

# Multicomponent Dark Matter and the Inert Doublet Model

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In collaboration with: Andrés Rivera and Guillermo Palacio



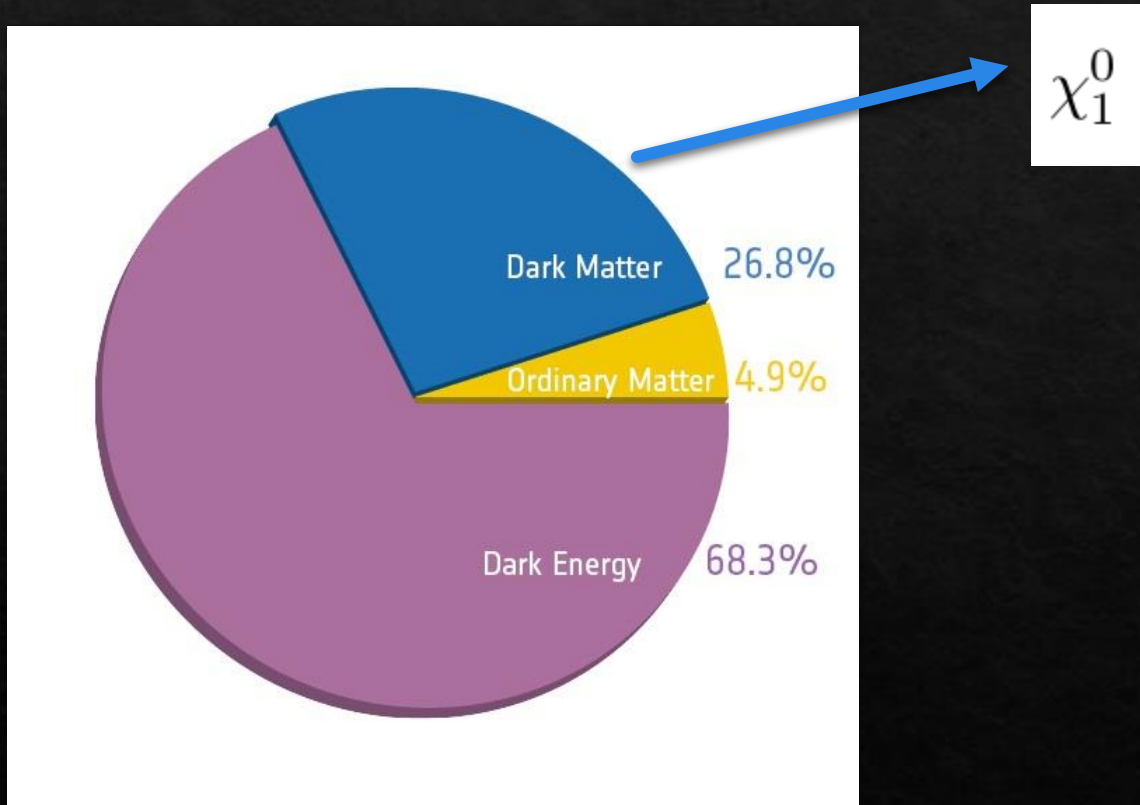
COMHEP 2020



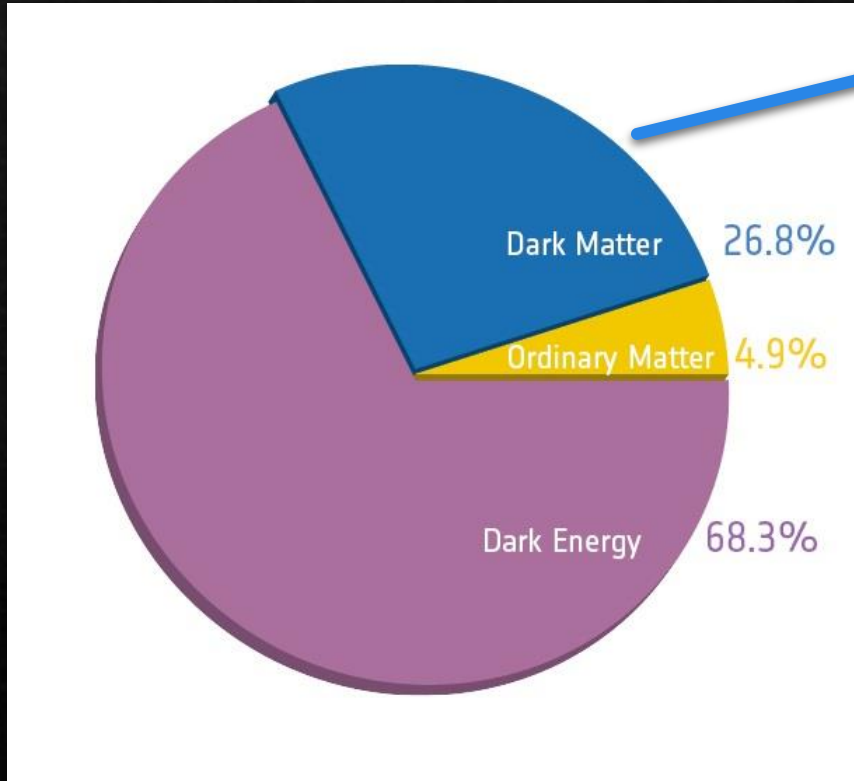
# Outline

- ◇ Motivation
- ◇ The models
- ◇ Phenomenology:
  - Relic density
  - Direct detection
  - Indirect detection
- ◇ Conclusions

# Motivation for Multicomponent Dark Matter



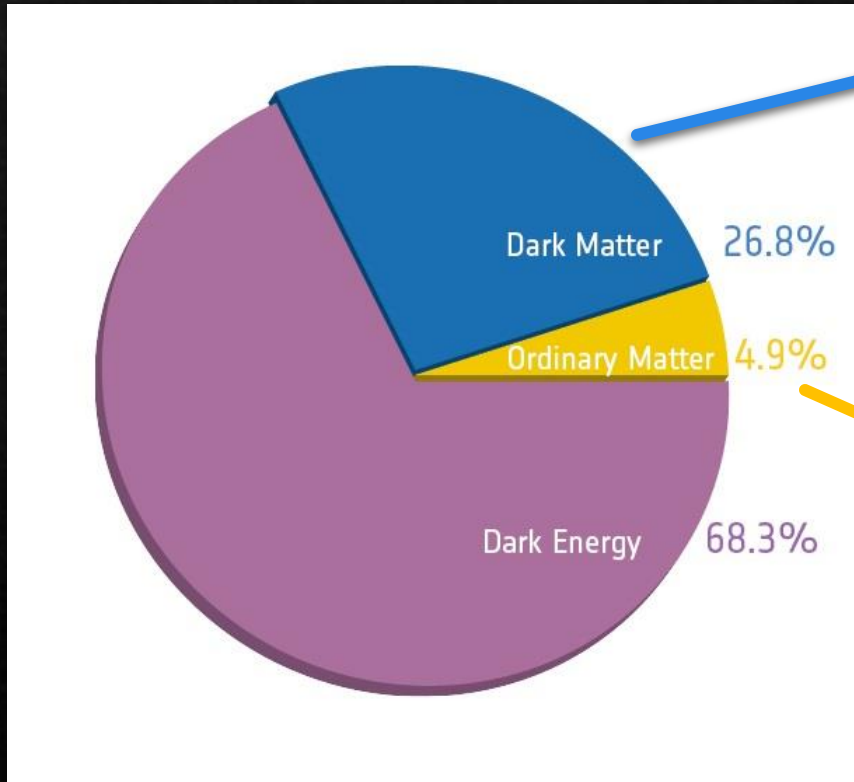
# Motivation for Multicomponent Dark Matter



$\chi_1^0$   
 $\nu_e \nu_\mu \nu_\tau$



# Motivation for Multicomponent Dark Matter



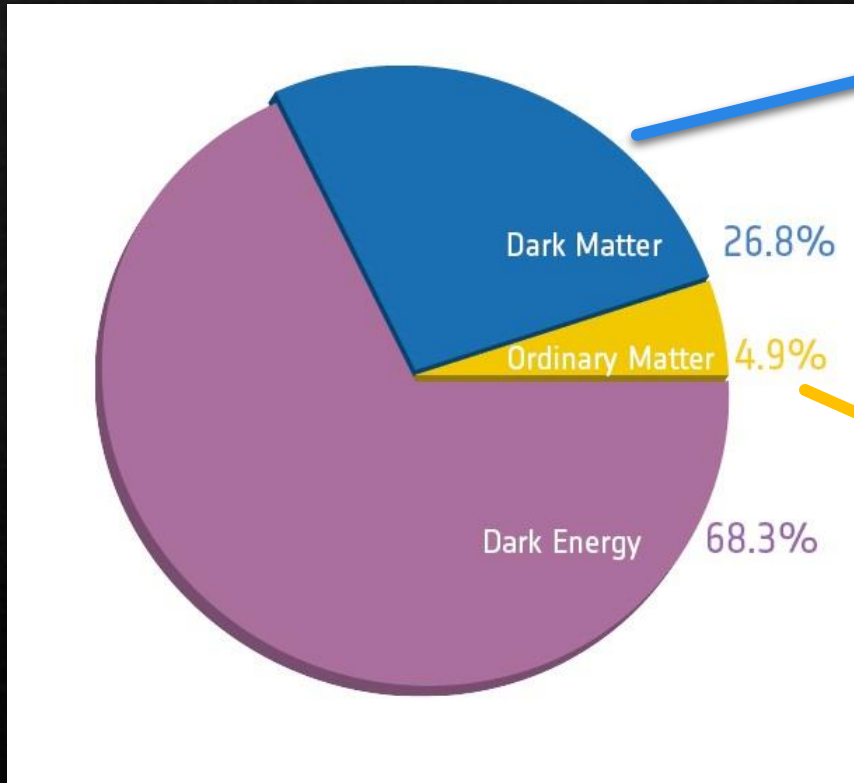
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### The Standard Model of Particle Physics

	FERMIONS (matter particles)			BOSONS (force carriers)	
QUARKS	$u$ up	$c$ charm	$t$ top	$g$ gluon	$H$ Higgs boson
	$d$ down	$s$ strange	$b$ bottom	$\gamma$ photon	
	$e$ electron	$\mu$ muon	$\tau$ tau	$Z^0$ Z boson	
LEPTONS	$\nu_e$ electron neutrino	$\nu_\mu$ muon neutrino	$\nu_\tau$ tau neutrino	$W^\pm$ W boson	

sciencealert

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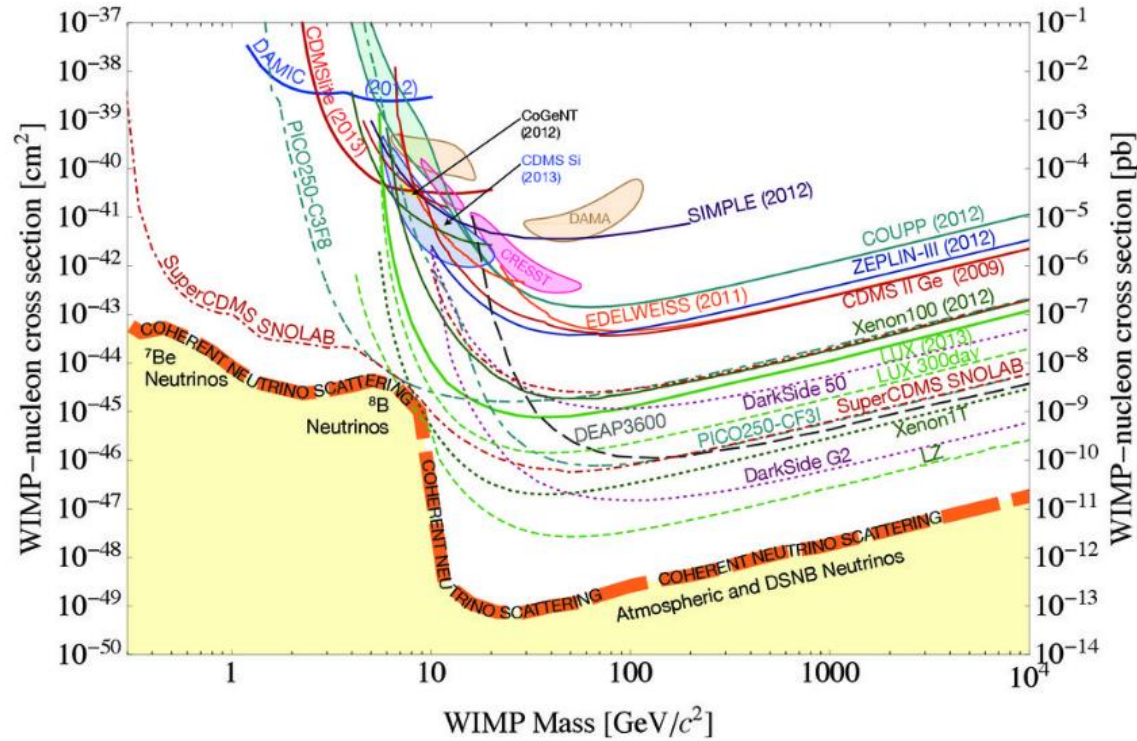
DM could be made up of a myriad of particles

### The Standard Model of Particle Physics

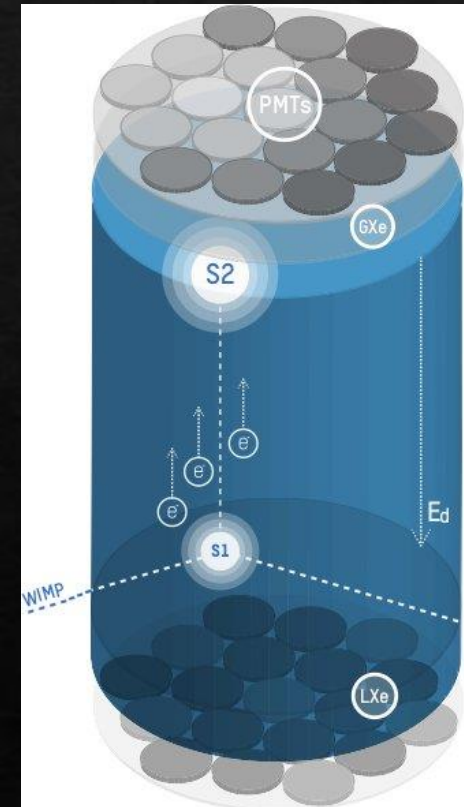
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# Advantages of Multicomponent dark sectors: Direct detection constraints

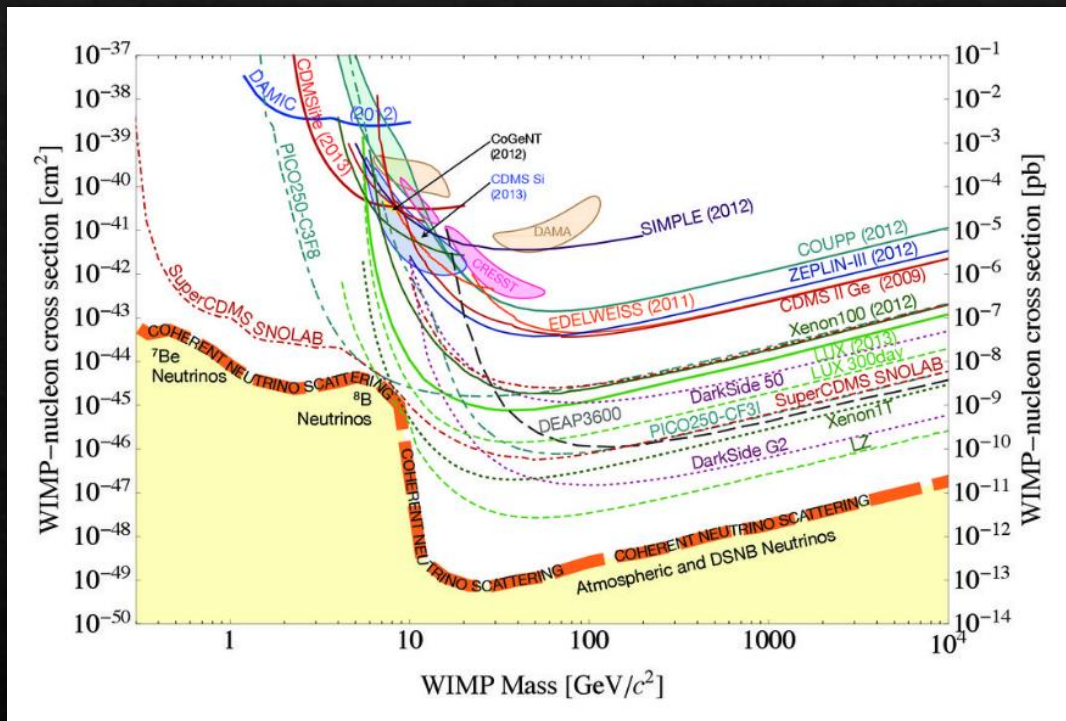


WIMP direct detection cross section is getting more constrained





# Advantages of Multicomponent dark sectors: Direct detection constraints



WIMP direct detection cross section is getting more constrained

Multicomponent dark matter is less constrained because direct detection restrictions must be rescaled



# The Inert Doublet Model

- ◇ Extend the scalar sector with a new doublet and a new  $Z_2$  symmetry:

New gauge interactions

New Higgs interactions

$$H_2 = \begin{pmatrix} H^+ \\ \frac{H^0 + iA^0}{\sqrt{2}} \end{pmatrix}$$

$$\mathcal{V}_I = \lambda_3 |H_1|^2 |H_2|^2 + \lambda_4 |H_1^\dagger H_2|^2 + \frac{\lambda_5}{2} [(H_1^\dagger H_2)^2 + \text{h.c.}]$$

$$m_{H^0}^2 = \mu_2^2 + \lambda_L v^2$$

$$\lambda_L = \frac{\lambda_3 + \lambda_4 + \lambda_5}{2}$$

DM candidate

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DM candidate

Interesting phenomenology  
but hard to probe around the  
electroweak scale!

# Fermion sector, the recipe

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- ◇ Singlet-doublet
- ◇ Doublet-triplet
- ◇ Singlet-triplet

# The fermion sector

Singlet-Doublet

$\Psi, N$

Doublet-Triplet

$\Psi, \Sigma$

Singlet-Triplet

$N, \Sigma$



# The fermion sector

Singlet-Doublet

$\Psi, N$

Doublet-Triplet

$\Psi, \Sigma$

Singlet-Triplet

$N, \Sigma, \Omega$

New scalar that mixes with the Higgs, not scalar DM

# The fermion sector

Model 1:  
Singlet-Doublet

$$\Psi, N$$

Model 2:  
Doublet-Triplet

$$\Psi, \Sigma$$

Model 3:  
Singlet-Triplet

$$N, \Sigma, \Omega$$

Fermion Mixing

$$\chi_1^0$$

Dark matter candidate  
plus other fermionic  
guys

New scalar that  
mixes with the  
Higgs, not scalar  
DM

# Phenomenology

- ◇ Check theoretical constraints for all models.

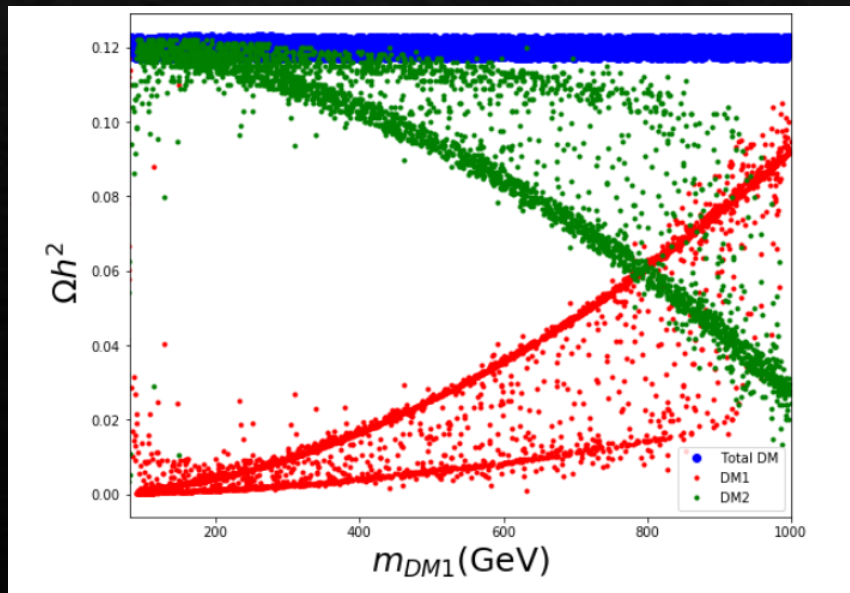


# Phenomenology

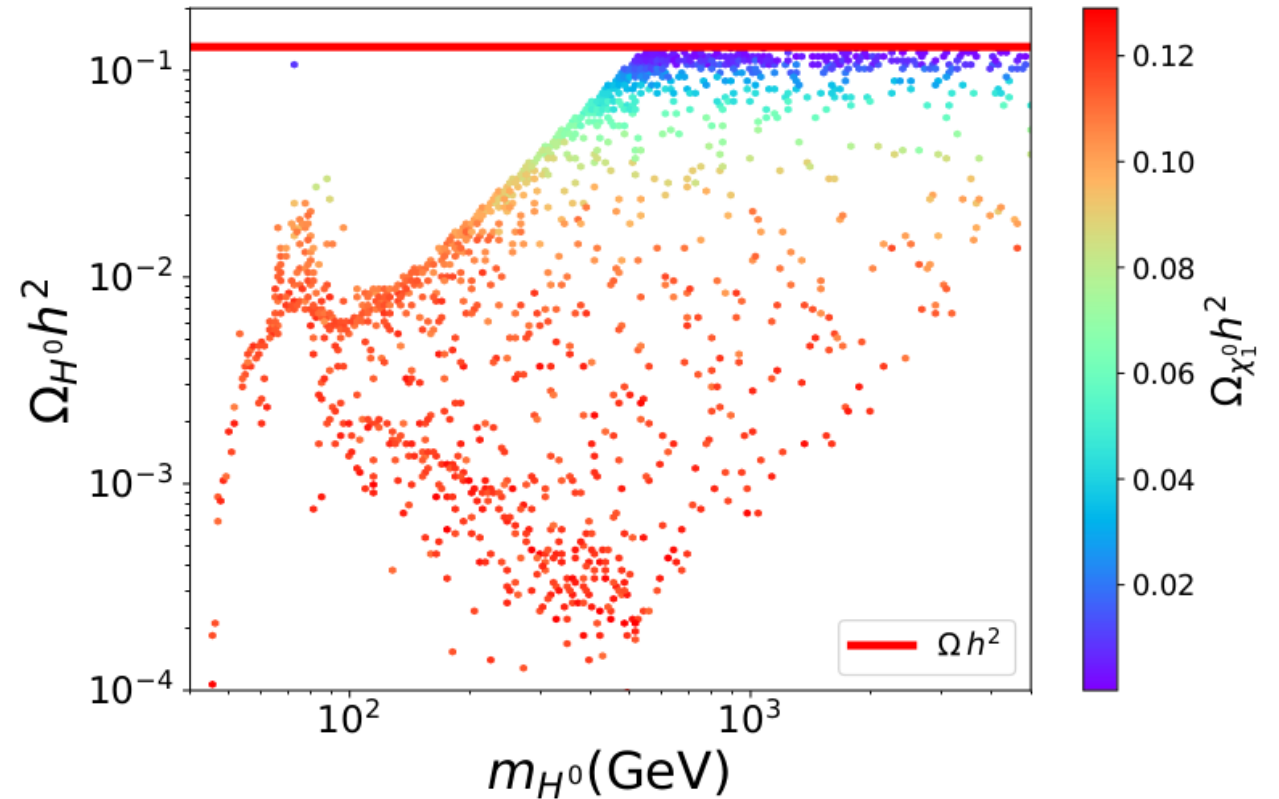
- ◆ Check theoretical constraints for all models.

Implement all models in **SARAH** use  
**SPheno**, **MicrOmegas** and **MadGraph**.  
Perform checks!

# Relic density Model 1: Singlet Doublet

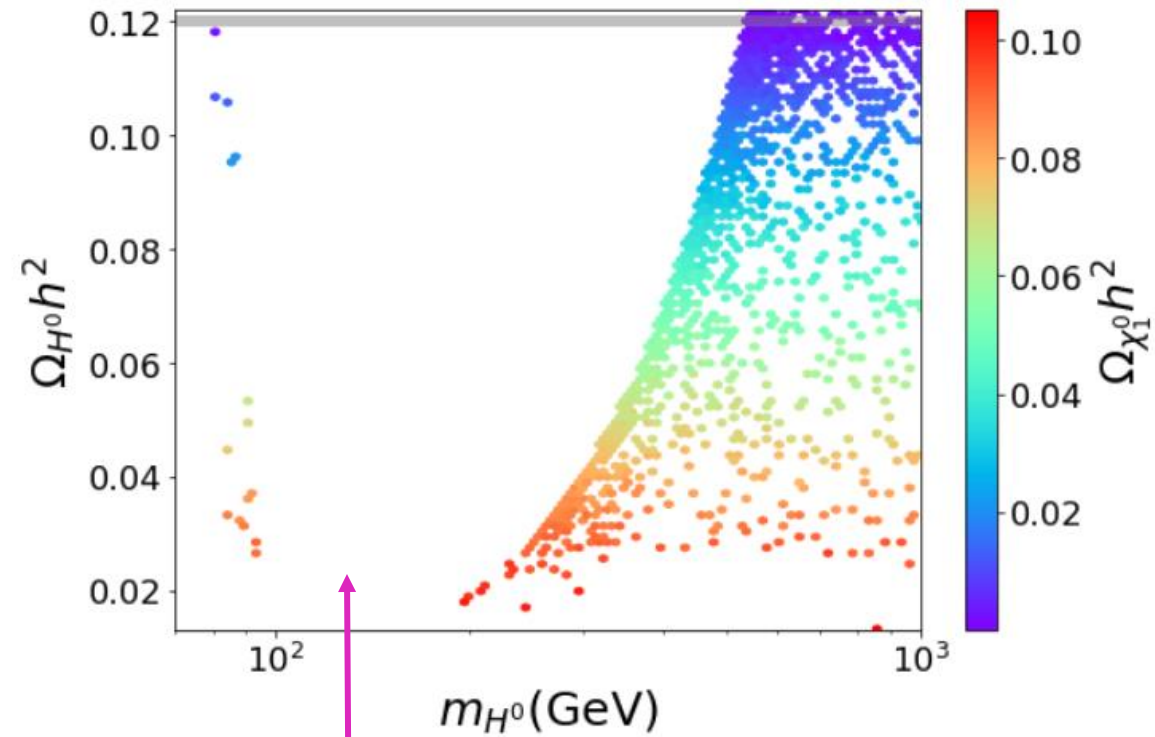


All points yield the correct relic abundance (amount of dark matter)



# Relic density Model 2: Doublet Triplet

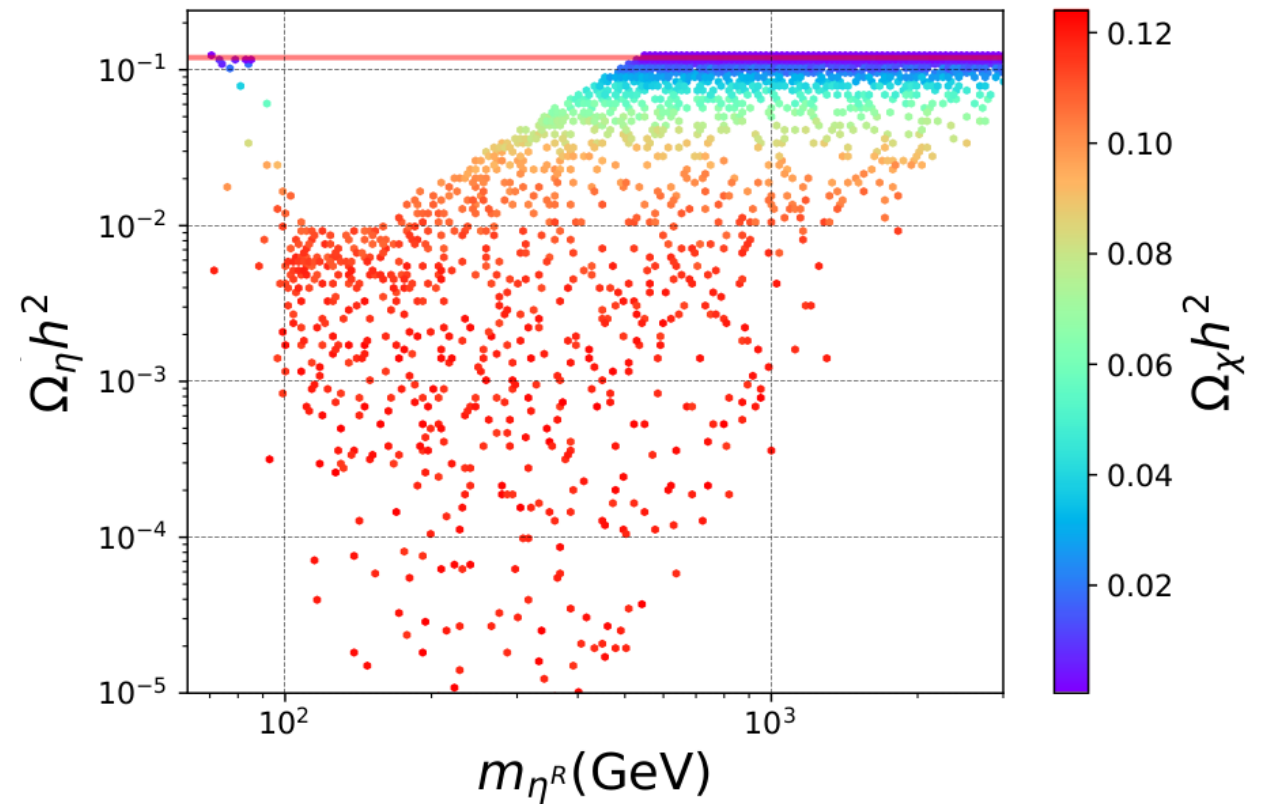
All points yield the correct relic abundance (amount of dark matter)



Due to large gauge interactions from the fermion sector and scalar sector it is not possible to recover this region

Relic density  
Model 3:  
Singlet Triplet

All points yield the correct relic abundance (amount of dark matter)





# Relic density



DM1 scalar



DM2 fermion

Relic density

Can they communicate?



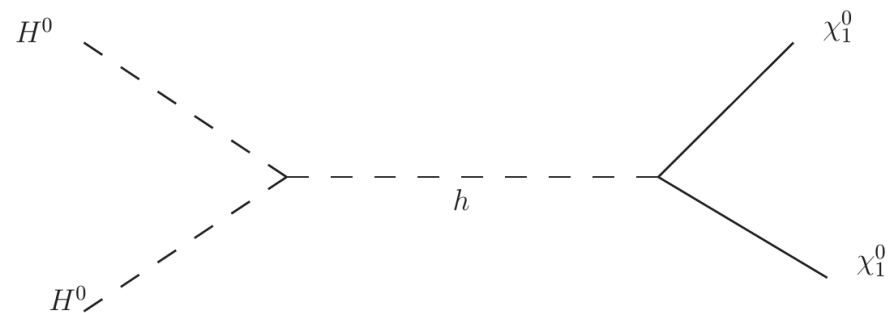
DM1 scalar



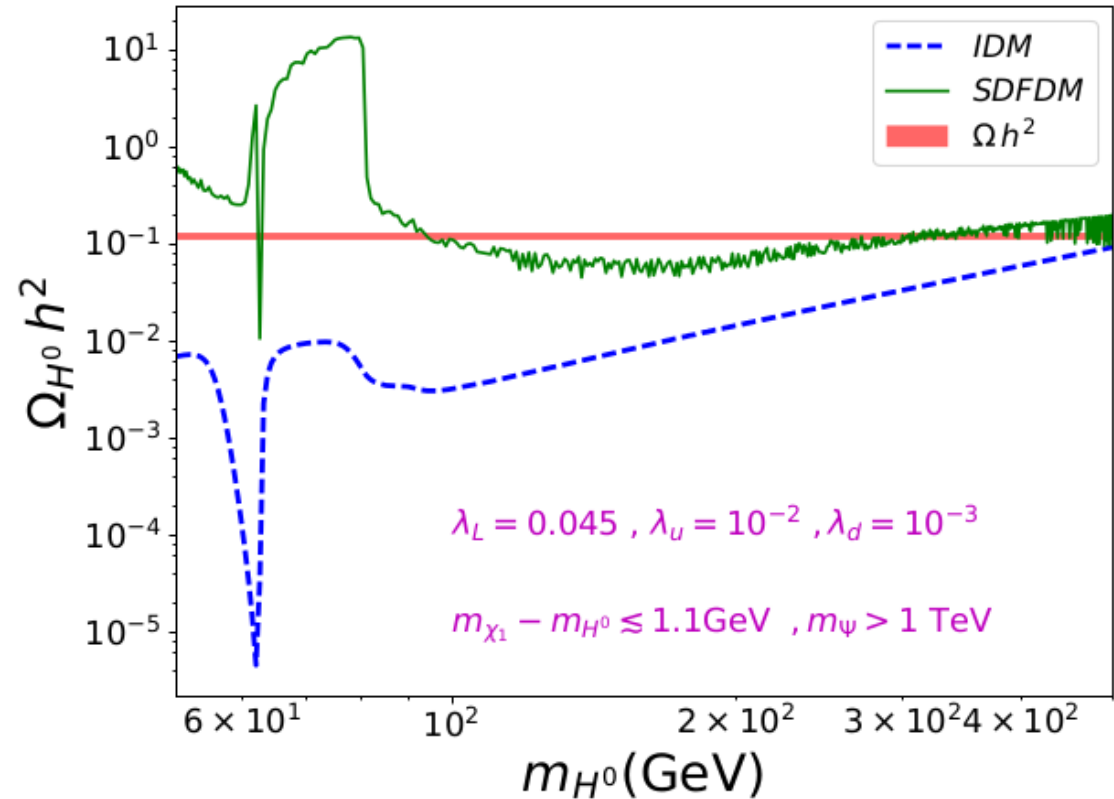
DM2 fermion

# Relic density Dark Matter conversion

Can they communicate?



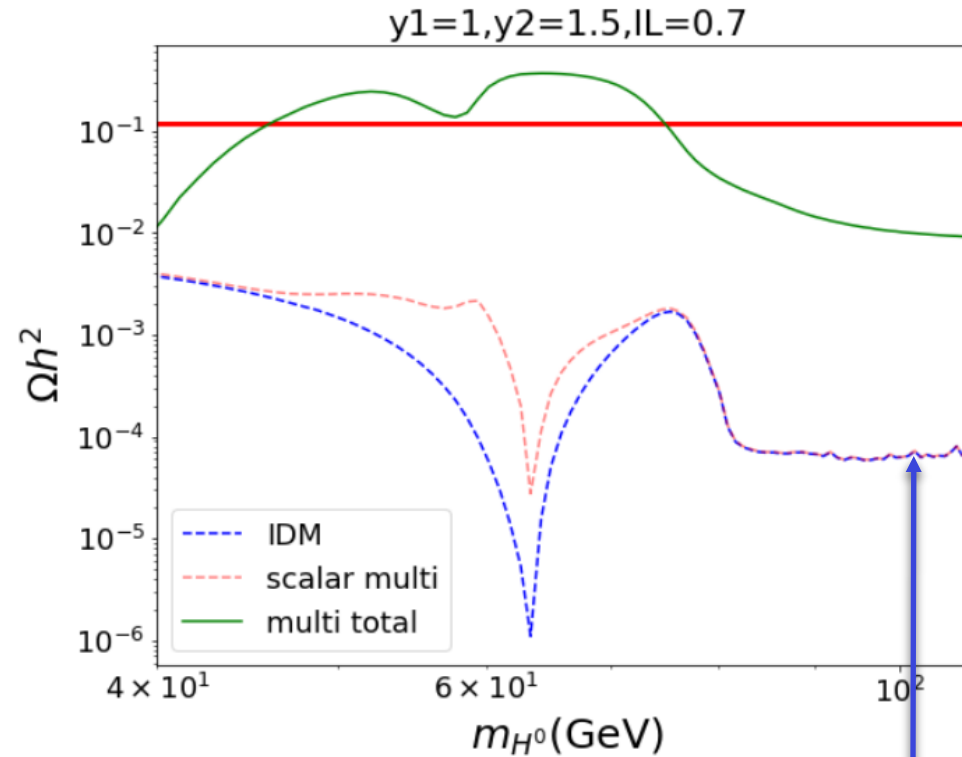
# Dark matter conversion singlet-doublet



Similar to the Singlet-  
Triplet



# Dark matter conversion doublet-triplet



Gauge interactions  
dominate over the  
Higgs interactions

# Direct Detection

- ◇ In multi-component dark sectors, direct detection is suppressed because it depends on the local relic abundance of the dark matter candidate and limits must be reinterpreted due to an expected different nucleon recoil rate.

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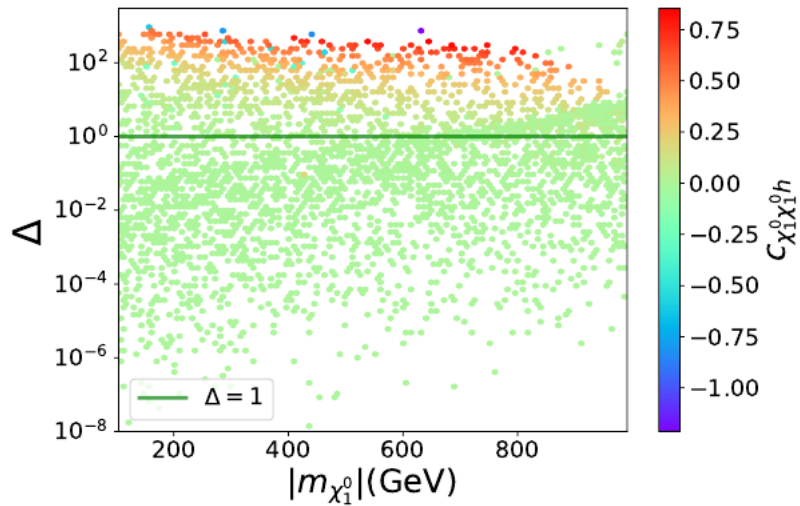
$$R < R_{\text{exp}}$$

$$\Delta = \frac{\sigma_{H^0}^{SI}}{\sigma_{X_e}^{SI}(M_{H^0})} \left( \frac{\Omega_{H^0}}{\Omega} \right) + \left( \frac{\sigma_{\chi_1^0}^{SI}}{\sigma_{X_e}^{SI}(m_{\chi_1^0})} + \frac{\sigma_{\chi_1^0}^{SD}}{\sigma_{X_e}^{SD}(m_{\chi_1^0})} \right) \left( \frac{\Omega_{\chi_1^0}}{\Omega} \right) < 1$$

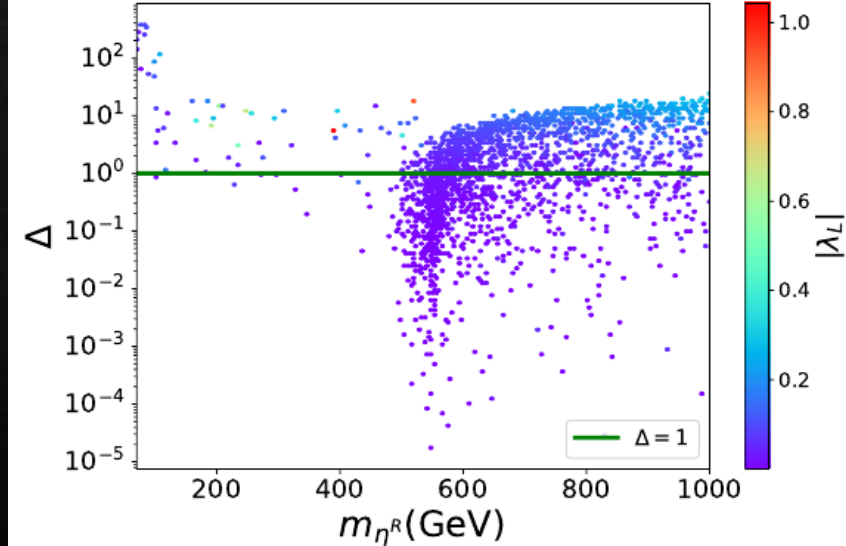


# Direct Detection limits from Xenon1T

## Doublet-Triplet

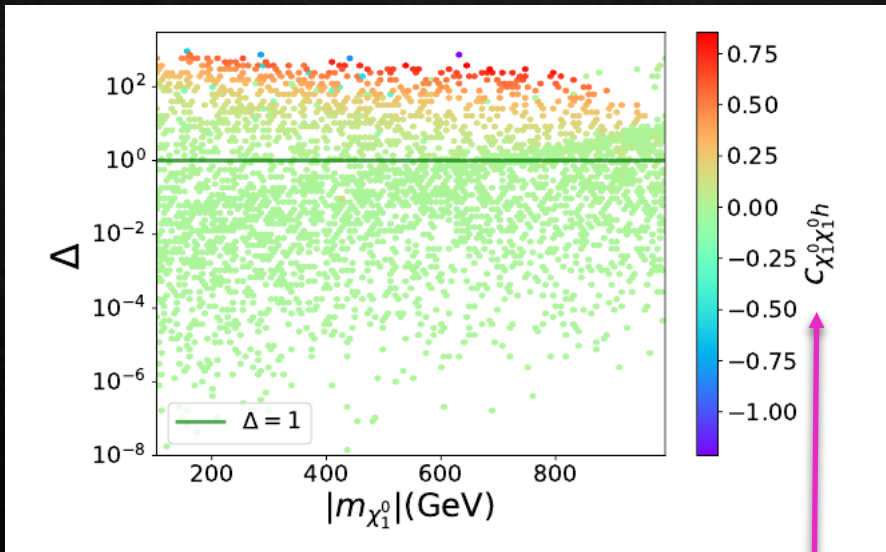


## Singlet-Triplet



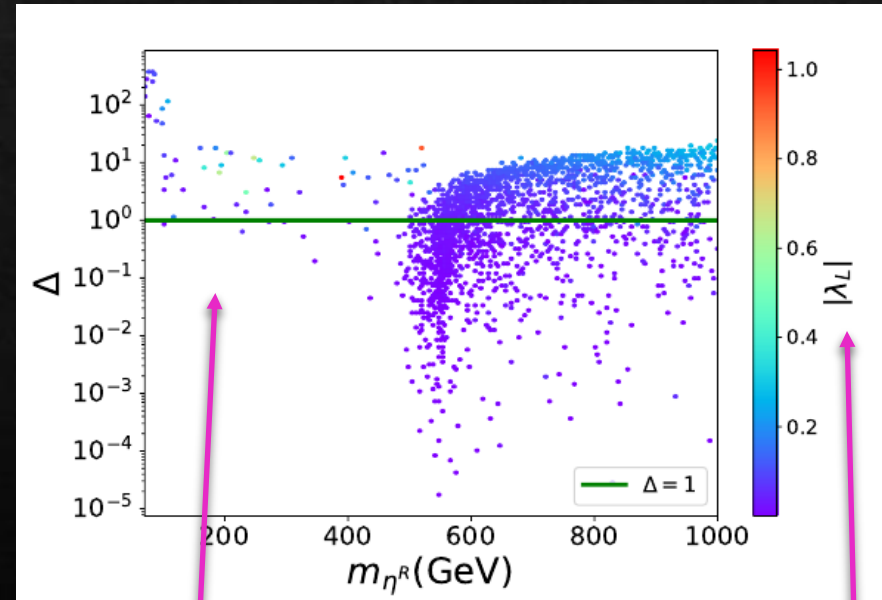
# Direct Detection limits from Xenon1T

## Doublet-Triplet



Higgs coupling to fermion  
DM  $< 0.08$

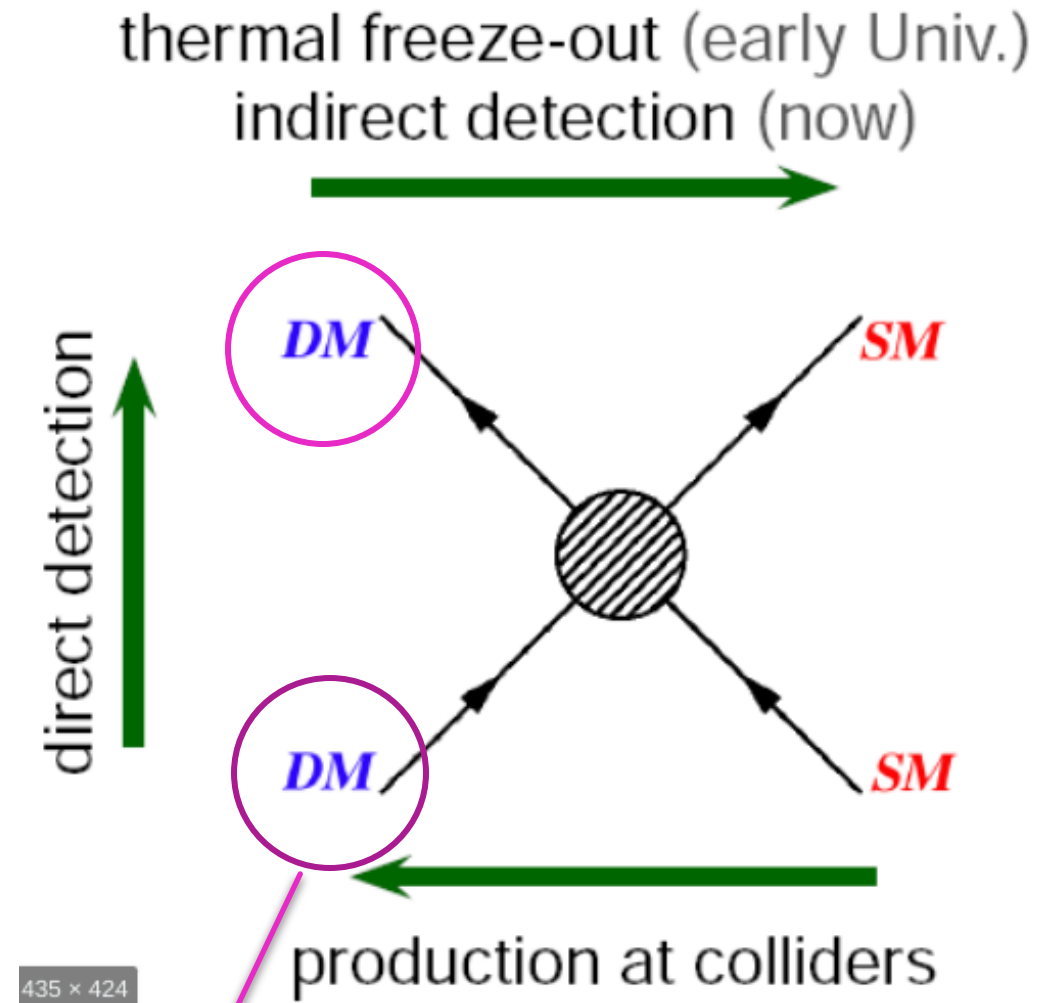
## Singlet-Triplet



Restrictions on the scalar  
mass

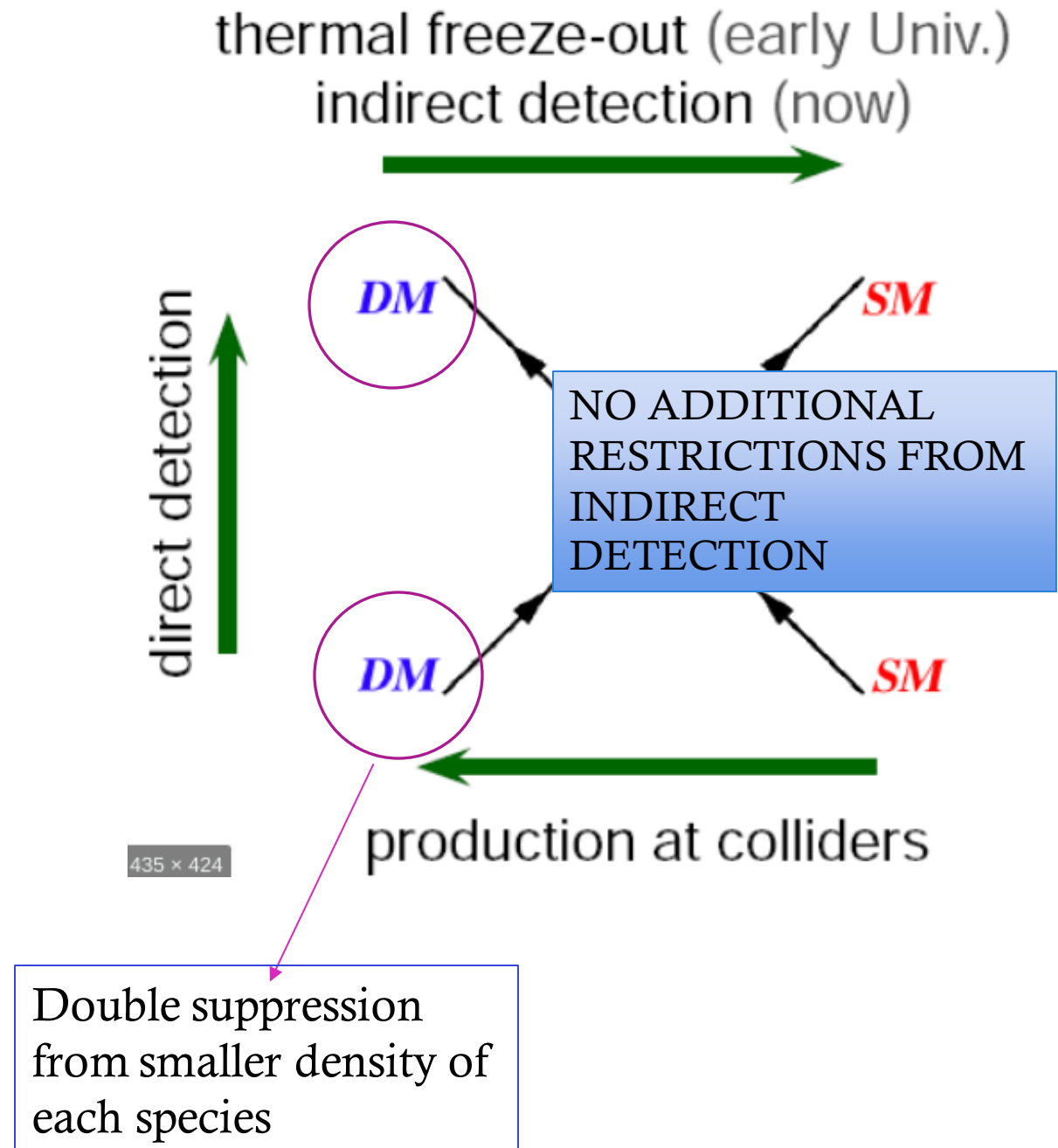
Higgs coupling to scalar  
DM  $< 0.2$

# Indirect detection



Double suppression  
from smaller density of  
each species

# Indirect detection





# Conclusions

- ◆ Multicomponent dark sectors are interesting scenarios of Dark Matter.
- ◆ For the Inert Doublet Model it is possible to recover the intermediate regime when another stable fermionic Dark Matter candidate is included.
- ◆ It is possible for the two DM candidates to communicate through the Higgs portal, thus altering each other's relic density
- ◆ Direct detection places constraints in the Higgs coupling to both scalars and fermions for two of the models while the thermally averaged cross sections of the models are out of bounds of current experiments.



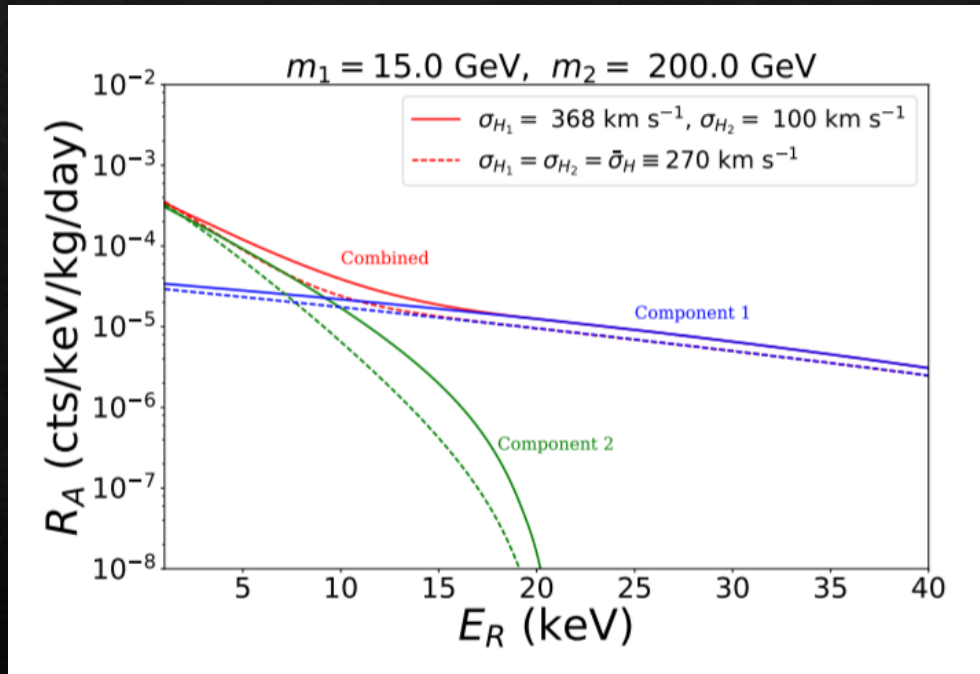
Thank you



# Bibliography

- [1] On the direct detection of multi-component dark matter implications of the relic abundance, JCAP, Herrero-Garcia, Juan and Scaffidi, Andre and White, Martin and Williams, Anthony G.
- [2] Dark Matter Search Results from a One Tonne $\times$ Year Exposure of XENON1T, XENON collaboration.
- [3] Exploring new models in all detail with SARAH, F. Staub
- [4] SPheno 3.1: extensions including flavour, CP-phases and models beyond the MSSM, W. Porod, F. Staub.

# Advantages of Multicomponent dark sectors: Direct detection constraints



It is even possible to differentiate between the dark matter candidates



# DM conversion through gauge interactions

