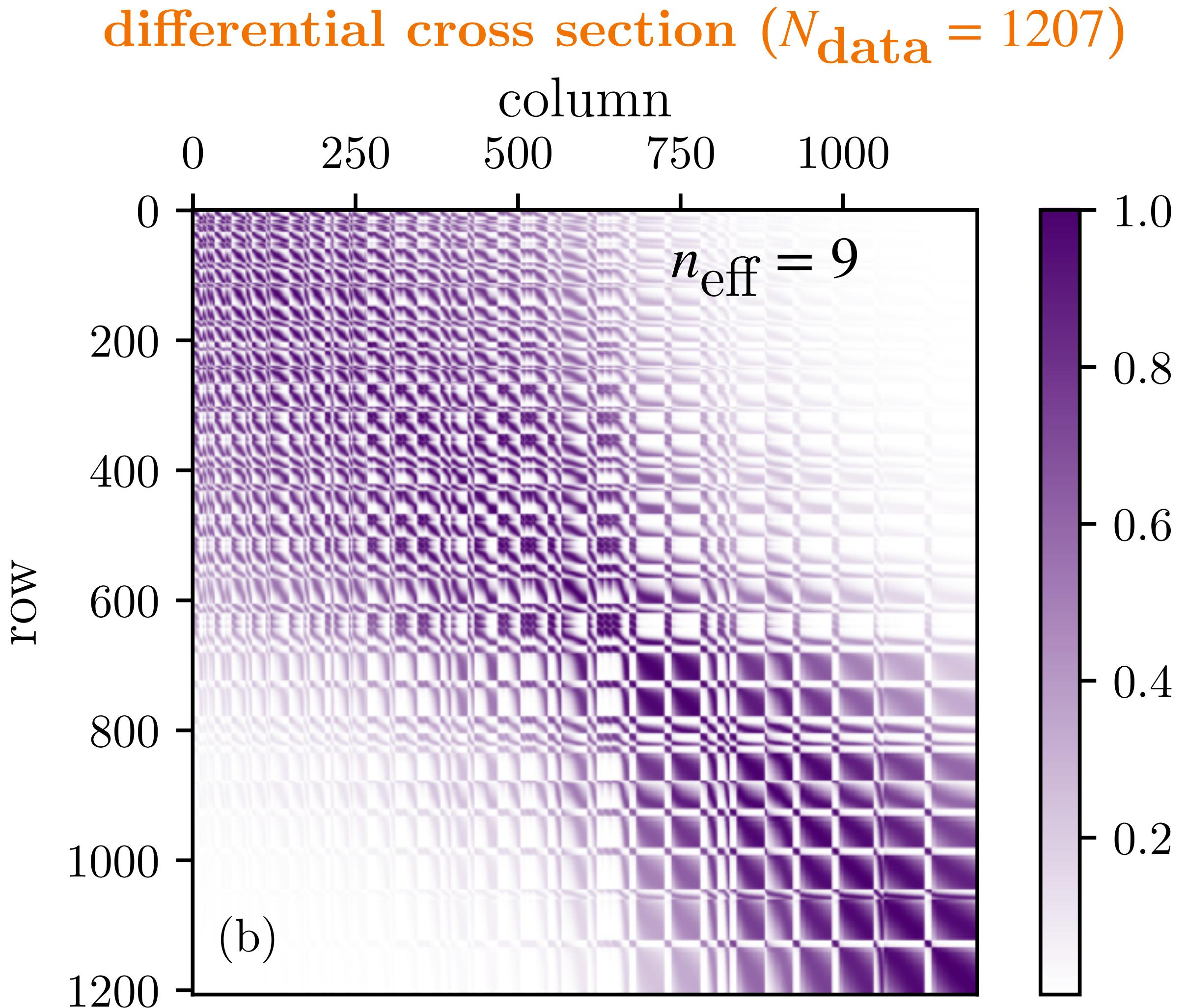
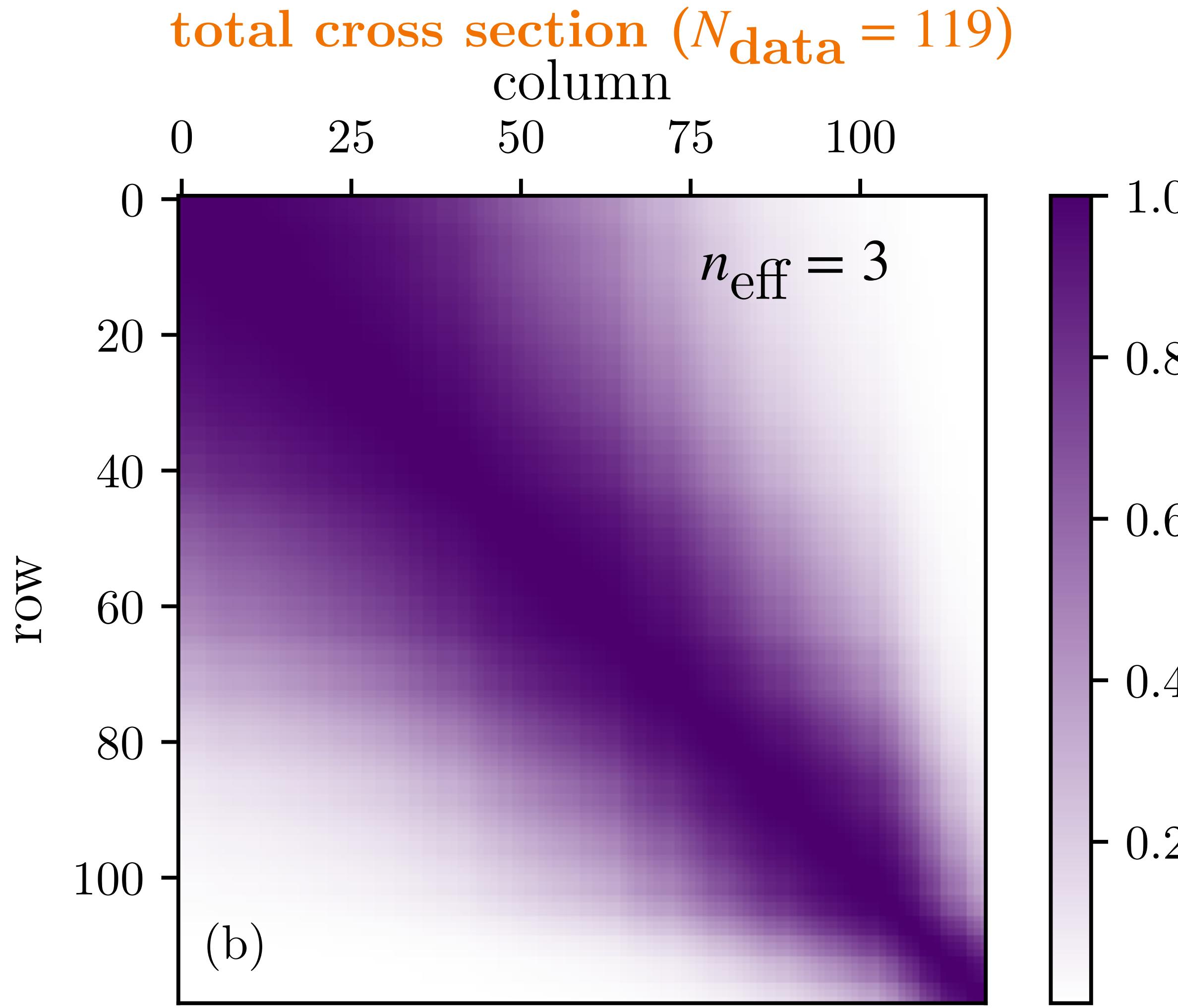


We emulate all ~ 2000 cross sections in \mathcal{D} .
Takes 1s (non-threaded + Google JAX)
to evaluate likelihood and its derivatives.

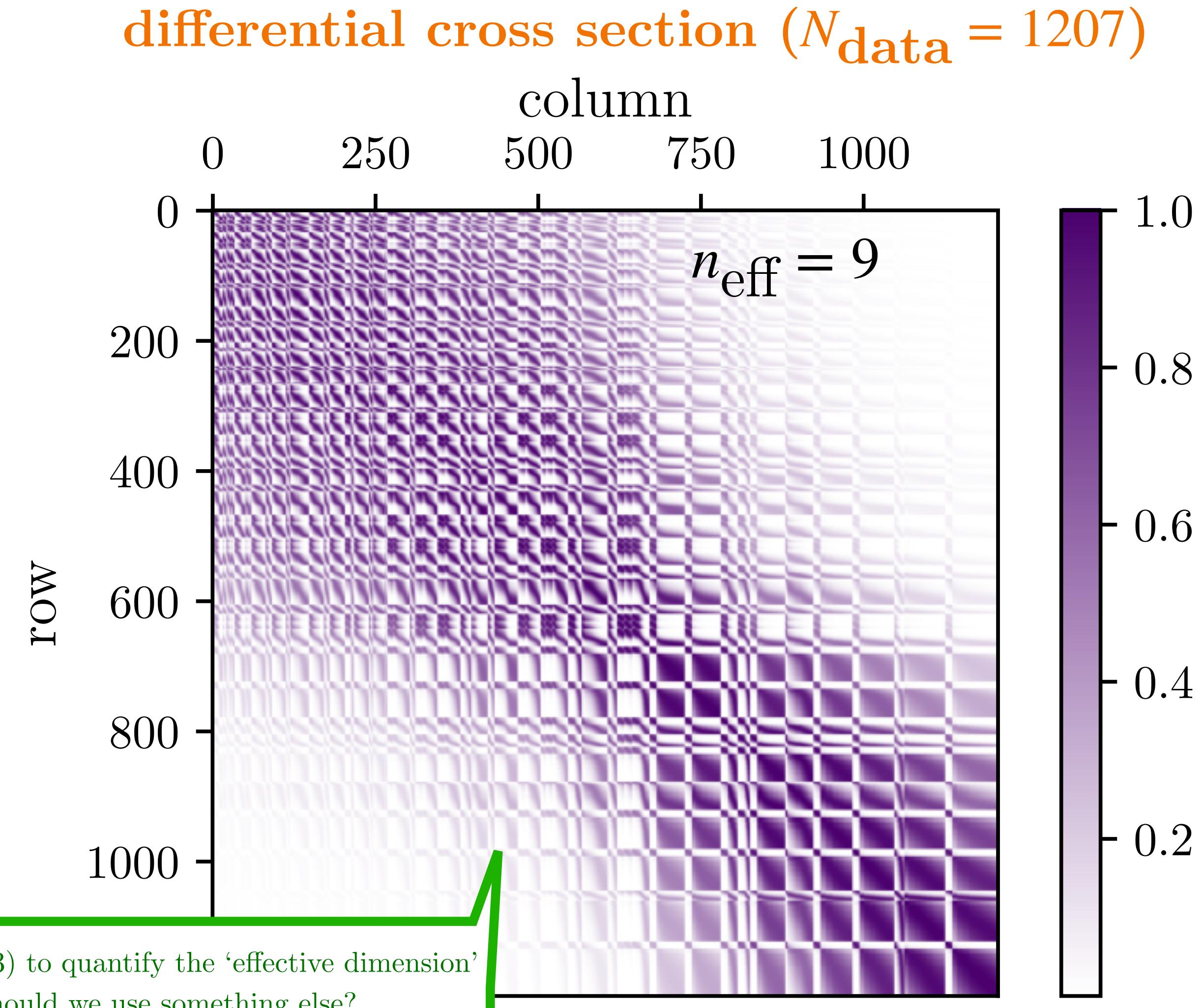
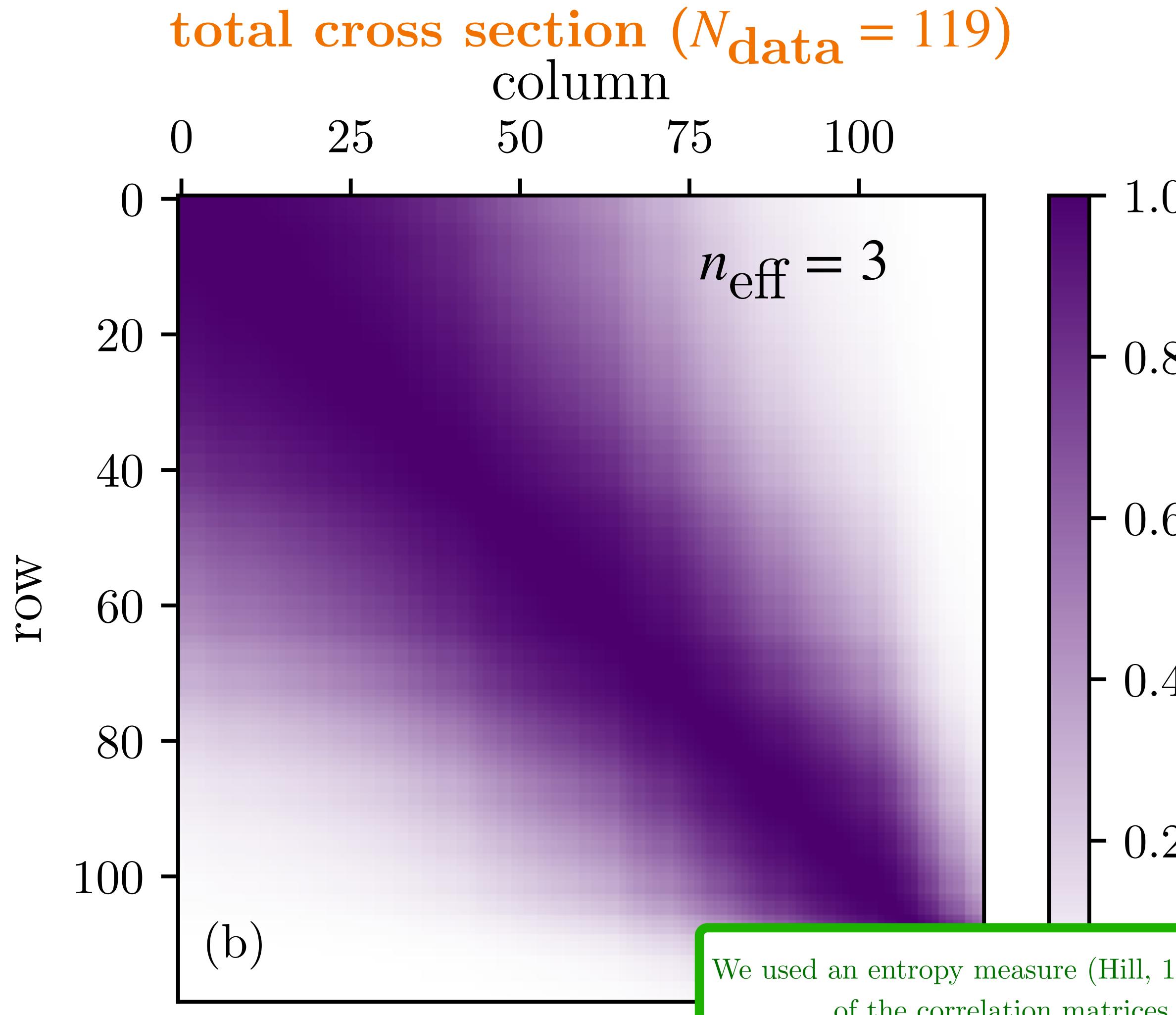


See R. Furnstahl's talk for more on emulators

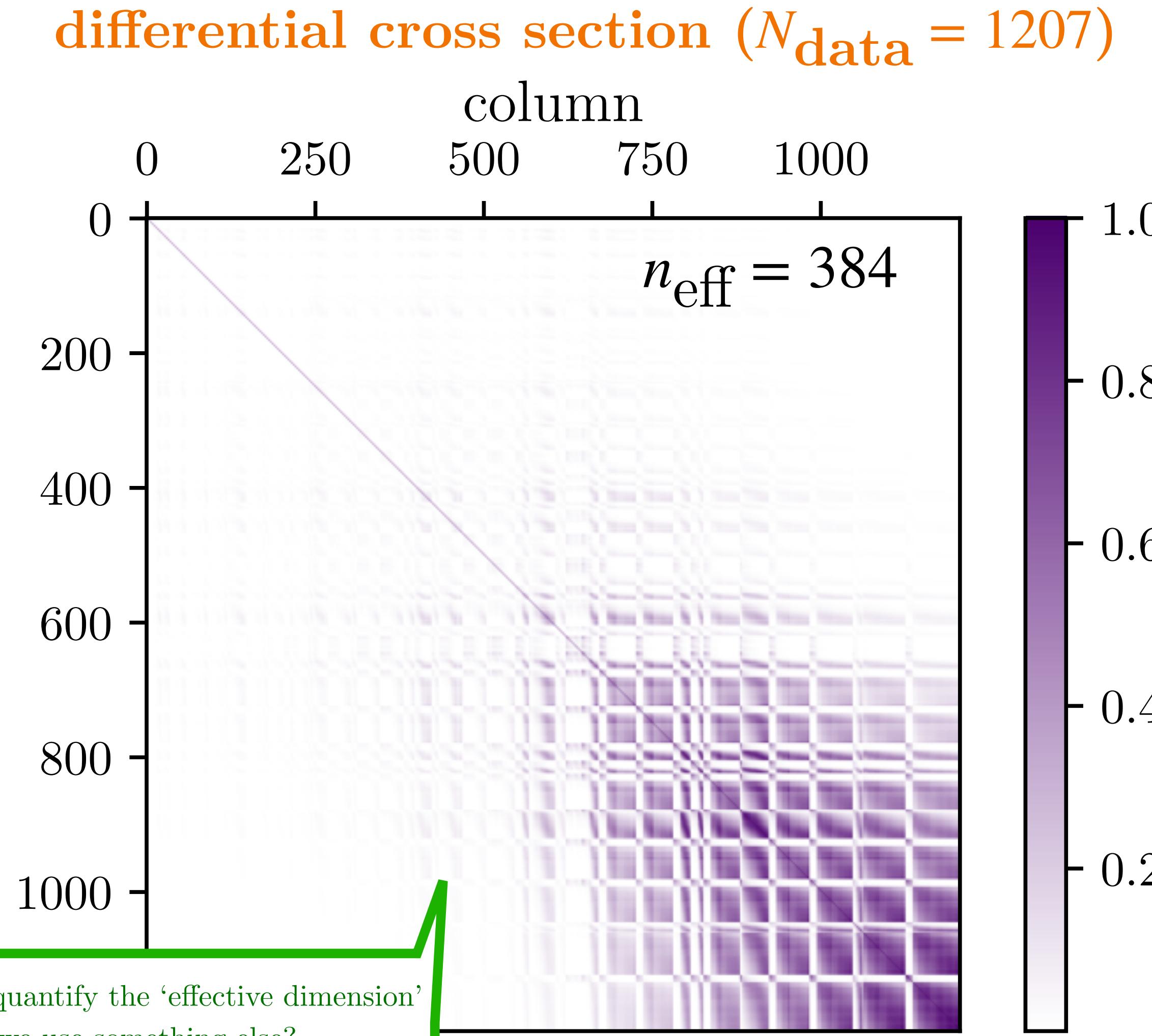
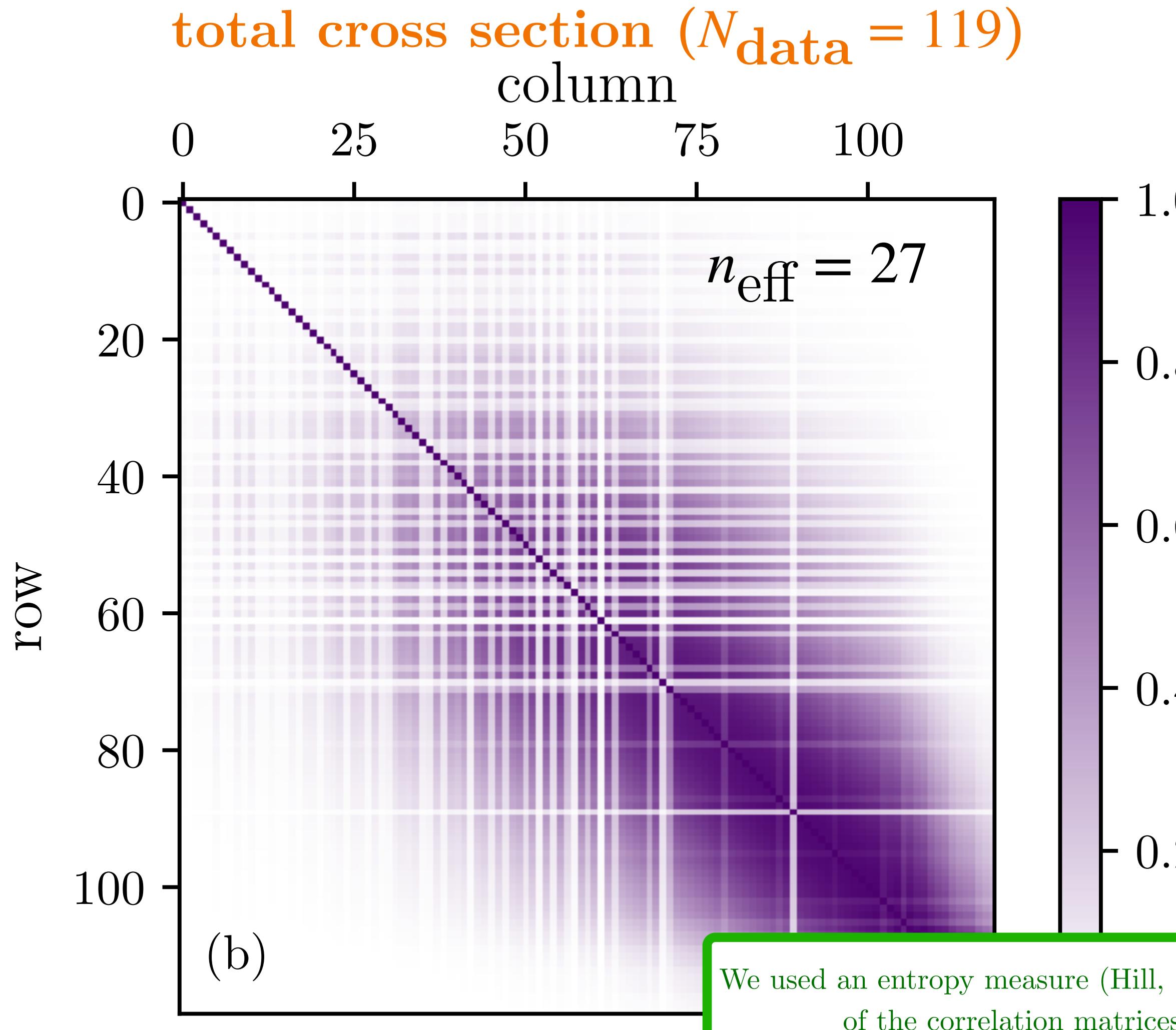
Effects of a small marginal variance in the model discrepancy



Effects of a small marginal variance in the model discrepancy



Effects of a small marginal variance in the model discrepancy



Bayes' rule: from likelihood & prior to posterior

- Collect N data points that we gather in a data vector D
- To explain the data, propose some model M , depending on parameters $\vec{\alpha}$
- Apply Bayes' rule

$$p(\vec{\alpha} | D, M, I) = \frac{p(D | \vec{\alpha}, M, I) \cdot p(\vec{\alpha} | M, I)}{p(D | M, I)}$$

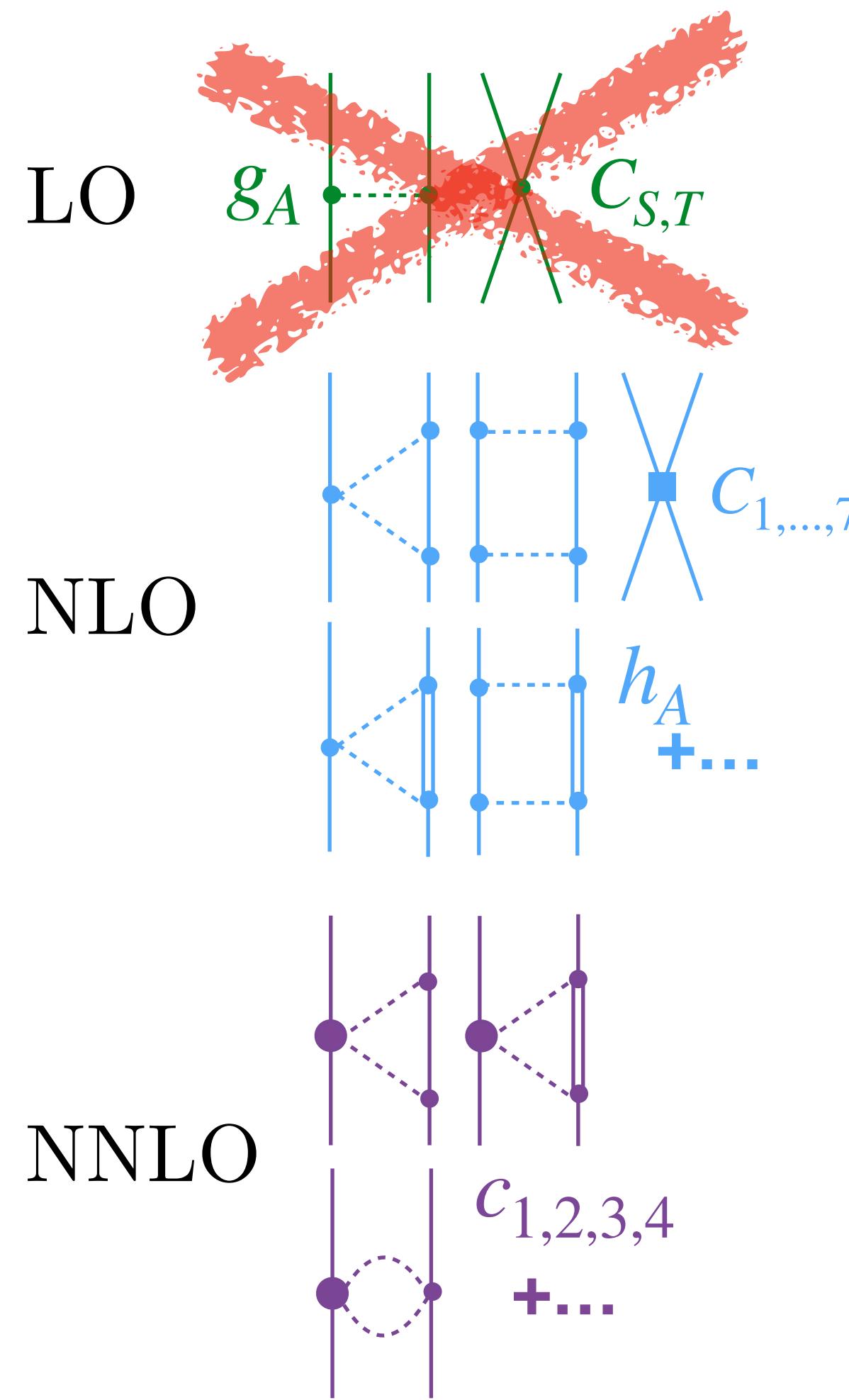
Marginal likelihood



most likely not Rev. T. Bayes

- The **prior** encodes our knowledge about the parameter values before analyzing the data
- The **likelihood** is the probability of the data given a set of parameters
- The **marginal likelihood** (or model evidence) provides normalization of the posterior
- The **posterior** is the complete inference and resulting probability density for the parameters $\vec{\alpha}$

Setting up the prior for the model params'



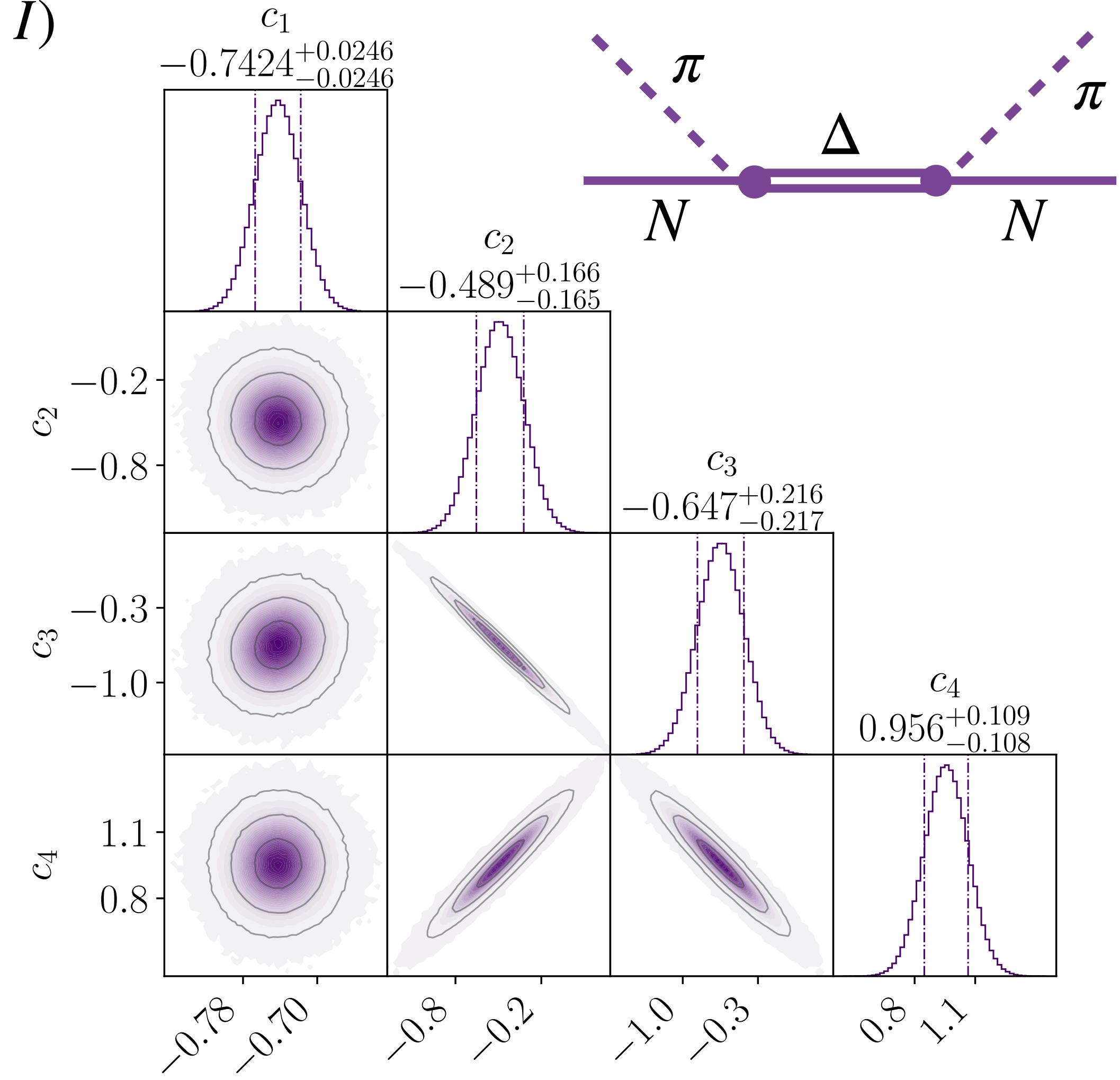
$$p(\vec{\alpha} | I) = p(\vec{\alpha}_{NN} | I) \cdot p(\vec{\alpha}_{\pi N} | I)$$

NN contacts: iid normal

$$\begin{aligned} p(\vec{\alpha} | I) &= p(\vec{\alpha}_{NN} | I) \\ &= \mathcal{N}(\mathbf{0}, \mathbf{1} \cdot 5^2) \end{aligned}$$

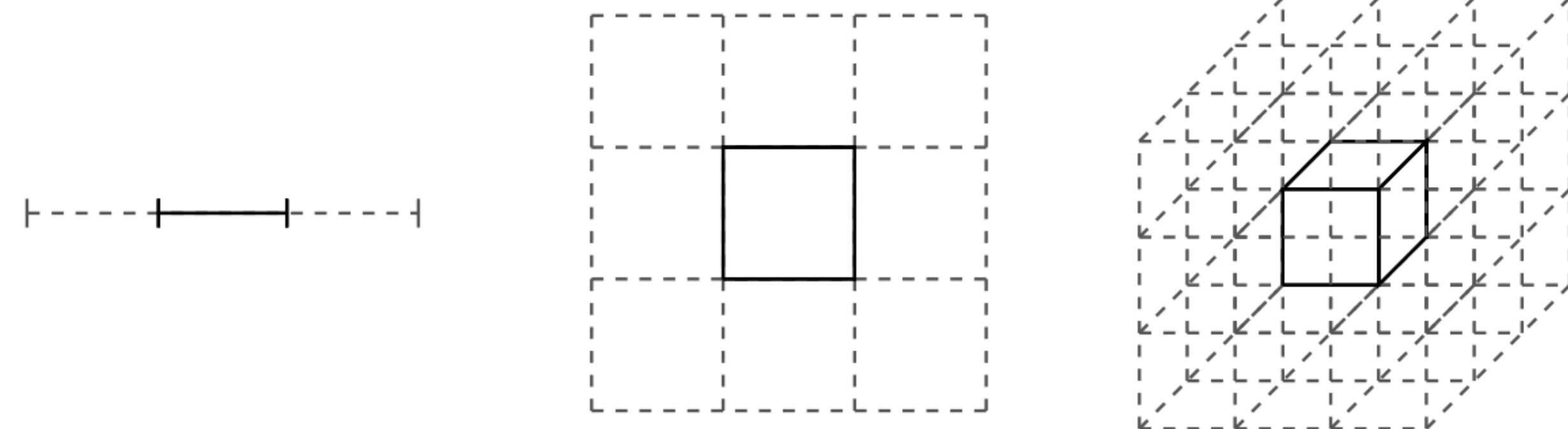
πN exchange: Roy-Steiner

$$p(\vec{\alpha} | I) = p(\vec{\alpha}_{\pi N} | I)$$





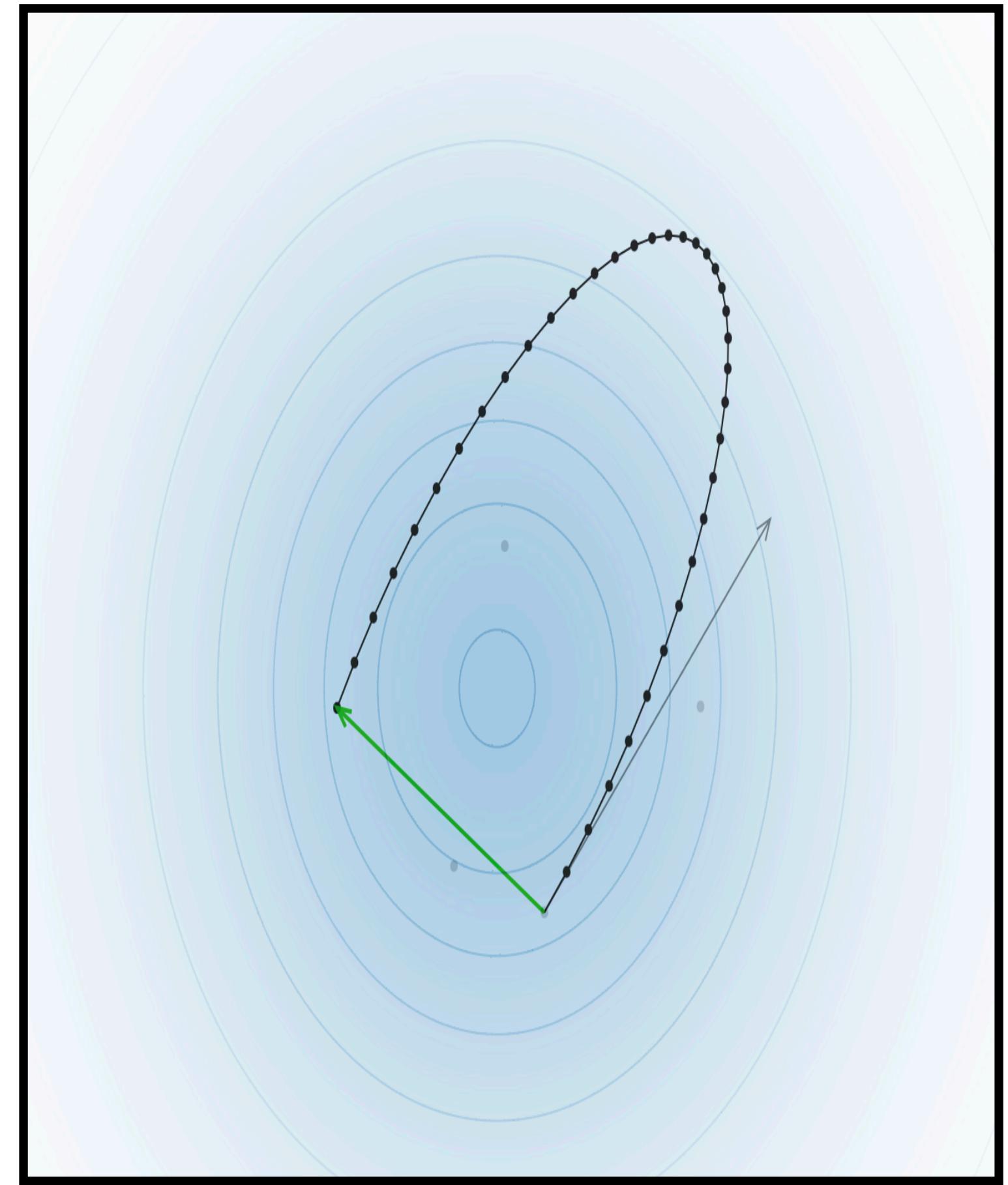
Hamiltonian Monte Carlo



credit: M. Betancourt arXiv:1701.02434

The LEC posterior is often multivariate (30 np/pp LECs at N3LO). Naive “guess and check” (random walk metropolis) will fail exponentially. We use Hamiltonian Monte Carlo (HMC) to take long jumps in parameter space while staying in regions with high probability mass.

See D. Mondal's talk for more on HMC



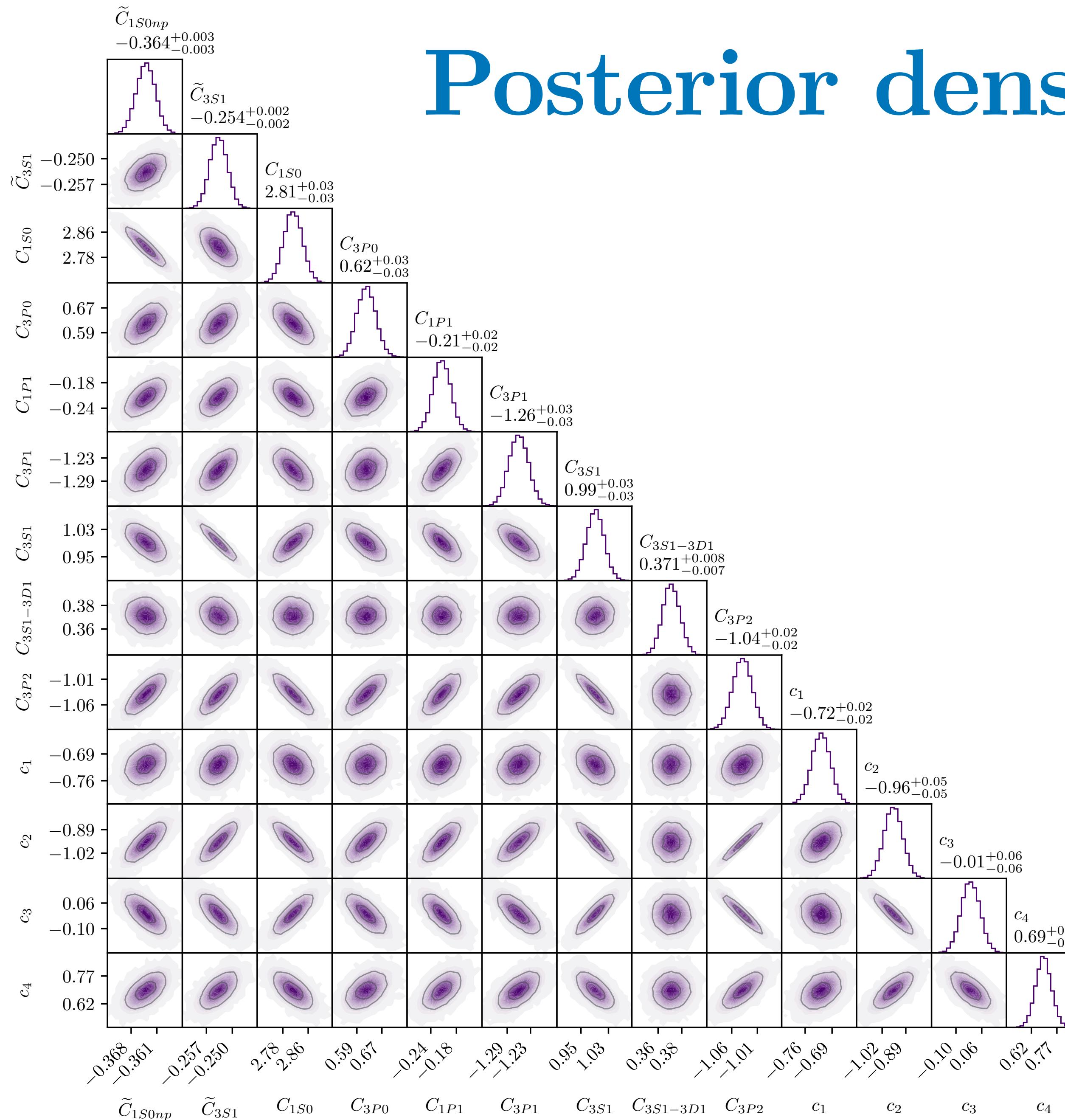
credit: <https://chi-feng.github.io/mcmc-demo/>

S. Duane, et al. Phys. Lett B **195**, 216 (1986)

I. Svensson, et al. Phys. Rev. C **105**, 014004(2022)

I. Svensson, et al. Phys. Rev. C **107**, 014001(2023)

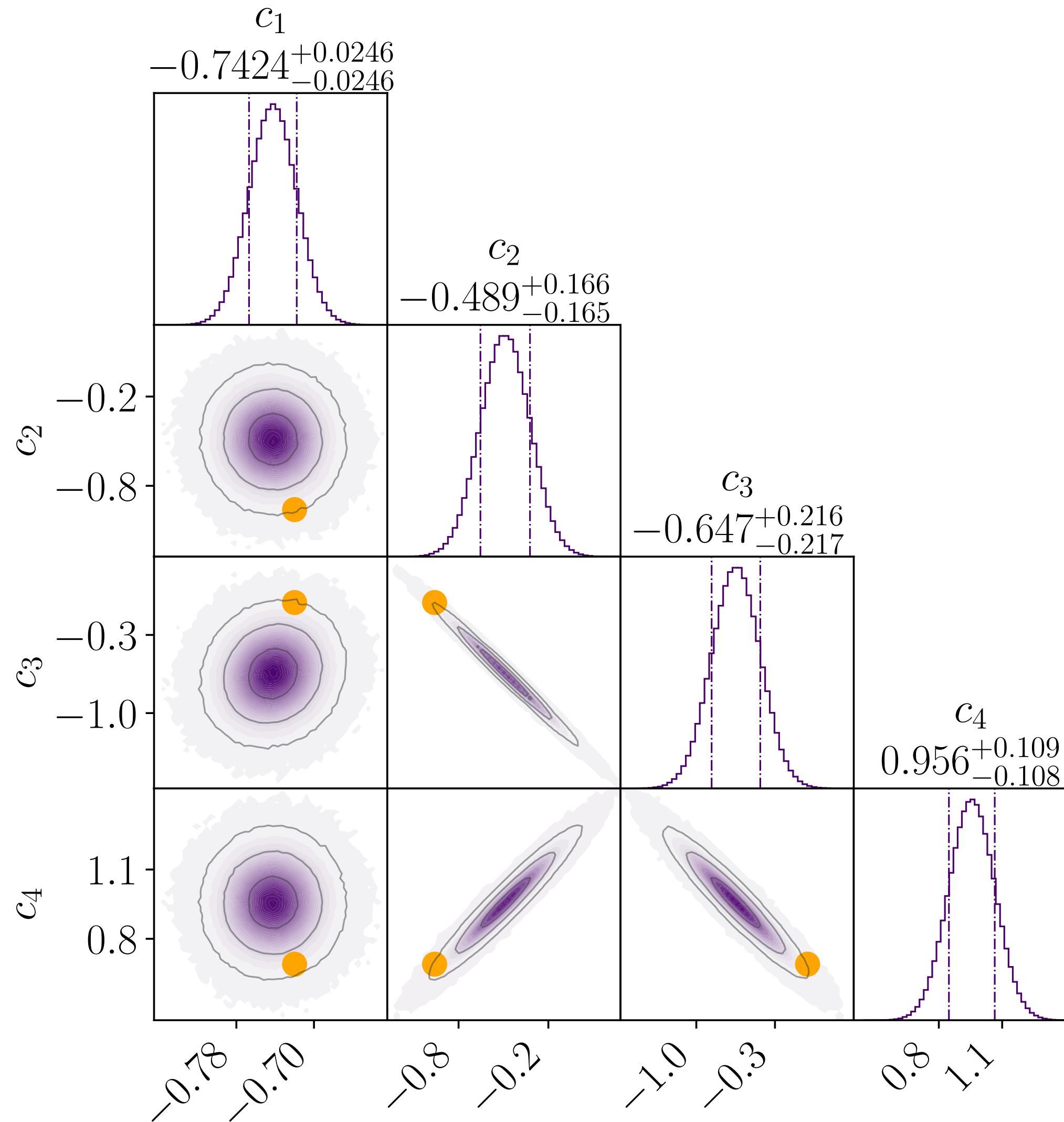
Posterior density (correlated δM)



credible ranges for the model parameters increase by a factor 1.5 when accounting for correlations in the model discrepancy

np data reduces the prior credible ranges for the πN parameters c_1, c_2, c_3, c_4 by a factor 3-5

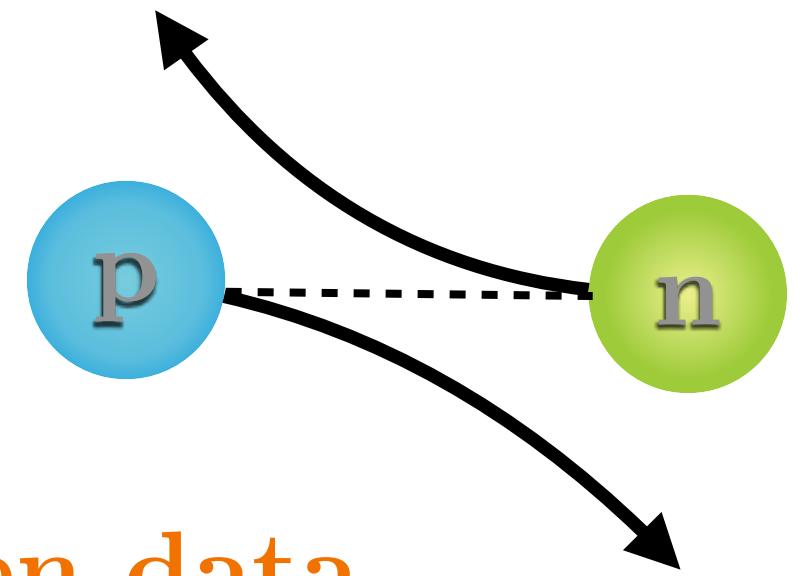
The πN prior vs posterior MAP



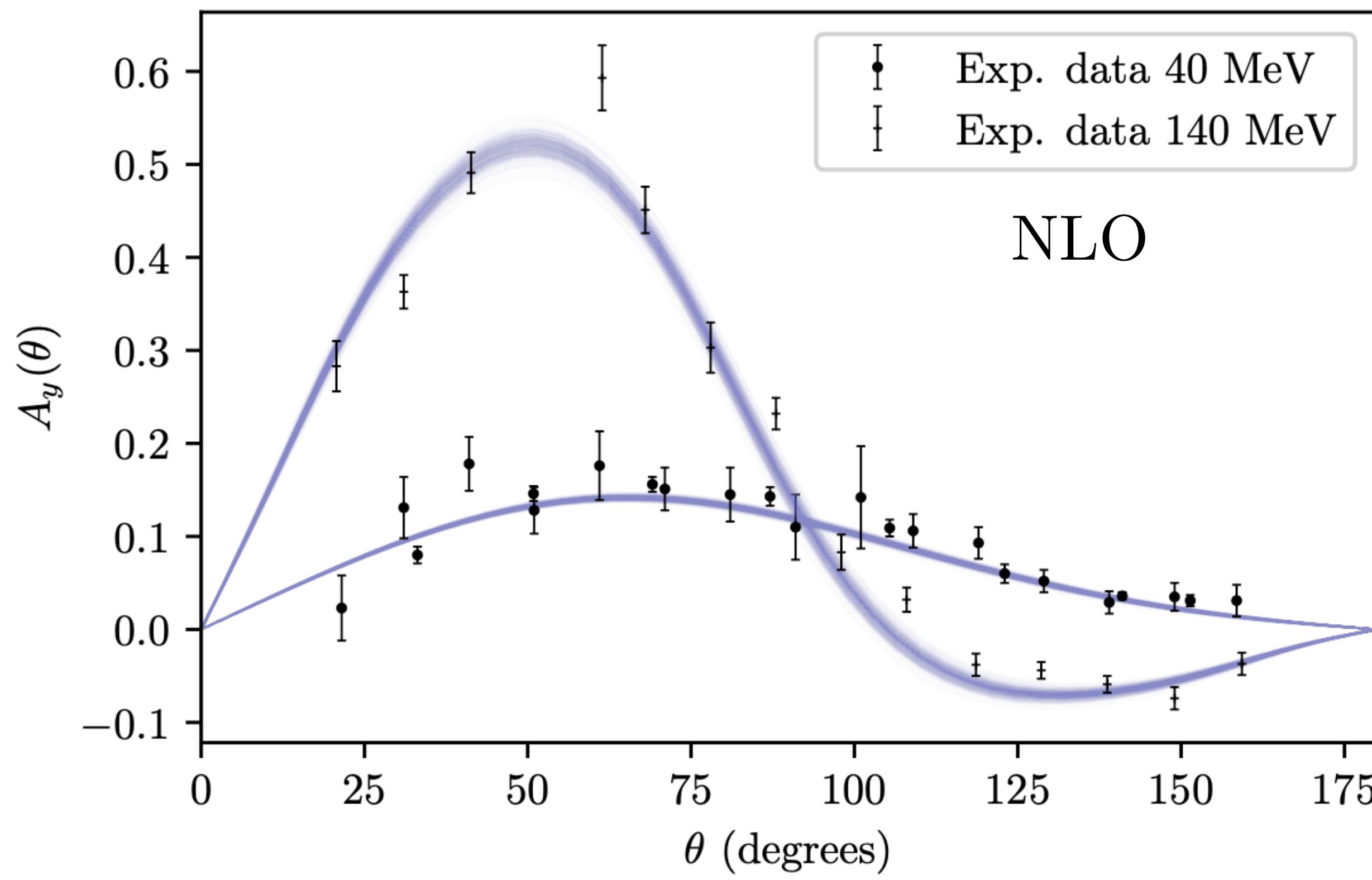
The Mahalanobis distance ($D_{MD}^2 = 9.95$) between the πN posterior MAP and prior (normal) mean is far enough to be outside 95% of the prior probability mass.

We need a better understanding of the model discrepancy and the underlying EFT.

Posterior predictive distributions

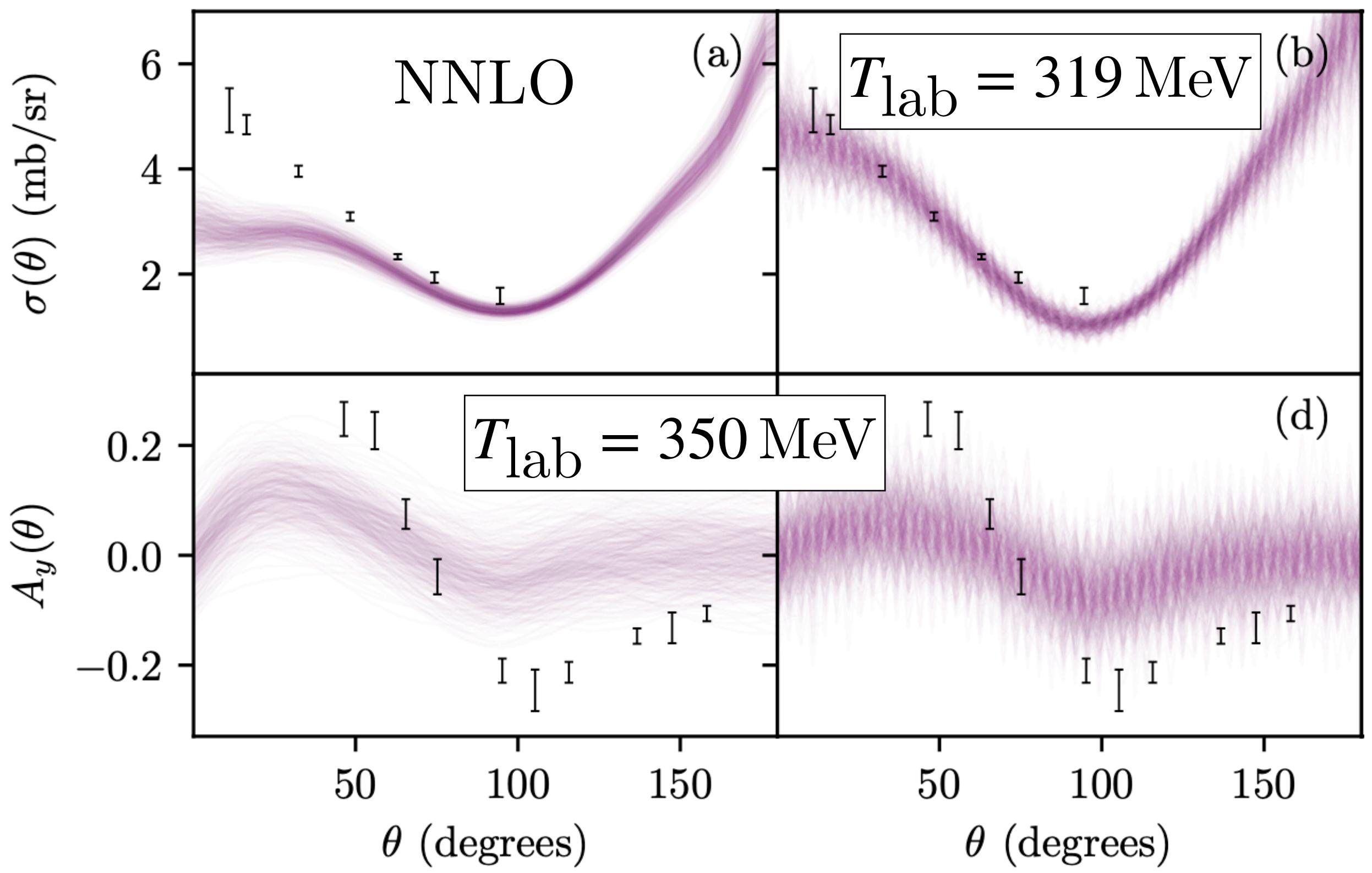


seen data



unseen data

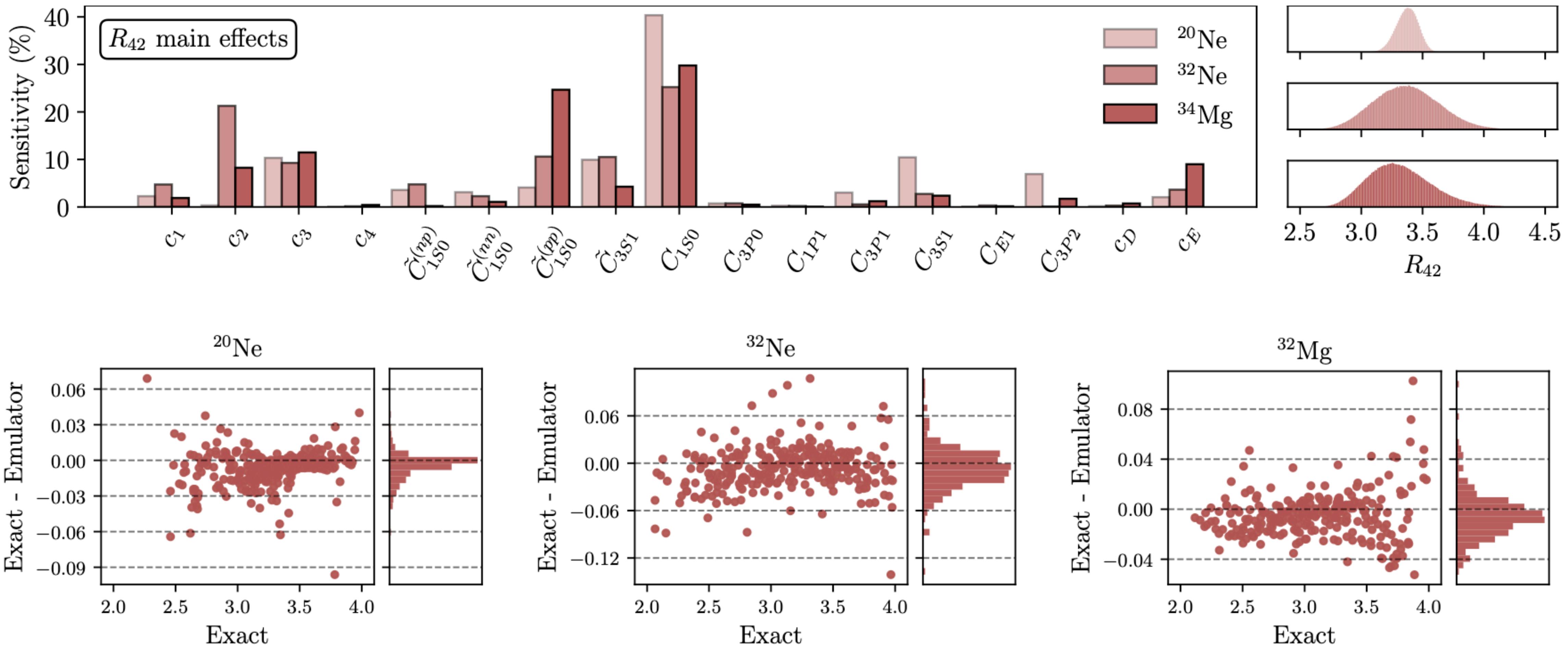
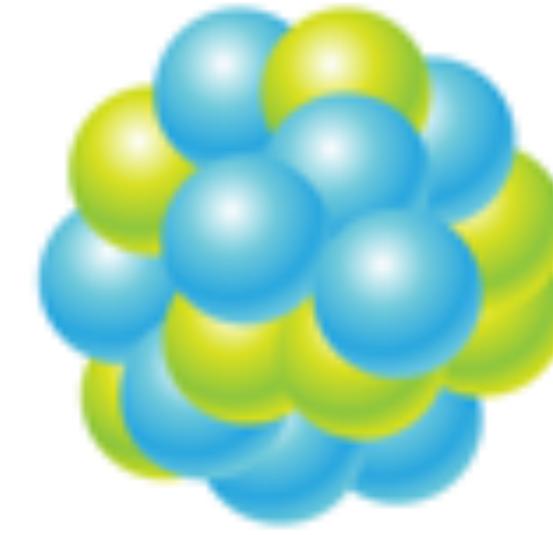
(some of the ‘bad ones’)



Global sensitivity analysis

Emulating excited state energies of deformed atomic nuclei

Nuclear Few-
and Many-Body
Problems

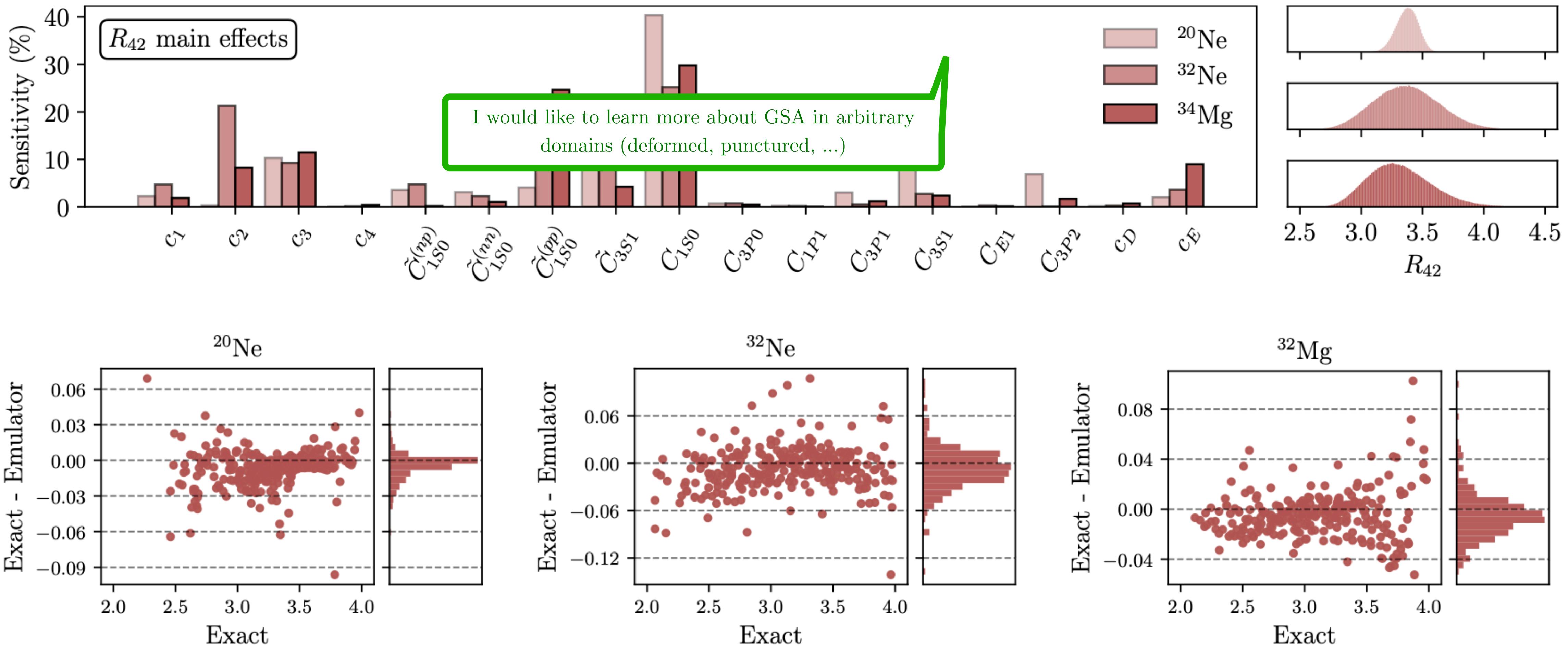
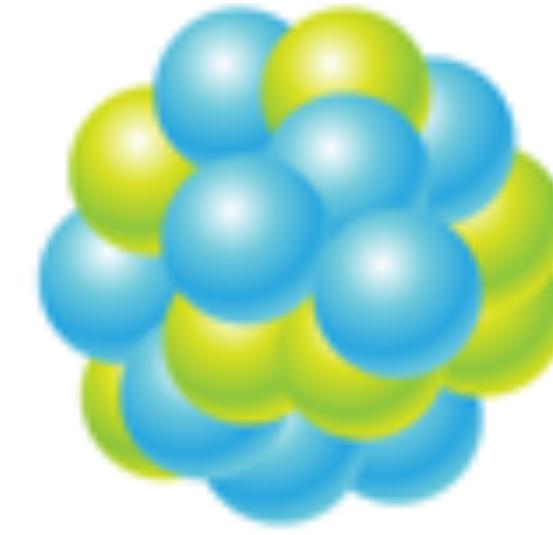


See K. Becker's talk for more on GSA

Global sensitivity analysis

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See K. Becker's talk for more on GSA

- we developed a Gaussian process to quantify the correlated model discrepancy (truncation error) of Δ -full chiral effective field theory at NLO and NNLO in np scattering. Much of the essential machinery is in place (open source!).
- we found that the *correlated* discrepancy has a small impact in Δ -full EFT up to NNLO. Small difference between NLO and NNLO -> underestimated prediction errors.
- problems abound:
 - tension between πN and NN sectors. (maybe just misspecified discrepancy)
 - sampling in high-dimensional spaces is always challenging.
 - do we have a sensible chiral EFT expansion? (likely not)
 - always a challenge to solve the many-body Schrödinger equation & quantify the associated method errors.
- emulators work like a charm and open up for many exciting and useful Bayesian analyses in *ab initio* nuclear physics!

Thanks for your attention!