

Light Dark Matter at the DUNE Near Detector

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Outline

- •Dark Photon + Light Fermion/Scalar DM
- •Thermal Relic DM
- **•**DUNE-PRISM
- •On- and off-axis searches for DM

- Mono-neutrino Events and Neutrinophilic DM
 - •Extending the SM with (B-L) Symmetry
 - Large missing-momentum events at the DUNE Near Detector
 - •Searching for thermal relic neutrinophilic DM at DUNE



Light Dark Photon + Dark Matter [1903.10505]

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Dark Photon + Light DM Model of Interest

 $\mathscr{L} \supset -\frac{\varepsilon}{2} F^{\mu\nu} F'_{\mu\nu} + \frac{M_{A'}^2}{2} A'_{\mu} A^{\prime\mu} + \overline{\chi} i \gamma^{\mu} \left(\partial_{\mu} - i g_D A'_{\mu}\right) \chi - M_{\chi} \overline{\chi} \chi.$

Or

In a fixed-target environment, many neutral mesons that can decay $\mathfrak{m} \to \gamma \gamma$ are produced. With suitable masses, they can decay instead by



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(or similar with scalar DM)







Where in parameter space are we looking?





Detection Signature

•Some fraction of the produced DM will travel in the direction of the DUNE Near Detector.

• Via kinetic mixing with the standard model photon, the DM can scatter off nucleons or electrons in the detector.







Let's talk about Backgrounds

- Signal looks identical to neutrino-nucleus neutral-current scattering or $\nu_{\mu}e^- \rightarrow \nu_{\mu}e^-$
- If just performing a counting experiment, this strategy will be background dominated.
- •Going beyond a counting experiment is difficult shape of the neutrino flux (in energy space) is constrained using $\nu_{\mu}e^- \rightarrow \nu_{\mu}e^-$ measurements!
- •Our proposed solution: utilize the DUNE-PRISM proposal.





DUNE-PRISM



Proposal: move the near detector transverse to the beam direction. Relative flux at different off-axis positions is predictable from meson decay kinematics. Neutrino beam is focused (coming from charged mesons), but DM flux is unfocused!



Signal-to-Background as off-axis angle increases



This shape should be well-predicted (meson decay kinematics), even if the energy spectrum of the neutrinos cannot be constrained as well.

Antineutrino Mode



Vetoing the ν_e CCQE Background with Kinematics



Even if final-state proton is lost in CCQE event, kinematics of the electron may allow us to effectively veto all of these events.





Results – Counting Experiments on- and off-axis, Fermionic DM



Results with Scalar DM are similar, although relic density targets are closer to being achievable. Nucleon scattering significantly weaker (see [1903.10505]).





Broader, two parameter search



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Neutrinophilic Dark Matter [1802.00009], [1901.01259]

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If (B-L) is a good symmetry of nature, what are the consequences?

- Neutrinos are Dirac Particles

•Extensions beyond the standard model are non-trivial. Let's focus on adding a scalar ϕ







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Only renormalizable operator allowed:

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 $\mathscr{L} \supset \frac{\lambda_c^{\eta}}{2} \nu_i^c \nu_j^c \phi^{\dagger}$







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Extending to higher operator dimension:

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 $\mathscr{L} \supset \frac{\lambda_c^{\prime \prime}}{2} \nu_i^c \nu_j^c \phi^{\dagger} + \frac{1}{\Lambda_{\alpha\beta}^2} \left(L_{\alpha} H \right) \left(L_{\beta} H \right) \phi + \text{h.c.}$







If (B-L) is a good symmetry of nature, what are the consequences?

Neutrinos are Dirac Particles

After EWSB:

 $\mathscr{L} \longrightarrow \frac{\lambda_c^{ij}}{2} \nu_i^c \nu_j^c \phi^{\dagger} + \frac{\lambda_{\alpha\beta}}{2} \nu_{\alpha} \nu_{\beta} \phi + \frac{\lambda_{\alpha\beta}}{\nu} \nu_{\alpha} \nu_{\beta} \phi h + \mathcal{O}(h^2),$

•Extensions beyond the standard model are non-trivial. Let's focus on adding a scalar ϕ

 $\Lambda^2_{\alpha\beta}$





Most constraints: rare decays when ϕ is kinematically accessible: Higgs, charged mesons, and double-beta-like decays.



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•Rare meson decays, e.g. $K^+ \to \mu^+ \overline{\nu}_{\mu} \phi$ as a contribution to the search for $K^+ \to \mu^+ \nu_{\mu} \nu \overline{\nu}$





Most constraints: rare decays when ϕ is kinematically accessible: Higgs, charged mesons, and double-beta-like decays.

• " ϕ -full double-beta decay" as a contribution to neutrino-full $2\nu\beta\beta$







Constraints in muon-coupled Channel



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An Aside, MiniBooNE/LSND



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An Aside, MiniBooNE/LSND



 $m_{\phi} \sim 50$ MeV, with flavor-violating vertex, could explain MiniBooNE!

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Neutrino Beamsstrahlung – Monov Events

explore what happens when it's kinematically accessible in scattering.

SM CC Interaction

Previous constraints relied on the mediator being kinematically accessible in a decay. Let's

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Neutrino Beamsstrahlung – Monov Events

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Beamsstrahlung

Key Characteristics of Mono ν

- •Wrong-sign lepton generated: neutrino beam produces positively-charged lepton.
- ϕ carries away significant energy reconstructing the event as a 2 to 2 scattering provides a small neutrino energy.
- ϕ carries away significant transverse momentum — reconstructed visible energy has large missing momentum.

Signal Distributions

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Signal distribution depends on scalar mass. Lines encompass 90% of signal events.

Background distribution should live on y-axis.

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Exploiting Wrong-Sign Leptons

Two opportunities for determining muon charge in the DUNE Near Detector:

- Positively-charged muons always decay to a Michel electron (a). Only 25% of negatively-charged muons do (b). Our signal then can be events that produce a Michel electron (when operating in neutrino mode).
- •With the proposed Gas Argon TPC, muons (energies above ~1 GeV) that reach the gas (c) may be charge-identified due to the magnetic field.

DUNE Sensitivity Reach (10 years Data Collection)

500 MeV bins in reconstructed energy, 125 MeV bins in transverse momentum. (95% CL)

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 $\frac{1}{6}y_3\chi^3\phi$

 $\frac{1}{2}y_S\chi^2\phi$

 $\frac{1}{2} y_F \bar{\chi}^c \chi \phi$

 $y_I \chi_1 \chi_2 \phi$

The Dark Matter Connection

particle(s)?

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We assumed that ϕ had (B-L) charge 2. What if it is a mediator to some dark matter

Scalar DM with (B-L) charge -1

Fermion DM with (B-L) charge -1

Scalar DM with (B-L) charge -2/3

Inelastic Fermion DM with (B-L) charges q, q-2.

particle(s)?

The Dark Matter Connection

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DUNE Reaching DM Targets Assumptions: DM couples to ϕ as strongly as neutrinos, $m_{\phi} = 10m_{DM}$

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Conclusions

- DUNE will allow us to search for thermal DM in a variety of interesting ways.
 - DUNE-PRISM can do more than cross section measurements.
 - •Focusing of neutrinos vs. DM provides interesting event spectra at on- and off-axis locations.
 - •Sensitivity can reach theoretically motivated targets, especially when considering scalar DM
 - Inspired by collider mono-jet searches, mono-neutrino searches at the DUNE ND can allow for sensitivity to neutrinophilic scalars.
 - •Scalar emission allows for events with large missing transverse momentum, low reconstructed neutrino energy, and (in this model) wrong-sign final-state muons.
 - •Light dark matter may couple to the same mediator DUNE can probe relic DM with this search strategy.

