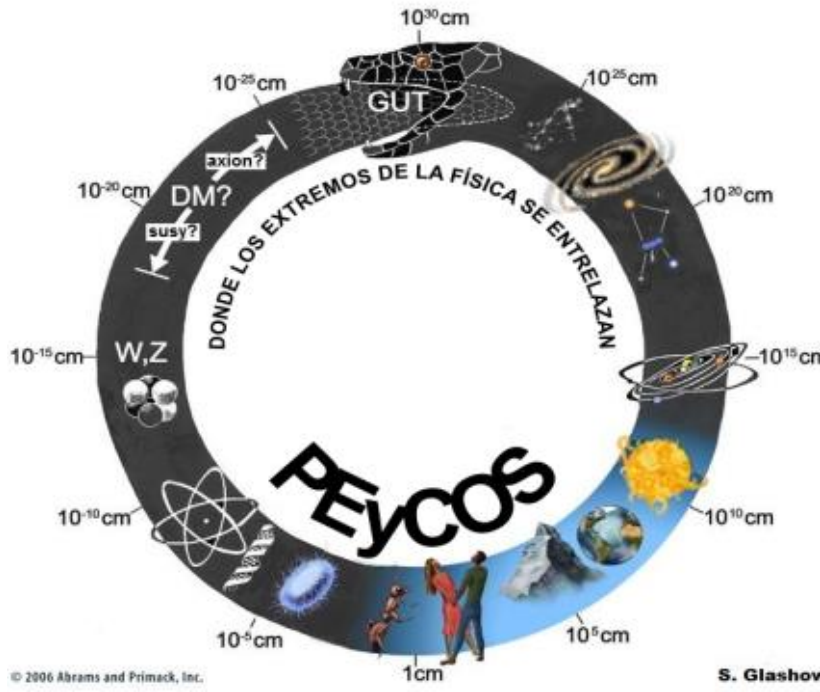


# THE NEUTRINO NATURE THROUGH CROSS SECTIONS IN THE LEFT-RIGHT SYMMETRIC MODEL




4<sup>th</sup> Colombian Meeting on High Energy Physics

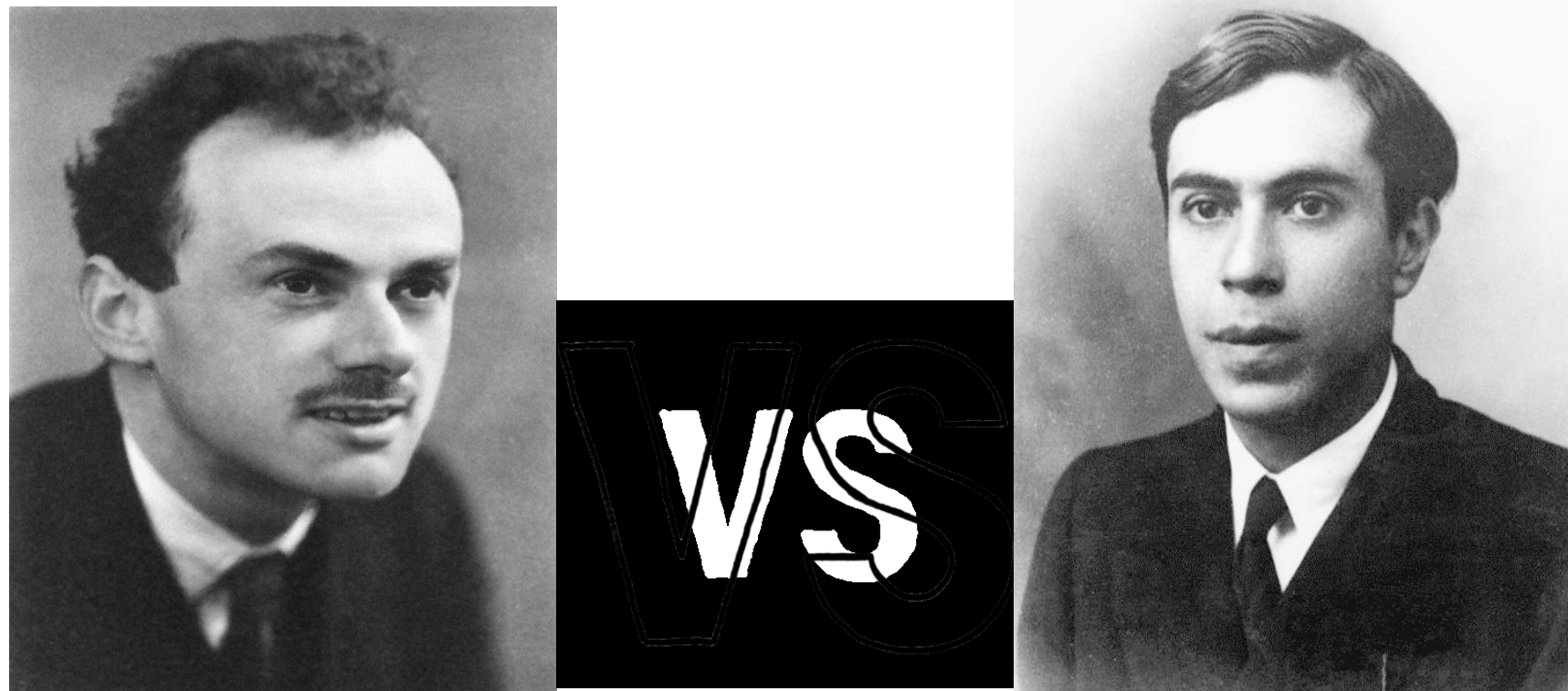
Frank Bula Martínez



# DIRAC OR MAJORANA NATURE

DIRAC	MAJORANA
$\nu \neq \bar{\nu}$	$\nu = \bar{\nu}$

- Every particle has its own antiparticle
- Neutrinos are uncharged, they could be their own antiparticles
- If neutrinos are Majorana particles, the leptonic number is not conserved
- Massive neutrinos  Physics beyond the SM



# THE LEFT-RIGHT SYMMETRIC MODEL

$$SU(3)_C \otimes SU(2)_L \otimes SU(2)_R \otimes U(1)_Y$$

$$ST_L \quad | \quad ST_R$$

Mirror symmetry is introduced in the SM, replacing the right singlet states by doublet states.

The new symmetric group  $SU(2)_R$  includes 3 non-ST bosons:  $W_R^\pm$ , and the  $Z_R$

These new bosons interact similarly like the other weak bosons do.

The Lagrangian for leptons contains Left-Right charged current (CC), and neutral current (CN) interactions.

$$\mathcal{L}_l = \mathcal{L}_{L,R}^{CC} + \mathcal{L}_{L,R}^{CN} \quad (1)$$

# CC INTERACTION

For the first leptonic family:

$$\mathcal{L}_{L,R}^{CC} = -\frac{g}{2\sqrt{2}} [\bar{\nu}_e(1 - \gamma_5)\gamma_\mu W_L^\mu e + \bar{\nu}_e(1 + \gamma_5)\gamma_\mu W_R^\mu e + H.c] \quad (2)$$

$W_L^\mu$  is the standar weak boson field.

$W_R^\mu$  is the new weak boson field for right CC interactions.

In the same way as  $W_L$ ,  $W_R$  has positive or negative charge, but is heavier than the particles that have been discovered so far.

# CN INTERACTION

For the first leptonic family:

$$\begin{aligned} \mathcal{L}_{L,R}^{CN} = & -\frac{g}{\cos \theta_W} \left( \frac{1}{2} \bar{\nu}_{eL} \not{Z} \nu_{eL} + \frac{\sin^2 \theta' \cos \theta_W}{2 \cos \theta'} \bar{\nu}_{eL} \not{Z}' \nu_{eL} - \left( \frac{1}{2} - \sin^2 \theta_W \right) \bar{e}_L \not{Z} e_L \right. \\ & + \frac{\sin^2 \theta' \cos \theta_W}{2 \cos \theta'} \bar{e}_L \not{Z}' e_L + \frac{\cos \theta_W}{2 \cos \theta'} \bar{\nu}_{eR} \not{Z}' \nu_{eR} + \sin^2 \theta_W \bar{e}_R \not{Z}' e_R + \sin^2 \theta_W \bar{e}_R \not{Z} e_R \\ & \left. - g \sin \theta_W \bar{e} \not{A} e \right) \end{aligned} \quad (3)$$

–  $\sin \theta' = \tan \theta_W$  New mixing angle

$\theta_W$  is the Weinberg's angle

$Z'$  is the new weak boson field for right CN interactions

# CN INTERACTION

For the first leptonic family:

$$\begin{aligned} \mathcal{L}_{L,R}^{CN} = & -\frac{g}{\cos \theta_W} \left( \frac{1}{2} \bar{\nu}_{eL} \not{Z} \nu_{eL} + \frac{\sin^2 \theta' \cos \theta_W}{2 \cos \theta'} \bar{\nu}_{eL} \not{Z}' \nu_{eL} - \left( \frac{1}{2} - \sin^2 \theta_W \right) \bar{e}_L \not{Z} e_L \right. \\ & + \frac{\sin^2 \theta' \cos \theta_W}{2 \cos \theta'} \bar{e}_L \not{Z}' e_L + \frac{\cos \theta_W}{2 \cos \theta'} \bar{\nu}_{eR} \not{Z}' \nu_{eR} + \sin^2 \theta_W \bar{e}_R \not{Z}' e_R + \sin^2 \theta_W \bar{e}_R \not{Z} e_R \Big) \\ & - g \sin \theta_W \bar{e} \not{A} e \end{aligned} \quad (3)$$

–  $\sin \theta' = \tan \theta_W$  New mixing angle

$\theta_W$  is the Weinberg's angle

$Z'$  is the new weak boson field for right CN interactions

**Charge of the electron has not changed!**

From the Lagrangian we can extract the coupling constants of some leptons with the  $Z$  and  $Z'$  bosons

$Z$

$$g_L^\nu = \frac{1}{2}$$

$$g_R^\nu = 0$$

$$g_L^e = -\frac{1}{2} + \sin^2 \theta_W$$

$$g_R^e = \sin^2 \theta_W$$

SM coupling constants have not changed

$Z'$

$$b_L^\nu = \frac{\sin^2 \theta' \cos \theta_W}{2 \cos \theta'}$$

$$b_R^\nu = \frac{\cos \theta_W}{2 \cos \theta'}$$

$$b_L^e = \frac{\sin^2 \theta' \cos \theta_W}{2 \cos \theta'}$$

$$b_R^e = -\frac{1}{2} + \sin^2 \theta_W$$

The new boson  $Z'$  is uncharged as  $Z$

Is heavier than the particles that have been discovered so far.

# Is not common to study the $\nu\nu \rightarrow \nu\nu$ interactions

- Bilenky, M., Bilenky, S. M., & Santamaria, A. Invisible width of the Z-boson and “secret”  $\nu$ - $\nu$  interactions. *Physics Letters B*, 301(2-3), 287–291. doi:10.1016/0370-2693(93)90703-k (2013)
- Ioka, K., & Murase, K. *IceCube PeV-EeV neutrinos and secret interactions of neutrinos*. *Progress of Theoretical and Experimental Physics*, 2014(6), 61E01–0. doi:10.1093/ptep/ptu090 (2014).
- Kolb, E. W., & Turner, M. S. *Supernova 1987A and the secret interactions of neutrinos*. *Physical Review D*, 36(10), 2895–2900. doi:10.1103/physrevd.36.2895 (1987).
- E. Flowers, P. Sutherland. *Neutrino-neutrino scattering and supernovae*. *Astrophysical Journal*, vol. 208, Aug. 15, 1976, pt. 2, p. L19-L21. (1976)....

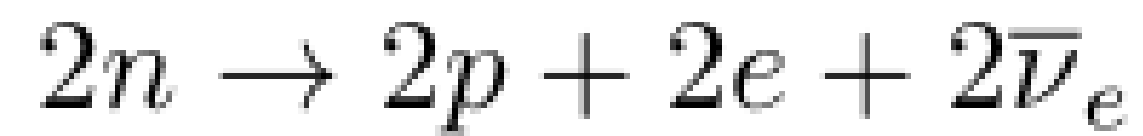


# Is not common to study the $\nu\nu \rightarrow \nu\nu$ interactions

- Bilenky, M., Bilenky, S. M., & Santamaria, A. Invisible width of the Z-boson and “secret”  $\nu$ - $\nu$  interactions. *Physics Letters B*, 301(2-3), 287–291. doi:10.1016/0370-2693(93)90703-k (2013)
- Ioka, K., & Murase, K. *IceCube PeV-EeV neutrinos and secret interactions of neutrinos*. *Progress of Theoretical and Experimental Physics*, 2014(6), 61E01–0. doi:10.1093/ptep/ptu090 (2014).
- Kolb, E. W., & Turner, M. S. *Supernova 1987A and the secret interactions of neutrinos*. *Physical Review D*, 36(10), 2895–2900. doi:10.1103/physrevd.36.2895 (1987).
- E. Flowers, P. Sutherland. *Neutrino-neutrino scattering and supernovae*. *Astrophysical Journal*, vol. 208, Aug. 15, 1976, pt. 2, p. L19-L21. (1976)....

Is extremely weak

$2\nu\beta\beta$



DIRAC

$0\nu\beta\beta$



MAJORANA

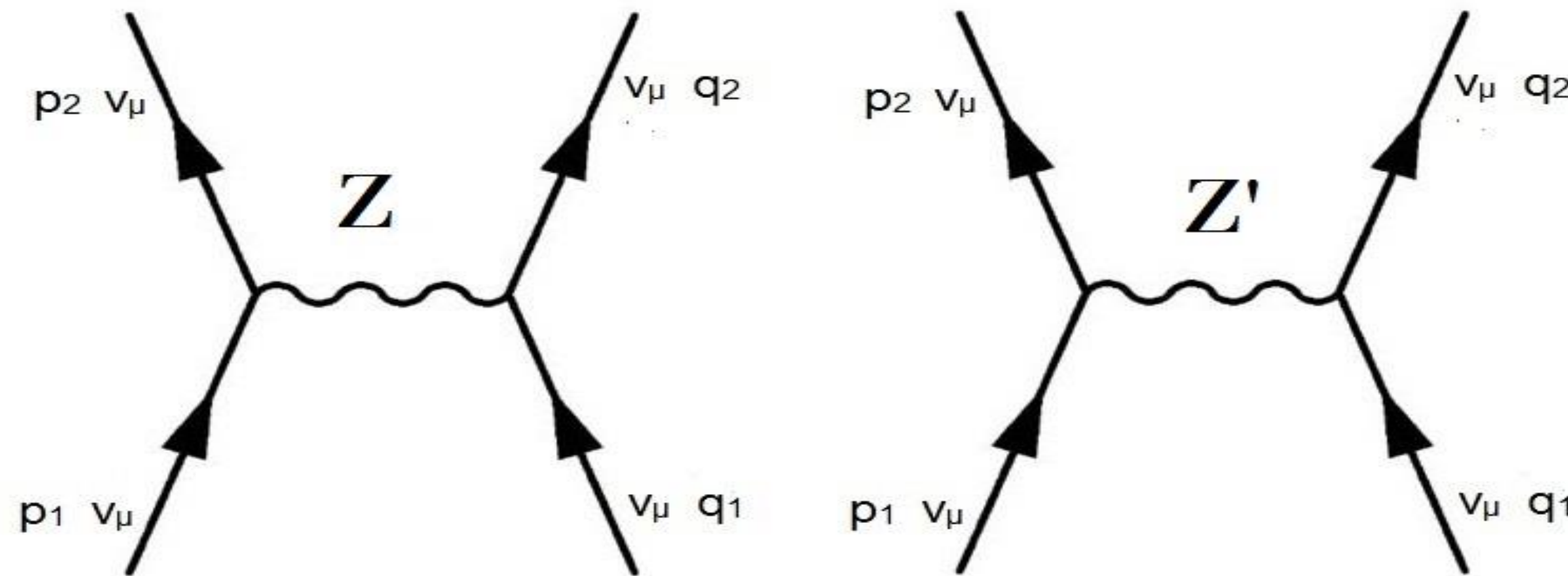


[Universo cuántico, los misterios del universo, “antipartículas, esas cosas raras”  
<https://universocuantico.wordpress.com/2009/04/18/antiparticulas-esas-cosas-raras-i/>]

One reason to study the neutrino nature is the asymmetry between matter and antimatter in the universe



Proposal for study the neutrino cross section



$$\mathcal{M} = \mathcal{M}_Z + \mathcal{M}_{Z'} \quad (4)$$

The  $\nu_\mu + \nu_\mu \rightarrow \nu_\mu + \nu_\mu$  cross section is calculated for both natures in the LRSM

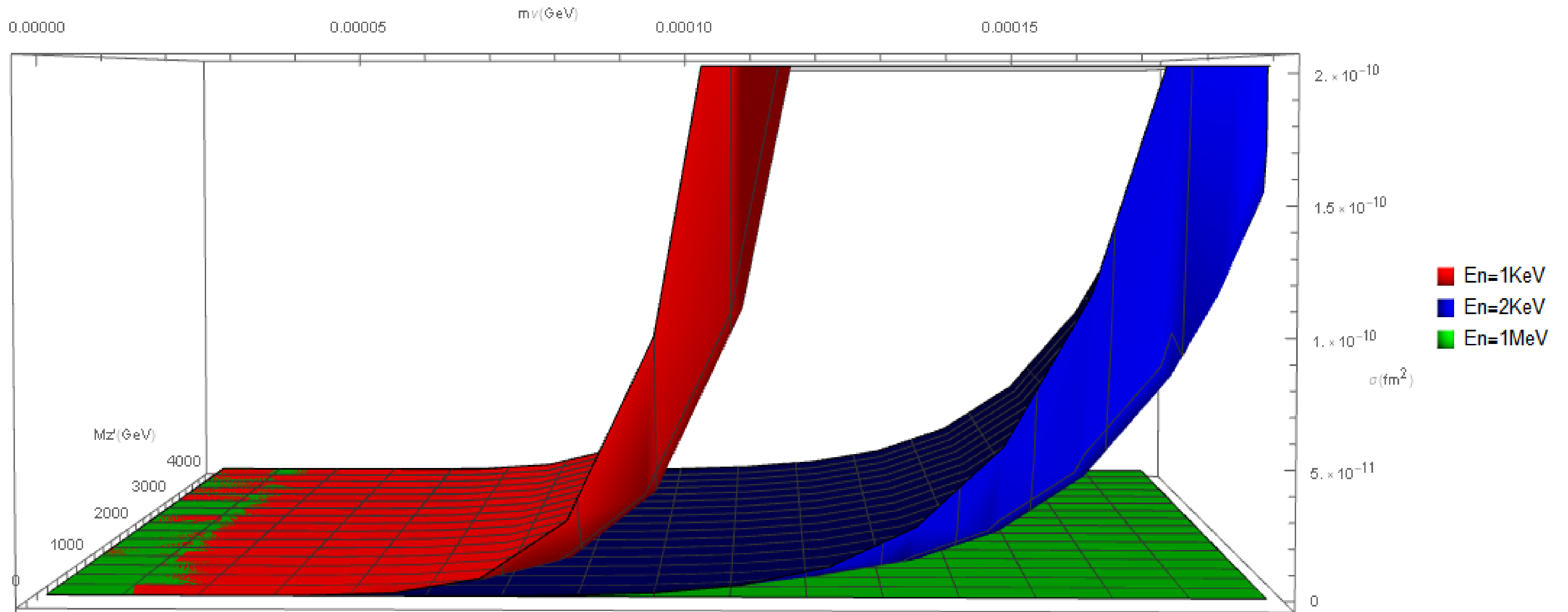
**DIRAC** 
$$\sigma_D(E_\nu, m_\nu, M_{Z'}) = \frac{(m_\nu^2 - 4E_\nu^2)^4(\beta + 3\gamma) - 32\rho m_\nu^2 E_\nu^2(16E_\nu^4 - 20m_\nu^2 E_\nu^2 + m_\nu^4)}{768\pi E_\nu^6} \quad (5)$$

**MAJORANA** 
$$\sigma_M(E_\nu, m_\nu, M_{Z'}) = G_F^2 (g_L^{\nu 2} + b_L^{\nu 2})^2 \frac{1280E_\nu^8 - 768m_\nu^2 E_\nu^6 + 736m_\nu^4 E_\nu^4 - 48m_\nu^6 E_\nu^2 + 5m_\nu^8}{512\pi E_\nu^6} \quad (6)$$

**This model includes massive neutrinos**

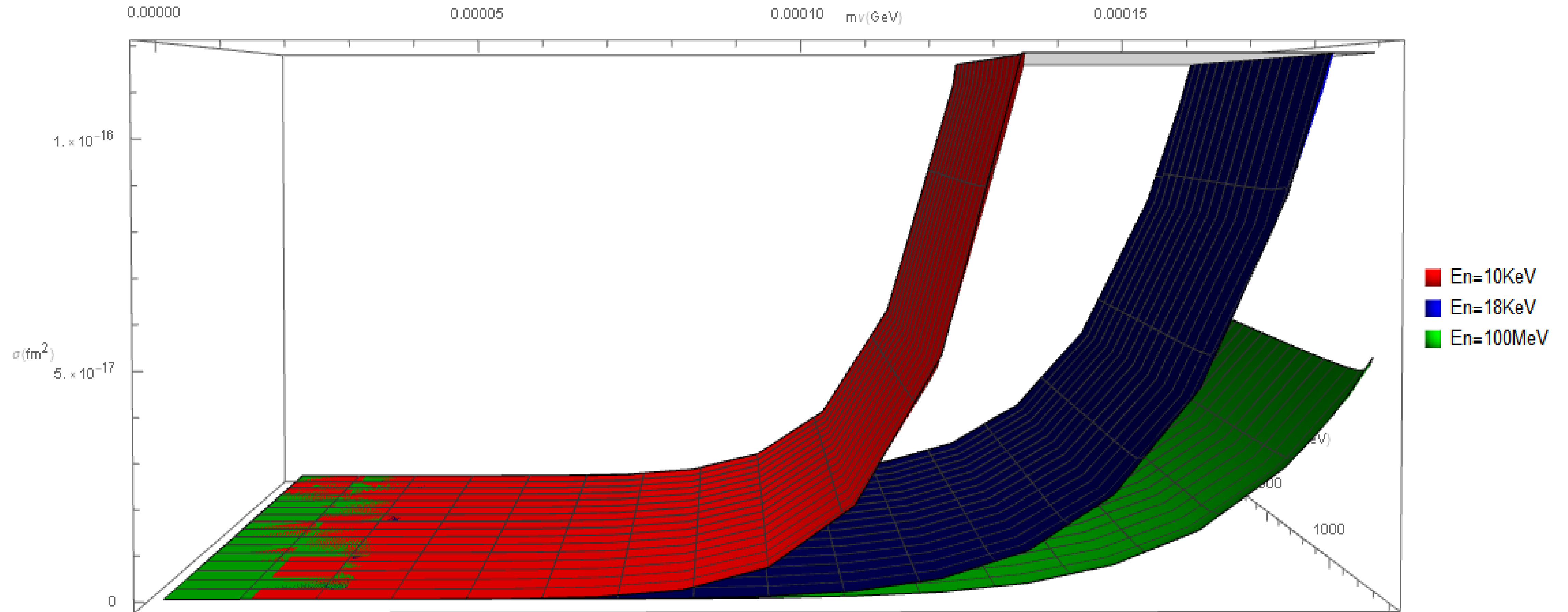
$\rho, \gamma, \beta$  are different combinations of the coupling constants

$$\sigma_D - \sigma_M$$



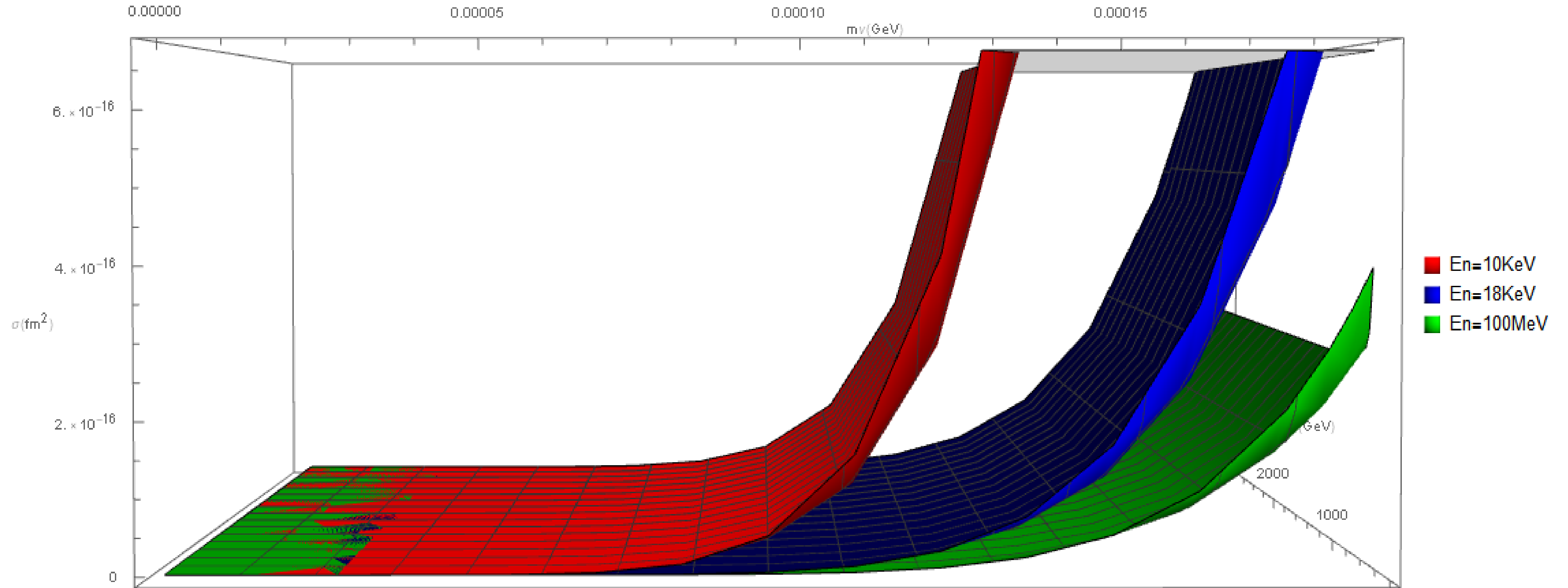
Cross Section Difference is bigger for low energy neutrinos

# Majorana Neutrino Mass



Cross section depends aswell of the neutrino mass

# Dirac Neutrino Mass



# Conclusion

- Neutrinos nature can be determined by their cross section
- Low energy neutrinos (KeV) are useful for this task, because at high energies region, Majorana and Dirac Natures, have similar cross section
- New symmetries can provide more information about neutrino masses



[www.uniatlantico.edu.co](http://www.uniatlantico.edu.co) | **#YocreoenlaUA** | VIGILADA MINEDUCACIÓN