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Modelling very high energy GRB afterglow emission

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Detection of delayed sub-TeV photons from Gamma-Ray Bursts (GRBs) by MAGIC and HESS has proved the promising future of GRB afterglow studies with the Cherenkov Telescope Array, the next-generation ground-based gamma-ray astronomy observatory. With the unprecedented sensitivity of CTA, afterglow detection rates are expected to increase dramatically in the coming decade. We embark on exploring the multi-dimensional afterglow parameter space to see the detectability of sub-TeV photons by CTA. Sub-TeV emission is always due to the self-Compton process. We find that jets with high kinetic energy decelerating into a dense ambient medium are better candidates for CTA. We apply our results in the context of shortduration GRBs and counterparts to Neutron Star mergers from the local universe.

Session

Astroparticle Physics and Cosmology

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