

# **SENSITIVITY OF COSMOLOGICAL DATA TO THE NEUTRINO MASS**

*from*

## **HIERARCHY**

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**TeVPA 2017, Columbus - OH**

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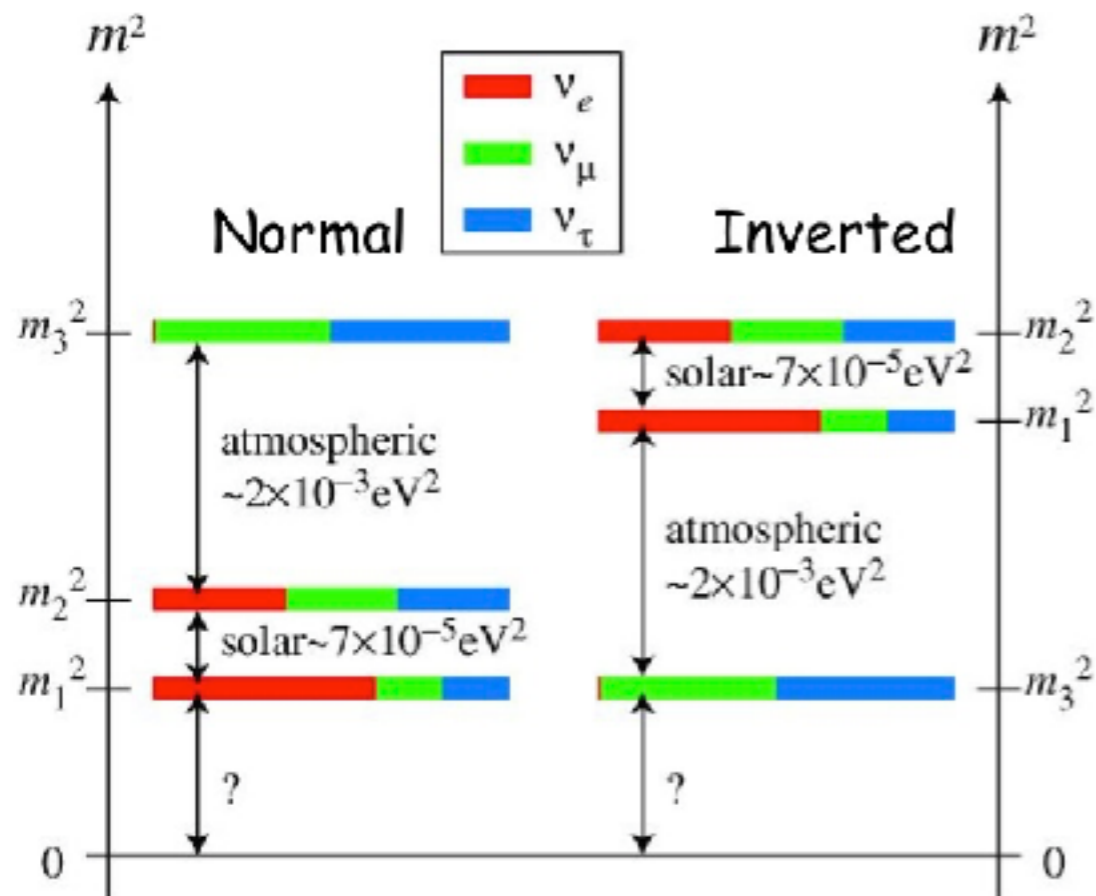
**based on Gerbino,Lattanzi,Mena,Freese,2017**

# What we know, from the outside

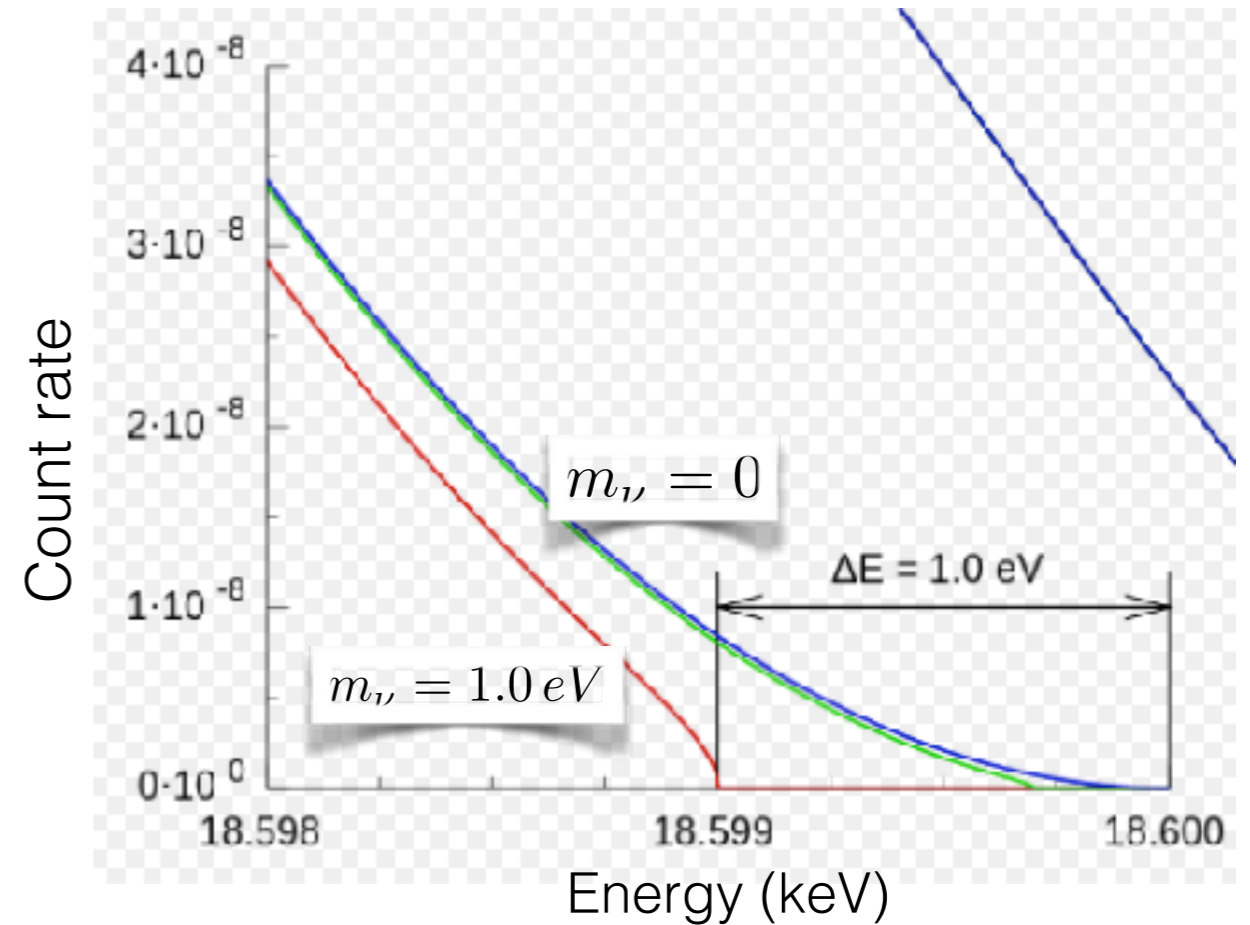
How do they behave?

Neutrinos oscillate, so they are massive

$$0.06 \text{ eV} < \Sigma m_\nu < 6 \text{ eV}$$

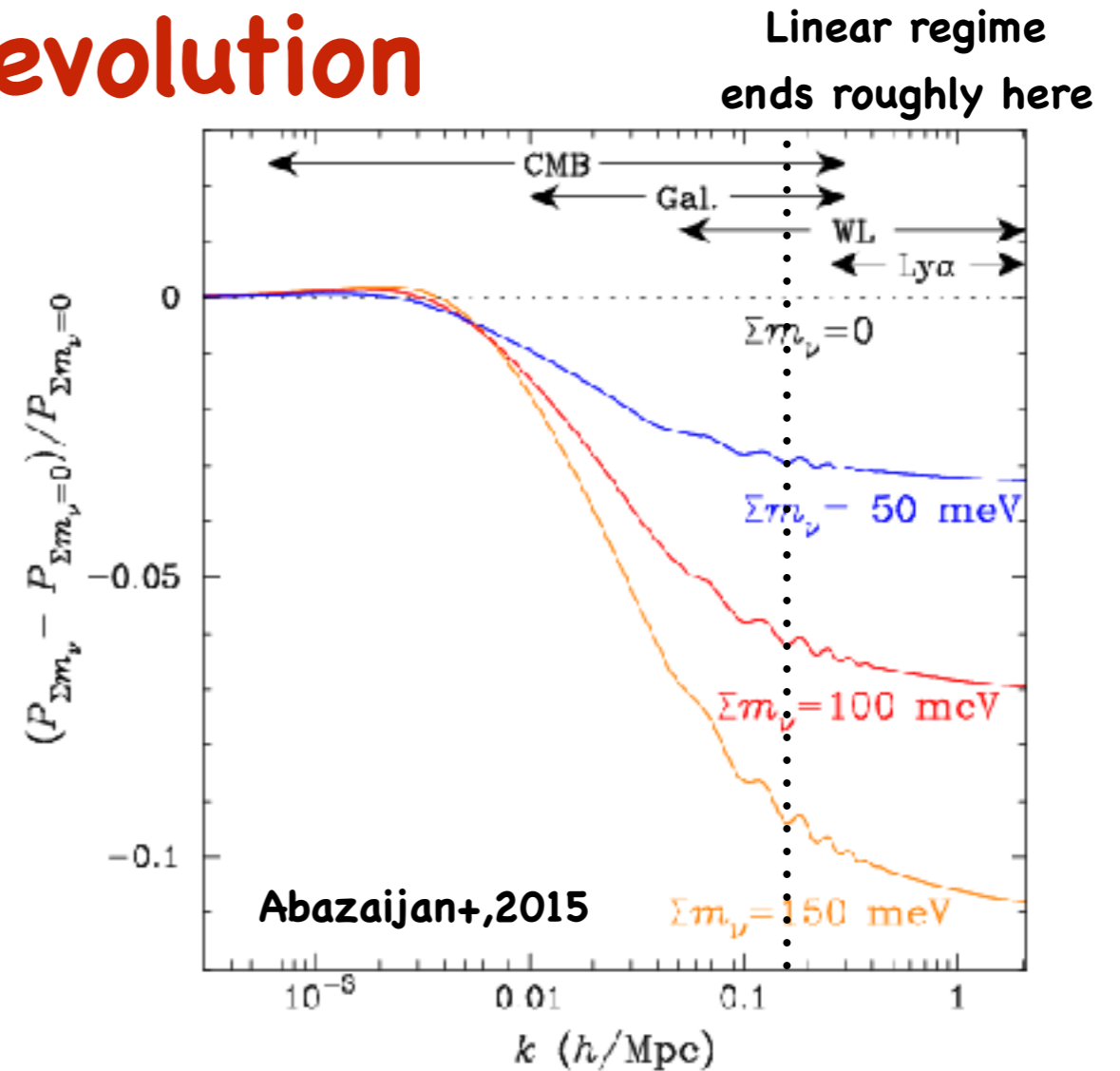
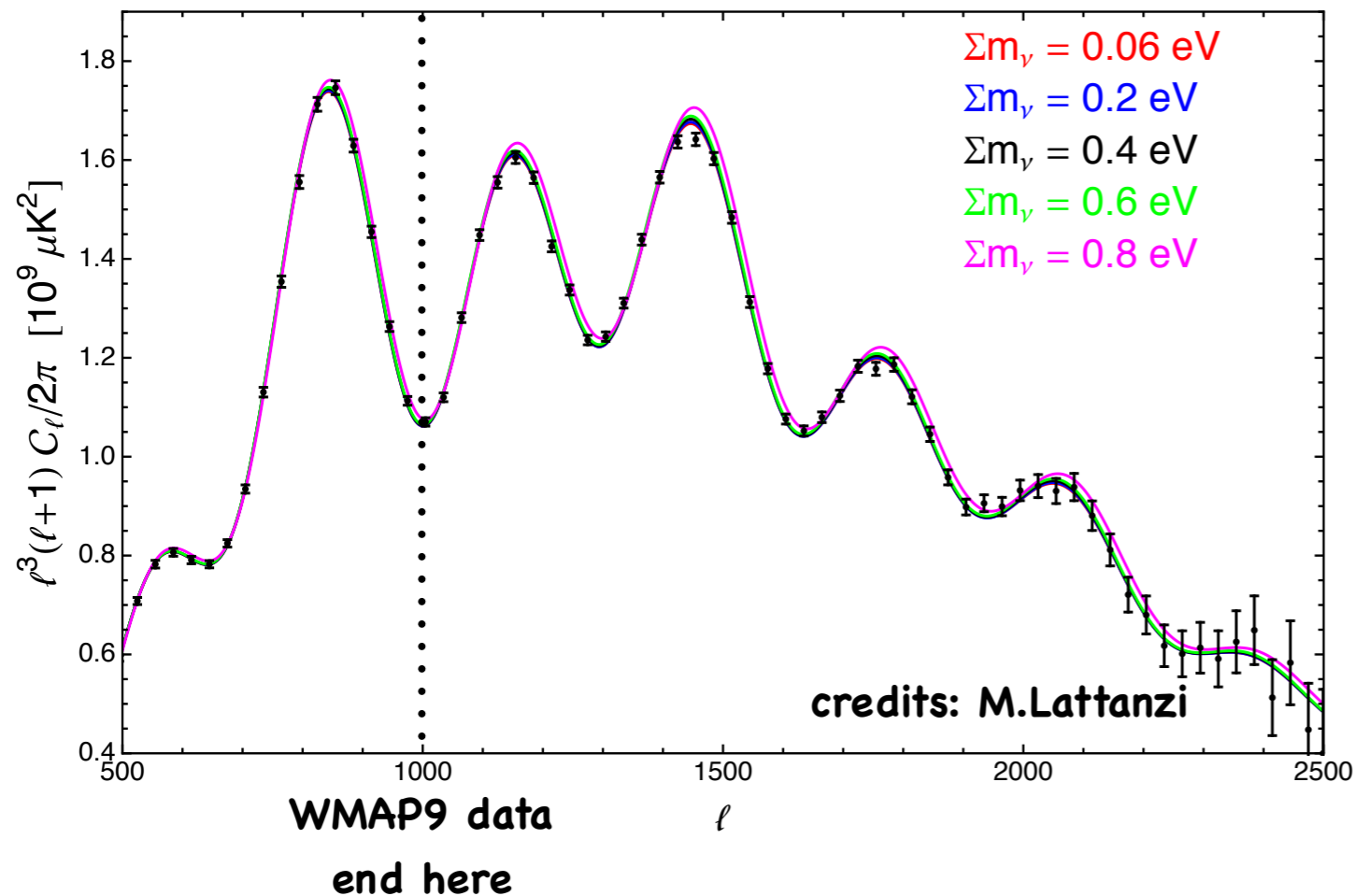


**Lower bound**  
from oscillation experiments



**Upper bound**  
from kinematic measurements

# Massive neutrinos alter background and perturbation evolution



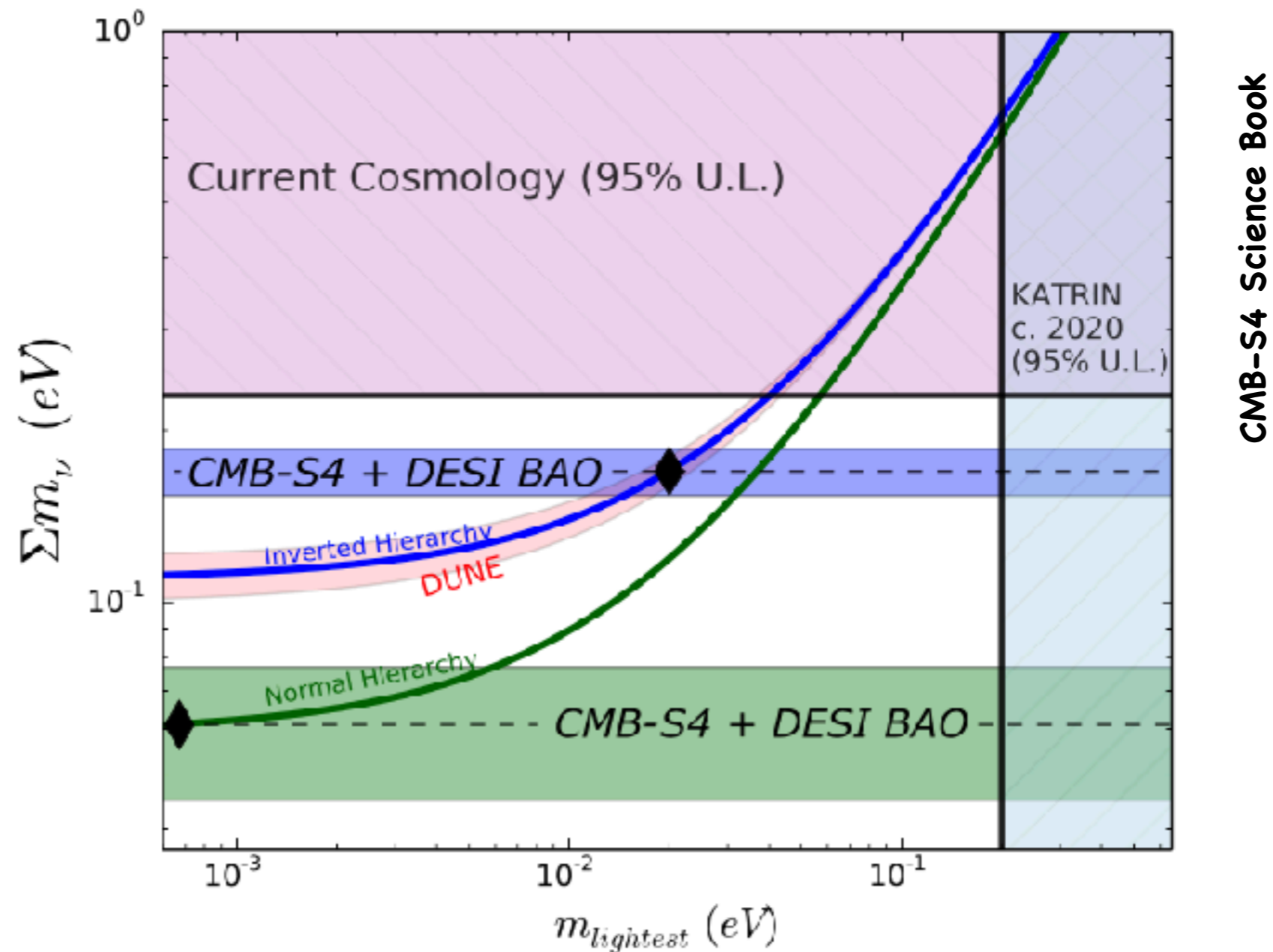
**3 active families, sub-eV masses**

**Relativistic at early times, non-relativistic today**

**(Almost) peculiar effects on cosmological observables**

**See M. Lattanzi & A. van Engelen's talks**

# Joint constraints on $M_{\nu}$ - future

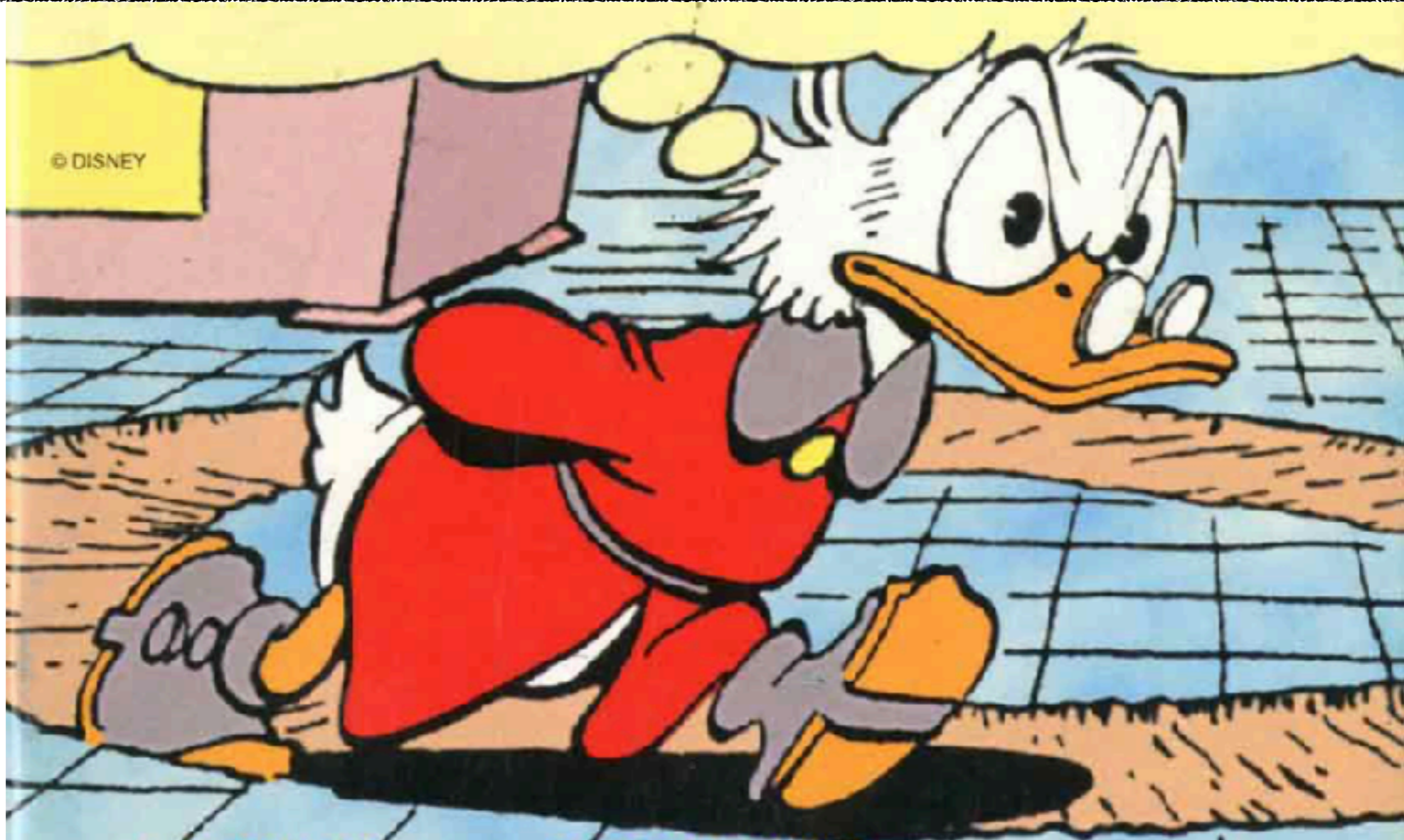


CMB-S4 Science Book

**~3sigma detection**

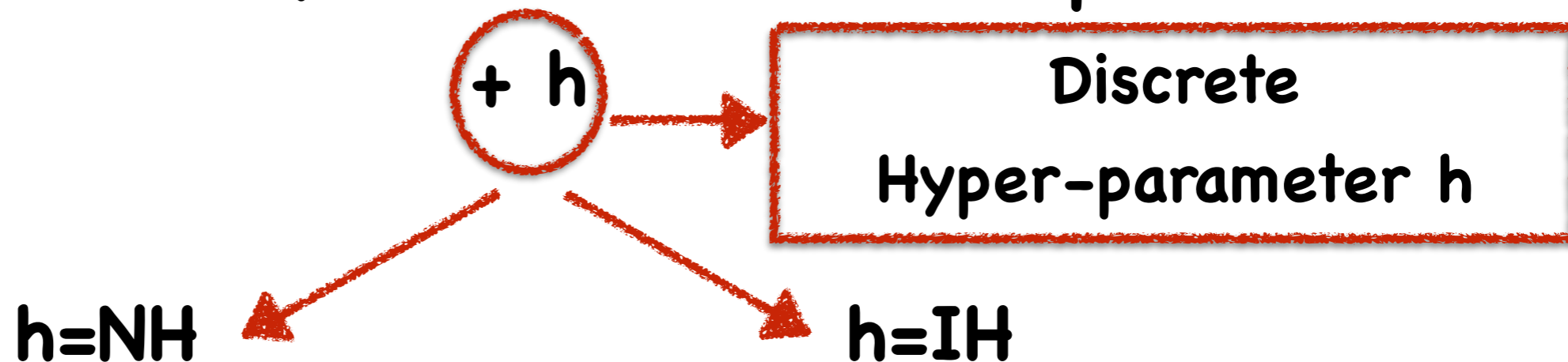
**in the minimal mass scenario with S4 surveys**

- Will we be able to discriminate the hierarchy?
- Can we provide a statistically robust answer?



# The proposed method

Bayesian MCMC, Mnu + other cosmo params



## NORMAL HIERARCHY

$$m_{\nu,1} = m_{\text{light}}$$

$$m_{\nu,2} = \sqrt{m_1^2 + \Delta m_{12}^2}$$

$$m_{\nu,3} = \sqrt{m_1^2 + \Delta m_{13}^2}$$

## INVERTED HIERARCHY

$$m_{\nu,3} = m_{\text{light}}$$

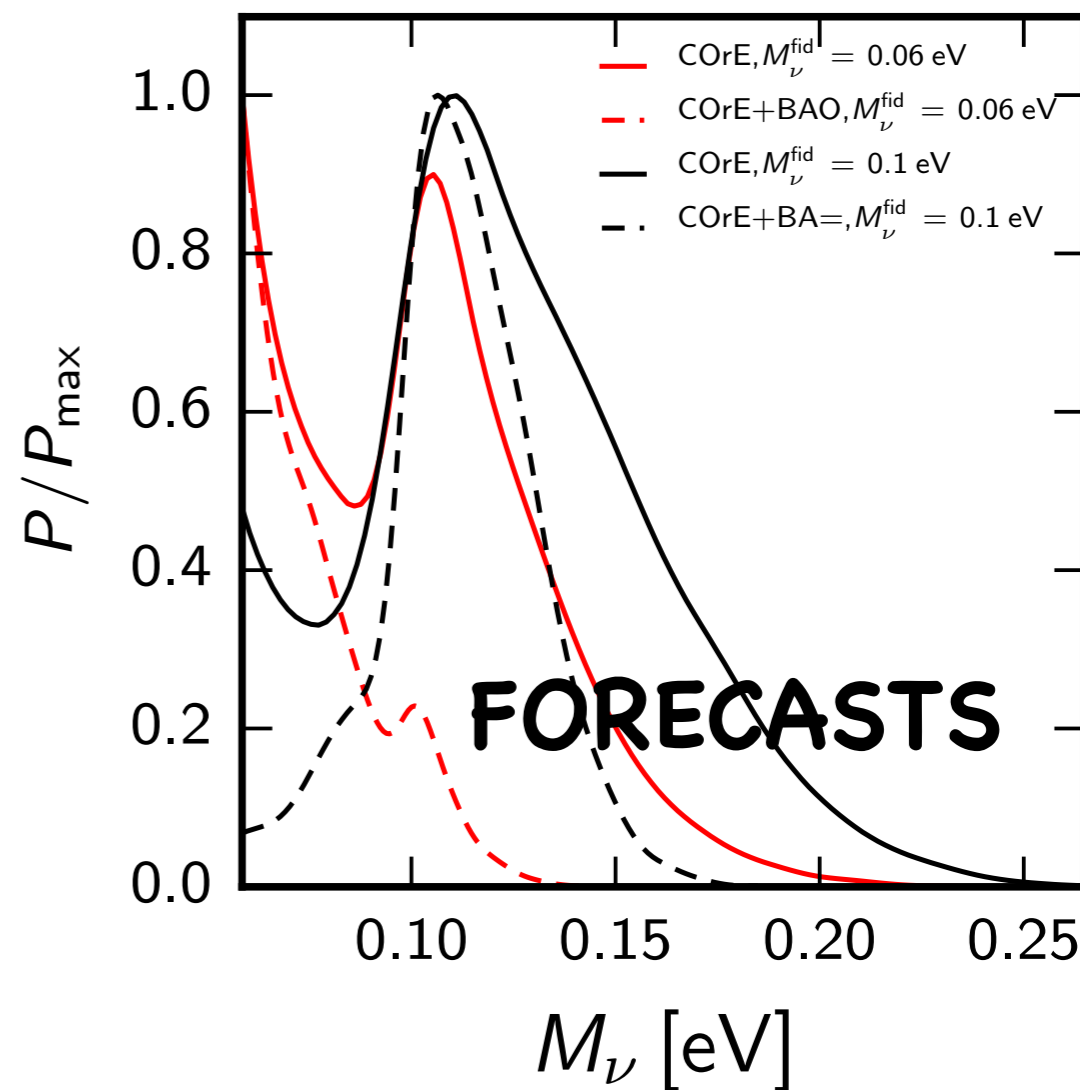
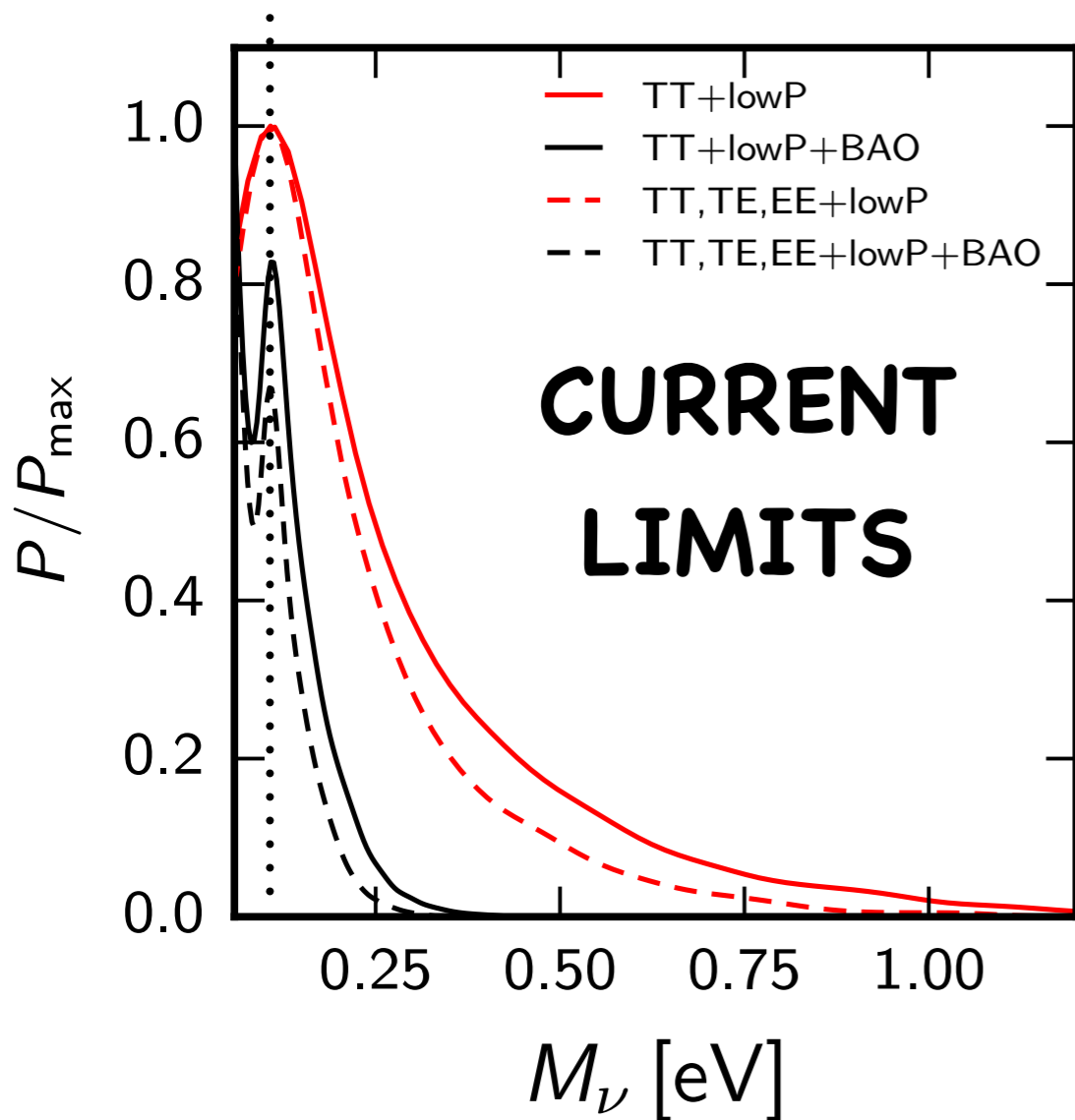
$$m_{\nu,1} = \sqrt{m_3^2 + \Delta m_{13}^2}$$

$$m_{\nu,2} = \sqrt{m_1^2 + \Delta m_{12}^2}$$

Advantages:

- neutrinos modelled with exact mass spectrum
- information from oscillations taken into account
- quantifies sensitivity to the hierarchy
- takes into account uncertainties related to the hierarchy

# Sensitivity to the hierarchy



$$\mathcal{P}(h = NH) : \mathcal{P}(h = IH)$$

..... **3:2**

$$\mathcal{P}(h = NH) : \mathcal{P}(h = IH)$$

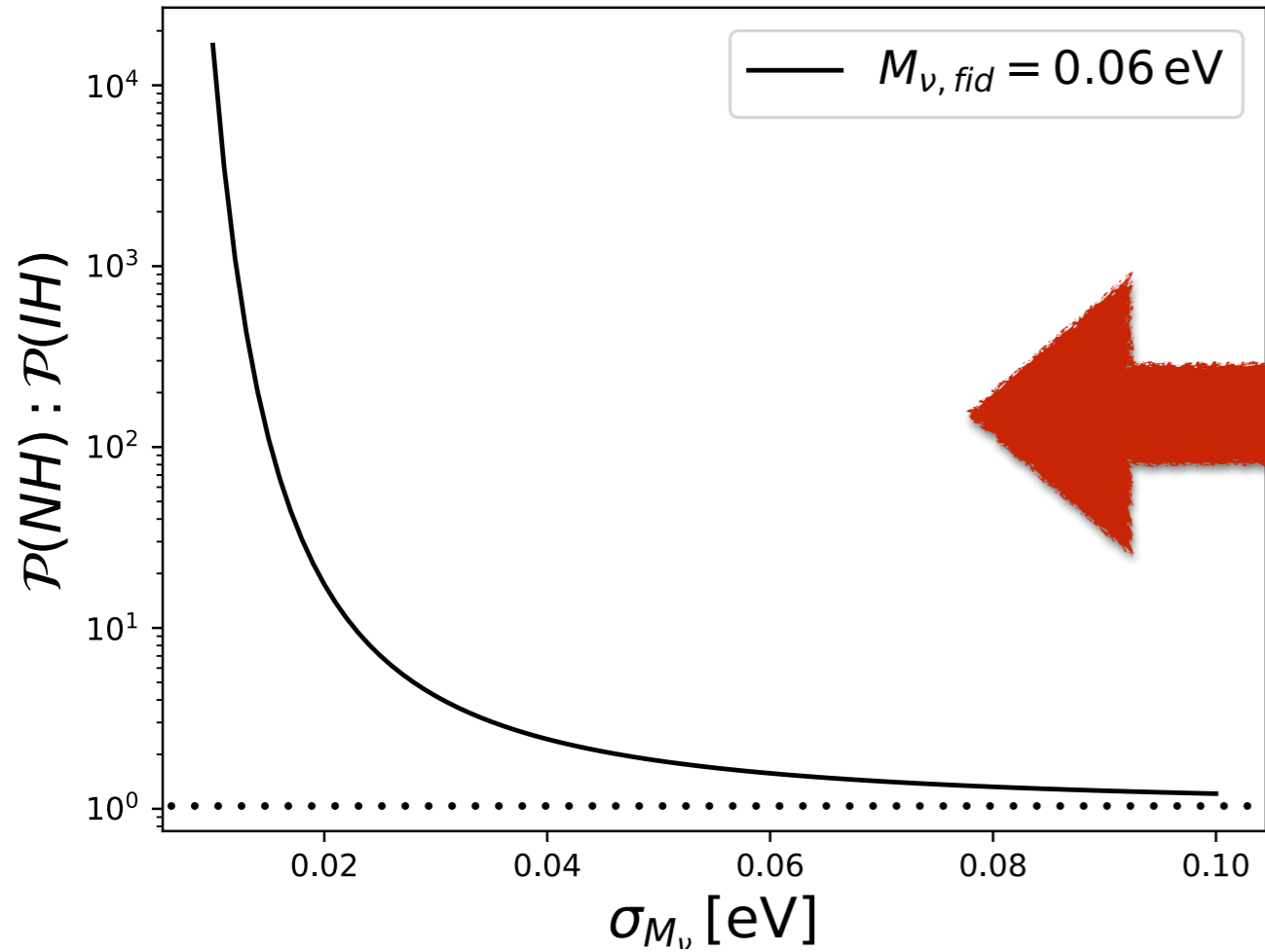
..... **0.06eV mass -> 9:1**

..... **0.1eV mass -> 1:1**

Gerbino, Lattanzi, Mena, Freese 2017

See also Hannestad&Schwetz 2016,  
Couchot et al 2017, Capozzi et al 2017

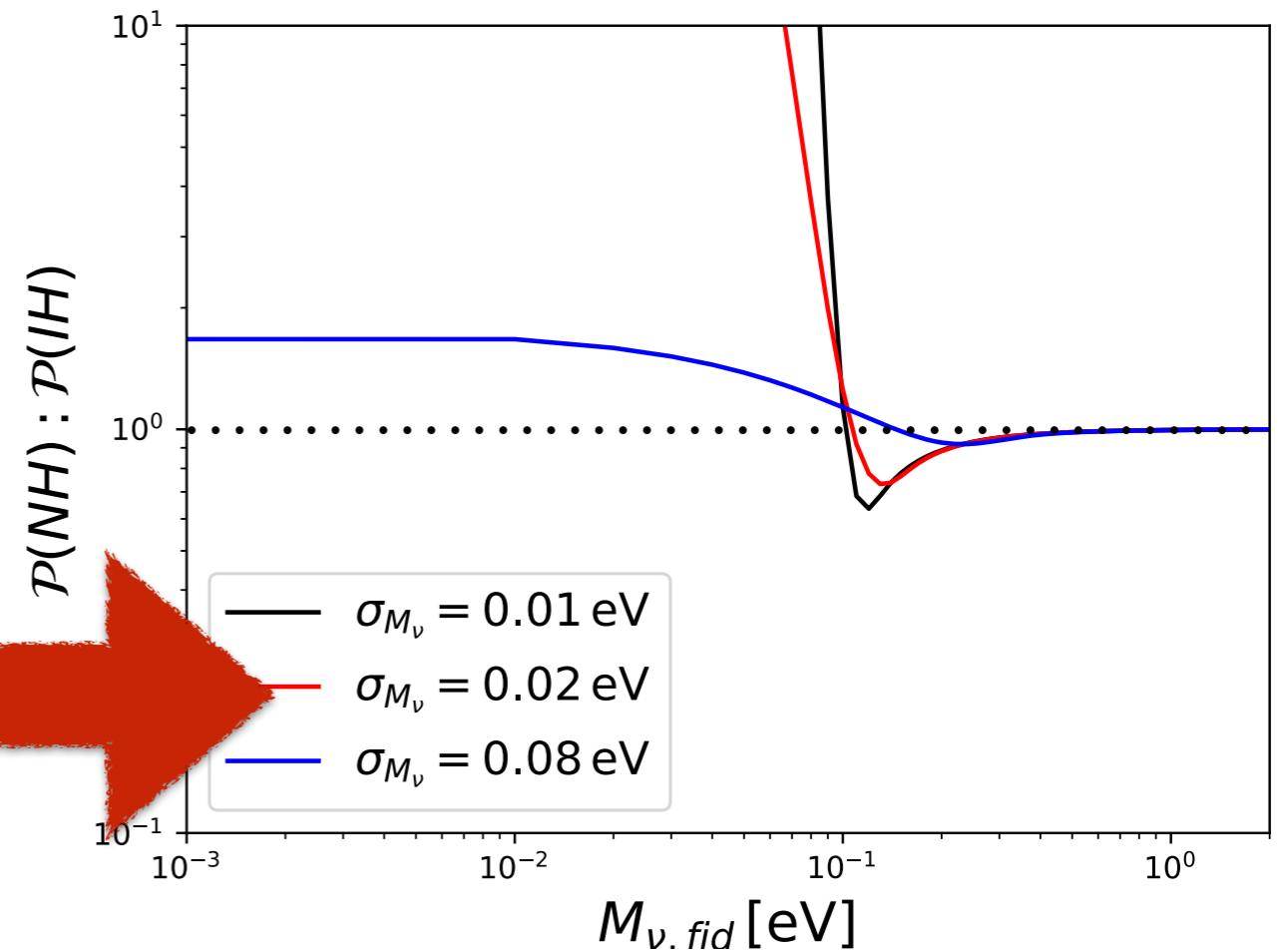
# Sensitivity to the hierarchy



**Fixed fiducial**  
**As sigma increases,**  
**NH is as likely as IH**

**Fixed sigma**

**NH favoured for  $M_{\nu} < 0.1$  eV,**  
**IH favoured for  $M_{\nu} \sim 0.1$  eV,**  
**no preference for  $M_{\nu} > 0.1$  eV**





# CONCLUSIONS

- Tight bounds on neutrino mass from cosmology
- Inverted hierarchy in trouble: how much?
- By introducing an hyper-parameter we can: 1) easily account for exact neutrino mass spectra; 2) quantify sensitivity to the hierarchy; 3) take into account uncertainty due to imperfect knowledge of the hierarchy
- NH favoured 5:3 by current data
- NH favoured 10:1 by future measurements, if the mass is minimal
- **sensitivity driven by prior choice**

**BACK-UP**

# The standard method

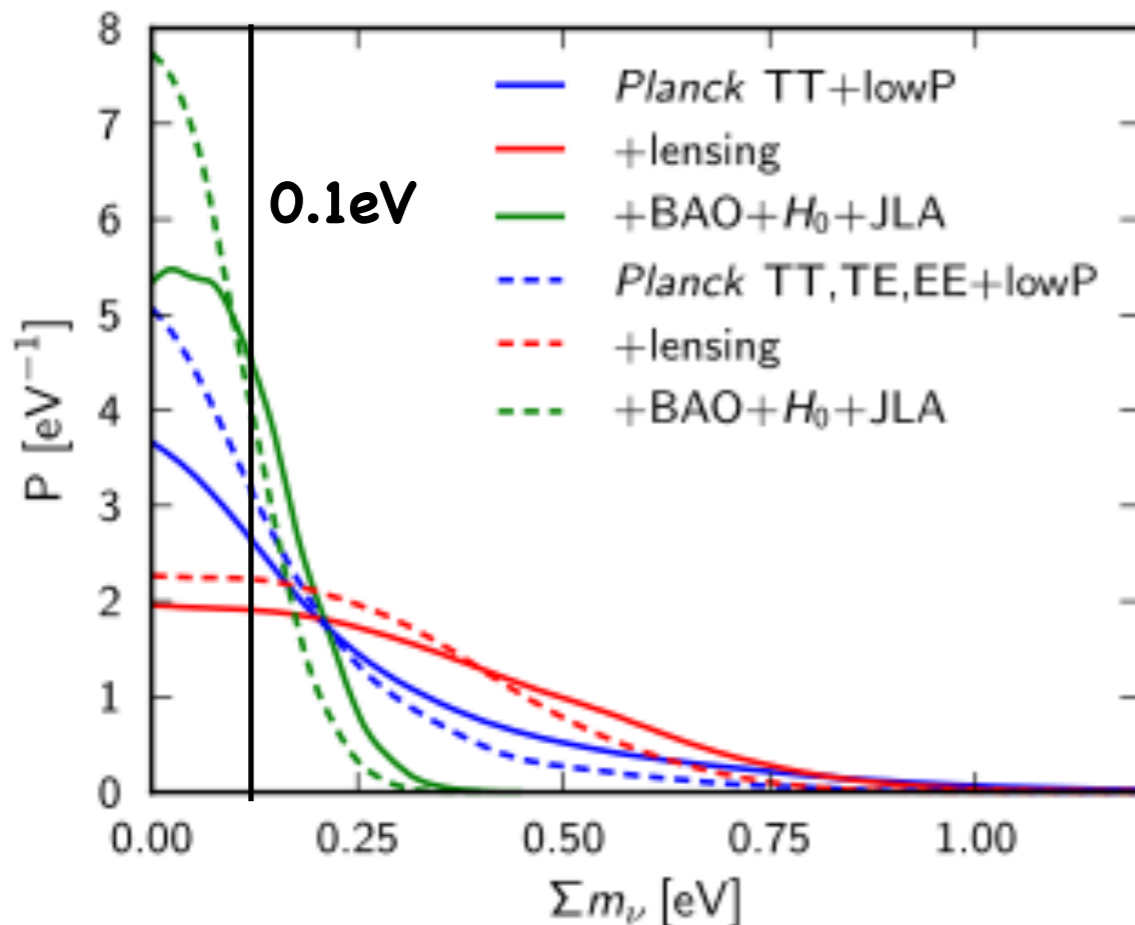
$M_{\nu}$

(+other cosmological parameters)



$$m_{\nu,i} = M_{\nu}/3, \quad i = 1, 2, 3$$

Degenerate spectrum:



Different authors obtain upper bounds from current data approaching the “critical” value of 0.1 eV. These results suggest that IH starts to get under pressure from cosmology. [...] Such a claim should be based on a proper statistical analysis. The question to be answered is, whether the hypothesis of IH can be rejected with some confidence against NH.

(Hannestad&Schwetz,2016)

# Joint constraints on $M_{\nu}$ - present

Vagnozzi, Giusarma, Mena, Freese, MG, Ho, Lattanzi 2017

$$M_{\nu} \equiv \sum_i m_{\nu,i}$$

Planck TT+lowP+BAO:

$M_{\nu} < 0.2 \text{ eV}$

Planck TT+lowP+Pk:

$M_{\nu} < 0.3 \text{ eV}$

Planck TT+lowP+BAO+Pk:

$M_{\nu} < 0.25 \text{ eV}$

Compilation of CMB and LSS data

95% CL

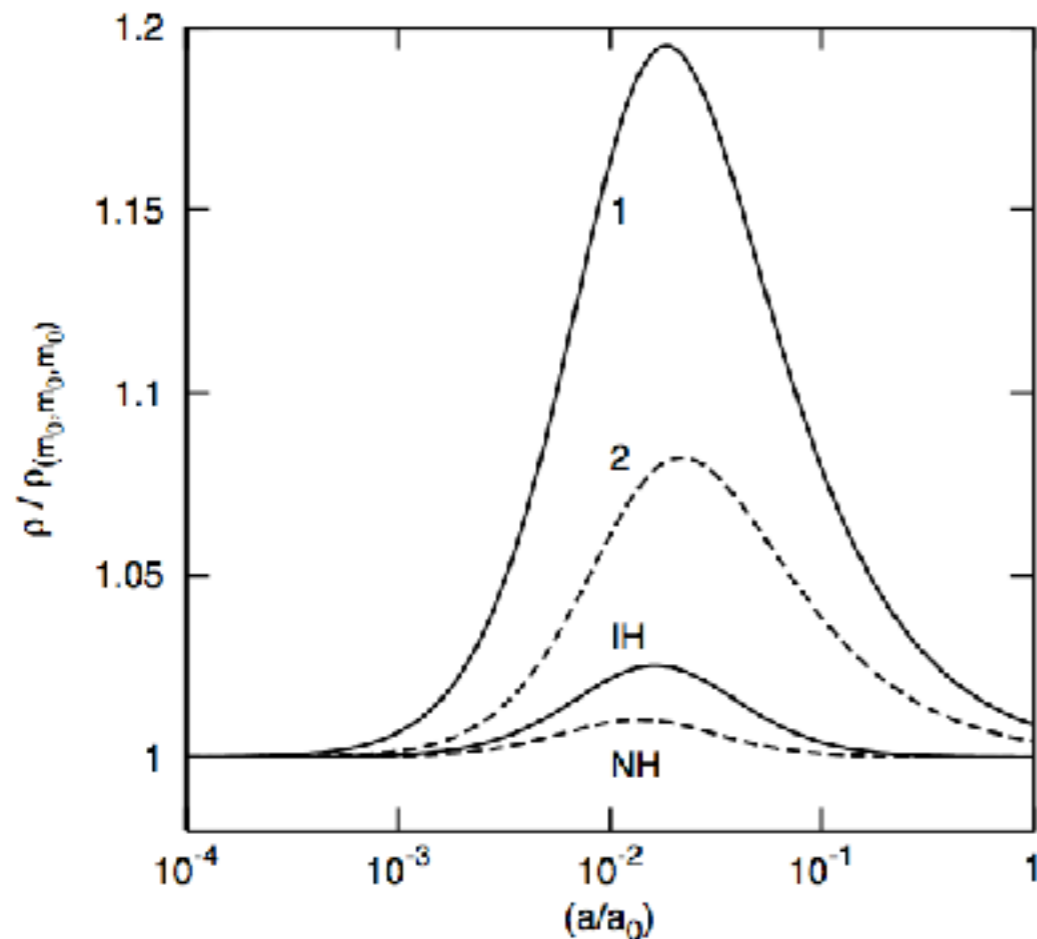
major improvement with a better measurement of the optical depth and/or use of CMB small scale polarisation:  
wait for Planck legacy release!

Take home message: tight, yet robust bounds

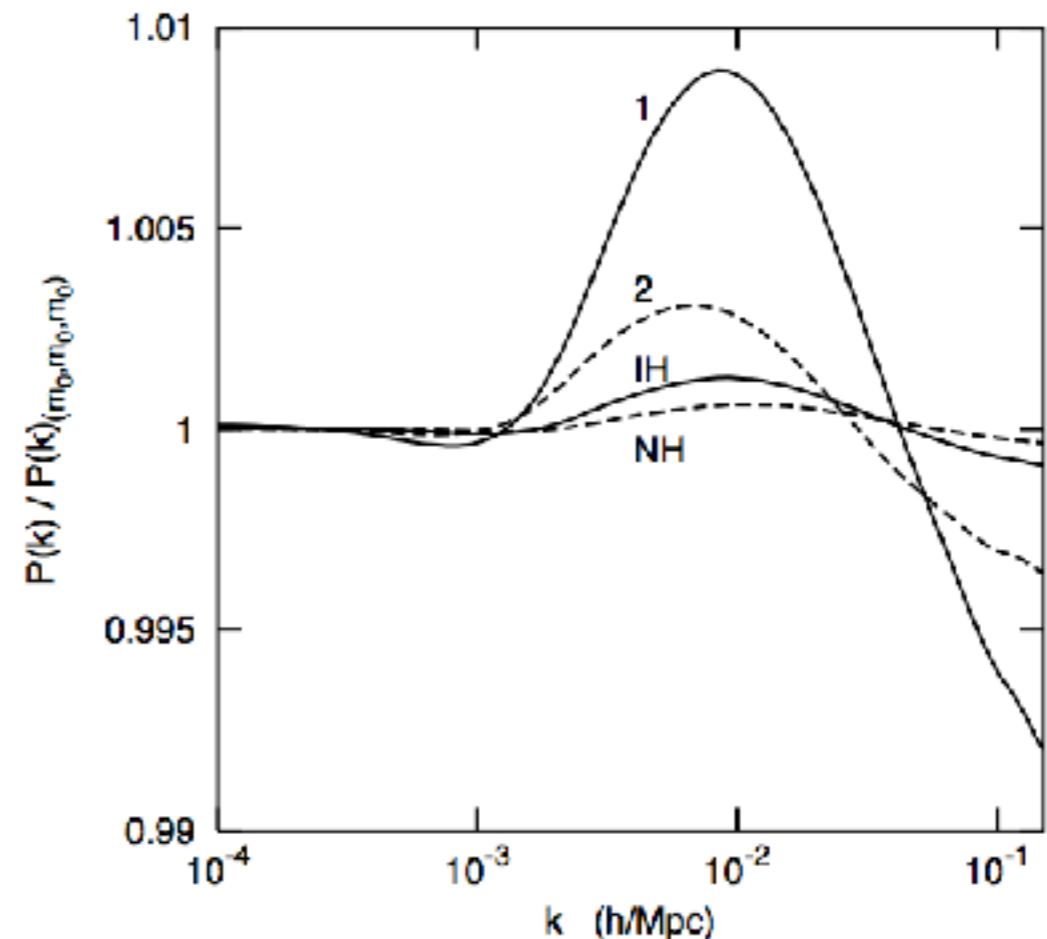
# Sensitivity to the hierarchy

Physical effects due to different distribution of the sum of the masses for the 2 hierarchies

Total nu energy density



Matter power spectrum

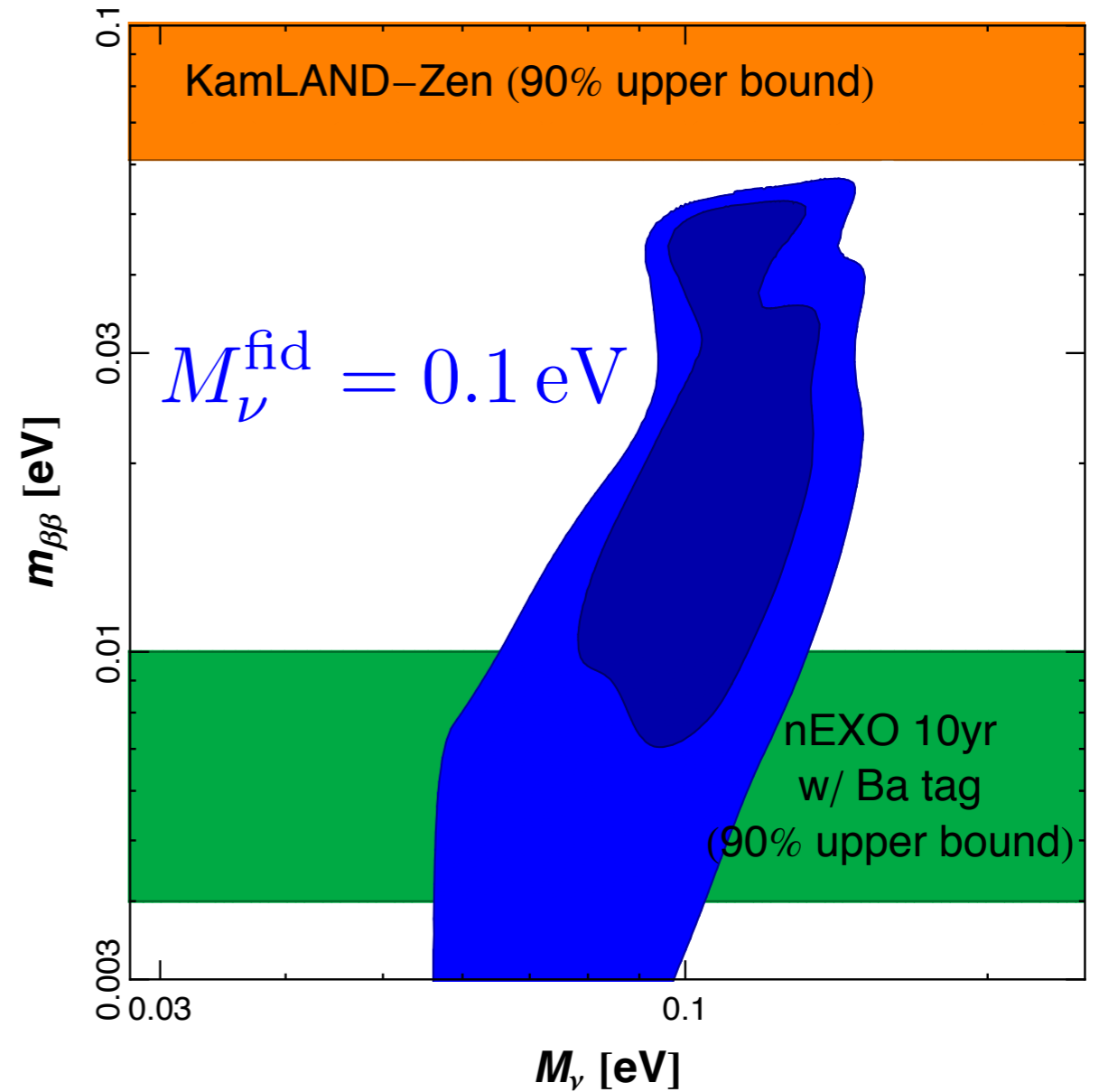
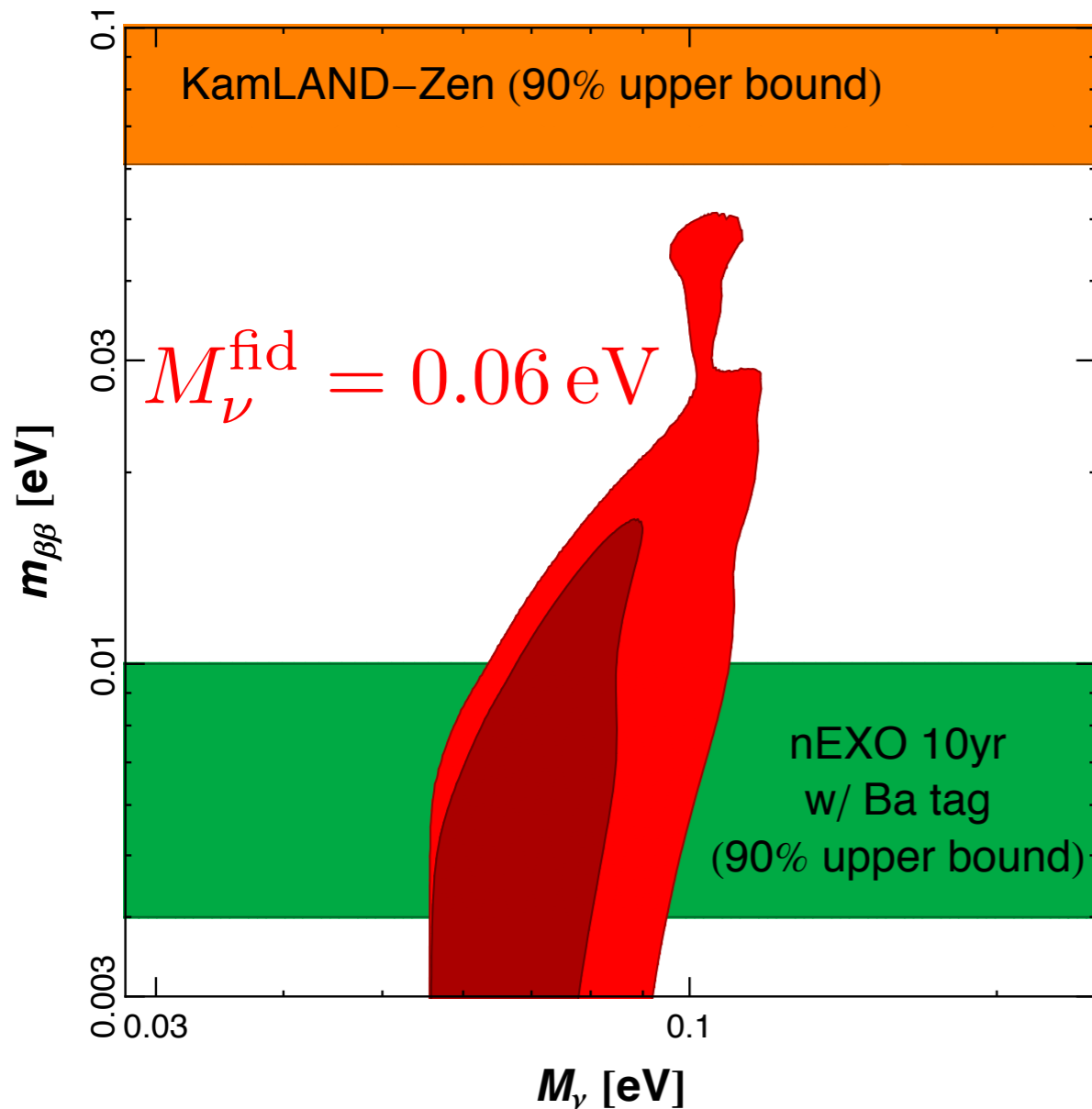


Lesgourgues&Pastor, 2006

Are current (and future) data sensitive to these effects?

How much?

# Sensitivity to the hierarchy



Gerbino, Lattanzi, Mena, Freese 2017

If  $M_\nu = 0.1 \text{ eV}$ ,  $\sigma(m_{\beta\beta}) \sim 10 \text{ meV}$  could guarantee

On2b measurement

On2b could in turn help unravel the hierarchy (wip, extending the results in Gerbino+2015 in the hierarchical bayesian context)