



A Unitarity Bound for Type-1 Seesaw Models

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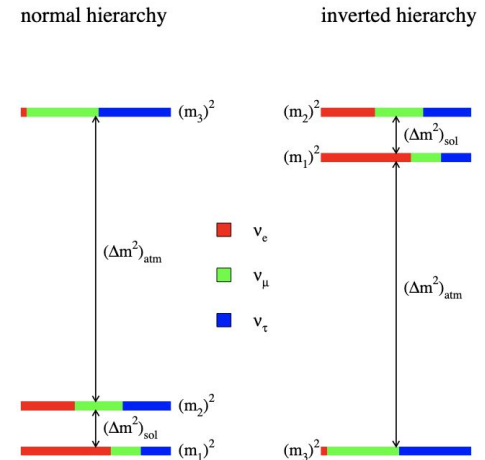
Outline:

1. Neutrino Oscillation and Mass Models overview
2. Simulating New Physics In Madgraph
3. Bolster analysis with Unitarity Bound

Neutrino Masses

- Neutrinos must have distinct masses to oscillate between lepton flavor eigenstates
- Mass eigenstates different from Flavor eigenstates
- Mass states evolve over distances
- Transform between mass and flavor bases with Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix
- Current data constrains “mixing angles”, yielding two possible “mass orderings”
- Still unknown: Origin of Neutrino Masses

$$|\nu_i(t, \vec{x})\rangle = e^{-ip_i x} |\nu_i(0, \vec{0})\rangle.$$





Theory: Type 1 Seesaw Models

- Lagrangian:
$$\mathcal{L}_N = -\bar{L} Y_\nu^D \tilde{H} N_R - \frac{1}{2} \overline{(N^c)}_L M_R N_R + \text{H.c.}$$
- Mass Matrix:
$$\mathbb{N}^\dagger \begin{pmatrix} 0 & m_D \\ m_D^T & M \end{pmatrix} \mathbb{N}^* = \begin{pmatrix} m_\nu & 0 \\ 0 & M_N \end{pmatrix}$$
- The diagonalization matrix elements go as: $m_\nu \approx -m_D M_R^{-1} m_D^T$, $M_N \approx M_R$
- Elements of unitary diagonalization matrix yield mixing between SM and Majorana states



Type 1 Coupling to Standard Model

- In the lepton flavor basis, the coupling to the standard model takes the form:

$$\mathcal{L} \supset \frac{gU_\ell}{\sqrt{2}} (W_\mu \bar{l}_L \gamma^\mu N + \text{h.c.}) - \frac{gU_\ell}{2 \cos \theta_w} Z_\mu (\bar{\nu}_L \gamma^\mu N + \bar{N} \gamma^\mu \nu_L)$$



Features of Type 1:

- Mixing between Majorana Neutrinos and leptons must be small (empirically)
- Majorana Neutrinos can undergo lepton number-violating processes like neutrinoless double beta decay
- Large Majorana mass might be required to reproduce observed neutrino masses

Inverse Seesaw Extension

- Modification of Type 1 Seesaw
- Lagrangian:
 - X field is a Majorana singlet
 - μ_X is small
 - Smallness of μ_X allows for smaller mediator masses
- New mass matrix:
 - M_R is a diagonal 3x3 matrix with masses separated by:
 - Diagonalization yields mass eigenstates for each generation:

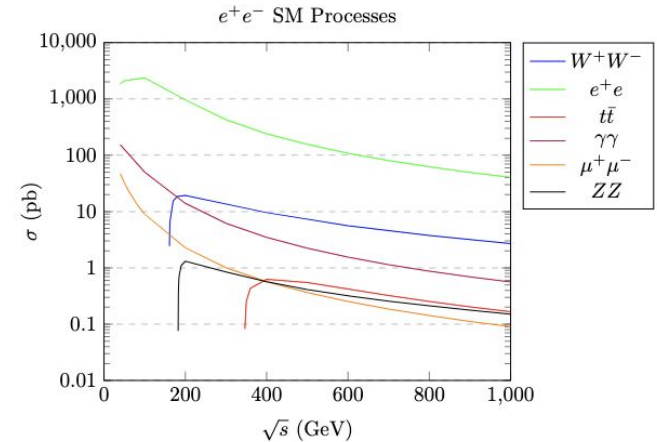
$$\mathcal{L}_{\text{ISS}} = -Y_\nu^{ij} \bar{L}_i \tilde{\Phi} \nu_{Rj} - M_R^{ij} \bar{\nu}_{Ri}^C X_j - \frac{1}{2} \mu_X^{ij} \bar{X}_i^C X_j + h.c.$$

$$\begin{pmatrix} 0_{3 \times 3} & m_{D_{3 \times 3}} & 0_{3 \times 3} \\ m_{D_{3 \times 3}}^T & 0_{3 \times 3} & M_R \\ 0_{3 \times 3} & M_R^T & \mu_X \end{pmatrix}$$

$$m_\nu \simeq \frac{m_D^2}{m_D^2 + M_R^2} \mu_X, \quad m_{N_1, N_2} \simeq \sqrt{M_R^2 + m_D^2} \mp \frac{M_R^2 \mu_X}{2(m_D^2 + M_R^2)}$$

Process Simulation Method:

- Tools:
 - Feynrules - Mathematica package that will take model parameters and generate Feynman Rules readable by an event generation tool
 - Madgraph - MonteCarlo event generation tool
- Example SM Processes:
- Goal: Use these tools to analyze BSM processes for specific collider conditions (e.g. pp collision at 13 TeV)



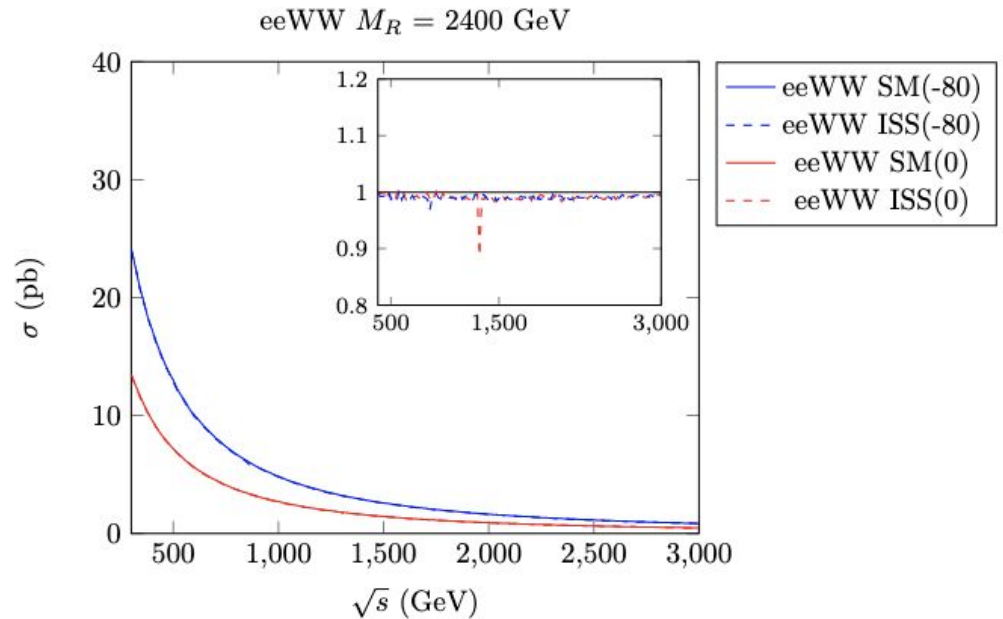


Testing Inverse Seesaw at a lepton collider (1)

- Consider two processes: $e^+ e^- \rightarrow W^+ W^-$ and $e^+ e^- \rightarrow W^+ W^- H$
- Fix $M_R = 2400$ GeV (all massive neutrinos have the same mass)
- Use numerically generated mass parameters and mixing angles
- Generate cross sections using Madgraph, scan over energy

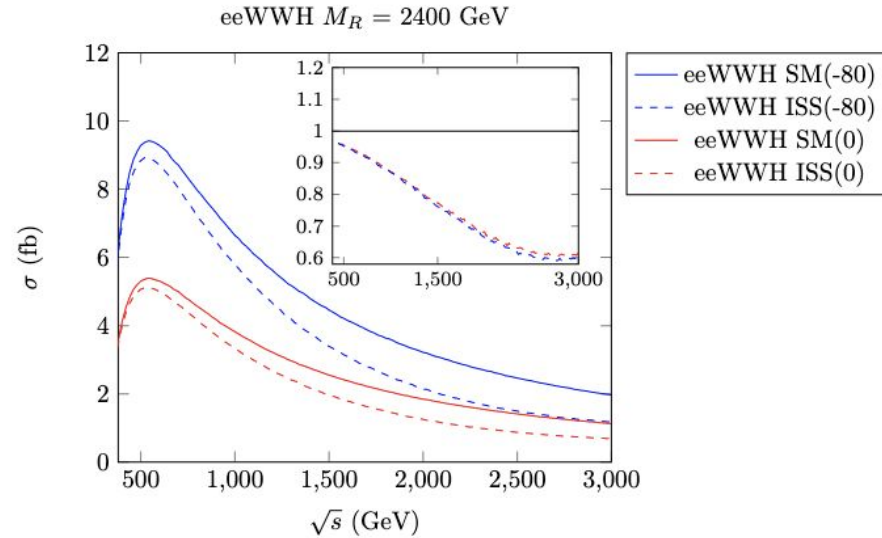
Testing Inverse Seesaw at a lepton collider (2)

- Cross section is larger, but signal is not significant



Testing Inverse Seesaw at a lepton collider (3)

- Discernible signal
 - Drawback: Third-order process suppresses the cross section significantly





Unitarity and the S-wave constraint

- Unitarity

- Requirement for the time evolution operator
- Deviation from unitarity:
 - Inner product no longer preserved
 - Non-physical matrix elements

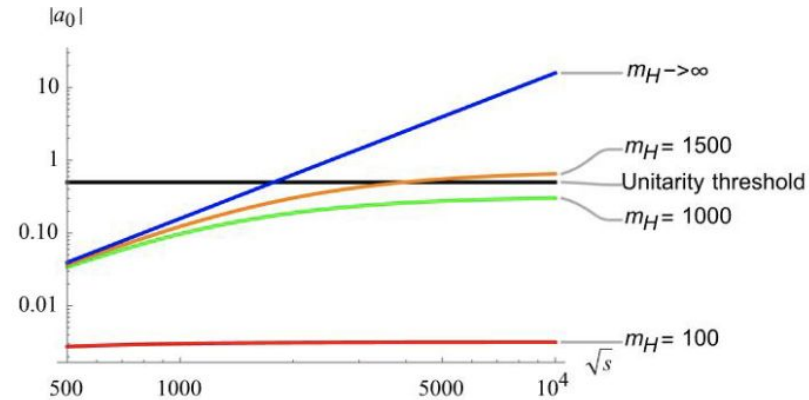
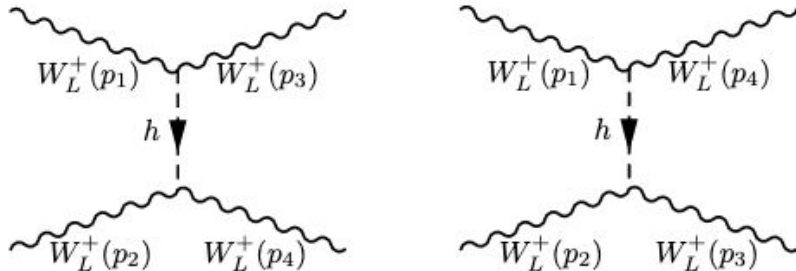
- Angular Momentum Spherical Wave Decomposition: $a_l(s) = \frac{1}{64\pi} \int_{-1}^1 (i\mathcal{M}(s, t)) P_l(\cos \theta) d \cos \theta$

- S-wave for $l = 0$

- Thus: $a_0 = \frac{1}{64\pi} \int_{-1}^1 d \cos \theta (-i\mathcal{M})$

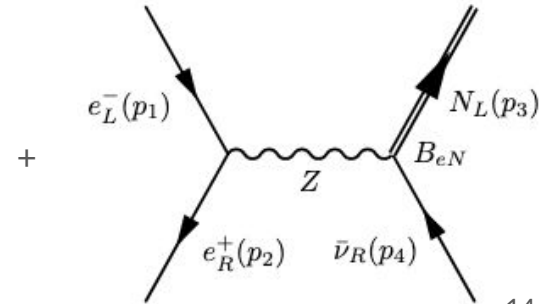
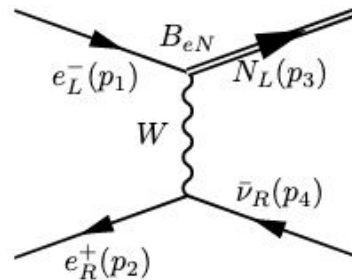
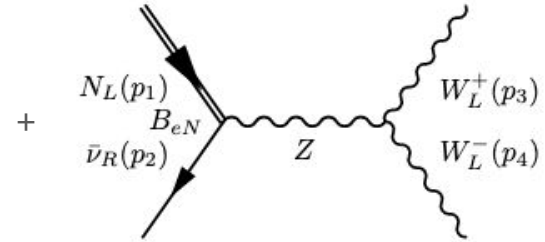
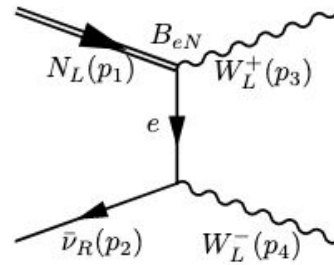
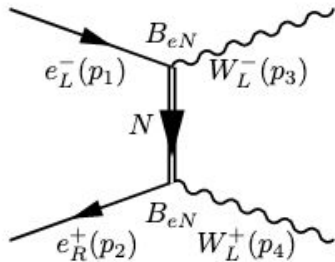
Lesson from the Past - Bound on Higgs Mass:

- Assume large Higgs mass
- Main contribution from t, u Higgs mediator channels
 - Longitudinal vector bosons have bad high energy behavior
- Take the high-energy limit ($\sqrt{s} \rightarrow \infty$)
- Upper bound on Higgs mass is $O(1 \text{ TeV})$

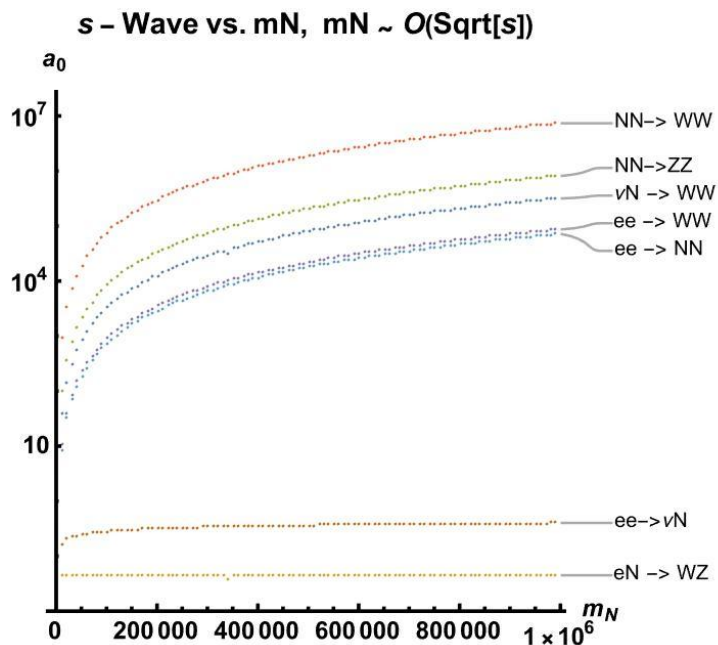


Bound on Neutrino Mixing:

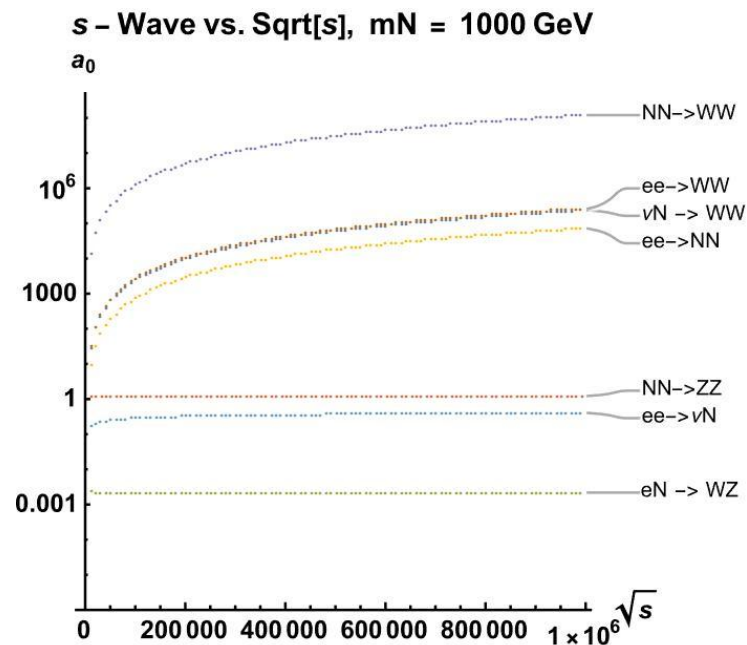
- Assume general Type 1-like model
- Assume general neutrino mixing
- So far, considered processes such as:
 - (among others to be shown)
- Take Longitudinal Vector Bosons



Preliminary Results:



NN \rightarrow ZZ (in red) is interesting!





Next Steps:

- Analyze the matrix elements to determine a competitive unitarity bound
- Use the new unitarity bound for neutrino masses to motivate a collider search.
- Use the new bound as input parameters for a madgraph simulation
- Compare this new data against current constrained phase space
- Conduct a detailed analysis looking at various different upcoming collider and neutrino experiments



References:

- [1] Baglio, J., Weiland, C., “The Triple Higgs Coupling: A New Probe of Low-Scale Seesaw Models”. [arXiv:1612.06403](https://arxiv.org/abs/1612.06403)
- [2] Cai, Y., Han, T., Ruiz, R., “Lepton Number Violation: Seesaw Models and Their Collider Tests”. [arXiv:1711.02180](https://arxiv.org/abs/1711.02180)
- [3] Cárcamo Hernández A.E., King, S. F. “Littles Inverse Seesaw Model”. [arXiv:1903.02565](https://arxiv.org/abs/1903.02565)



Backup 1: Type 2 Seesaw

- Lagrangian: $\Delta\mathcal{L}_{II}^m = -\bar{L}^c Y_\nu i\sigma_2 \Delta_L L + \text{H.c.}, \quad \Delta\mathcal{L}_{H\Delta_L} \ni \mu H^T i\sigma_2 \Delta_L^\dagger H + \text{H.c.}$
- Scalar Triplet Higgs acquires vev through the coupling to the Higgs mechanism
- Consequently generates left-handed Majorana mass
- Model results in seven total Higgses (singly and doubly charged)
- M_u has direct connections with collider physics, neutrino mass measurements, and neutrino oscillation experiments since Higgses couple to gauge bosons directly



Backup 2: Type 3 Seesaw

- Add SU(2)_L Triplet Leptons transforming as (1,3,0)
- Coupling to the SM Leptons and Higgs Doublet takes form:
- Lagrangian implies mixing between SM and triplet leptons
- Diagonalizing the mass eigenstates yields masses for heavy and light neutrinos and leptons:
- Heavy charged leptons can be searched for in colliders through charged and neutral Drell-Yan scattering

$$\mathcal{L}_Y = -Y_\Sigma \bar{L} \Sigma_R^c i\sigma^2 H^* + \text{H.c.}$$

$$m_\nu \approx \frac{Y_\Sigma^2 v_0^2}{2M_\Sigma}, \quad M_N \approx M_\Sigma,$$

$$m_l - m_l \frac{Y_\Sigma^2 v_0^2}{2M_\Sigma^2} \approx m_l, \quad M_E \approx M_\Sigma.$$

Backup 3: Type 1 Decay Width Calculations

$$\begin{aligned}
 \Gamma^{lW_L} &= \frac{g^2}{64\pi M_W^2} |V_{l4}|^2 m_N^3 \left(1 - \frac{m_W^2}{m_N^2}\right)^2 &= \frac{g^2}{64\pi M_W^2} m_{\nu_l} m_N^2 \left(1 - \frac{m_W^2}{m_N^2}\right)^2 \\
 \Gamma^{lZ_L} &= \frac{g^2}{64\pi^2 M_W^2} |V_{l4}|^2 m_N^3 \left(1 - \frac{m_Z^2}{m_N^2}\right)^2 &= \frac{g^2}{64\pi^2 M_W^2} m_{\nu_l} m_N^2 \left(1 - \frac{m_Z^2}{m_N^2}\right)^2 \\
 \Gamma^{lW_T} &= \frac{g^2}{32\pi} |V_{l4}|^2 m_N \left(1 - \frac{m_W^2}{m_N^2}\right)^2 &= \frac{g^2}{32\pi} m_{\nu_l} \left(1 - \frac{m_W^2}{m_N^2}\right)^2 \\
 \Gamma^{lZ_T} &= \frac{g^2}{32\pi \cos^2(\theta_W)} |V_{l4}|^2 m_N \left(1 - \frac{m_Z^2}{m_N^2}\right)^2 &= \frac{g^2}{32\pi \cos^2(\theta_W)} m_{\nu_l} \left(1 - \frac{m_Z^2}{m_N^2}\right)^2 \\
 \Gamma^{\nu_l H} &= \frac{g^2}{64\pi M_W^2} |V_{l4}|^2 m_N^3 \left(1 - \frac{m_H^2}{m_N^2}\right)^2 &= \frac{g^2}{64\pi M_W^2} m_{\nu_l} m_N^2 \left(1 - \frac{m_H^2}{m_N^2}\right)^2
 \end{aligned}$$

Backup 4: Neutrino Unitarity Bound Calculation

$e^+e^- \rightarrow W^+W^-$

- Matrix element:
- Kinematics:

$$i\mathcal{M} = \frac{g^2 |B_N|^2}{2((l_1 - k_-)^2 - m_{N4}^2)} \bar{v}_L(l_2) \not{\epsilon}_L^*(k_+) (l_1 - \not{k}_- + m_{N4}) \not{\epsilon}_L^*(k_-) u_L(l_1)$$

$$l_1^\mu = (E, 0, 0, E)^T$$

$$l_2^\mu = (E, 0, 0, -E)^T$$

$$k_-^\mu = (E, -k \sin \theta, 0, -k \cos \theta)^T$$

$$k_+^\mu = (E, k \sin \theta, 0, k \cos \theta)^T$$

$$\epsilon_{L\mu}^*(k_-) = \frac{1}{m_W} (k, E \sin \theta, 0, E \cos \theta)^T$$

$$\epsilon_{L\mu}^*(k_+) = \frac{1}{m_W} (k, -E \sin \theta, 0, -E \cos \theta)^T.$$

- Evaluating the terms in the Chiral basis yields:

$$i\mathcal{M} = -\frac{|B_N|^2 s^2}{4v^2 \left(\left(\frac{gv}{2} \right)^2 - \frac{s}{2} (1 + \beta \cos \theta) - m_N^2 \right)} (\beta^3 + \beta - 2 \cos \theta) \sin \theta$$



Backup 4: Neutrino Unitarity Bound Continued

- High-energy limit of amplitude:
$$i\mathcal{M} = \frac{|B_N|^2 s}{v^2} \frac{1 - \cos \theta}{1 + \cos \theta} \sin \theta$$
- Using unitarity constraint:
$$|B_N| \leq \frac{8v}{3\sqrt{s}}$$



Backup 5: Neutrino Oscillation Calculation

- To see why neutrino oscillations must result from a mass, consider high energy limit

- $E \sim p + m^2/(2p) \sim E + m^2/(2E)$
- Take $t \sim L$ (length)
- Obtain:

$$|\nu_j(L)\rangle = e^{-i\left(\frac{m_j^2 L}{2E}\right)} |\nu_j(0)\rangle$$

- Moreover,
 - Probability is proportional to squared mass difference
- Neutrino lepton flavor oscillation probabilities are a direct result from the small mass difference

Backup 6: Inverse Seesaw Coupling to Standard Model

- In Feynman-t'Hooft Gauge and mass basis, the SM interactions with the neutrinos are:
- B is proportionate to the mixing between neutrinos and leptons, while C is proportionate to the mixing between neutrinos and neutrinos

$$\begin{aligned} \mathcal{L}_{\text{int}}^Z &= -\frac{g_2}{4 \cos \theta_W} \bar{n}_i \not{Z} [C_{ij} P_L - C_{ij}^* P_R] n_j, \\ \mathcal{L}_{\text{int}}^H &= -\frac{g_2}{4m_W} H \bar{n}_i [(C_{ij} m_{n_i} + C_{ij}^* m_{n_j}) P_L + (C_{ij} m_{n_j} + C_{ij}^* m_{n_i}) P_R] n_j, \\ \mathcal{L}_{\text{int}}^{G^0} &= \frac{ig_2}{4m_W} G^0 \bar{n}_i [-(C_{ij} m_{n_i} + C_{ij}^* m_{n_j}) P_L + (C_{ij} m_{n_j} + C_{ij}^* m_{n_i}) P_R] n_j \\ \mathcal{L}_{\text{int}}^{W^\pm} &= -\frac{g_2}{\sqrt{2}} \bar{l}_i B_{ij} W^\pm P_L n_j + h.c., \\ \mathcal{L}_{\text{int}}^{G^\pm} &= \frac{-g_2}{\sqrt{2}m_W} G^\pm [\bar{l}_i B_{ij} (m_{l_i} P_L - m_{n_j} P_R) n_j] + h.c., \end{aligned}$$

Backup 7: Inverse Seesaw Parameters

- Use these mixings to compute the massive neutrino decay widths.
- Example decay width formulae:

$$\Gamma^{eW_L} = \frac{g^2}{64\pi M_W^2} |B_{eN_4}|^2 m_{N_4}^3 \left(1 - \frac{M_W^2}{m_{N_4}^2}\right)^2$$

$$\Gamma^{\mu W_L} = \frac{g^2}{64\pi M_W^2} |B_{\mu N_4}|^2 m_{N_4}^3 \left(1 - \frac{M_W^2}{m_{N_4}^2}\right)^2$$

$$\Gamma^{\tau W_L} = \frac{g^2}{64\pi M_W^2} |B_{\tau N_4}|^2 m_{N_4}^3 \left(1 - \frac{M_W^2}{m_{N_4}^2}\right)^2$$

$$\Gamma^{eW_T} = \frac{g^2}{32\pi} |B_{eN_4}|^2 m_{N_4} \left(1 - \frac{M_W^2}{m_{N_4}^2}\right)^2$$

$$\Gamma^{\mu W_T} = \frac{g^2}{32\pi} |B_{\mu N_4}|^2 m_{N_4} \left(1 - \frac{m_W^2}{m_{N_4}^2}\right)^2$$


$$\Gamma^{\tau W_T} = \frac{g^2}{32\pi} |B_{\tau N_4}|^2 m_{N_4} \left(1 - \frac{m_W^2}{m_{N_4}^2}\right)^2$$



Backup 8: Numerical Inverse Seesaw Mass Parameters

- Mass Parameters are and mixing angles computed numerically from a single M_R input
 - Fix Yukawa coupling to Dirac mass as 1
 - To enforce unitarity conditions, set:
 - Viz. analysis by Weiland and Baglio, 2017
 - Use Casas-Ibarra Parametrization adapted to the Inverse Seesaw model.
 - Method of diagonalizing the new mass matrix by using the PMNS matrix and measured neutrino masses
 - Back-calculate μ_X with empirical constraints and M_R input
 - End result: Diagonalized masses and mixing angles

$$M_{R_1} = 1.51M_R, \quad M_{R_2} = 3.59M_R, \quad M_{R_3} = M_R;$$



Backup 9: Analytical Inverse Seesaw Mass parameters

- Diagonalizing the mass matrix to get $M_{\{Ni\}}$ and $m_{\{ui\}}$ eigenmasses yields:

$$\begin{aligned}d_1 &\approx \frac{M_{R1}m_{\nu 3}}{m_D\mu_1} & e_1 &\approx \frac{M_{R1}}{m_D} \sqrt{\frac{m_{\nu 2}}{\mu_5}} & f_1 &\approx \frac{M_{R1}}{m_D} \sqrt{\frac{m_{\nu 1}}{\mu_3}} \\d_2 &\approx \frac{M_{R2}m_{\nu 3}}{m_D\mu_2} & e_2 &\approx \frac{M_{R2}}{m_D} \sqrt{\frac{m_{\nu 2}}{\mu_4}} & f_2 &\approx \frac{M_{R2}}{m_D} \sqrt{\frac{m_{\nu 1}}{\mu_5}} \\d_3 &\approx \frac{M_3m_{\nu 3}}{m_D\mu_3} & e_3 &\approx \frac{M_{R3}}{m_D} \sqrt{\frac{m_{\nu 2}}{\mu_2}} & f_3 &\approx \frac{M_{R3}}{m_D} \sqrt{\frac{m_{\nu 1}}{\mu_6}}\end{aligned}$$

- f_i represents the mixing between the electron and the first massive pseudo-dirac Neutrino mass