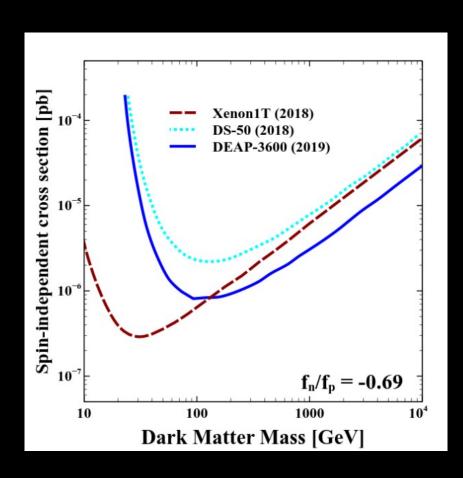
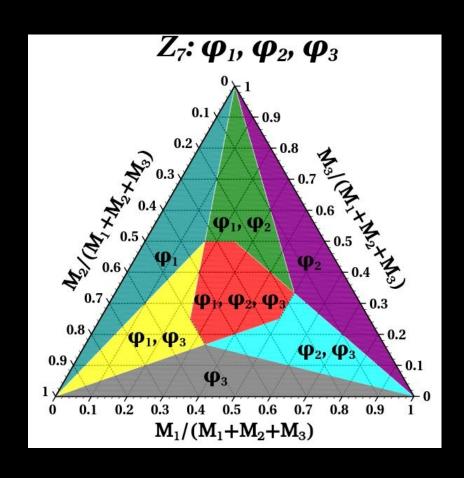
# Dark Matter: challenges and open problems





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#### We don't know much about the dark matter particle

#### What we know:

It's stable, neutral and "cold"

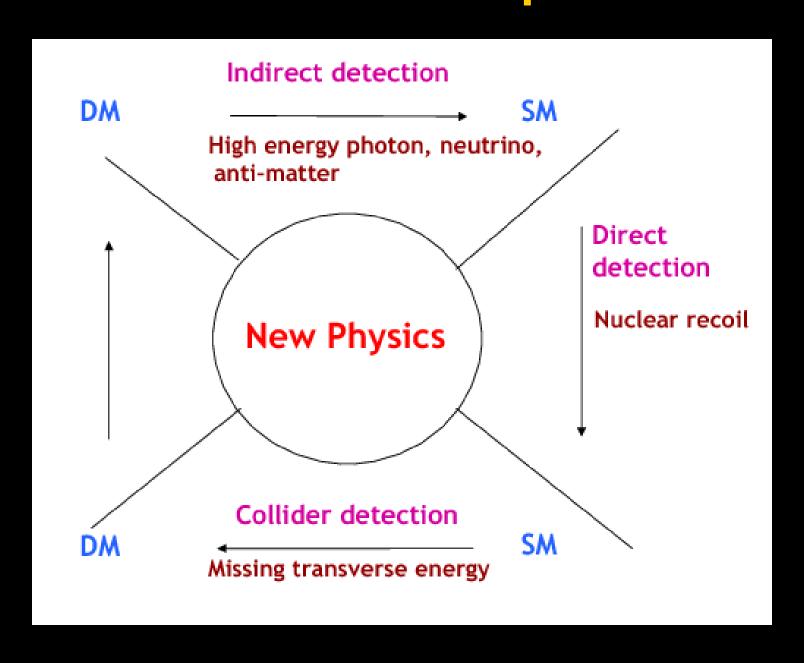
It does not interact much

It's a relic from the early Universe

#### What we don't know:

Its mass, its spin. Is it real or complex? Why is it stable? Its quantum numbers. How can it be detected? Are there several of them? Is it connected to other problems? How was it produced? How is it distributed? ...

## During the next few years we may be able to answer some of these questions

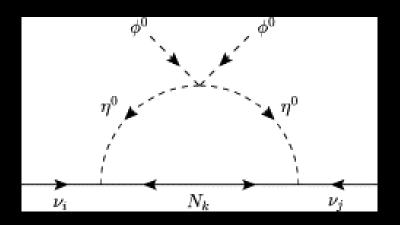


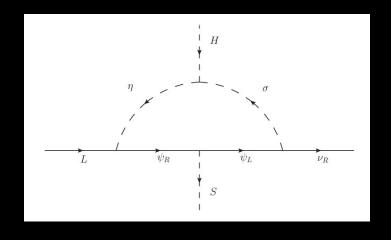
#### Is dark matter related to neutrinos?

They both require new physics

v-masses at 1-loop from DM

Also for Dirac vmasses At the TeV scale?





#### Dark matter may be related to neutrinos via anomaly cancellation

Similar to quarks and leptons in the SM

A new model based on a B-L symmetry

The DM consists of two Dirac particles

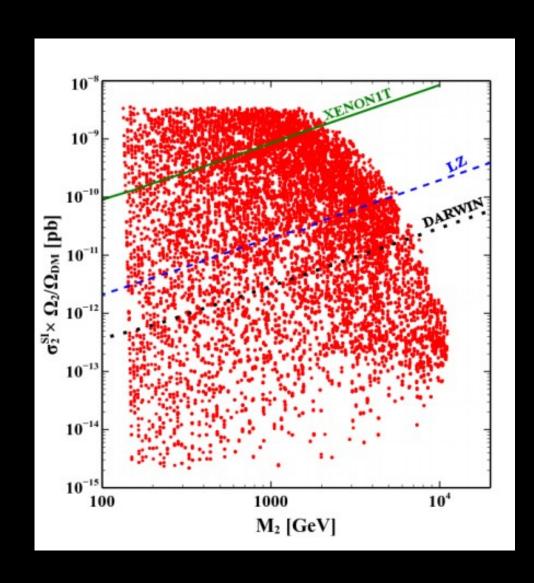
Particles	$U(1)_{B-L}$	$(SU(3)_c, SU(2)_L, U(1)_Y)$
$Q_{Li}$	1/3	( <b>3</b> , <b>2</b> , 1/6)
$u_{Ri}$	1/3	(3, 1, 2/3)
$d_{Ri}$	1/3	$(\overline{3}, 1, -1/3)$
$L_i$	-1	(1, 2, -1/2)
$e_{Ri}$	-1	(1, 1, -1)
$N_{R1}$	-1	( <b>1</b> , <b>1</b> , 0)
$N_{R2}$	-1	( <b>1</b> , <b>1</b> , 0)
$\xi_L$	10/7	( <b>1</b> , <b>1</b> , 0)
$\eta_R$	-4/7	( <b>1</b> , <b>1</b> , 0)
$\zeta_R$	-2/7	( <b>1</b> , <b>1</b> , 0)
$\chi_L$	-9/7	( <b>1</b> , <b>1</b> , 0)
H	0	(1, 2, 1/2)
$\phi_1$	1	( <b>1</b> , <b>1</b> , 0)
$\phi_2$	2	( <b>1</b> , <b>1</b> , 0)

### This model can be tested at the LHC and in direct detection experiments

Z' (B-L) searches at the LHC

**Current and future DD experiments** 

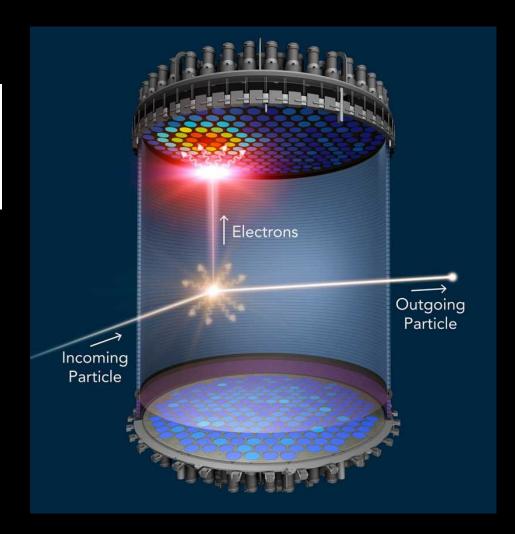
It also predicts a massless v



## Does the dark matter particle couple equally to protons and neutrons?

$$\sigma_0^{SI} = \sigma_p \cdot \frac{\mu_A^2}{\mu_p^2} \cdot [Z \cdot f^p + (A - Z) \cdot f^n]^2$$

 $f_p = f_n$  is the usual assumption



But it does not have to be so

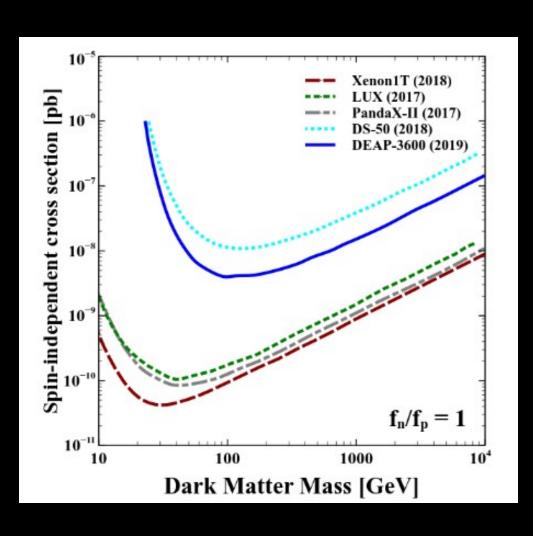
### Several simple dark matter models give rise to isospin-violating DM

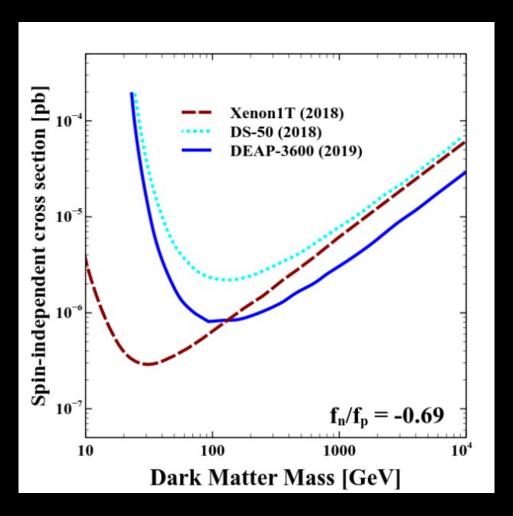
Dark-Photon Mediation:  $f_n/f_p=0$ 

Z Mediation:  $f_n/f_p = -12.5$ 

Light Squark-Mediation:  $f_n/f_p = 0.7 (u)$  or 1.5 (d)

### The interpretation of the experimental limits depend on this assumption





Future signals may allow to test if the coupling is really the same

#### Do dark matter interactions violate P or CP?

3 possibilities for spin-0 DM

Scalar, pseudoscalar, or CPV

DM decaying into gamma-ray lines

$$\mathcal{L} = \frac{a}{M_{pl}} \varphi B_{\mu\nu} B^{\mu\nu} + \frac{b}{M_{pl}} \varphi B_{\mu\nu} \widetilde{B}^{\mu\nu} + \frac{c}{M_{pl}} \varphi W^{I}_{\mu\nu} W^{I\mu\nu} + \frac{d}{M_{pl}} \varphi W^{I}_{\mu\nu} \widetilde{W}^{I\mu\nu}$$

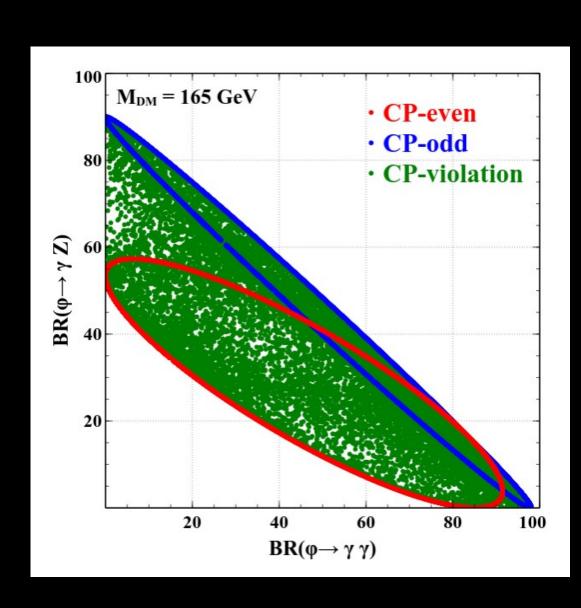
Via dim-5 operators

## The correlations between different decays depend on the CP properties of the DM

The CP nature may be determined

For annihilating DM is not so simple

Perhaps at the LHC, or DD?



#### How many dark matter particles are there?

Just one is the simplest option

But several are allowed

**Guaranteeing the stability is tricky** 

**Even for one DM** particle

We lack a theoretical framework

# Z<sub>N</sub> (N≥4) symmetries can simultaneously stabilize several (scalar) DM particles

They often appear as remnants of U(1)

The stability depends on the masses

