



# Completing & Improving the TeV Cosmic-Ray Sky with HAWC & IceCube

---

TeVPA 2017  
Columbus, Ohio

Dan Fiorino<sup>\*d</sup>  
Juan Carlos Díaz Velez<sup>ab</sup>  
Paolo Desiati<sup>b</sup>  
Markus Ahlers<sup>c</sup>

<sup>\*</sup>Centro Universitario de los Valles, Universidad de Guadalajara, Guadalajara, Jalisco, México  
<sup>b</sup>Wisconsin IceCube Particle Astrophysics Center (WIPAC) and Department of Physics, University of Wisconsin-Madison, Madison, WI 53706, USA  
<sup>c</sup>Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark  
<sup>d</sup>Department of Physics, University of Maryland, College Park, MD, USA



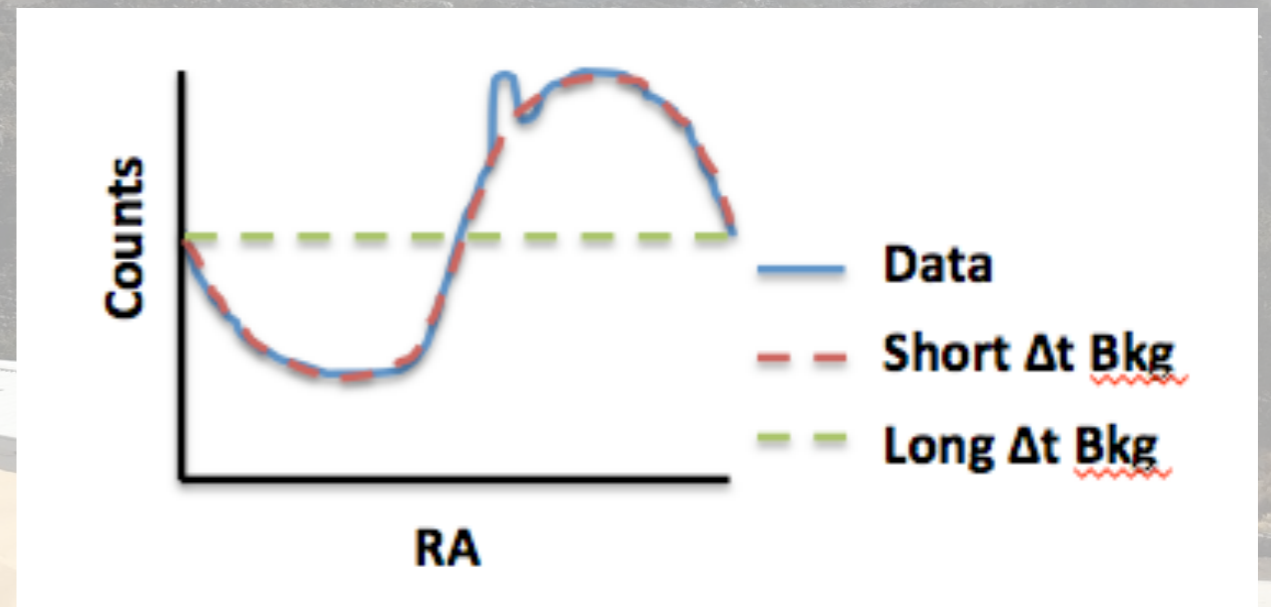
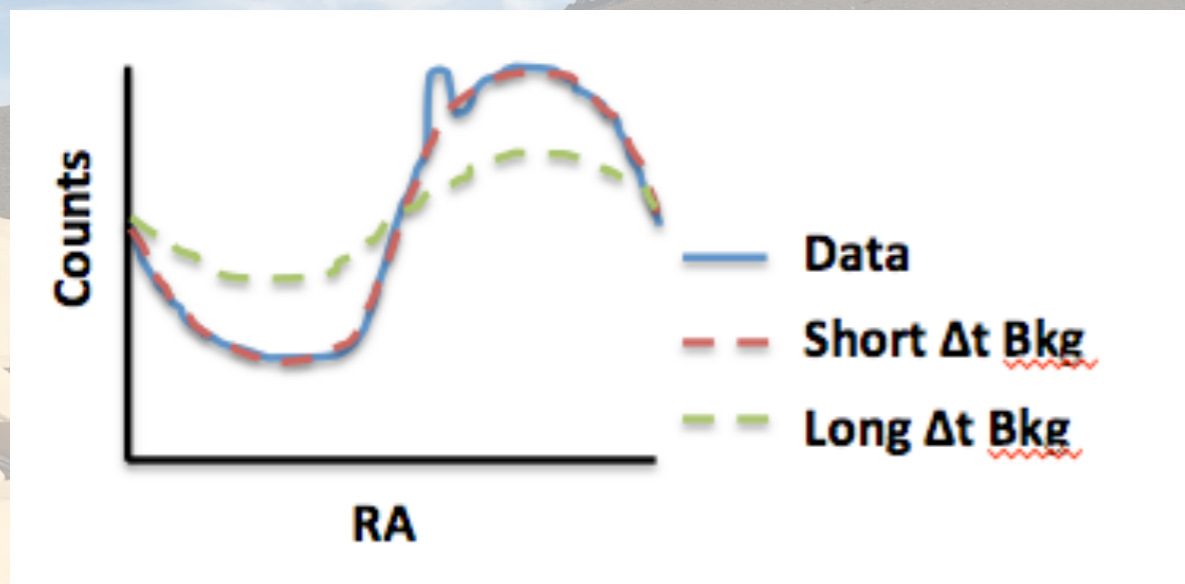
# The TeV Cosmic-Ray Sky

- Difficult to measure
  - Only ground-based detectors have detections
  - Background and weak signal are intertwined spatially
  - Ground detectors are currently limited to only measuring anisotropies in direction of Earth's rotation (RA)
- Hard to interpret
  - Need complete sky to properly measure
  - Missing anisotropy along Declination ( $m=0$  modes)
  - Are deficits and excesses equally important?
- Recent Developments
  - Maximum-likelihood technique which uses direct integration to find unbiased background
  - Cosmic-ray energy estimation for ground detectors is evolving. Not just using number of hit detectors
  - Complete two-dimensional descriptions



# Unbiased Background Estimation

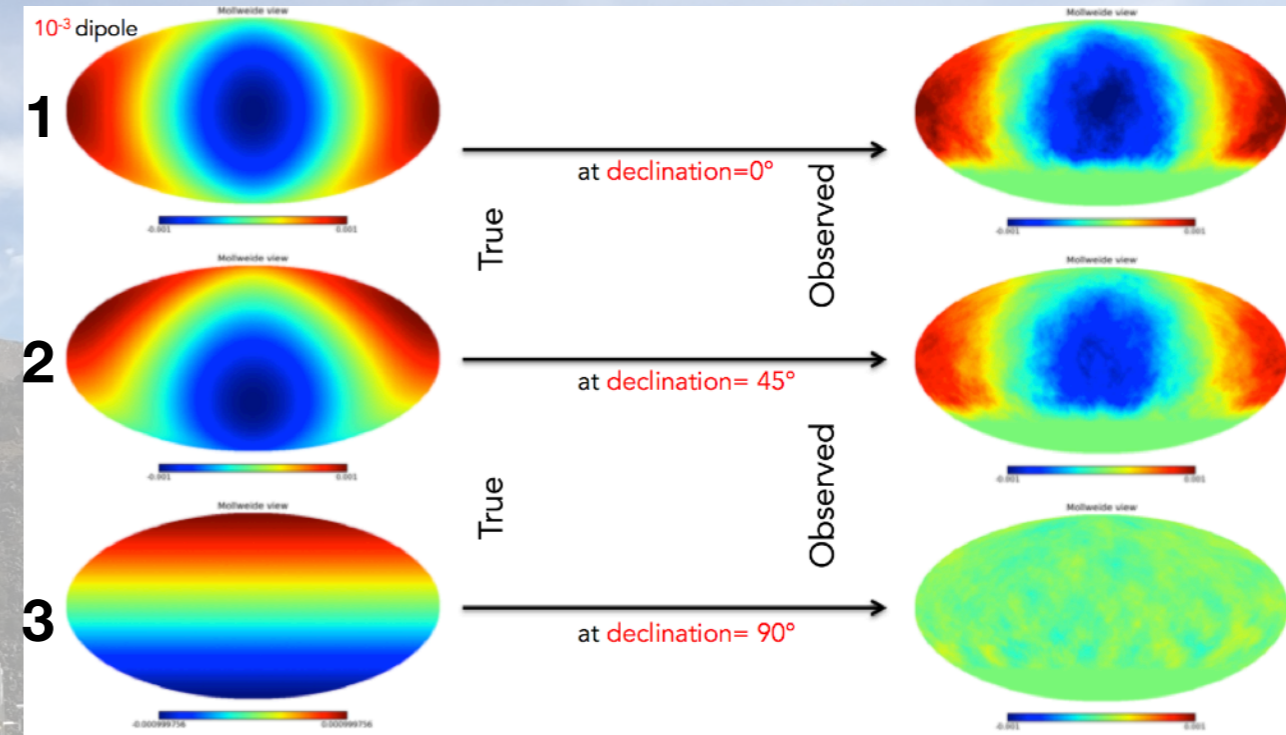
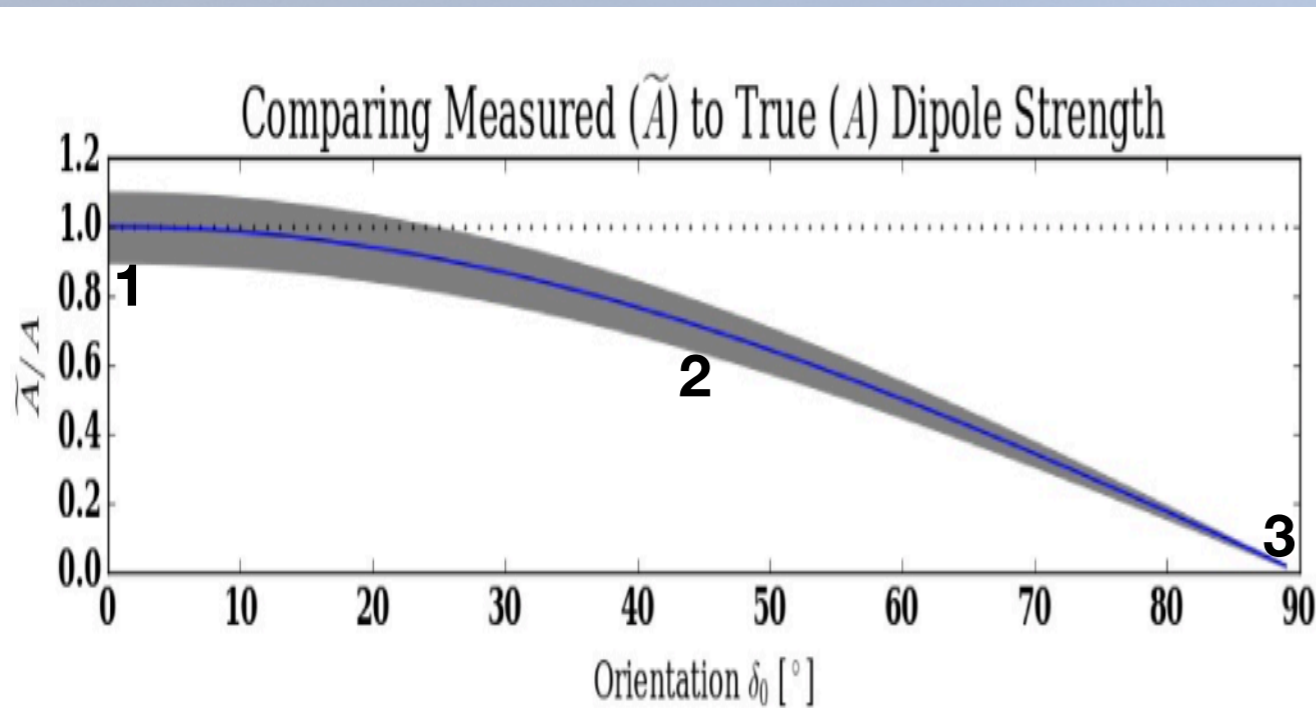
- For a detector with a constant acceptance as a function of time, the expected result of an isotropic flux would be ... well ... isotropic!
- Adding a dipolar signal to the isotropic flux will bias background estimation methods that are based upon data (e.g. time-scrambling, direct integration) and over long periods (Long  $\Delta t$ ).



- With an iterative approach, this bias can be removed (right diagram) by fitting the detector response and anisotropy simultaneously.



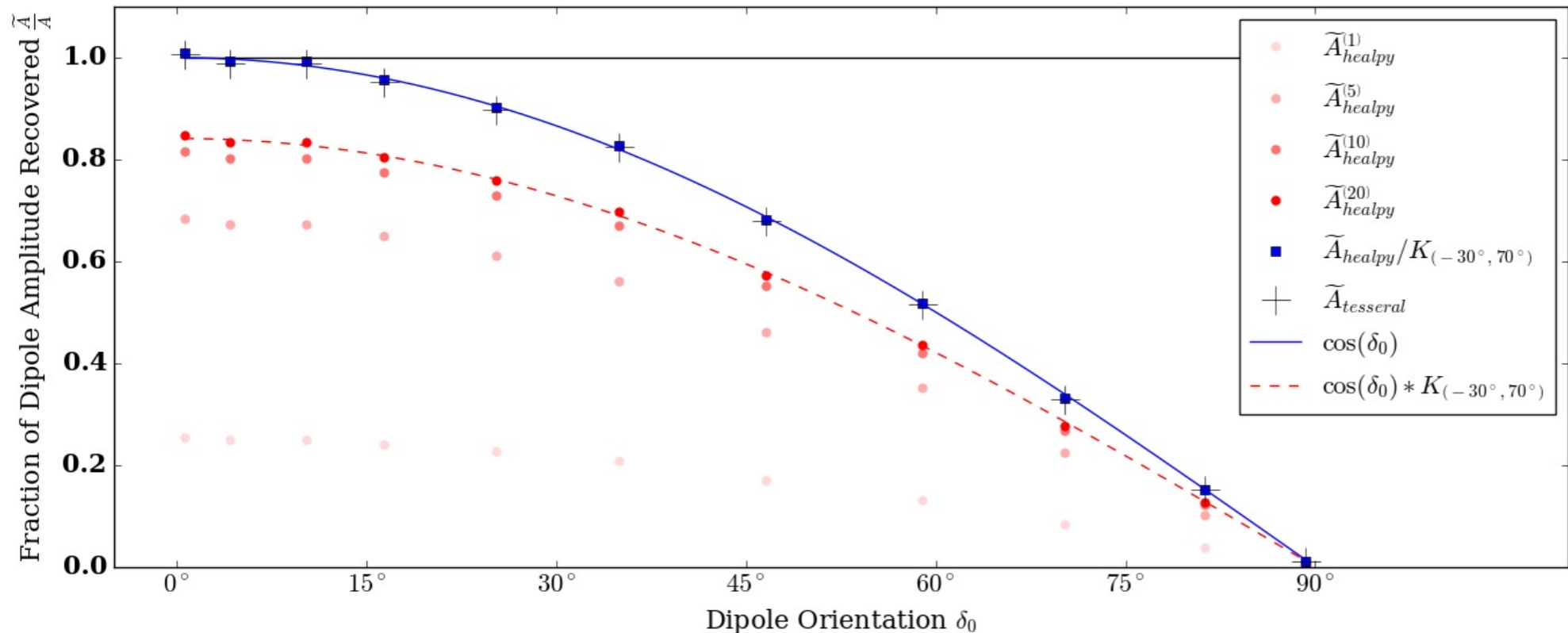
# Limitations (Ground-based)



- Experiments are limited currently to measuring anisotropies in direction of Earth rotation (RA)
- For a pure dipole anisotropy, this greatly affects the measurable strength
- Measured strength goes to zero as the orientation goes to the poles (90°)



# Unbiased Background / Limitations



- Simulated dipole anisotropy reconstructed using method of (M. Ahlers et al. 2016)
- Blue line is the best one can do with ground-based instrument

- Method improves with iteration (light to dark red)

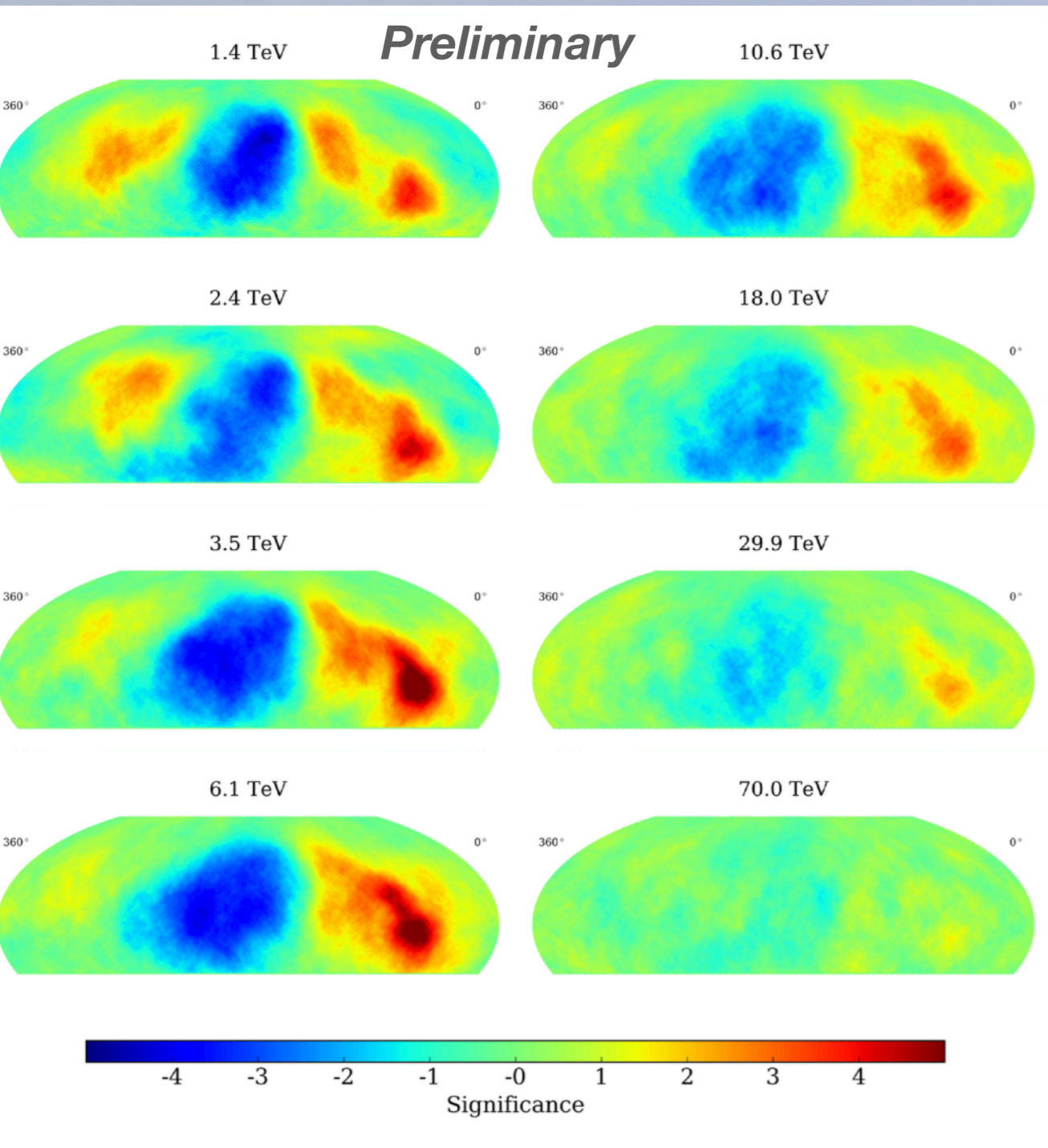
- healpy fit must be corrected because of limited sky coverage (blue)

- Red/blue points show fit from healpy software

**healpy:** HEALPix + Python  
**HEALPix:** sphere pixelation + sky map routines



# HAWC Results



- Background method applied to 400 full, sidereal days of HAWC data
- Applied strong cut (2% pass) to reach unprecedented energy resolution for this measurement. Harsh cuts provide highly diagonalized mixing matrix (verified on CR spectrum/Moon shadow).
- Described using full 2D dipole fit (truncated)

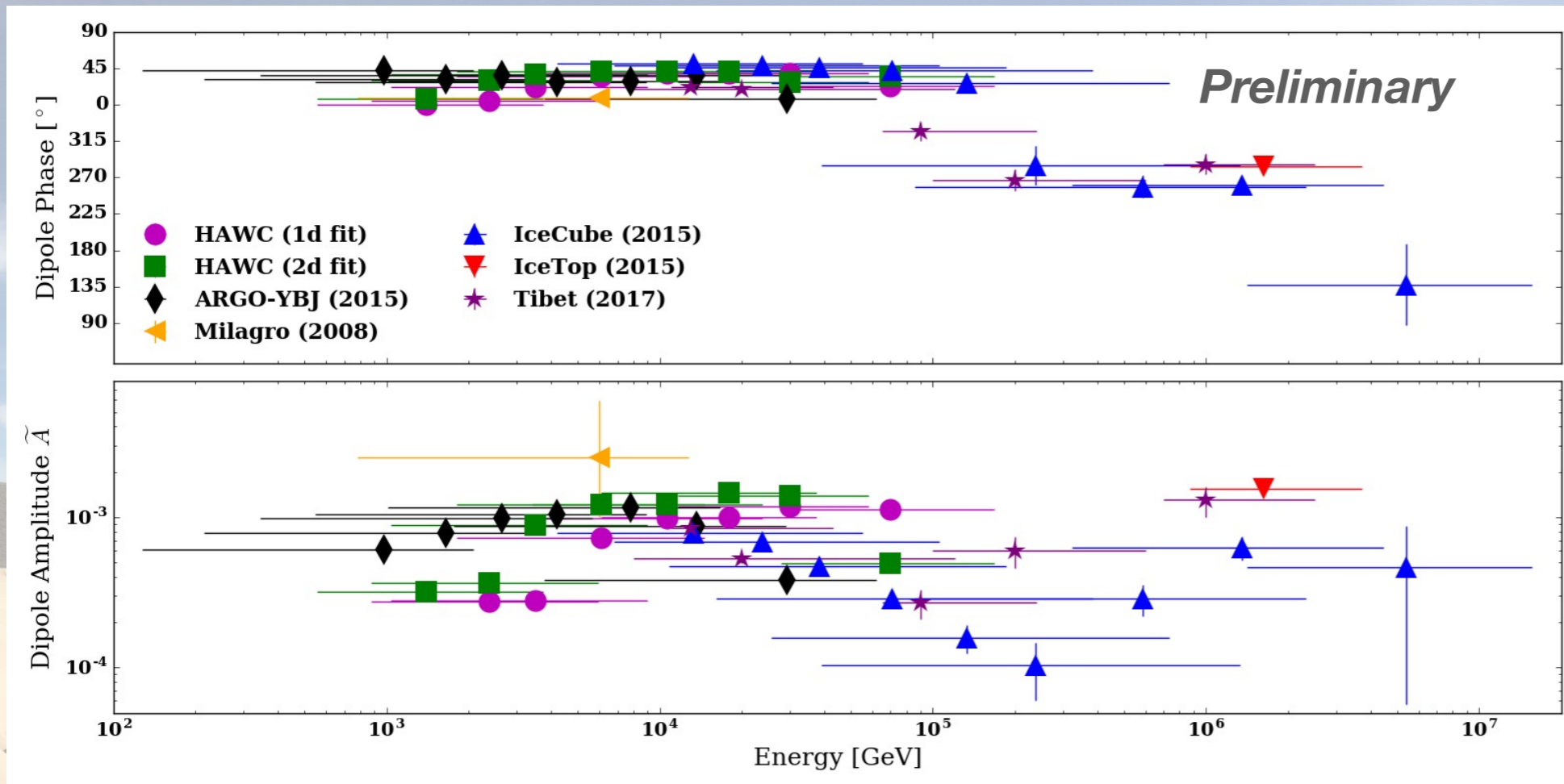
Energy [TeV]	Amplitude [ $\times 10^{-4}$ ]	Phase	$a_{1,1}$ [ $\times 10^{-4}$ ]	$a_{1,-1}$ [ $\times 10^{-4}$ ]
1.4 $\left(\begin{smallmatrix} -0.8 \\ +2.4 \end{smallmatrix}\right)$	$7.5 \pm 0.6$	$33.3^\circ \pm 4.4^\circ$	$-18.1 \pm 1.7$	$-11.9 \pm 1.7$
2.4 $\left(\begin{smallmatrix} -1.5 \\ +3.6 \end{smallmatrix}\right)$	$6.4 \pm 0.3$	$46.6^\circ \pm 2.2^\circ$	$-12.7 \pm 0.7$	$-13.5 \pm 0.7$
3.5 $\left(\begin{smallmatrix} -2.5 \\ +5.5 \end{smallmatrix}\right)$	$10.2 \pm 0.2$	$42.8^\circ \pm 1.4^\circ$	$-21.6 \pm 0.7$	$-20.1 \pm 0.7$
6.1 $\left(\begin{smallmatrix} -4.3 \\ +8.5 \end{smallmatrix}\right)$	$13.2 \pm 0.3$	$44.0^\circ \pm 1.2^\circ$	$-27.5 \pm 0.8$	$-26.6 \pm 0.8$
10.6 $\left(\begin{smallmatrix} -7.2 \\ +13.1 \end{smallmatrix}\right)$	$13.2 \pm 0.4$	$45.0^\circ \pm 1.6^\circ$	$-27.1 \pm 1.1$	$-27.1 \pm 1.1$
18.0 $\left(\begin{smallmatrix} -11.9 \\ +19.6 \end{smallmatrix}\right)$	$15.6 \pm 0.5$	$42.9^\circ \pm 1.8^\circ$	$-33.0 \pm 1.4$	$-30.7 \pm 1.4$
29.9 $\left(\begin{smallmatrix} -18.2 \\ +28.4 \end{smallmatrix}\right)$	$14.4 \pm 0.7$	$29.9^\circ \pm 2.6^\circ$	$-36.2 \pm 1.9$	$-20.8 \pm 1.9$
70.0 $\left(\begin{smallmatrix} -42.1 \\ +97.9 \end{smallmatrix}\right)$	$5.4 \pm 0.7$	$41.9^\circ \pm 7.0^\circ$	$-11.7 \pm 1.9$	$-10.4 \pm 1.9$







# Comparison to other experiments



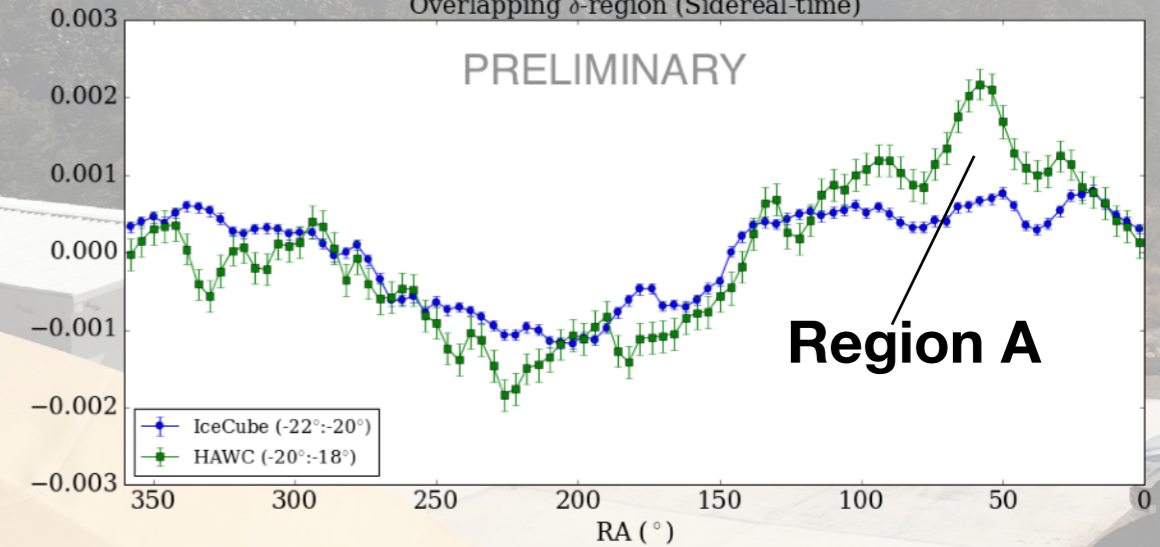
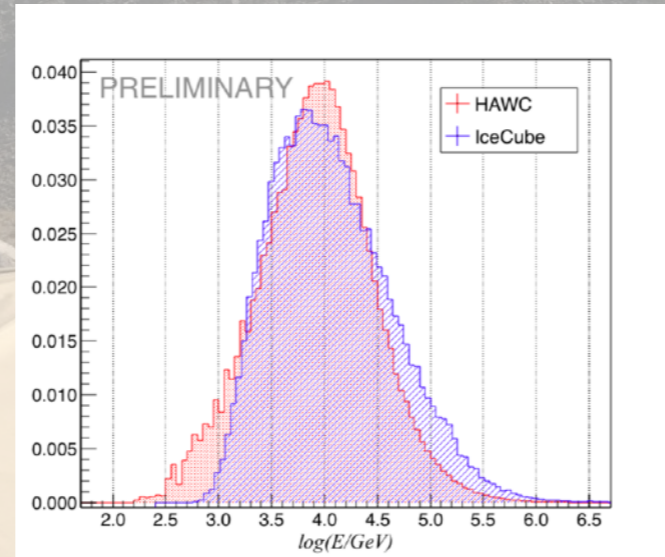
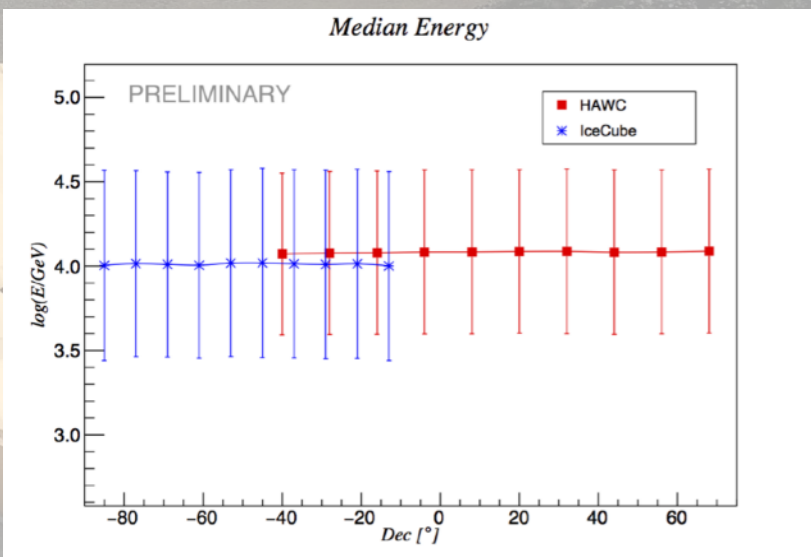
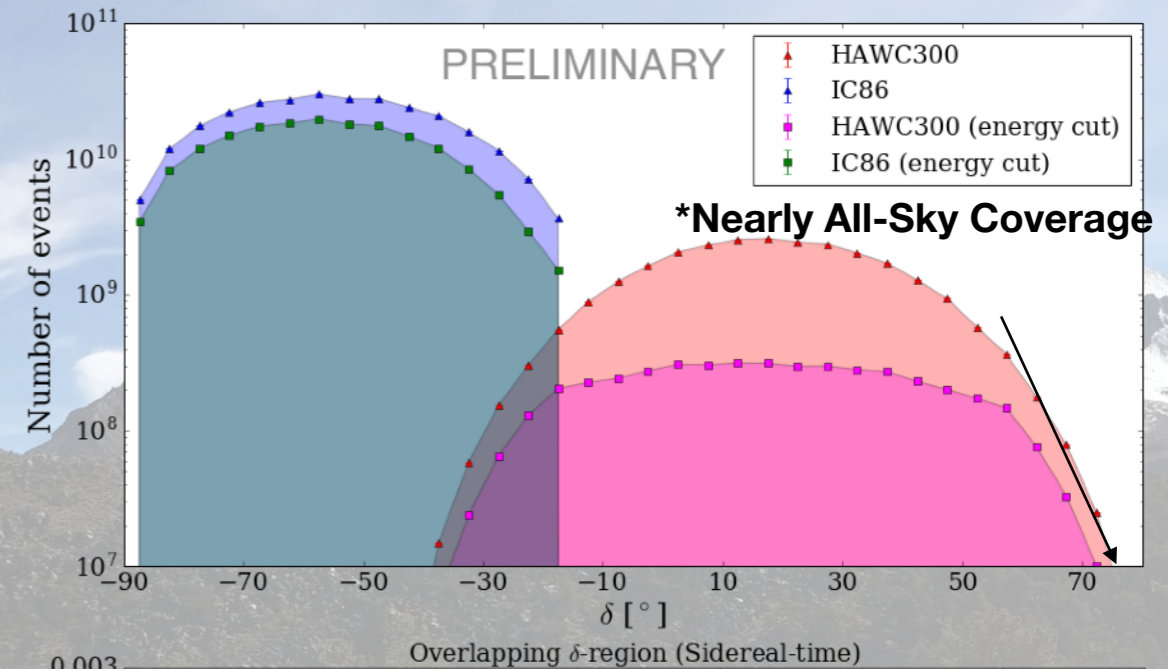
- Cause of Discrepancies?
- Energy scale?  $\sim 10\%$
- Use of 2D Fit?
- CR composition sensitivity? (IceCube is underground...)

- Compares favorable with other experiments when considering uncertainty in energy



# 1st Ever All-Sky\* TeV Cosmic-Ray Map

- Combining data from HAWC (North) and IceCube (South) gives almost full sky coverage
- Data is energy-matched to be meaningful
- Same method as with HAWC-only data



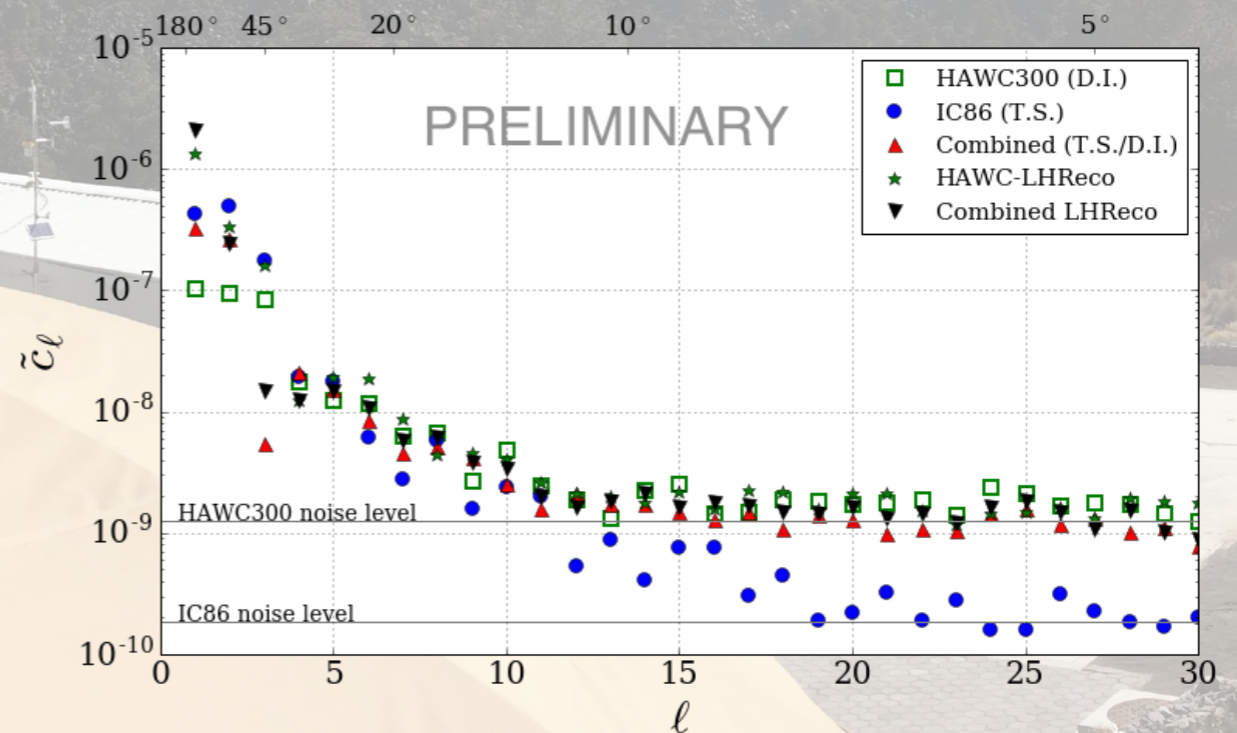
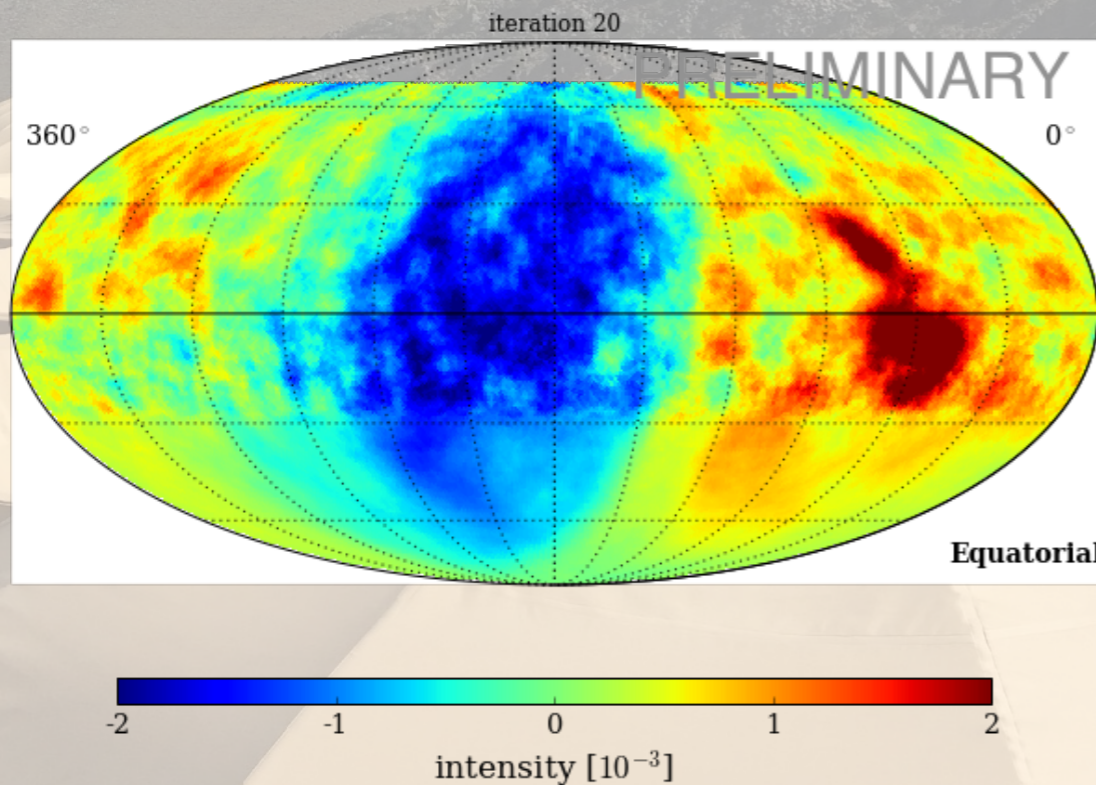
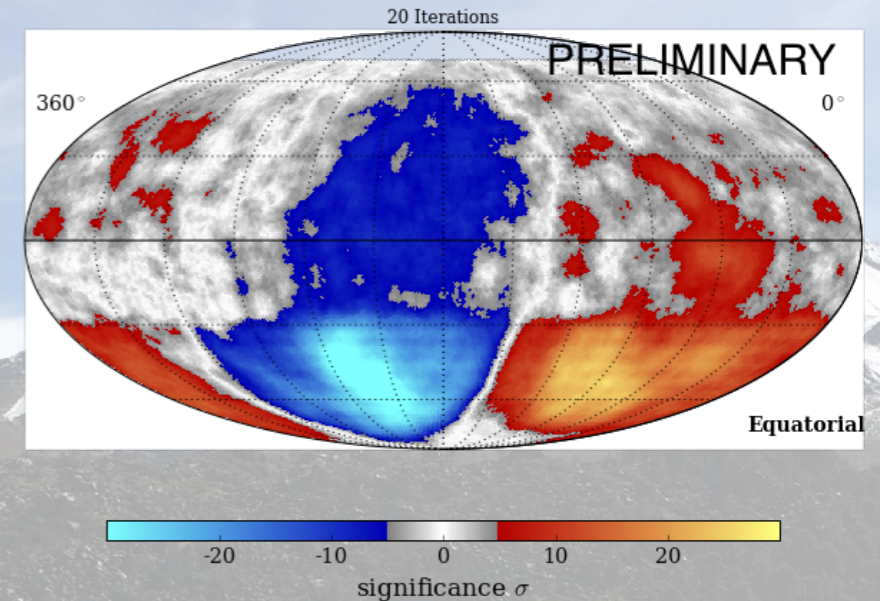
plots taken from: ICRC 2017 Contribution by Diaz Velez, J.C., et al. (PoS 539)  
 ‘Combined Analysis of Cosmic-Ray Anisotropy with 2 IceCube and HAWC’





# 1st Ever All-Sky\* TeV Cosmic-Ray Map

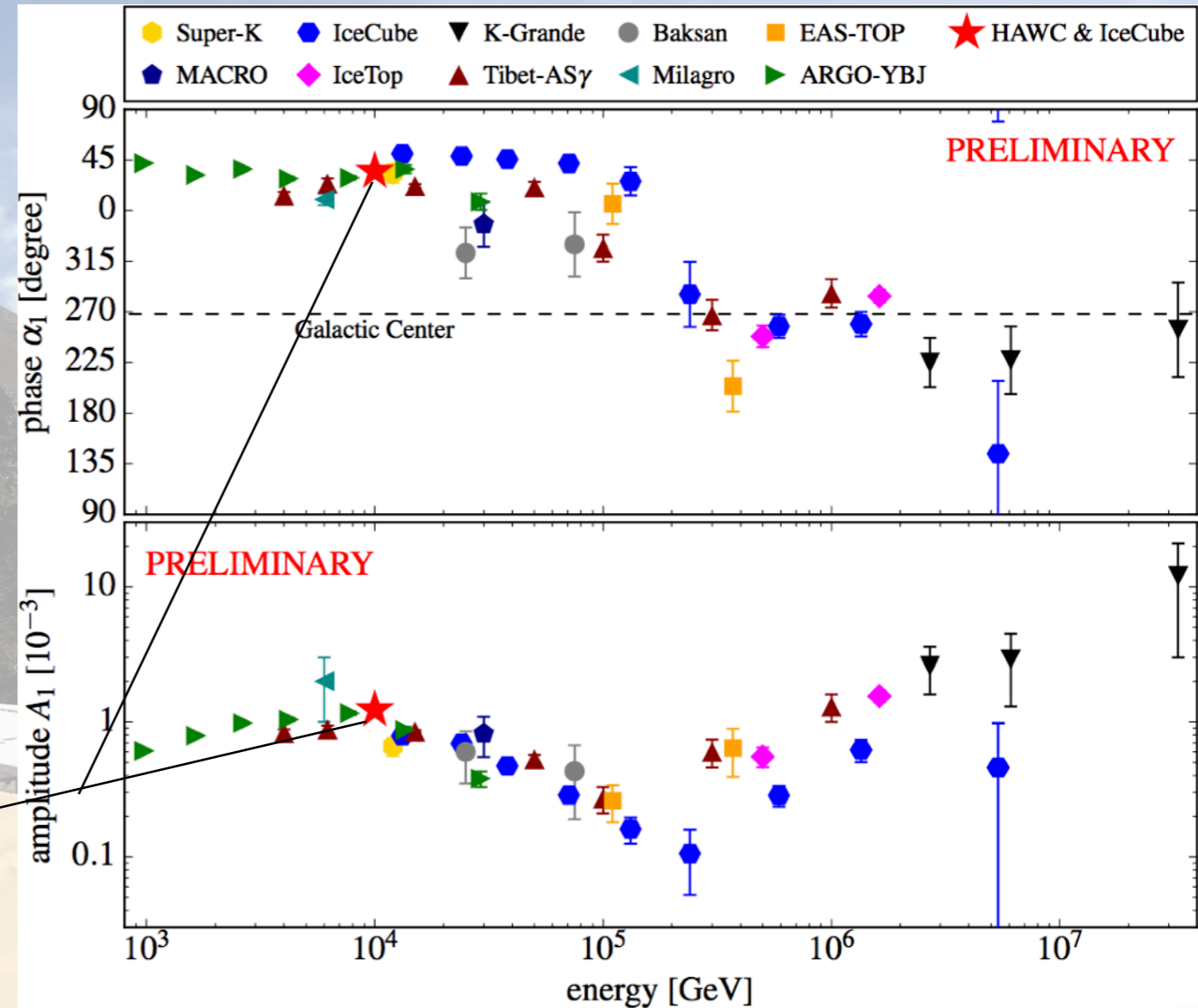
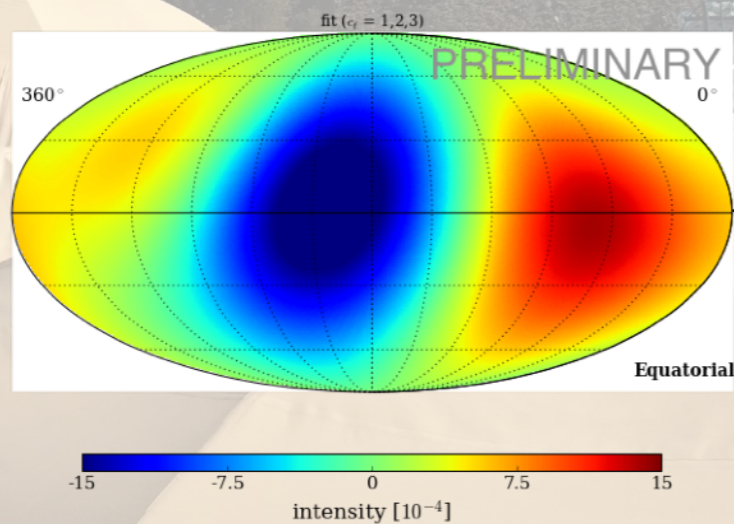
- Dipolar feature connects well,
- Region A does not continue South at these energies
- IceCube features are more significant (more data)
- Combined map has more power in the dipole due to better sky coverage (most of this comes from the HAWC data)





# 1st Ever All-Sky\* TeV Cosmic-Ray Map

- Combined map is then fit with truncated spherical harmonic series ( $l \leq 3$ )
- Dipole amplitude and phase can be compared against other experiments

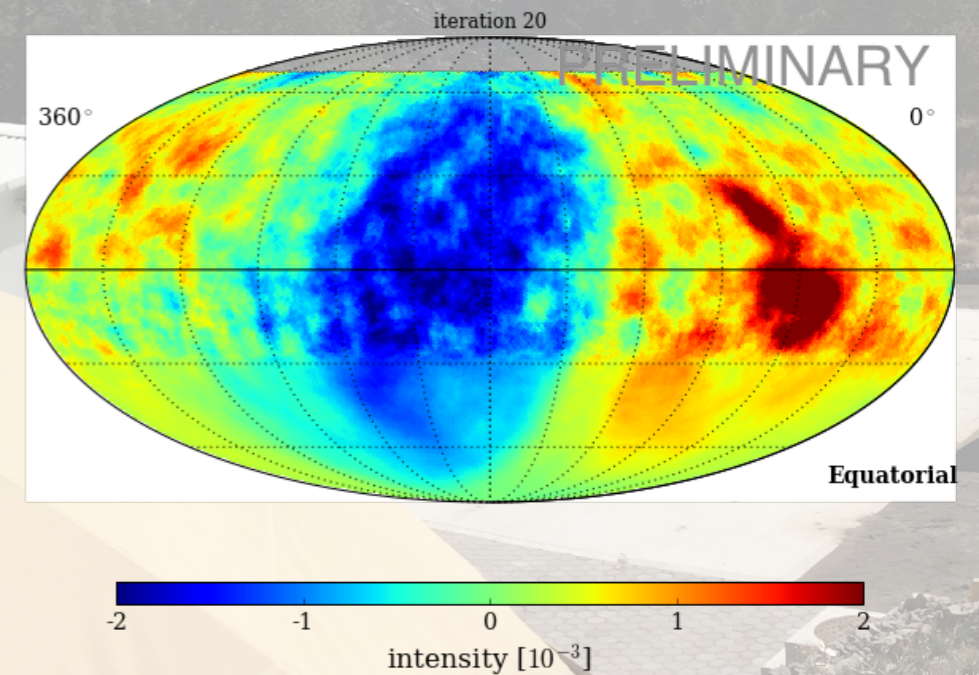
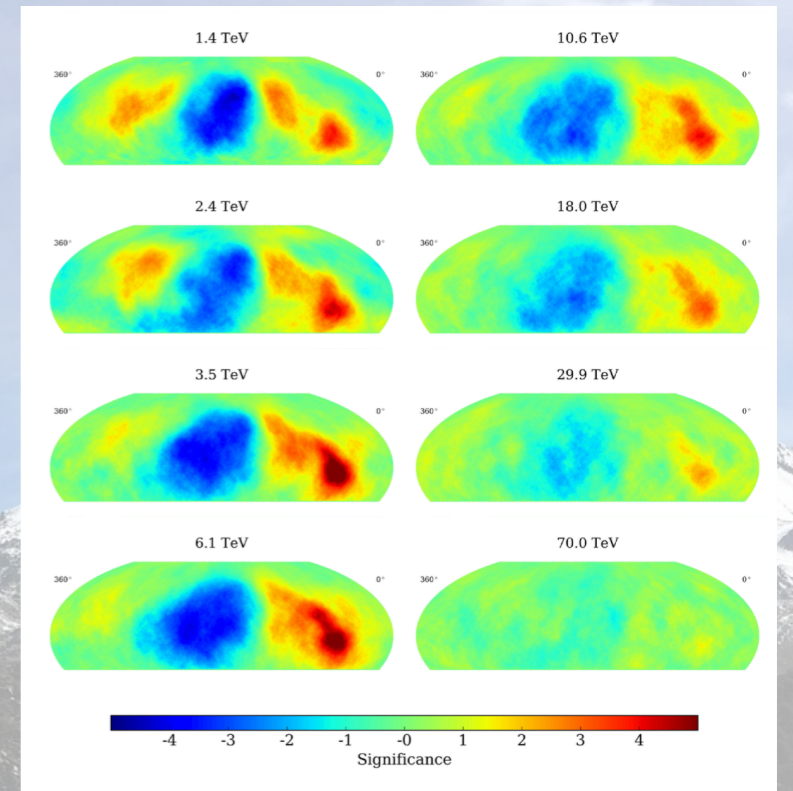


adopted from [M. Ahlers et al. ArXiv:1612.01873](https://arxiv.org/abs/1612.01873)



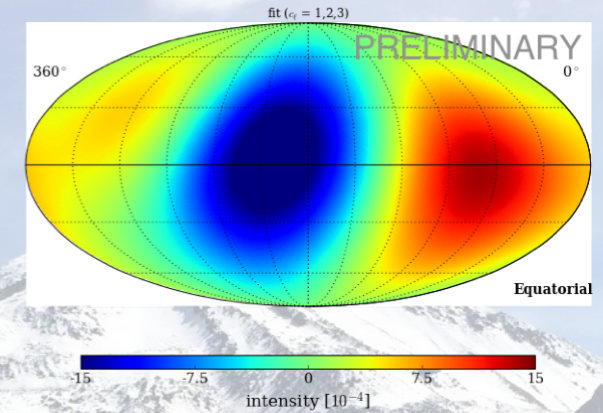
# Summary

- Our descriptions of the TeV Sky are improving!
  - Background techniques are mature
  - Energy estimation is maturing
- Results are being reported with full two-dimensionality
- A full-sky\* picture is available!

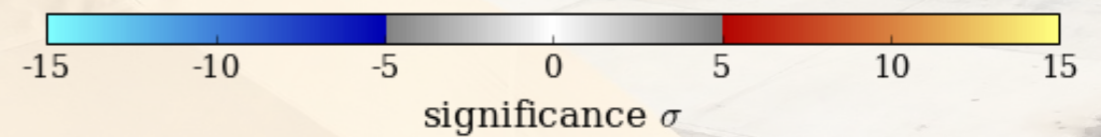
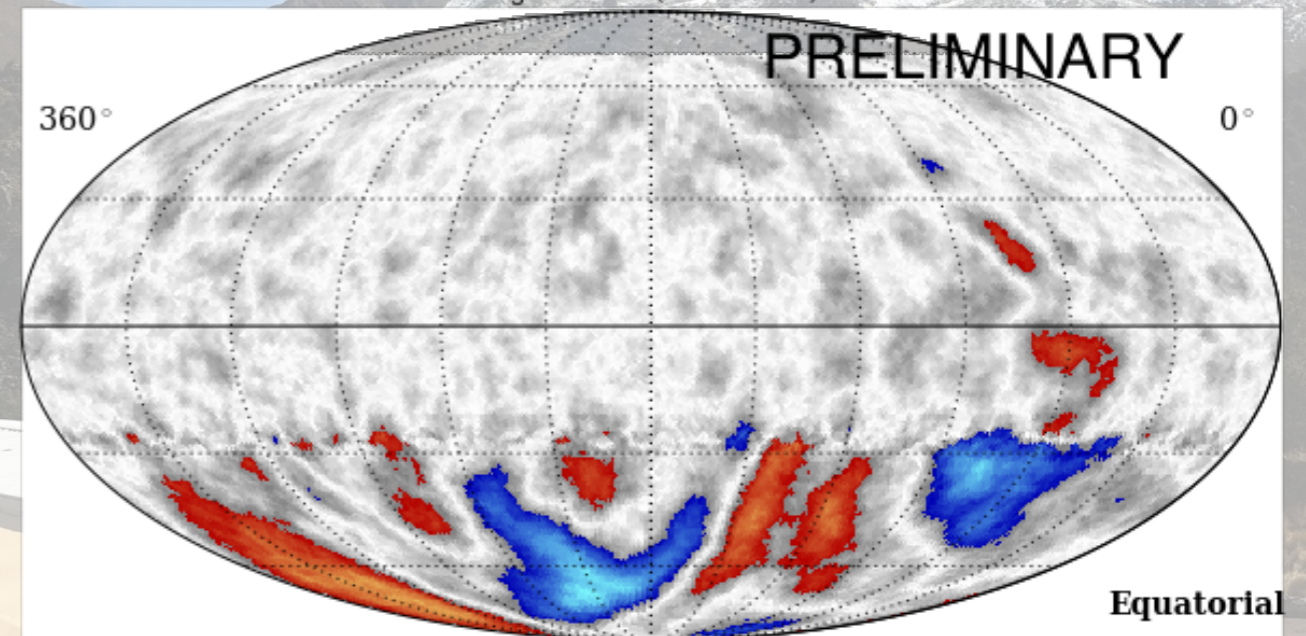




# Combined Sky Map Large-scale Removed (right)



Significance (small scale)



subtracted fit ( $c_i = 1, 2, 3$ )

