Scattering Amplitudes and Unitarity for Gravitationally Mediated Dark Matter en extra Dimensions

Kirkimaan Mohan

PPC 2022: XV International Conference on Interconnections between Particle Physics and Cosmology Washington University in St. Louis 8th June 2022

R. Sekhar Chivukula (UCSD), Dennis Foren (UCSD), Dipan Sengupta (Adelaide) and Elizabeth H. Simmons (UCSD) arXiv:1903.05650, arXiv:1910.06159, arXiv:2002.12458, arXiv:2104.08169, arXiv:2206:xxxxx

With

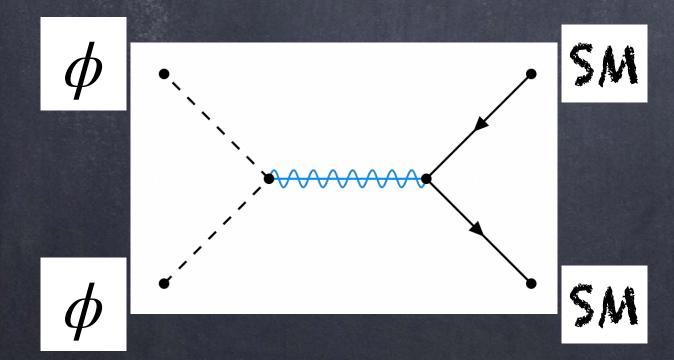




M. C. ELVAELCOM

Can Dark Matter only have gravitational interactions
Yes, but interactions are Planck scale suppressed.
Can Dark Matter only have gravitational interactions
and and still behave like a WIMP?
Yes!

Simple Example:





 $\mathcal{L}_{ ext{DM}} = rac{1}{2} \partial_\mu \phi \partial^\mu \phi - rac{1}{2} m^2 \phi^2 \; .$

 $h^{\mu
u}(T^{
m SM}_{\mu
u}+T^{
m DM}_{\mu
u})$

 $y = \pi r_c$



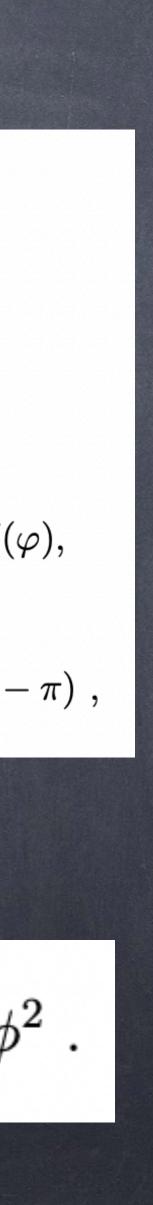
o tower of massive spin-2 states $h_{\mu\nu}$ o Radion (r) @ SM +DM (Localized on Tev) SM + DM couple to $\frac{1}{\Lambda}h^{\mu\nu}(T^{(SM)}_{\mu\nu} + T^{(DM)}_{\mu\nu})$ $A \simeq M_{pl} e^{-kr_c}$ (Usually a few TeV or more) $[G_{MN}] = \begin{pmatrix} g_{\mu\nu} \exp\left[-2\left(A(y) + \frac{e^{2A(y)}}{2\sqrt{6}}\kappa\,\hat{r}(x,y)\right)\right] \\ 0 & -\left(1 + \frac{e^{2A(y)}}{4}\right) \end{pmatrix}$

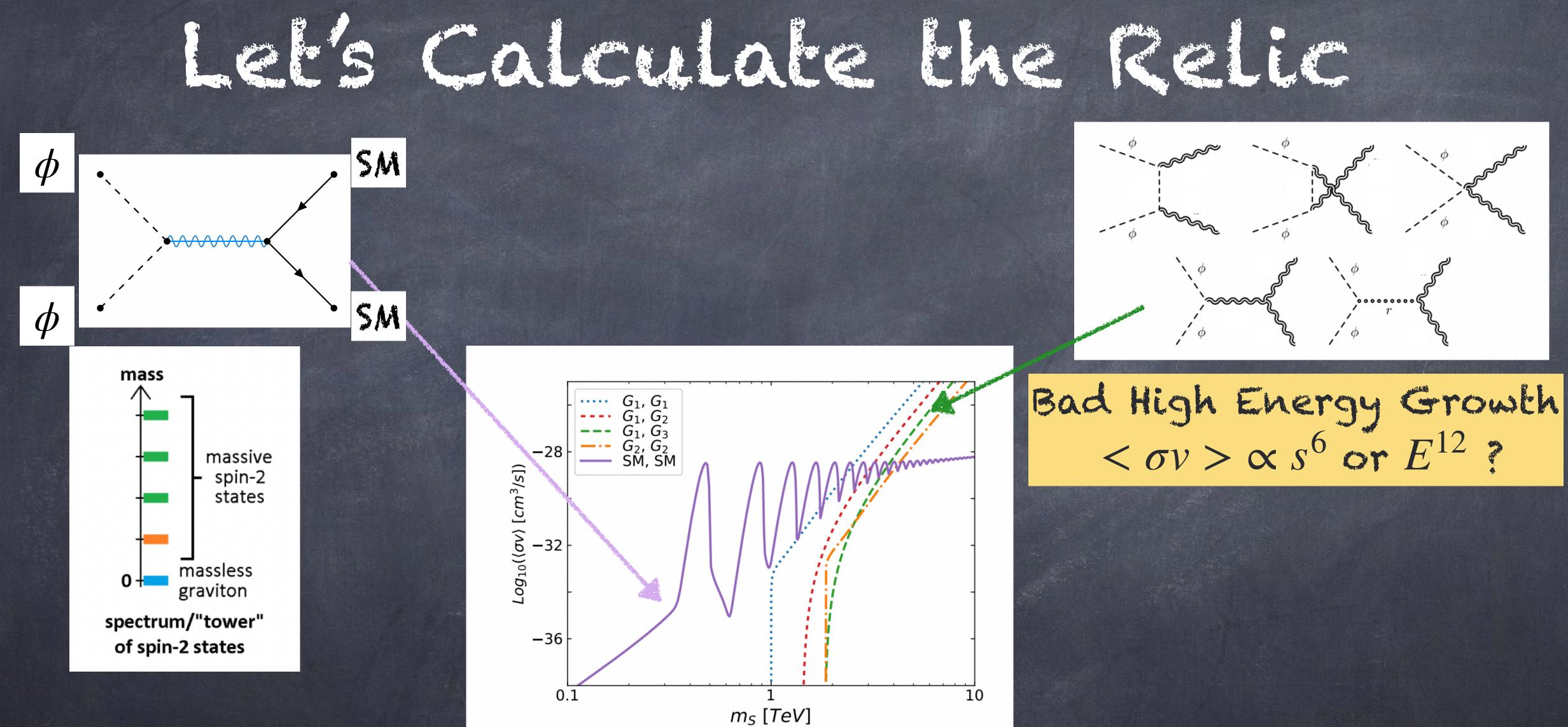
 $S = S_{\text{bulk}} + S_{\text{UV}} + S_{\text{IR}} ,$

$$egin{aligned} S_{ ext{bulk}} &= rac{1}{2}M_5^3 \int d^4x \int\limits_{-\pi}^{\pi} darphi \sqrt{G}(R-2\Lambda_B), \ S_{ ext{UV}} &= \int d^4x \int\limits_{-\pi}^{\pi} darphi \sqrt{-g_{ ext{UV}}}(-V_{ ext{UV}}+\mathcal{L}_{ ext{UV}})\delta \ S_{ ext{IR}} &= \int d^4x \int\limits_{-\pi}^{\pi} darphi \sqrt{-g_{ ext{IR}}}(-V_{ ext{IR}}+\mathcal{L}_{ ext{IR}})\delta(arphi) \end{aligned}$$

$$\left(\frac{0}{e^{2A(y)}}{\kappa \hat{r}(x,y)}\right)^2$$

 $\mathcal{L}_{\mathrm{DM}} = rac{1}{2} \partial_\mu \phi \partial^\mu \phi - rac{1}{2} m^2 \phi^2 \; .$



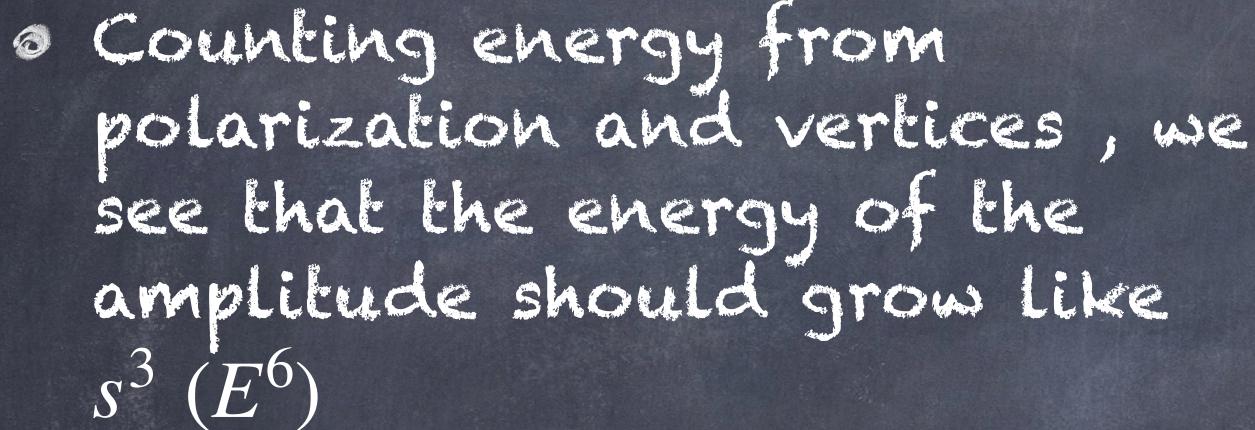


arXiv: 1511.03278, 1709.09688, 1907.04340, 2004.14403



4/14





 $\emptyset \Longrightarrow$ Unitarity is violated at a scale much smaller than A

Why the bad high energy growth? $\epsilon^0_{\mu\nu} o rac{k_\mu k_
u}{m^2}$ $\sim -\frac{\phi}{\Lambda} \xrightarrow{\sqrt{s}}$

 $\epsilon^0_{\mu\nu} \rightarrow \frac{k_\mu k_\nu}{\bar{k}}$





 $\emptyset \Longrightarrow Unitar.$ is violated at a scale mu . .: valler than A

Why the bad high energy growth?

 $\epsilon^0_{\mu\nu} o \frac{k_\mu k_\nu}{-}$





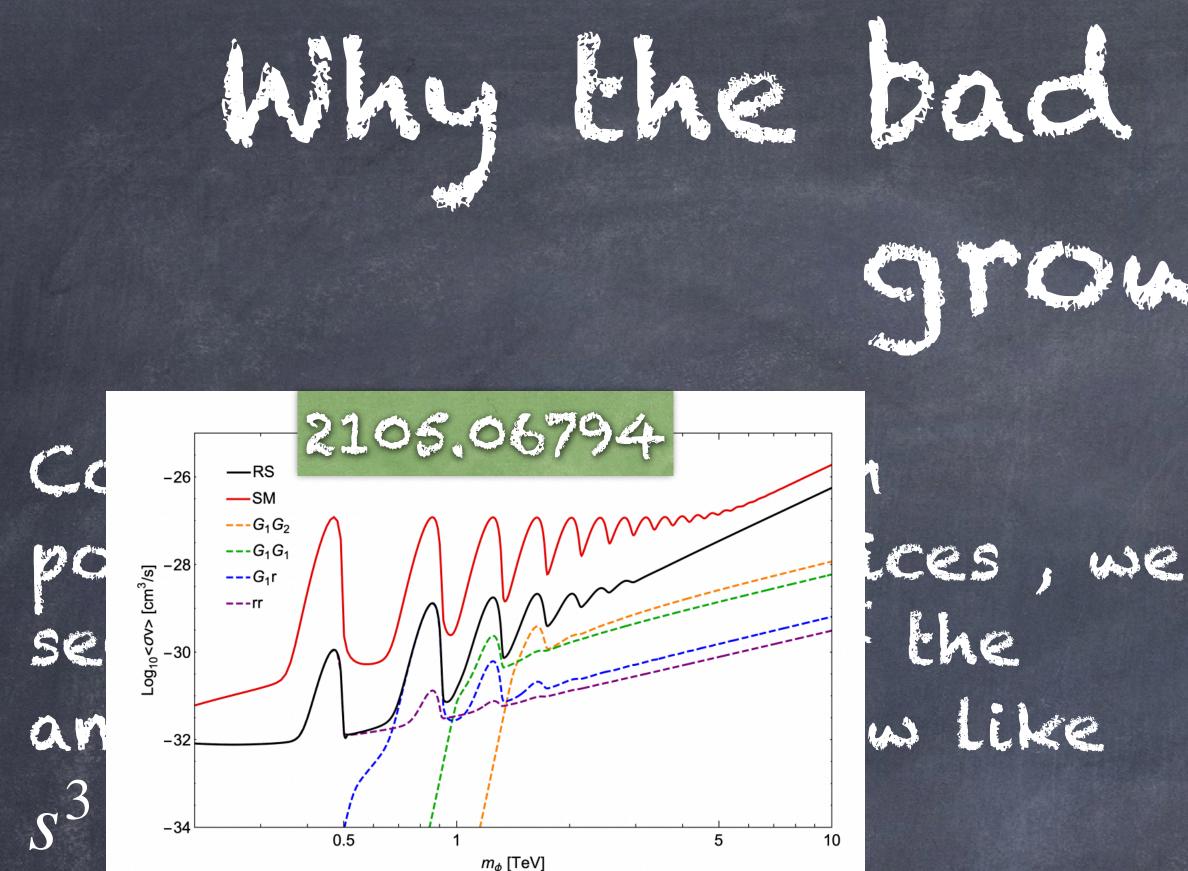
 $\emptyset \Longrightarrow Unitar.$ is violated at a scale my is saller than A

Why the bad high energy growth? $\begin{array}{ccc} & \phi & \frac{\sqrt{s}}{\Lambda} & & & \\ & & & \\ & & & & & \\ & & & & & \\ & & & & \\ & & & & & \\ & & & & & \\ & & & & & \\ & &$ $\epsilon^0_{\mu\nu} \rightarrow \frac{k_\mu k_\nu}{m^2}$

Diffeomorphism invariance protects amplitude from bad high energy growth! How?

 $h_{\mu\nu} \to h_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$





 $\emptyset \Longrightarrow Unitar$ scale mu ...

is violated at a aller than A

Why the bad high energy $\epsilon^0_{\mu\nu} o \frac{k_\mu k_\nu}{m^2}$ $\epsilon^0_{\mu\nu} \rightarrow \frac{k_\mu k_\nu}{k_\nu}$

Diffeomorphism invariance protects amplitude from bad high energy growth! How?

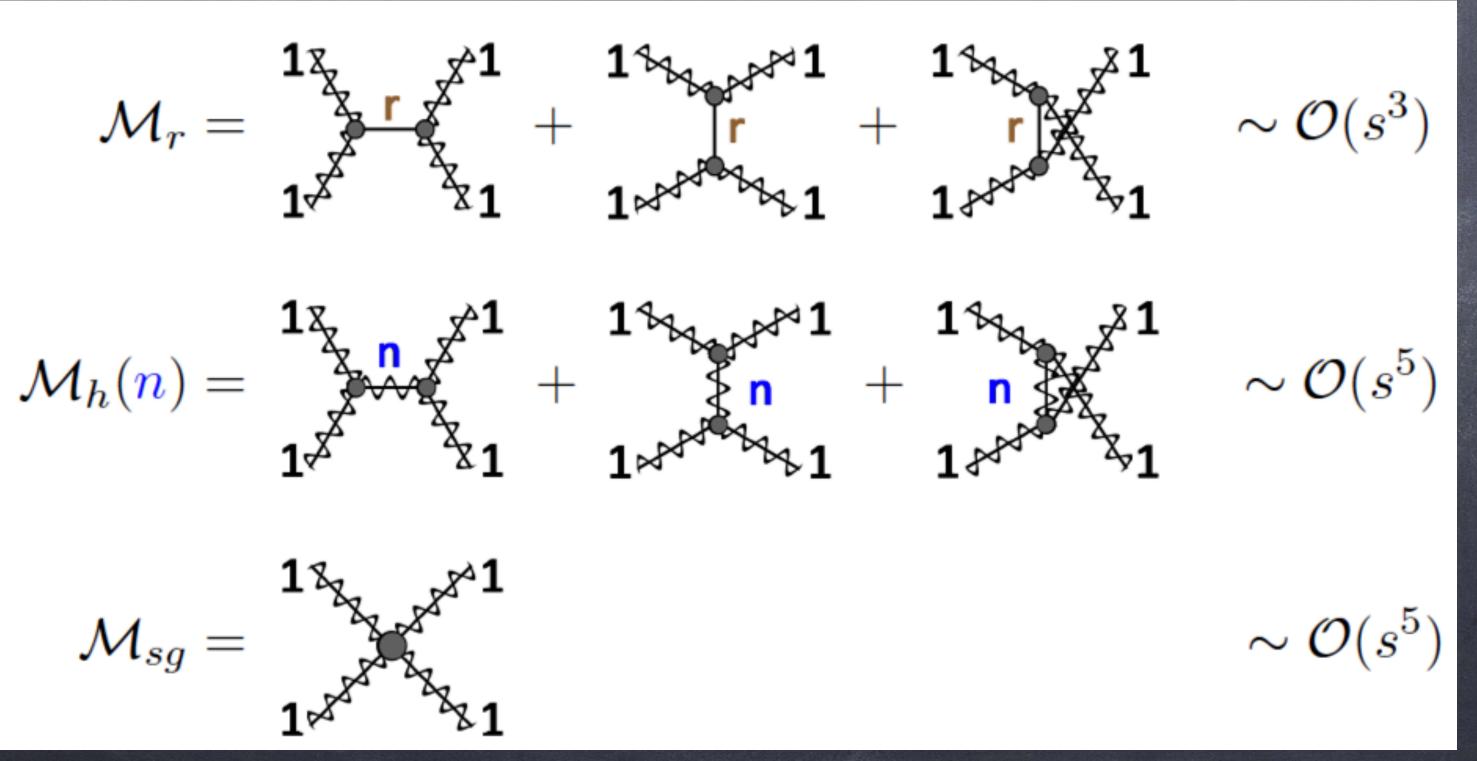
 $h_{\mu\nu} \to h_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$



Unitarity in Extra Dimensional Models

Unitarity in Extra Dimensions

Let's Look at $2 \rightarrow 2$ scattering of gravitons - worst high energy behavior



7/14

Freliminaties

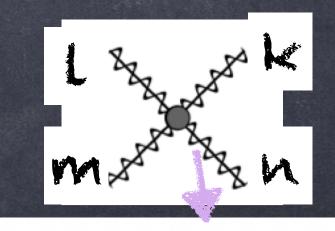
- @ Decompose the 5D graviton field into a lower of 4D fields
- a Determine the fifth dimensional wave function for each mode from the sturn Liouville problem
- @ Wave functions are orthogonal and complete
- @ Substitute wave functions into overlap integrals that set the couplings of the vertices

$$\mathbf{1}_{\mathbf{x}_{m}} \mathbf{1}_{\mathbf{x}_{m}} \mathbf{1}_{\mathbf{x}_{$$

$$\hat{h}_{\mu\nu}(x,y) = \frac{1}{\sqrt{\pi r_c}} \sum_{n=0}^{+\infty} \hat{h}_{\mu\nu}^{(n)}(x) \psi_n(\varphi) .$$

$$\partial_{\varphi} [e^{-4A} \, \partial_{\varphi} \psi_n] = -\mu_n^2 e^{-2A} \psi_n$$

$$\frac{1}{\pi} \int_{-\pi}^{+\pi} d\varphi \quad e^{-2A} \psi_m \,\psi_n = \delta_{m,n}$$
$$\delta(\varphi_2 - \varphi_1) = \frac{1}{\pi} e^{-2A} \sum_{j=0}^{+\infty} \psi_j(\varphi_1) \,\psi_j(\varphi_2)$$

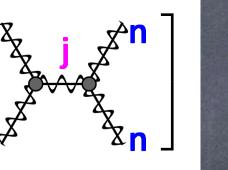


$$a_{klmn} \equiv rac{1}{\pi} \int_{-\pi}^{+\pi} d\varphi \ e^{-2A} \psi_k \psi_l \psi_m \psi_n$$

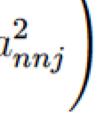


$$\mathcal{M} = \frac{n^{3}}{n^{4}} + \sum_{S,T,U} \left[\frac{n^{3}}{n^{4}} + \frac{n^{3}}{2n} + \frac{n^{3}}{n^{4}} + \frac{n^{3}}{2n} + \frac{n^{3}}{n^{4}} + \sum_{j>0} \frac{n^{3}}{n^{4}} + \frac{n^{3}}{n^{2}} + \frac{n^{3}}{n^{4}} +$$





$$\mathcal{M}(s, heta) = \sum_{\sigma \in \frac{1}{2}\mathbb{Z}} \overline{\mathcal{M}}^{(\sigma)}(heta) \cdot s^{\sigma}$$



 At each order of S, relations
 between overlap integrals ensure that the bad high energy growth vanishes

 \odot Example: $\sum_{i} a_{nnj}^2 = a_{nnnn}$

@ Possible to prove analytically using completeness and orthogonality

o Finally, as expected, we find that the amplitude grows as $O(s) \implies$ Unitarity violated at $\sim \Lambda \simeq M_{pl} \ e^{-kr_c}$

What about a stabilized Extra almansion?

@ In its original formulation RS consists of a massless radion => Casimir force that renders the extra-dimension unstable.

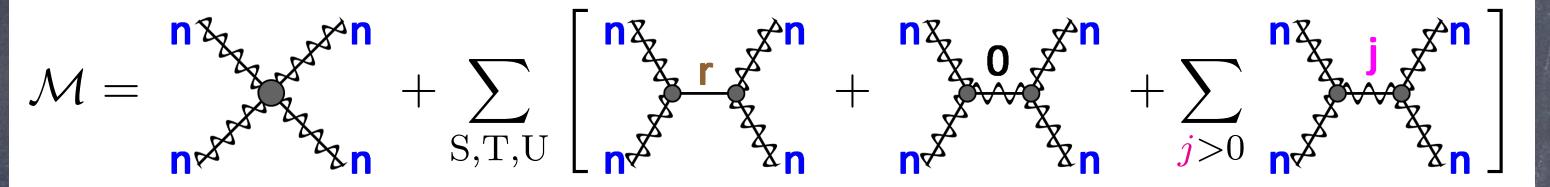
@ Alternately, a massless radion leads to a Brans-Dicke (Scalar-Tensor) theory of gravity and contradicts experiments when there is matter on the TeV Brane.

Goldberger-Wise mechanism can provide mass to the radion and stabilize the extra-dimension.

Goldberger, Wise 2000 De-Wolfe, Freedman, Karch, Gubser 2001 Peloso, Koffman 2001 Tanaka, Garriga 2001 Smolyakov, Volobuev 2002



Massive Radion and Scattering Amplitudes



- Naively adding a mass term by hand to the radio
 propagator messes up the sum rules and the amplitude grows as $O(s^2)$
- ø This too must cancel due to diffeomorphism invariance

11/14

$$\begin{split} \mathcal{L}_{5\mathrm{D}} &\equiv \mathcal{L}_{\mathrm{EH}} + \mathcal{L}_{\Phi\Phi} + \mathcal{L}_{\mathrm{pot}} , \\ \mathcal{L}_{5\mathrm{D}} &\equiv \sqrt{G} \Big[\frac{1}{2} \, \tilde{G}^{MN} \left(\partial_M \hat{\Phi} \right) \left(\partial_N \hat{\Phi} \right) \Big] . \\ \mathcal{L}_{\Phi\Phi} &\equiv \sqrt{G} \Big[\frac{1}{2} \, \tilde{G}^{MN} \left(\partial_M \hat{\Phi} \right) \left(\partial_N \hat{\Phi} \right) \Big] . \\ \mathcal{L}_{\mathrm{pot}} &\equiv - \frac{4}{\kappa^2} \Big[\sqrt{G} \, V[\hat{\Phi}] + \sqrt{\overline{G}} \, V_1[\hat{\Phi}] \delta(\varphi) + \sqrt{\overline{G}} \, V_2[\hat{\Phi}] \delta(\varphi - \pi) \Big] . \\ \hat{\Phi}(x, y) &\equiv \frac{1}{\kappa} \hat{\phi}(x, y) \equiv \frac{1}{\kappa} \Big[\phi_0(y) + \hat{f}(x, y) \Big] , \end{split}$$

AISS MESMACHESMA

Introduce a Bulk Scalar field that acquires a ver $\phi_0(y)$ that varies with the dimension 5 coordinate.

- This kinetically mixes with the radion field and is the source of its mass.
- We end up with two towers of scalar fields one corresponding to the modulus and the other to the bulk scalar. We can use the residual gauge freedom to collapse this into a single tower



o The radion is identified as the lightest mode of the lower of scalars.

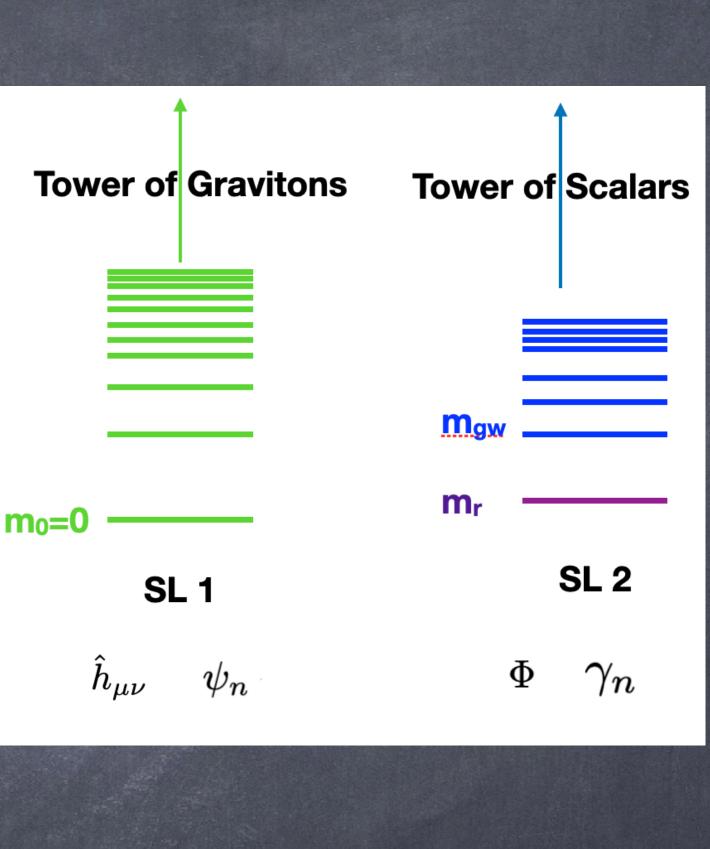
@ The order s² growth of the amplitude due to the mass of the radion is cancelled by the remainder of the scalar tower



 $\sum_{n=1}^{3} a_{n'n'(i)}^2 = \frac{5}{9} c_{n'n'n'n'} + \frac{1}{9} \mu_n^4 a_{nn0}^2 + \frac{1}{27} \mu_n^4 a_{nnn}$ 5 i=0





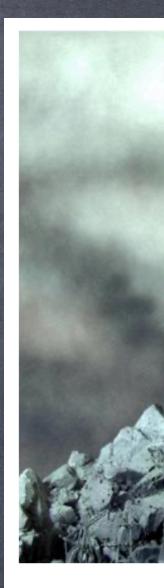






o The radion is identified as the Lightest mode of the lower of scalars.

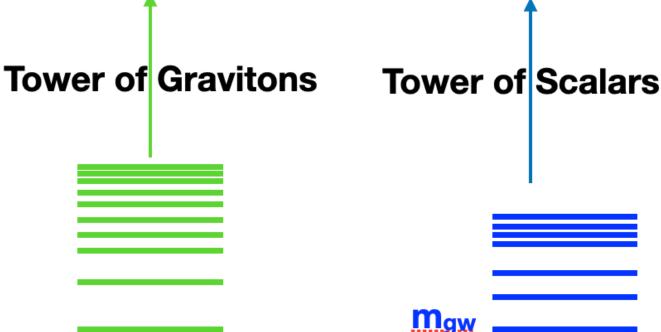
The order s² growth of the amplitude due to the mass of the radion is cancelled by the remainder of the scalar tower

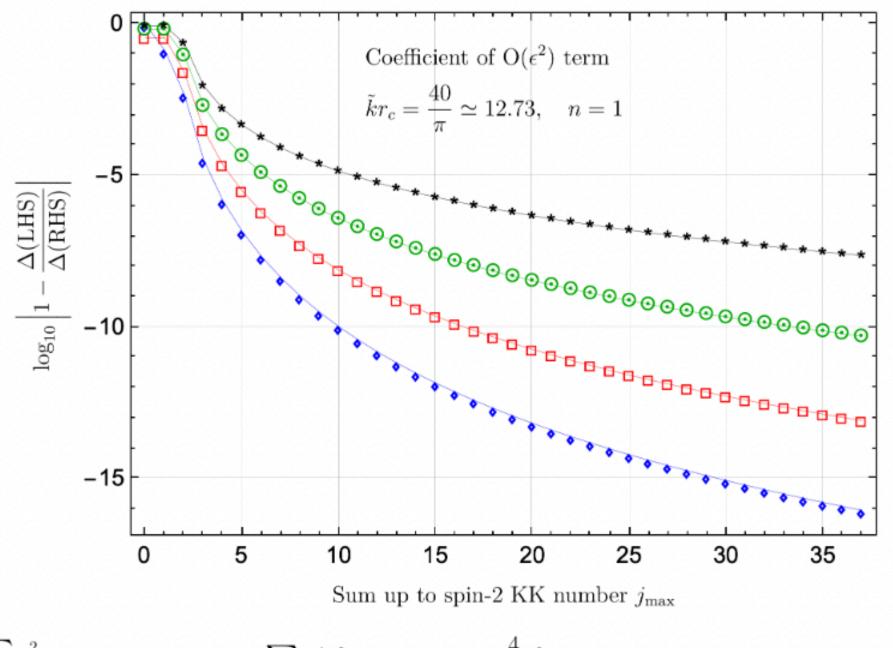


5 $\sum a_{n'n'(i)}^2 = \frac{5}{9}c_{n'n'n'} + \frac{1}{9}\mu_n^4 a_{nn0}^2 + \frac{1}{27}\mu_n^4 a_{nnn}$ i=0













- dimensional theories are tamed by diffeomorphism invariance.
- wavefunctions, couplings and masses- important for implementing in phenomenological applications.
 - @ Unstabilized model: Cancellations between single tower of states and couplings. a stabilized model: Cancellations between two towers Validity of 4D EFT - does not depend on modulus stabilization
 - @ 5D Planck Scale for flat stabilized models
 - $\Lambda \simeq M_{pl} e^{-kr_c}$ for warped RS model
- @ In order to demonstrate this

 - @ First calculation to involve full dynamics of a stabilized model.

Summary

a Apparent bad high energy growth of massive graviton scattering amplitudes in compactifield Extra

@ Uncovered sum rules that ensure this cancellation. This also provides a cross-check for our

Tetermined the Lagrangian of the stabilized model up to quartic order - this is non-trivial 14

