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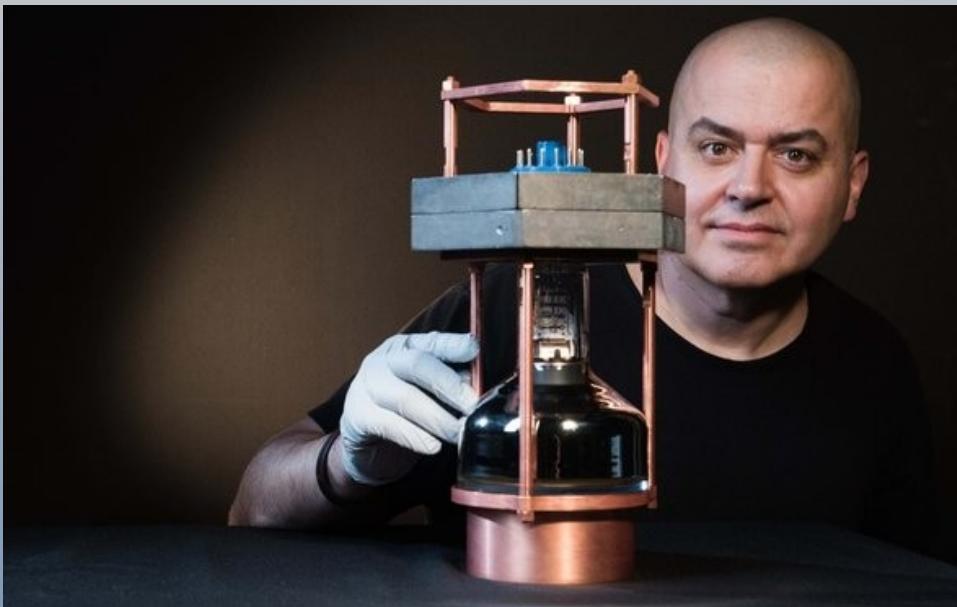
New physics in coherent neutrino scattering

Xun-Jie Xu

Max-Planck-Institut für Kernphysik, Heidelberg

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Talk at the NTN workshop on Neutrino Non-Standard Interactions:
<https://indico.cern.ch/event/812851/>



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1. Observation of Coherent Elastic Neutrino-Nucleus Scattering

COHERENT Collaboration (D. Akimov (Moscow, ITEP & Moscow Phys. Eng. Inst.) et al.).

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DOI: [10.1126/science.aao0990](https://doi.org/10.1126/science.aao0990)

e-Print: [arXiv:1708.01294 \[nucl-ex\]](https://arxiv.org/abs/1708.01294) | [PDF](#)

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[ADS Abstract Service](#); [Interactions.org article](#); [Link to Gizmodo article](#)

[Detailed record](#) - [Cited by 129 records](#) 100+

Outline

- Introduction to coherent neutrino scattering
 - What is it? Current status/Future potential
- New physics
 - NSI (Non-Standard Interaction)
 - SPVAT (S: 1, P: γ^5 , V: γ^μ , A: $\gamma^\mu\gamma^5$, T: $\sigma^{\mu\nu}$)
 - Light mediators
 - Sterile neutrinos
 - Neutrino magnetic moments
 - Dark matter
- Summary

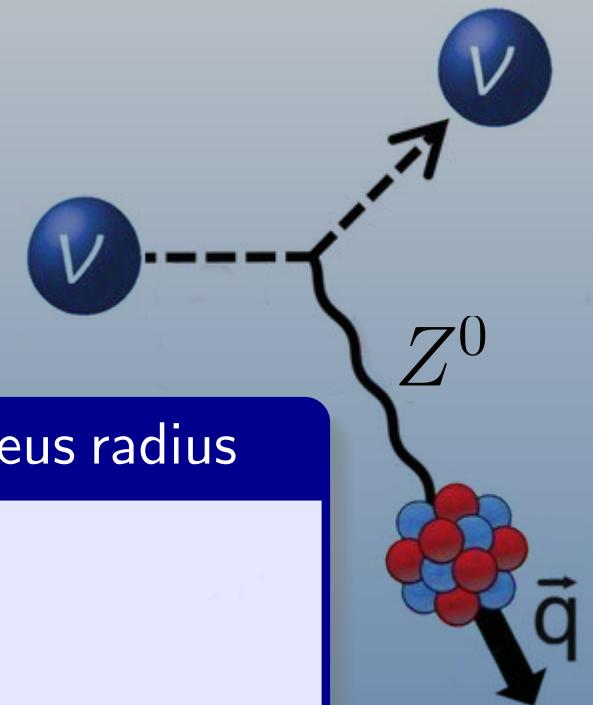
Coherent Elastic Neutrino-Nucleus Scattering

What is Coherent?

Without Coherency: neutrino wavelength \ll nucleus radius

Sum over cross sections

$$\sigma_{\text{tot}} = \sigma_p + \sigma_p + \cdots + \sigma_n + \sigma_n + \cdots$$



With Coherency: neutrino wavelength \gg nucleus radius

Sum over amplitude, then square

$$i\mathcal{M} = i\mathcal{M}_p + i\mathcal{M}_p + \cdots + i\mathcal{M}_n + i\mathcal{M}_n + \cdots$$

$$\sigma_{\text{tot}} = |i\mathcal{M}_p + i\mathcal{M}_p + \cdots + i\mathcal{M}_n + i\mathcal{M}_n + \cdots|^2$$

Cross section (full coherency)

$$\frac{d\sigma}{dT} = \frac{G_F^2 [N - (1 - 4s_W^2)Z]^2 M_{\text{nucleus}}}{4\pi} \left(1 - \frac{T}{T_{\max}}\right)$$

T : recoil energy, T_{\max} : max recoil, N & Z : neutron&proton numbers.

- Weak dependence on Z (proton number)

– because $1 - 4s_W^2 \approx 0$.

$$T_{\max} = \frac{2E_\nu^2}{M + 2E_\nu}$$

- Large cross section for large N

– $\frac{d\sigma}{dT} \propto N^2(N + Z)$, $\sigma \propto N^2$

- Although large, difficult to detect, T too low,

– T_{\max} determined by kinetics, 1 MeV neutrino $\Rightarrow T_{\max} \approx 0.1$ keV

- Modern tech: ultra-low threshold detection

– thanks to dark matter experiments

Coherency requires: $E_\nu < 50$ MeV.

Two suitable neutrino sources:

SNS

pros: higher $E_\nu \& T$, larger σ ,
 $T_{\text{thre}} = \mathcal{O}(10)$ keV enough

cons: $\downarrow T_{\text{thre}} \not\Rightarrow \uparrow N_{\text{tot}}$, com-
paratively low ν flux

Reactor neutrino

pros: ν flux high, high statis-
tics if 0.1 keV achieved

cons: needs very low T_{thre} ,
quenching factor unknown

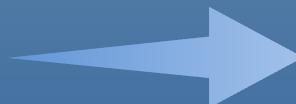
Example:

100 kg Ge

Threshold: 0.1 keV

1GW nuclear reactor

Distance: 10m

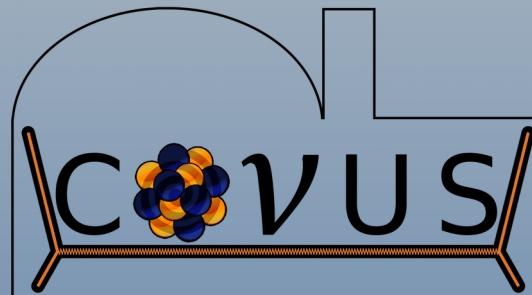


Event number:

$7.6 \times 10^6 / \text{yr}$

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$$7.6 \times 10^6 / \text{yr}$$



- Small, portable detector
- Precise measurement of neutrino interactions and EW parameters
 - E.g. $\sin^2 \theta_W = 0.238 \pm 0.0022$ (Lindner, et al, 1612.04150)
- Finding new physics

New physics

- NSI
 - 1804.03660, 1711.09773 , 1711.03521, 1708.04255 , 1708.02899, 1612.04150, 1805.01798, 1812.02778
- SPVAT (S: 1, P: γ^5 , V: γ^μ , A: $\gamma^\mu\gamma^5$)
 - 1806.07424, 1711.09773, 1612.04150, 1812.02778
- Sterile neutrinos
 - 1703.00054, 1711.09773, 1511.02834
- Light mediators (scalar, Z', dark photon...)
 - 1803.05466, 1803.01224, 1803.00060, 1802.05171, 1711.09773, 1711.04531, 1710.10889, 1612.06350, 1805.01798, 1508.07981, 1810.03626
- Neutrino electromagnetic properties
 - 1510.01684, 1706.02555, 1711.09773, 1805.01798, 1810.05606
- Dark matter
 - 1810.03626

New physics (NSI)

If mediated by a new gauge boson (e.g. Z'), integrated out \Rightarrow NSI.

Alternative: loop-induced NSI (I. Bischer, W. Rodejohann, X.J.X, 1807.08102).

NSI (Non-Standard Interaction)

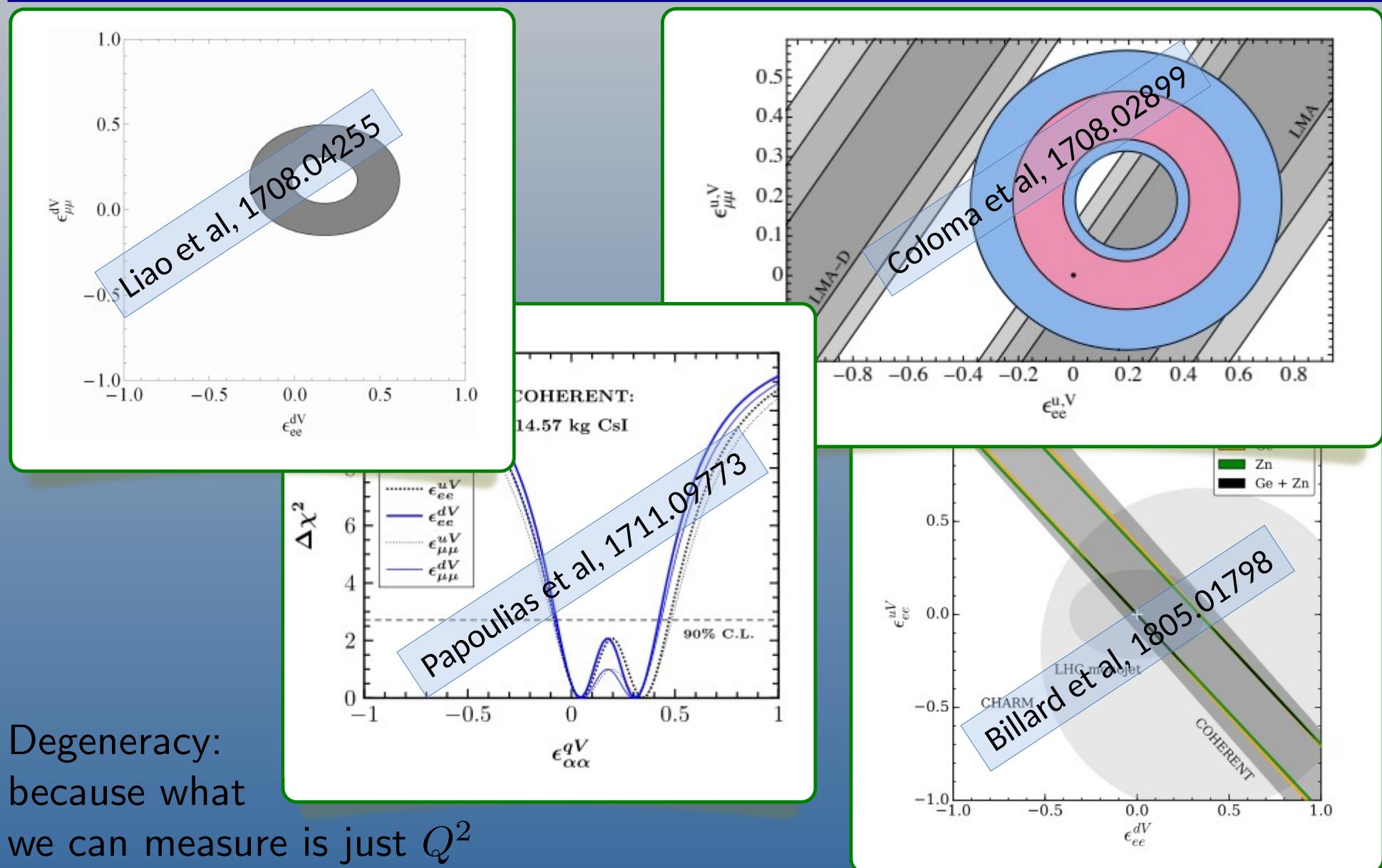
$$\mathcal{L} \supset \frac{G_F}{\sqrt{2}} \sum_{q=u,d} \bar{\nu}_\alpha \gamma^\mu (1 - \gamma^5) \nu_\beta \left[\bar{q} \gamma^\mu (\varepsilon_{\alpha\beta}^{qV} + \varepsilon_{\alpha\beta}^{qA} \gamma^5) q \right],$$

- Lepton Flavor Violation (LFV)
- Still V-A in $\bar{\nu}_\alpha \gamma^\mu (1 - \gamma^5) \nu_\beta$, because only left-handed ν
- the only change in $d\sigma/dT$, $\textcolor{red}{Q^2} : Q_{\text{SM}}^2 \rightarrow Q_{\text{NSI}}^2$

$$\frac{d\sigma}{dT} = \frac{G_F^2 \textcolor{red}{Q^2} M_{\text{nucleus}}}{4\pi} \left(1 - \frac{T}{T_{\max}} \right)$$

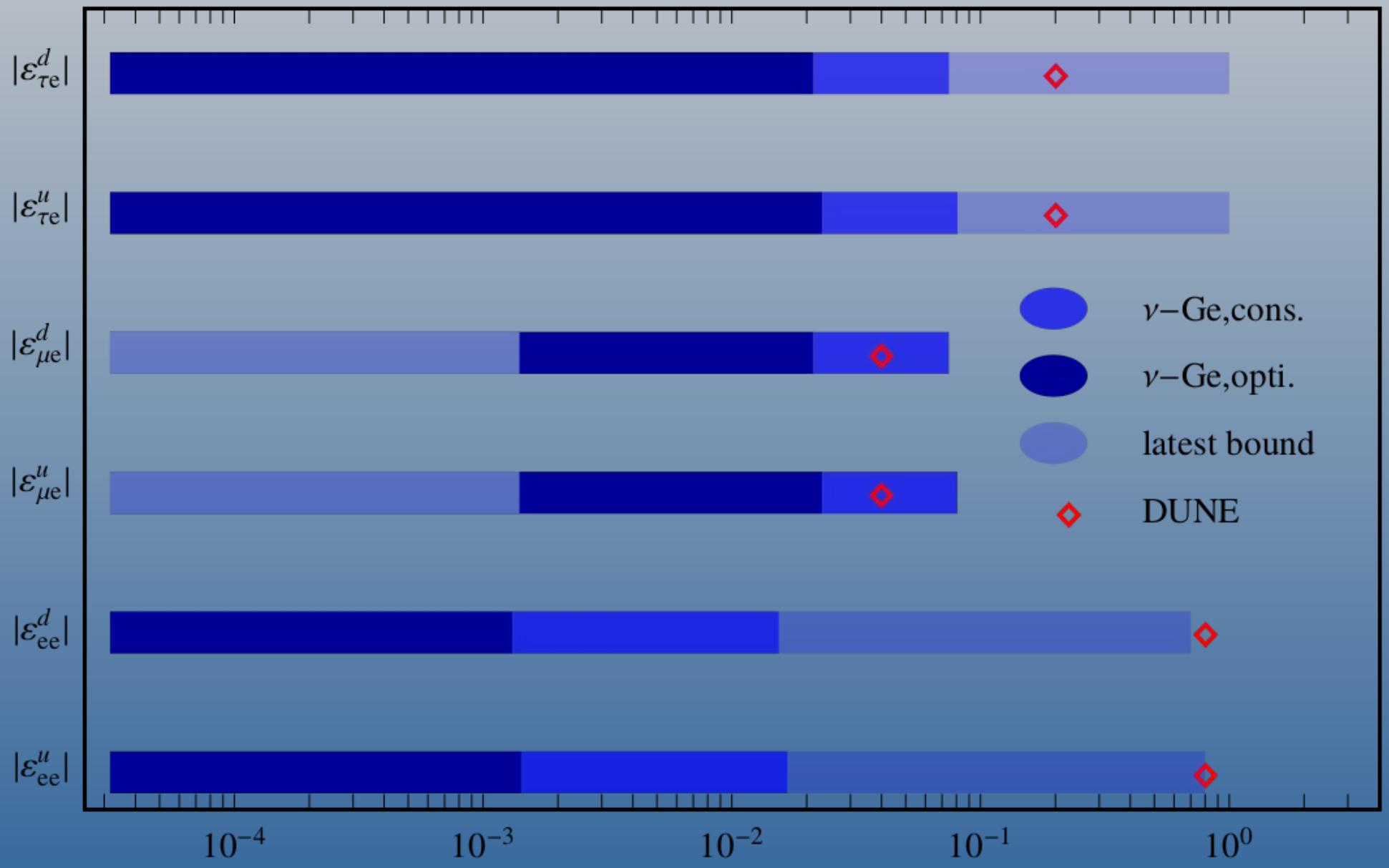
$$Q_{\text{NSI}}^2 = 4 \left[N(\cdots + \varepsilon_{ee}^{uV} + 2\varepsilon_{ee}^{dV}) + Z(\cdots + 2\varepsilon_{ee}^{uV} + \varepsilon_{ee}^{dV}) \right]^2 + \cdots$$

New physics (NSI)



$$Q_{\text{NSI}}^2 = 4 \left[N(\cdots + \varepsilon_{ee}^{uV} + 2\varepsilon_{ee}^{dV}) + Z(\cdots + 2\varepsilon_{ee}^{uV} + \varepsilon_{ee}^{dV}) \right]^2 + \cdots$$

New physics (NSI)



New physics (SPVAT)

If mediated by any kinds of forces, integrated out \Rightarrow SPVAT.

SPVAT (Scalar, Pseudo-S, Vector, Axial-V, Tensor

$$\mathcal{L} \supset \frac{G_F}{\sqrt{2}} \sum_{a=S,P,V,A,T} \bar{\nu} \Gamma^a \nu [\bar{\psi} \Gamma^a (C_a + D_a i \gamma^5) \psi] ,$$

$$\Gamma^a = \{1, i \gamma^5, \gamma^\mu, \gamma^\mu \gamma^5, \sigma^{\mu\nu} \equiv \frac{i}{2} [\gamma^\mu, \gamma^\nu]\}.$$

- Scalar (Pseudo-S) mediator $\Rightarrow 1 (i \gamma^5)$
- Charged scalar (Pseudo-S) mediator $\Rightarrow 1 (i \gamma^5) + \sigma^{\mu\nu}$
- Vector (Axial-V) mediator $\Rightarrow \gamma^\mu, \gamma^\mu \gamma^5$
- contains all possible Lorentz-invariant interactions
- involves ν_R . Light ν_R signal implies: ν is Dirac, not Majorana.

New physics (SPVAT)

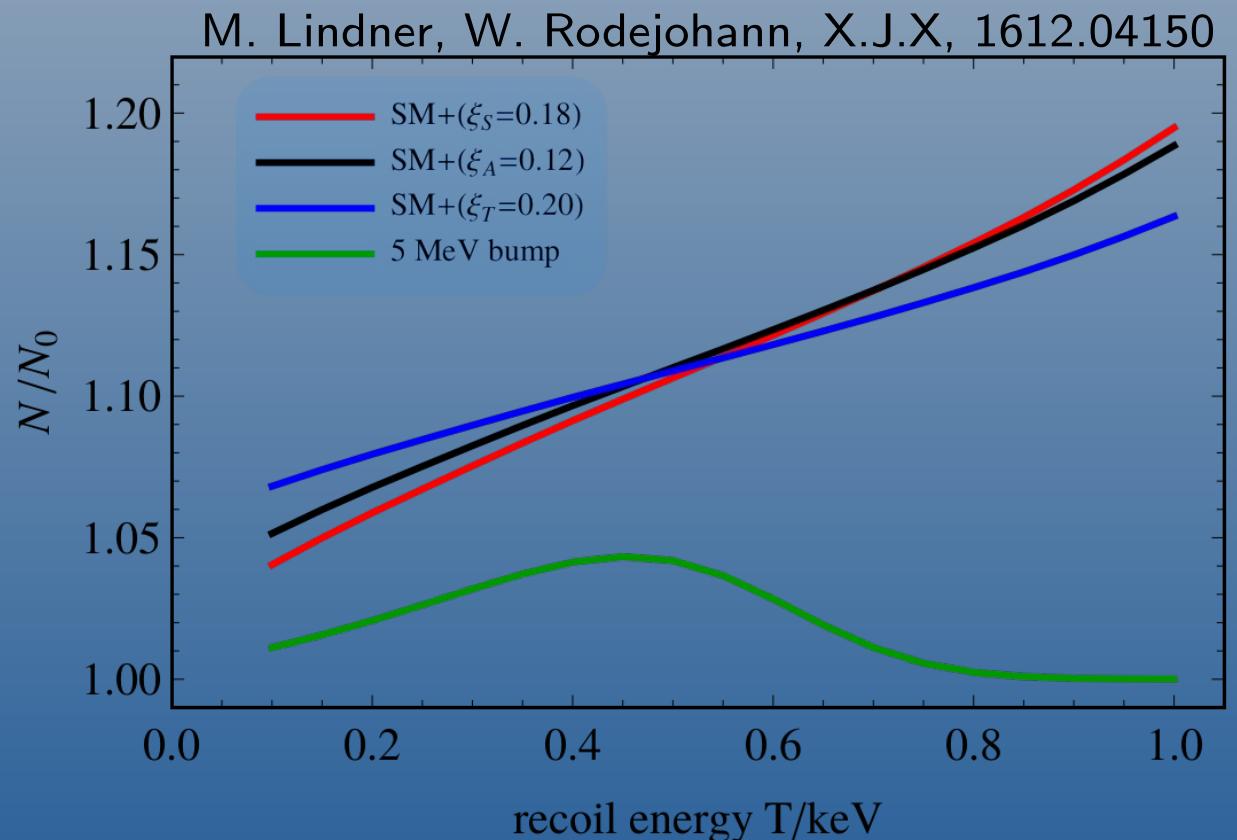
SM cross section:

$$\frac{d\sigma}{dT} = \frac{G_F^2 Q^2 M}{4\pi} \left(1 - \frac{T}{T_{\max}} \right)$$

SPVAT cross section:

$$\frac{d\sigma}{dT} = \frac{G_F^2 Q^2 M}{4\pi} \left([\dots] \times 1 - [\dots] \times \frac{T}{T_{\max}} \right)$$

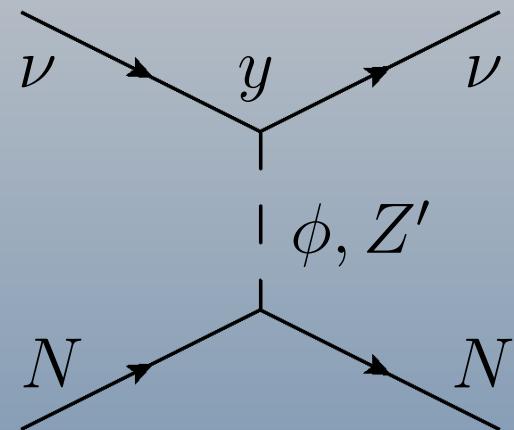
⇒ distortion of spectrum.



Light mediator (scalar, Z' , dark photon, etc.)

SM cross section:

$$\frac{d\sigma}{dT} = \frac{G_F^2 Q^2 M}{4\pi} \left(1 - \frac{T}{T_{\max}} \right)$$



Light mediator:

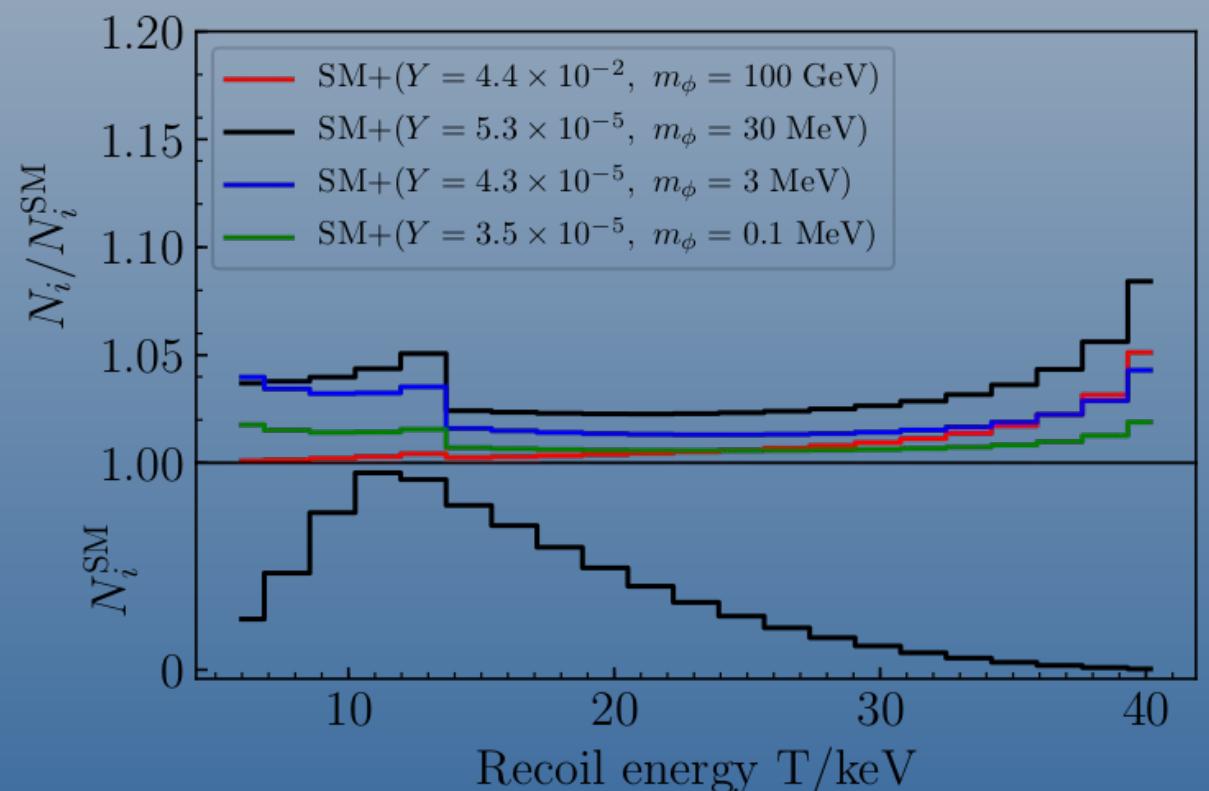
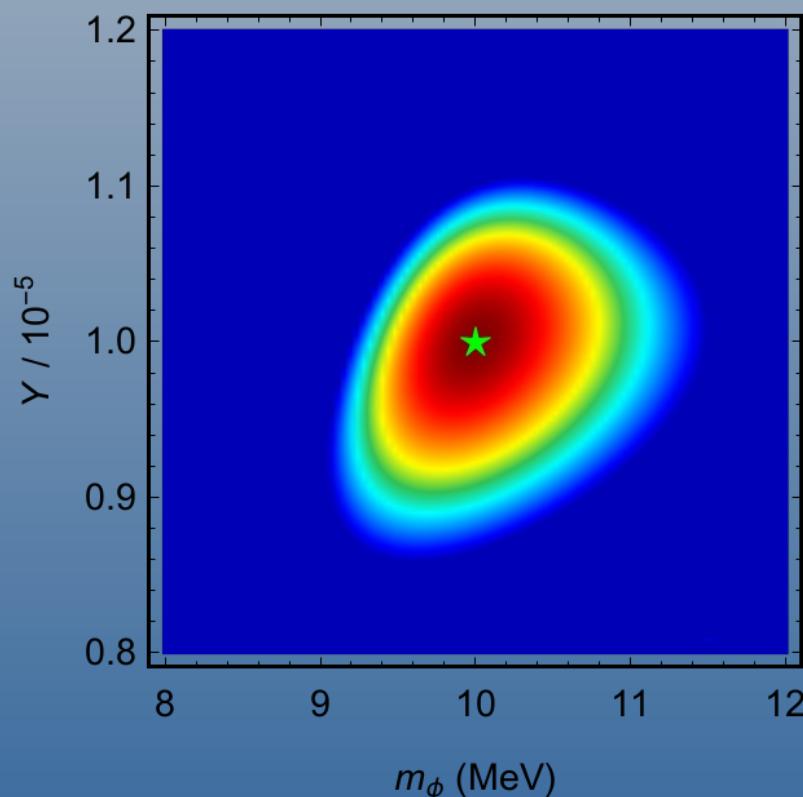
$$\frac{d\sigma}{dT} = \underbrace{\frac{y^4 M}{4\pi(2MT + m_\phi^2)^2}}_{\text{(i) effect of mediator mass}} \left(\underbrace{[\dots] \times 1 - [\dots] \times \frac{T}{T_{\max}}}_{\text{(ii) effect of mediator spin}} \right)$$

(i) effect of mediator mass (ii) effect of mediator spin

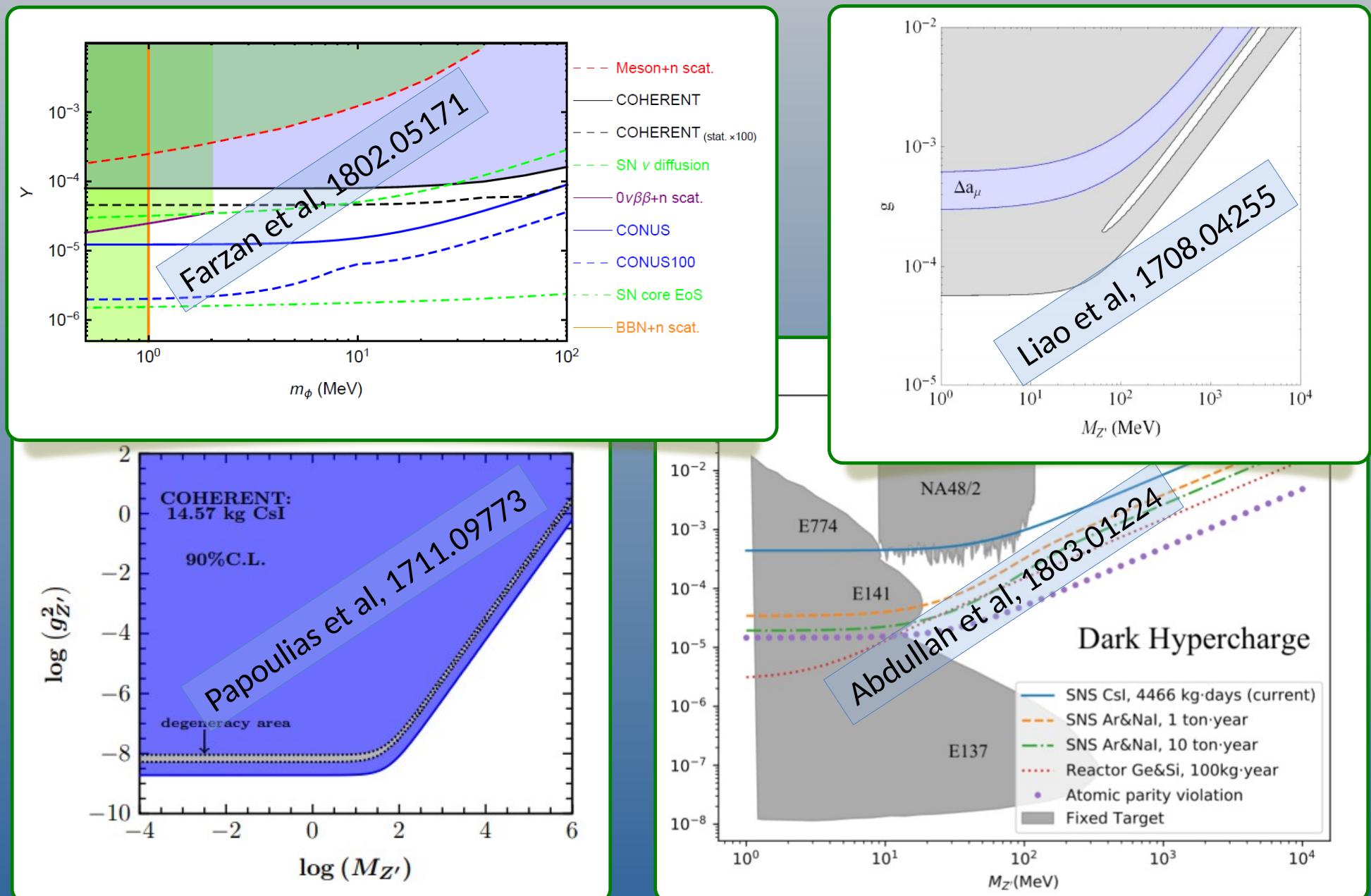
⇒ distortion of spectrum due to (i) and (ii).

Light mediator (scalar, Z' , dark photon, etc.)

Distortion of spectrum \Rightarrow distinct signal
 \Rightarrow to reconstruct the mediator mass & coupling

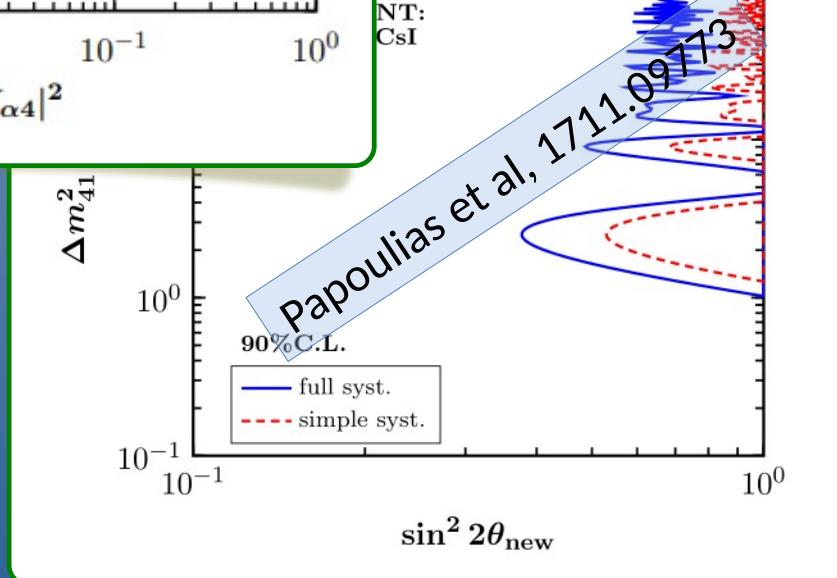
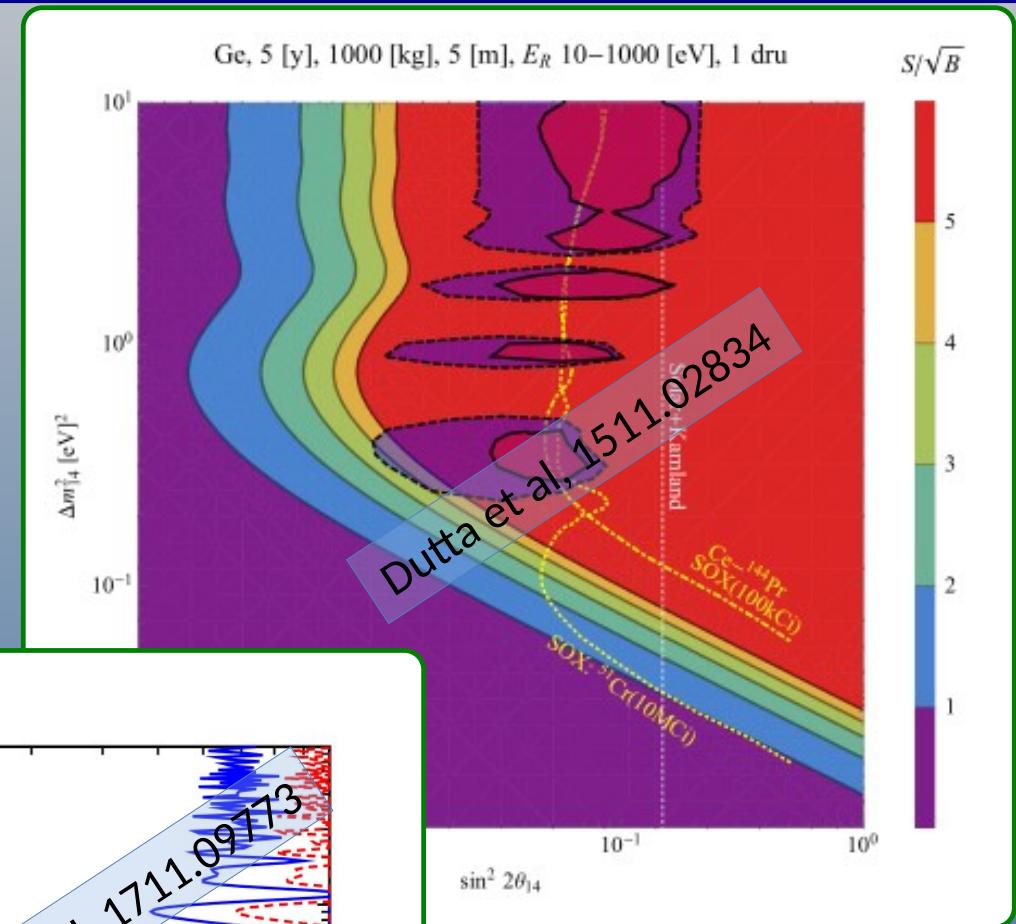
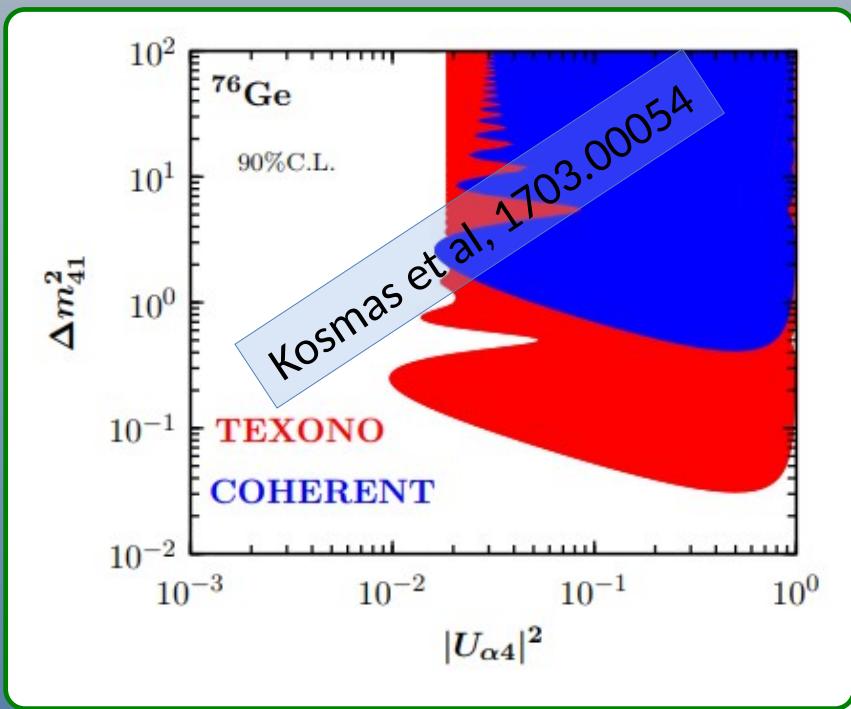


Light mediator (scalar, Z' , dark photon, etc.)

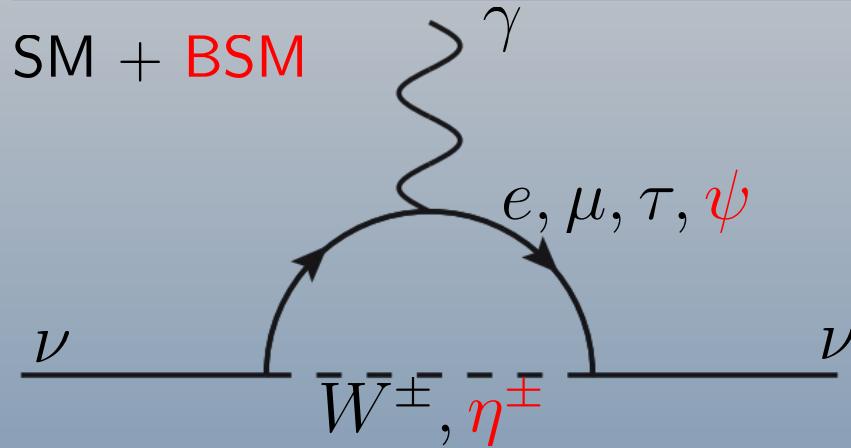


Sterile neutrinos

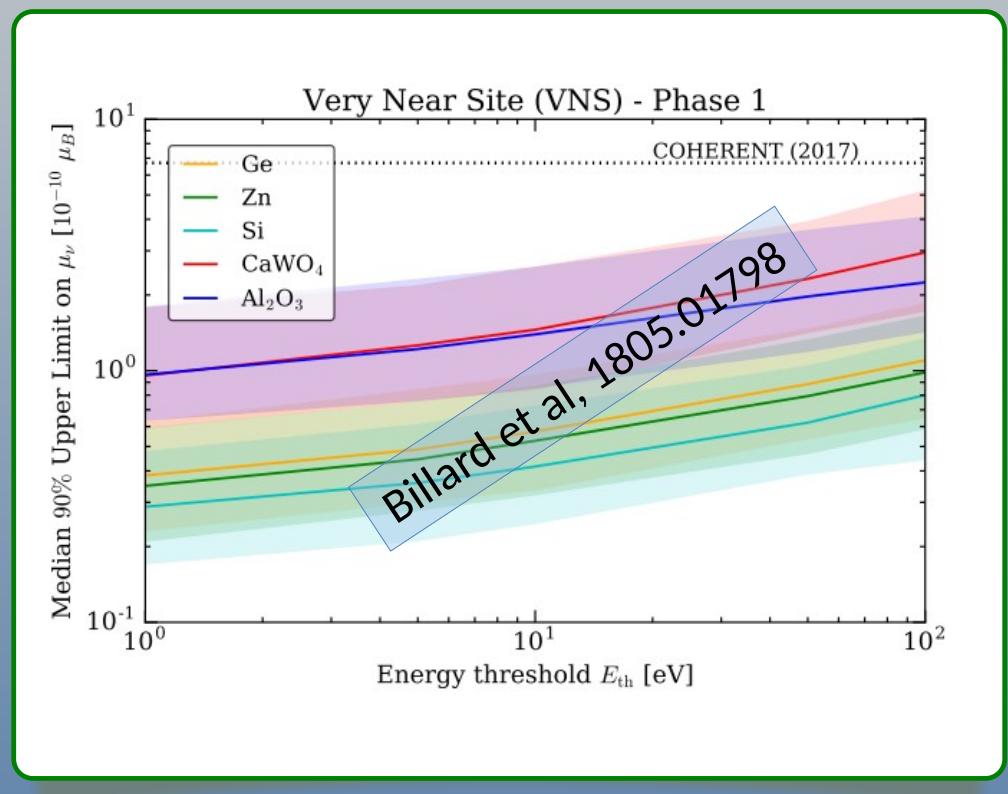
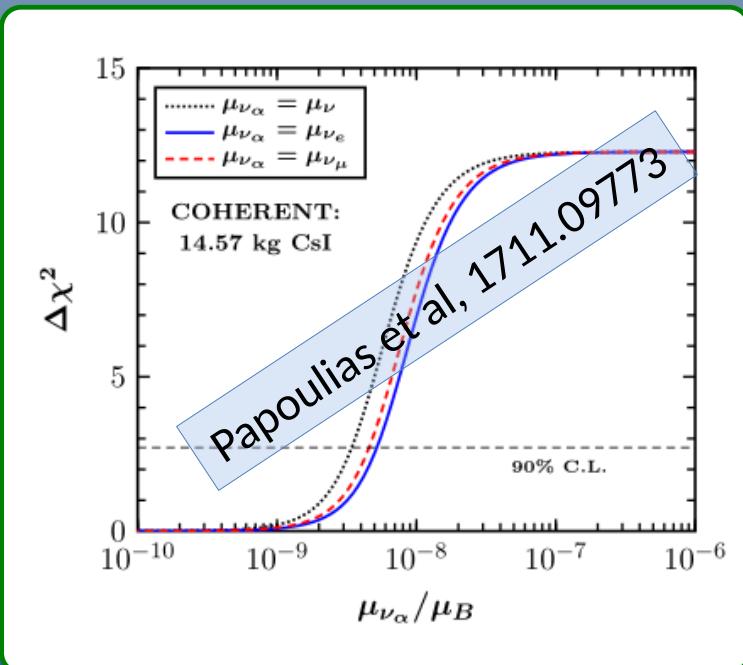
$$P_{ee} = 1 - \sin^2 2\theta_{14} \sin^2 \left(\frac{\Delta m_{41}^2 L}{4E_\nu} \right)$$



Neutrino magnetic moments



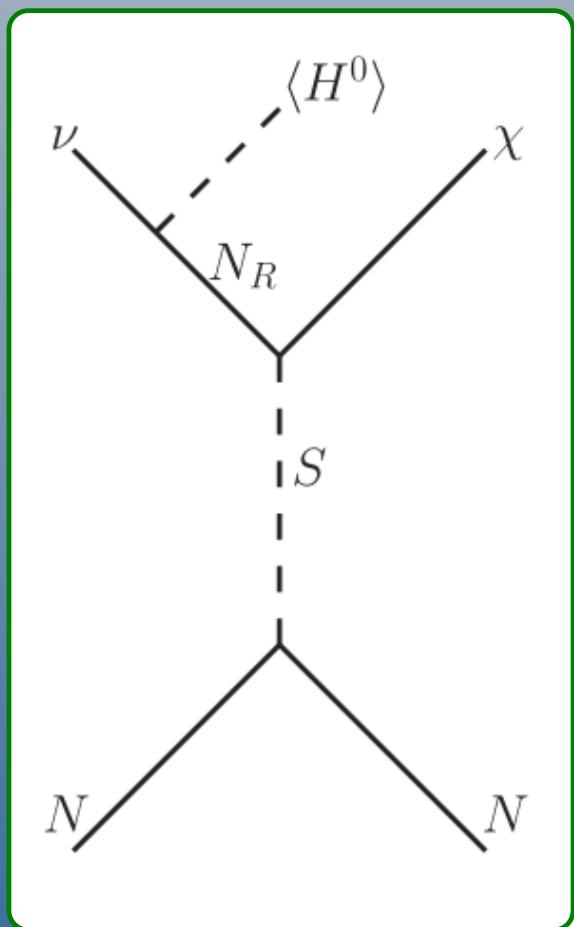
SM : $\mu_\nu \sim 10^{-20} \mu_B$
 chirality flipping $\Rightarrow \mu_\nu \uparrow$ to $10^{-9} \mu_B$
 [X.J.X, 1901.00482]



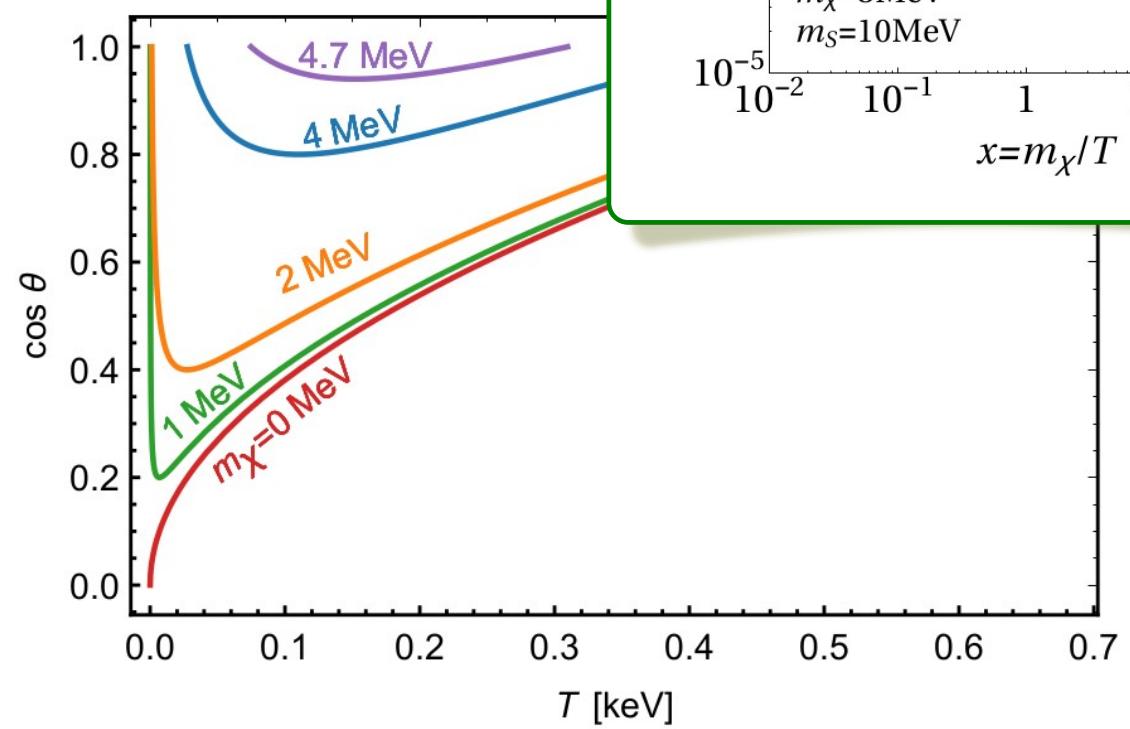
- $\nu + e$ scattering is better:
GEMMA: $\mu_\nu < 10^{-11} \mu_B$
- Also constraint on charge radius:
 $-8 \times 10^{-32} < \langle r^2 \rangle < 11 \times 10^{-32} \text{ cm}^2$
[M. Cadeddu, et al, 1810.05606]

Dark matter

BSM particles could appear in the final states.
Invisible \Rightarrow Dark matter?



V. Brdar, W. Rodejohann,
X.J.X, 1810.03626



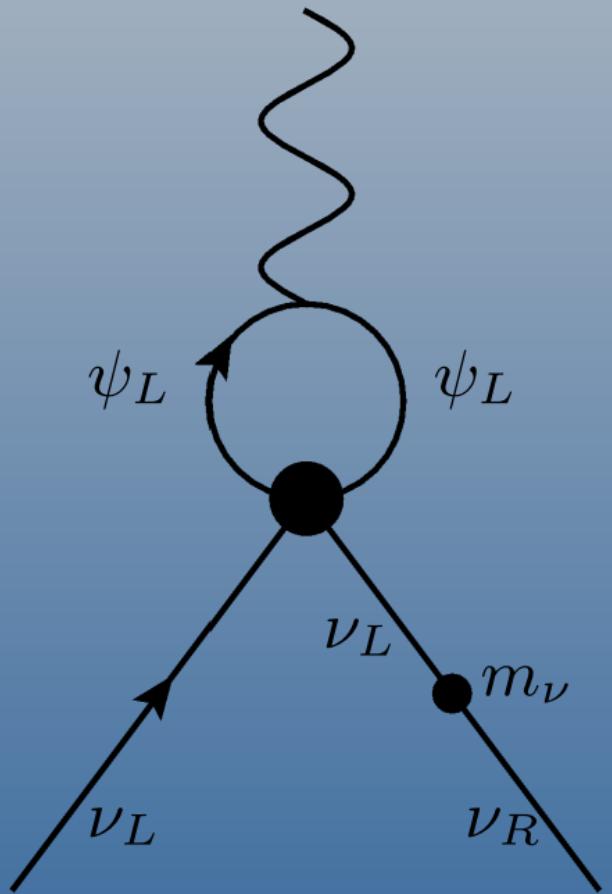
Summary

- Coherent neutrino scattering \Rightarrow a new channel to probe SM & BSM neutrino interactions;
- Recent experiments \Rightarrow constraints on NSI, SPVAT, sterile neutrinos, light mediators, neutrino magnetic moments;
- Future reactor-based experiments \Rightarrow very high statistics \Rightarrow more neutrino-related new physics.

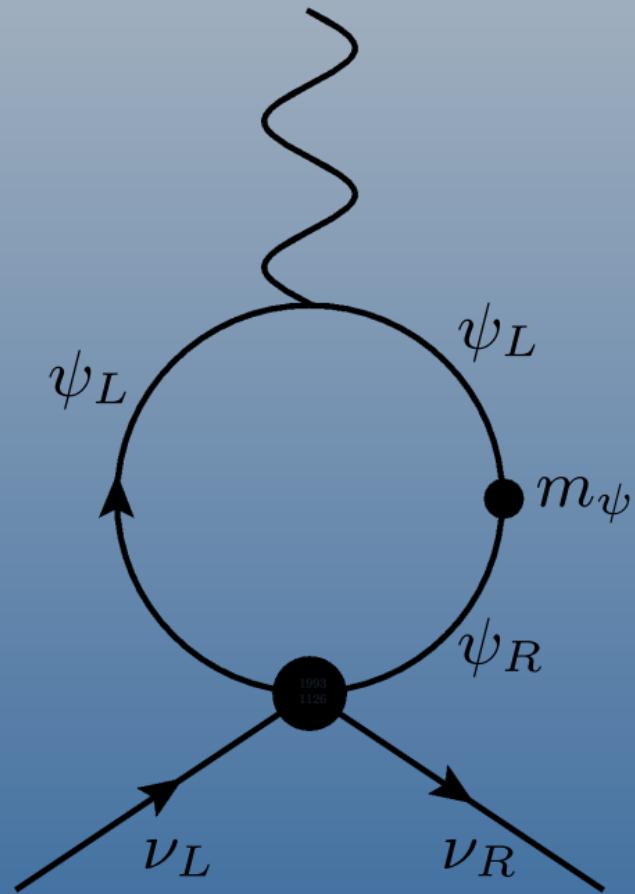
Backup

Backup

Chirality flipping and neutrino magnetic moment



(a): $\mu_\nu \propto m_\nu$



(b): $\mu_\nu \propto m_\psi$