



Search for Dark Matter Gamma-ray Emission from M31 with HAWC

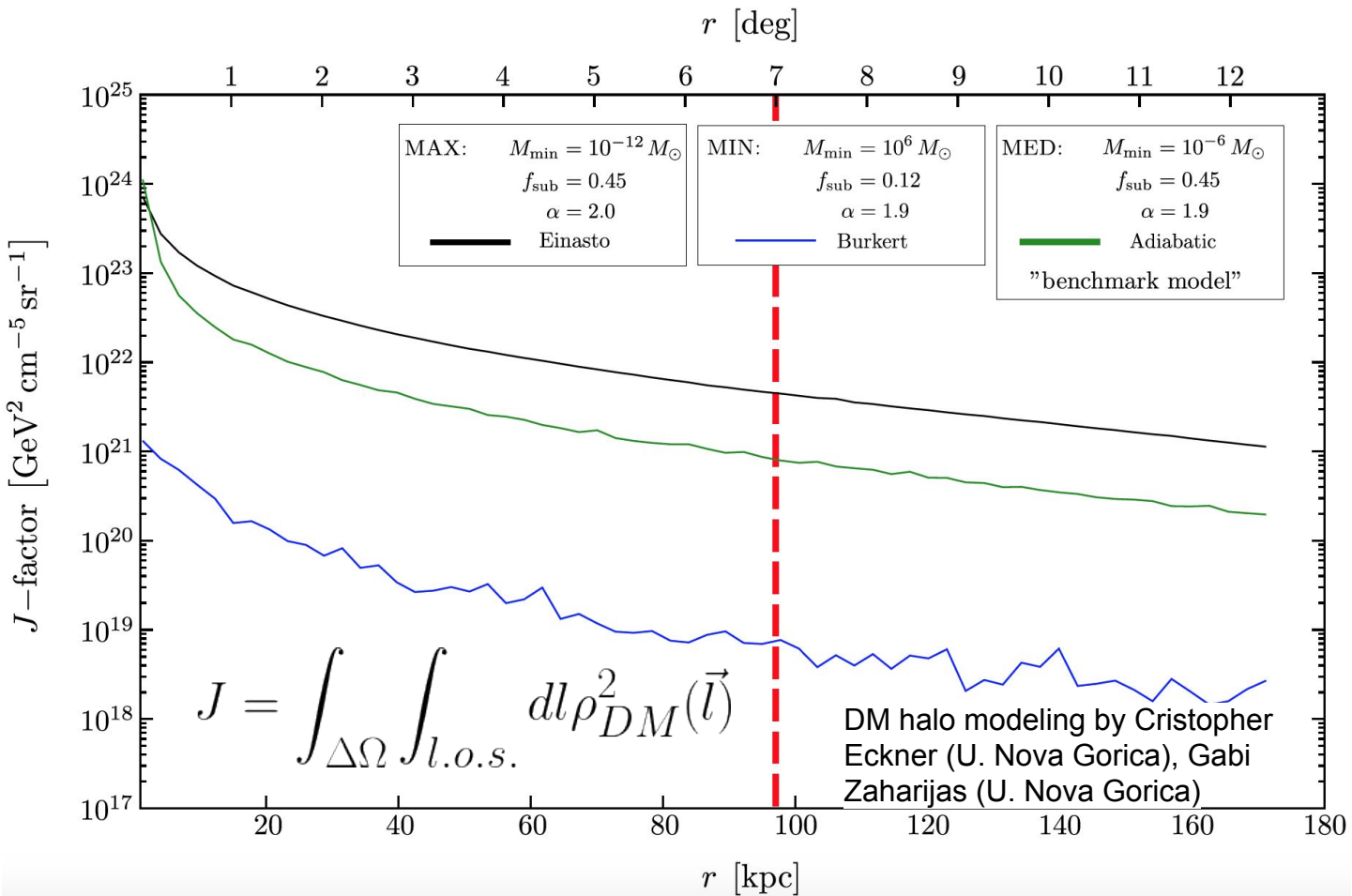
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TeVPA
Dark Matter Parallel
August 8, 2017

The Andromeda Galaxy (M31)

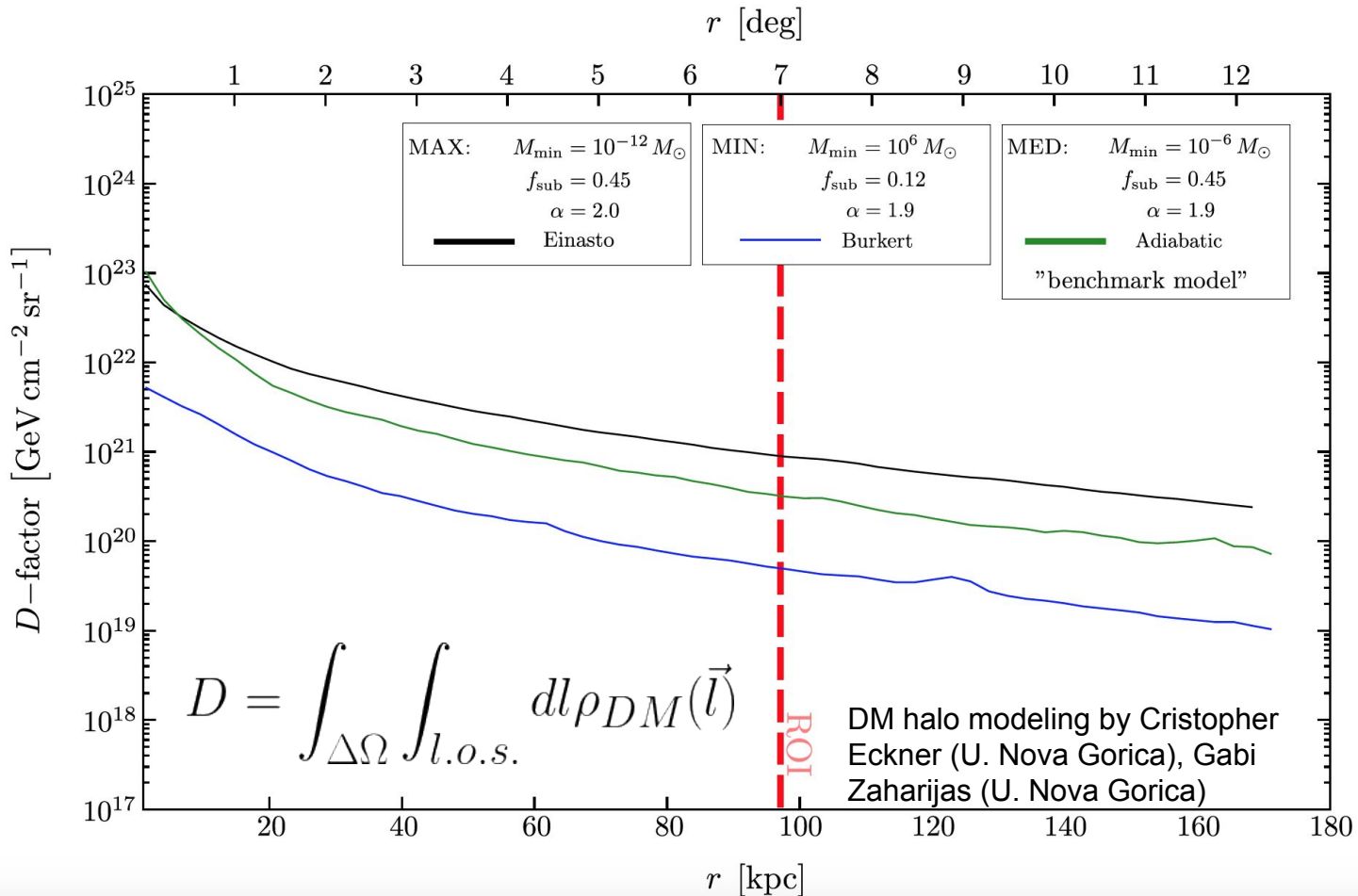


- **Nearby (~ 780 kpc) spiral Galaxy similar to the Milky Way**
- **Stellar rotation curves \rightarrow resides in a large dark matter halo**
- **Good target to search for gamma rays produced via dark matter annihilation or decay**



- Need to model both the smooth DM component and substructure
- Define MIN and MAX models and a realistic benchmark MED
- Smooth components come from Tamm+ (2012)
- Full halo modeled using CLUMPY software

The Andromeda Galaxy (M31)



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HAWC Observatory

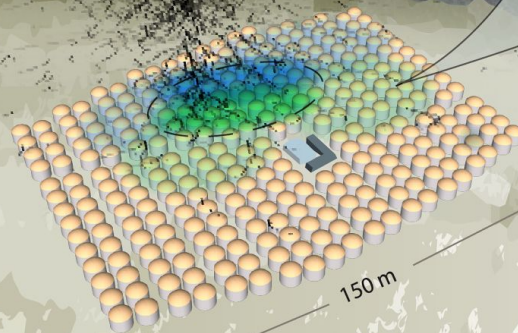
HAWC operates day and night, providing a large field of view for the observation of the highest energy gamma rays.



~17 rad lengths

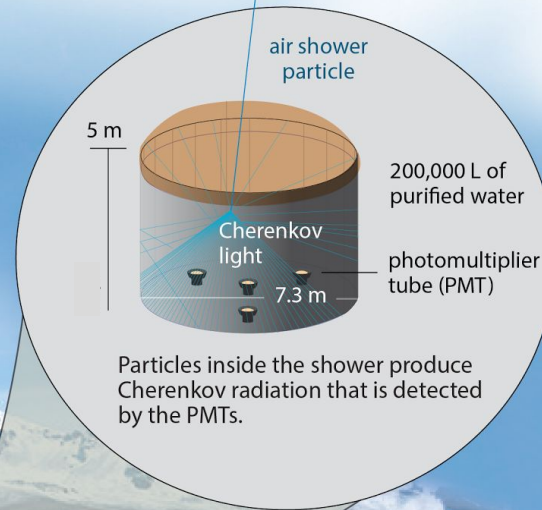
Pico de Orizaba
(5,626 m)

HAWC is located at 4,100 m above sea level, covering an area of 20,000 m².



Water Cherenkov tank

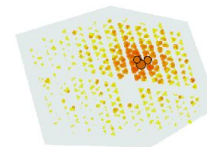
HAWC comprises an array of 300 tanks that record the particles created in gamma-ray and cosmic-ray showers.



Gamma rays vs cosmic rays

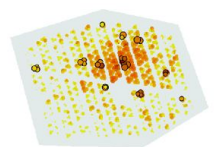
HAWC selects gamma rays from among a much more abundant background of cosmic rays.

gamma-ray shower

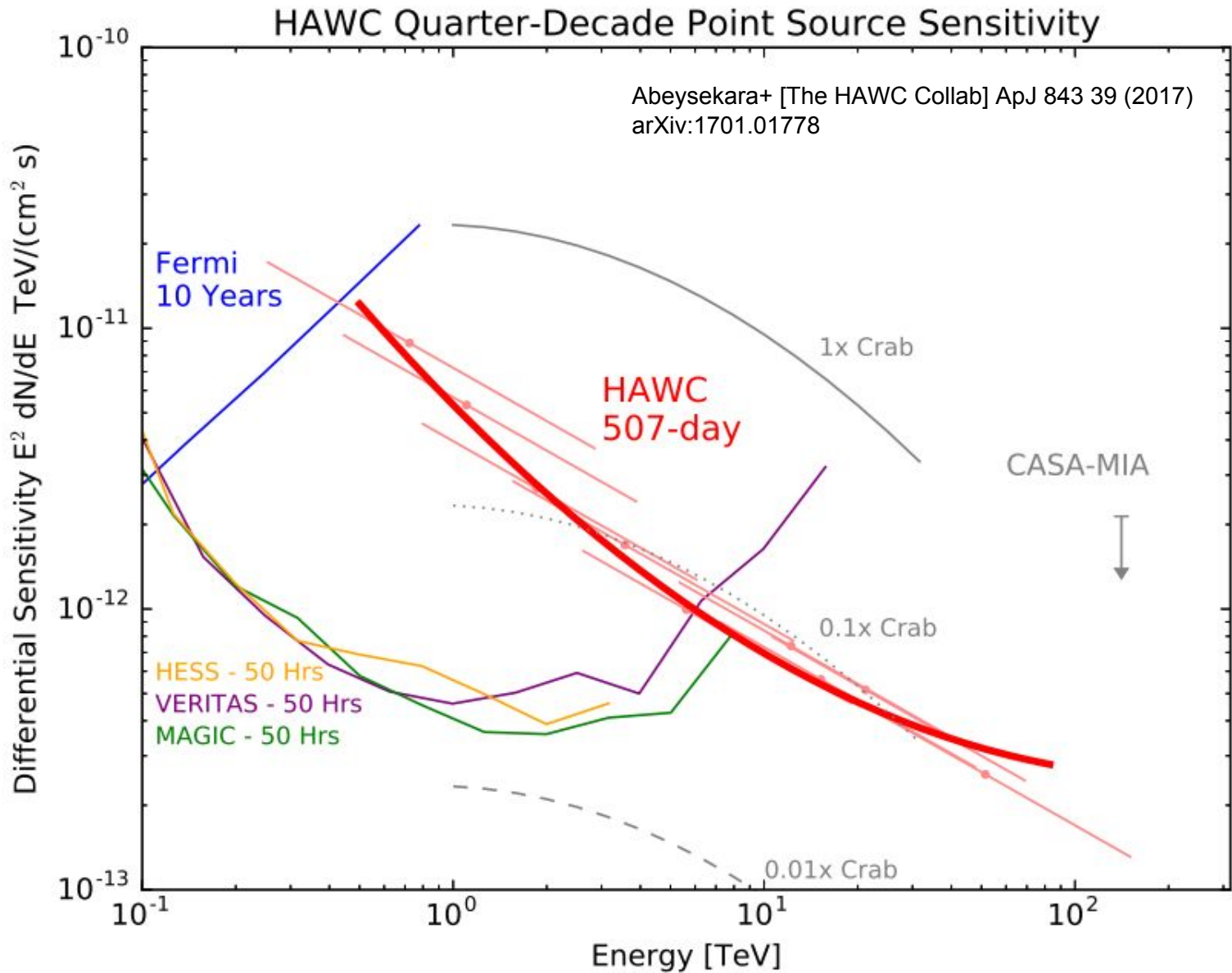


"hot" spots concentrate around the core

cosmic-ray shower



"hot" spots are more dispersed

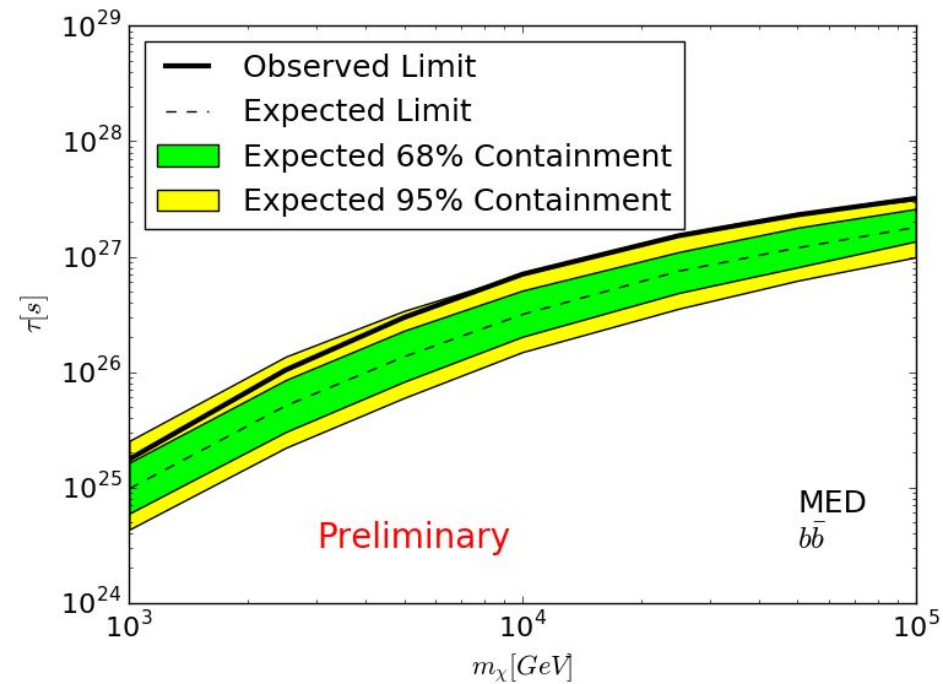
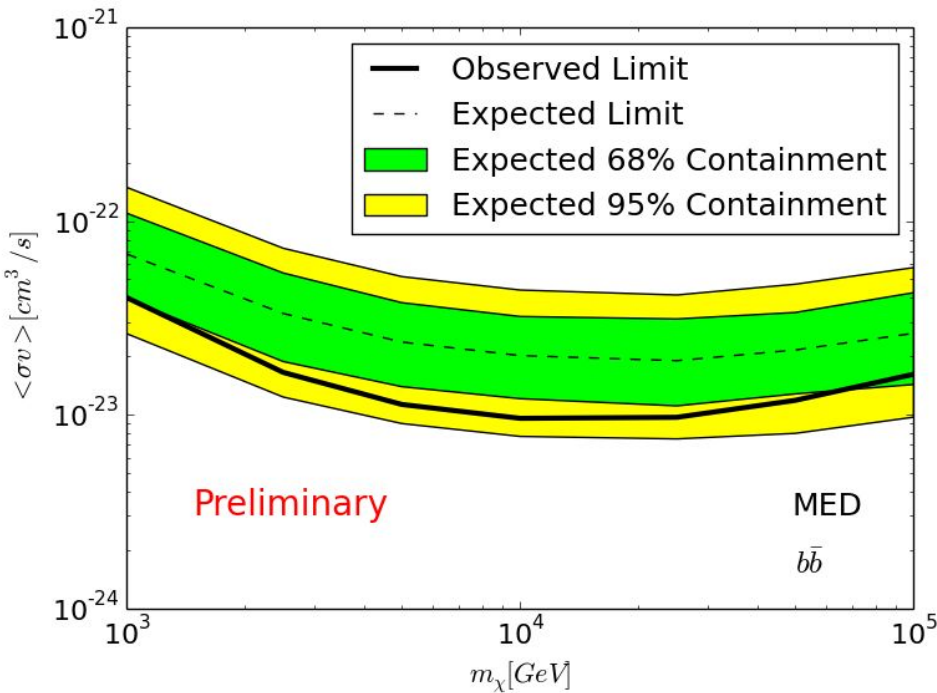


- 760 days of HAWC 300 data
- HAWC analysis is done in ‘fraction of available PMTs’ bins
 - Proxy for energy
 - See Abeysekara+ [The HAWC Collab] ApJ 843 39 (2017) arXiv:1701.01778
- Use the Multi-Mission Maximum Likelihood (3ML) software
 - Available on GitHub <https://github.com/giacomov/3ML>
 - See Vienello+ arXiv:1507.08343

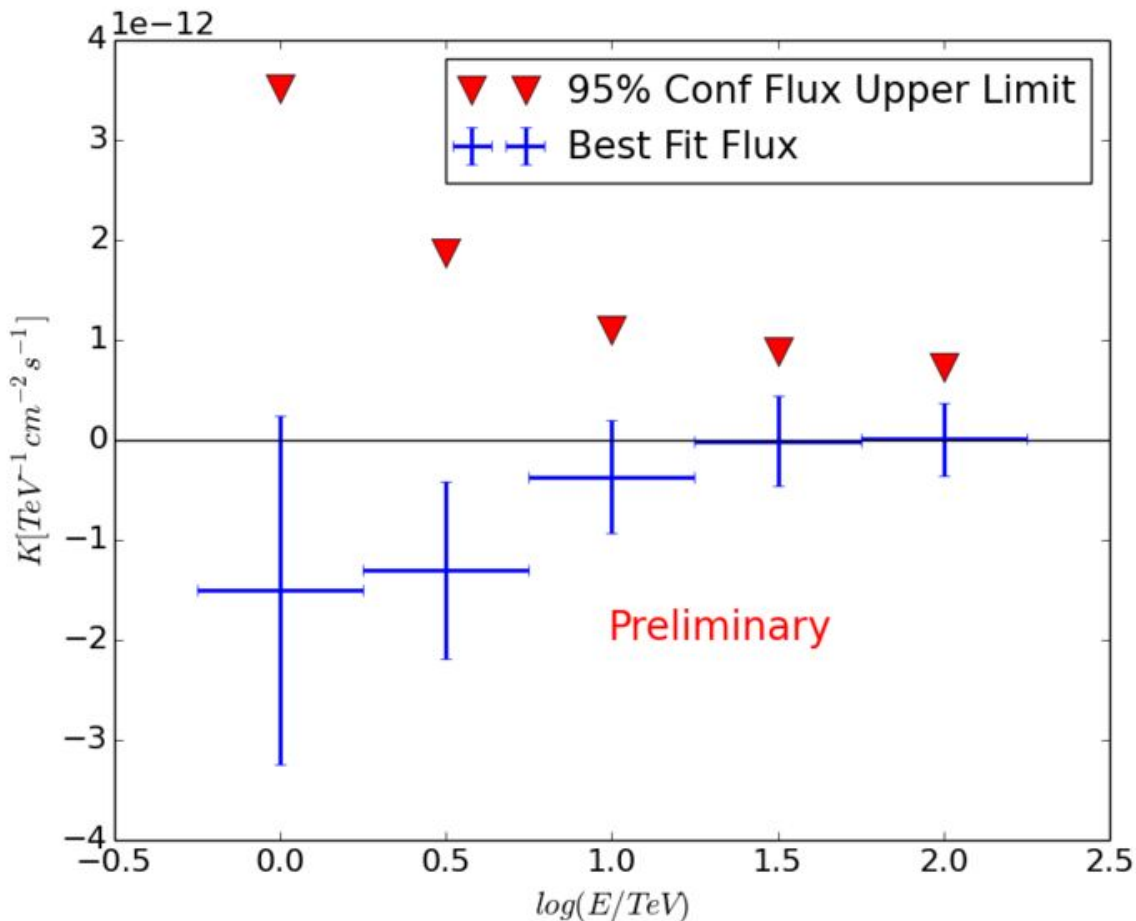
HAWC analysis bins

B	f_{hit}
1	6.7 - 10.5%
2	10.5 - 16.2%
3	16.2 - 24.7%
4	24.7 - 35.6%
5	35.6 - 48.5%
6	48.5 - 61.8%
7	61.8 - 74.0%
8	74.0 - 84.0%
9	84.0 - 100.0%

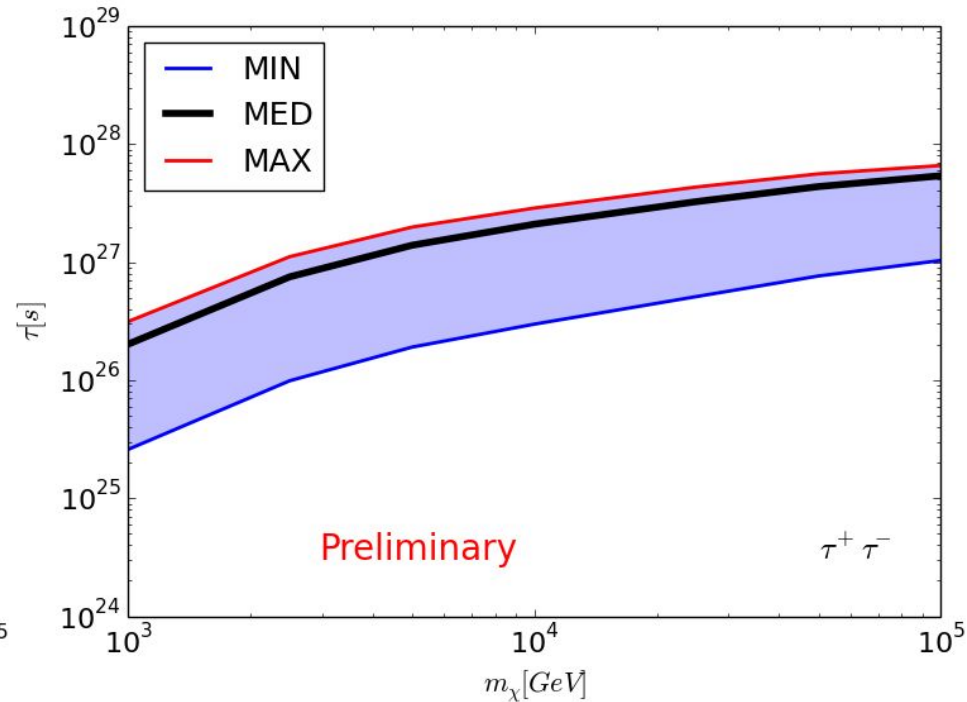
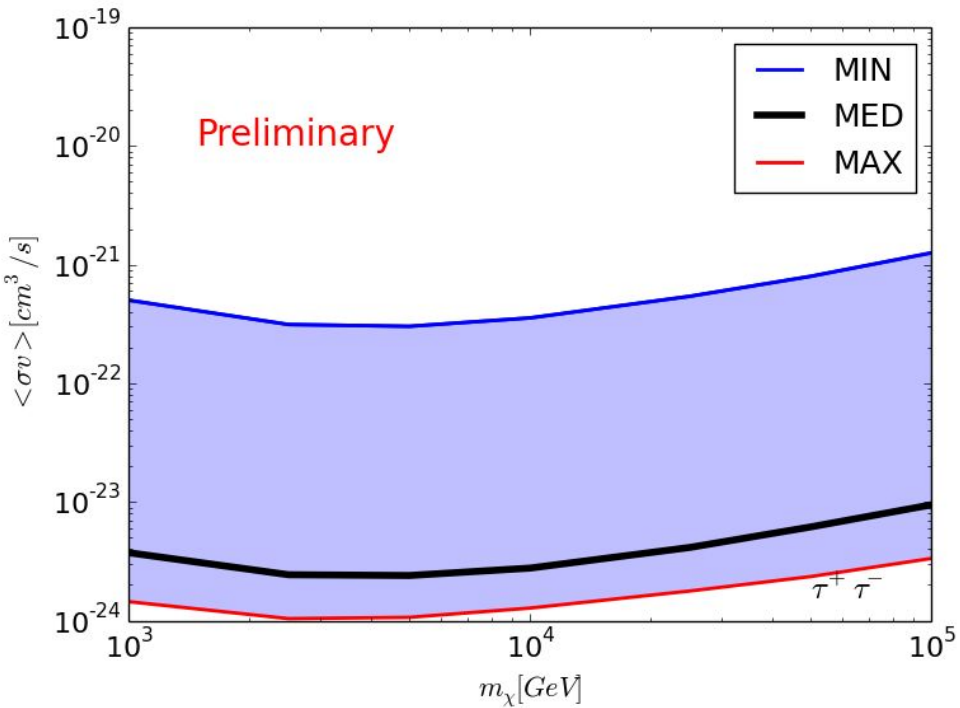
energy ↓



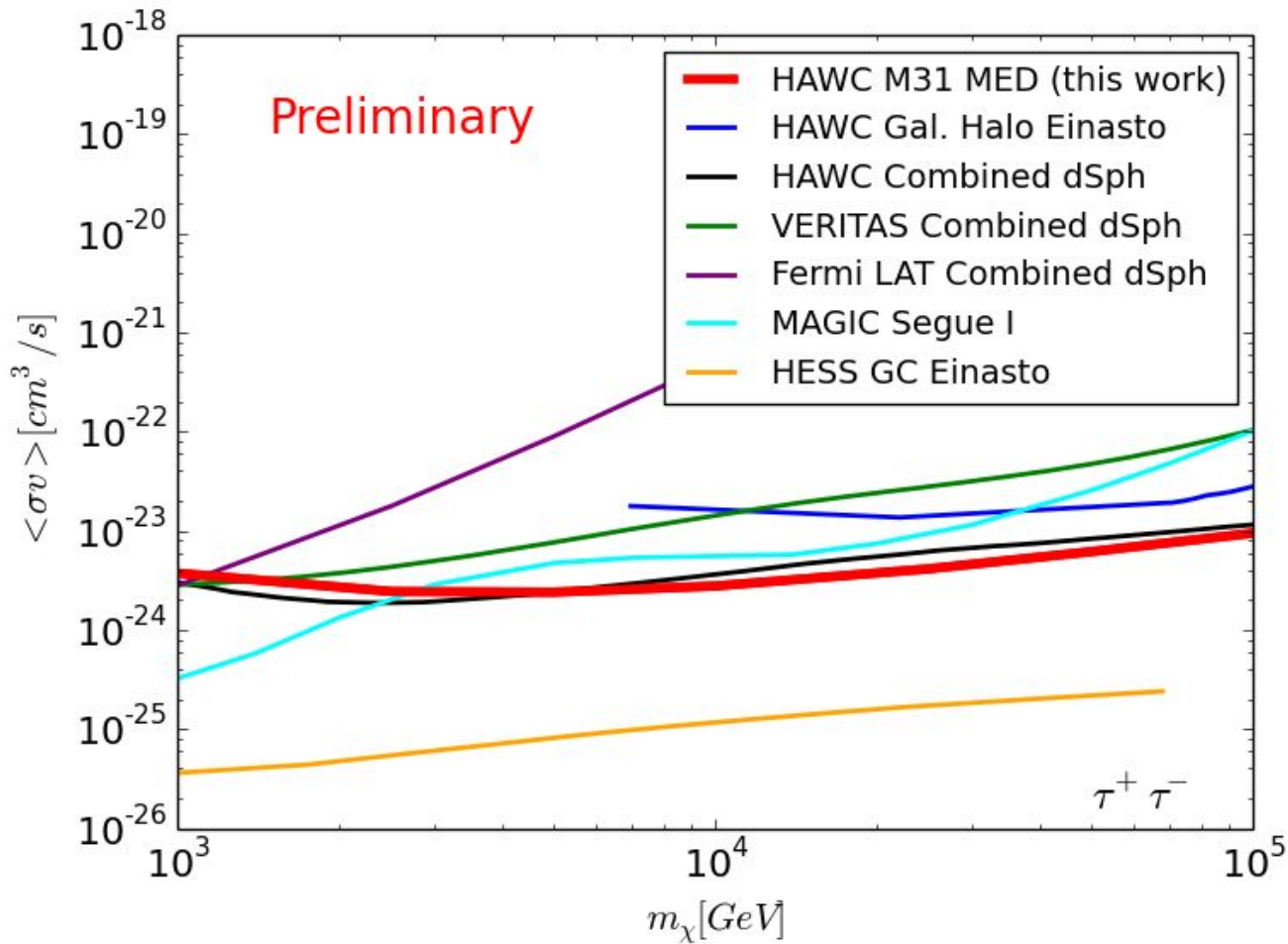
- No gamma-ray excess detected
- Limits set on DM annihilation cross section and decay lifetime
- Limits are 1 to 1.5 sigma below expectation



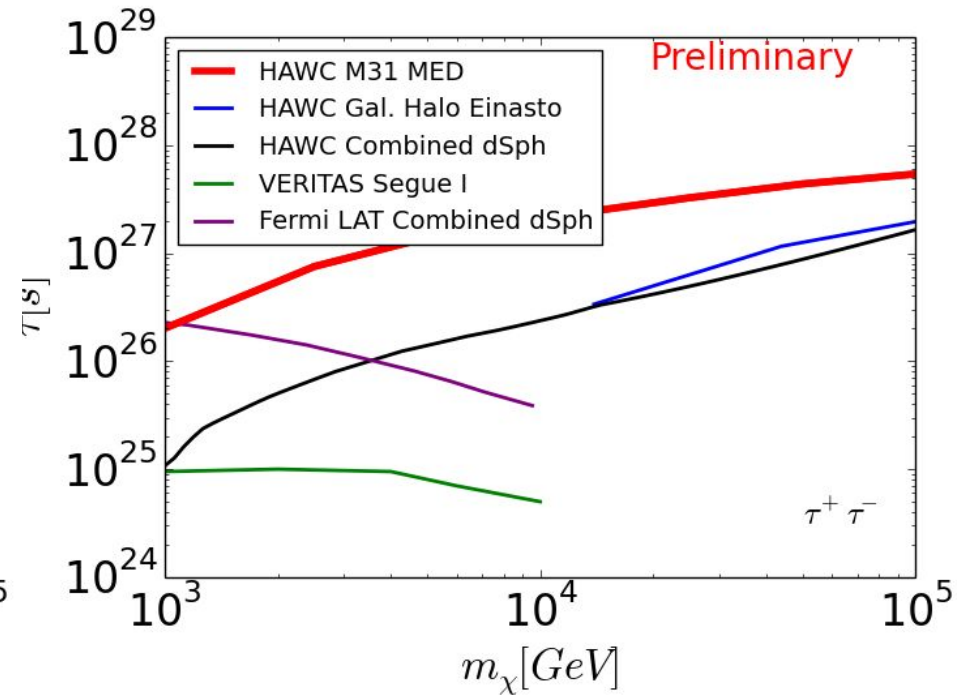
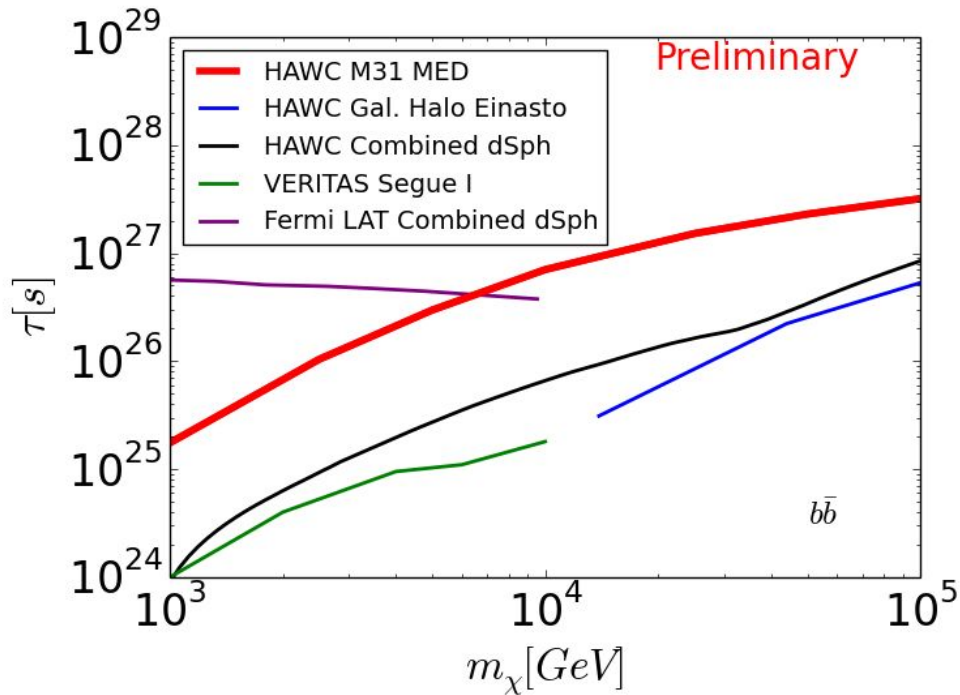
- Calculate quasi-differential DM model-independent flux limits
 - See [Aartsen+ arXiv:1702.06131](#) and [Albert+ arXiv:1706.01277](#)
- Find best fit normalization of powerlaw ($\Gamma=2$) restricted to half decade in $\log(E/\text{TeV})$
 - Calculate 95% CL limit to be where $\Delta\text{TS} = 2.71$



- No gamma-ray excess detected
- Limits set on DM annihilation cross section and decay lifetime
- Spread in limits between MIN and MAX models from DM halo modeling uncertainties



- HAWC M31 limits complement DM limits from other experiments
- Most constraining annihilation limits for mass > 70 TeV



- **HAWC M31 limits complement DM limits from other experiments**
- **Most constraining decay limits in $b\bar{b}$ for mass > 7 TeV**
- **Most constraining decay limits in $\tau^+\tau^-$ for mass > 1 TeV**

- **M31 being close by and in a large dark matter halo makes it a good target for indirect dark matter searches**
- **We find no gamma-ray excess in the direction of M31 in 760 days of HAWC data**
- **We calculate annihilation and decay limits using 3 different dark matter halo models: MIN, MED, MAX**
- **HAWC M31 limits complement dark matter limits obtained from other experiments**