

# Autonomous detection of air showers with the TREND radio-array

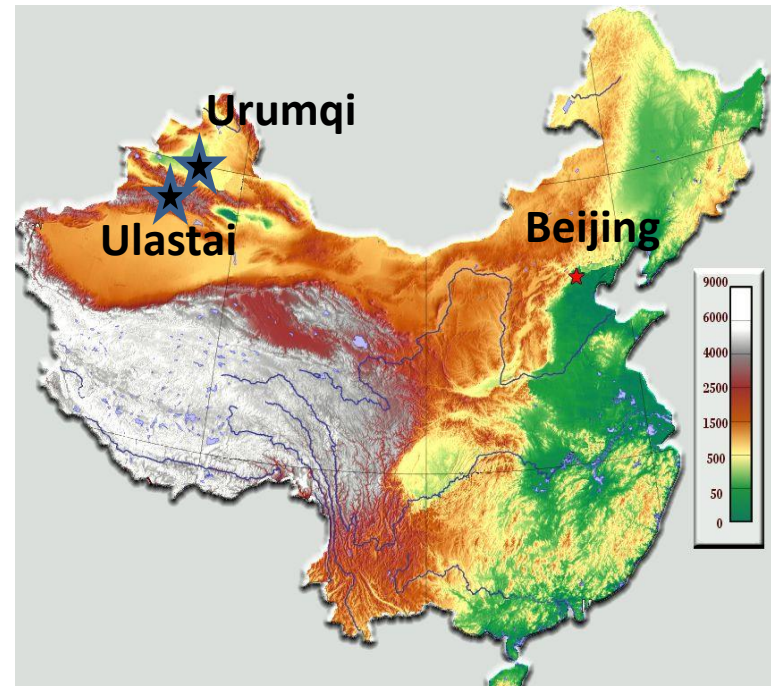
Olivier Martineau-Huynh

TeVPA2017, Columbus

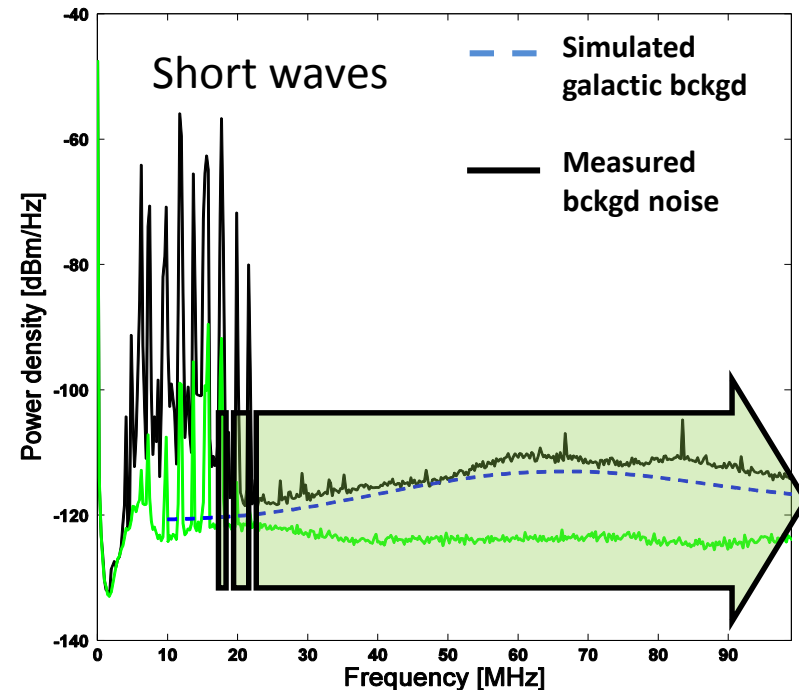
Aug. 10, 2017



# Autonomous radio-detection



- ☺ Cheap, light & robust: radio is an **appealing technique for giant detectors of HE air showers.**
- ☹ **Background sources** are a major challenge for ground-based radio experiments.
- **TREND proposal (2008):** tiny group from China (Wu XiangPing NAOC, Hu HongBo IHEP) & France (P. Lautridou Nantes, V. Niess Clermont-Ferrand, OM Paris/NAOC/IHEP) aiming at **demonstrating that self-triggered radio trigger + identification of air shower is possible.**
- Site: 21CMA radio interferometer, XinJiang Province, China



# TREND setup

50 1D antenna (1 polarisation) – trigger rate up to  $\sim 200\text{Hz}/\text{antenna}$  –  
transfert of analogic signal to DAQ room – on-the-fly digitization –  
trigger if signal  $> 6$  or  $8\sigma$  – record event if 4+ antenna triggers

DAQ periods :

- EW orientation 2011-2012
- NS orientation 2013-2014

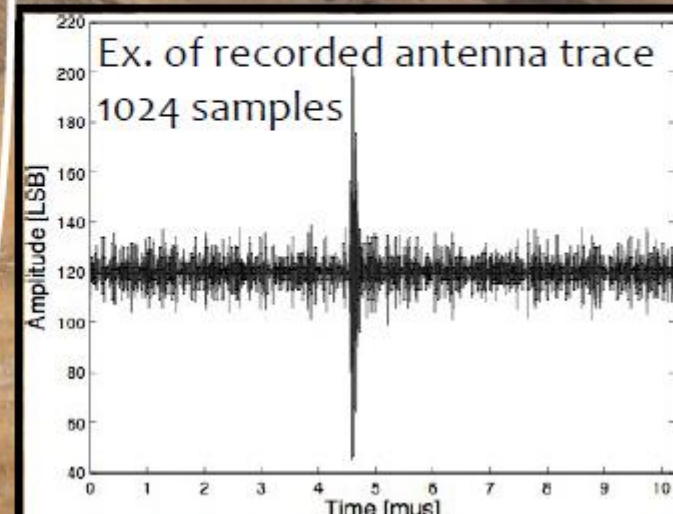
One postdoc  
lost in the  
field

Single polar  
antenna

@ pod  
level  
( $< 300\text{m}$ ):  
optical  
fiber

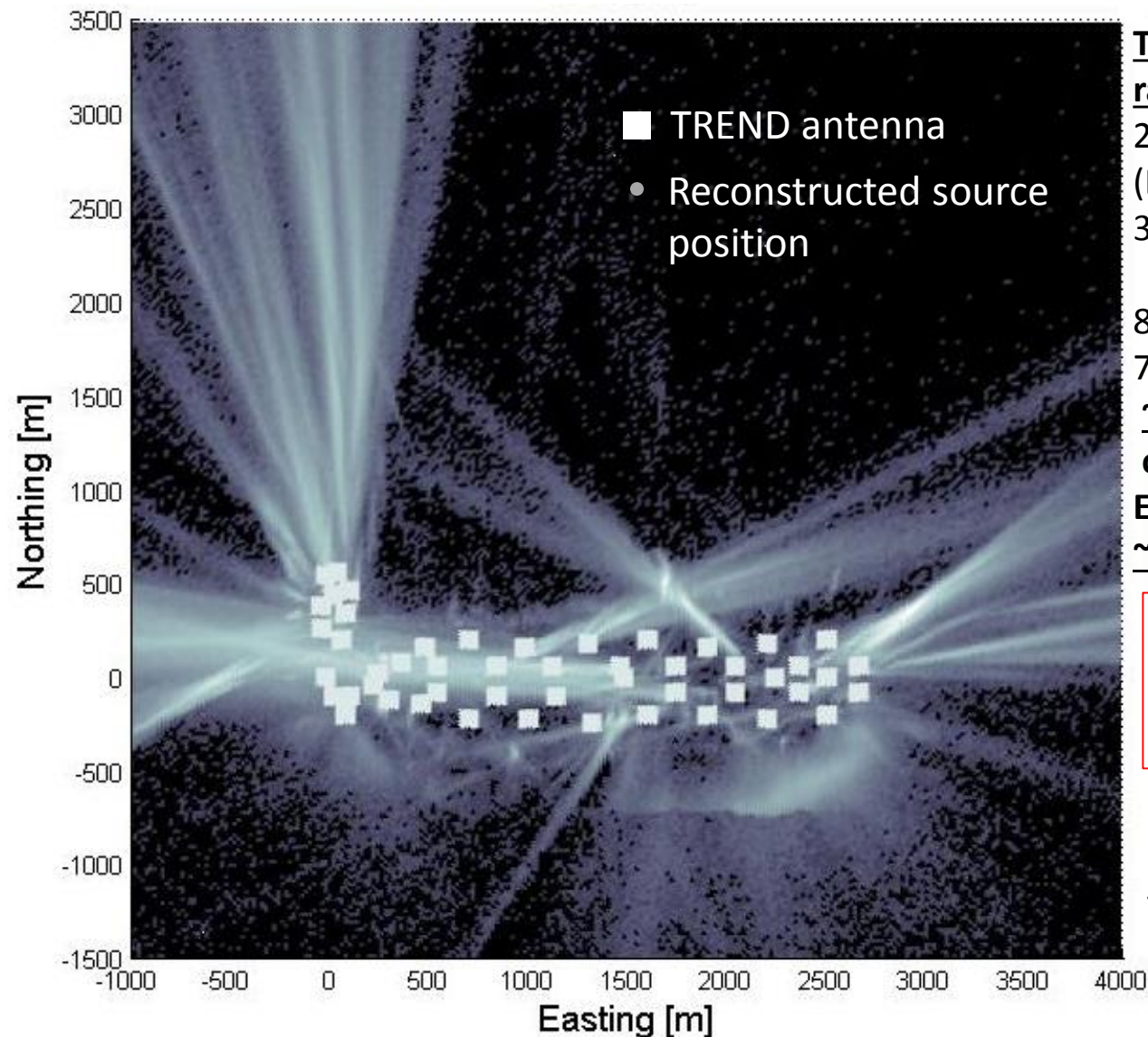
@ DAQ  
room( $< 2\text{km}$ ):  
digitization  
( $200\text{MS}/\text{s} + 8\text{bits}$ )  
+trigger  
+reccord if 4+ant

Ampli (64 dB) +  
filter (50-100MHz)



DAQ= Data AcQuisition

# Radio background



## TREND-50 antennas

### radio array:

2011-2012 data  
(E-W polarization)  
316 DAQ days analyzed

8  $10^9$  triggers recorded  $\Leftrightarrow$  **8TBy**

7  $10^8$  coincidences

~25Hz event rate

**over whole array (physical origin)**

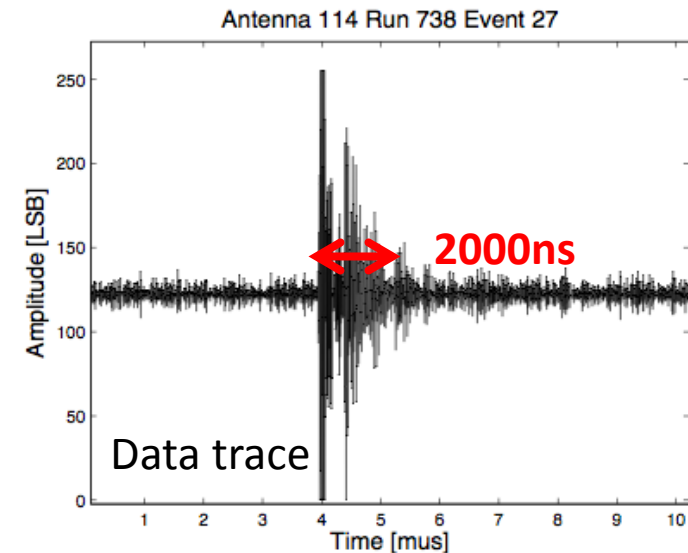
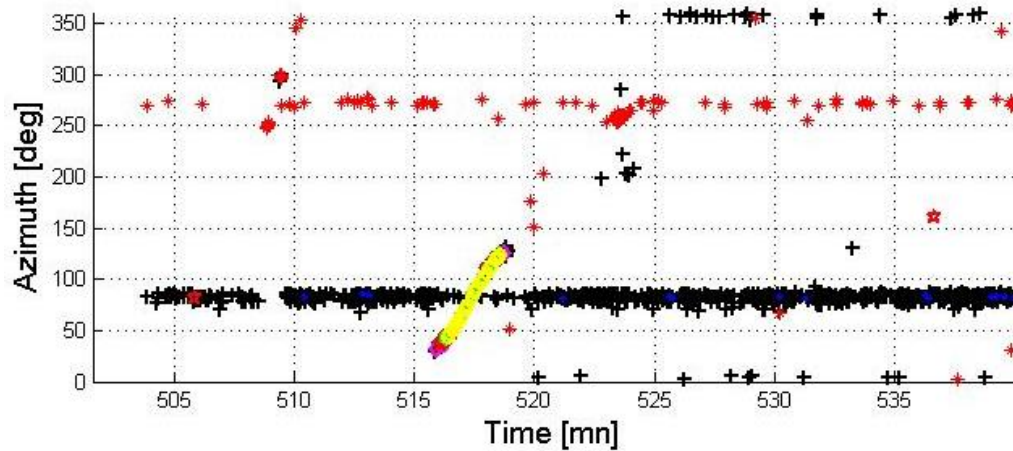
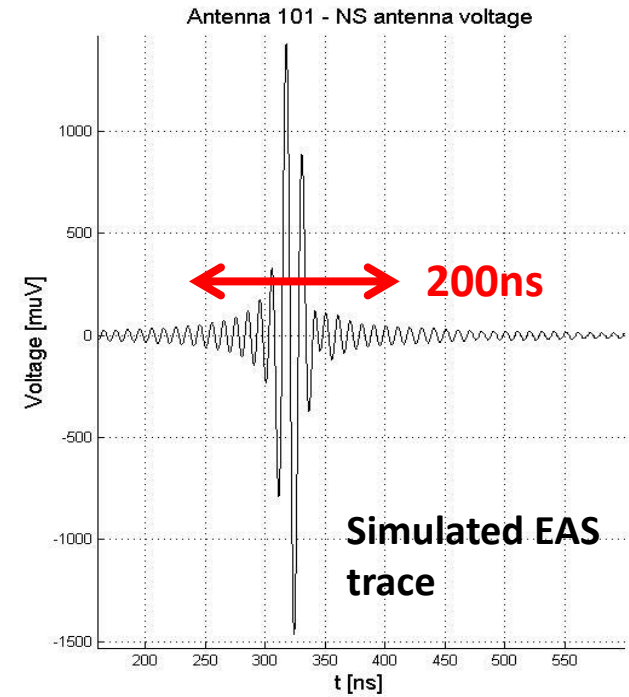
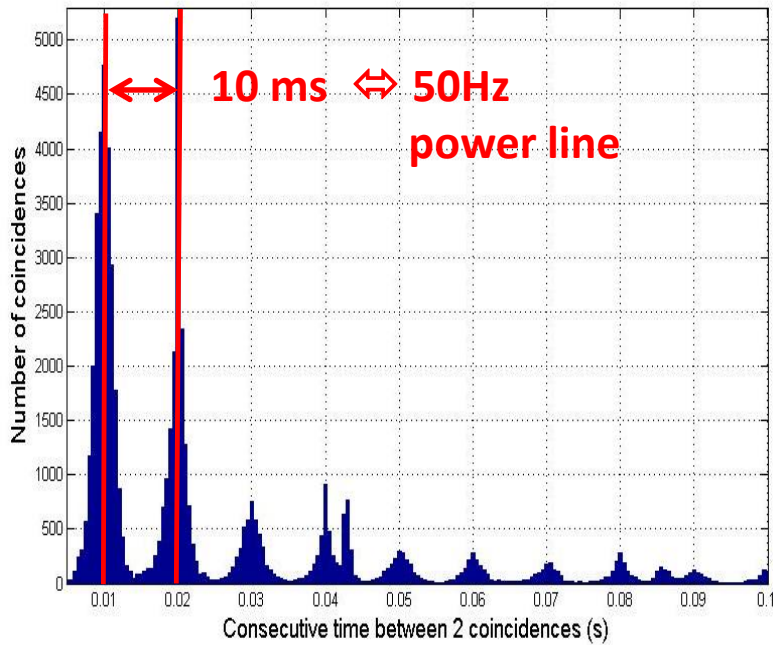
**Expected EAS trigger rate:**

~40/day

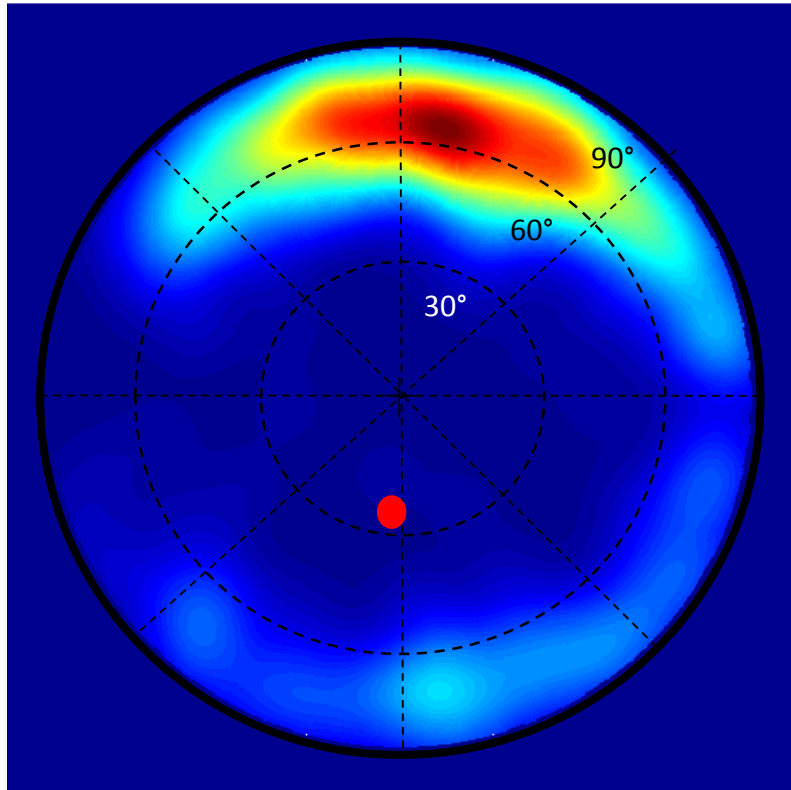
**Background rejection is a key issue for EAS radio-detection.**

Background sources: HV lines, radio emitters, train, cars, planes, thunderstorms...

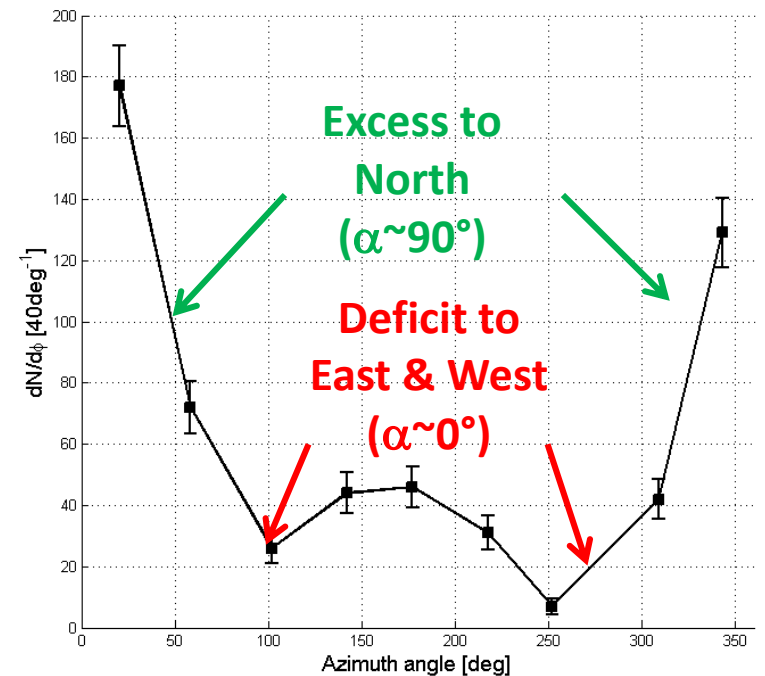
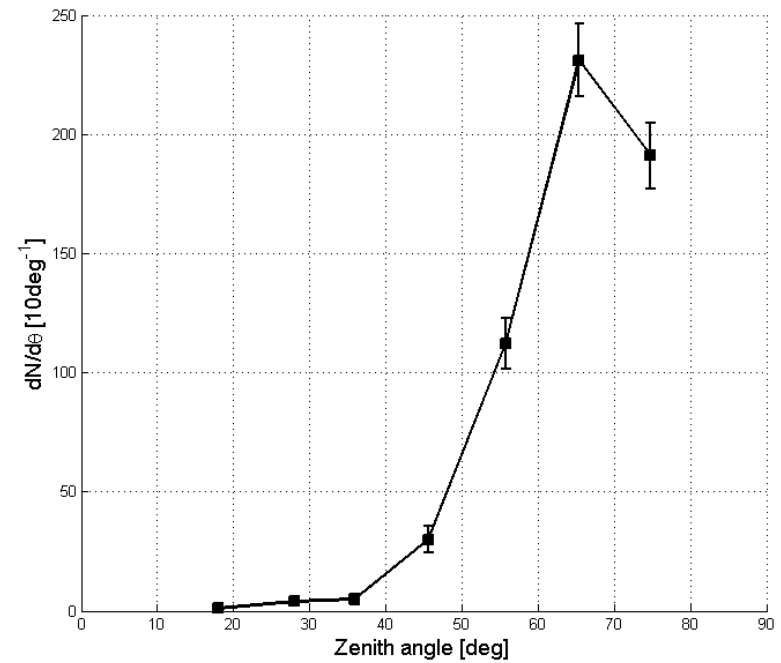
# Discriminating parameters



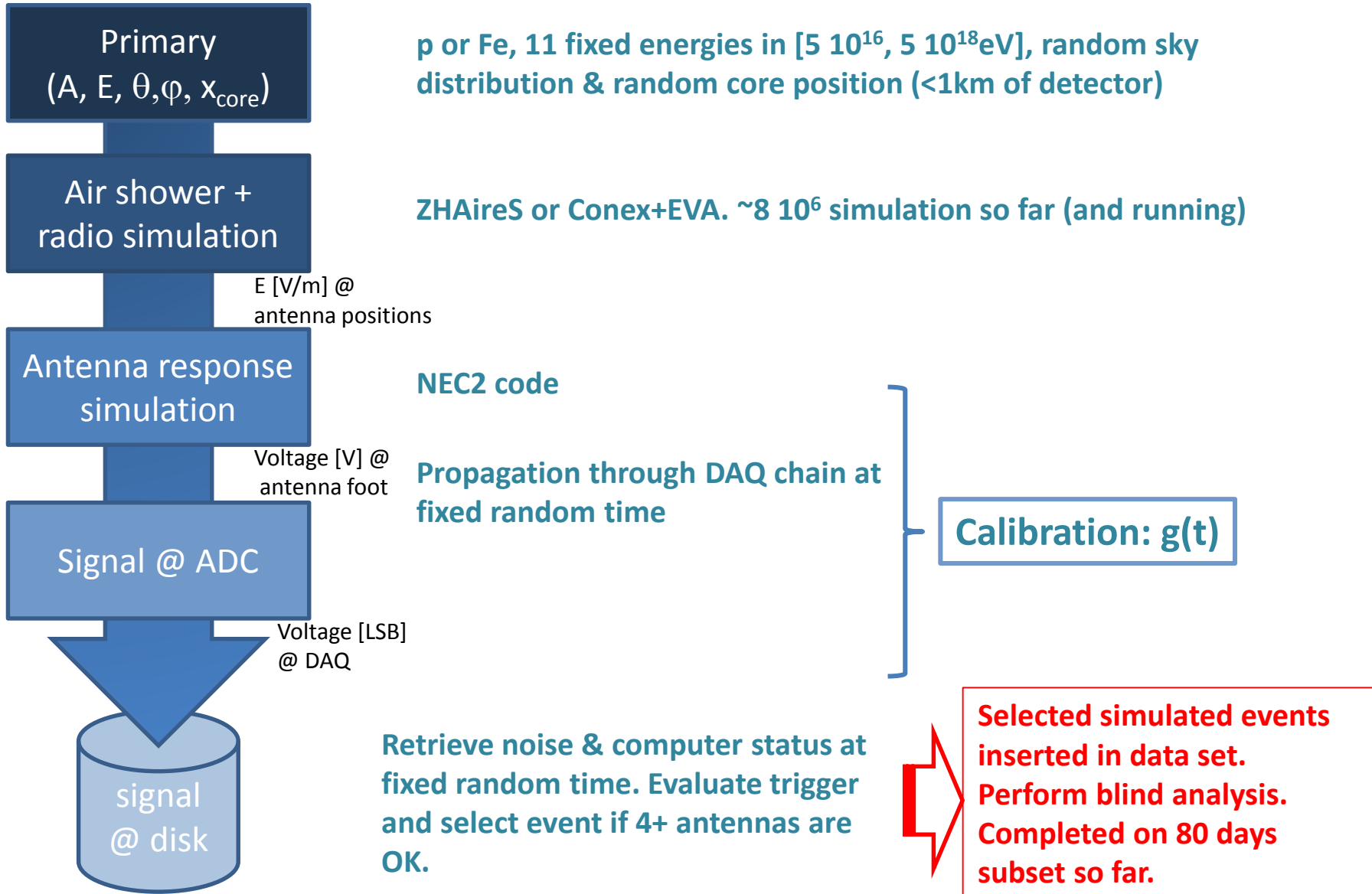
# TREND EAS candidates



2011-2012 data - 316 DAQ days  
EW polar  
574 EAS candidates

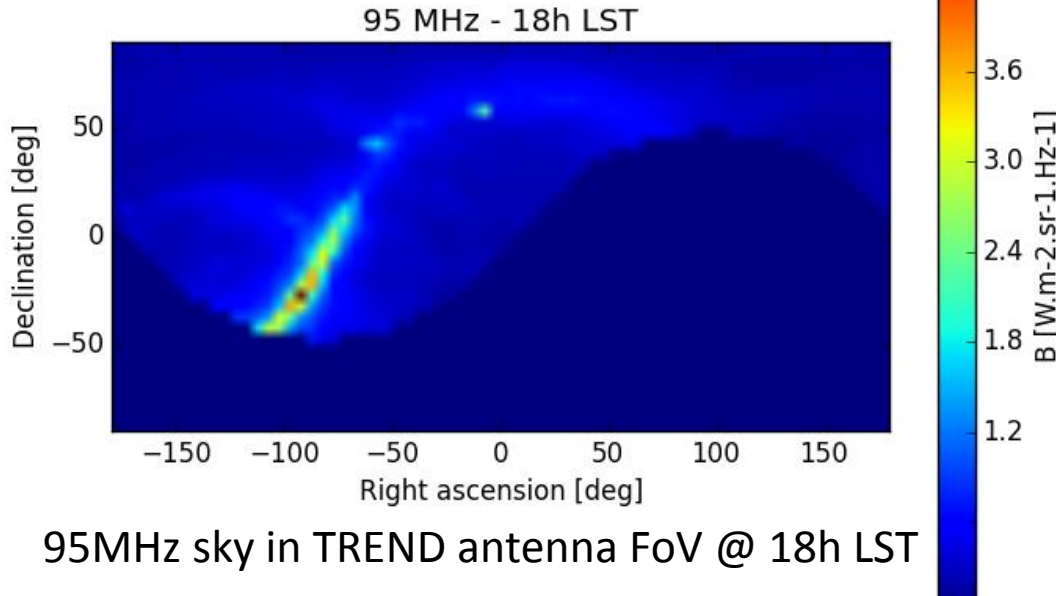


# TREND real-time event simulation



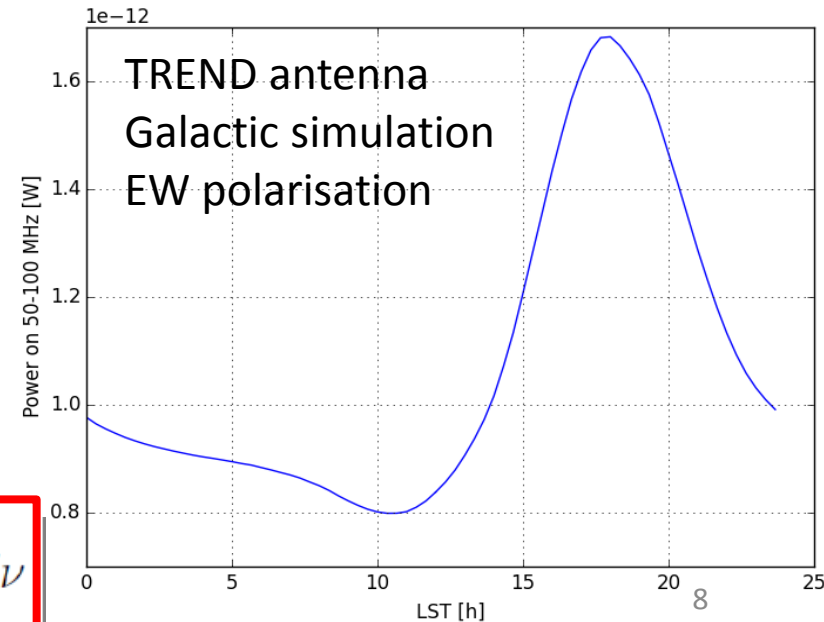
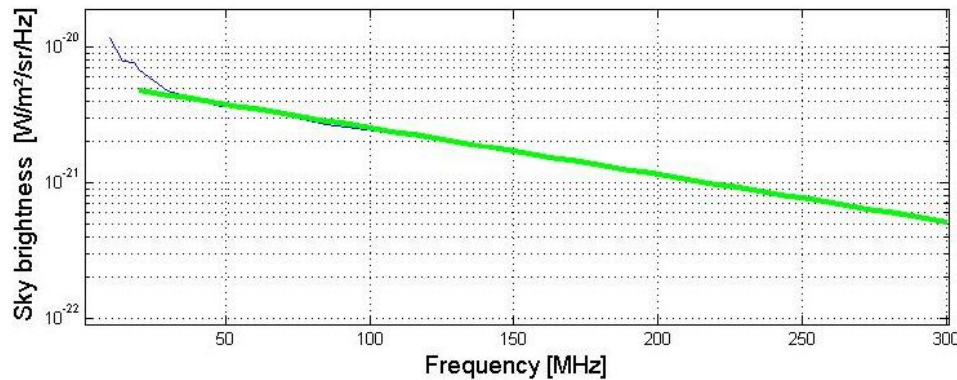
# TREND calibration

[S. Le Coz, V. Niess]



- In TREND frequency range, sky emission highly dominated by Galactic plane radiation.
- Very well known → provides a stable & periodic source for calibration & monitoring.

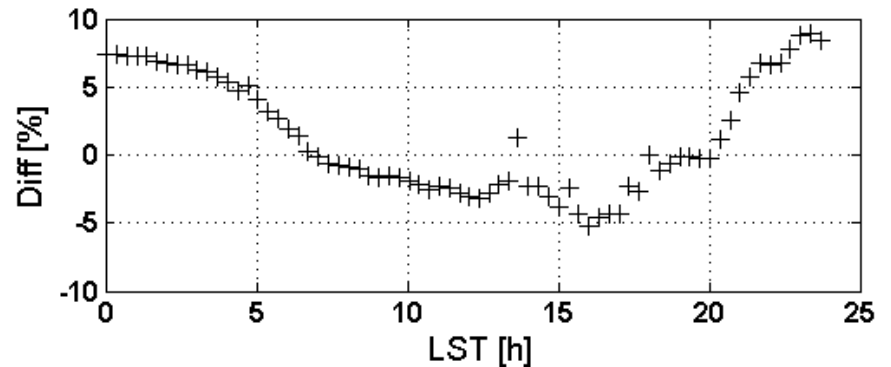
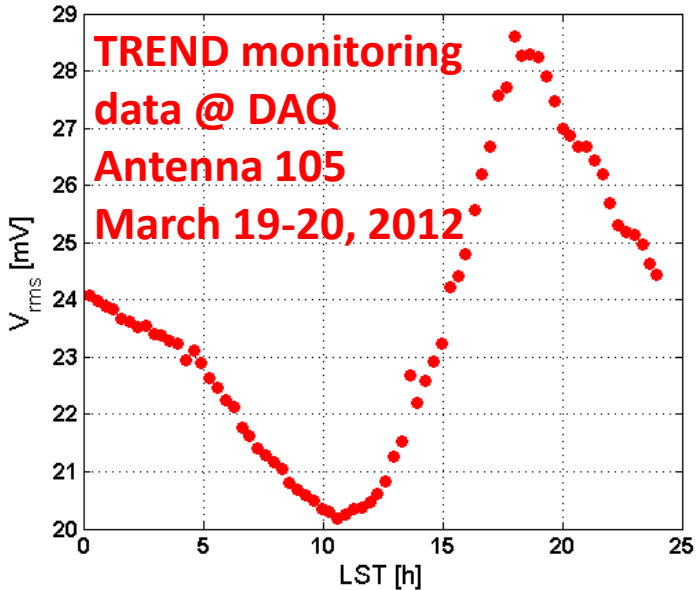
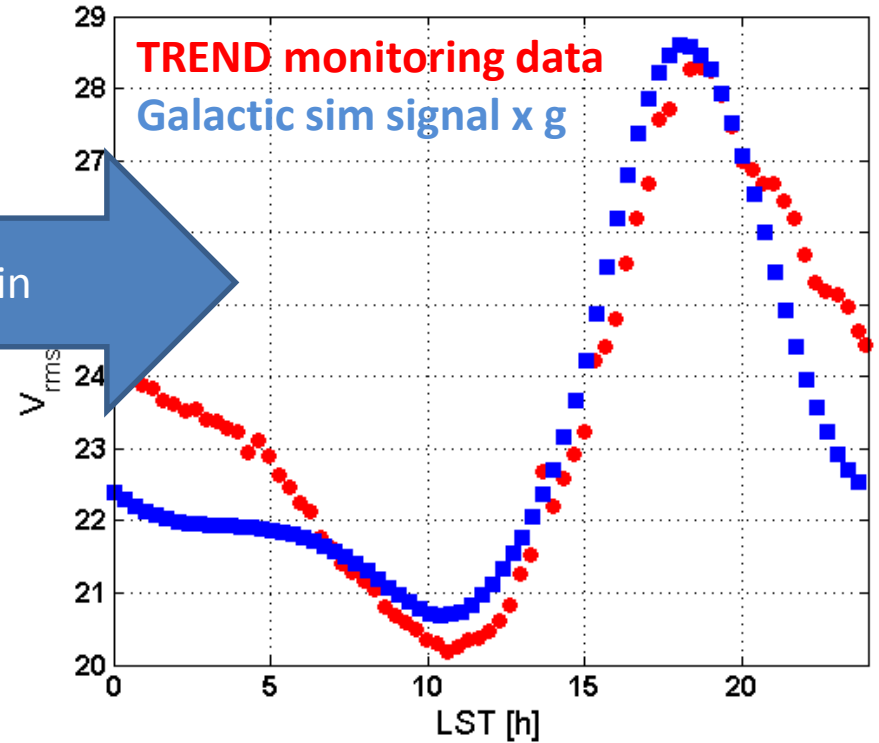
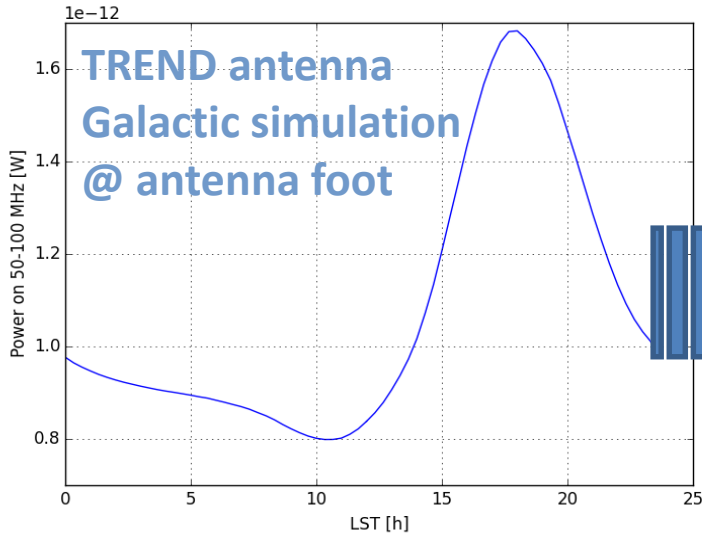
95MHz sky in TREND antenna FoV @ 18h LST



$$\langle V_{sky}^2 \rangle = \frac{R_L}{2} \int_{\Delta\nu} \int_{4\pi} B_\nu(\theta, \varphi) A_{eff}(\theta', \varphi') \sin \theta d\theta d\varphi d\nu$$



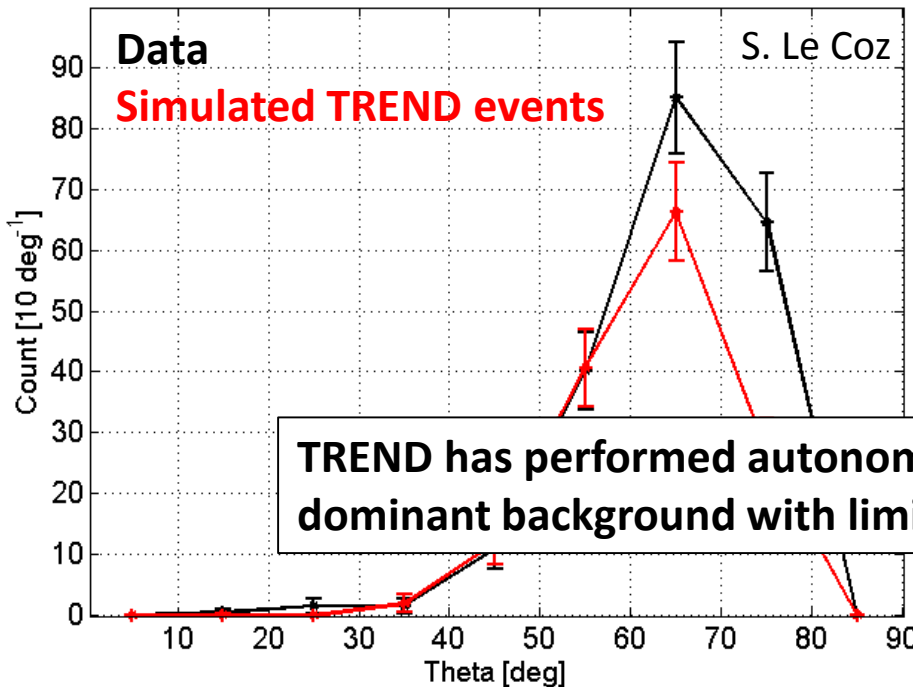
# TREND calibration



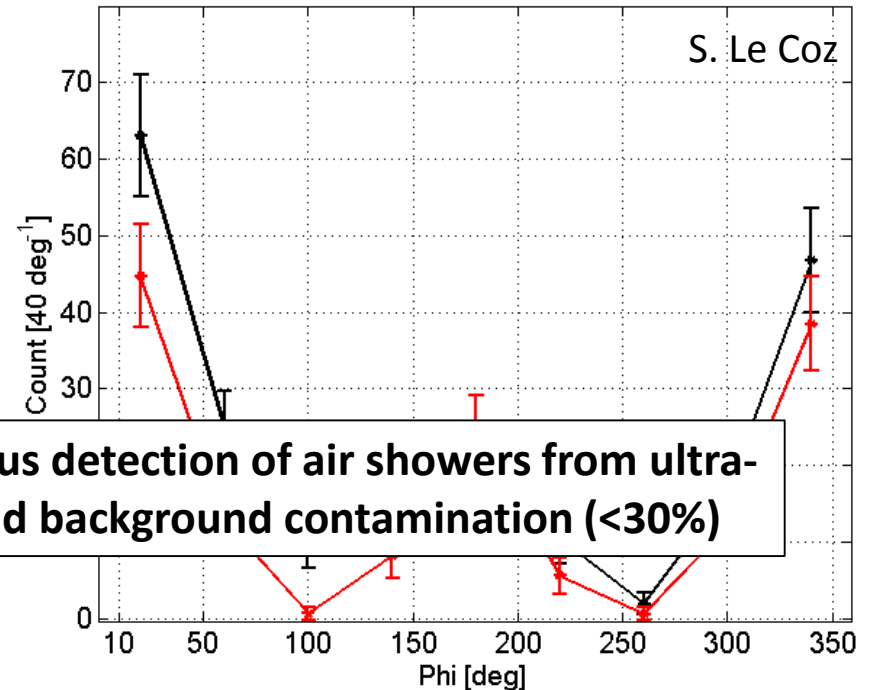
# Results

- Simulated data set allows computing detector aperture:
  - $S_{eff}(E, \theta, \varphi) = S_{sim}(E, \theta, \varphi) * \frac{N_{det}(E, \theta, \varphi)}{N_{sim}(E, \theta, \varphi)}$  effective area [m<sup>2</sup>]
  - $App(E) = \iint_{d\Omega} S_{eff} \cos \theta \sin \theta d\theta d\phi$  aperture [m<sup>2</sup>sr]
- Event number during period  $\Delta t=80$  days:
  - $N(\Delta t) = \int_E App(E) \frac{dN}{dE} dE * \Delta t$  with differential flux  $\frac{dN}{dE} = K_0 E^{-\alpha}$  [m<sup>-2</sup>sr<sup>-1</sup>s<sup>-1</sup>]
  - **N(80 days) = 177<sup>+49</sup><sub>-44</sub>** (205 experimental candidates in the corresponding subset)

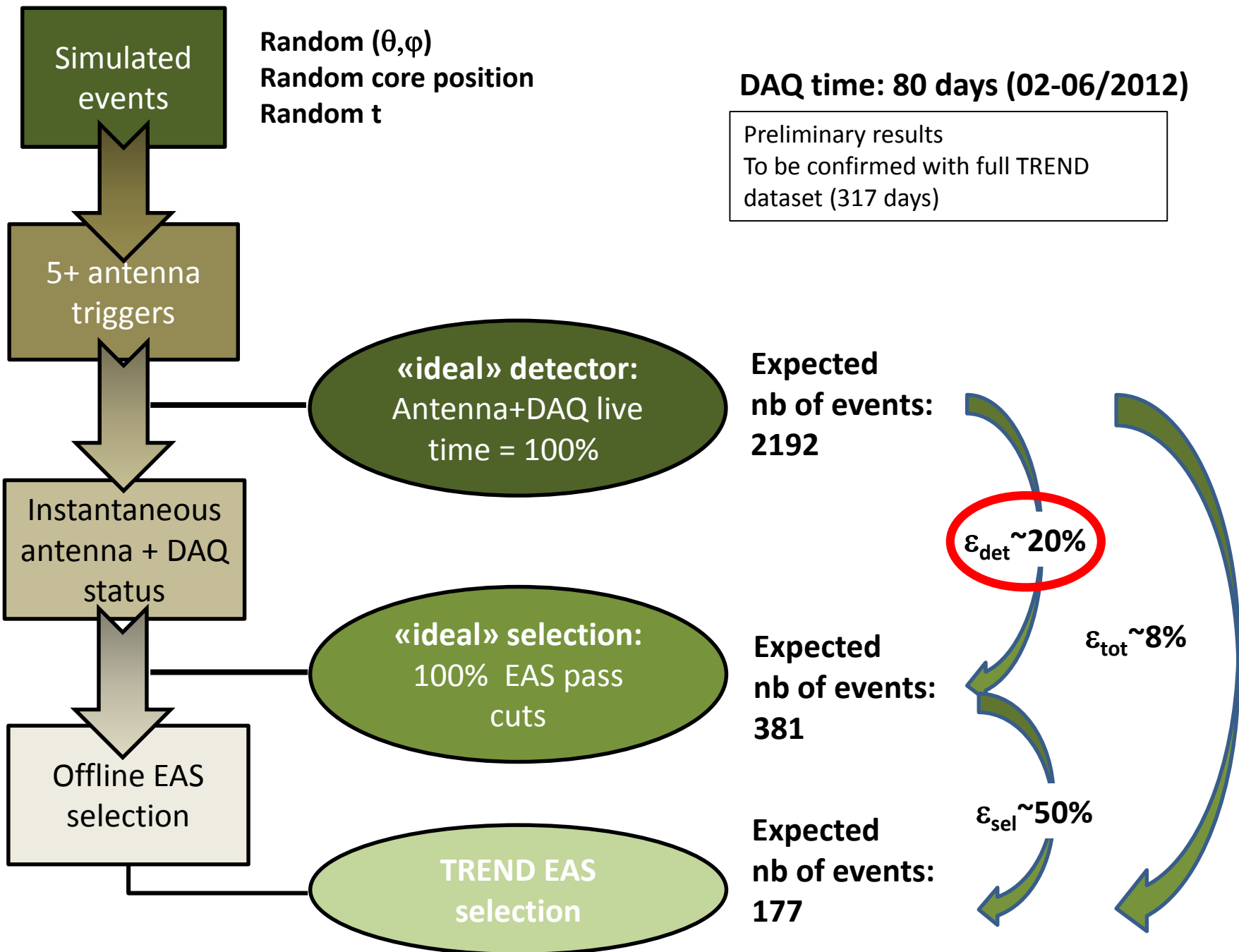
Zenith distribution



Azimuth distribution

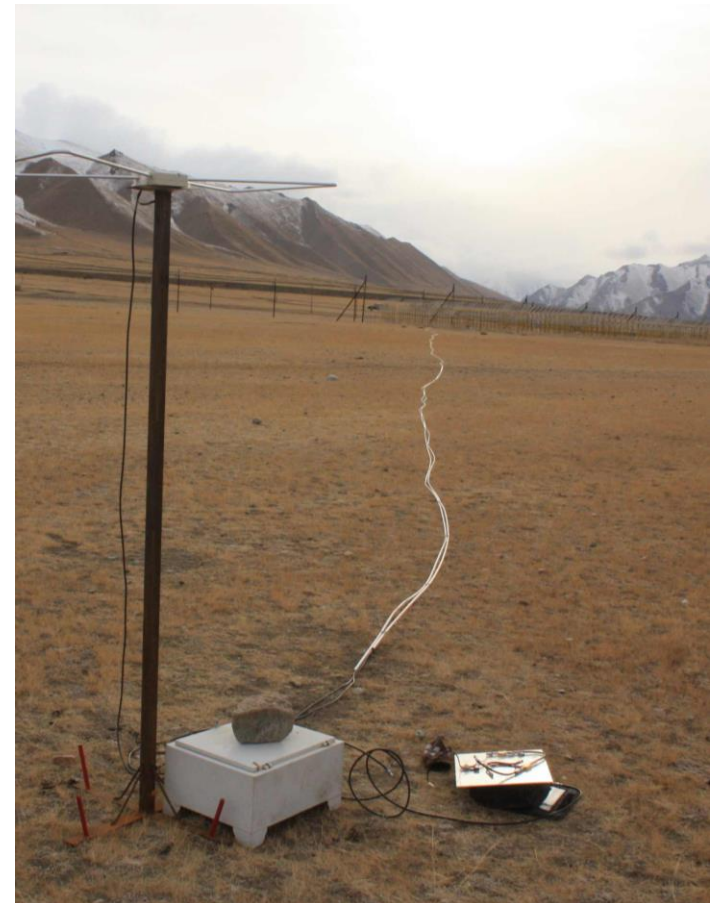
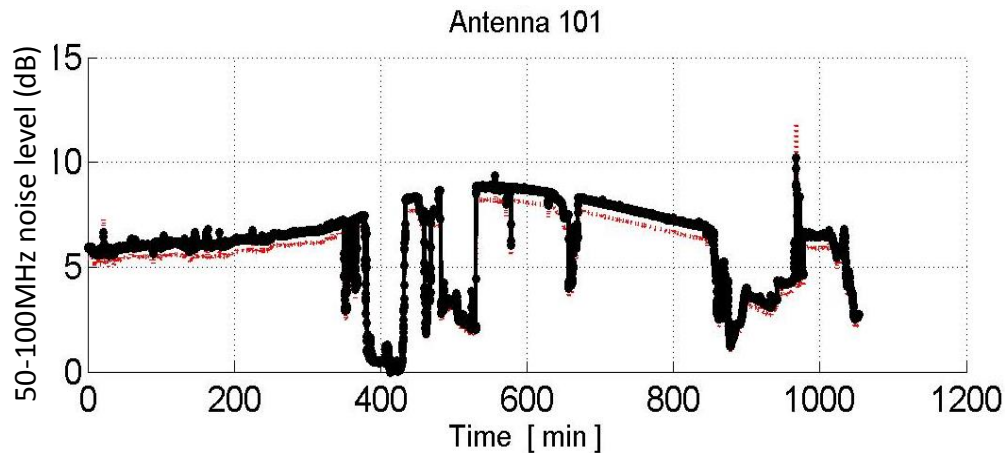


**TREND has performed autonomous detection of air showers from ultra-dominant background with limited background contamination (<30%)**



# TREND issues

- «You get what you pay for»: **system reliability** questionable
  - Sudden drops in gain
  - Aging (antennas, amplifiers, optical system, computers...)

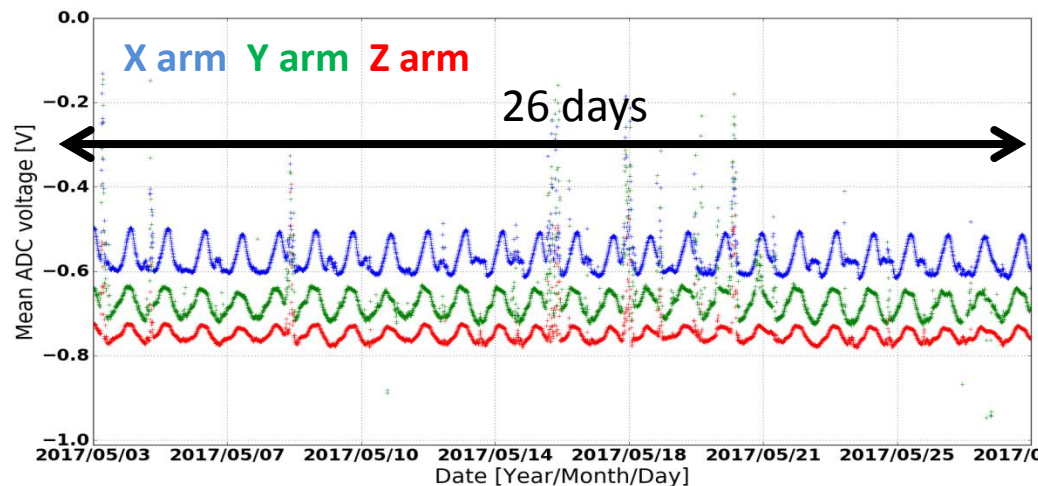
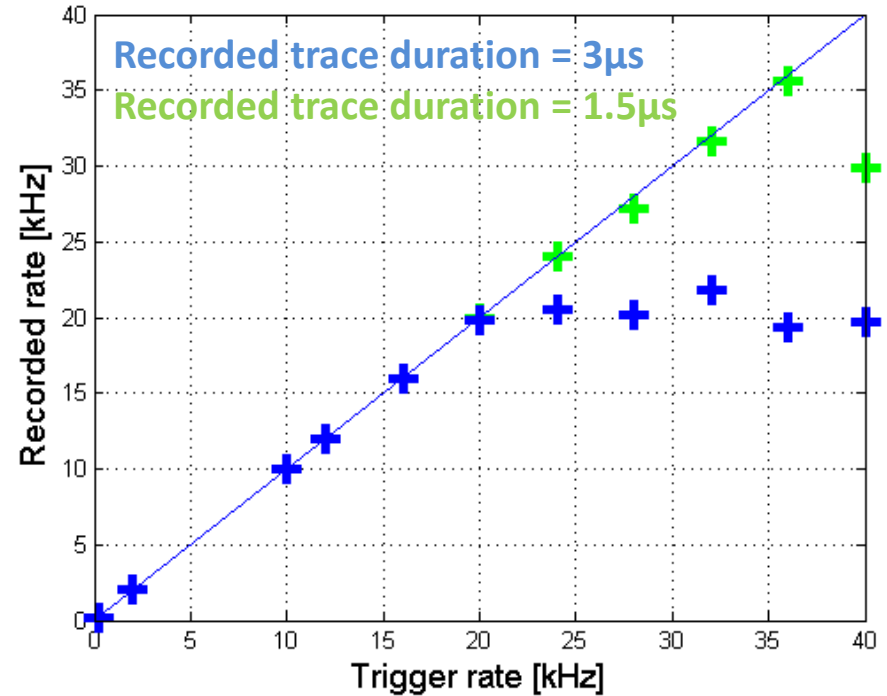


# Next steps

- **GRANDproto35**: dedicated setup of 35 antennas deployed on the TREND site. Goal: optimized detection efficiency & air shower purity

- Polarization measurement (3 arms antenna) brings additional tools for background rejection
- Dedicated electronics & DAQ (developped at LPNHE) should allow  $\sim 100\%$  live time.

➔ Expecting improved detection efficiency & background rejection compared to TREND.



# Conclusion

- TREND accomplished its initial goal: performing successful detection & identification of air shower with an autonomous radio array. Paper in preparation.
- Limited detection efficiency & sample purity due to flows in TREND setup. Should be fixed on GRANDproto35 (starting early 2018), a first step towards GRAND [Zilles, Friday PM].

谢谢



# Cut efficiency

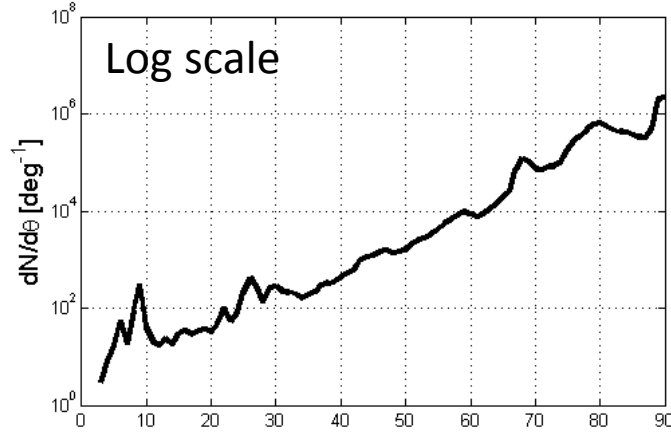
Cut	% survival	$N_{\text{coincs}} \text{ final}$	% survival [simu]	$N_{\text{coincs}} \text{ [simu]}$ (not E weighted)
« 50Hz » cut	8%	$6.3 \cdot 10^7$	82%	1526
Pulse duration	46%	$2.9 \cdot 10^7$	86%	1312
Multiplicity > 4	51%	$1.4 \cdot 10^7$	<b>80%</b>	1050
Valid reconstruction	67%	$1.0 \cdot 10^7$	99%	1040
Radius > 3000m	35%	$3.4 \cdot 10^6$	95%	988
$\Theta < 80^\circ$	30%	$1.0 \cdot 10^6$	98%	968
Trigger pattern/ Extension	32%	$3.3 \cdot 10^5$	88%	855
Neighbours (direction)	<b>0.3%</b>	1440	85%	725
Neighbours (time)	40%	<b>574</b>	92%	<b>668</b>

**Simulated EAS selection cuts survival rate: 37%** (46% if multiplicity cut excluded)

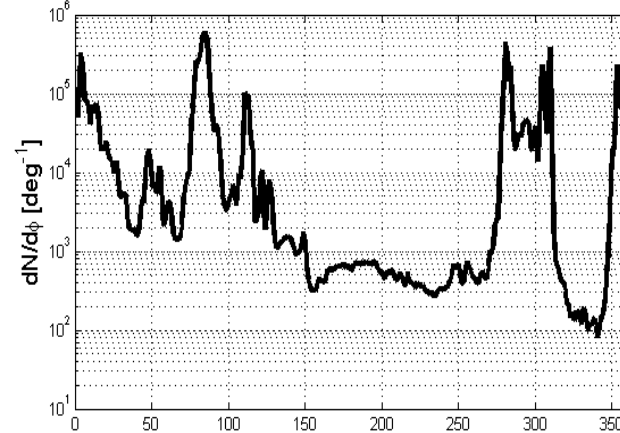


# Radio background

Zenith distribution

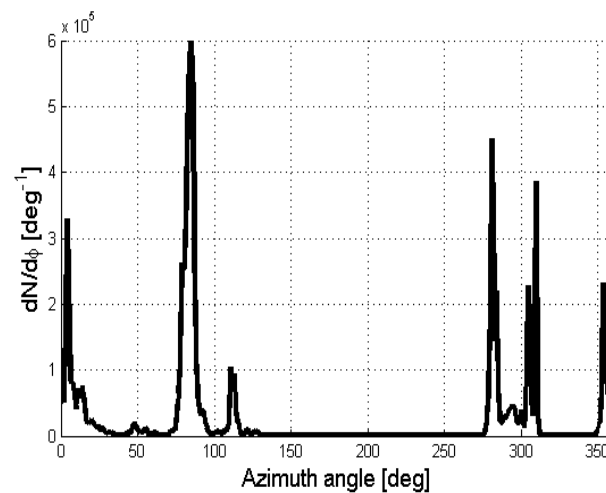
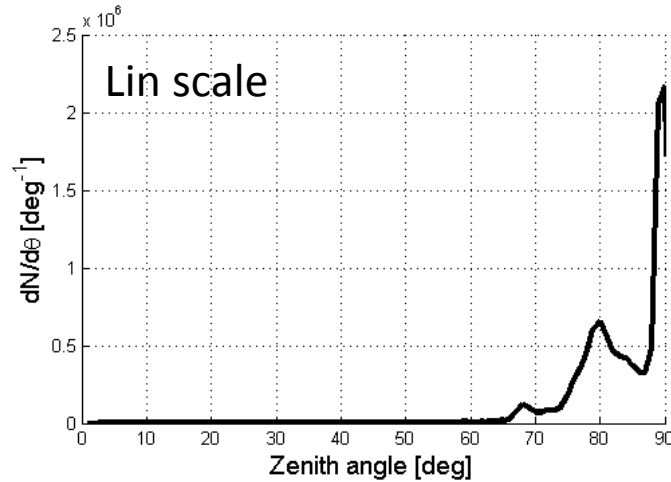


Azimuth distribution



- Here selecting plane wavefronts only (curvature radius > 500m)
- Most events coming from specific directions (N,E,W) along horizon (static noise sources)

North West South East



# Direction distribution

- Shapes of zenith and azimuth distributions of **experimental radio candidates** and **simulated EAS** are compatible.
- Excess of experimental data at large zenith angles & towards North & West ( $\Leftrightarrow$  standard background sources)
- ➔ Excess most likely corresponds to background events passing cuts.
- ➔ **Still to be refined, but already TREND has performed autonomous detection & identification of EAS, sample purity  $\sim 70\%$ .**

