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## **Distinguishing dark matter from millisecond pulsars with the Cherenkov Telescope Array**

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The leading explanation of the Fermi Galactic center gamma-ray excess is the extended emission from an unresolved population of millisecond pulsars (MSPs) in the Galactic bulge. Such a population would, along with the prompt gamma rays, also inject large quantities of electrons/positrons ( $e-e+$ ) into the interstellar medium. These  $e-e+$  could potentially inverse-Compton (IC) scatter ambient photons into gamma rays that fall within the sensitivity range of the upcoming Cherenkov Telescope Array (CTA). In this talk, I will highlight the unique capabilities of the Cherenkov Telescope Array to detect the expected IC signal from a putative population of 10 to 50 thousand MSPs in the center of our Galaxy. I will further show that, if an IC signal were detected, then CTA can successfully discriminate between a MSPs and a dark matter origin for the radiating  $e-e+$ .

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