

Dwarf Galaxy Population of a Nearby Star-Forming Galaxy and Implications for Dark Matter

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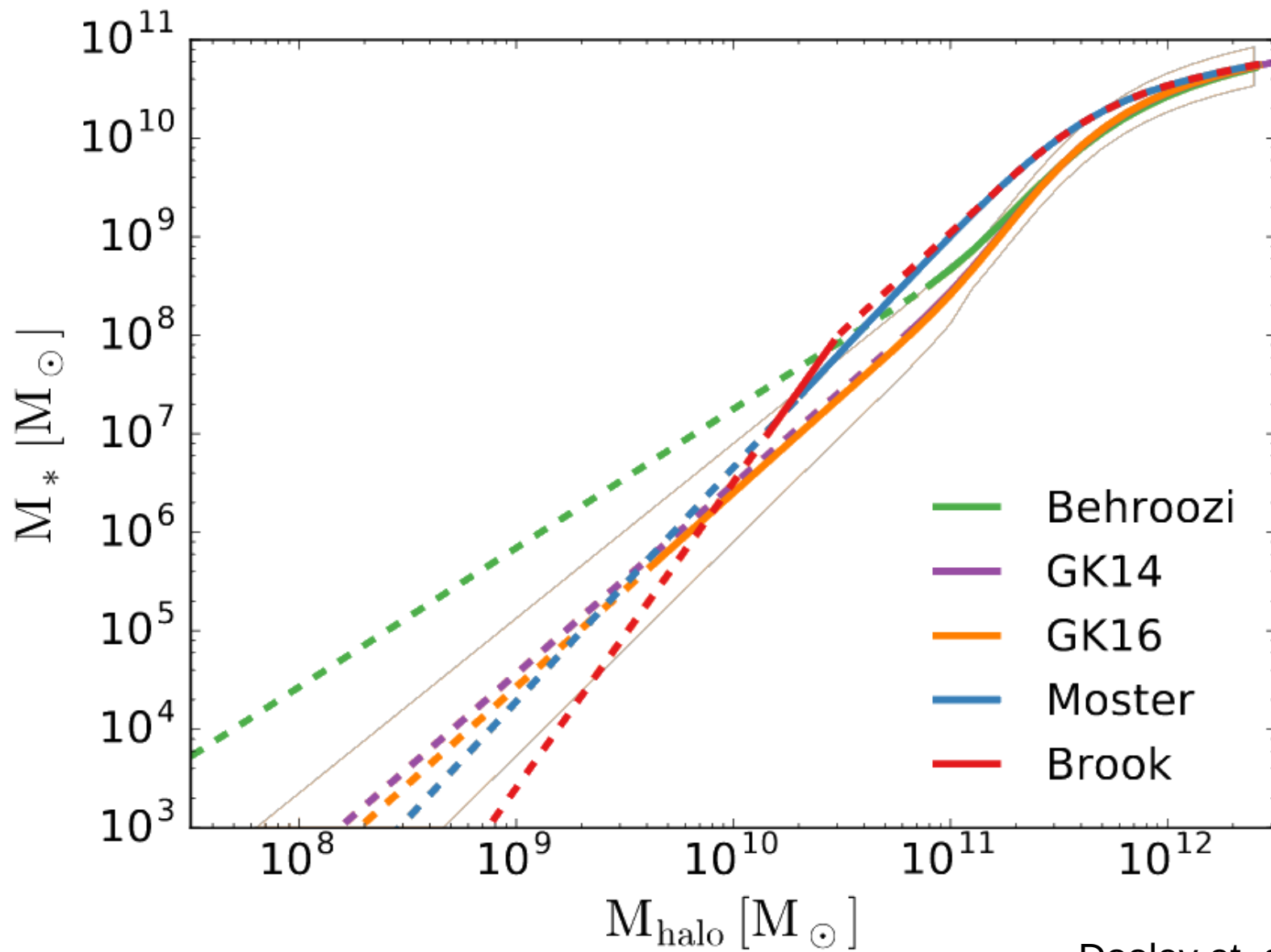


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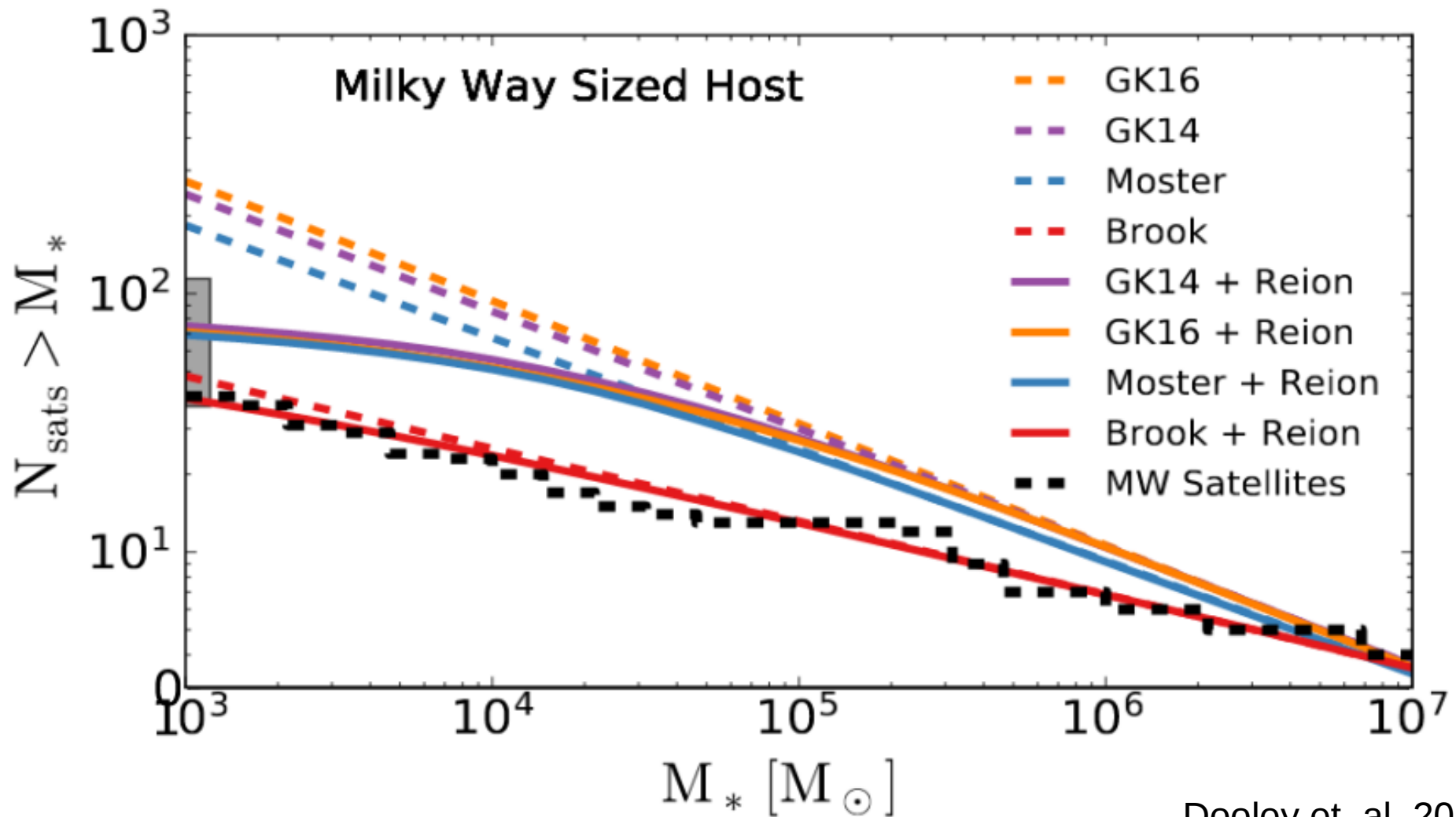
Image credit: NASA, ESA, and T. Brown and J. Tumlinson (STScI).

Stellar Mass to DM Halo Mass Relationship



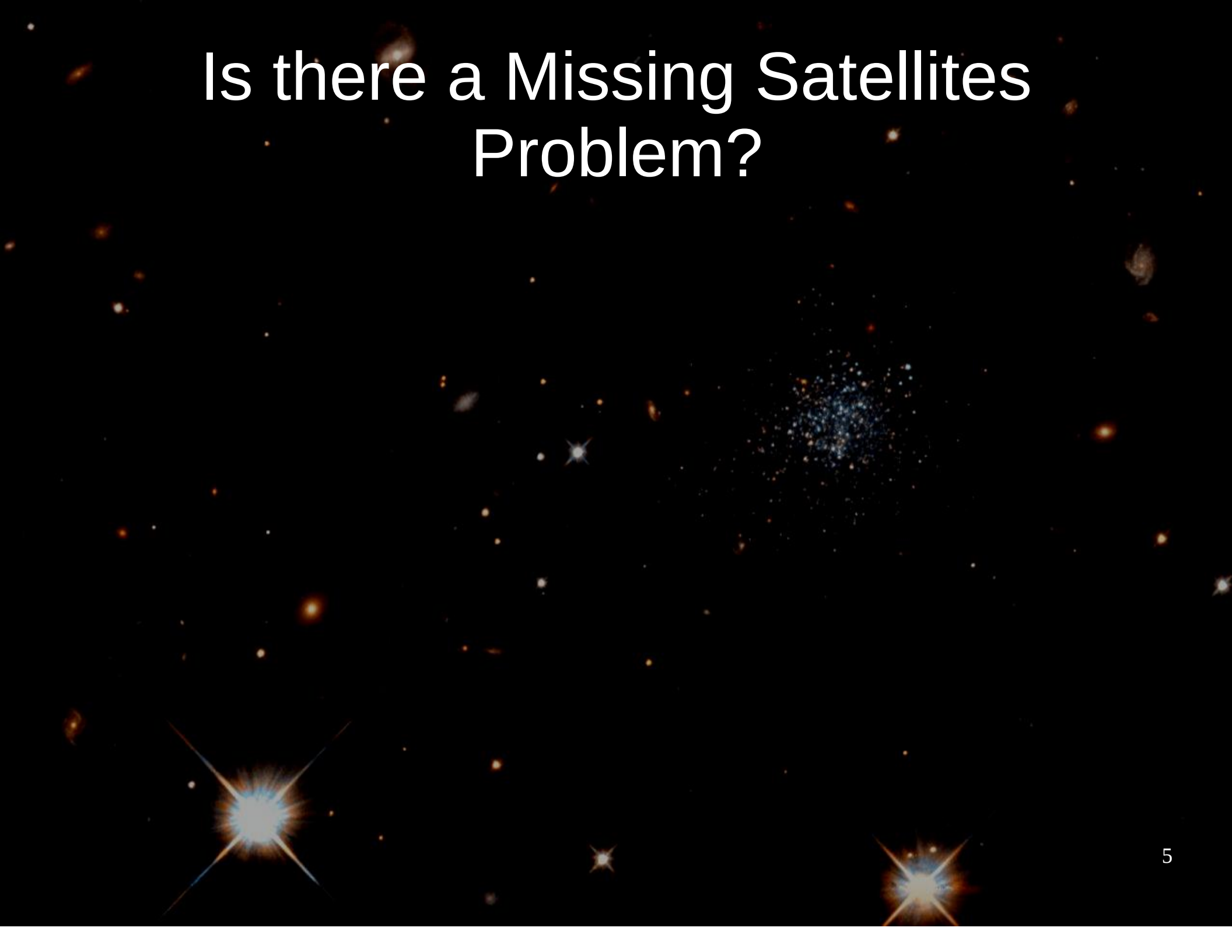
Dooley et. al. 2017

What We See



Dooley et. al. 2017

Is there a Missing Satellites Problem?



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- Internal effects

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- Environmental effects

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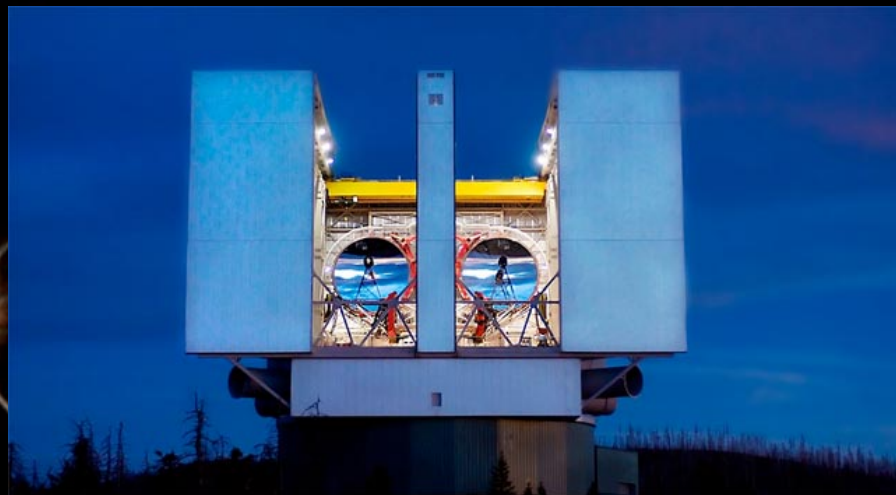
Is there a Missing Satellites Problem?

- Internal effects
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- Global effects
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We need more systems!

Our Survey

- “Search for Failed Supernovae with the LBT” dataset, Kochanek 2008
- 27 star-forming galaxies within 10 Mpc
- Range in luminosity from LMC to MW size
- R band magnitude limits fainter than 26 mag
- Probing 1% to 20% of virial volumes



N628/M74

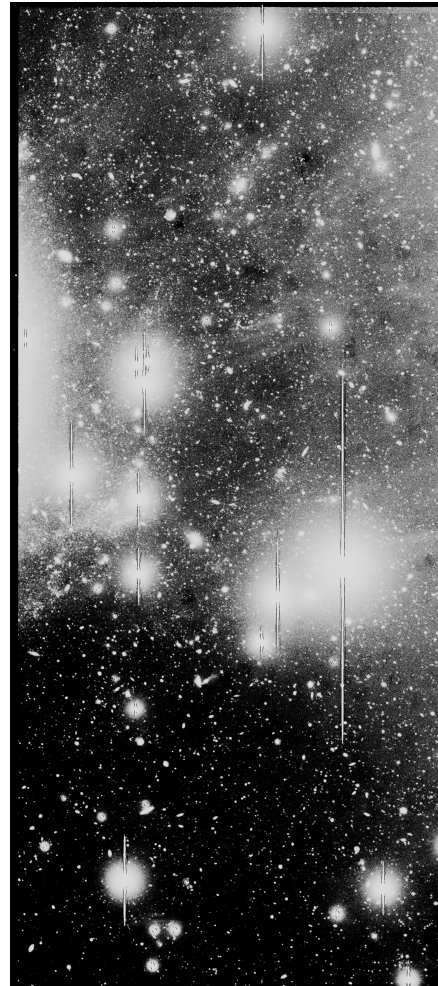
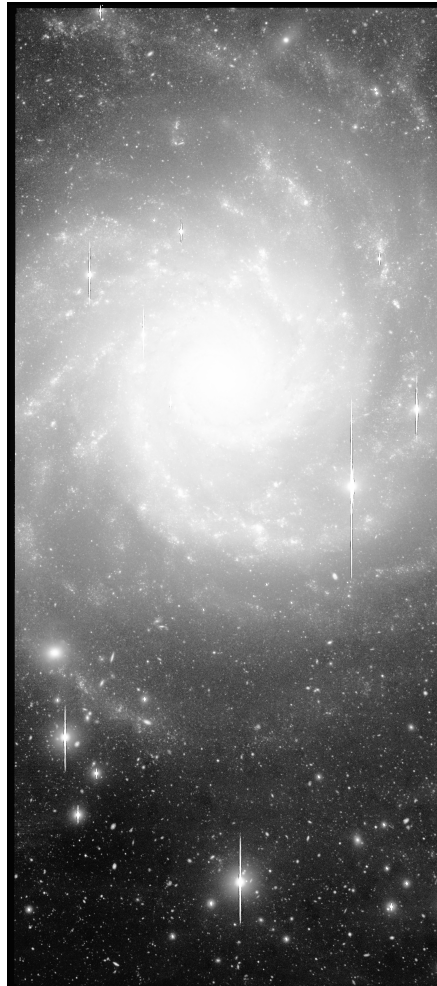
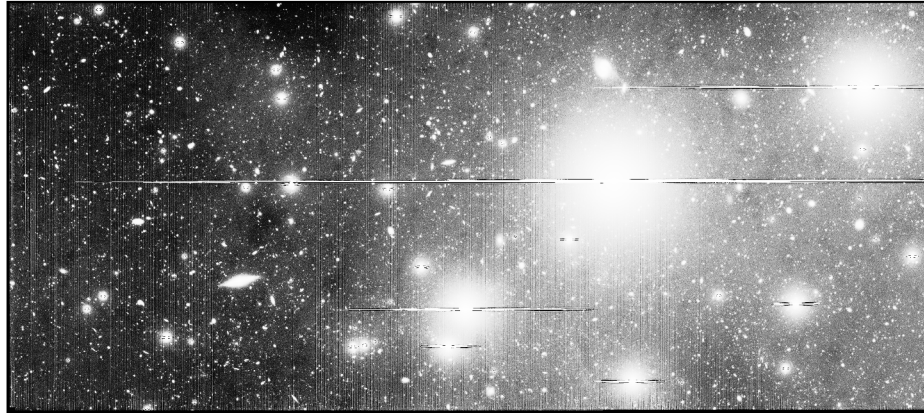
$$M_* \approx 1.3 \times 10^{10} M_\odot$$

~1/4 the stellar mass
of the Milky Way

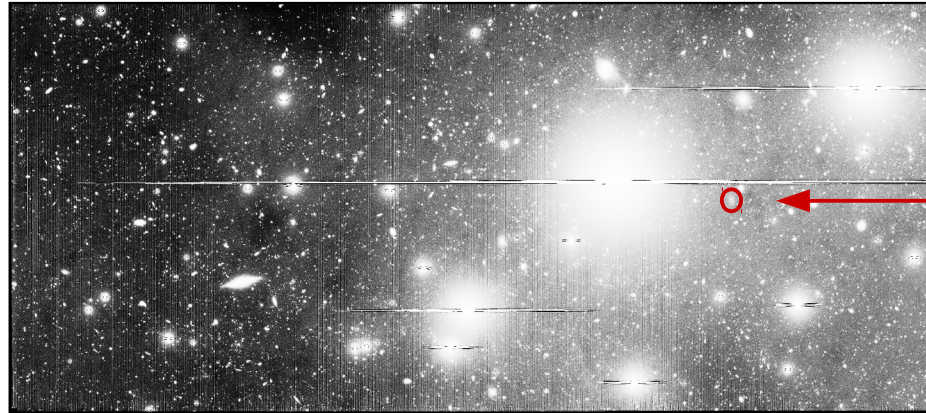
At 10 Mpc – farther
and one of largest
virial volumes we can
probe

Very deep imaging -
Fall target

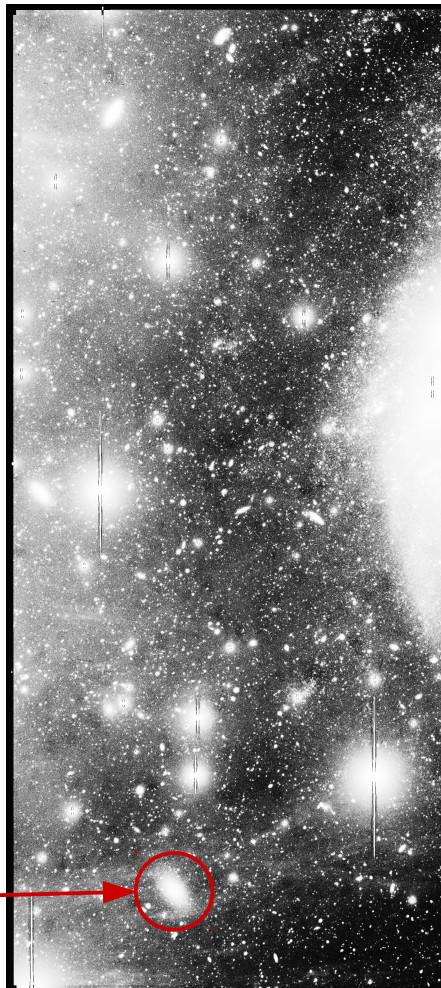
The Data



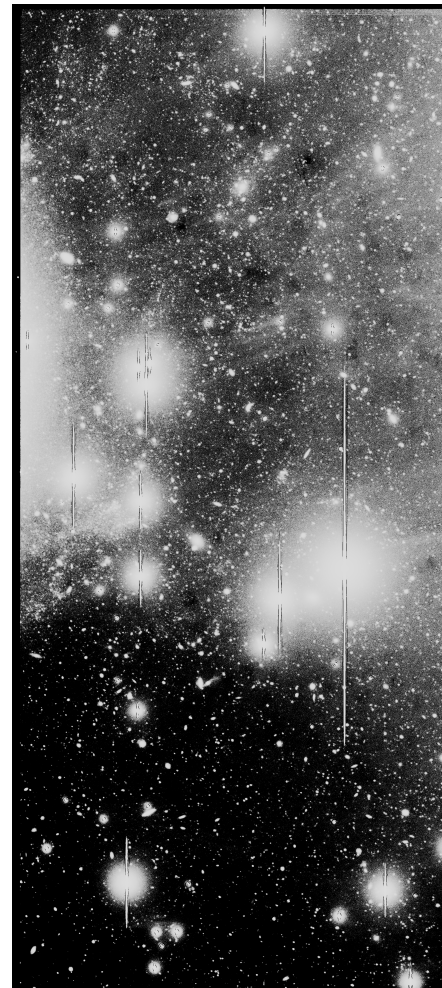
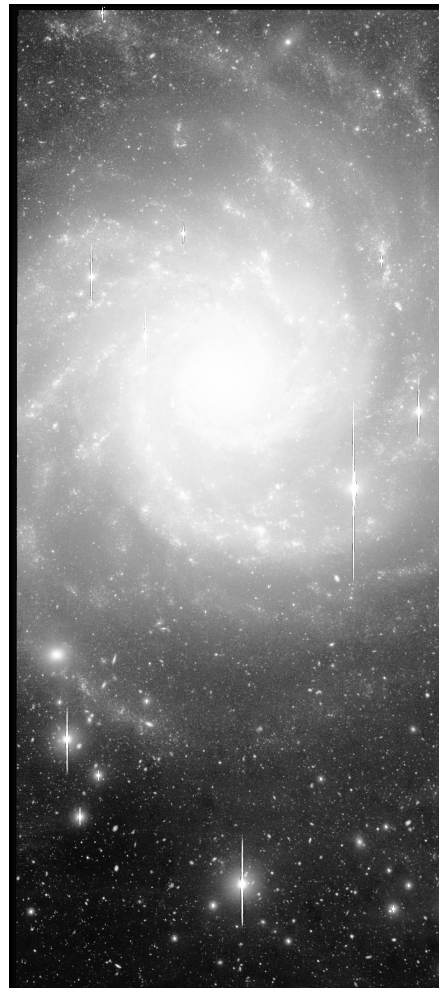
First Candidates



NGC-628-dwB



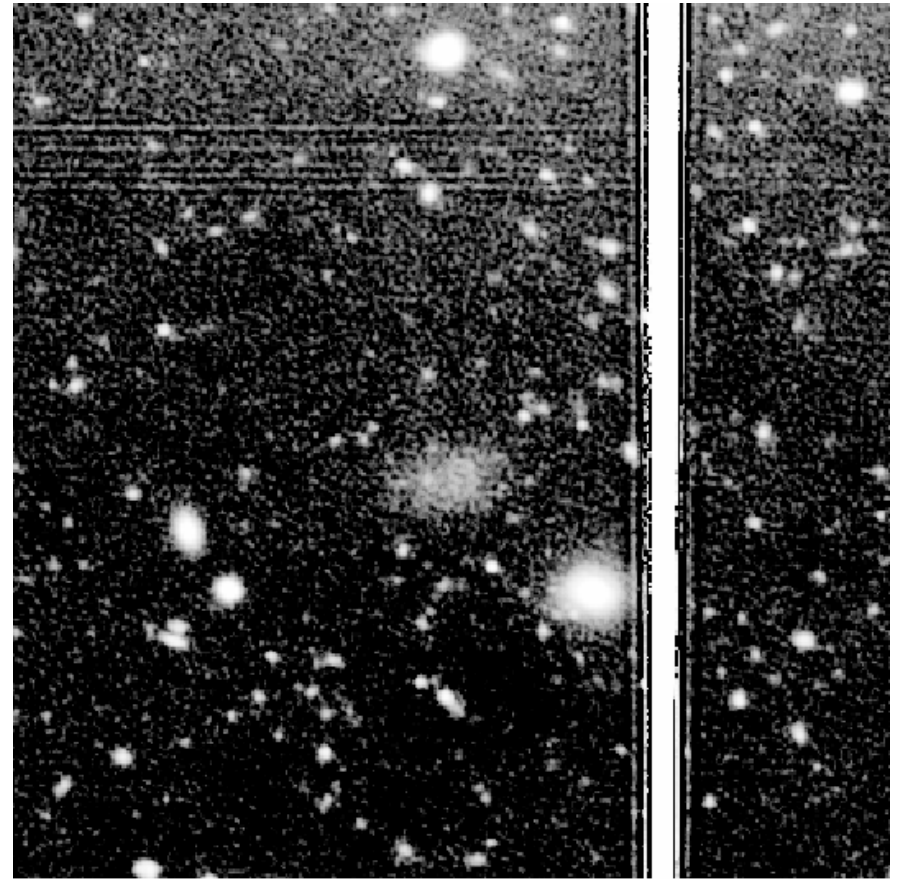
NGC-628-dwA



First Candidates



NGC-628-dwA m_B m_V
 18.9 18.2



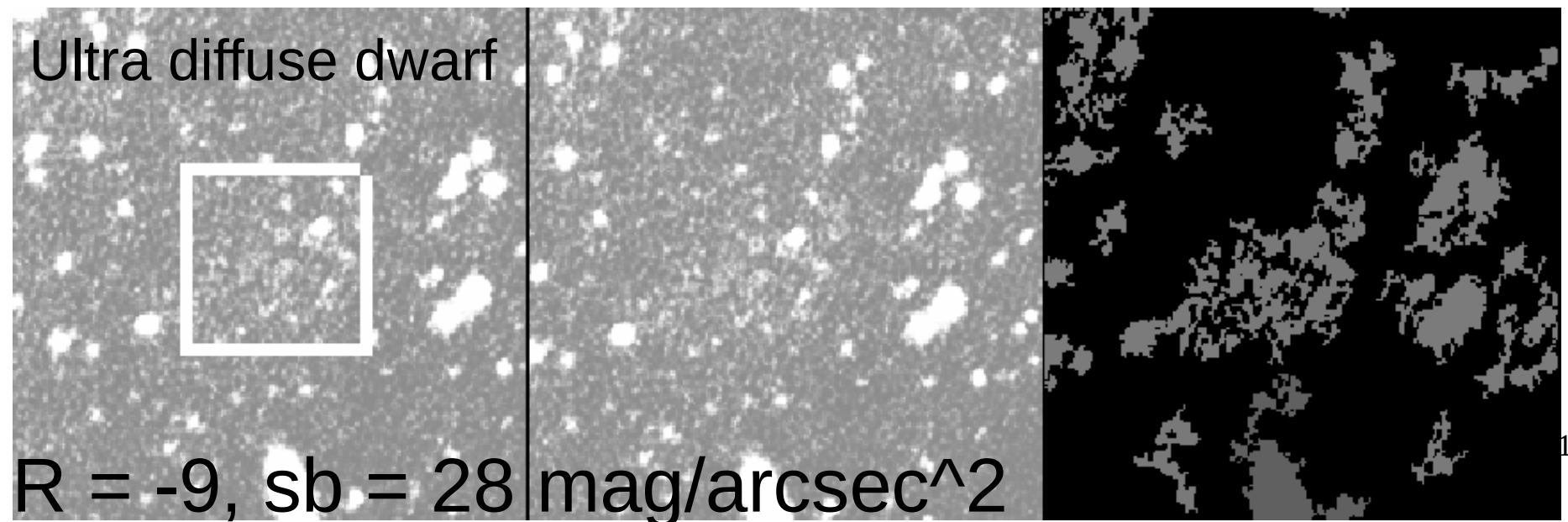
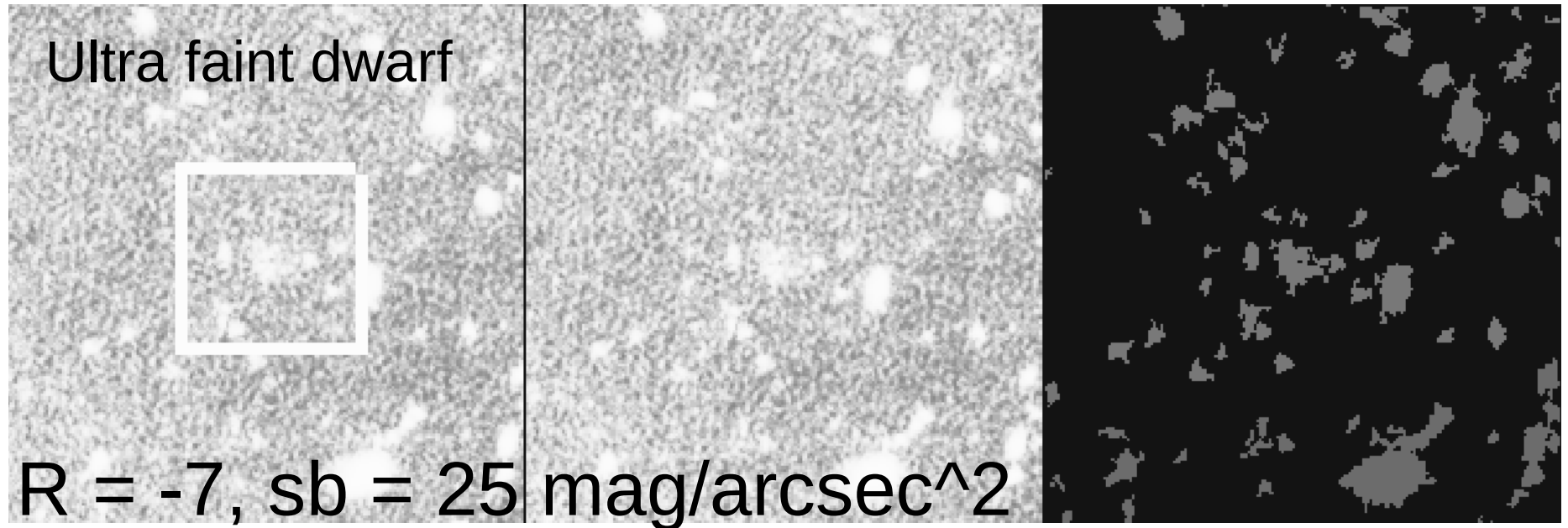
NGC-628-dwB m_B m_V
 > 22.9 22.1

Both red, non-star forming, gas poor and have small projected distances from host

Method

- Using Source Extractor software to identify objects
- Inserting and recovering fake galaxies
- Computing completeness
- Using properties of fake galaxies in data, finding new candidates

Sample of recovered galaxies



Conclusion

Characterizing 20 new systems down to ultra faint and ultra diffuse regimes

=> constrain the Stellar mass – DM halo relation

=> constrain DM models based on their predicted subhalo populations