

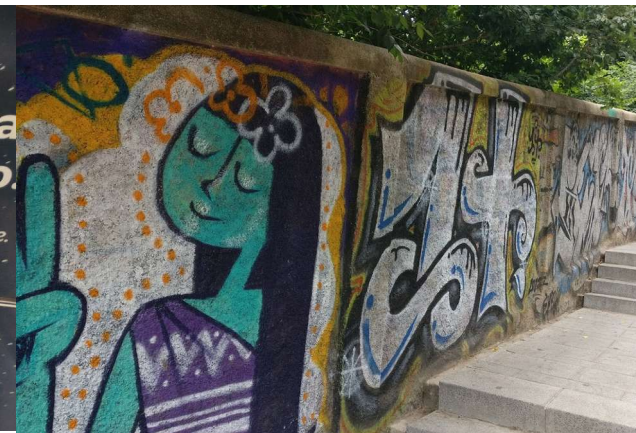
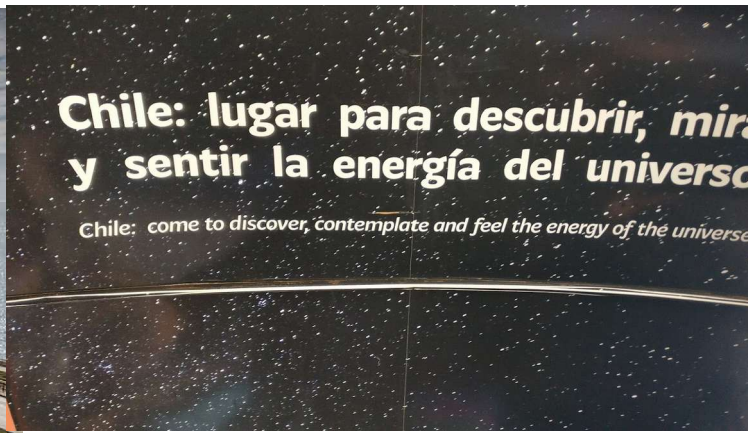
# Vector Dark Matter via a Fermionic Portal from a New Gauge Sector

Alexander Belyaev



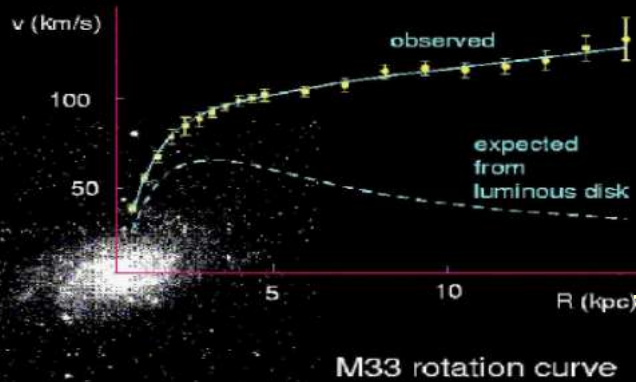
Southampton University & Rutherford Appleton Laboratory

2203.04681 and 2204.03510 AB, Luca Panizzi, Aldo Deandrea, Stefano Moretti and Nakorn Thongyoi

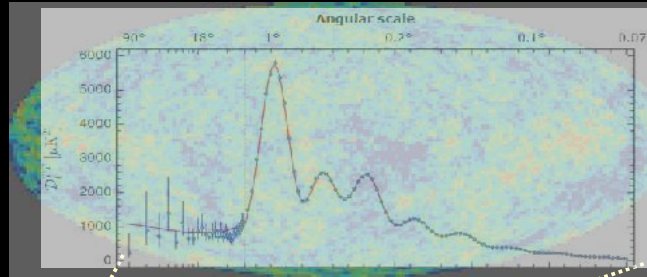


# The existence of Dark Matter is confirmed by several independent observations at cosmological scale

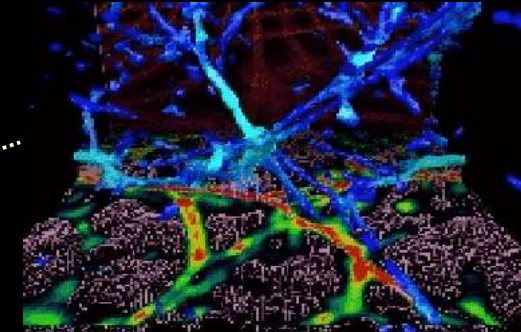
## Galactic rotation curves



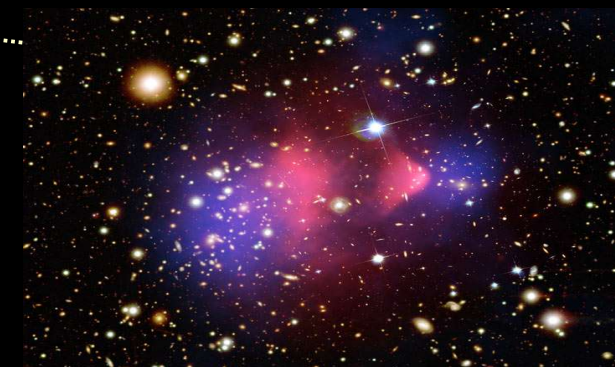
## CMB: WMAP and PLANCK



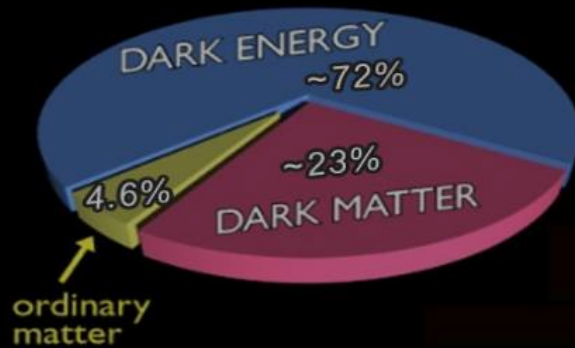
## Large Scale Structures



## Bullet cluster



## Gravitational lensing



# DM is very appealing even though we know almost nothing about it!

**Spin**

**Mass**

**Stable**

Yes

No

symmetry behind  
stability

**Couplings**

gravity

weak

higgs

quarks/gluons

leptons

New mediators

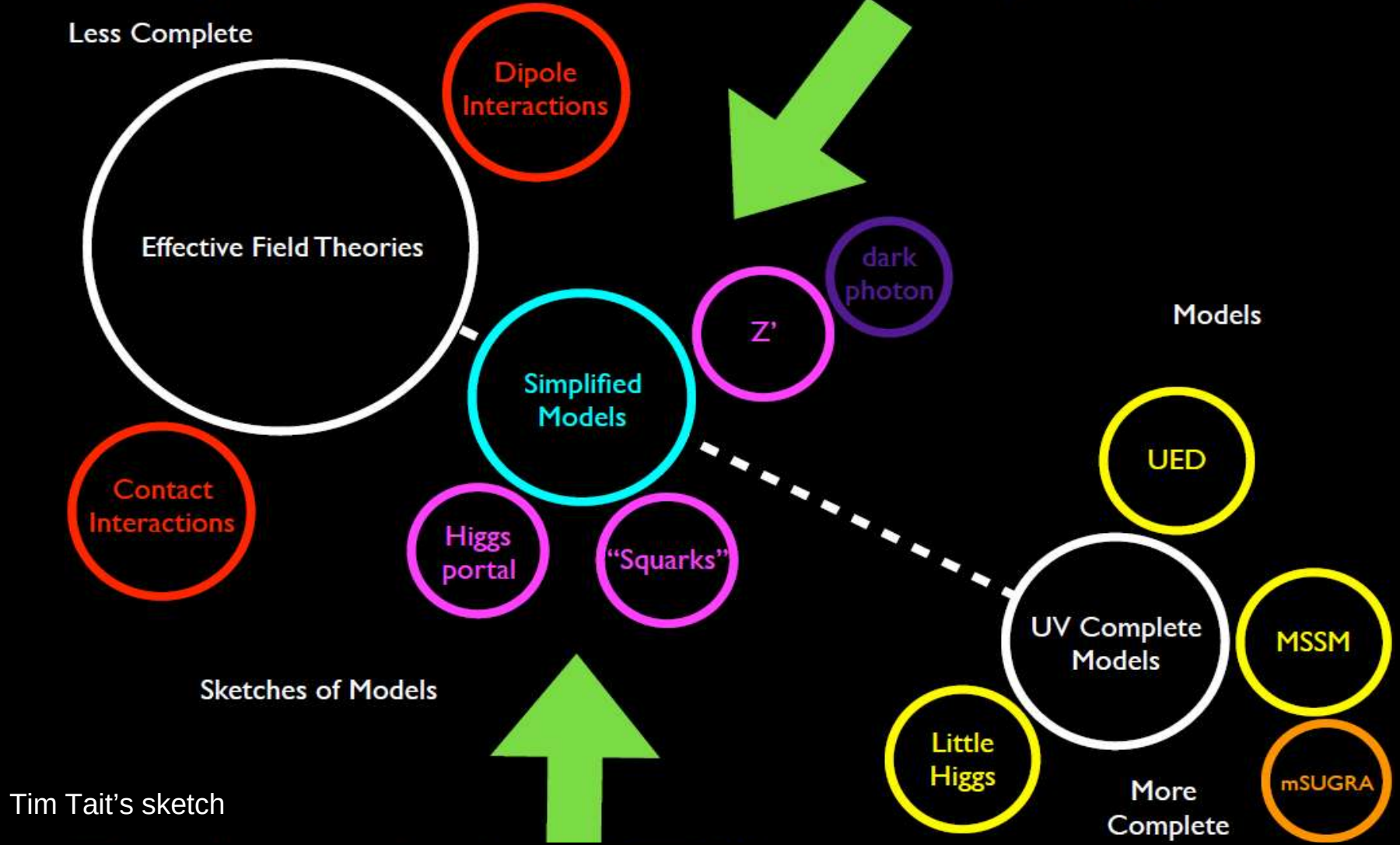
**Thermal relic**

Yes

No

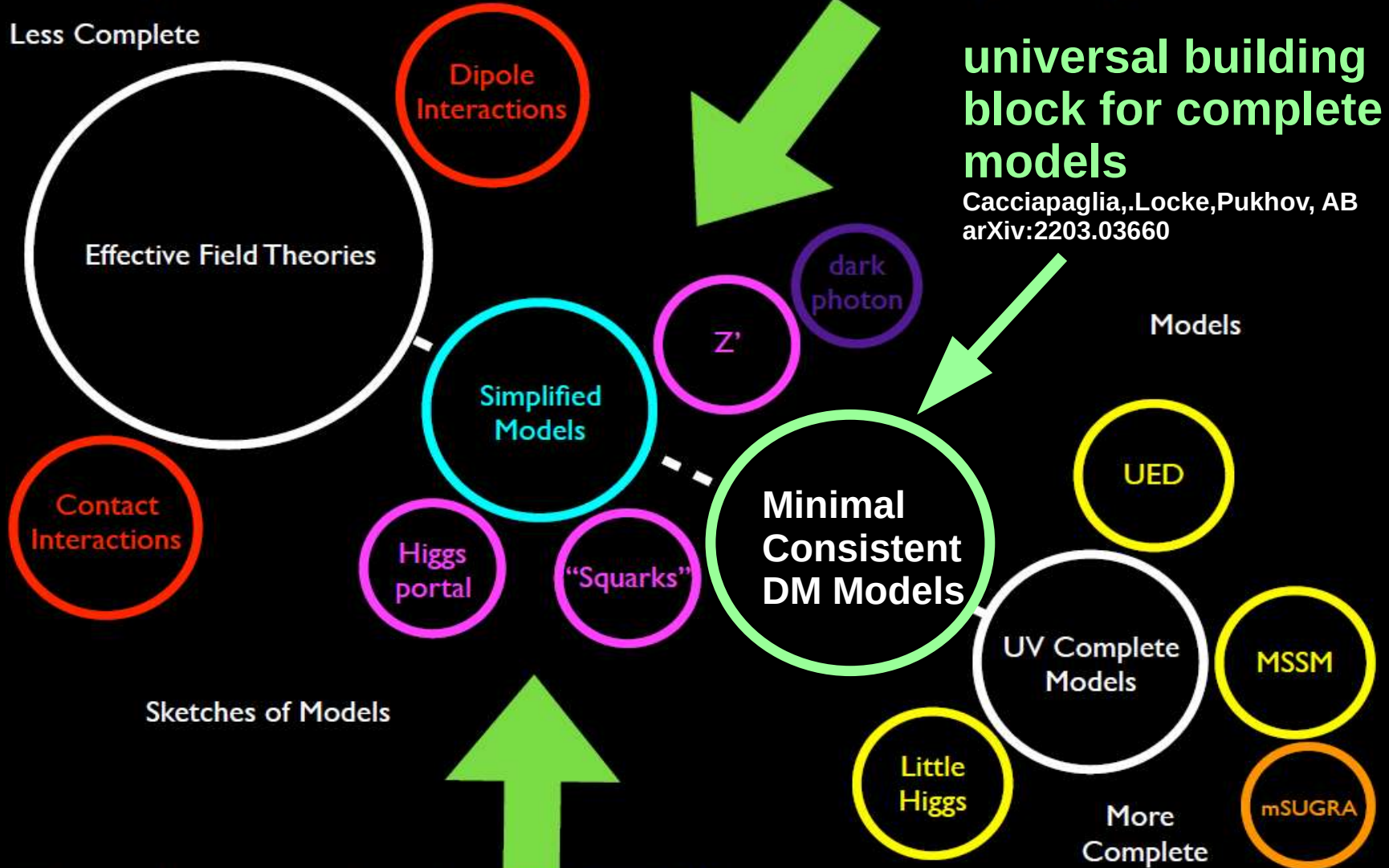


# Spectrum of Theory Space



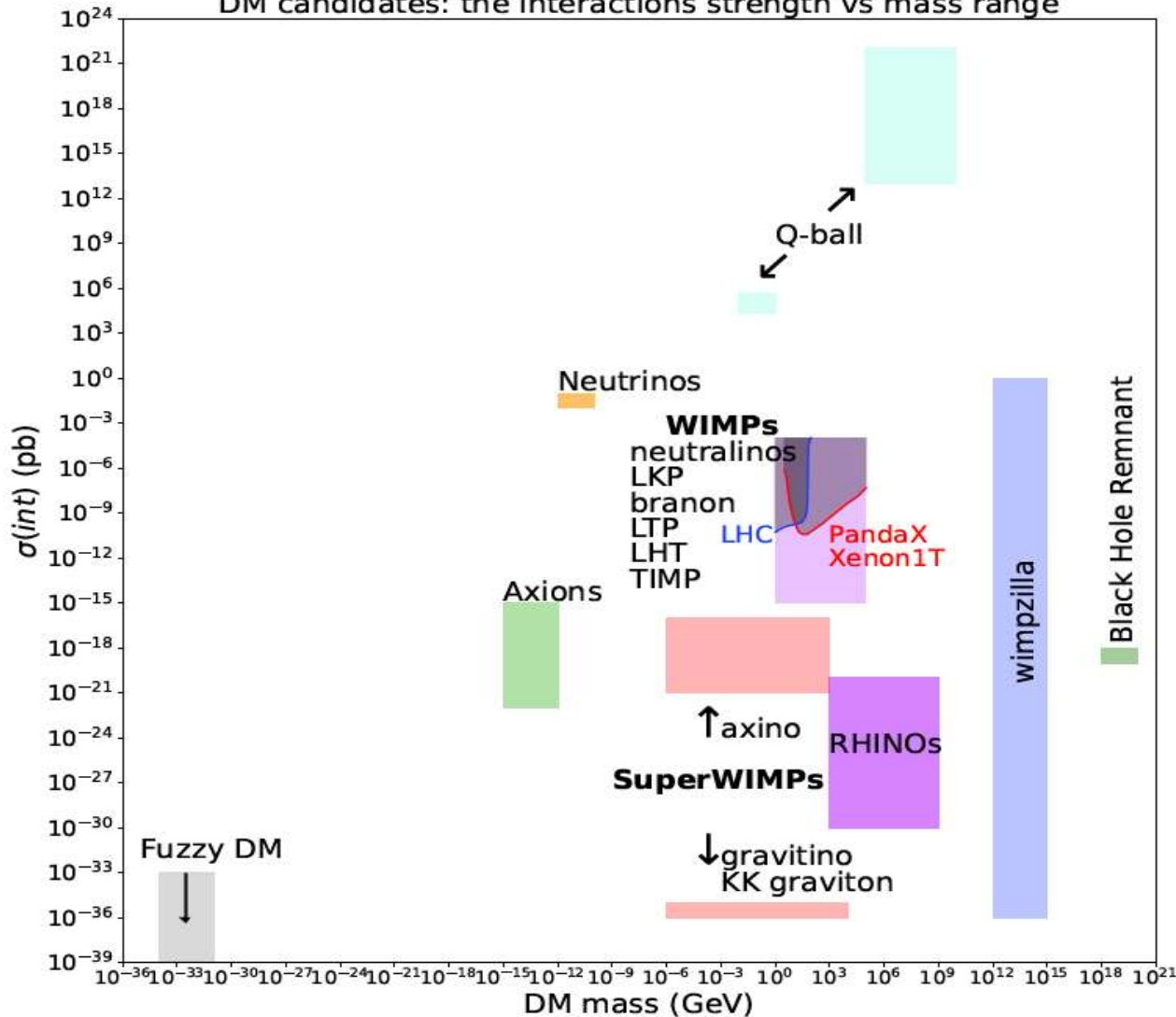
# Spectrum of Theory Space

Less Complete



# DM candidates: interaction vs mass

DM candidates: the interactions strength vs mass range



**Planck mass BH** remnants: tiny black holes protected by gravity effects [Chen '04] from decay via Hawking radiation

**Wimpzillas**: very massive non-thermal WIMPs [Kolb, Chung, Riotto '98]

**Q-balls**: topological solitons that occur in QFT [Coleman '86]

**EW scale WIMPs**, protected by parity – LSP, LKP, LTP particles

**SuperWIMPs**: electrically and color neutral DM interacting with much smaller strength (perhaps only gravitationally)

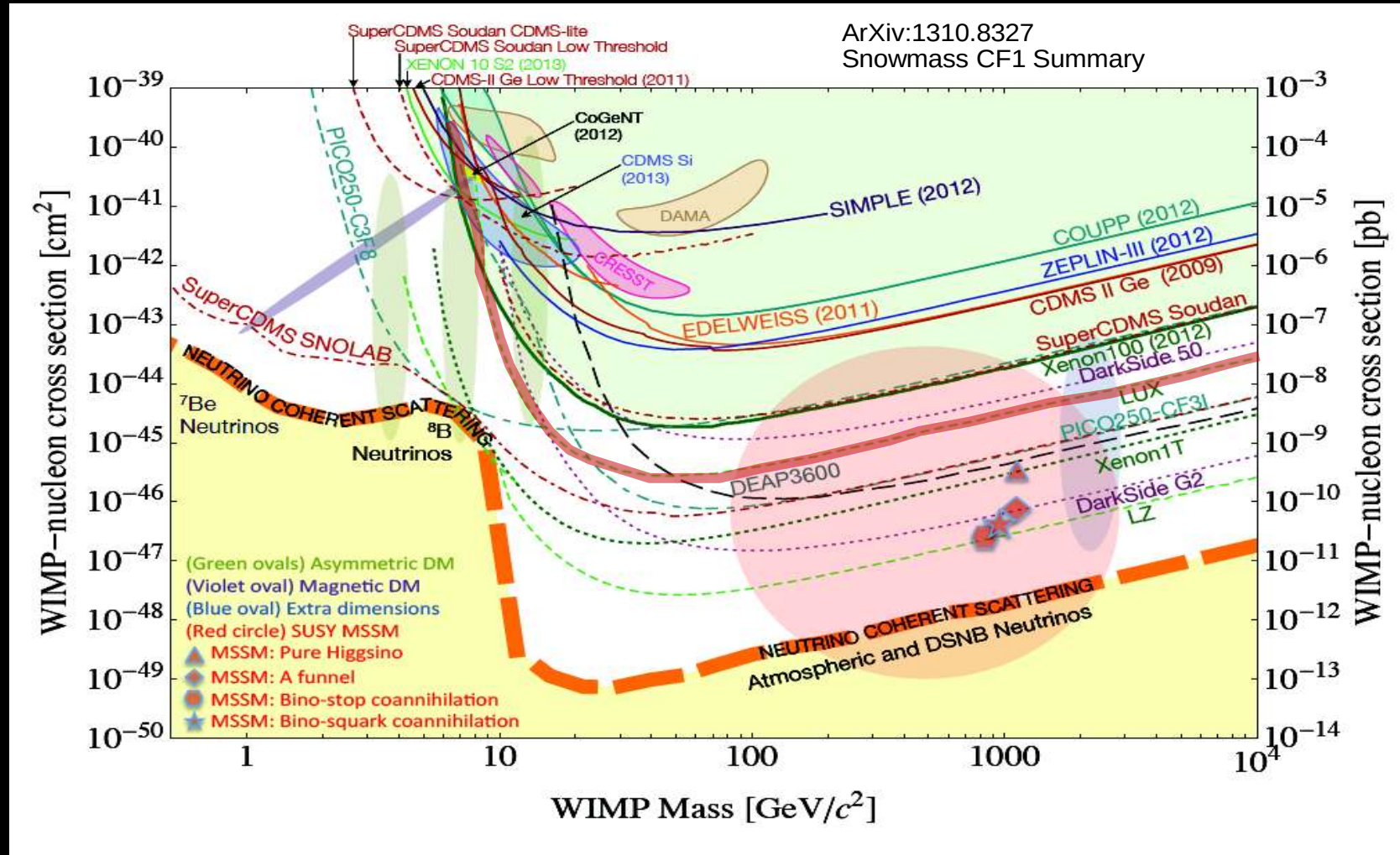
**Neutrinos**: usual neutrinos are too light-HDM, subdominant component only (to be consistent with large scale structures); but heavier gauge singlet neutrinos can be CDM

**Axions**: 
$$\frac{\theta_{QCD}}{32\pi^2} F^{\mu\nu} \tilde{F}^{\mu\nu}$$

$\theta_{QCD}$  is replaced by a quantum field, the potential energy allows the field to relax to near zero strength, axion as a consequence



# Direct DM search: testing DM interactions rate versus mass



- Improvements in technology have given impressive limits on properties of DM
- But nothing conclusive yet ... However DM can be around the corner !

# DM Observables: the power of WIMP

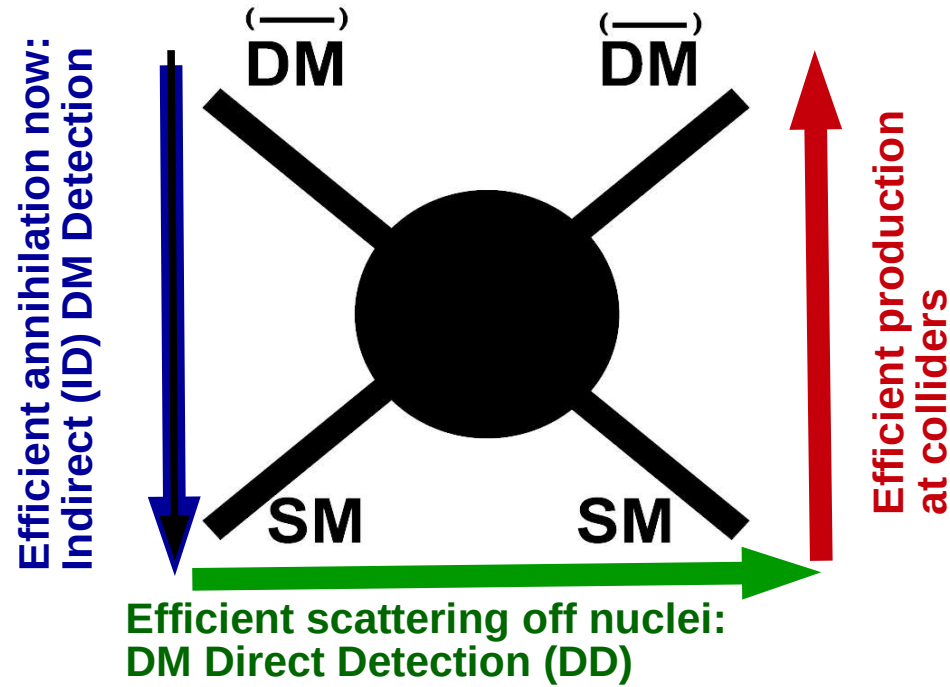
Correct Relic density: efficient (co) annihilation  
WMAP, Planck ; annihilation to photons can affect CMB

Signatures from neutralino annihilation in halo, core of the Earth and Sun

- photons,
- Anti-protons
- positrons,
- Neutrinos

Neutrino telescopes:

- Amanda
- Icecube
- Antares



LHC signatures

- mono-jet
- mono-photon
- mono-Z
- mono Higgs
- VBF+MET
- soft leptons+MET
- ....

**Note:** there is no 100% correlation between signatures above. For example, the high rate of annihilation does not always guarantee high rate for DD!

**Actually there is a great complementarity in this:**

- In case of NO DM Signal – we can efficiently exclude DM models
- In case of DM signal – we can efficiently determine the nature of DM



# Vector DM

- The abelian/non-abelian Vector DM with Higgs portal
  - $U(1)_D$  Group

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- $V_D^\mu \leftrightarrow -V_D^\mu$  Explicit  $Z_2$  symmetry plus a Higgs portal to provide the stability and the mass for VDM and connect it to the SM

$$\mathcal{L} \supset -\frac{1}{4}V_{\mu\nu}V^{\mu\nu} + (D_\mu\Phi)^\dagger (D^\mu\Phi) - V(\Phi) + \lambda_P |H|^2|\Phi|^2$$

with  $D_\mu\Phi \equiv \partial_\mu\Phi - gQ_\Phi V_\mu\Phi$

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so one has  $m_V^2 = g^2 Q_\Phi^2 v_\phi^2$



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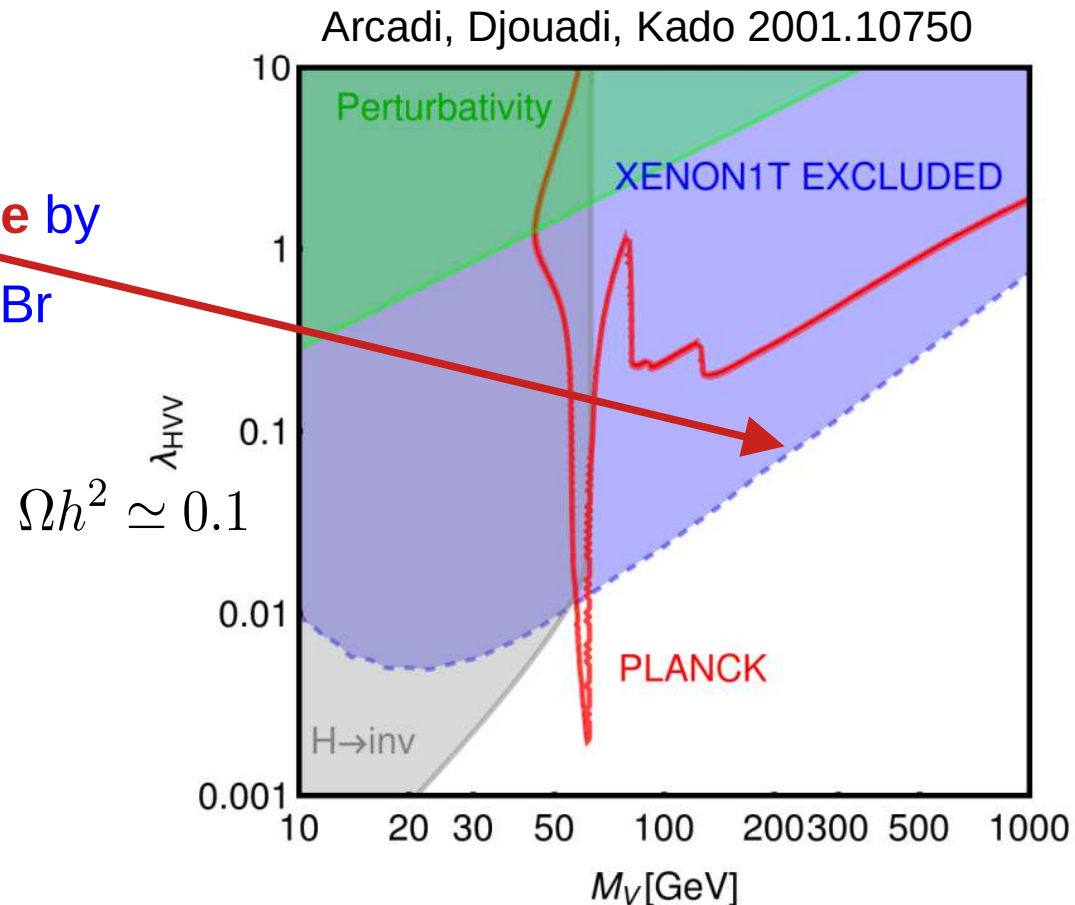
- Quite a few papers:

Lebedev, Lee, Mambrini 1111.4482,  
 Baek, Ko, Park, Senaha 1212.2131  
 DiFranzo, Fox, Tait 1512.06853

Farzan, Akbarieh 1207.4272  
 Duch, Grzadkowski, McGarrie 1506.08805  
 .....

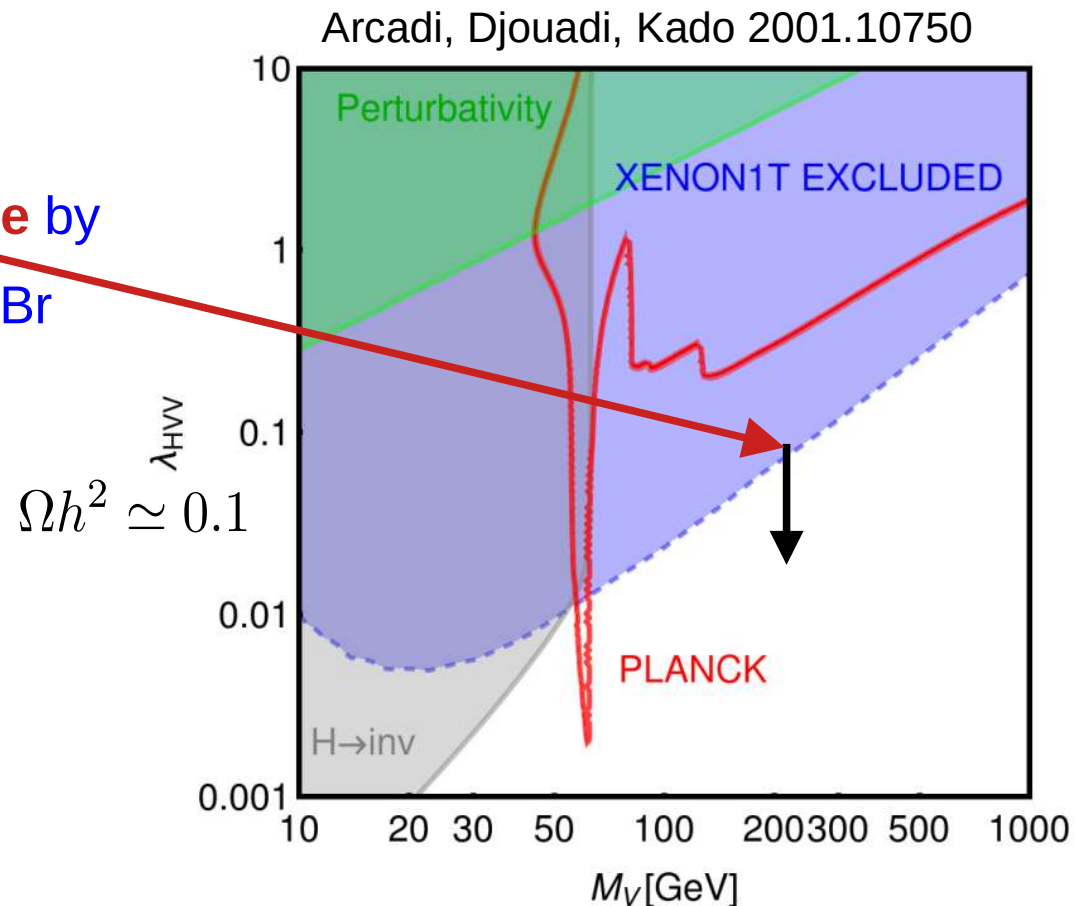
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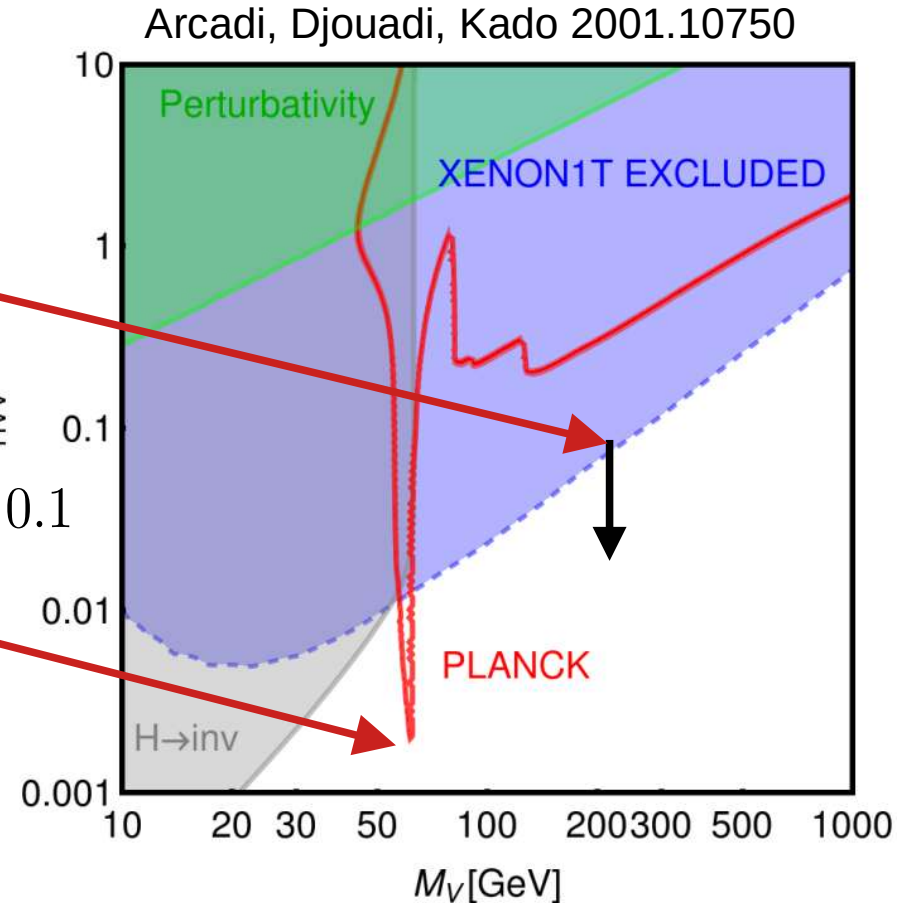
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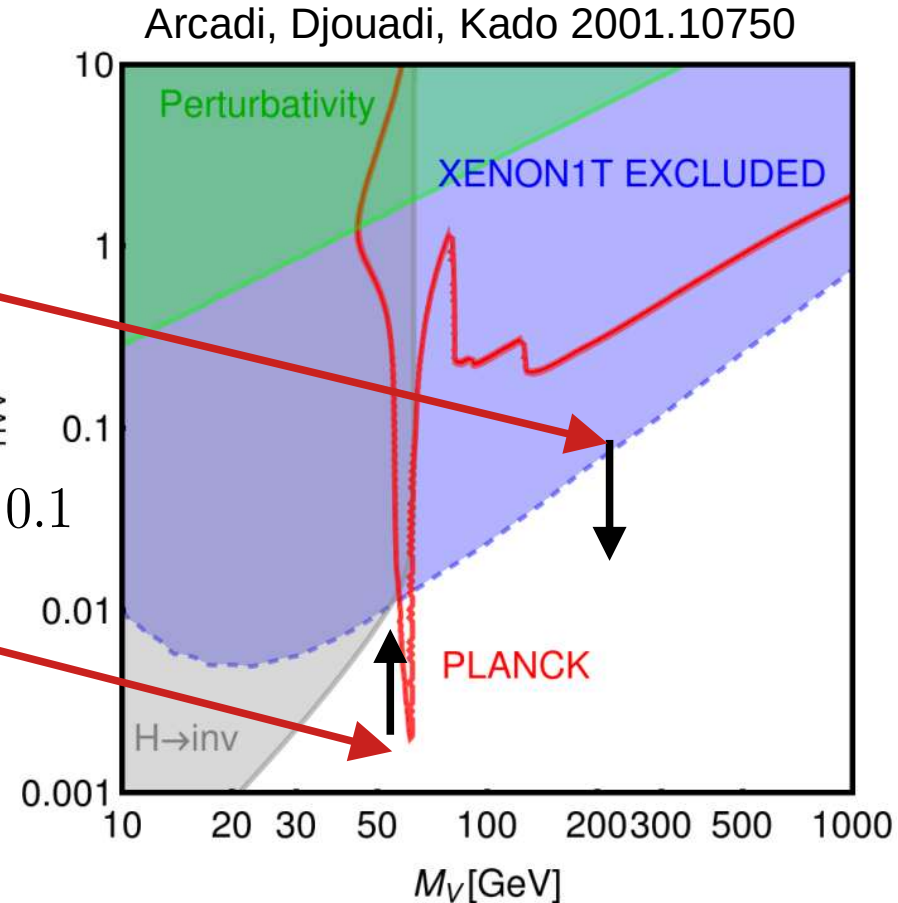
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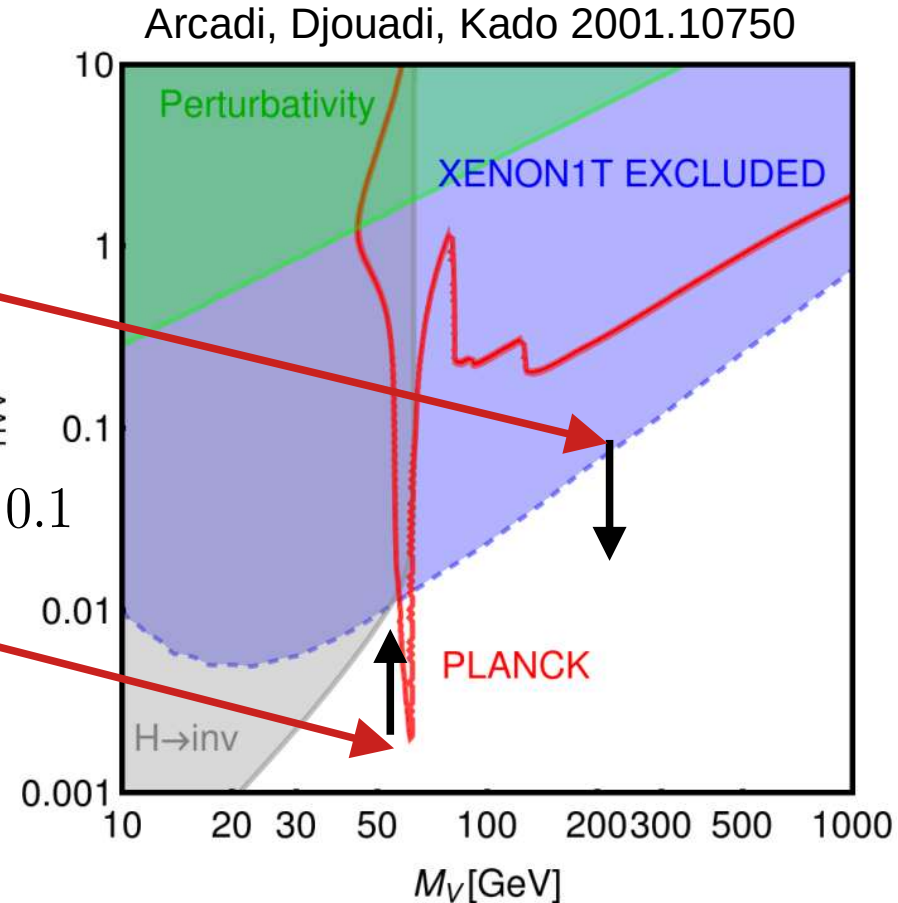
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- The Higgs portal VDM parameter space is very limited by interplay of collider, DD and DM relic density





# Vector DM with the Higgs portal

- Non-abelian case

- Generalisation to SU(N) case:

Gross, Lebedev, Mambrini 1505.07480

SSB with N-1 complex scalar N-plets in fundamental rep of SU(N) – gives mass to VDM and predicts  $(N-1)^2$  scalars

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- electroweakly interacting non-abelian vector dark matter:

Abea, Fujiwara, Hisano, Matsushita 2004.00884

$SU(2)_0 \times SU(2)_1 \times SU(2)_2 \times U(1)_Y$  :  $SU(2)_0 \leftrightarrow SU(2)_2$  symmetry provides stability for VDM, so there are VDM triplet + vector triplet of unstable W'/Z' bosons

$$\begin{aligned} V_{\text{scalar}} = & m^2 H^\dagger H + m_\Phi^2 \text{tr} \left( \Phi_1^\dagger \Phi_1 \right) + m_\Phi^2 \text{tr} \left( \Phi_2^\dagger \Phi_2 \right) \\ & + \lambda (H^\dagger H)^2 + \lambda_\Phi \left( \text{tr} \left( \Phi_1^\dagger \Phi_1 \right) \right)^2 + \lambda_\Phi \left( \text{tr} \left( \Phi_2^\dagger \Phi_2 \right) \right)^2 \\ & + \lambda_{h\Phi} H^\dagger H \text{tr} \left( \Phi_1^\dagger \Phi_1 \right) + \lambda_{h\Phi} H^\dagger H \text{tr} \left( \Phi_2^\dagger \Phi_2 \right) + \lambda_{12} \text{tr} \left( \Phi_1^\dagger \Phi_1 \right) \text{tr} \left( \Phi_2^\dagger \Phi_2 \right) \end{aligned}$$

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quite a non-minimal model



# Vector like (VL) fermion Portal for Vector DM

- Higgs portal is very-well studied and the parameter space for minimal scenarios is almost excluded
- **We are driven by curiosity and simplicity to find an alternative portal for Vector Dark Matter**

## SM + three ingredients:

- $SU(2)_D$  new (dark) non-abelian new gauge group  $V_\mu^D$
- Complex scalar doublet charged under  $SU(2)_D$   $\Phi_D$
- Vector-Like fermion doublet of  $SU(2)_D$   $\Psi$

# Vector like (VL) fermion Portal for Vector DM

- The general form of the Yukawa terms of the new fermion sector reads

$$-\mathcal{L}_f = M_\Psi \bar{\Psi} \Psi + (y' \bar{\Psi}_L \Phi_D f_R^{\text{SM}} + y'' \bar{\Psi}_L \Phi_D^c f_R^{\text{SM}} + h.c) ,$$

where  $\Phi_D^c = i\tau_2 \Phi^*$ , while  $y'$  and  $y''$  are new Yukawas, connecting SM fermions and new VL fermions

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- **Problem:** the presence of both  $y'$  and  $y''$  breaks the stability of gauge bosons, since it breaks global  $SU(2)$  in the dark sector
- If we assign the “dark charge” to the components of the doublets, e.g.  $Q_D = T_D^3 + Y_D$  and require its conservation, we will get
  - $SU(2)_D \times U(1)_{\text{glob}} \rightarrow U(1)_{\text{glob}}^d$  pattern of dark sector breaking
  - $\mathbb{Z}_2$  Subgroup can be defined as  $\mathbb{Z}_2 : (-1)^{Q_D}$
  - for  $\Phi_D$  we choose, e.g.  $Y_D = 1/2$ , then  $y''$  is eliminated, stabilizing VDM



# Vector like fermion Portal for Vector DM

- So, we have:  $SU(2)_D \times U(1)_{\text{glob}} \rightarrow U(1)_{\text{glob}}^d$ ,  $\mathbb{Z}_2 : (-1)^{Q_D}$ ,  $Q_D = T_D^3 + Y_D$

$Y_D = 1/2$  for the doublet and  $Y_D = 0$  for the triplet

- Different components of the doublet and triplet will have 0/1 dark charge, so +/-  $\mathbb{Z}_2$

- two scalar degrees of the doublet (i.e upper part of the doublet) are  $\mathbb{Z}_2$  - odd – they become longitudinal component of DM  
the lower part of scalar doublet is  $\mathbb{Z}_2$ -even, it contains vev

$$\Phi_D = \begin{pmatrix} \varphi_{D+1/2}^0 \\ \varphi_{D-1/2}^0 \end{pmatrix} \rightarrow \langle \Phi_D \rangle = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_D \end{pmatrix}$$

- this means that one of the components of the vector triplet is  $\mathbb{Z}_2$ -even

- the term, connecting dark scalar and VL fermion and SM RH fermion:

$$y' \bar{\Psi}_L \Phi_D f_R^{\text{SM}}$$

one component of VL fermion doublet is  $\mathbb{Z}_2$ -even and the other -  $\mathbb{Z}_2$ -odd

	$SU(2)_L$	$U(1)_Y$	$SU(2)_D$	$\mathbb{Z}_2/Q_D$
$\Phi_D = \begin{pmatrix} \varphi_{D+1/2}^0 \\ \varphi_{D-1/2}^0 \end{pmatrix}$	1	0	2	-/ +1 +/ 0
$\Psi = \begin{pmatrix} \psi_D \\ \psi \end{pmatrix}$	1	$Q_{EM}$	2	-/ +1 +/ 0
$V_\mu^D = \begin{pmatrix} V_{D+\mu}^0 \\ V_{D0\mu}^0 \\ V_{D-\mu}^0 \end{pmatrix}$	1	0	3	-/ +1 +/ 0 -/ -1

# Building VLF Portal for Vector DM: $V_{D+}^0 / V_{D-}^0$ . Dark Matter

$$SU(2)_D \quad V_\mu^D = \begin{pmatrix} V_{D+}^0 \\ V_{D0}^0 \\ V_{D-}^0 \end{pmatrix} \quad \Phi_D = \begin{pmatrix} \varphi_{D+\frac{1}{2}}^0 \\ \varphi_{D-\frac{1}{2}}^0 \end{pmatrix} \quad \Psi = \begin{pmatrix} \psi_D \\ \psi \end{pmatrix}$$

$$\mathbb{Z}_2 : \{+, -\}$$

The only\*  $\mathbb{Z}_2$ -odd neutral massive particles are the D-charged gauge bosons  $V_{D\pm}^0$

→ dark matter

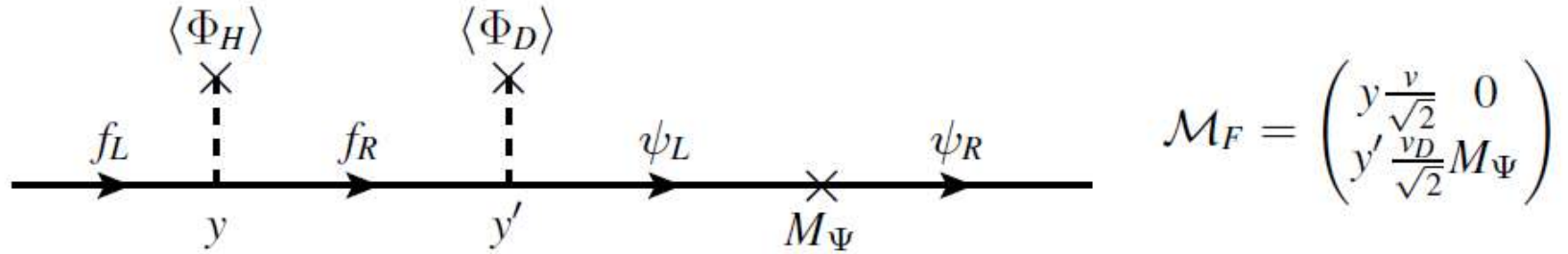
\* unless  $\Psi$  is a neutrino partner

$$SU(2)_L \times U(1)_Y \quad V_\mu = \begin{pmatrix} W^+ \\ W_3 \\ W^- \end{pmatrix}, B_\mu \quad \Phi_H = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix} \quad \begin{pmatrix} u \\ d \end{pmatrix}_L \quad \begin{pmatrix} \nu \\ e \end{pmatrix}_L \quad \begin{matrix} u_R \\ d_R \end{matrix} \quad \begin{matrix} e_R \\ \psi_D \end{matrix} \quad \psi$$

$$\begin{aligned} \mathcal{L} = & -\frac{1}{4}(W_{\mu\nu}^i)^2 - \frac{1}{4}(B_{\mu\nu})^2 + |D_\mu \Phi_H|^2 + \mu^2 \Phi_H^\dagger \Phi_H - \lambda(\Phi_H^\dagger \Phi_H)^2 + \bar{f}^{\text{SM}} i \not{\partial} f^{\text{SM}} - (y \bar{f}_L^{\text{SM}} \Phi_H f_R^{\text{SM}} + h.c.) \\ & - \frac{1}{4}(V_{\mu\nu}^{Di})^2 + |D_\mu \Phi_D|^2 + \mu_D^2 \Phi_D^\dagger \Phi_D - \lambda_D(\Phi_D^\dagger \Phi_D)^2 + \bar{\Psi} i \not{\partial} \Psi - M_\Psi \bar{\Psi} \Psi - (y' \bar{\Psi}_L \Phi_D f_R^{\text{SM}} + h.c.) \\ & - \lambda_{\Phi_H \Phi_D} \Phi_H^\dagger \Phi_H \Phi_D^\dagger \Phi_D \end{aligned}$$

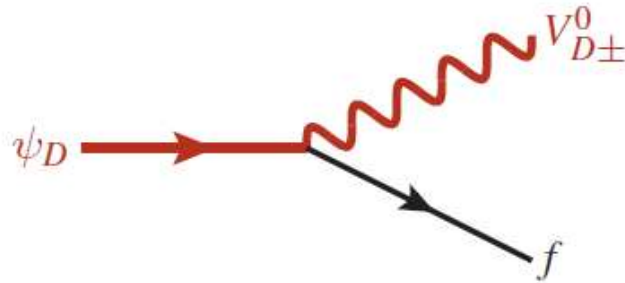
# VLF portal: $\mathbb{Z}_2$ -even fermions – RH SM ones and VL ones – mix

$$-\mathcal{L}_f = (y \bar{f}_L^{\text{SM}} \Phi_H f_R^{\text{SM}} + y' \bar{\Psi}_L \Phi_D f_R^{\text{SM}} + h.c.) + M_\Psi \bar{\Psi} \Psi \quad \text{with} \quad \Psi = \begin{pmatrix} \psi_D \\ \psi \end{pmatrix}$$



$\mathbb{Z}_2$ -odd  $\psi_D$  is DM-SM mediator

$\mathbb{Z}_2$ -even  $\psi$  mixes with SM

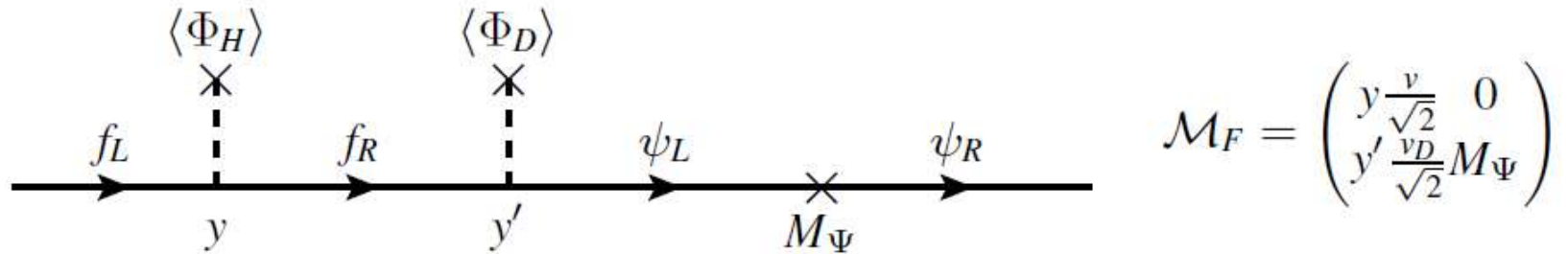


$$\begin{pmatrix} f^{\text{SM}} \\ \psi \end{pmatrix}_{L,R} = \begin{pmatrix} \cos \theta_{fL,R} & \sin \theta_{fL,R} \\ -\sin \theta_{fL,R} & \cos \theta_{fL,R} \end{pmatrix} \begin{pmatrix} f \\ F \end{pmatrix}_{L,R}$$

The hierarchy between mass eigenstates is always  $m_f < m_\psi \leq m_F$

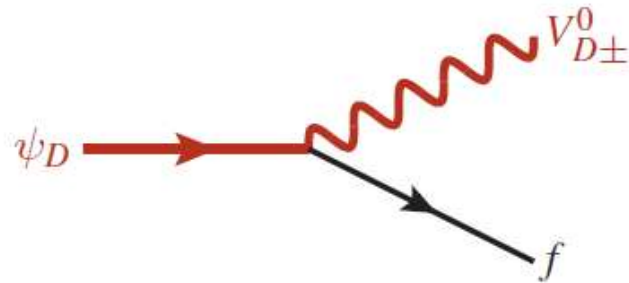
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Potential to introduce flavour structure(s) with VL fermions, including VL leptons to explain various flavour anomalies, including  $(g-2)_\mu$  !

# The gauge sector: $v' / v_D$ radiative mass split, no tree-level $V' - Z$ mixing

- At tree-level:  $m_{V_{D\pm}^0} = m_{V_{D0}^0} = \frac{g_D}{2} v_D$

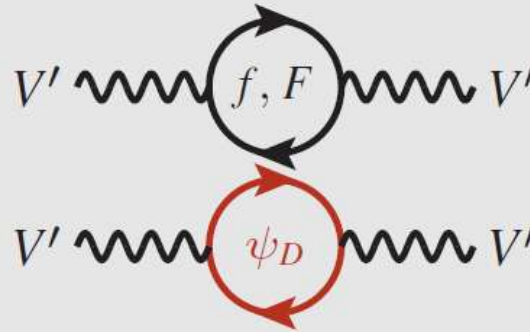


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- At loop-level:

Different loop corrections:  
 ( $V_{D\pm}^0 \equiv V_D$  and  $V_{D0}^0 \equiv V'$ )



Similar diagrams appear for  
 Kinetic mixing (backup slides)

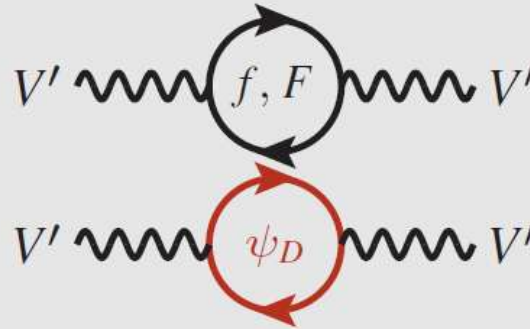
$$m_{V_D} - m_{V'} \simeq \frac{g_D^2}{32\pi^2} \frac{m_F^2 - m_{\psi_D}^2}{m_{V_D}} > 0 \quad \text{for } m_F \gg m_f, m_{V_D}$$

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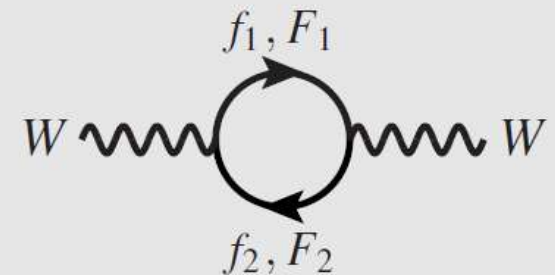
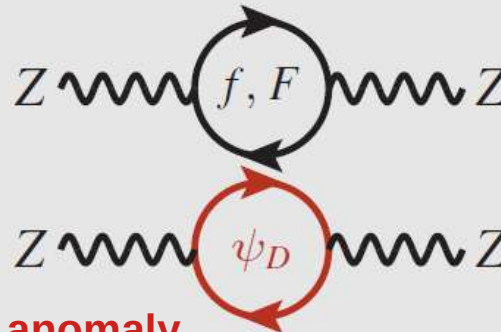


Similar diagrams appear for  
 Kinetic mixing (backup slides)

$$m_{V_D} - m_{V'} \simeq \frac{g_D^2}{32\pi^2} \frac{m_F^2 - m_{\psi_D}^2}{m_{V_D}} > 0 \quad \text{for } m_F \gg m_f, m_{V_D}$$

- Effect for W/Z boson masses

Modifications to SM  
 different for Z and W



Potential to explain W-boson mass anomaly

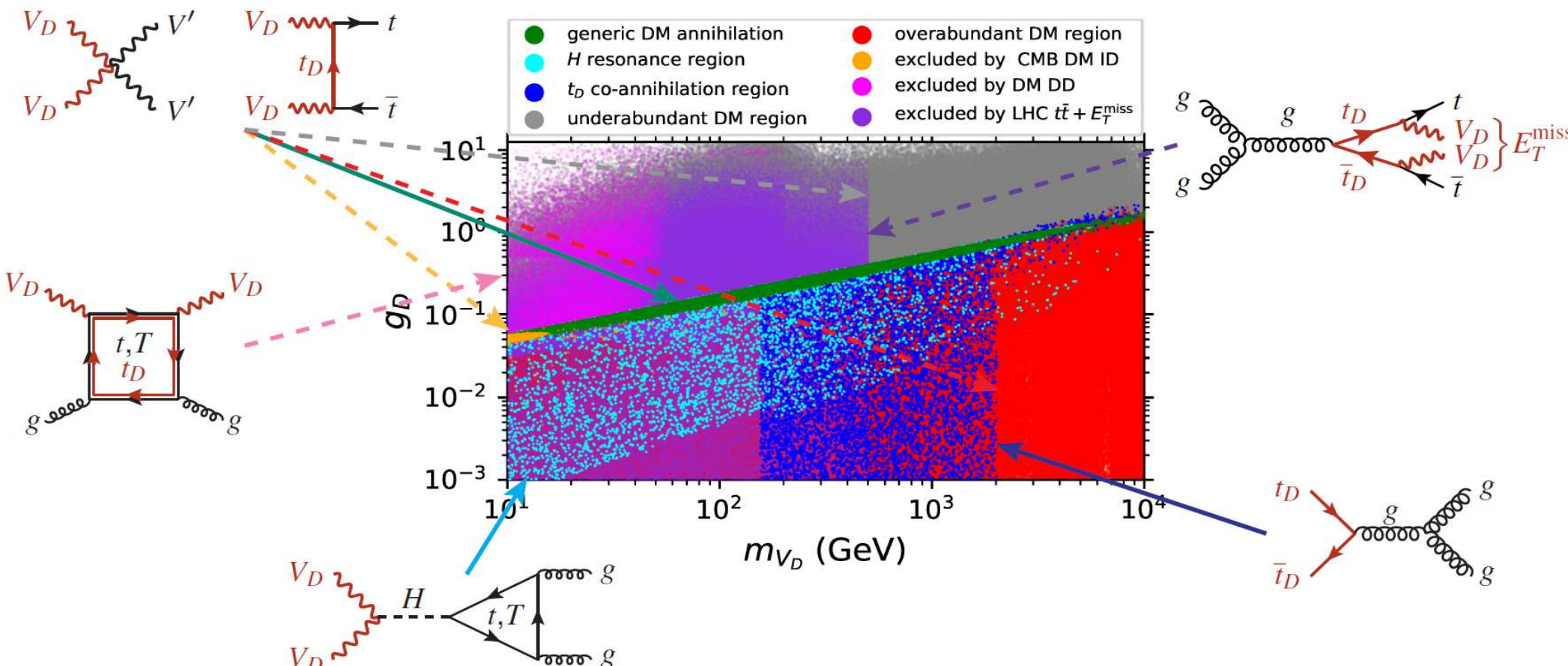
# Minimal VL top portal VDM: VL top portal without higgs portal mixing

The VL fermion is composed of top partners and there is no mixing between scalars

$$\Psi = \begin{pmatrix} t_D \\ T \end{pmatrix} \quad \text{with} \quad m_t < m_{t_D} \leq m_T$$

$$\sin \theta_S = 0$$

5D parameter space:  $g_D, m_{V_D}, m_H, m_T, m_{t_D}$





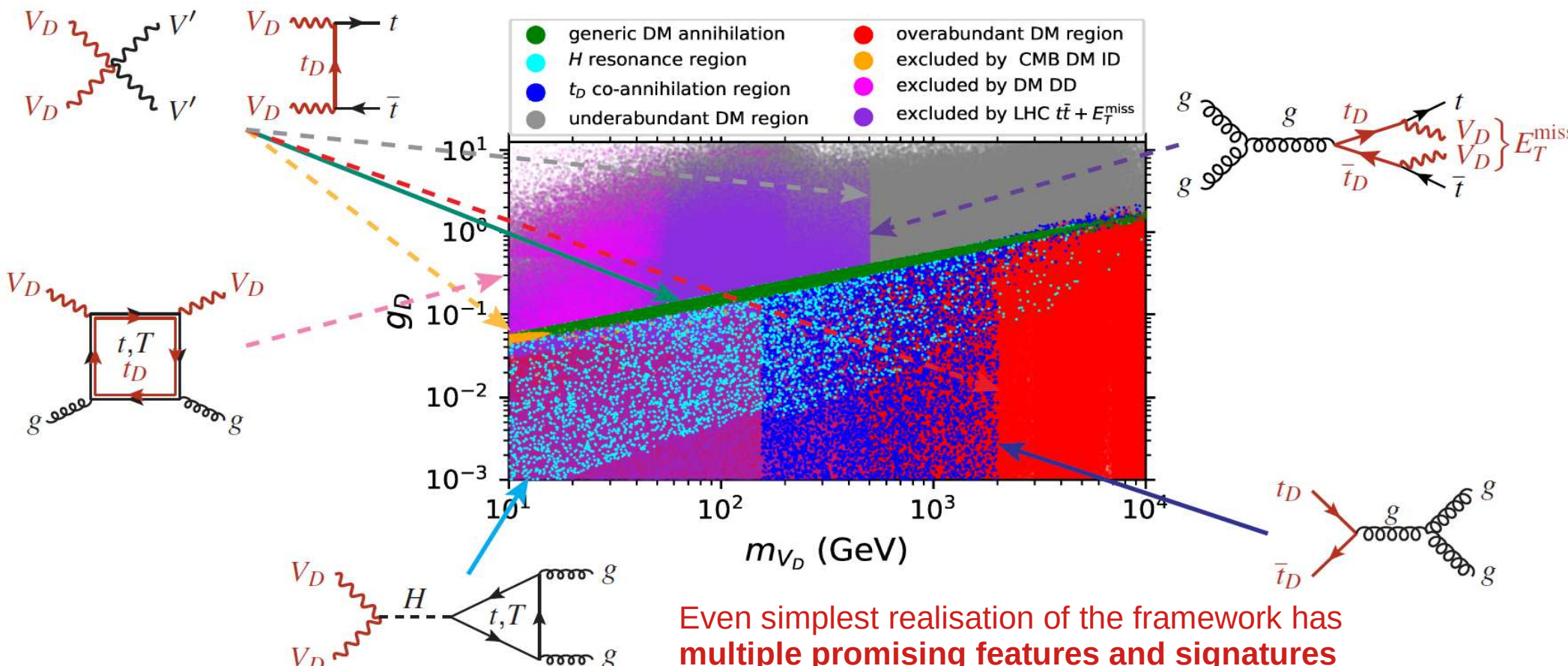
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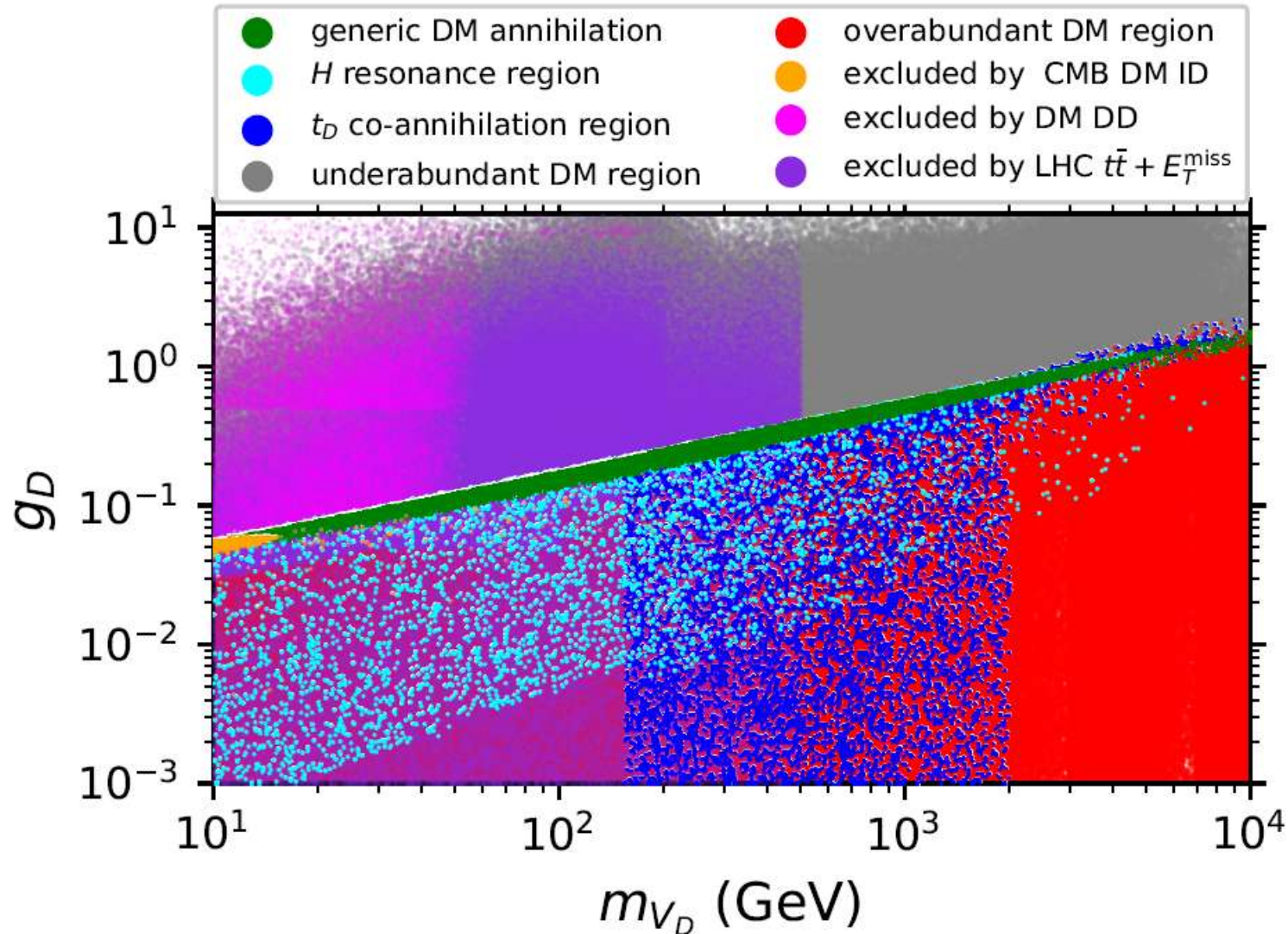
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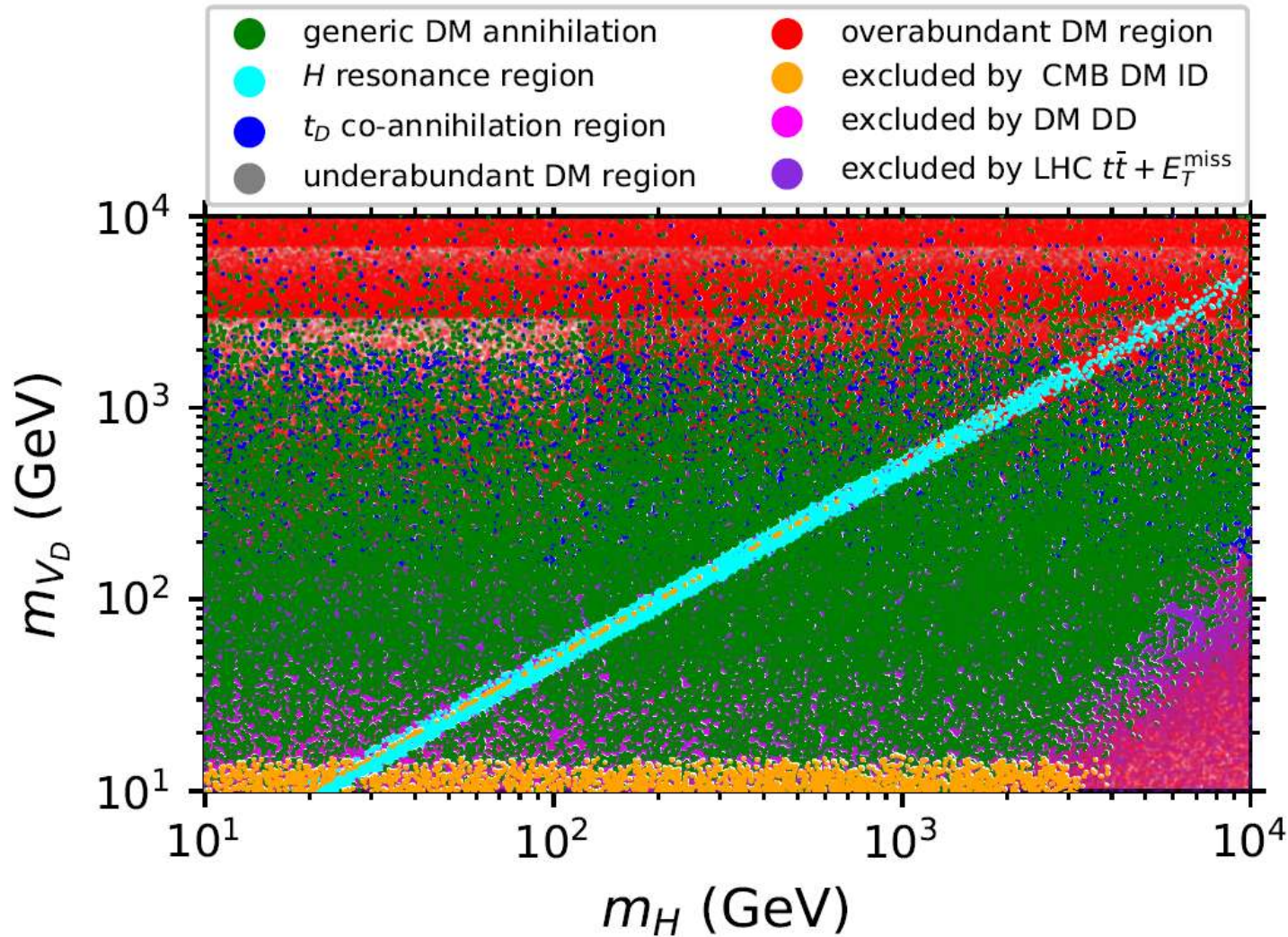


# Minimal VL top portal VDM: projections of 5D scan in $g_D, m_{V_D}, m_H, m_T, m_{t_D}$

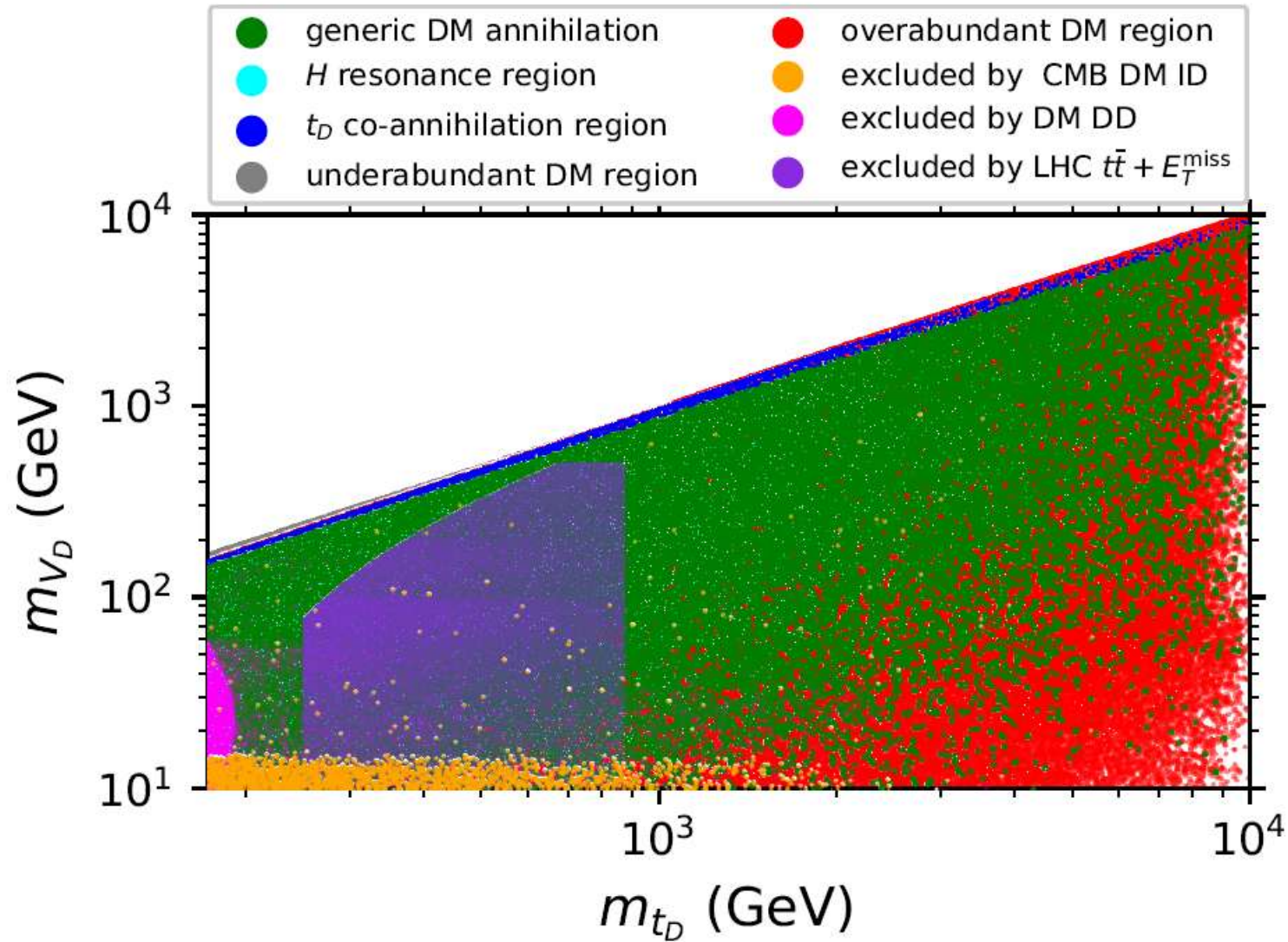




# Minimal VL top portal VDM: projections of 5D scan in $g_D, m_{V_D}, m_H, m_T, m_{t_D}$



# Minimal VL top portal VDM: projections of 5D scan in $g_D, m_{V_D}, m_H, m_T, m_{t_D}$



# Minimal VL top portal VDM: collider signatures

Process	Representative diagrams
mono-jet (only loop)	$E_T^{\text{miss}} + \text{jet from ISR or from loop}$
$t\bar{t} + E_T^{\text{miss}}$	$E_T^{\text{miss}}$
$t\bar{t}t\bar{t}$	
$hV'$ and $V'V'$ (only loop)	



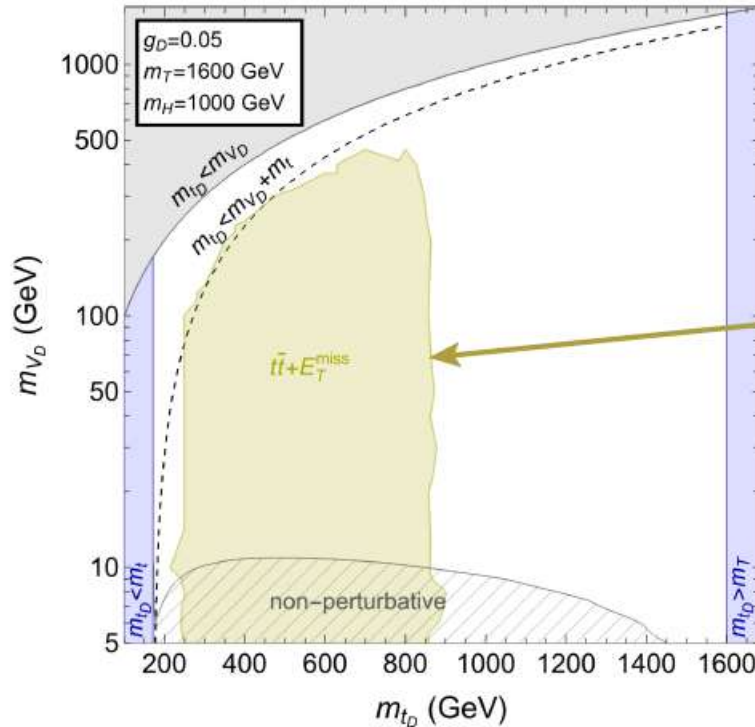
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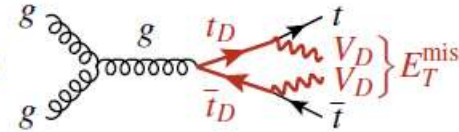
$$\sin \theta_S = 0$$

Representative benchmarks:  $\left\{ \begin{array}{l} g_D = 0.05, 0.5 \\ m_T = 1600 \text{ GeV} \\ m_H = 1000 \text{ GeV} \end{array} \right\}$  strong or weak cosmological constraints  
heavy enough to evade LHC constraints



Mediator mass bounded from below and above  
Light DM in non-perturbative region

LHC constrains  $m_{t_D}$  for  $m_{t_D} - m_{V_D} \gtrsim m_t$   
(bounds almost independent on  $g_D, m_T$  and  $m_H$ )



Recast

A. M. Sirunyan *et al.* [CMS], Search for top squarks and dark matter particles in opposite-charge dilepton final states at  $\sqrt{s} = 13$  TeV, Phys. Rev. D 97 (2018) no.3, 032009, arXiv:1711.00752 [hep-ex]

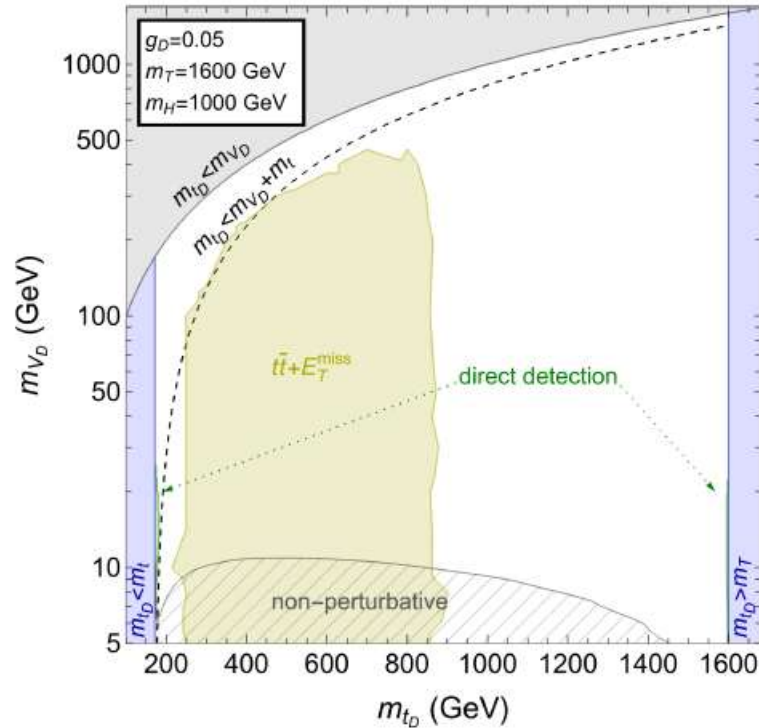
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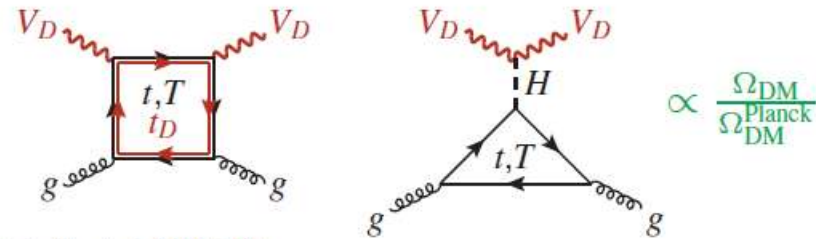
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Very weak direct detection constraints  
(mostly for  $m_{t_D} \sim m_t$  or  $m_{t_D} \sim m_T$  and light DM)



E. Aprile *et al.* [XENON],  
Dark Matter Search Results from a One Ton-Year Exposure of XENON1T,  
Phys. Rev. Lett. 121 (2018) no.11, 111302, arXiv:1805.12562 [astro-ph.CO]

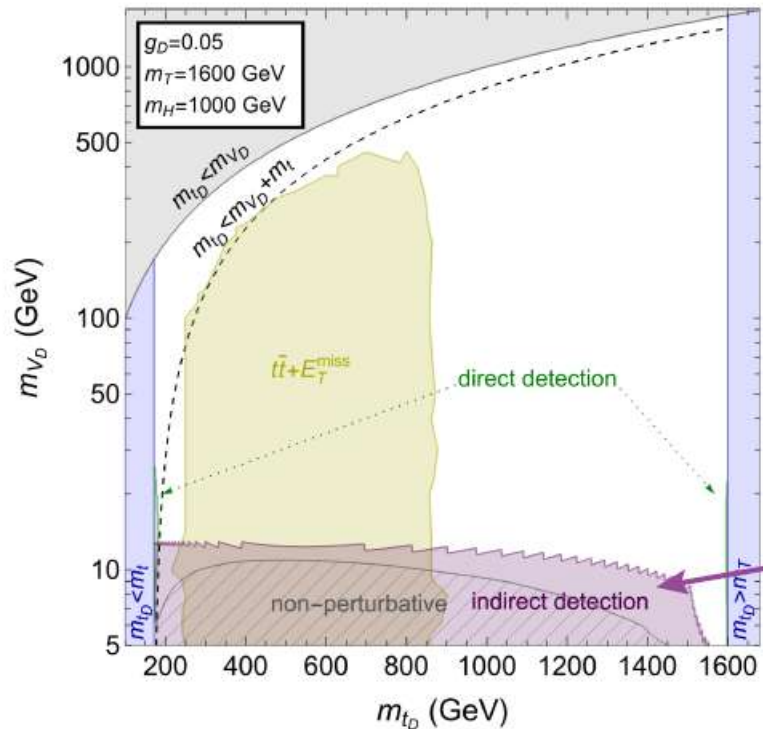
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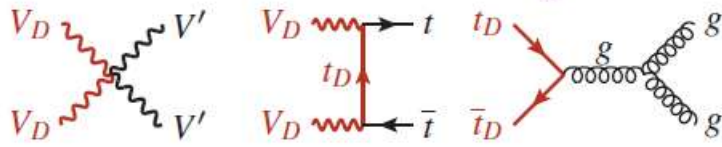


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Indirect detection constrains light DM



$$\propto \left( \frac{\Omega_{DM}}{\Omega_{Planck}^{DM}} \right)^2$$

N. Aghanim *et al.* [Planck],  
Planck 2018 results. VI. Cosmological parameters,  
Astron. Astrophys. 641 (2020), A6, arXiv:1807.06209 [astro-ph.CO]



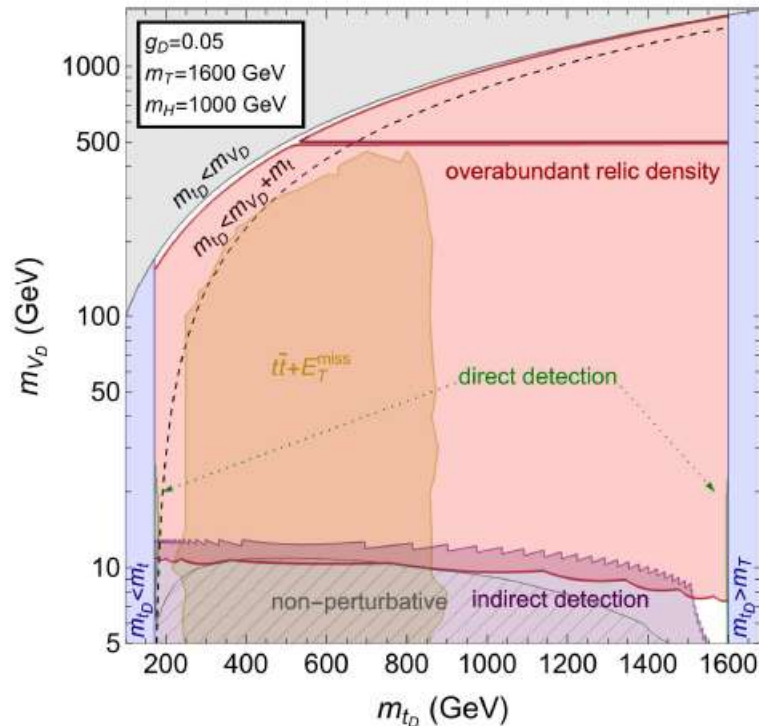
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Strong constrain from relic density

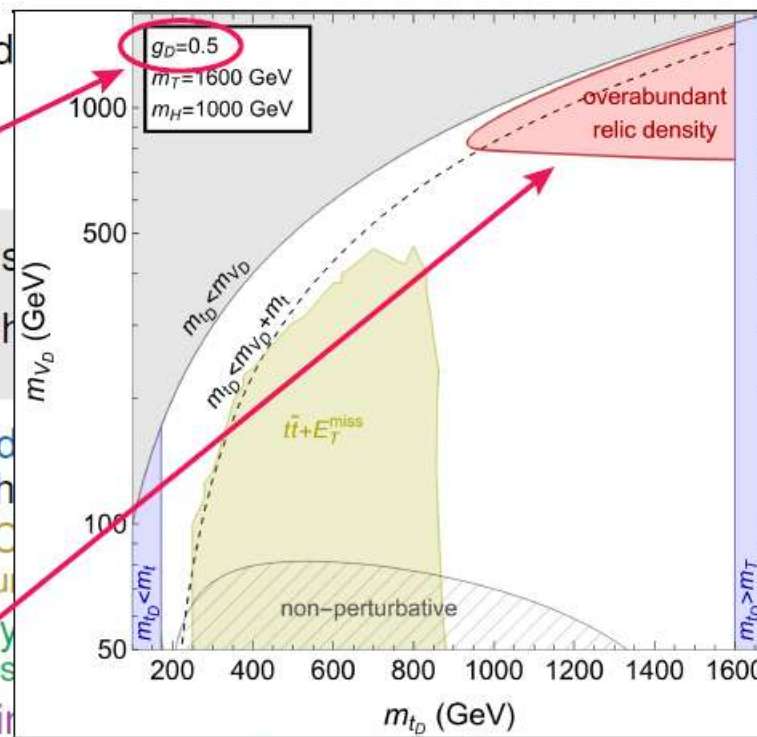
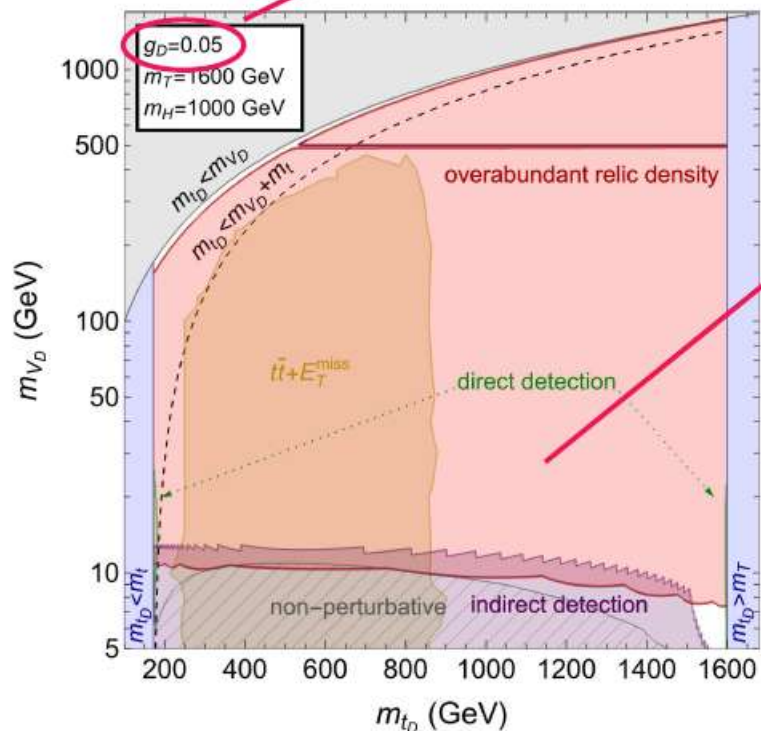
→ the model “lives” on the red contours ( $\Omega_{\text{DM}}^{\text{Planck}}$ )

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The VL fermion is composed of top partners and

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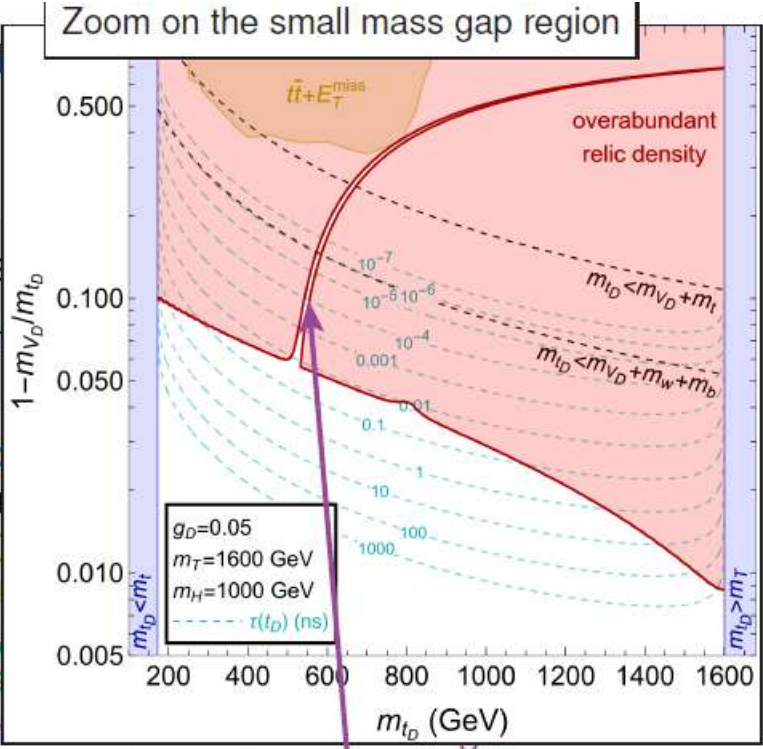
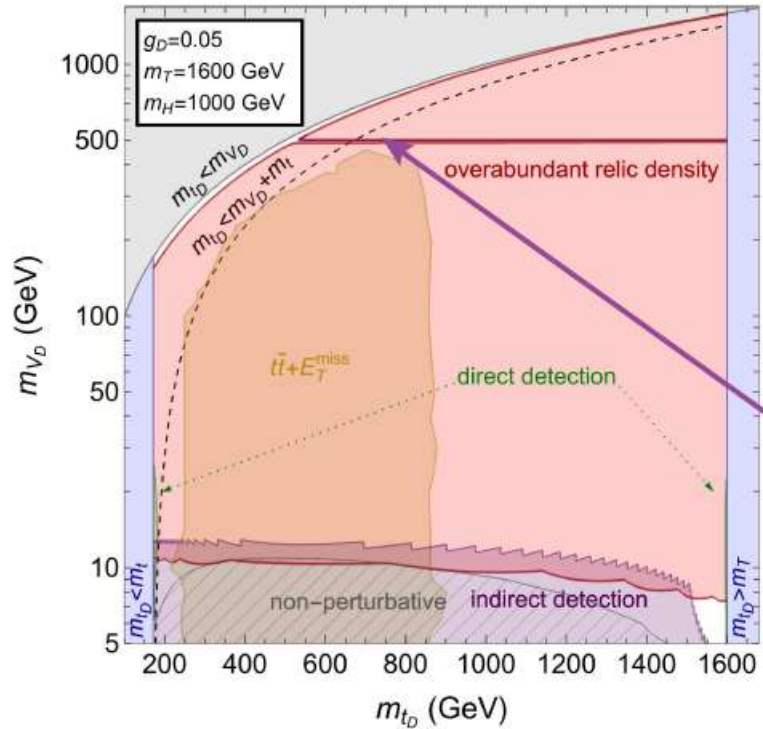
- Strong constrain from relic density
- the model “lives” on the red contours ( $\Omega_{DM}^{Planck}$ )
  - overabundant region shrinks for larger  $g_D$
  - and ID constraints vanish

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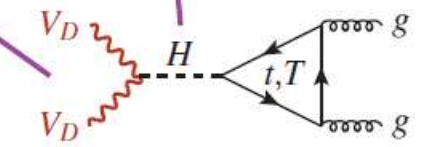
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- effective (co-)annihilation processes
- on the  $H_D$  pole, exclusion from ID





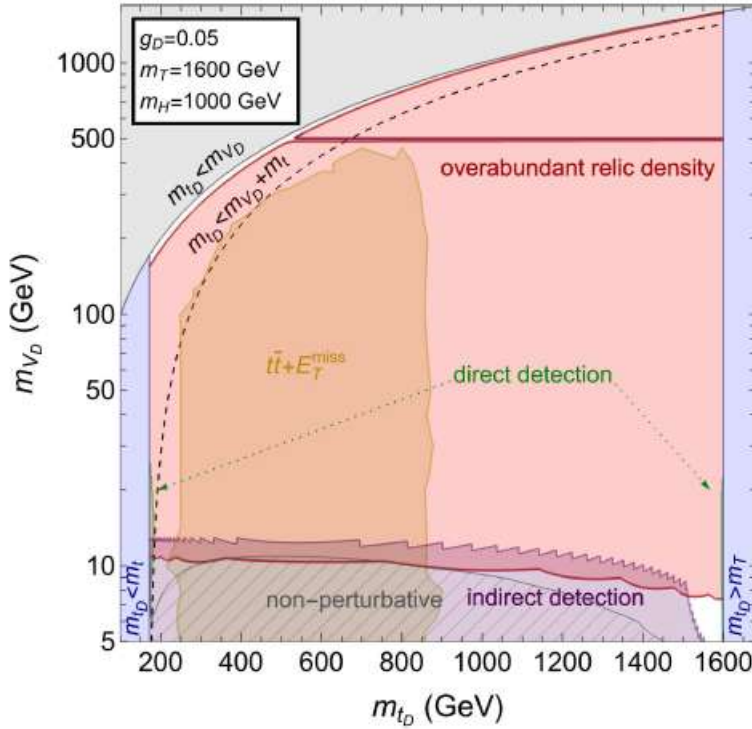
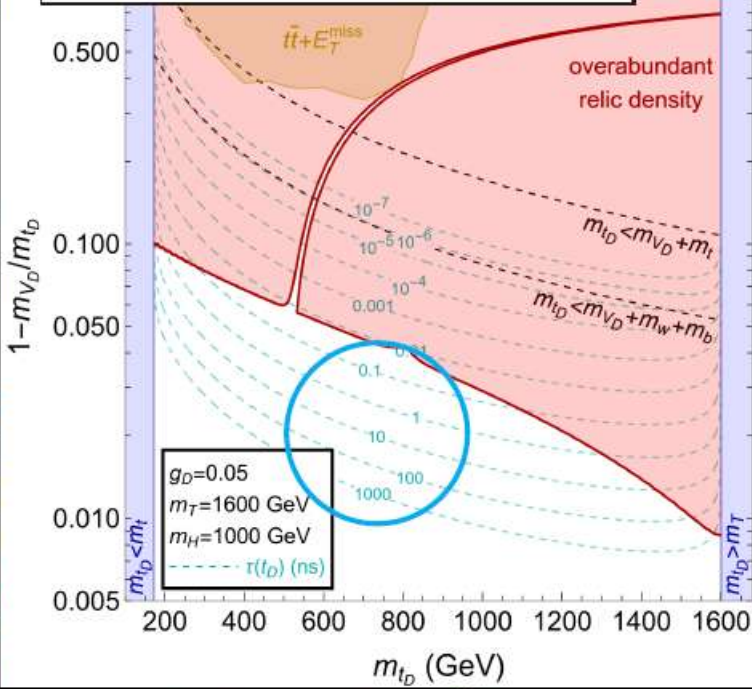
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Zoom on the small mass gap region



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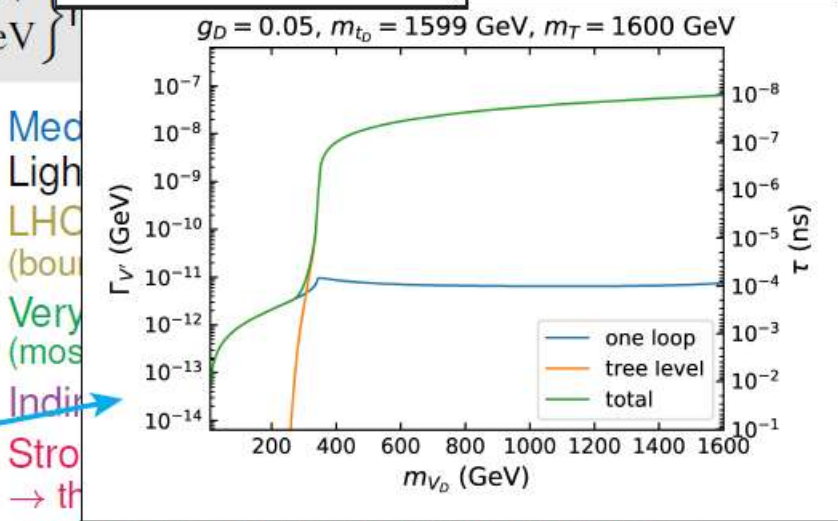
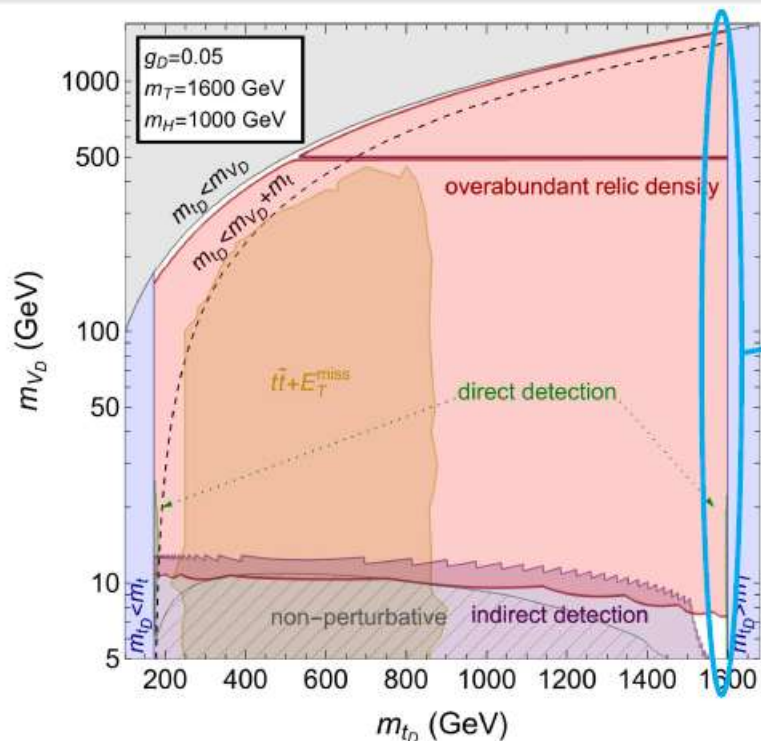
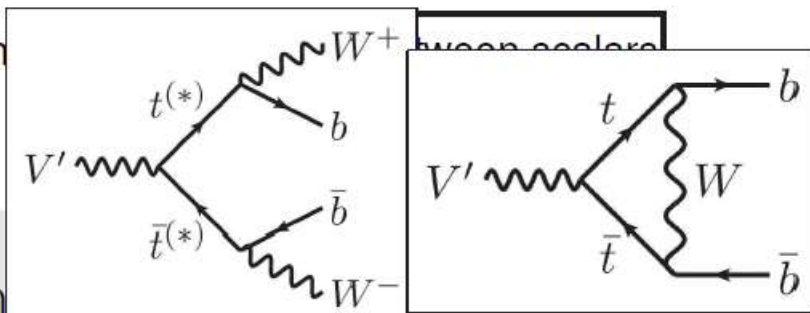
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The mediator  $t_D$  can be long lived, and  $V'$  too

# Summary on Fermion Portal Vector Dark Matter (FPVDM)

- FPVDM is a new framework which does not require the Higgs portal
- Has new features with new collider and cosmological implications
- Simple realisation of the top sector – several promising predictions and signatures
  - great potential to explain dark matter
  - collider signatures:  $t\bar{t} + \text{miss}$ ,  $Z'$ ,  $Z'H$ , long-lived  $Z'$

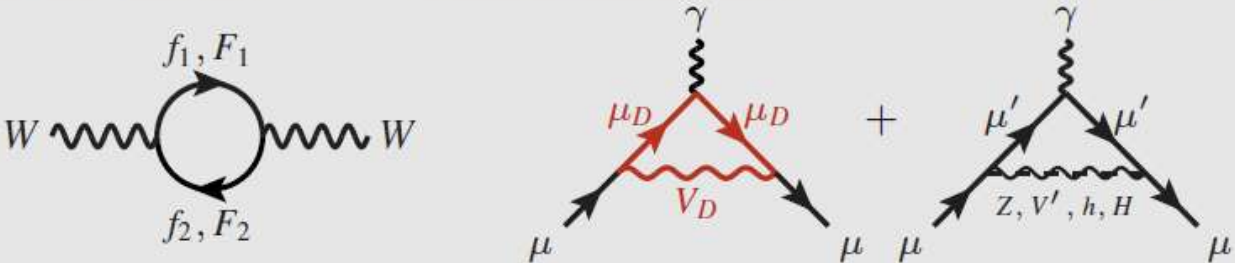


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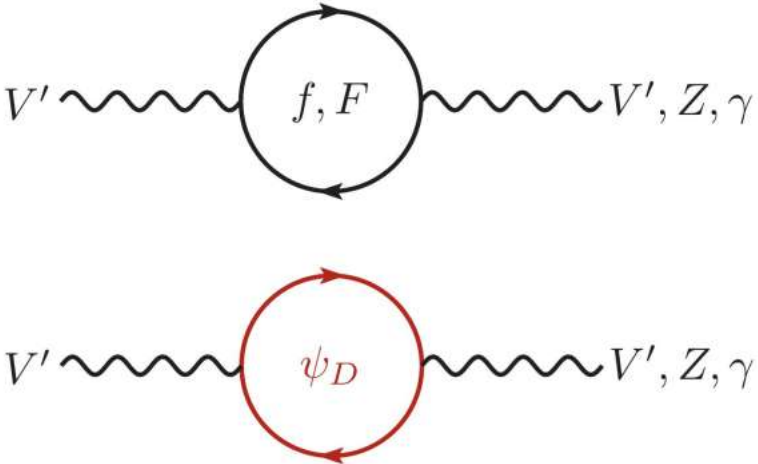
# Outlook

→ Different realizations to study current anomalies (LFU,  $(g - 2)_\mu$ ,  $m_W \dots$ )

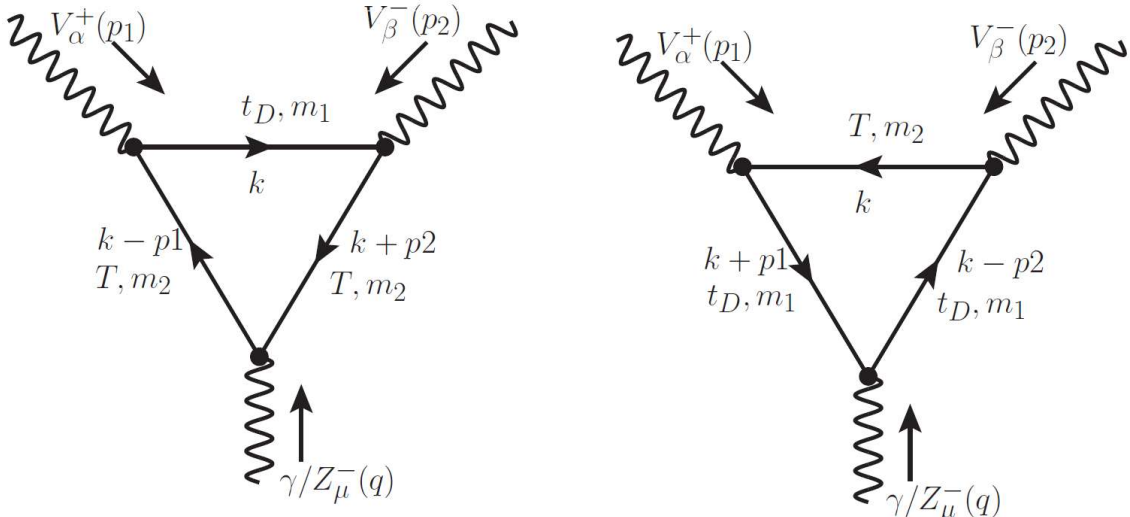


- Study of different theoretical embeddings
- Further analysis of cosmological implications and scenarios for future colliders

## V' Kinetic mixing



## DM interaction with Z/gamma at loop level



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# Thank you!

# Backup slides

# Gauging the global $U(1)$

A dark electroweak sector

Extend the dark sector with a  $U(1)_{YD}$  (dark hypercharge). Same scalars  $\Phi_H$  and  $\Phi_D$ .

$$\mathcal{G} = \mathcal{G}_{\text{SM}} \times \mathcal{G}_D = SU(2)_L \times U(1)_Y \times SU(2)_D \times U(1)_{YD} \longrightarrow U(1)_{\text{EM}} \times U(1)_D$$

**Conserved charge** from the unbroken  $U(1)_D$  symmetry:  $Q_D = T_{3D} + Y_D$

One assumption: SM fields do not carry  $Q_D$  charge

The only  $Q_D$ -charged state is  $V_{D\pm}^0 \equiv W_D \longrightarrow$  stable  $\longrightarrow$  **DM candidate**

Renormalizable, gauge-invariant kinetic mixing between  $U(1)_Y$  and  $U(1)_{YD}$  can be generated

$$-\mathcal{L}_{\text{KM}} = \frac{1}{4}B_{\mu\nu}B^{\mu\nu} + \frac{1}{4}B_{D\mu\nu}B_D^{\mu\nu} + \frac{\varepsilon}{2}B_{\mu\nu}B_D^{\mu\nu} \quad \begin{pmatrix} B^\mu \\ B_{D0}^{0\mu} \end{pmatrix} = \begin{pmatrix} \frac{1}{\sqrt{1-\varepsilon^2}} & 0 \\ -\frac{\varepsilon^2}{\sqrt{1-\varepsilon^2}} & 1 \end{pmatrix} \begin{pmatrix} \cos \theta_k & -\sin \theta_k \\ \sin \theta_k & \cos \theta_k \end{pmatrix} \begin{pmatrix} B_1^\mu \\ B_2^\mu \end{pmatrix}$$

Mixing between all  $Q$ - and  $Q_D$ -neutral bosons

$$\begin{cases} m_\gamma = 0 \\ m_{\gamma_D} = 0 \end{cases} \quad \begin{cases} m_Z^2 = \frac{v^2}{4} \left[ g^2 + g'^2 \left( 1 + \frac{(g^2 + g'^2)v^2 - g_D^2 v_D^2}{(g^2 + g'^2)v^2 - (g_D^2 + g_D'^2)v_D^2} \varepsilon^2 \right) \right] + \mathcal{O}(\varepsilon^4) \\ m_{Z'}^2 = \frac{v_D^2}{4} \left[ g_D^2 + g_D'^2 \left( 1 + \frac{g^2 v^2 - (g_D^2 + g_D'^2)v_D^2}{(g^2 + g'^2)v^2 - (g_D^2 + g_D'^2)v_D^2} \varepsilon^2 \right) \right] + \mathcal{O}(\varepsilon^4) \end{cases}$$

**2 massless and 2 massive vectors**

**Connections with dark-photon phenomenology**



# The scalar sector: when the higgs portal is absent, the interactions become minimal

EW + Dark symmetry breaking  $\longrightarrow$

$\begin{cases} v = \pm \sqrt{\frac{4\lambda_D\mu^2 - 2\lambda_{\Phi_H\Phi_D}\mu_D^2}{4\lambda\lambda_D - \lambda_{\Phi_H\Phi_D}^2}} \\ v_D = \pm \sqrt{\frac{4\lambda\mu_D^2 - 2\lambda_{\Phi_H\Phi_D}\mu^2}{4\lambda\lambda_D - \lambda_{\Phi_H\Phi_D}^2}} \end{cases}$	<p style="text-align: center;">Including Higgs portal</p>	$\begin{cases} v = \pm \sqrt{\frac{\mu^2}{\lambda}} \\ v_D = \pm \sqrt{\frac{\mu_D^2}{\lambda_D}} \end{cases}$
	<p style="text-align: center;">Without Higgs portal</p>	

8 degrees of freedom, 6 massive gauge bosons, 2 physical scalars  $h, H$

$$\mathcal{M}_S = \begin{pmatrix} \lambda v^2 & \frac{\lambda_{\Phi_H\Phi_D}}{2} v v_D \\ \frac{\lambda_{\Phi_H\Phi_D}}{2} v v_D & \lambda_D v_D^2 \end{pmatrix} \quad \sin \theta_S = \sqrt{2 \frac{m_H^2 v^2 \lambda - m_h^2 v_D^2 \lambda_D}{m_H^4 - m_h^4}}$$

$$m_{h,H}^2 = \lambda v^2 + \lambda_D v_D^2 \mp \sqrt{(\lambda v^2 - \lambda_D v_D^2)^2 + \lambda_{\Phi_H\Phi_D}^2 v^2 v_D^2}$$

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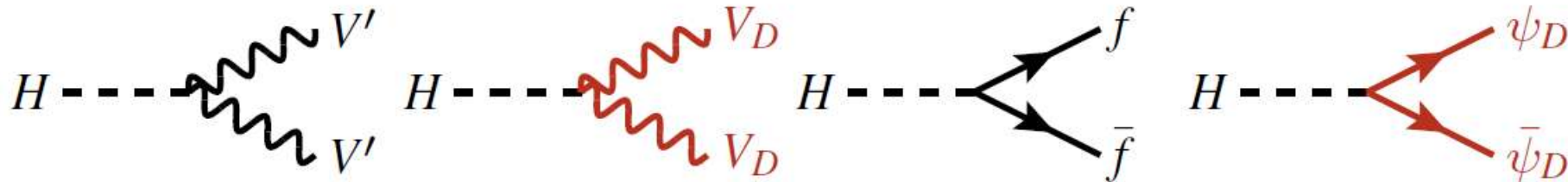
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If no Higgs portal, the interactions of the new scalar  $H$  are limited to:



# VL portal VDM: the summary of particle content

Vectors				
	$SU(2)_L$	$U(1)_Y$	$SU(2)_D$	$\mathbb{Z}_2$
$W_\mu = \begin{pmatrix} W_\mu^+ \\ W_\mu^3 \\ W_\mu^- \end{pmatrix}$	3	0	1	+
$B_\mu$	1	0	1	+
$V_\mu^D = \begin{pmatrix} V_{D+\mu}^0 \\ V_{D0\mu}^0 \\ V_{D-\mu}^0 \end{pmatrix}$	1	0	3	-

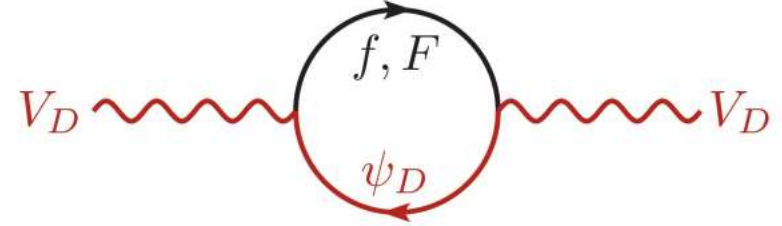
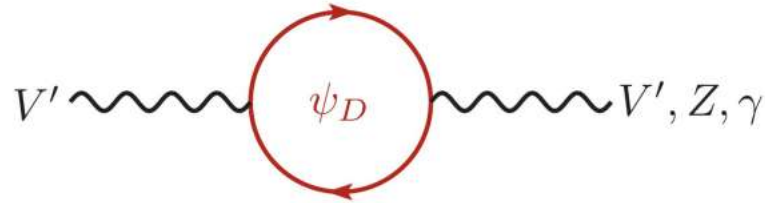
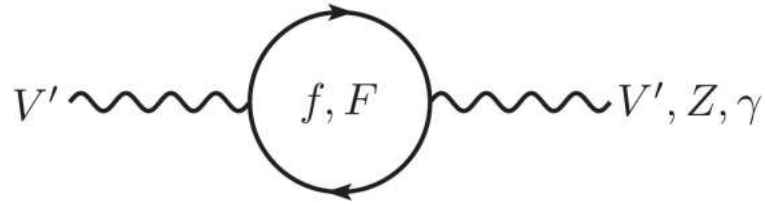
  

Scalars				
	$SU(2)_L$	$U(1)_Y$	$SU(2)_D$	$\mathbb{Z}_2$
$\Phi_H = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix}$	2	1/2	1	+
$\Phi_D = \begin{pmatrix} \varphi_{D+\frac{1}{2}}^0 \\ \varphi_{D-\frac{1}{2}}^0 \end{pmatrix}$	1	0	2	-

Fermions				
	$SU(2)_L$	$U(1)_Y$	$SU(2)_D$	$\mathbb{Z}_2$
$f_L^{\text{SM}} = \begin{pmatrix} f_{u,\nu}^{\text{SM}} \\ f_{d,\ell}^{\text{SM}} \end{pmatrix}_L$	2	$\frac{1}{6}, -\frac{1}{2}$	1	+
$u_R^{\text{SM}}, \nu_R^{\text{SM}}$	1	$\frac{2}{3}, 0$	1	+
$d_R^{\text{SM}}, \ell_R^{\text{SM}}$	1	$-\frac{1}{3}, -1$	1	+
$\Psi = \begin{pmatrix} \psi^D \\ \psi \end{pmatrix}$	1	$Q$	2	-

# Kinetic Mixing in FPVDM models

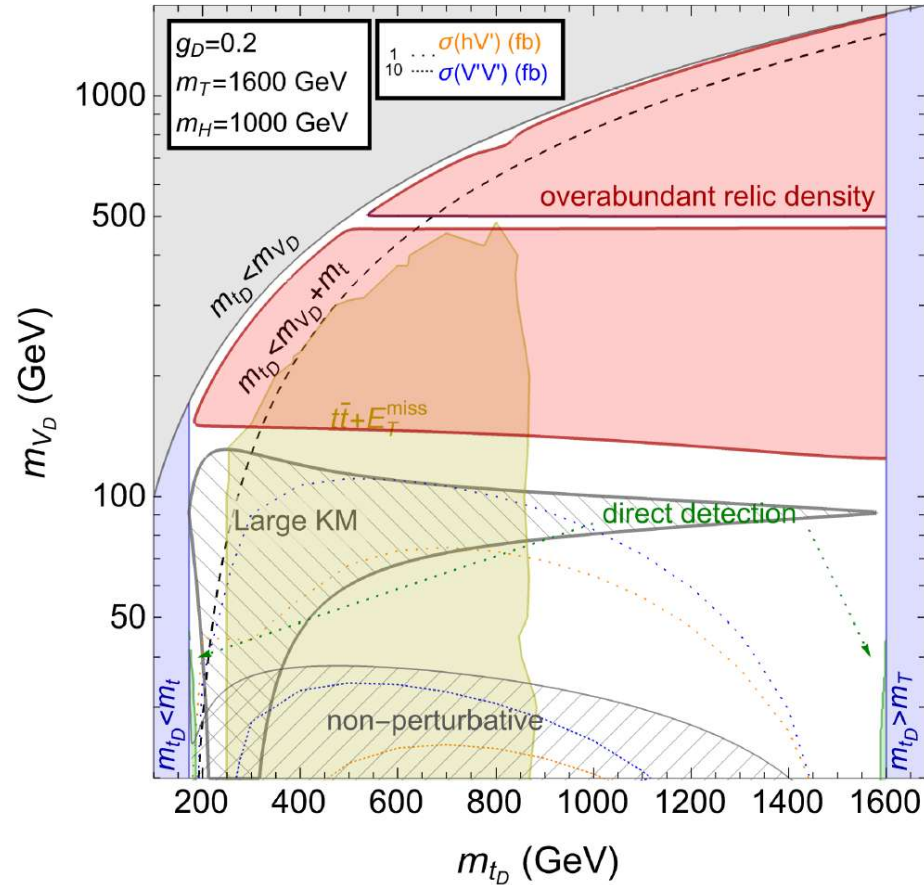
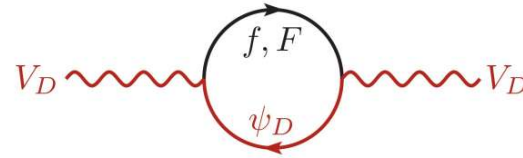
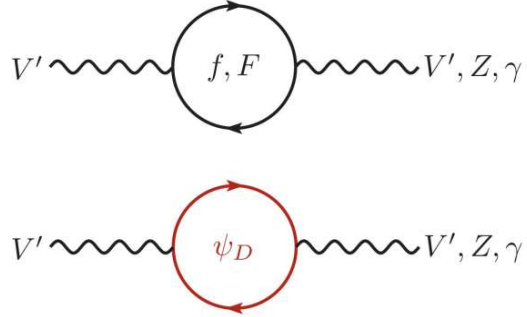


$$V^{\text{KM}} = \begin{pmatrix} 1 & 0 & -\frac{\epsilon_{AV}}{\sqrt{1-\epsilon_{AV}^2-\epsilon_{ZV}^2}} \\ 0 & 1 & -\frac{\epsilon_{ZV}}{\sqrt{1-\epsilon_{AV}^2-\epsilon_{ZV}^2}} \\ 0 & 0 & \frac{1}{\sqrt{1-\epsilon_{AV}^2-\epsilon_{ZV}^2}} \end{pmatrix}$$

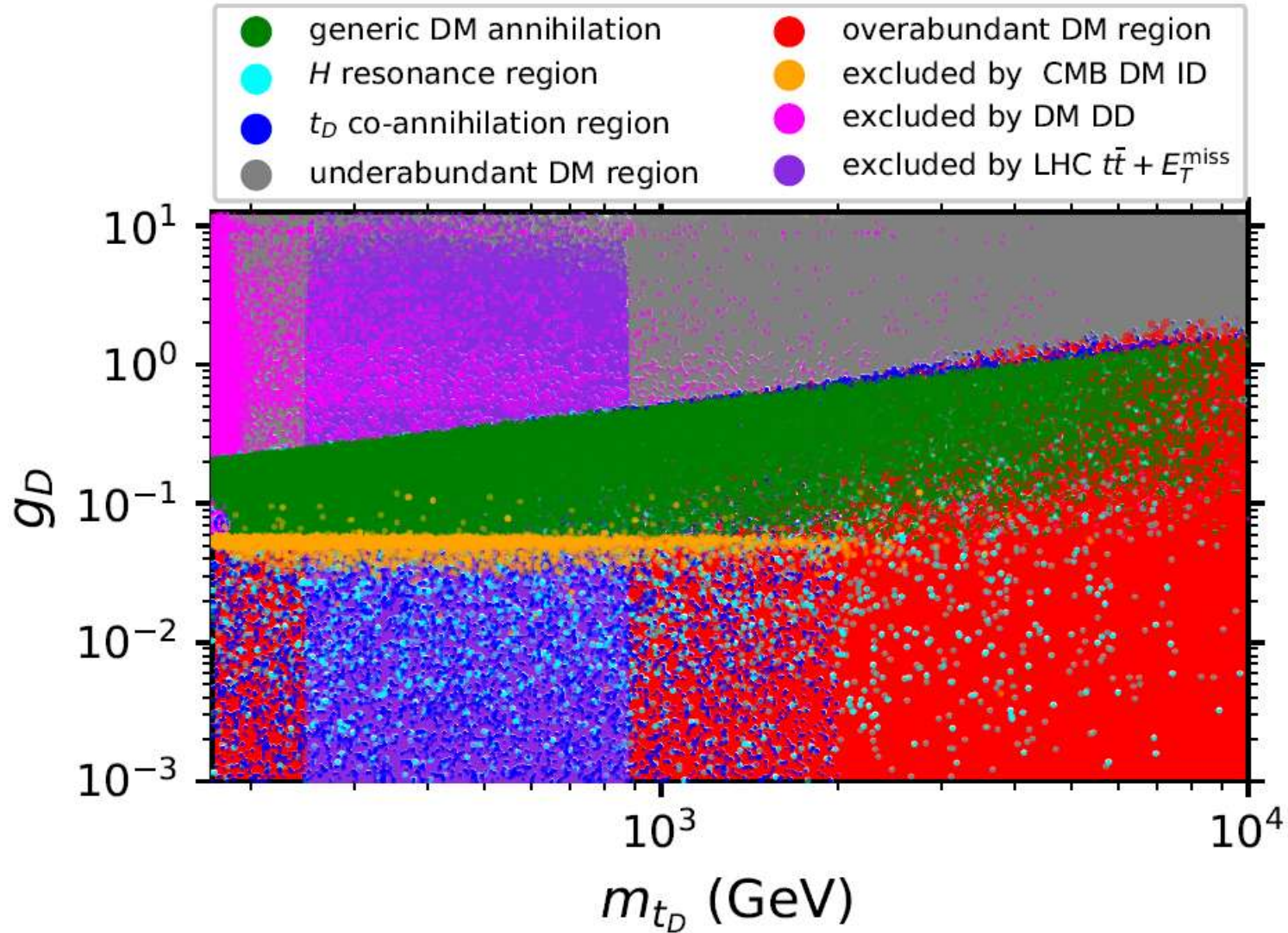
$$\epsilon_{ZV} = \frac{gg_D}{16\pi^2 c_w} \left( \mathcal{F}_{qT1+qL}^{ZV}(r_f, r_{\psi_D}) + Q_f s_W^2 \mathcal{F}_{qT2}^{ZV}(r_f, r_{\psi_D}) \right)$$

$$\epsilon_{AV} = \frac{g_D e Q_f}{4\pi^2} \mathcal{F}^{AV}(r_f, r_{\psi_D})$$

# Kinetic Mixing in FPVDM models







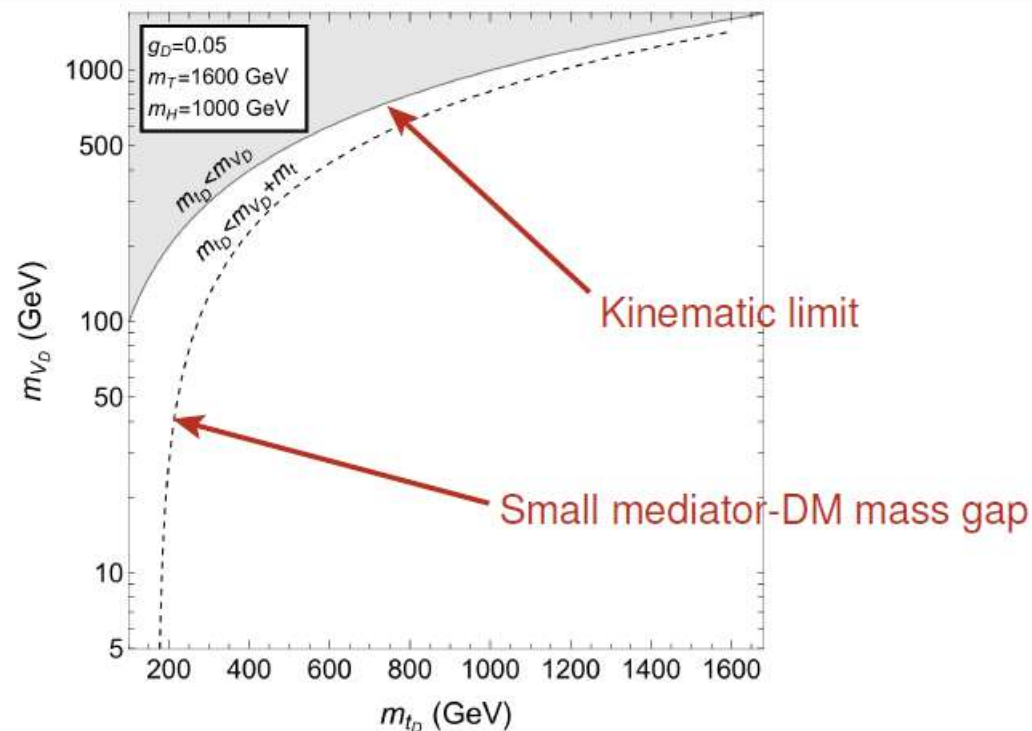
# Minimal VL top portal VDM: details of 2D space for chosen benchmarks

The VL fermion is composed of top partners and there is no mixing between scalars

$$\Psi = \begin{pmatrix} t_D \\ T \end{pmatrix} \quad \text{with} \quad m_t < m_{t_D} \leq m_T$$

$$\sin \theta_S = 0$$

Representative benchmarks:  $\begin{cases} g_D = 0.05, 0.5 & \text{strong or weak cosmological constraints} \\ m_T = 1600 \text{ GeV} \\ m_H = 1000 \text{ GeV} \end{cases}$  heavy enough to evade LHC constraints



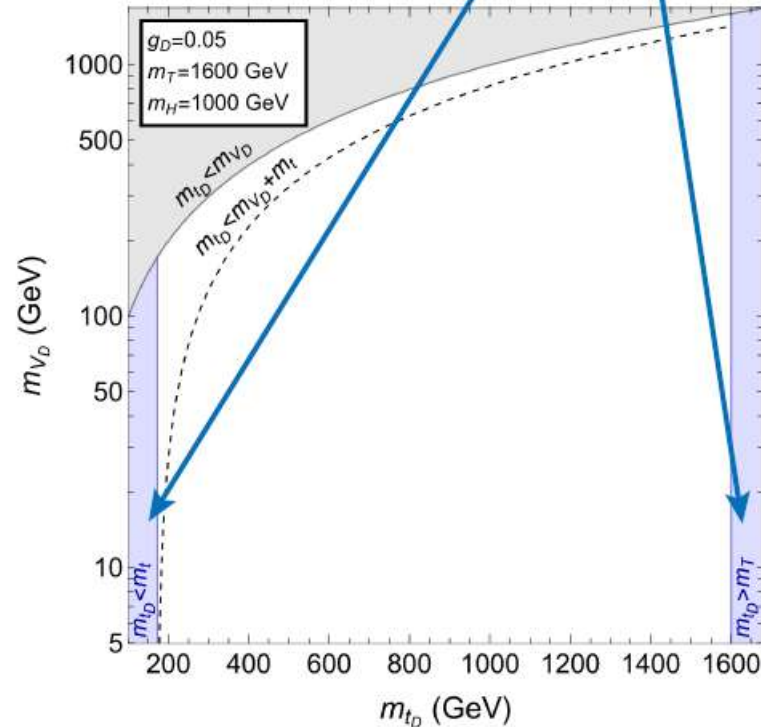
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Mediator mass bounded from below and above

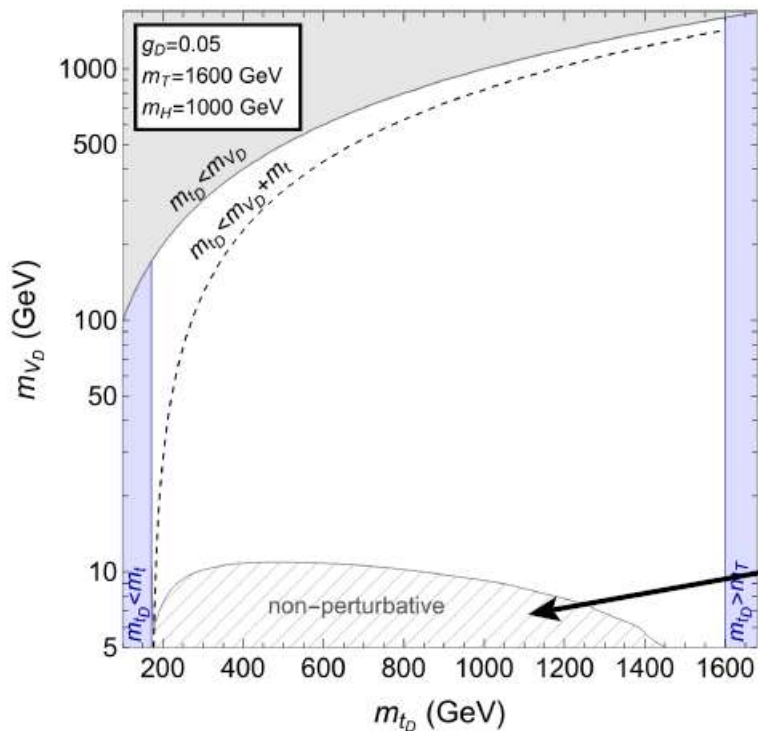
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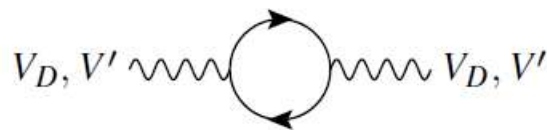
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Mediator mass bounded from below and above  
Light DM in non-perturbative region



$$\frac{m_V^{\text{pole}} - m_V}{m_V} > 50\%$$



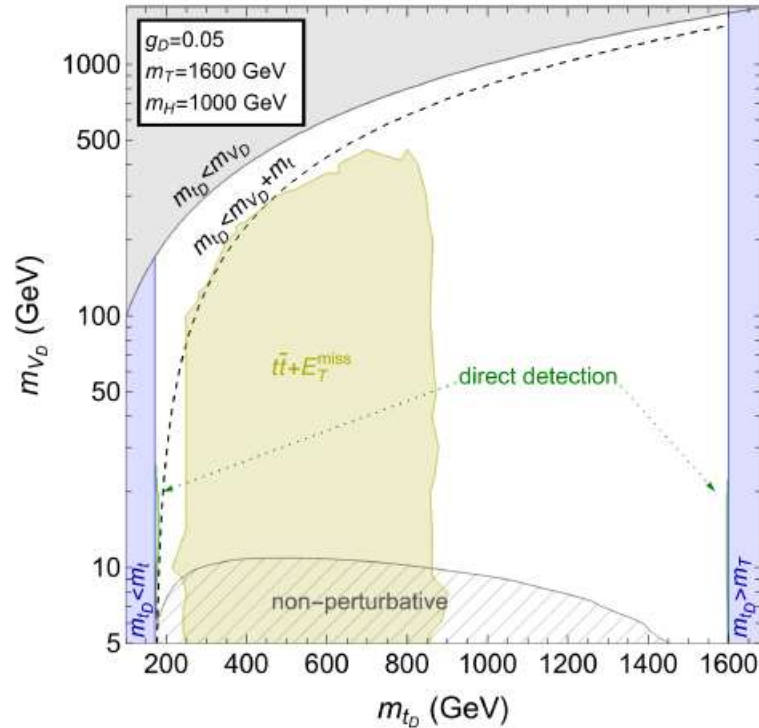
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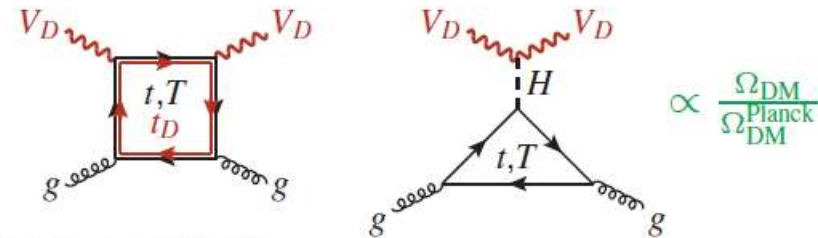
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Very weak direct detection constraints  
(mostly for  $m_{t_D} \sim m_t$  or  $m_{t_D} \sim m_T$  and light DM)



E. Aprile *et al.* [XENON],  
Dark Matter Search Results from a One Ton-Year Exposure of XENON1T,  
Phys. Rev. Lett. 121 (2018) no.11, 111302, arXiv:1805.12562 [astro-ph.CO]



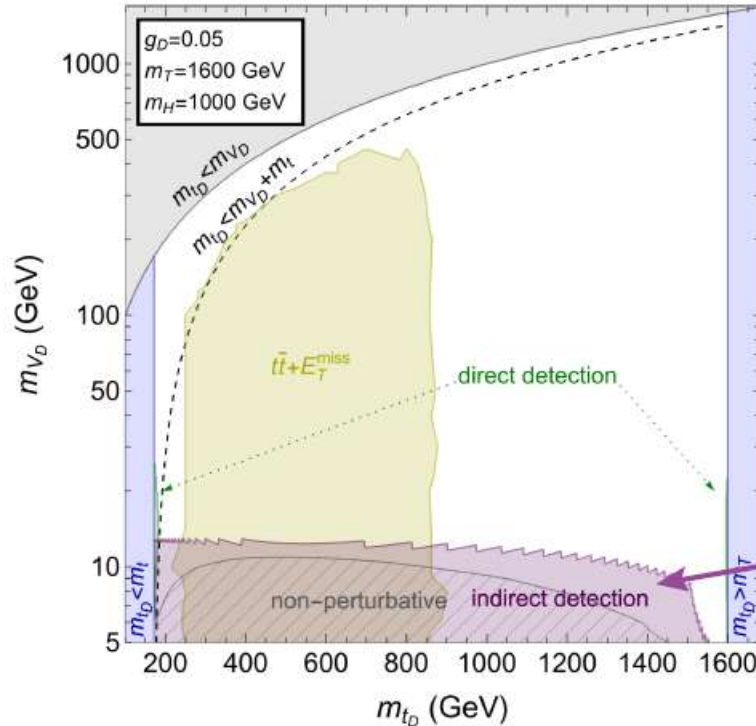
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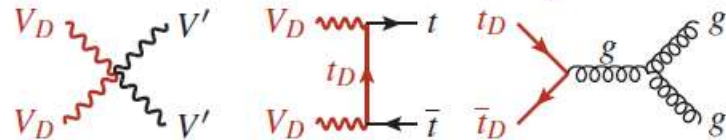


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Indirect detection constrains light DM



$$\propto \left( \frac{\Omega_{DM}}{\Omega_{Planck}^{DM}} \right)^2$$

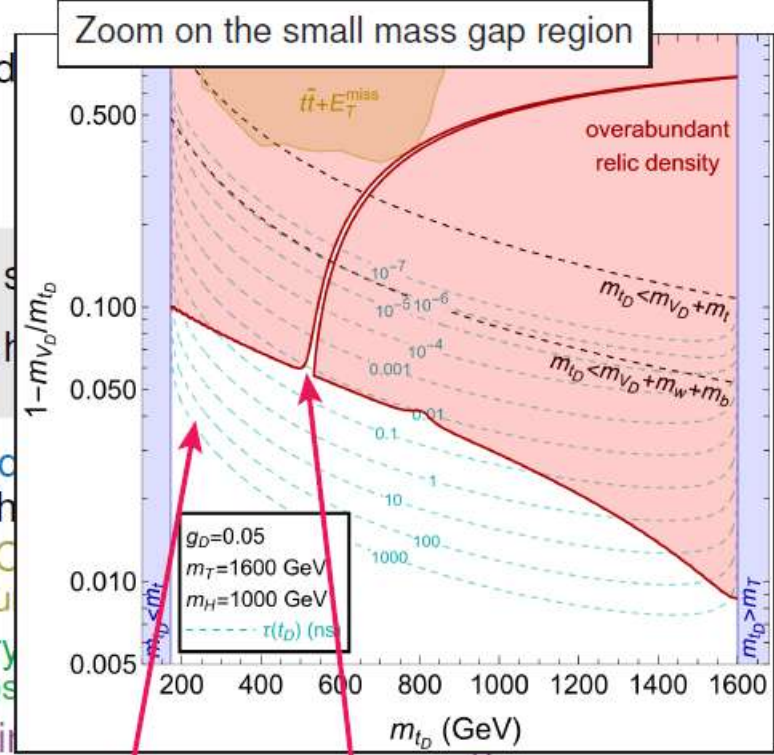
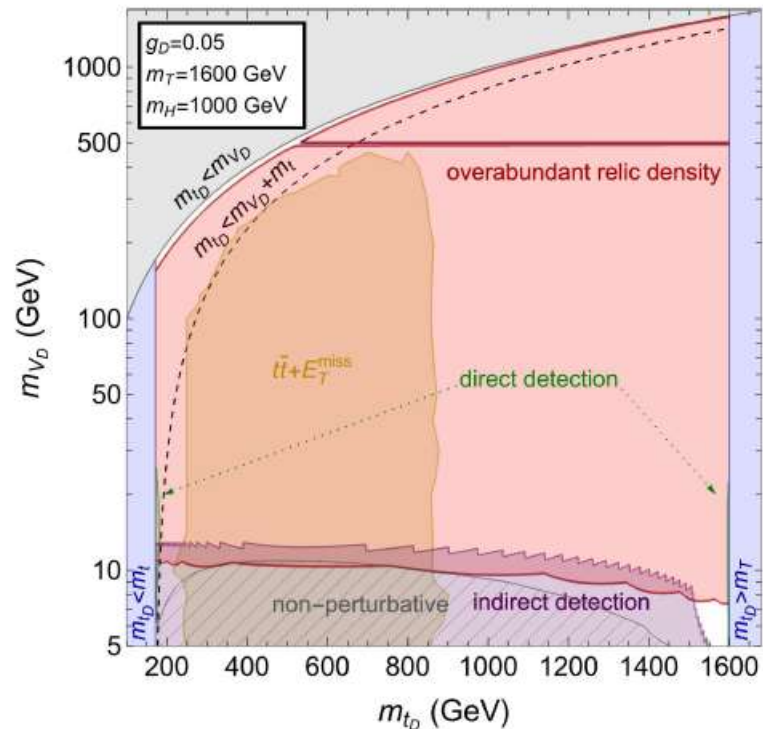
N. Aghanim *et al.* [Planck],  
Planck 2018 results. VI. Cosmological parameters,  
Astron. Astrophys. 641 (2020), A6, arXiv:1807.06209 [astro-ph.CO]

# Minimal VL top portal VDM: details of 2D space for chosen benchmarks

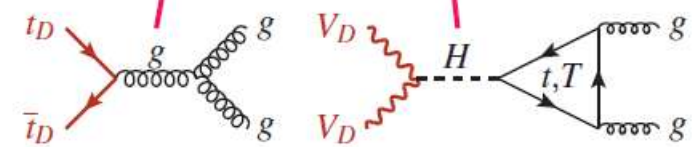
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Strong constrain from relic density  
 → the model “lives” on the red contours ( $\Omega_{DM}^{Planck}$ )  
 → overabundant region shrinks for larger  $g_D$   
 → and ID constraints vanish  
 → effective (co-)annihilation processes



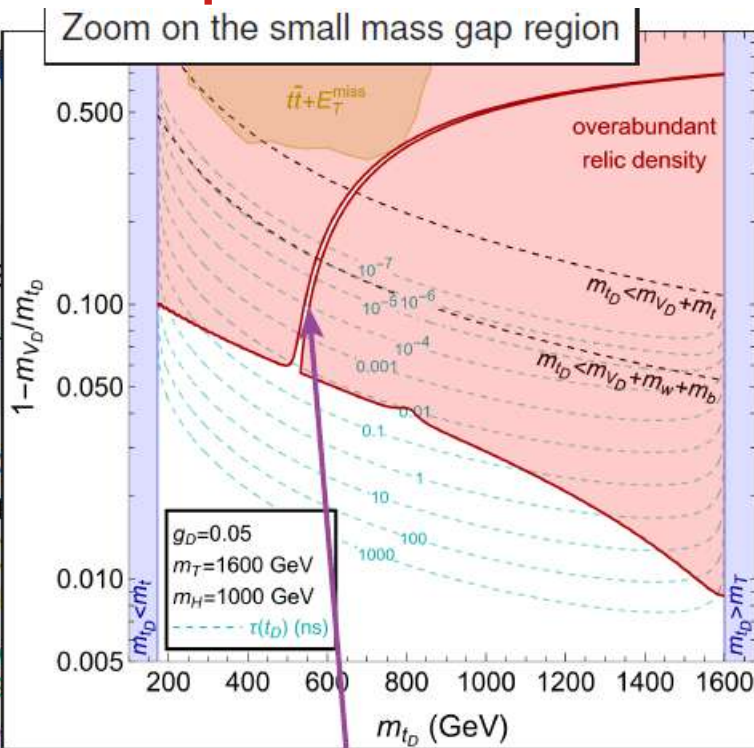
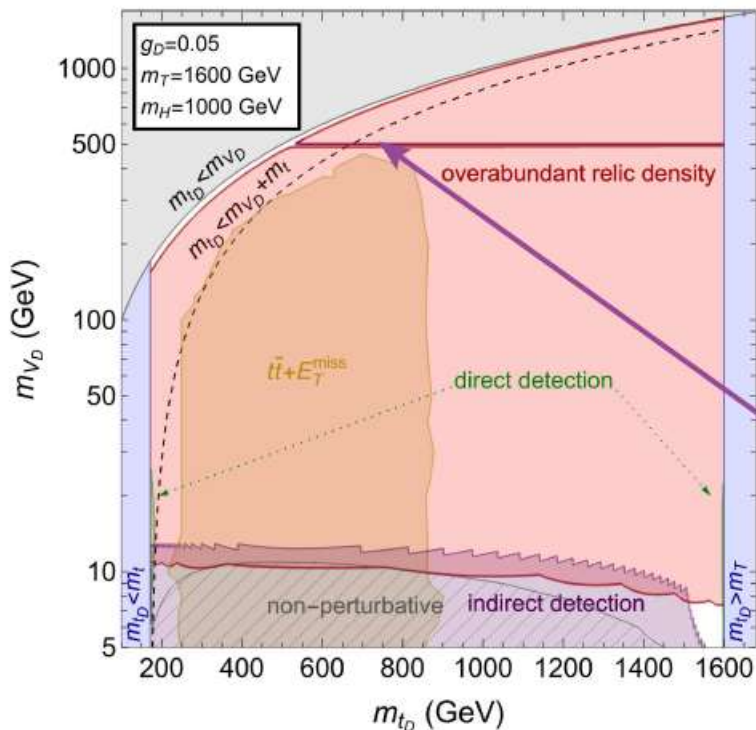


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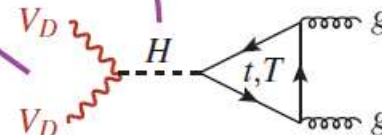
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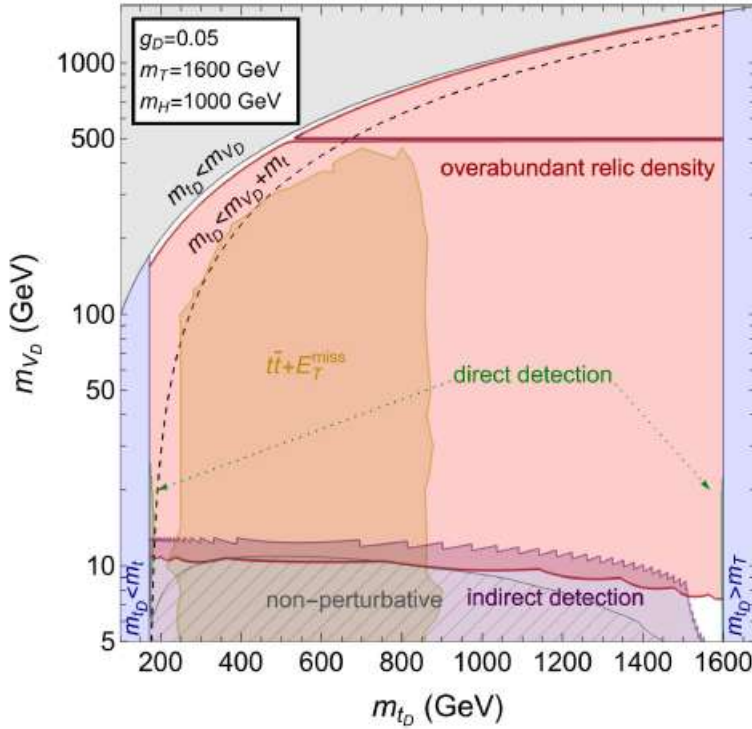
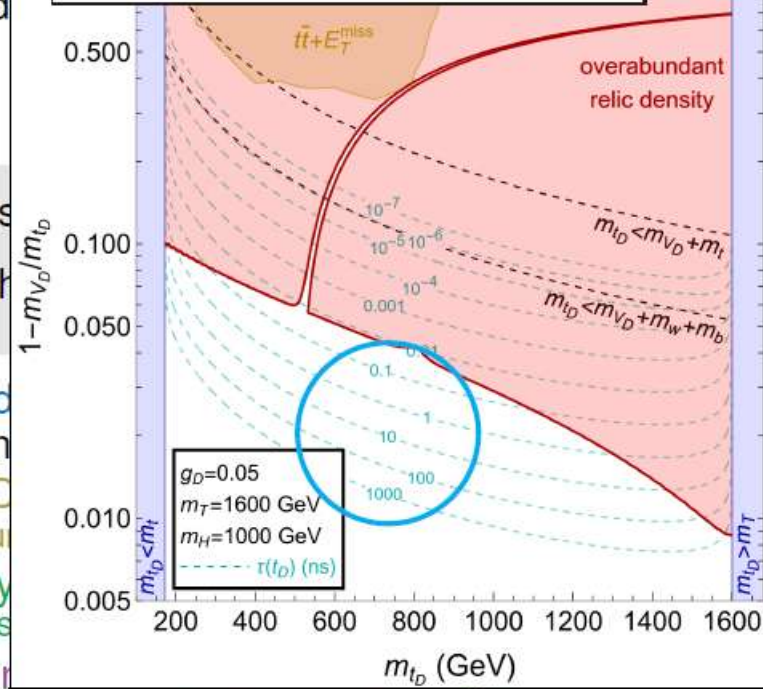
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Zoom on the small mass gap region



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- The mediator  $t_D$  can be long lived



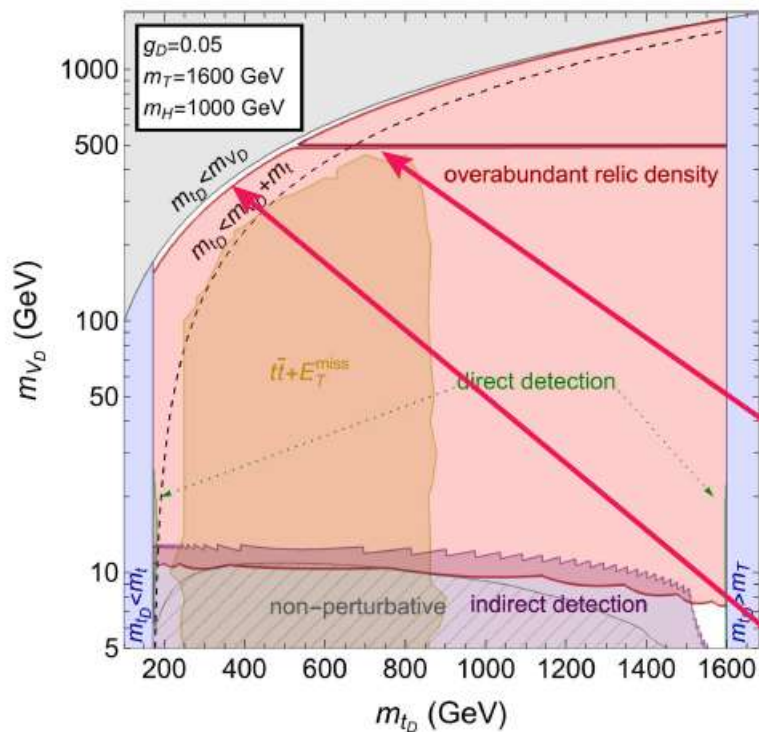
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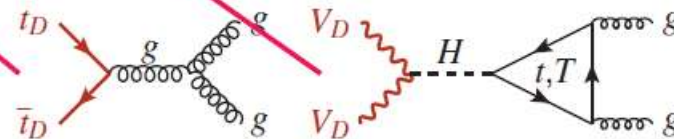
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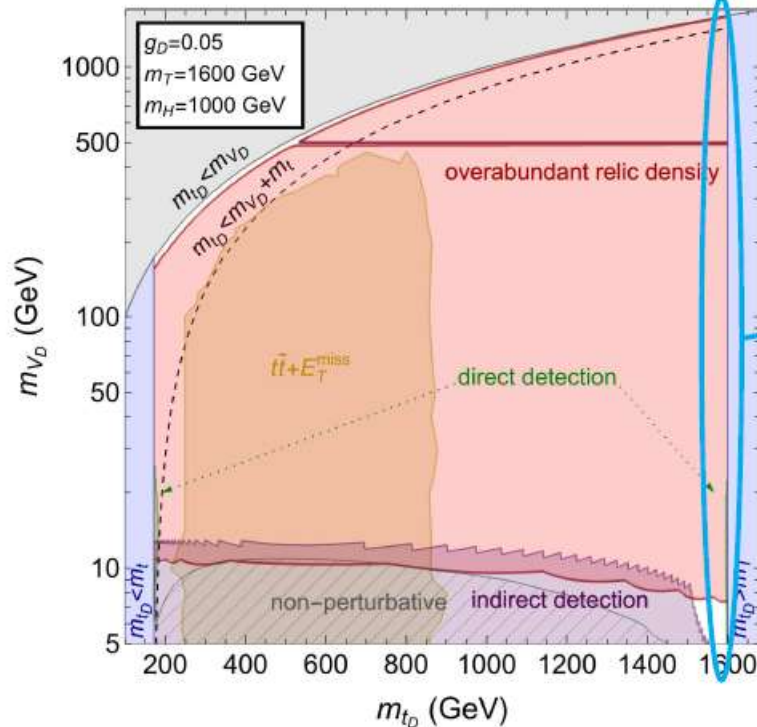
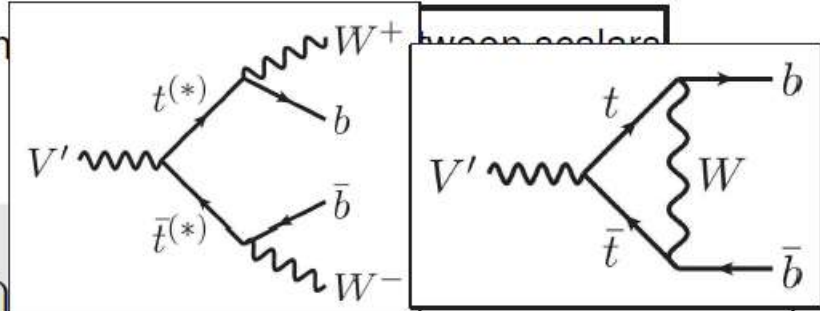


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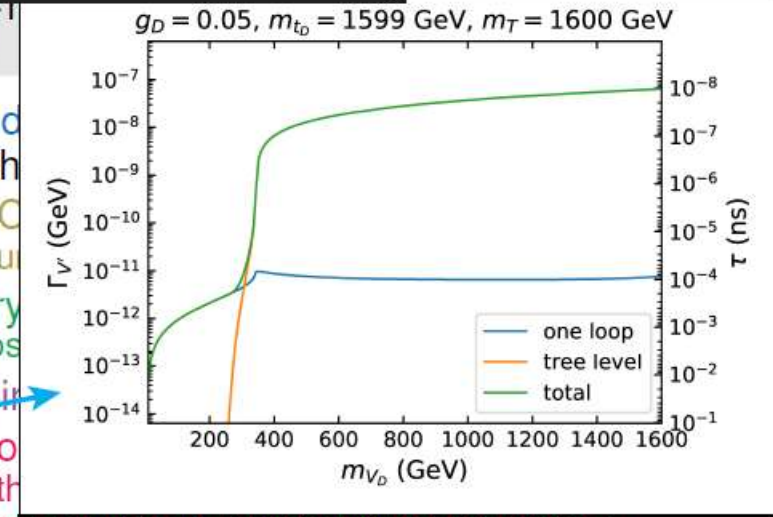
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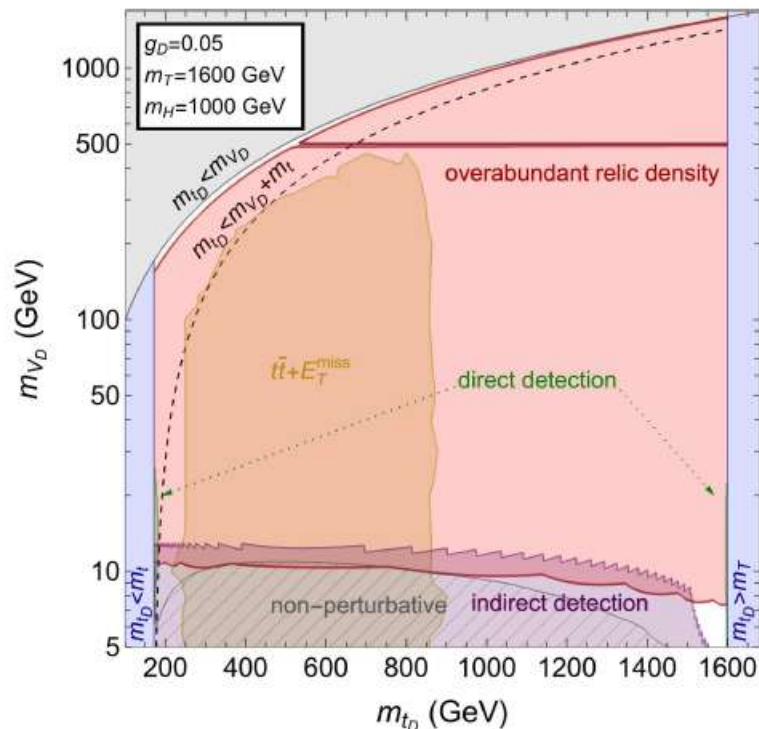
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just a simple realization of the model template  
**multiple features and signatures**

