

Precision Electroweak Tensions and a Wide Dark Photon

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with Aaron Pierce and Keisuke Harigaya



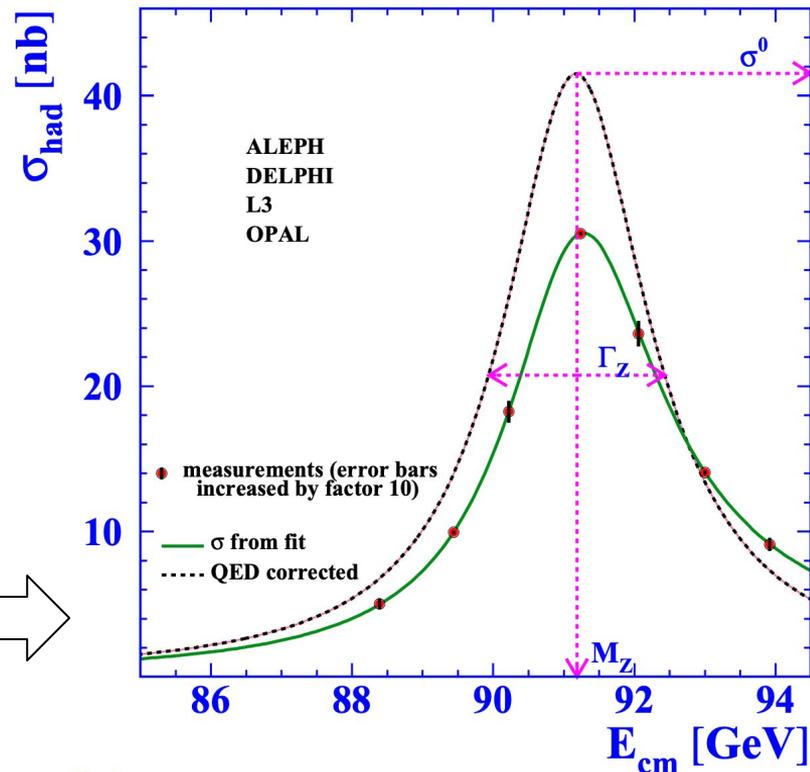
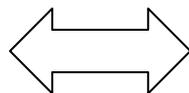
Outline

- Precision Electroweak (PEW) Analysis
 - Oblique Parameters STU
- Kinetic Mixing and the Dark Photon
- Results

The Precision Electroweak Fit

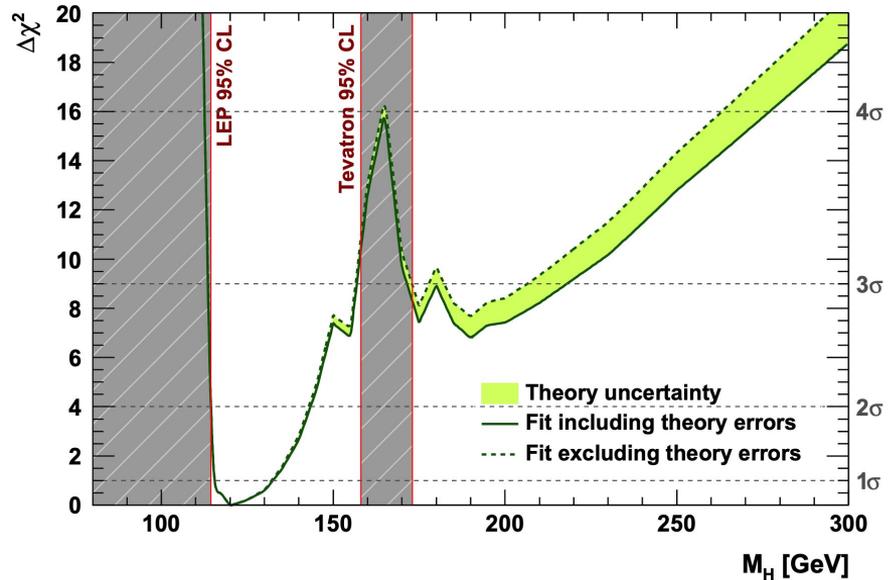
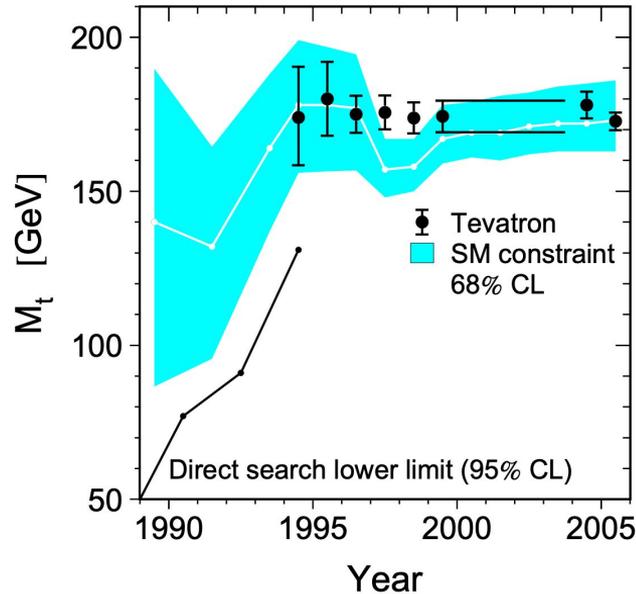
- Precision Measurements
- Computation of observables
- Fit tells us:
 - If the model works
 - Values for parameters

$$\mathcal{O} = \mathcal{O}(M_z, M_h, M_t, \alpha_s, \Delta\alpha_{had}^{(5)})$$



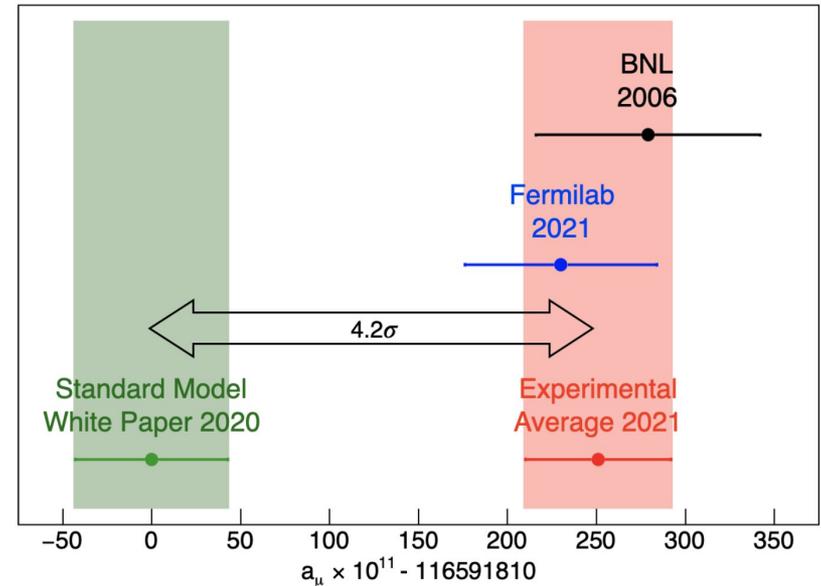
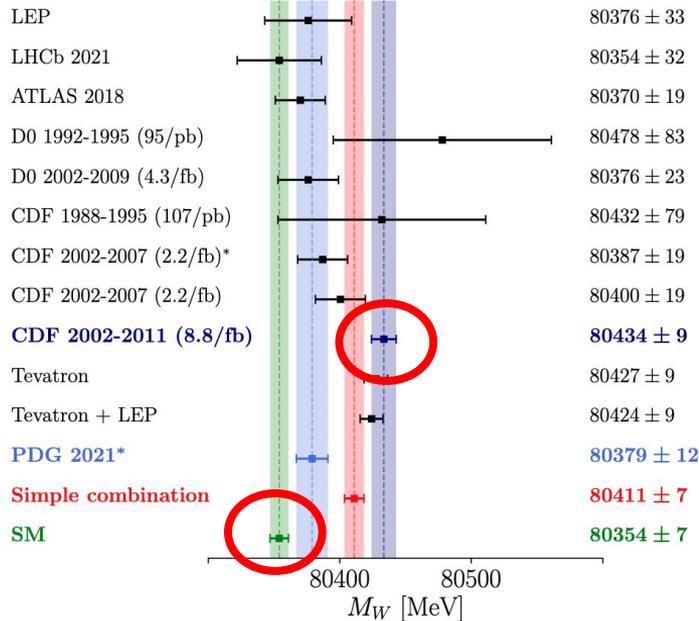
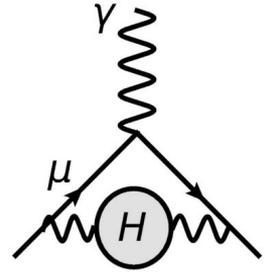
The Precision Electroweak Fit - Successes

- Successful predictions of the top and higgs masses

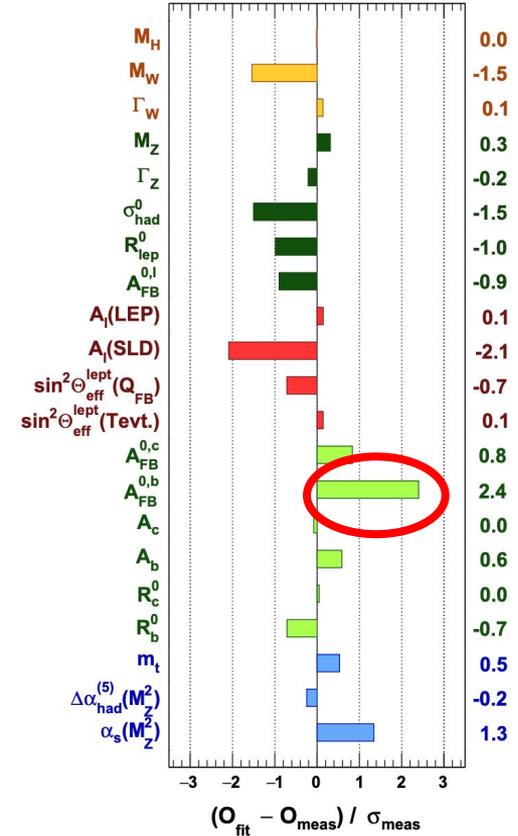
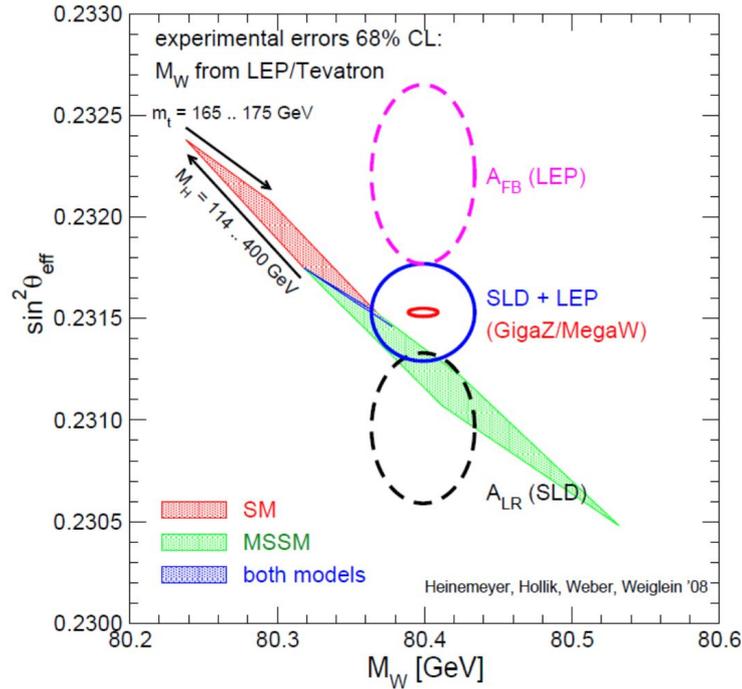


Precision Electroweak Fit - Tensions

- Things look pretty good, but a few tensions remain



Precision Electroweak Fit - Tensions (2)



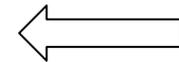
Precision Electroweak Analysis + BSM Physics

- One Option: Modify SM and recompute all observables
 - Problem: Not very efficient
- Better Option: Find a generic way to parameterize BSM effects
 - Oblique Parameters - Capture modifications to gauge boson propagators
 - S, T, U

$$S \leftrightarrow (h^\dagger \sigma^a h) W_{\mu\nu}^a B^{\mu\nu}$$

$$T \leftrightarrow (h^\dagger D^\mu h)(D_\mu h^\dagger h)$$

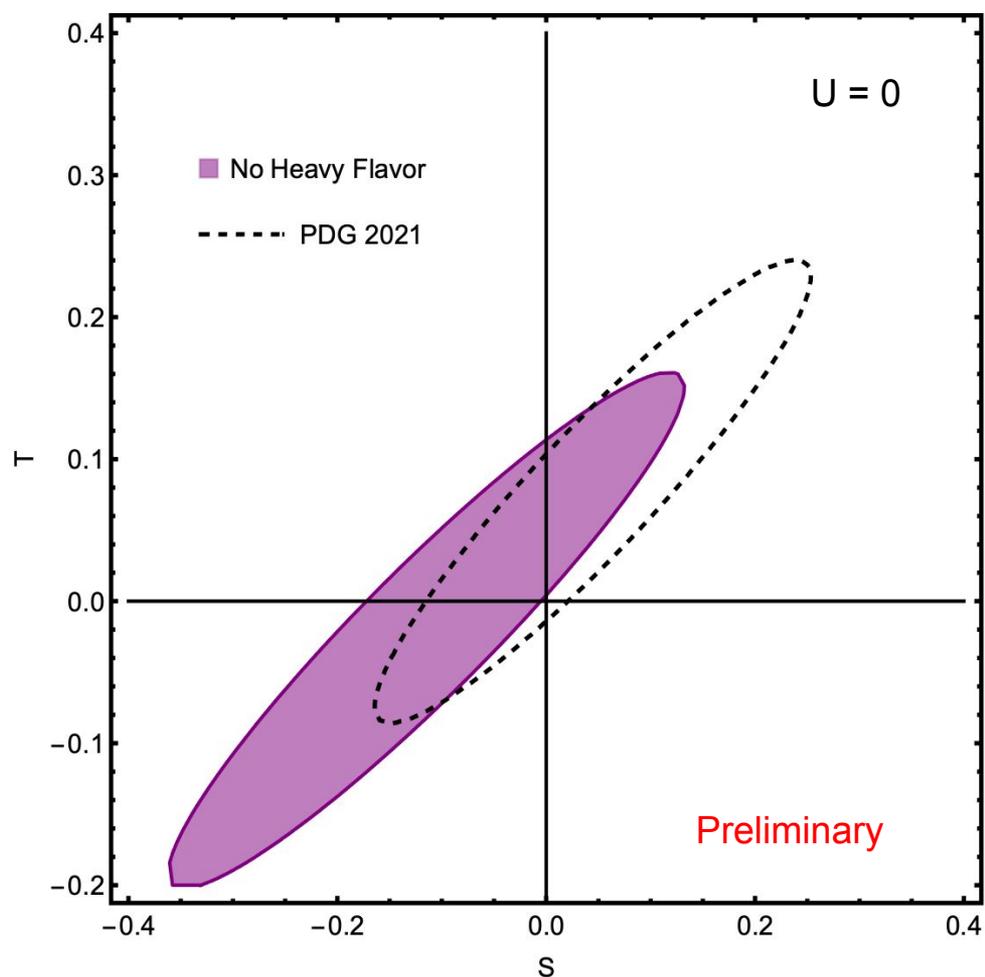
$$U \leftrightarrow (h^\dagger W^{a\mu\nu} h)(h^\dagger W_{\mu\nu}^a h)$$



Dim 8

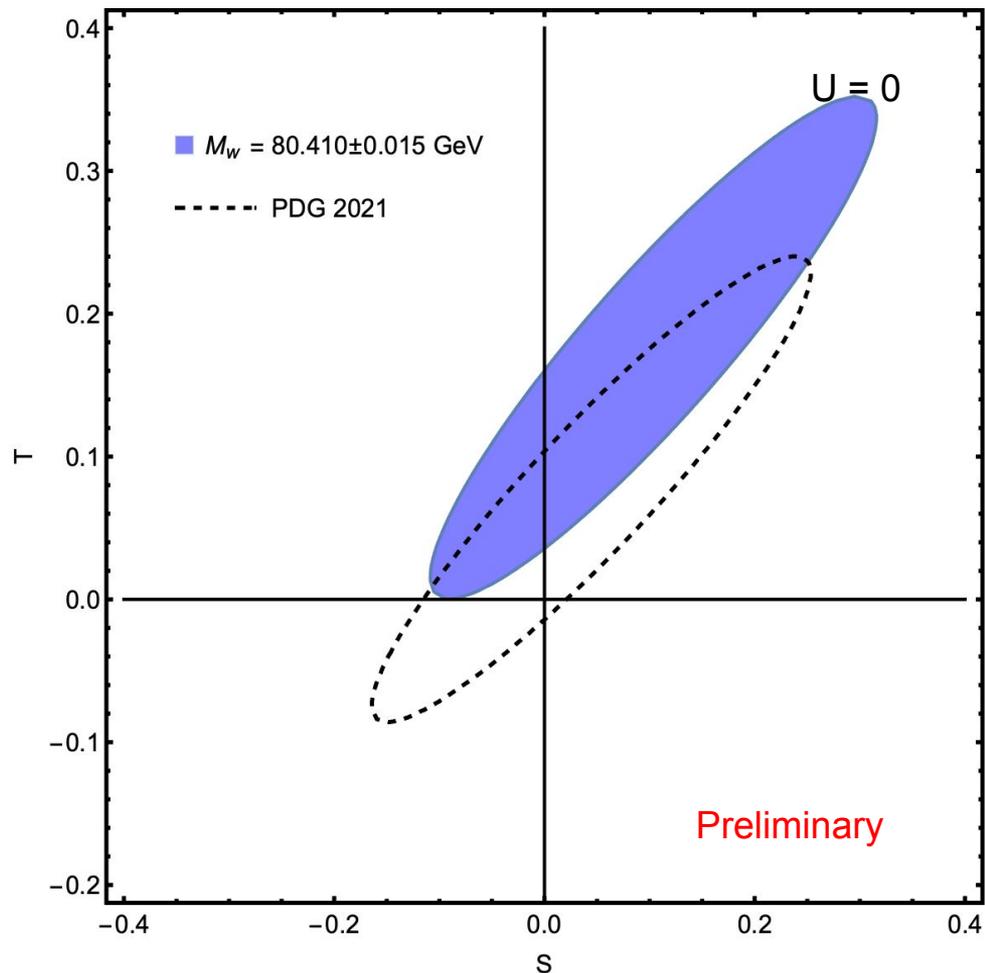
Oblique Fits

$$\chi^2(\theta = M_z, M_h, M_t, \alpha_s, \Delta\alpha_{had}^{(5)}, S, T)$$
$$= (\mathbf{y} - \boldsymbol{\mu}(\theta))^T V^{-1} (\mathbf{y} - \boldsymbol{\mu}(\theta))$$



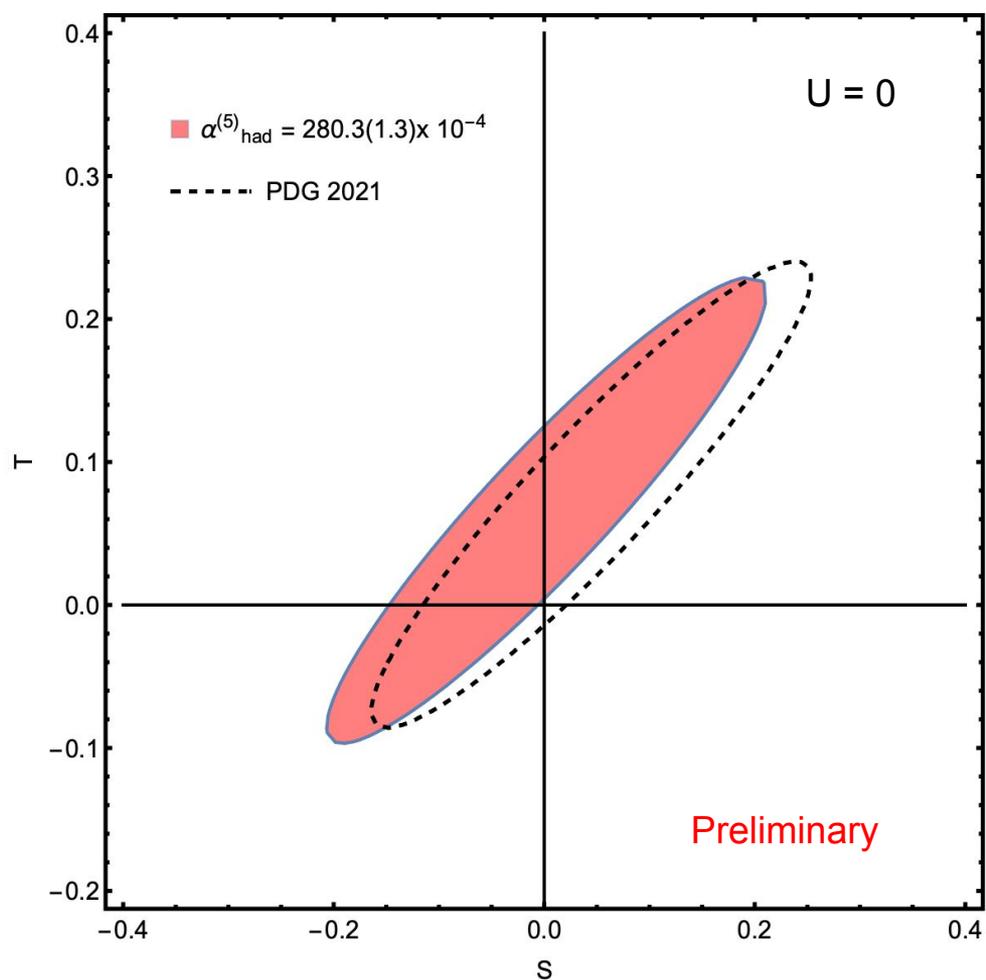
Oblique Fits

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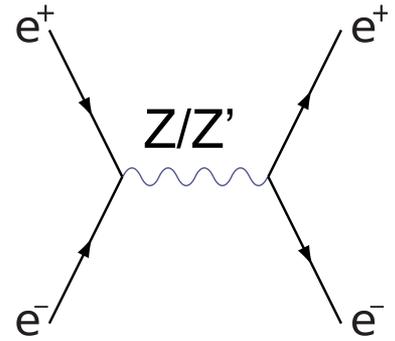
Oblique Fits

$$\chi^2(\theta = M_z, M_h, M_t, \alpha_s, \Delta\alpha_{had}^{(5)}, S, T)$$
$$= (\mathbf{y} - \boldsymbol{\mu}(\theta))^T V^{-1} (\mathbf{y} - \boldsymbol{\mu}(\theta))$$



Kinetic Mixing

- Add a new spontaneously broken U(1) symmetry
 - “Dark Photon”, “Dark Z”, “Z”
- Simple and well-motivated extension to the SM
- Renormalizable coupling between new field strength and hypercharge



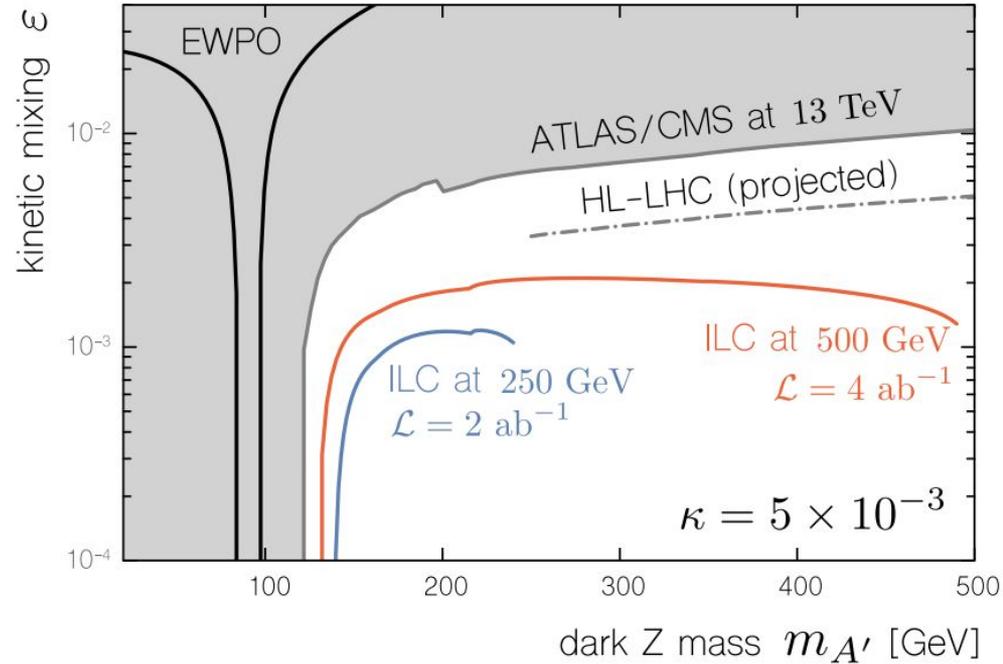
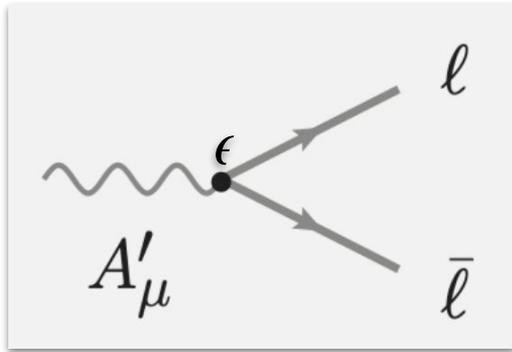
$$\mathcal{L} \supset -\frac{1}{4}\hat{B}_{\mu\nu}\hat{B}^{\mu\nu} - \frac{1}{4}\hat{Z}_{D\mu\nu}\hat{Z}_D^{\mu\nu} + \frac{\epsilon}{2\cos\theta_w}\hat{Z}_{D\mu\nu}\hat{B}^{\mu\nu} + \frac{1}{2}M_{D,0}^2\hat{Z}_{D\mu}\hat{Z}_D^\mu$$

$$\alpha S = 4\xi c_W^2 s_W \tan\chi \quad \alpha T = \xi^2 \left(\frac{M_{Z_2}^2}{M_{Z_1}^2} - 1 \right) + 2\xi s_W \tan\chi$$

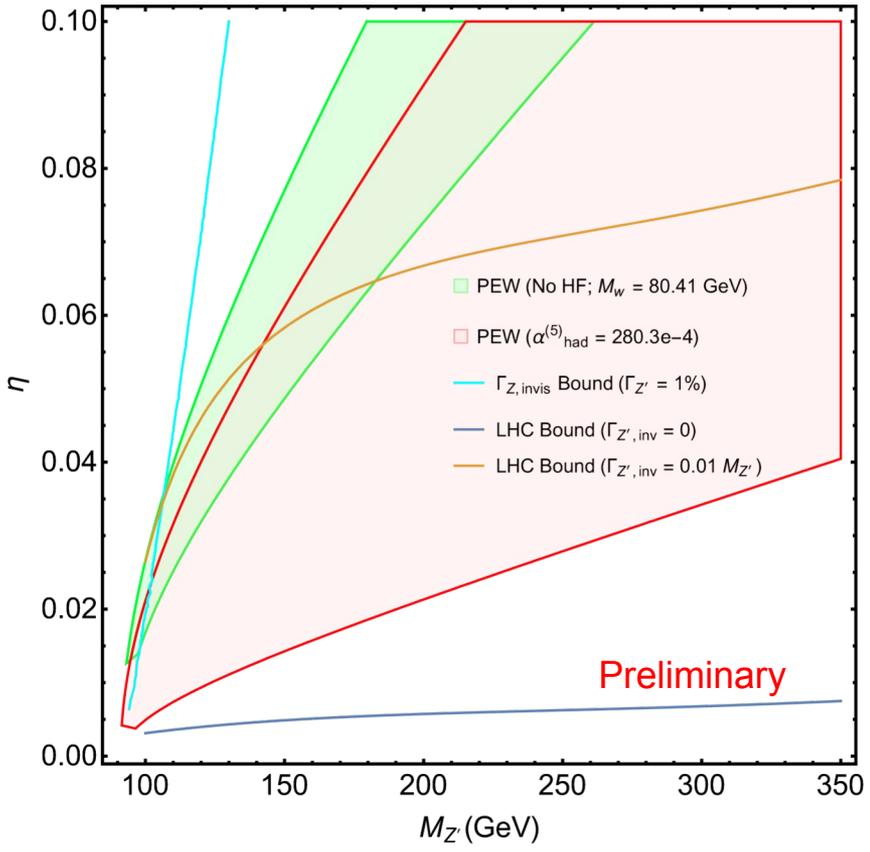
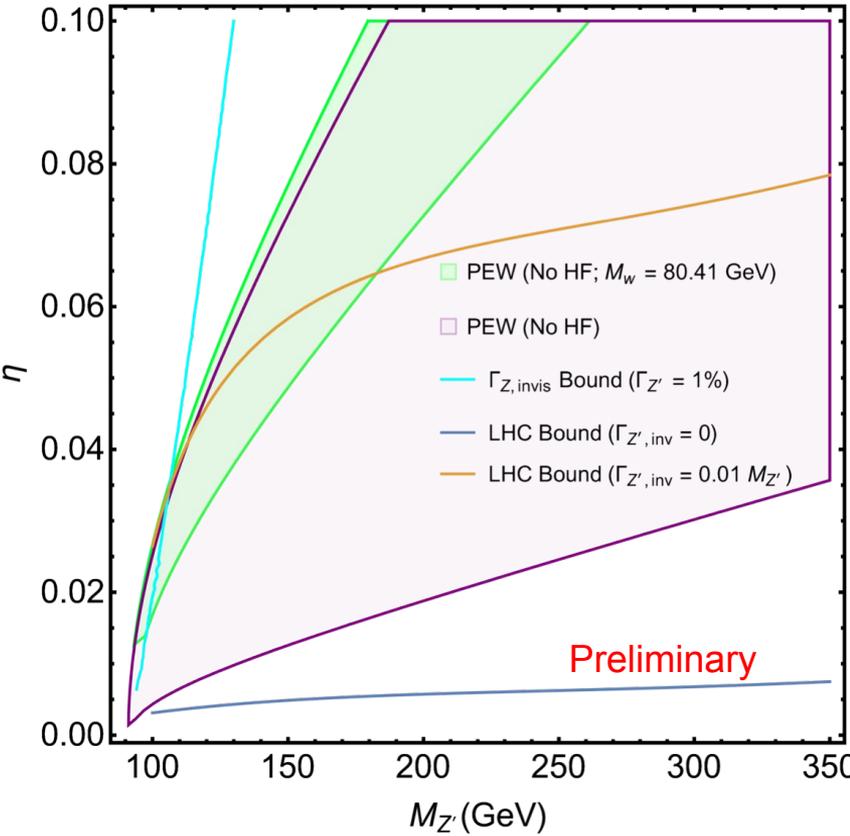
Kinetic Mixing - Collider Bounds and Width Effects

$$\sigma(q\bar{q} \rightarrow l^+l^-) \sim \epsilon^2$$

$$\Gamma(Z' \rightarrow \text{SM}) \sim \epsilon^2$$



Dark Photon Parameter Space



Summary

- Existing tensions motivate the exploration of different data combinations in the PEW fit
- S-T preferred region for different data combinations
 - Connection to kinetic mixing
- Complementarity between PEW and collider bounds
- Interesting target for future LHC searches

Thank You!

Backup Slides

Oblique Parameters

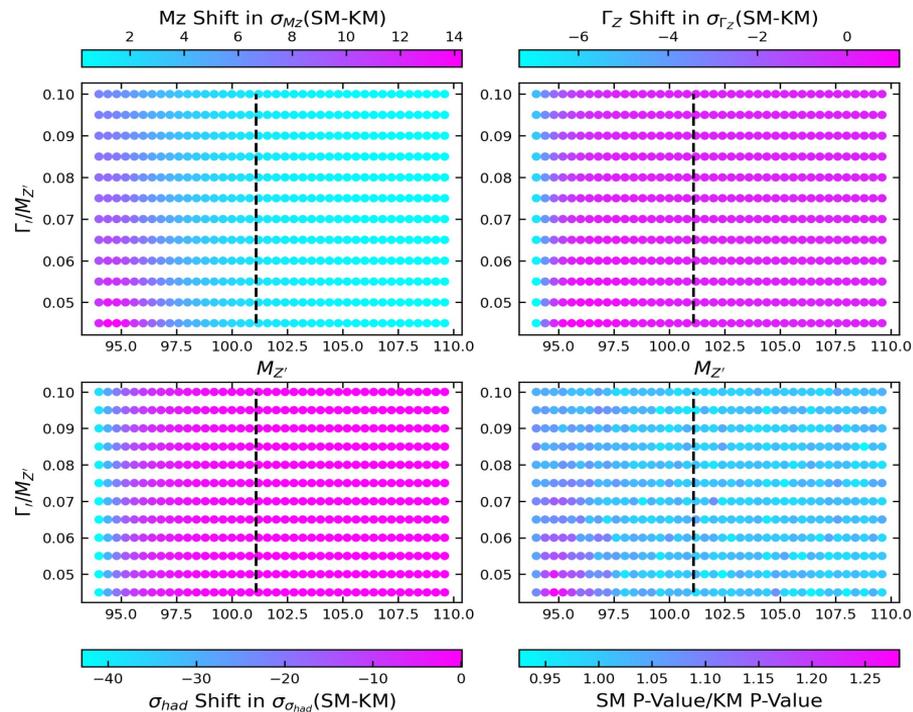
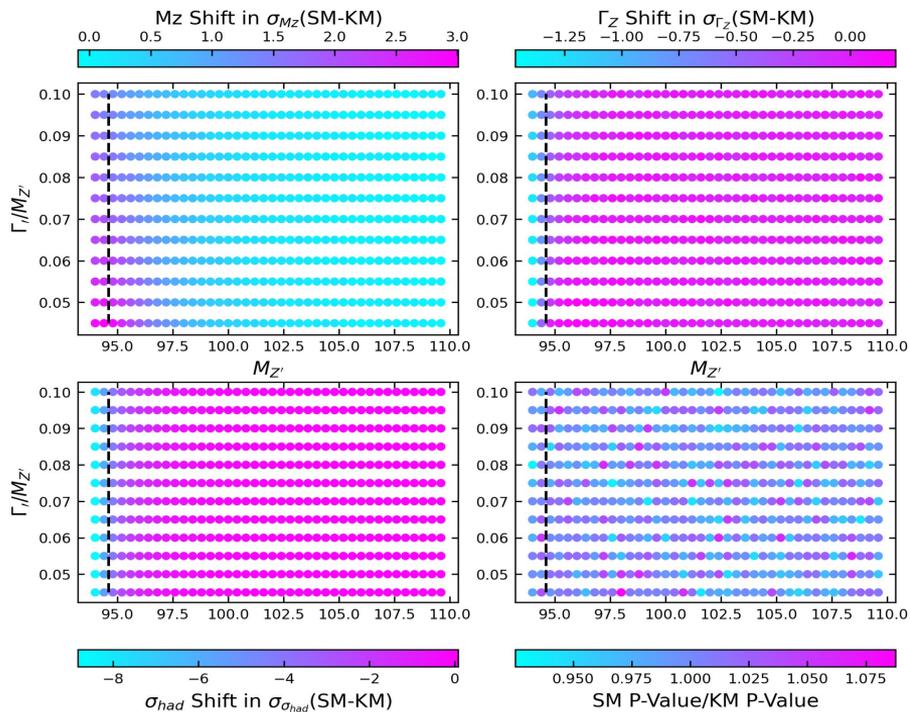
$$\begin{aligned}
 T &\equiv \frac{1}{\alpha(m_Z)} \left(\frac{\Pi_{WW}(0)}{m_W^2} - \frac{\Pi_{ZZ}(0)}{m_Z^2} \right) = \frac{\rho - 1}{\alpha(m_Z)} \\
 S &\equiv \frac{4c^2 s^2}{\alpha(m_Z)} \left(\frac{\Pi_{ZZ}(m_Z^2) - \Pi_{ZZ}(0)}{m_Z^2} - \frac{c^2 - s^2}{cs} \frac{\Pi_{Z\gamma}(m_Z^2)}{m_Z^2} - \frac{\Pi_{\gamma\gamma}(m_Z^2)}{m_Z^2} \right) \\
 U &\equiv \frac{4s^2}{\alpha(m_Z)} \left(\frac{\Pi_{WW}(m_W^2) - \Pi_{WW}(0)}{m_W^2} - \frac{c}{s} \frac{\Pi_{Z\gamma}(m_Z^2)}{m_Z^2} - \frac{\Pi_{\gamma\gamma}(m_Z^2)}{m_Z^2} \right) - S
 \end{aligned}$$

$$\begin{aligned}
 \mathcal{L}_{\text{eff}} &= \mathcal{L}_{SM}(\tilde{e}_i) + \hat{\mathcal{L}}_{\text{new}}, & \alpha S &= 4s_w^2 c_w^2 \left(A - C - \frac{c_w^2 - s_w^2}{c_w s_w} G \right), \\
 \hat{\mathcal{L}}_{\text{new}} &= -\frac{A}{4} \hat{F}_{\mu\nu} \hat{F}^{\mu\nu} - \frac{B}{2} \hat{W}_{\mu\nu}^\dagger \hat{W}^{\mu\nu} - \frac{C}{4} \hat{Z}_{\mu\nu} \hat{Z}^{\mu\nu} + \frac{G}{2} \hat{F}_{\mu\nu} \hat{Z}^{\mu\nu}, & \alpha T &= w - z, \\
 & - w \tilde{m}_W^2 \hat{W}_\mu^\dagger \hat{W}^\mu - \frac{z}{2} \tilde{m}_Z^2 \hat{Z}_\mu \hat{Z}^\mu. & \alpha U &= 4s_w^4 \left(A - \frac{1}{s_w^2} B + \frac{c_w^2}{s_w^2} C - 2 \frac{c_w}{s_w} G \right).
 \end{aligned}$$

What about interference or effects on Z-pole fit?

$\eta = 0.018$

$\eta = 0.044$



What about interference or effects on Z-pole fit?

