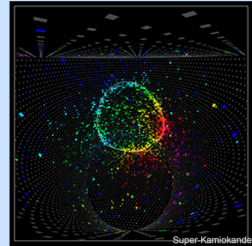
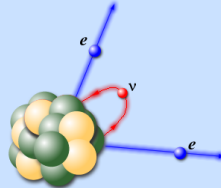
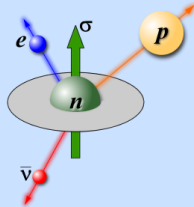


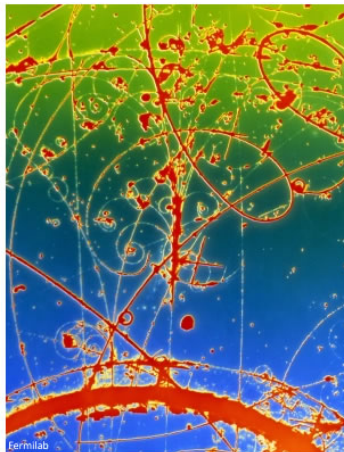
# Defying Lorentz invariance with neutrinos

Jorge S. Diaz |

HEP CHILE, 07.01.2016



- Lorentz and CPT violation
- Searching for Lorentz and CPT violation
  - **neutrino oscillations**
  - **beta decays**
- Summary



- Cornerstone of modern physics.
- Symmetry that underlies Special Relativity.
- Laws of physics are independent of speed and direction of propagation.
- Linked to CPT symmetry (relating properties of matter and antimatter).
- Established experiments indicate that nature is Lorentz invariant (so far).



Einstein & Lorentz (1921)

# Lorentz violation

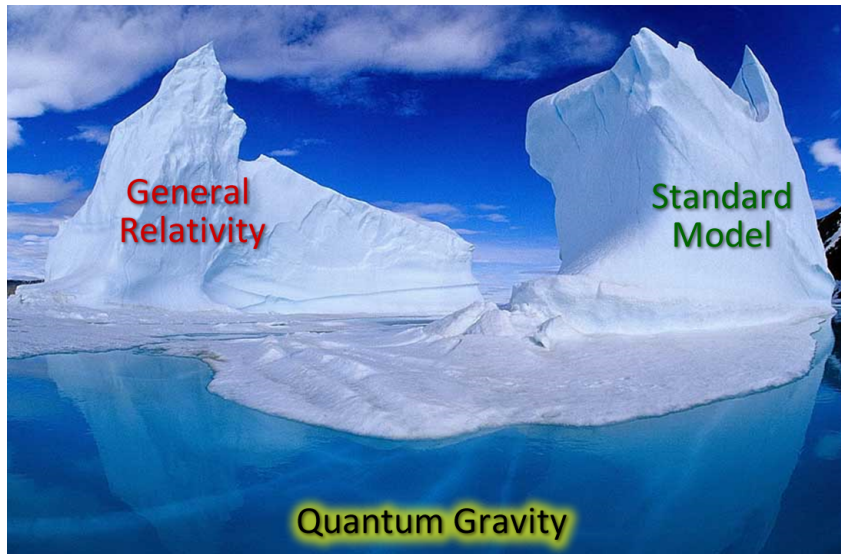
- Last 25 years, growing interest in the possibility that Lorentz symmetry may not be exact.
- Quantum gravity candidates involve the breaking of Lorentz symmetry.
- Lorentz symmetry is a basic building block of GR and the SM. Anything this fundamental should be tested.
- New era of high-precision measurements.



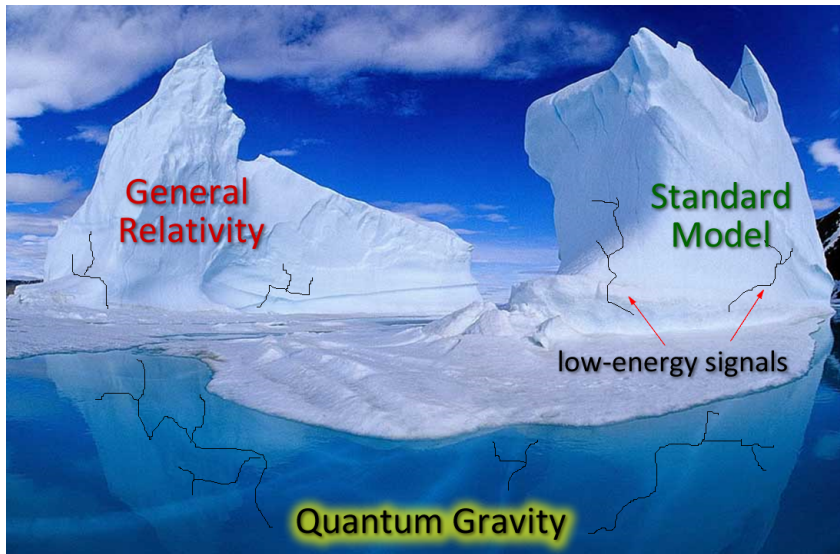


GR and the SM are expected to merge  
at the Planck scale

# Lorentz violation

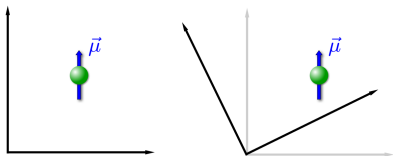


# Lorentz violation



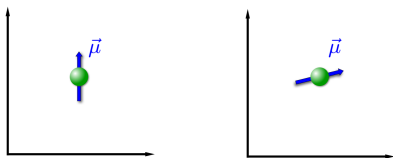
# Lorentz transformations

## Observer transformation



coordinate invariance

## Particle transformation



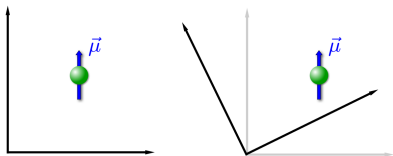
symmetry



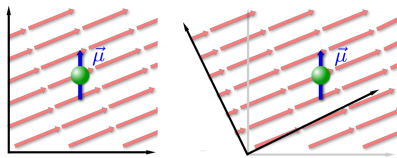
# Lorentz transformations

## Observer transformation

$$U = -\vec{\mu} \cdot \vec{B} = -\vec{\mu}' \cdot \vec{B}'$$

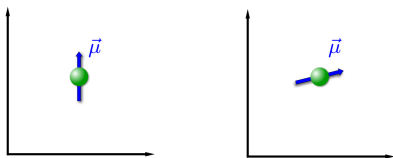


coordinate invariance



coordinate invariance

## Particle transformation

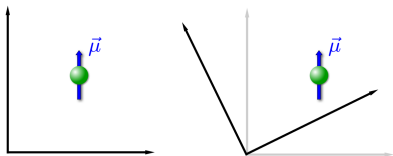


symmetry

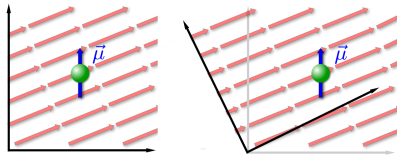
# Lorentz transformations

## Observer transformation

$$U = -\vec{\mu} \cdot \vec{B} = -\vec{\mu}' \cdot \vec{B}'$$



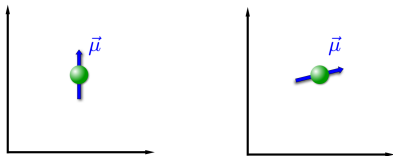
coordinate invariance



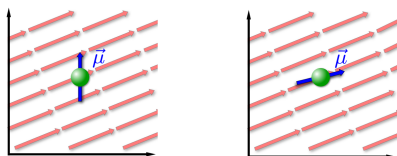
coordinate invariance

## Particle transformation

$$U = -\vec{\mu} \cdot \vec{B} \neq -\vec{\mu}' \cdot \vec{B}$$



symmetry

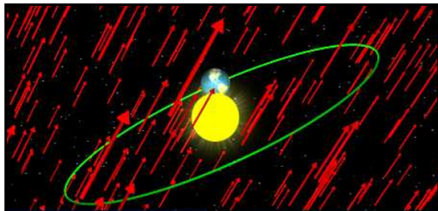


broken symmetry

# Standard-Model Extension (SME)

$$\text{SME} = \text{Standard Model coupled to General Relativity} + \text{all possible terms that break Lorentz symmetry}$$

Colladay & Kostelecký, PRD 55, 6760 (1997)  
Colladay & Kostelecký, PRD 58, 116002 (1998)  
Kostelecký, PRD 69, 105009 (2004)



- general framework to search for Lorentz violation
- defined experimental signatures

example (from fermion sector):

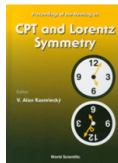
$$\mathcal{L}_{\text{LV}} \supset a_\mu (\bar{\psi} \gamma^\mu \psi)$$

- Standard fields
- Controlling coefficients
- Observer scalars
- CPT violation included (no  $m \neq \bar{m}$  terms)

# SME: theory & experiment playground

## Studies of CPT and Lorentz violation involve:

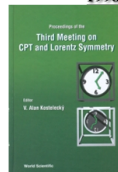
- beta decay
- neutrino oscillations
- oscillations and decays of K, B, D mesons
- particle-antiparticle comparisons
- matter interferometry
- birefringence and dispersion from cosmological sources
- clock-comparison measurements
- CMB polarization
- collider experiments
- electromagnetic resonant cavities
- equivalence principle
- gauge and Higgs particles
- high-energy astrophysical observations
- laboratory and gravimetric tests of gravity
- post-newtonian gravity in the solar system and beyond
- second- and third-generation particles
- space-based missions
- spectroscopy of hydrogen and antihydrogen
- spin-polarized matter



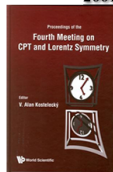
1998



2001



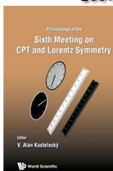
2004



2007



2010



2013

# Neutrinos in the SME

## effective hamiltonian

$$H_{\text{eff}} = \left( \begin{array}{c|c} h_0 & 0 \\ \hline 0 & h_0^* \end{array} \right) + \left( \begin{array}{c|c} \delta h_{\nu\nu} & \delta h_{\nu\bar{\nu}} \\ \hline \delta h_{\bar{\nu}\nu} & \delta h_{\bar{\nu}\bar{\nu}} \end{array} \right)$$

Kostelecký & Mewes, PRD 69, 016005 (2004)

← 6 × 6 matrix

Minimal\* neutrino 3 × 3 block:

$$H_{ab}^{\nu} = |\mathbf{p}| \delta_{ab} + \frac{m_{ab}^2}{2|\mathbf{p}|} + (a_L)_{ab}^{\alpha} \hat{p}_{\alpha} - (c_L)_{ab}^{\alpha\beta} \hat{p}_{\alpha} \hat{p}_{\beta} |\mathbf{p}|, \quad a, b = e, \mu, \tau; \hat{p}^{\alpha} = (1; \hat{\mathbf{p}})$$

## Novel effects

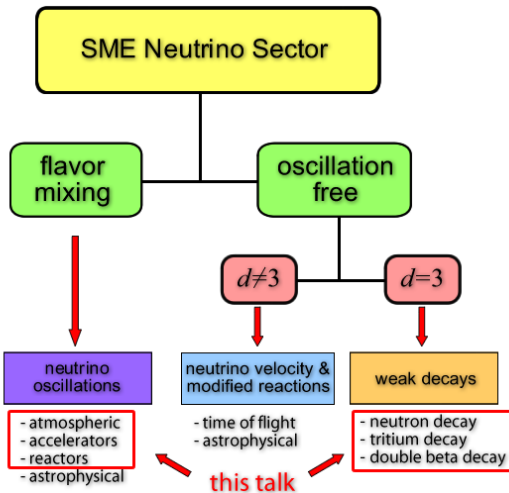
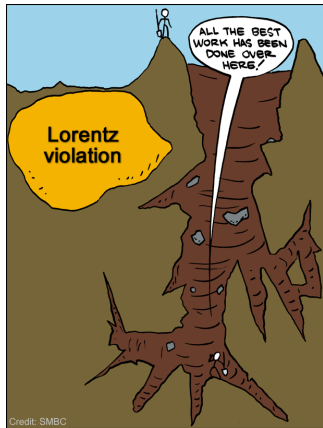
- unconventional energy dependence
- direction dependence
- sidereal time dependence
- CPT violation
- $\nu - \bar{\nu}$  mixing

Higher derivatives appear by including operators of arbitrary dimension  $d$

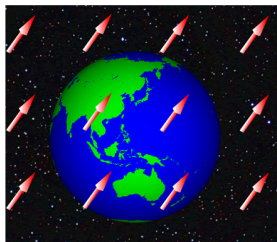
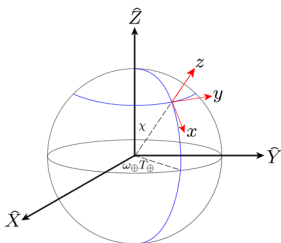
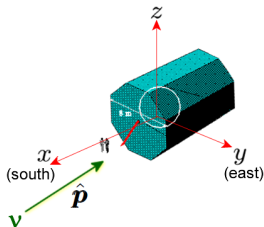
Kostelecký & Mewes, PRD 85, 096005 (2012)

\*: operators of dimension three and four only

## Complementarity between experiments



Kostelecký & Mewes, PRD 70, 076002 (2004)  
 JSD, Kostelecký & Mewes, PRD 80, 076007 (2009)



**Sidereal variation** of the oscillation probability:

$$\begin{aligned}
 P_{\nu_b \rightarrow \nu_a} = & (P_C)_{ab} + (P_{A_s})_{ab} \sin \omega_{\oplus} T_{\oplus} + (P_{A_c})_{ab} \cos \omega_{\oplus} T_{\oplus} \\
 & + (P_{B_s})_{ab} \sin 2\omega_{\oplus} T_{\oplus} + (P_{B_c})_{ab} \cos 2\omega_{\oplus} T_{\oplus} \\
 & + \dots
 \end{aligned}$$



# LV neutrino oscillations

**Example:** accelerator neutrinos

$$P_{\nu_{\mu} \rightarrow \nu_{\tau}} = P_{\nu_{\mu} \rightarrow \nu_{\tau}}^{(0)} + P_{\nu_{\mu} \rightarrow \nu_{\tau}}^{(1)}$$

JSD, Kostelecký & Mewes, PRD 80, 076007 (2009)  
MINOS, PRL 105, 151601 (2010)



## Search for Lorentz Invariance and CPT Violation with the MINOS Far Detector (MINOS Collaboration)

In the SME,  $P_{\mu\tau}^{(1)}$  is given by [8]

$$\begin{aligned}
 P_{\mu\tau}^{(1)} = & 2L\{(P_C^{(1)})_{\tau\mu} + (P_{\mathcal{A}_c}^{(1)})_{\tau\mu} \sin\omega_\oplus T_\oplus \\
 & + (P_{\mathcal{A}_c}^{(1)})_{\tau\mu} \cos\omega_\oplus T_\oplus + (P_{B_c}^{(1)})_{\tau\mu} \sin 2\omega_\oplus T_\oplus \\
 & + (P_{B_c}^{(1)})_{\tau\mu} \cos 2\omega_\oplus T_\oplus\}, \quad (1)
 \end{aligned}$$

where  $L = 735$  km is the distance from neutrino production in the NuMI beam to the MINOS FD [2],  $T_\oplus$  is the local sidereal time (LST) at neutrino detection, and the coefficients  $(P_C^{(1)})_{\tau\mu}$ ,  $(P_{\mathcal{A}_c}^{(1)})_{\tau\mu}$ ,  $(P_{\mathcal{A}_c}^{(1)})_{\tau\mu}$ ,  $(P_{B_c}^{(1)})_{\tau\mu}$ , and  $(P_{B_c}^{(1)})_{\tau\mu}$  contain the LV and CPTV information.



Fermilab

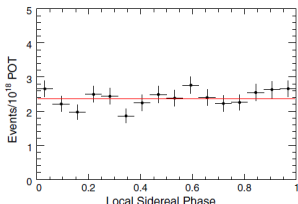


TABLE III. 99.7% C.L. limits on SME coefficients for  $\nu_\mu \rightarrow \nu_\tau$ ;  $(a_L)_{\mu\tau}^\alpha$  have units [GeV];  $(c_L)_{\mu\tau}^{\alpha\beta}$  are unitless.

Coeff.	Limit	Coeff.	Limit
$(a_L)_{\mu\tau}^X$	$5.9 \times 10^{-23}$	$(a_L)_{\mu\tau}^Y$	$6.1 \times 10^{-23}$
$(c_L)_{\mu\tau}^{TX}$	$0.5 \times 10^{-23}$	$(c_L)_{\mu\tau}^{TY}$	$0.5 \times 10^{-23}$
$(c_L)_{\mu\tau}^{XX}$	$2.5 \times 10^{-23}$	$(c_L)_{\mu\tau}^{YY}$	$2.4 \times 10^{-23}$
$(c_L)_{\mu\tau}^{XY}$	$1.2 \times 10^{-23}$	$(c_L)_{\mu\tau}^{YZ}$	$0.7 \times 10^{-23}$
$(c_L)_{\mu\tau}^{XZ}$	$0.7 \times 10^{-23}$	...	...

# LV neutrino oscillations

**Example:** reactor antineutrinos

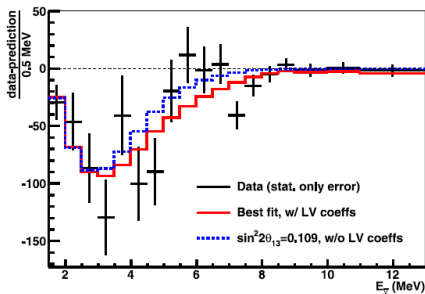
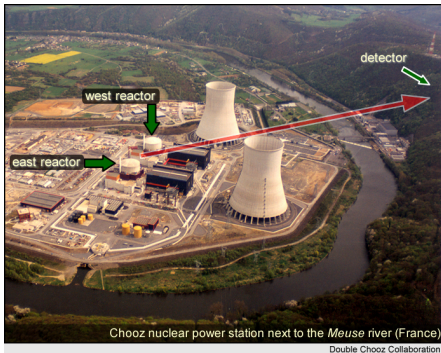
JSD, Katori, Spitz & Conrad, PLB 727, 412 (2013)

$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = P_{\bar{\nu}_e \rightarrow \bar{\nu}_e}^{(0)} - P_{\bar{\nu}_e \rightarrow \nu_x}^{(2)}$$



Credit: QuantumDiaries.org

# LV neutrino oscillations



upper limits (90% CL)

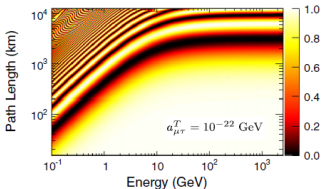
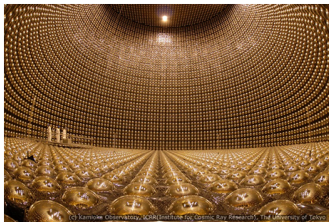
-	$ \tilde{\delta}_{e\bar{e}}^{ZT}  < 9.7 \times 10^{-18}$	$ \tilde{g}_{e\bar{e}}^{ZZ}  < 3.3 \times 10^{-17}$
-	$ \tilde{\delta}_{\mu\bar{\mu}}^{ZT}  < 2.3 \times 10^{-16}$	$ \tilde{g}_{\mu\bar{\mu}}^{ZZ}  < 8.1 \times 10^{-16}$
-	$ \tilde{\delta}_{\tau\bar{\tau}}^{ZT}  < 2.3 \times 10^{-16}$	$ \tilde{g}_{\tau\bar{\tau}}^{ZZ}  < 8.1 \times 10^{-16}$
$ \tilde{H}_{e\bar{\mu}}^Z  < 1.4 \times 10^{-19}$	$ \tilde{\delta}_{e\bar{\mu}}^{ZT}  < 2.7 \times 10^{-17}$	$ \tilde{g}_{e\bar{\mu}}^{ZZ}  < 9.3 \times 10^{-17}$
$ \tilde{H}_{e\bar{\tau}}^Z  < 1.4 \times 10^{-19}$	$ \tilde{\delta}_{e\bar{\tau}}^{ZT}  < 2.7 \times 10^{-17}$	$ \tilde{g}_{e\bar{\tau}}^{ZZ}  < 9.3 \times 10^{-17}$
$ \tilde{H}_{\mu\bar{\tau}}^Z  < 1.7 \times 10^{-18}$	$ \tilde{\delta}_{\mu\bar{\tau}}^{ZT}  < 4.4 \times 10^{-16}$	$ \tilde{g}_{\mu\bar{\tau}}^{ZZ}  < 1.5 \times 10^{-15}$

# LV neutrino oscillations

**Example:** atmospheric neutrinos

Super-Kamiokande, PRD 91, 052003 (2015)

$$P_{\nu_\mu \rightarrow \nu_\mu} = P_{\nu_\mu \rightarrow \nu_\mu}^{(0)} + P_{\nu_\mu \rightarrow \nu_\mu}^{(1)}$$



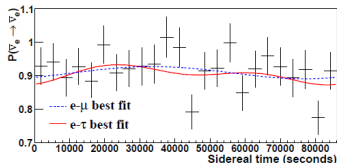
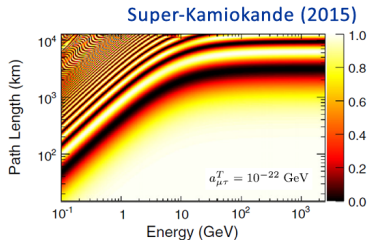
LV parameter	Limit at 95% C.L.	
$e\mu$	$\text{Re}(a^T)$	$1.8 \times 10^{-23}$ GeV
	$\text{Im}(a^T)$	$1.8 \times 10^{-23}$ GeV
	$\text{Re}(c^{TT})$	$8.0 \times 10^{-27}$
	$\text{Im}(c^{TT})$	$8.0 \times 10^{-27}$
$e\tau$	$\text{Re}(a^T)$	$4.1 \times 10^{-23}$ GeV
	$\text{Im}(a^T)$	$2.8 \times 10^{-23}$ GeV
	$\text{Re}(c^{TT})$	$9.3 \times 10^{-25}$
	$\text{Im}(c^{TT})$	$1.0 \times 10^{-24}$
$\mu\tau$	$\text{Re}(a^T)$	$6.5 \times 10^{-24}$ GeV
	$\text{Im}(a^T)$	$5.1 \times 10^{-24}$ GeV
	$\text{Re}(c^{TT})$	$4.4 \times 10^{-27}$
	$\text{Im}(c^{TT})$	$4.2 \times 10^{-27}$

# LV neutrino oscillations

## Experimental searches

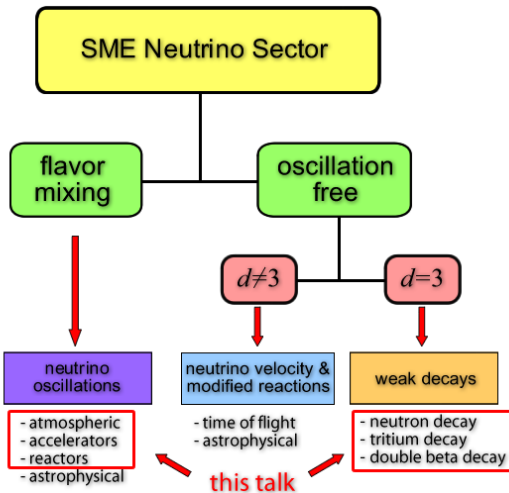
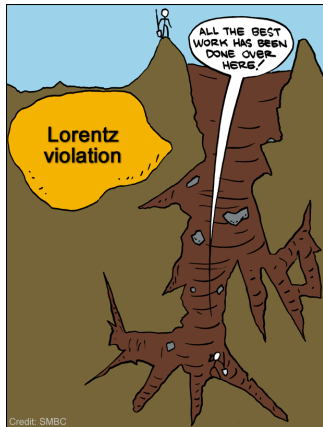
- **LSND** PRD 72, 076004 (2005)
- **MINOS** PRL 101, 151601 (2008)
- **IceCube** PRD 82, 112003 (2010)
- **MINOS** PRL 105, 151601 (2010)
- **MINOS** PRD 85, 031101 (2012)
- **Double Chooz** PRD 86, 112009 (2012)
- **MiniBooNE** PLB 718, 1303 (2013)
- **Rebel & Mufson** AP 48, 78 (2013)
- **Conrad, JSD, Katori, Spitz** PLB 727, 412 (2013)
- **Super-Kamiokande** PRD 91, 052003 (2015)
- ...

Kostelecký & Mewes, PRD 70, 076002 (2004)  
JSD, Kostelecký & Mewes, PRD 80, 076007 (2009)



Double Chooz (2012)

## Complementarity between experiments



## Theoretical considerations:

- antineutrino spinors get modified
- modified antineutrino phase space
- coefficients:  $(\mathbf{a}_{\text{of}}^{(3)})_{jm}$

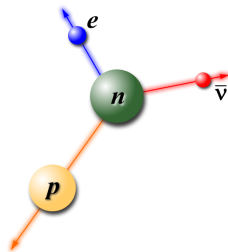
$$d\Gamma \propto E\omega \left\{ 1 + a \frac{\mathbf{p} \cdot \tilde{\mathbf{q}}}{E\omega} + A \frac{\hat{\mathbf{n}} \cdot \mathbf{p}}{E} + B \frac{\hat{\mathbf{n}} \cdot \tilde{\mathbf{q}}}{\omega} + D \frac{\hat{\mathbf{n}} \cdot (\mathbf{p} \times \tilde{\mathbf{q}})}{E\omega} \right\} \frac{d^3p}{2E} \frac{d^3q}{2\omega} \delta(E + \omega - E_0)$$

## Observable effects

- spectrum distortion:  $(\mathbf{a}_{\text{of}}^{(3)})_{00}$
- modified angular correlations:  $(\mathbf{a}_{\text{of}}^{(3)})_{10}, (\mathbf{a}_{\text{of}}^{(3)})_{11}$

JSD, Kostelecký & Lehnert, PRD 88, 071902 (2013)

JSD, Adv.HEP 2014, 305298 (2014)



## LV in W boson

Vos et al., PRC 92, 052501 (2015)

Vos et al., PRC 91, 038501 (2015)

Vos et al., Rev. Mod. Phys. (2015)

Vos et al., PLB 729, 112 (2014)

Noordmans et al., PRD 89, 101702 (2014)

Noordmans, PRC 87, 055502 (2013)

Altschul, PRD 88, 076015 (2013)

Dijck et al., PRD 88, 07190 (2013)

Noordmans et al., PRL 111, 171601 (2013)

Müller et al., PRD 88, 071901 (2013)



# LV tritium decay

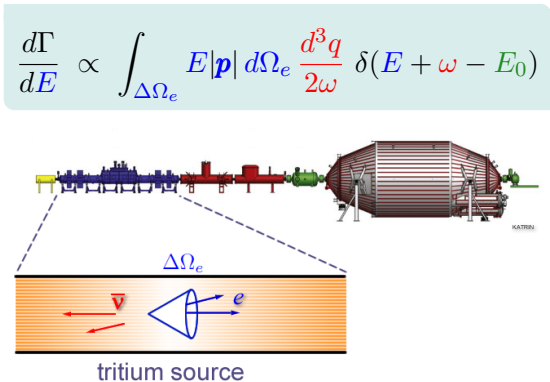
## Endpoint measurements

JSD, Kostelecký & Lehnert, PRD 88, 071902 (2013)

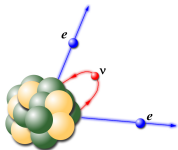
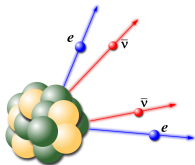
JSD, Adv.HEP 2014, 305298 (2014)



- Mainz
- Troitsk
- KATRIN



# LV double beta decay



$$\Gamma \propto \int |\overline{\mathcal{M}^{2\nu}}|^2 F(Z, E_1) F(Z, E_2) \times \delta(E_1 + E_2 + \omega_1 + \omega_2 - \Delta M) \times \frac{d^3 p_1}{2E_1} \frac{d^3 p_2}{2E_2} \frac{d^3 q_1}{2\omega_1} \frac{d^3 q_2}{2\omega_2}$$

$$\Gamma \propto \int |\overline{\mathcal{M}^{0\nu}}|^2 F(Z, E_1) F(Z, E_2) \times \delta(E_1 + E_2 - \Delta M) \frac{d^3 p_1}{2E_1} \frac{d^3 p_2}{2E_2}$$

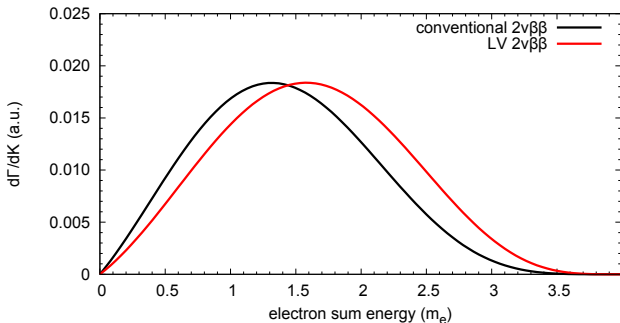
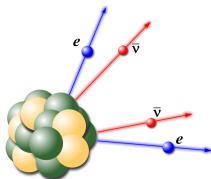
# LV double beta decay: $2\nu\beta\beta$

## Isotropic spectral distortion

- directionality is irrelevant
- unconventional energy dependence modifies the sum electron spectrum

JSD, Kostelecký & Lehnert, PRD 88, 071902 (2013)

JSD, PRD 89, 036002 (2014)



# LV double beta decay: $0\nu\beta\beta$

## Half-life gets modified

- neutrino propagator modifies the effective mass

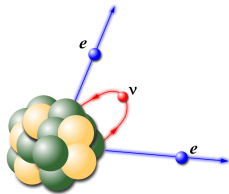
$$\frac{1}{T_{1/2}} = G(Z, Q) |M^{0\nu}|^2 m^2$$

$$m^2 \rightarrow m^2 + m \frac{g}{R} + \left(\frac{g}{R}\right)^2$$

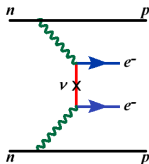
$R$ : nuclear radius

- neutrinoless double beta decay can occur for massless neutrinos

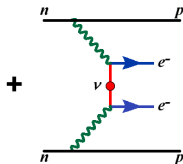
JSD, Kostelecký & Lehnert, PRD 88, 071902 (2013)  
JSD, PRD 89, 036002 (2014)



Majorana  
mass coupling



SME Majorana coupling  
Lorentz & CPT violation



+

# LV double beta decay: $0\nu\beta\beta$

## $\beta\beta$ angular correlation

JSD, PRD 89, 036002 (2014)

- unique sensitivity for experiments with tracking system
- angular correlation

$$\frac{d\Gamma}{dx_1 dx_2} = \frac{\Gamma}{4} (1 - K x_1 x_2)$$

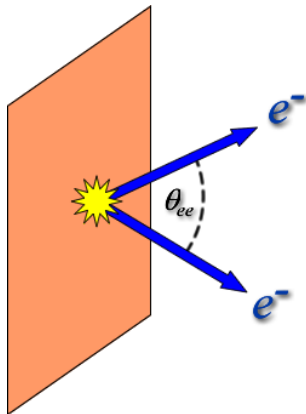
- forward-backward asymmetry

$$\mathcal{A} = \frac{N_+ - N_-}{N_+ + N_-} \propto K,$$

with  $N_{\pm}$  events with  $\theta_{ee} \gtrless 90^\circ$

Majorana-mass case:  $\mathcal{A}$  is constant

LV case:  $\mathcal{A}$  depends on location and can vary sidereal time



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*“Today we say that the law of relativity is supposed to be true at all energies, but someday somebody may come along and say how stupid we were.”*

R.P. Feynman