

First measurement of weak mixing angle in direct detection

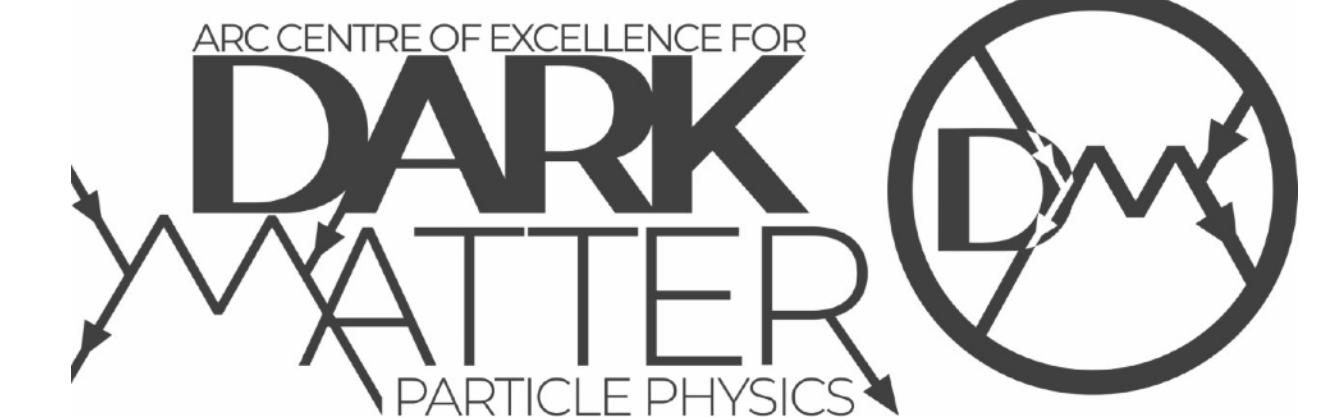
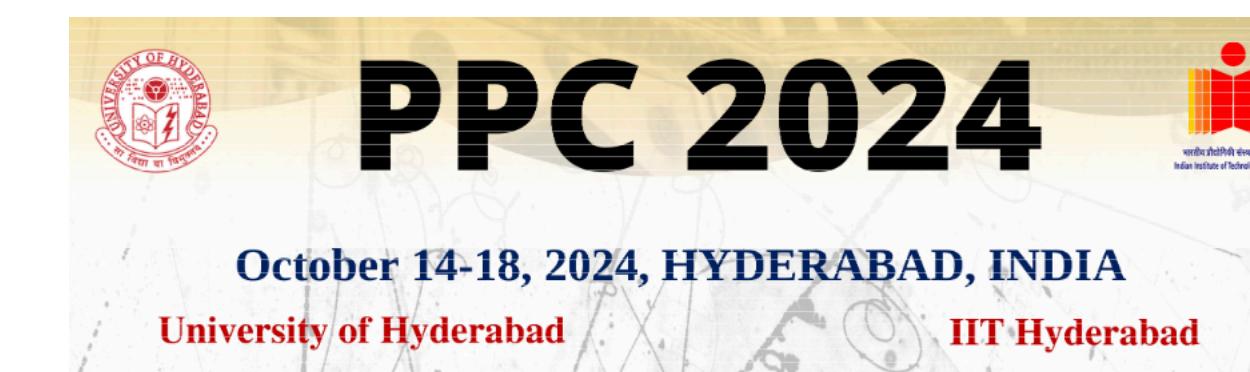
Tarak Nath Maity
University of Sydney

Based on:

[TNM](#), C Boehm; 2409.04385



THE UNIVERSITY OF
SYDNEY



What are we doing?

Have we ever **tested** the **Standard Model** of particle physics in **sub-MeV energy regime** before?

Perhaps not!

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Weak mixing angle

$$A^\mu = B_0^\mu \cos \theta_W + W_0^\mu \sin \theta_W$$

$$Z^\mu = W_0^\mu \cos \theta_W - B_0^\mu \sin \theta_W$$

$$\sin^2 \theta_W = \frac{g'^2}{g^2 + g'^2}$$

g : SU(2)_L gauge coupling

g' : U(1)_Y gauge coupling

Quantum correction



Running of weak mixing angle

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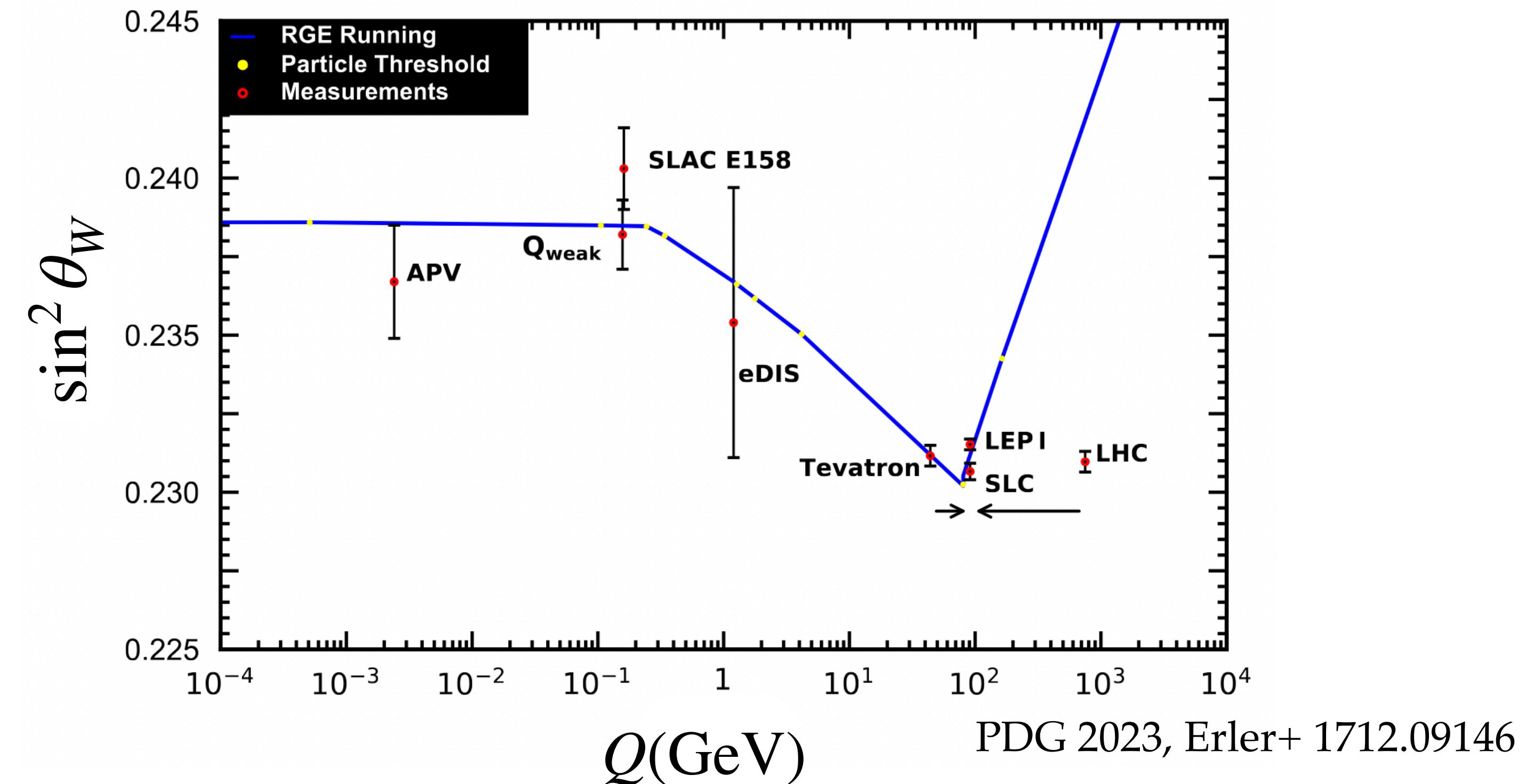
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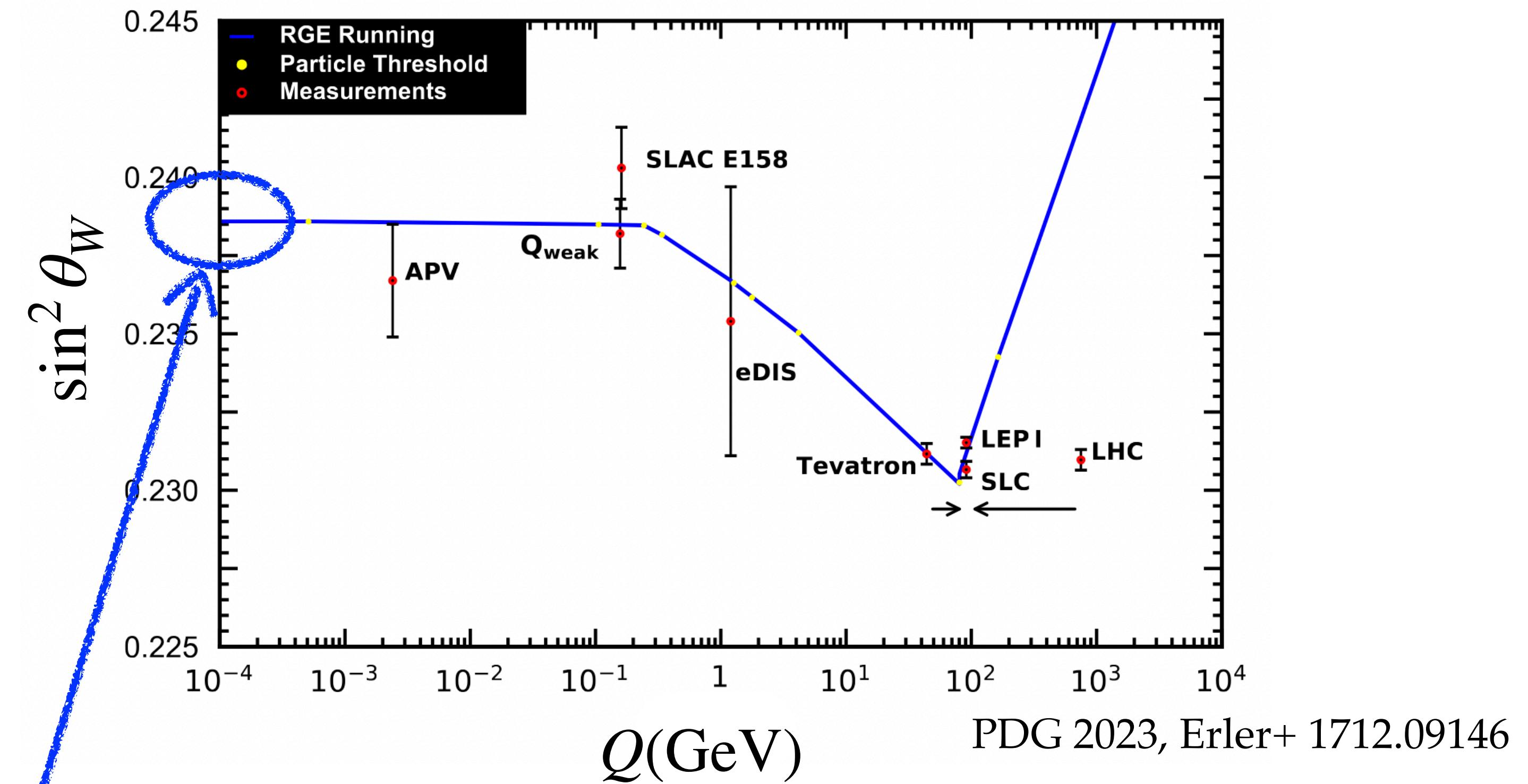
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This talk: direct detection
experiments can measure here

TNM, Boehm; 2409.04385

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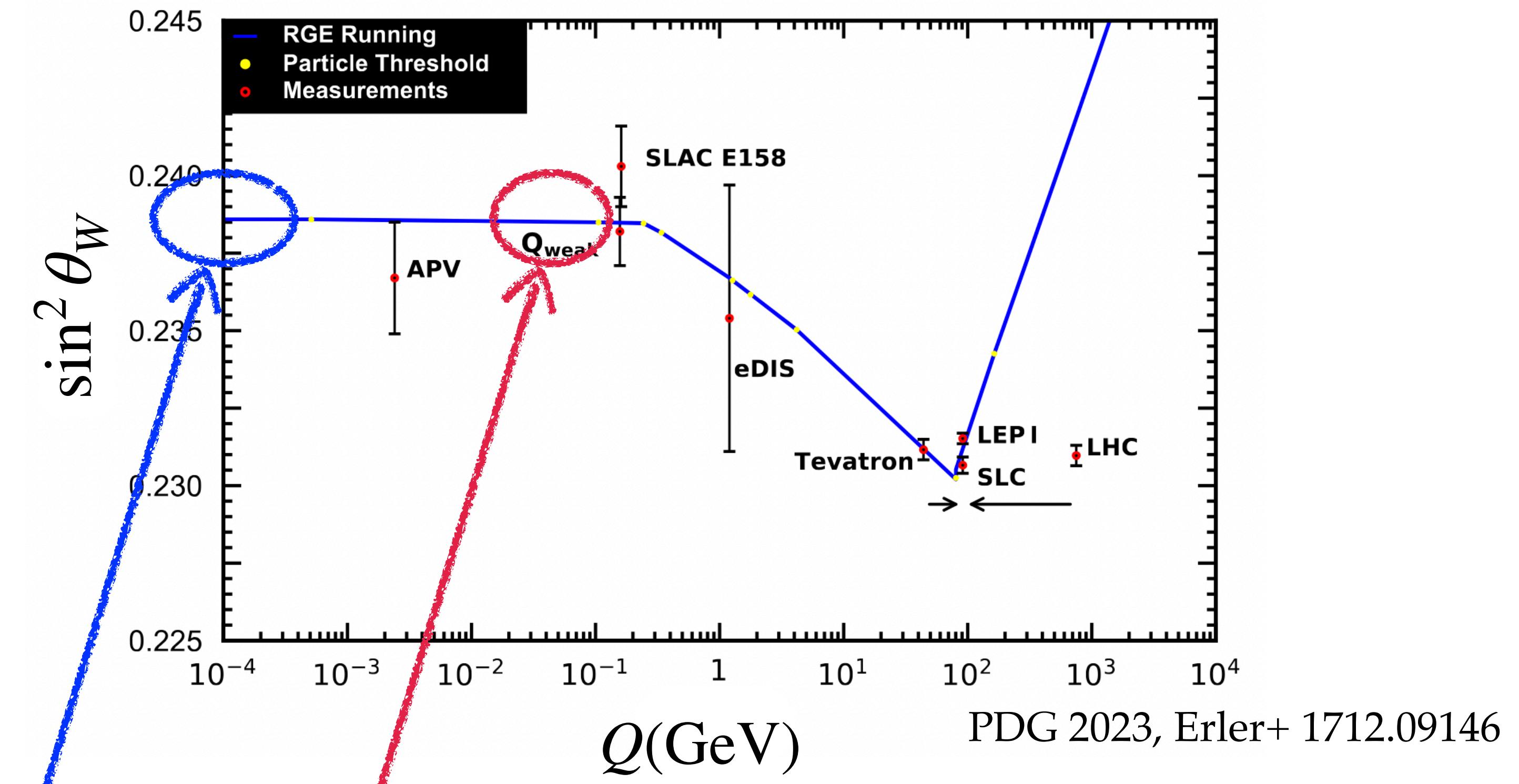
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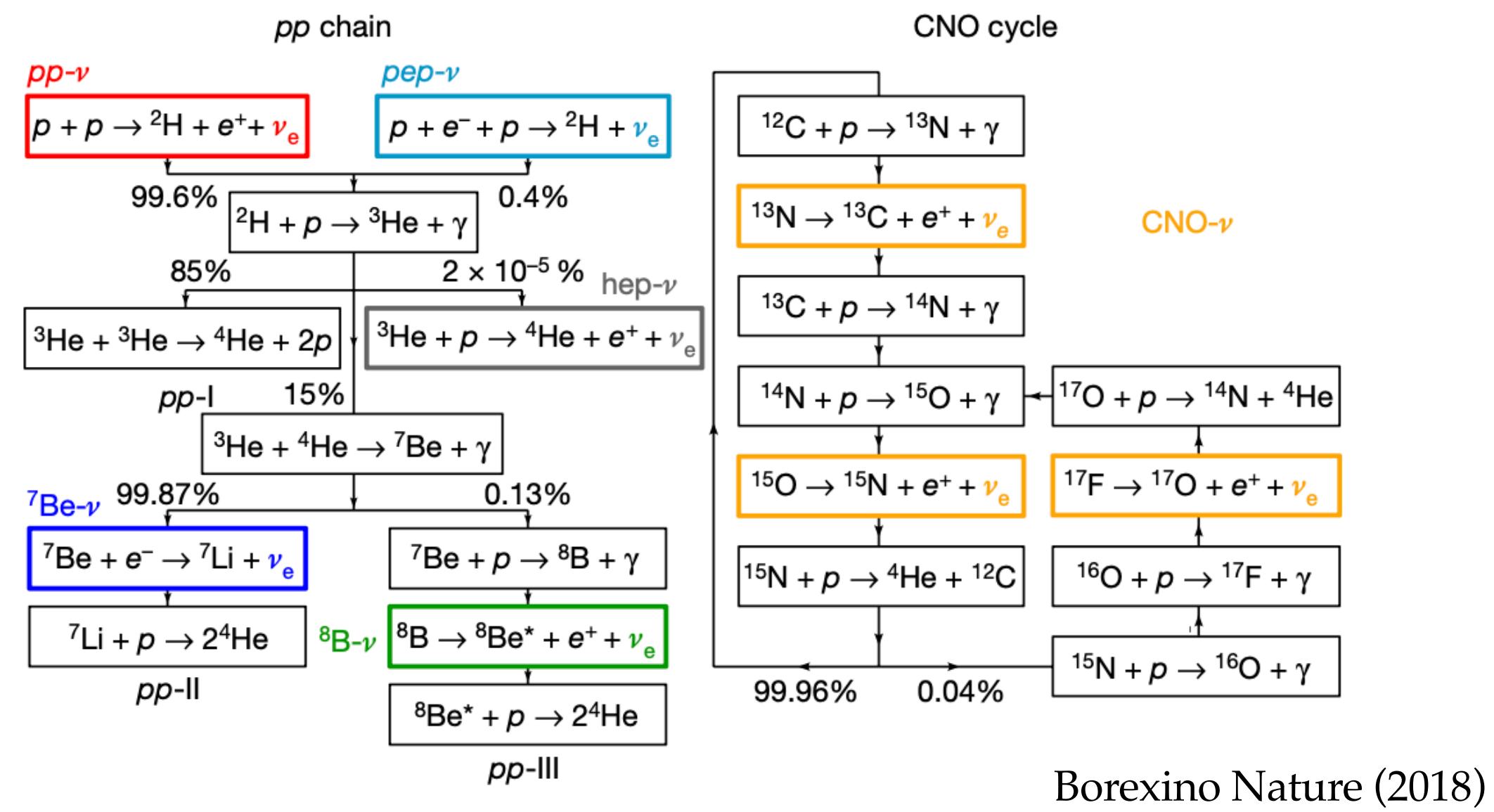
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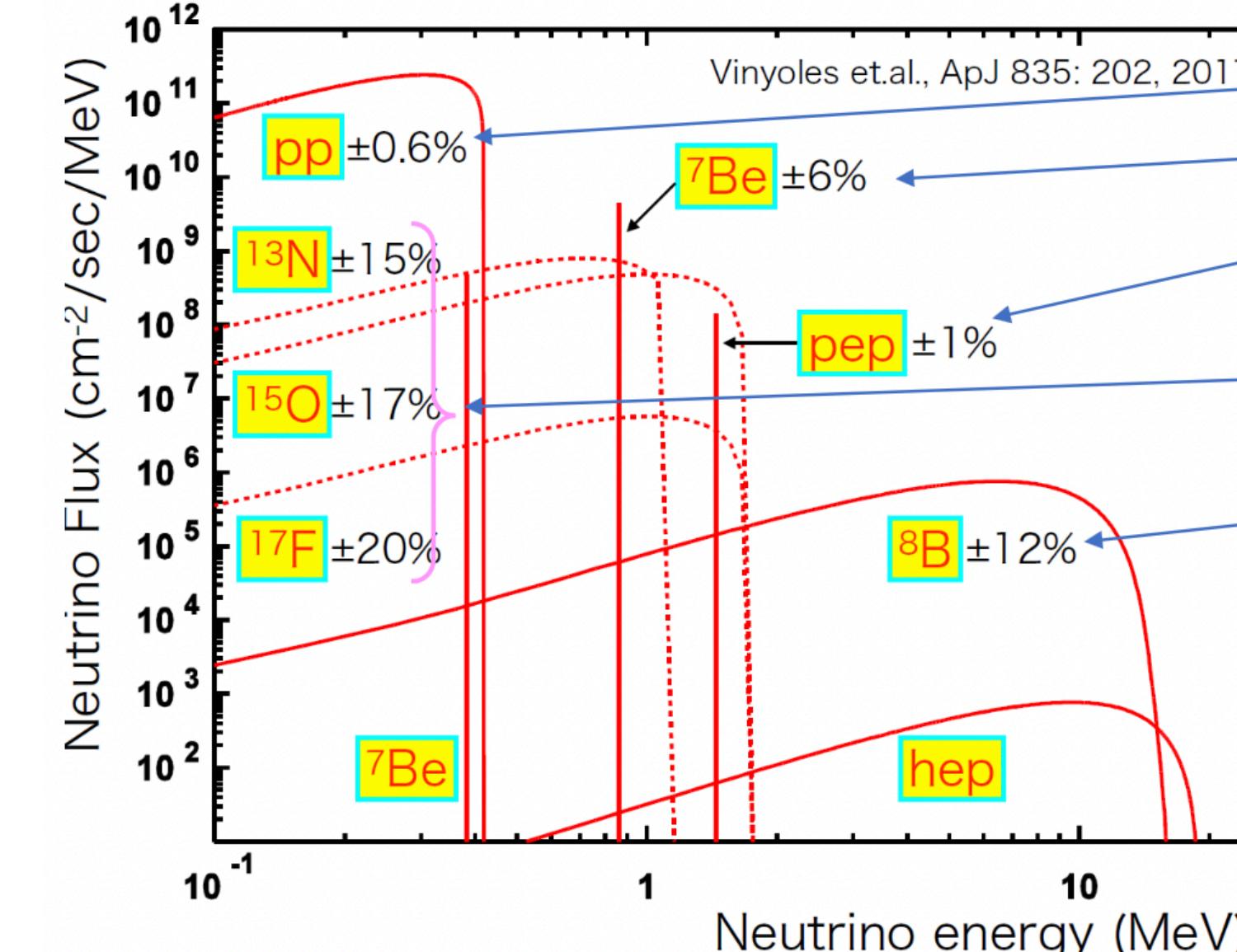
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How are we doing?

By observing Solar neutrinos



Energy spectra w/ SSM uncertainties



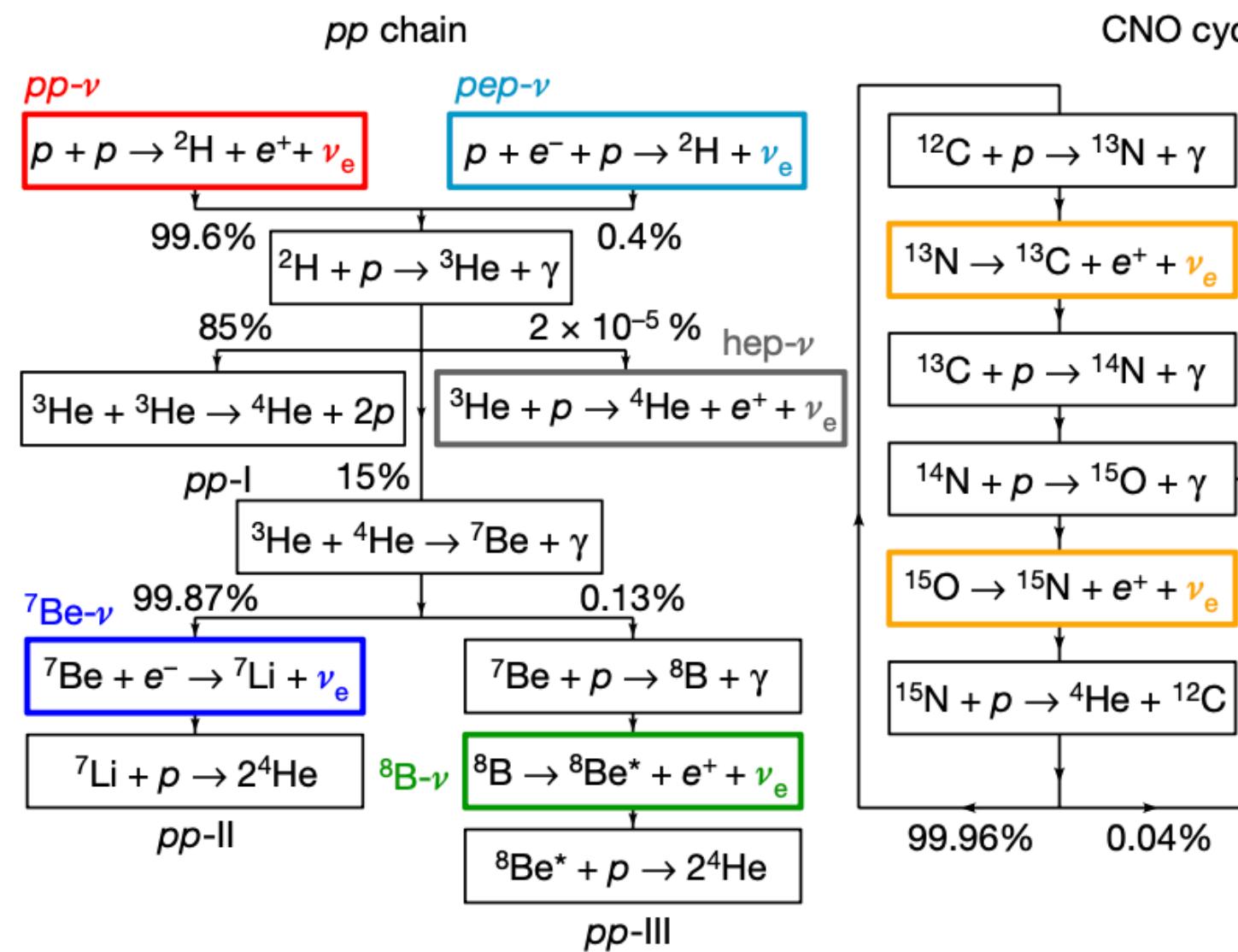
Measured flux errors



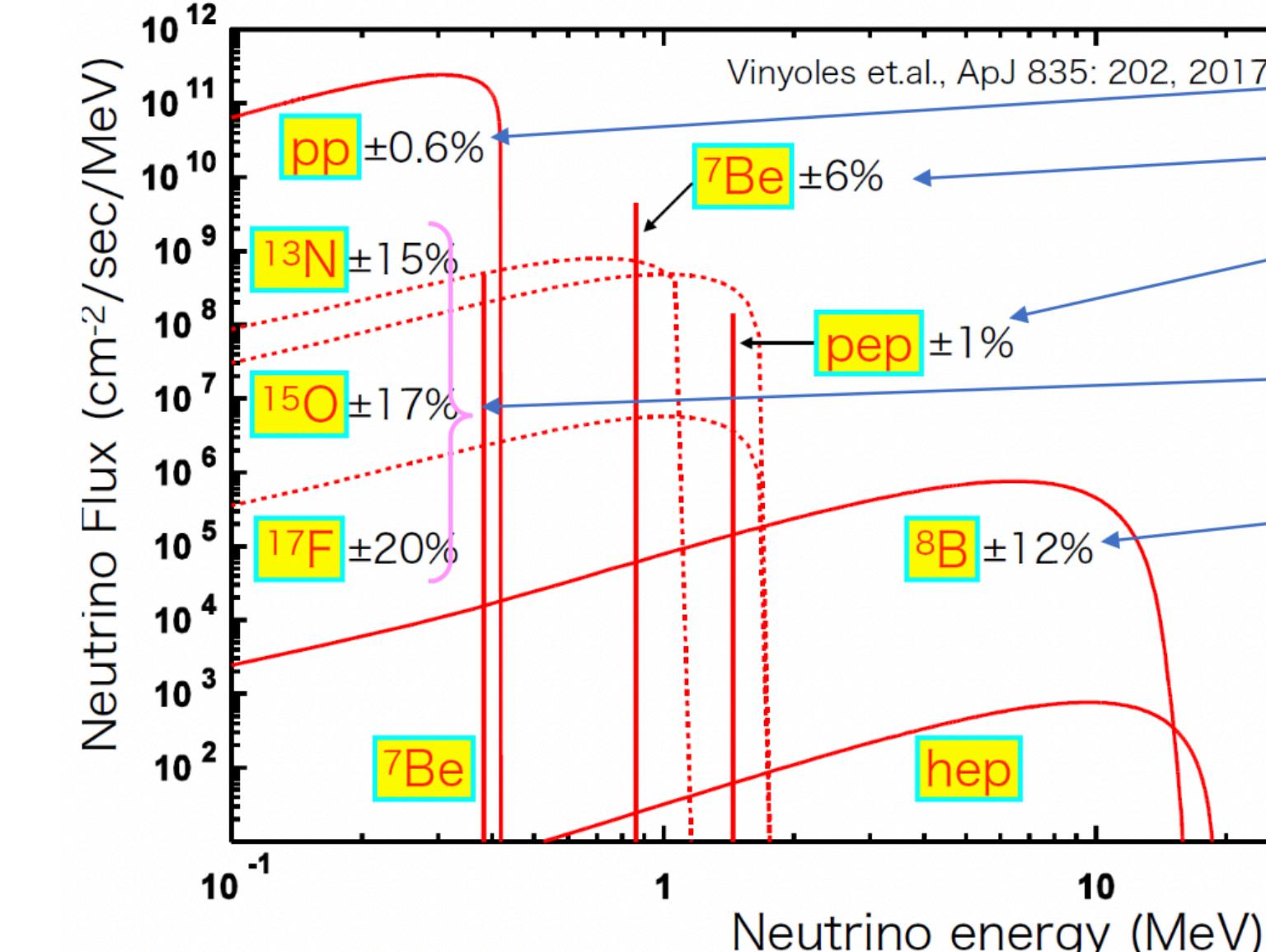
Sekiya TAUP 2023
SNO nucl-ex/0204008
Bergstrom+ 1601.00972

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Energy spectra w/ SSM uncertainties



Measured flux errors



$\pm 10\%$

$\pm 2.7\%$

$\pm 18\%$

+42%

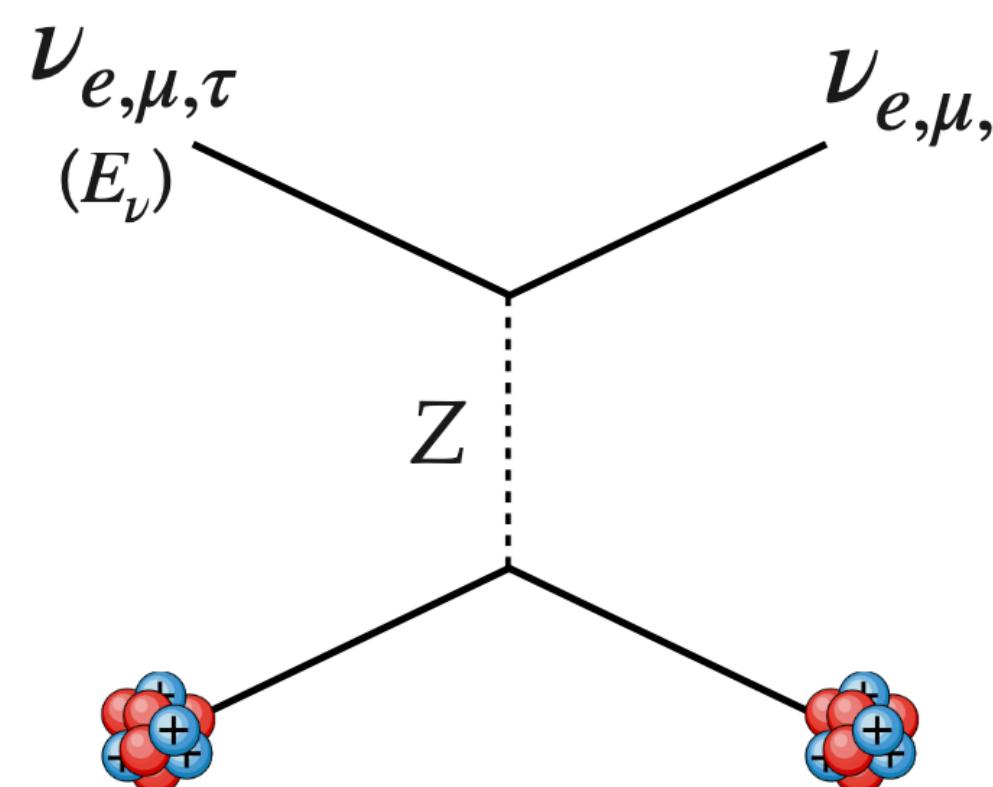
-10%

$\pm 2\%$

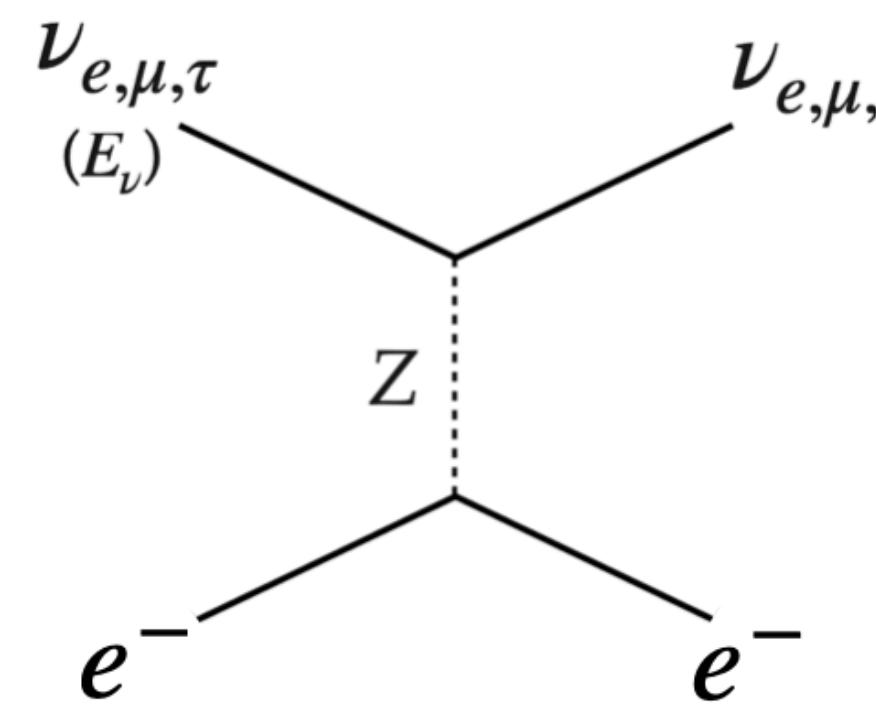
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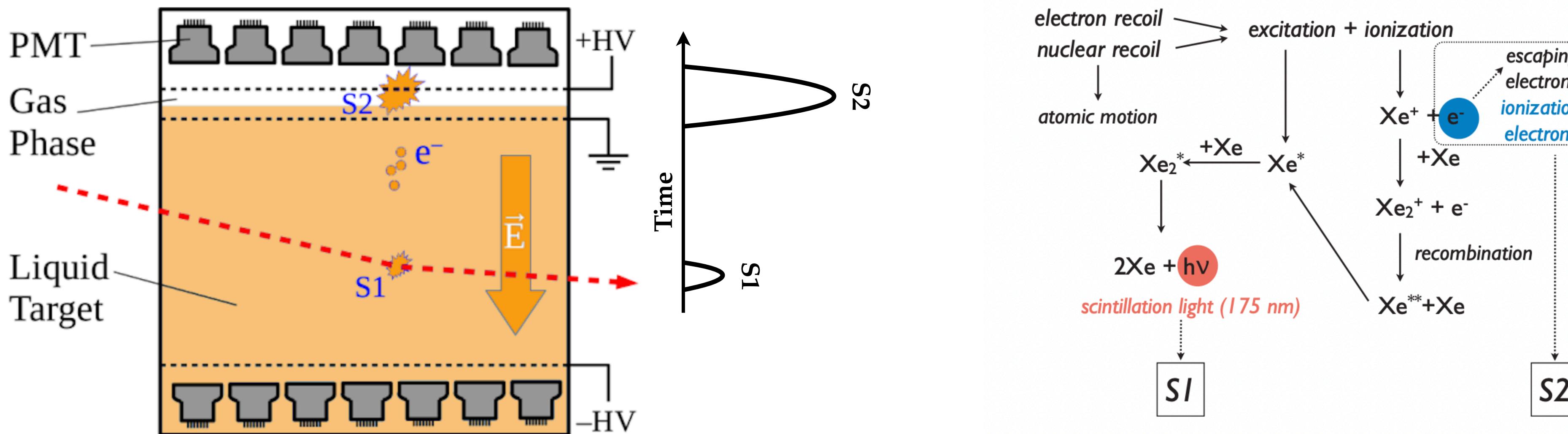
Coherent elastic neutrino-nucleus scattering (CE ν NS)



Neutrino-electron scattering

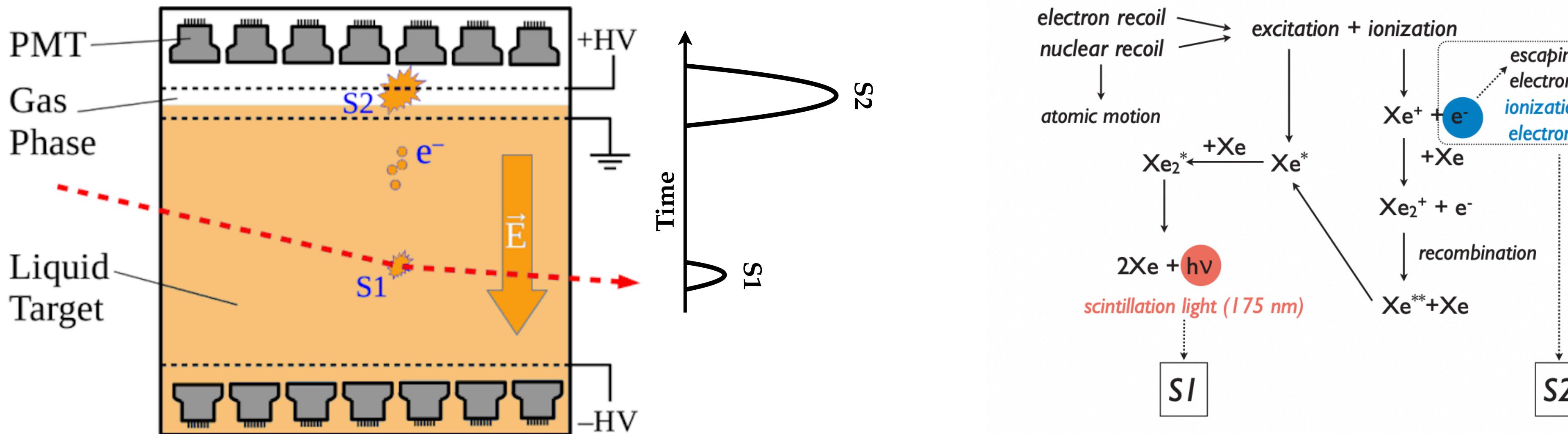
Observing neutrinos in DD(Xe)

XENONnT, PandaX-4T, LZ ...



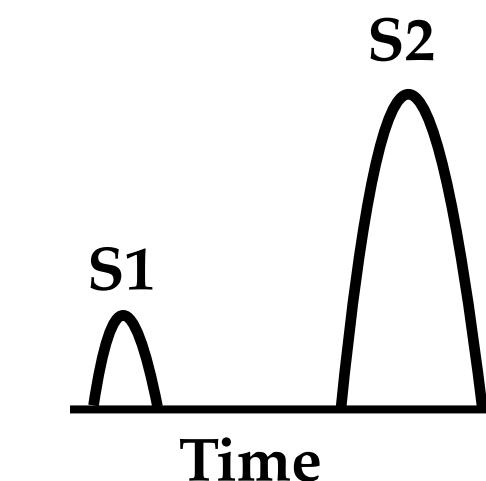
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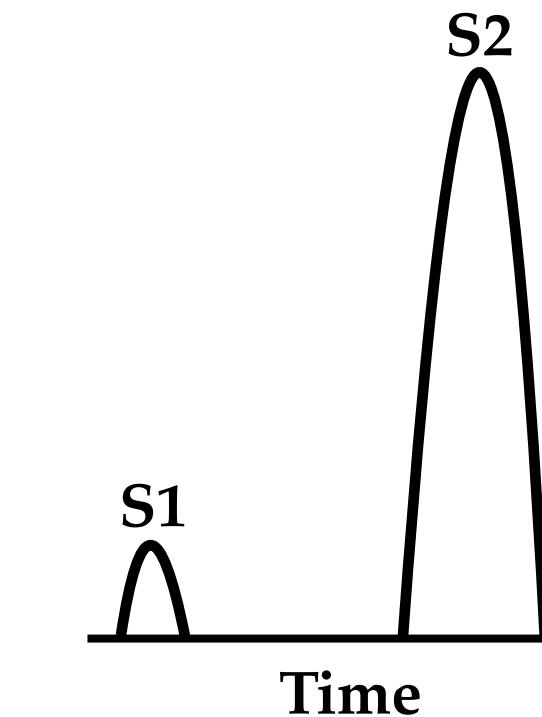


S1-S2 only analysis

Nuclear recoil



Electron recoil

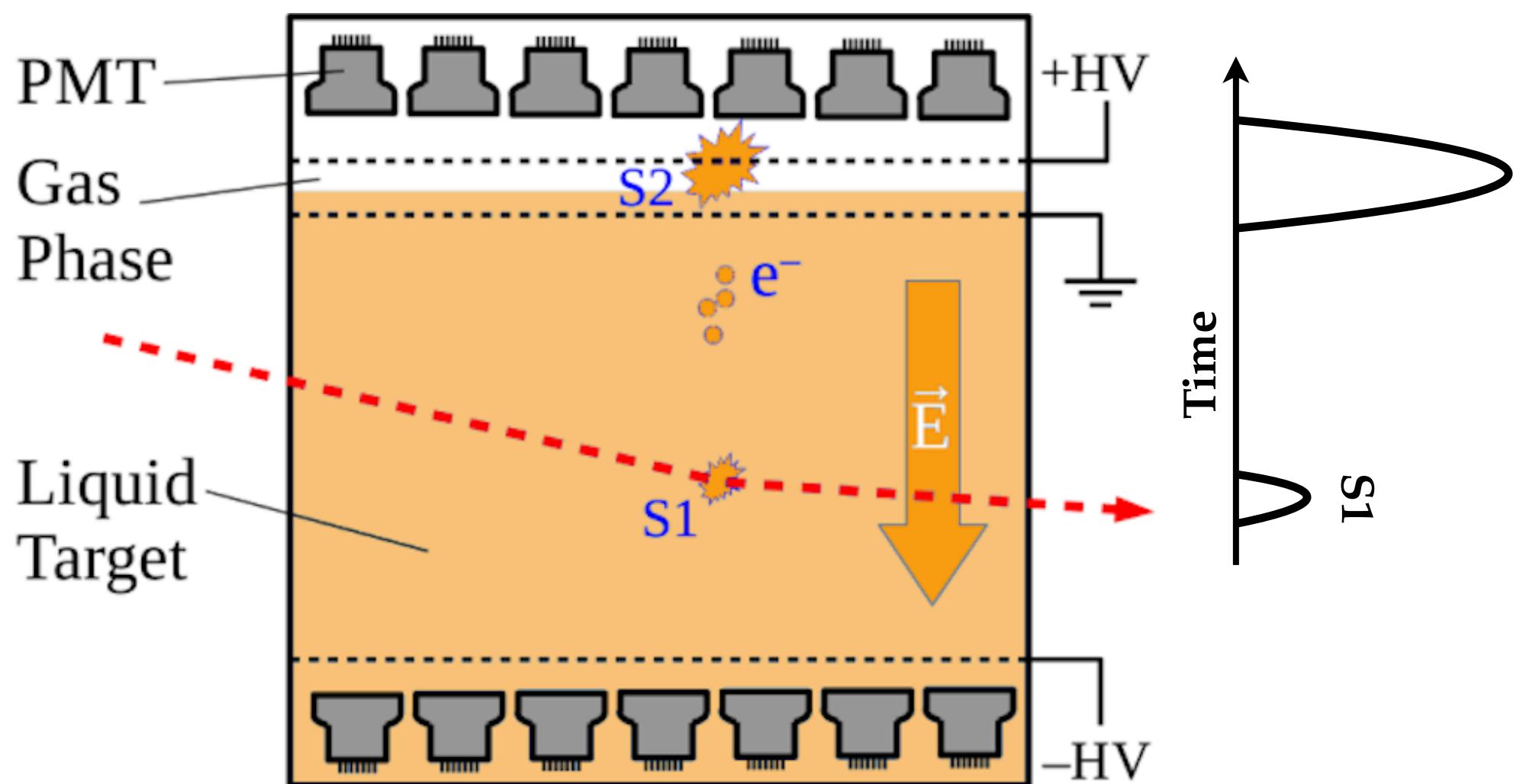


S2/S1 ratio - can distinguish - nuclear and electron recoil

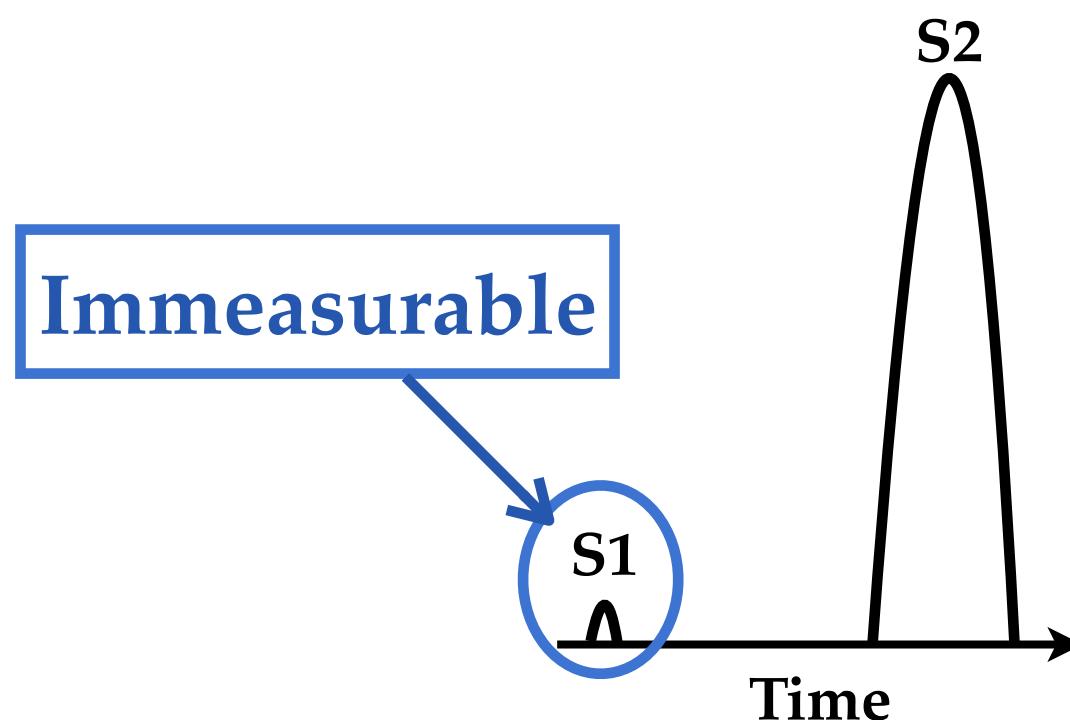
$$E_{\text{recoil}} \gtrsim 0.5 \text{ keV}$$

Observing neutrinos in DD(Xe)

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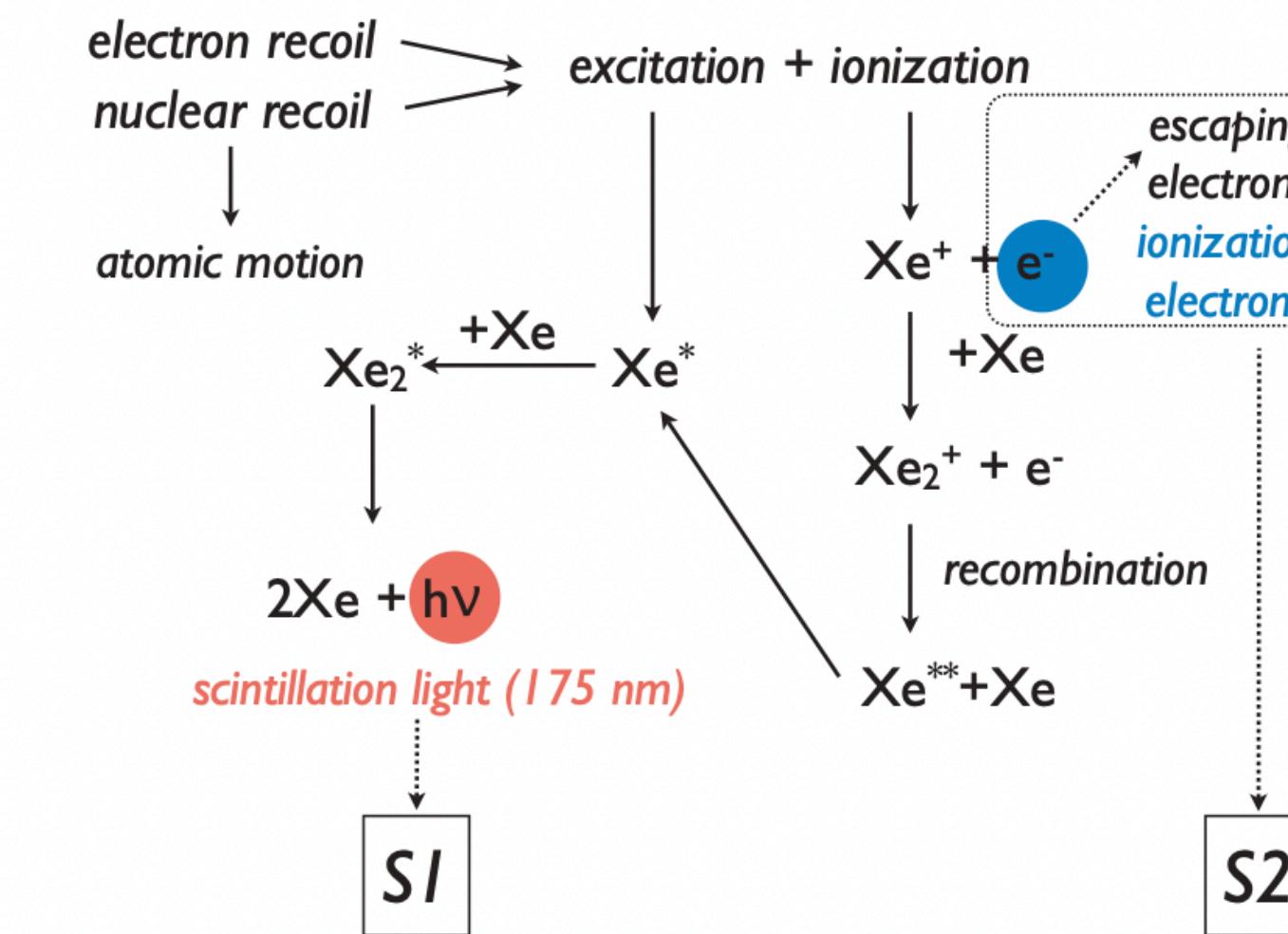


S2 only analysis



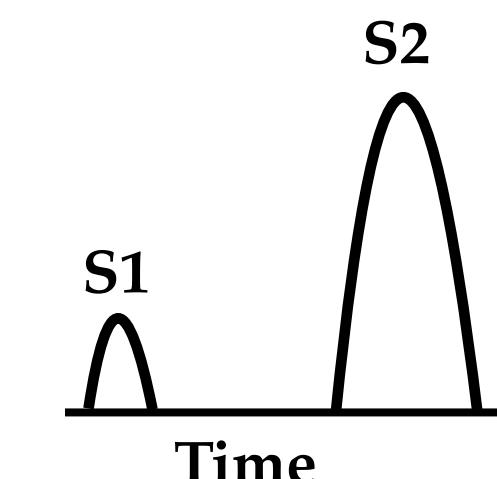
No S2/S1 ratio - can't distinguish - nuclear and electron recoil

$$E_{\text{recoil}} \lesssim 0.5 \text{ keV}$$

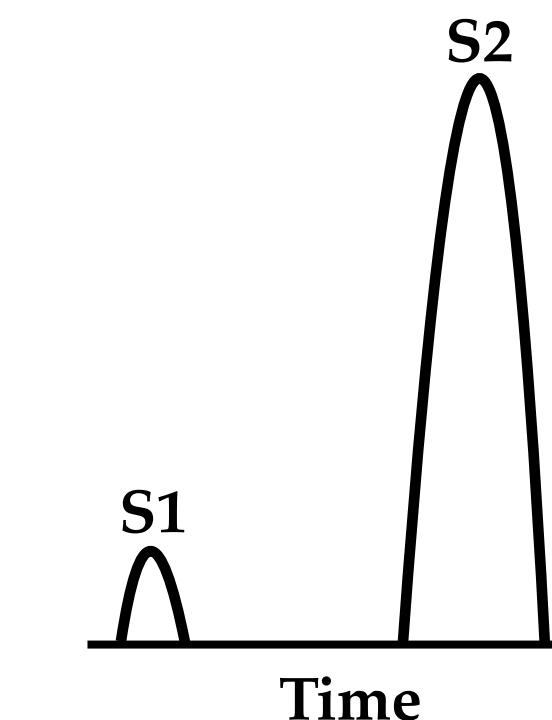


S1-S2 only analysis

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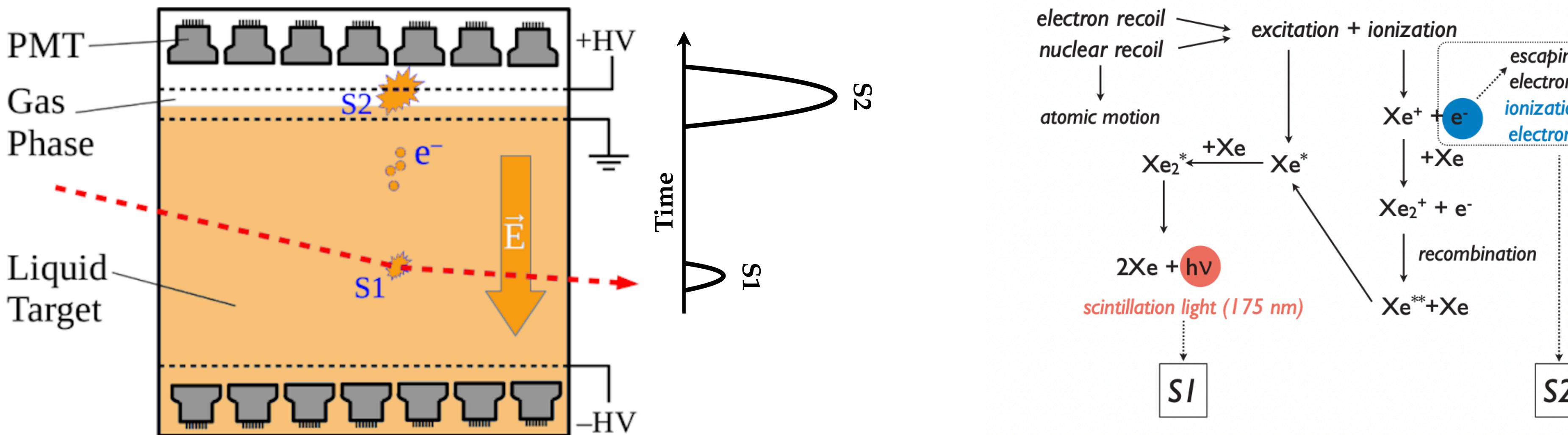


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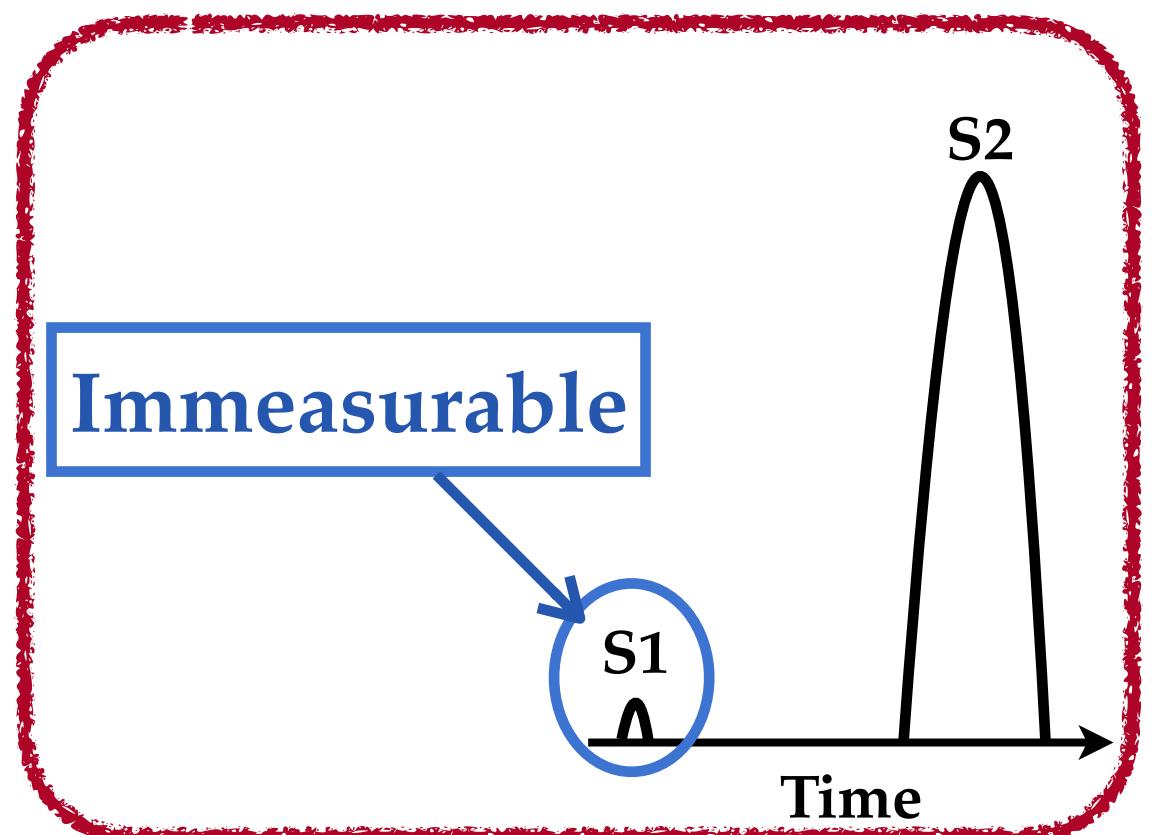
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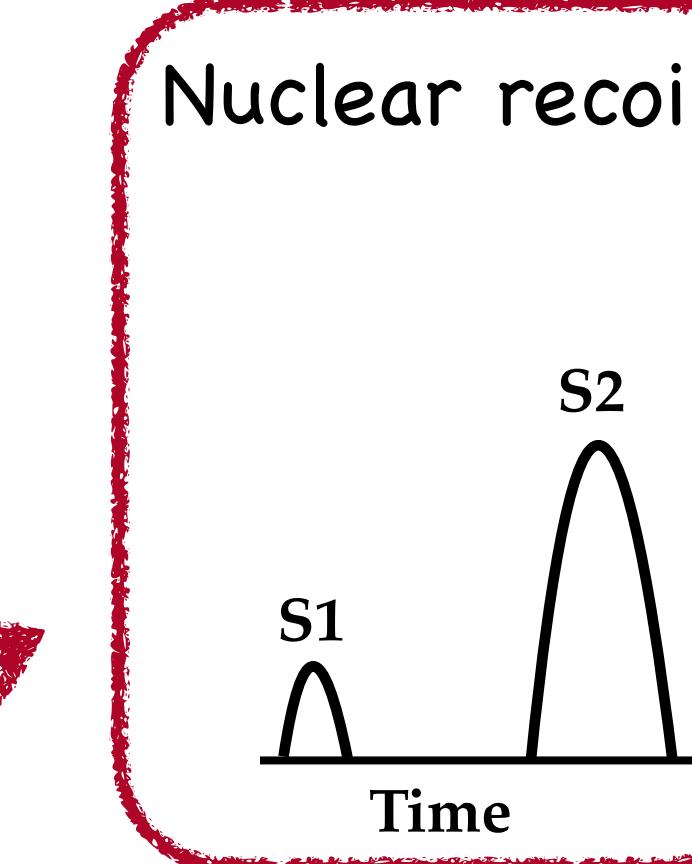
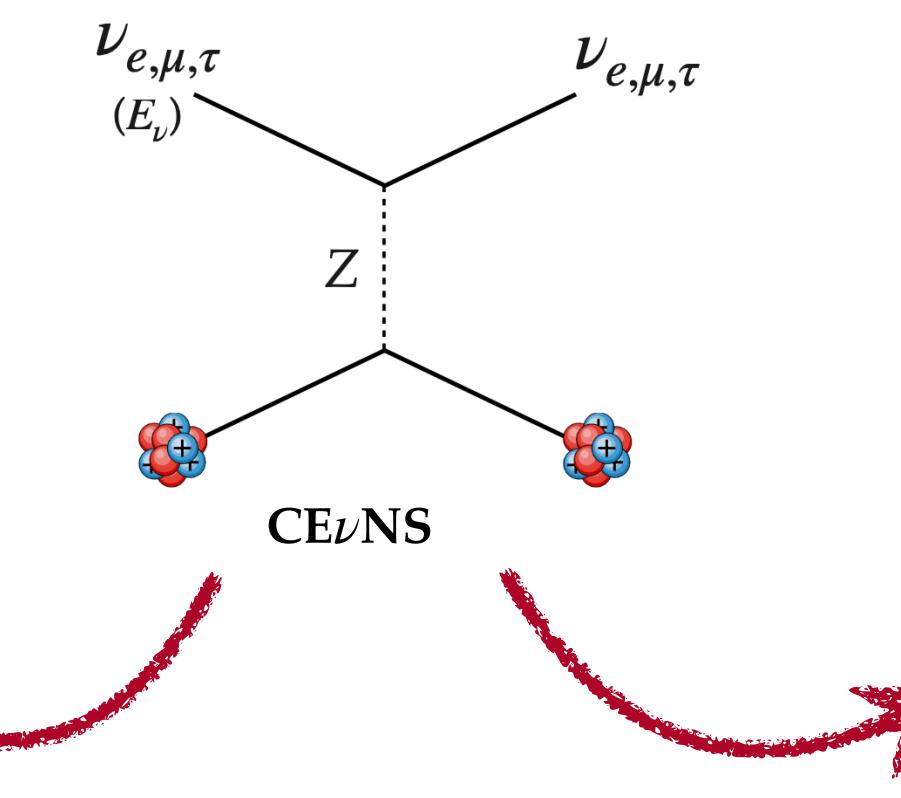


S_2 only analysis



$$E_{\text{recoil}} \lesssim 0.5 \text{ keV}$$

S_1 - S_2 only analysis

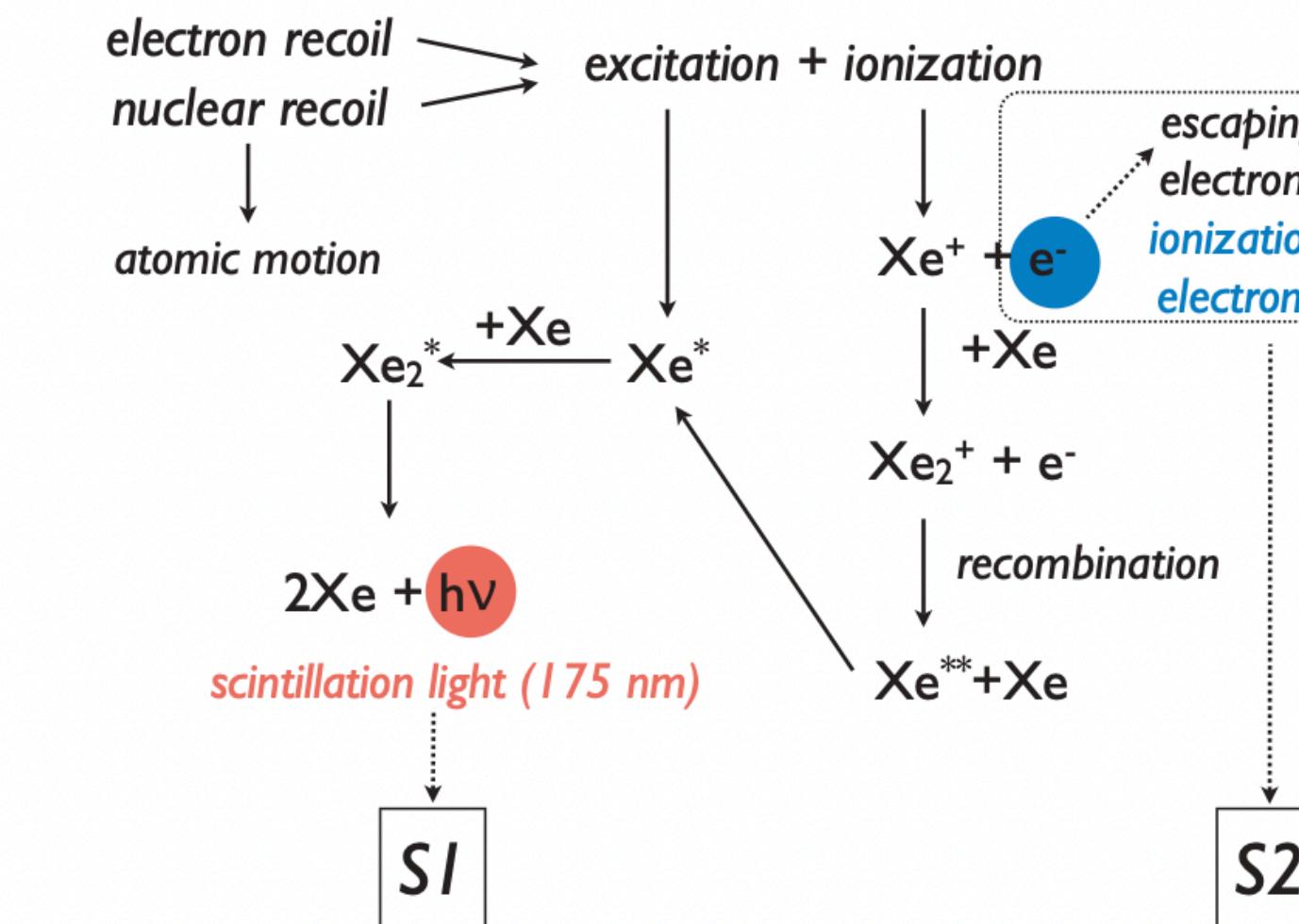
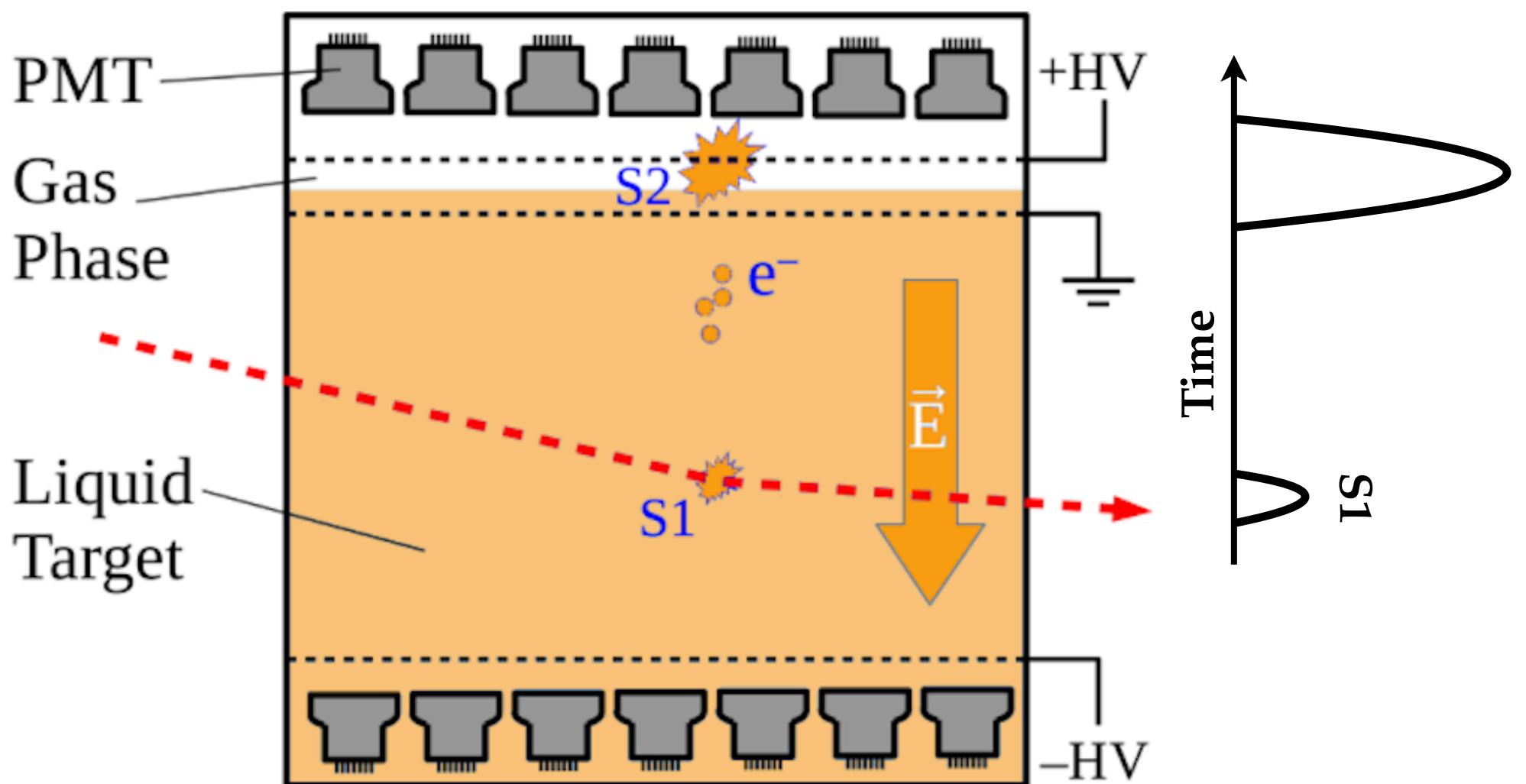


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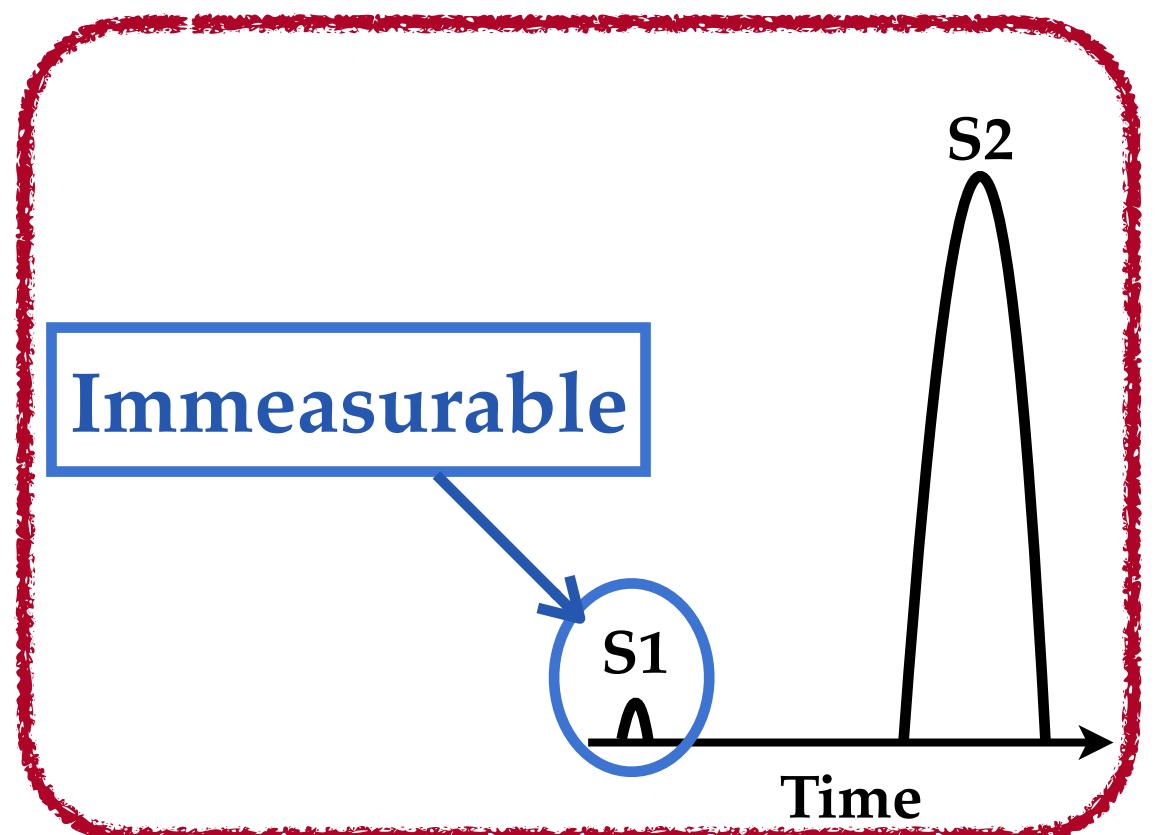
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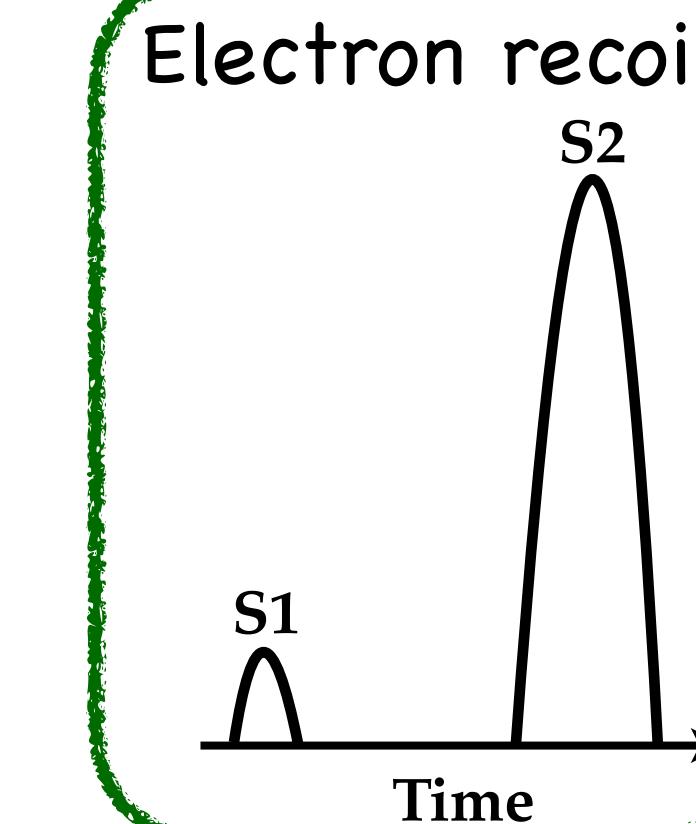
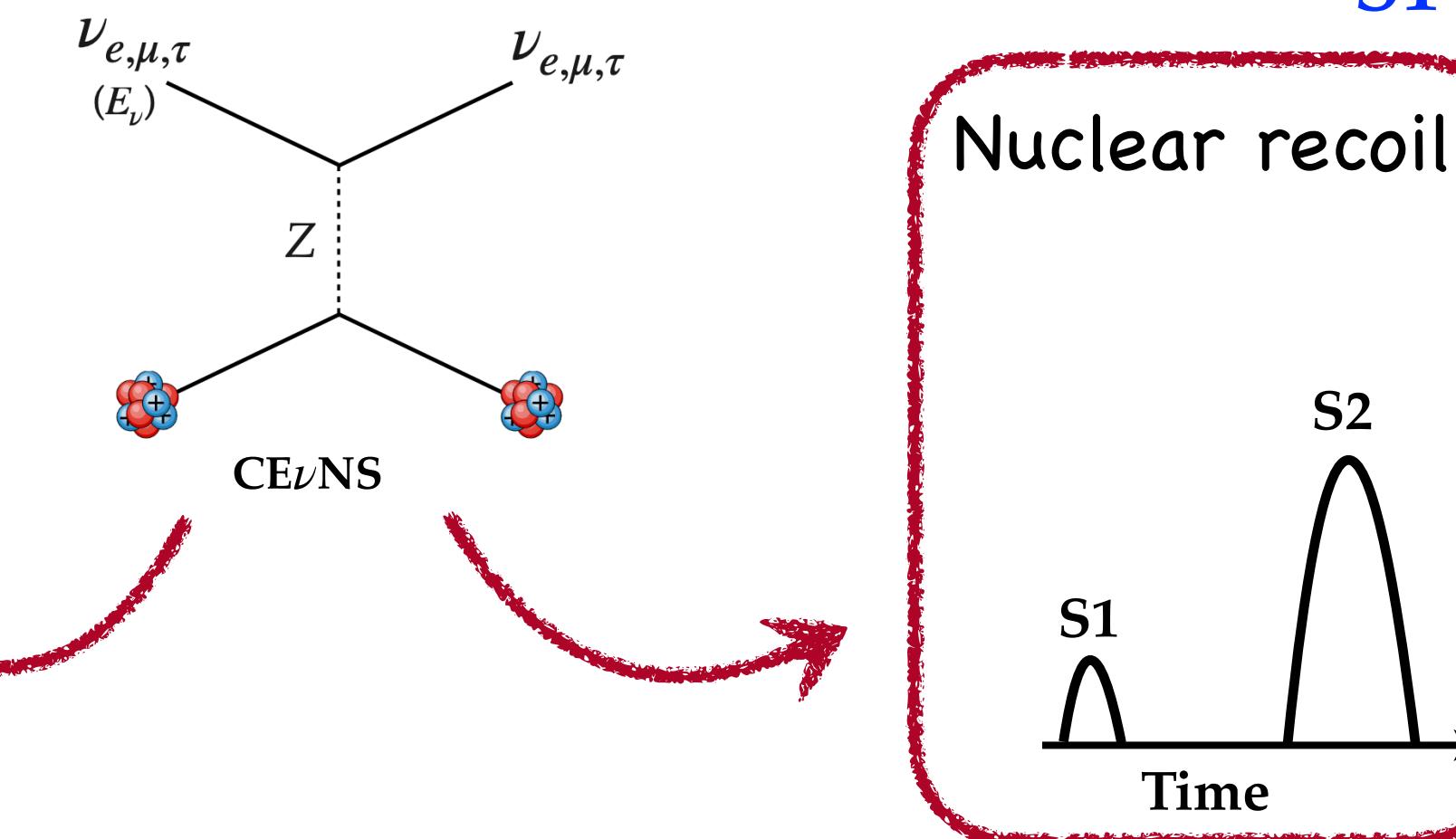
XENONnT, PandaX-4T, LZ ...



S2 only analysis



S1-S2 only analysis



No S2/S1 ratio - can't distinguish - nuclear and electron recoil

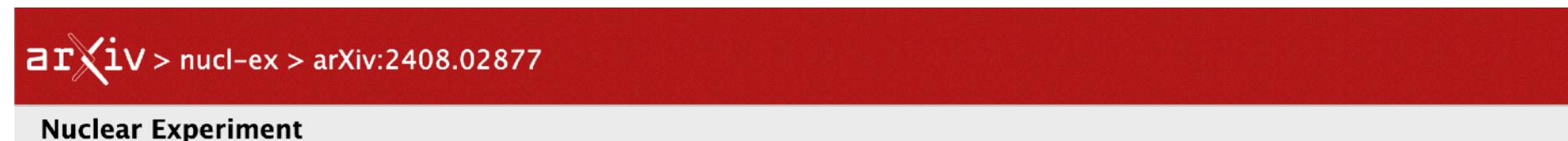
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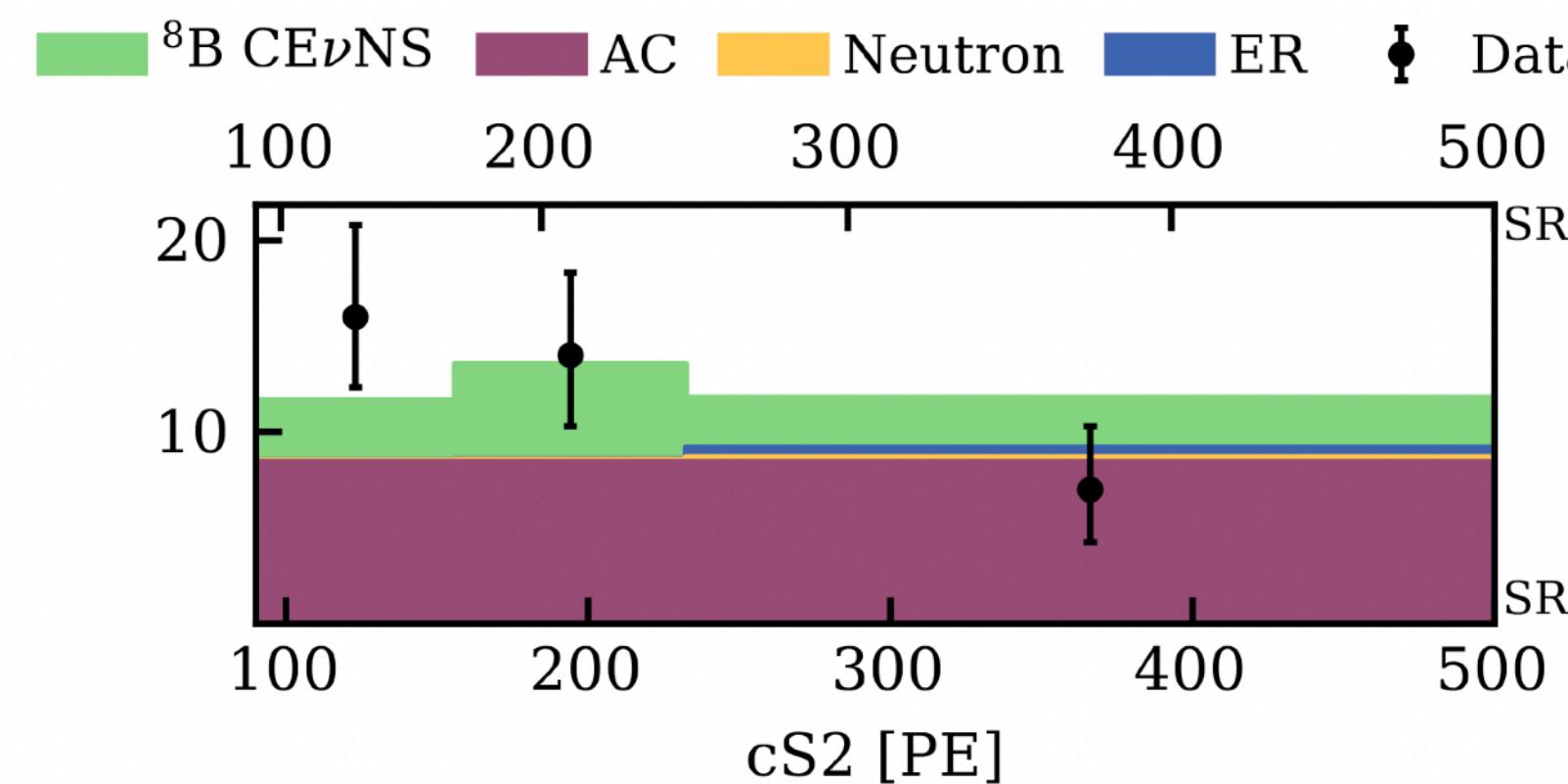
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Neutrino events at DD? nuclear recoil

XENONnT

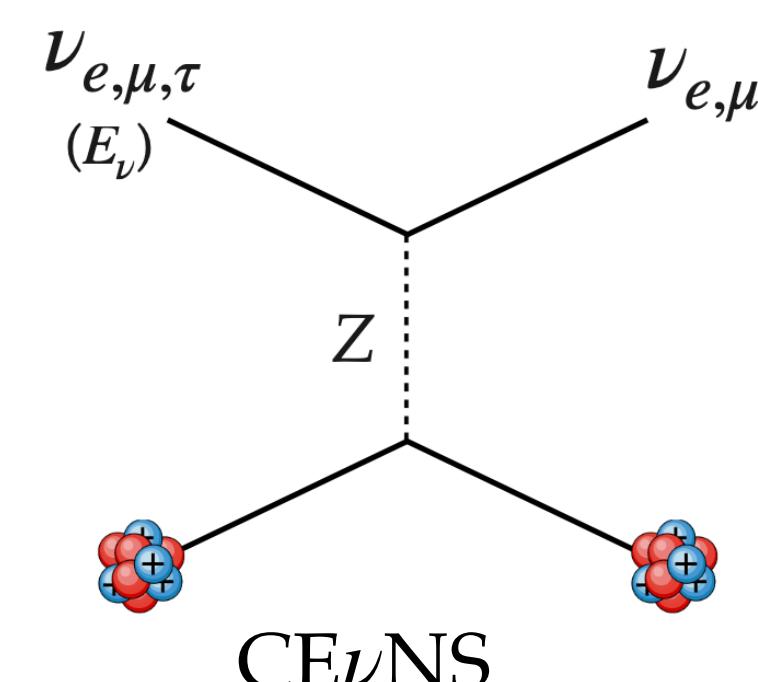


First Measurement of Solar ${}^8\text{B}$ Neutrinos via Coherent Elastic Neutrino-Nucleus Scattering with XENONnT

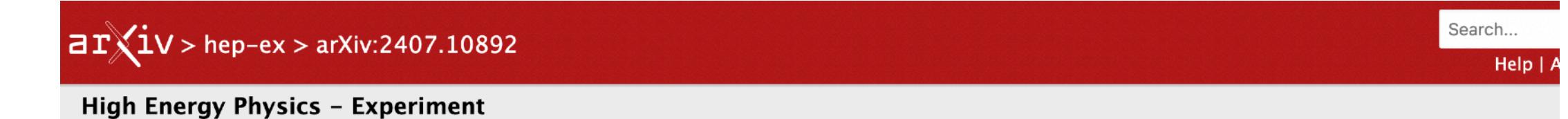


Observed events: $10.7^{+3.7}_{-4.2}$ (S1-S2 analysis)

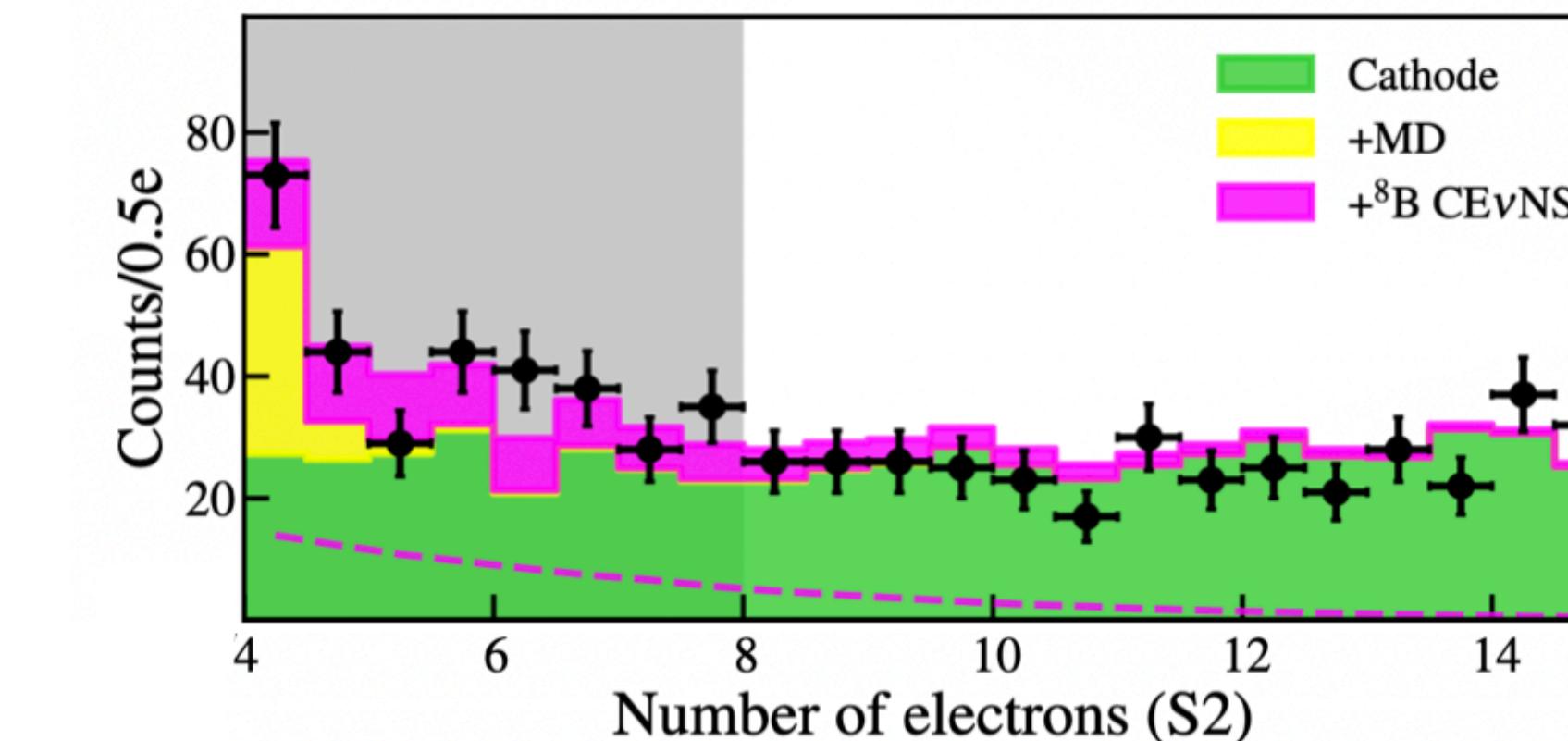
Statistical significance: 2.73σ



PandaX-4T



First Indication of Solar ${}^8\text{B}$ Neutrino Flux through Coherent Elastic Neutrino-Nucleus Scattering in PandaX-4T



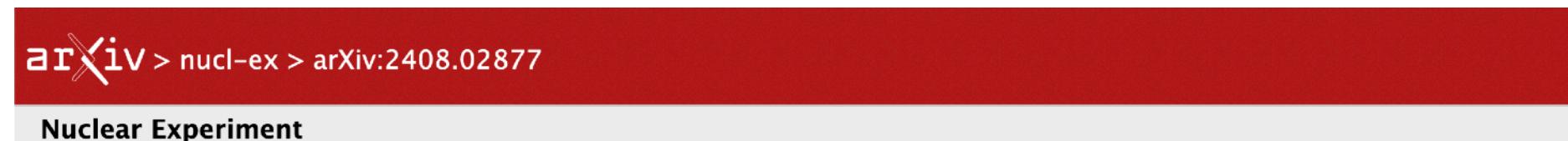
Observed events: 3.5 ± 1.3 (S1-S2 analysis)

Observed events: 78 ± 28 (S2-only analysis)

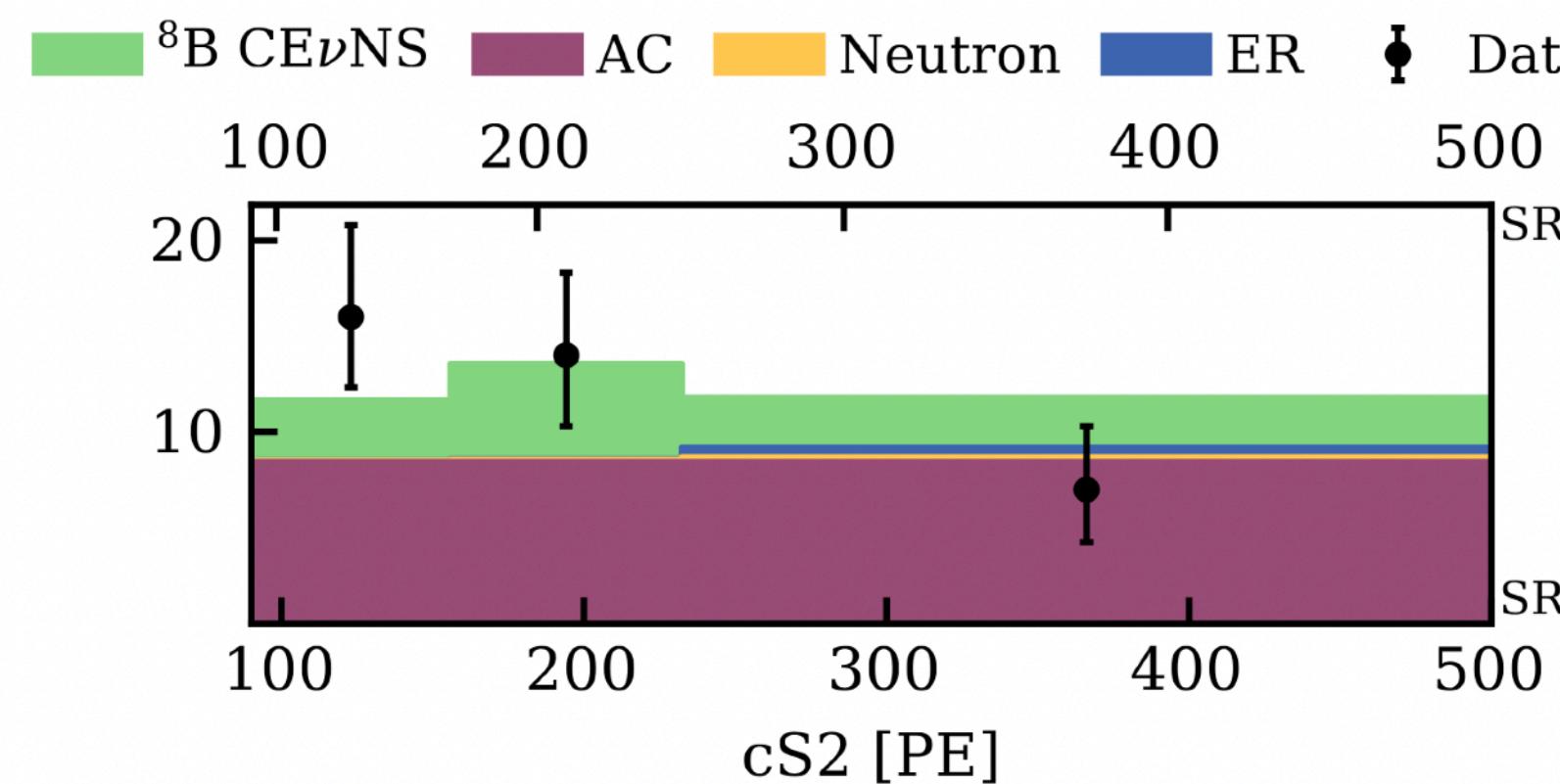
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Neutrino events at DD? nuclear recoil

XENONnT

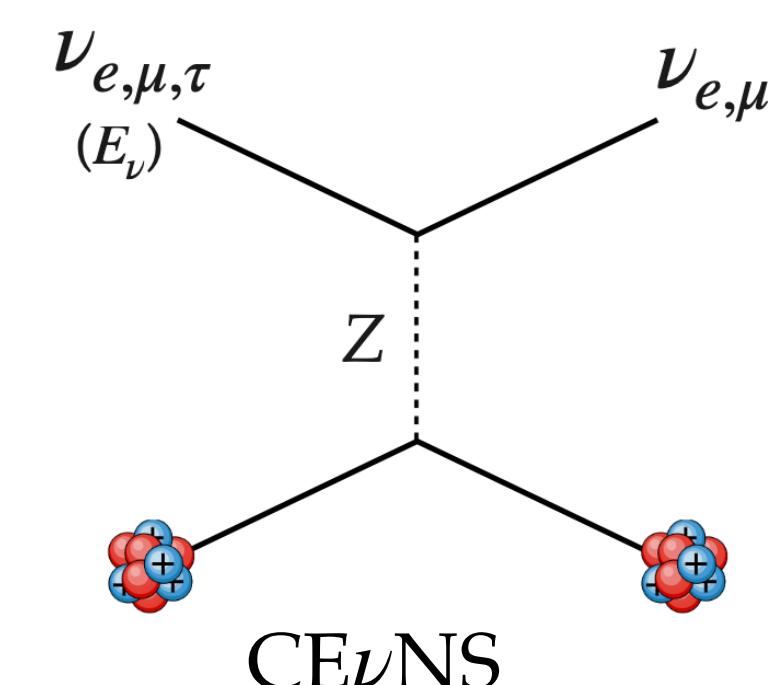


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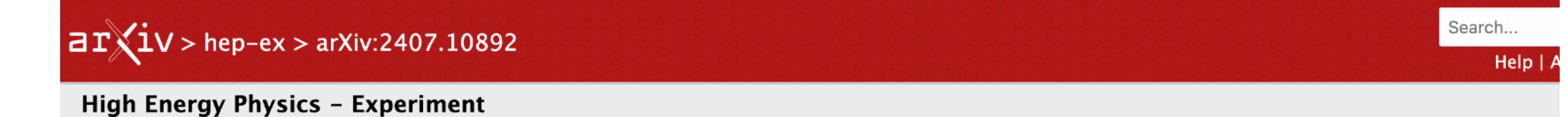


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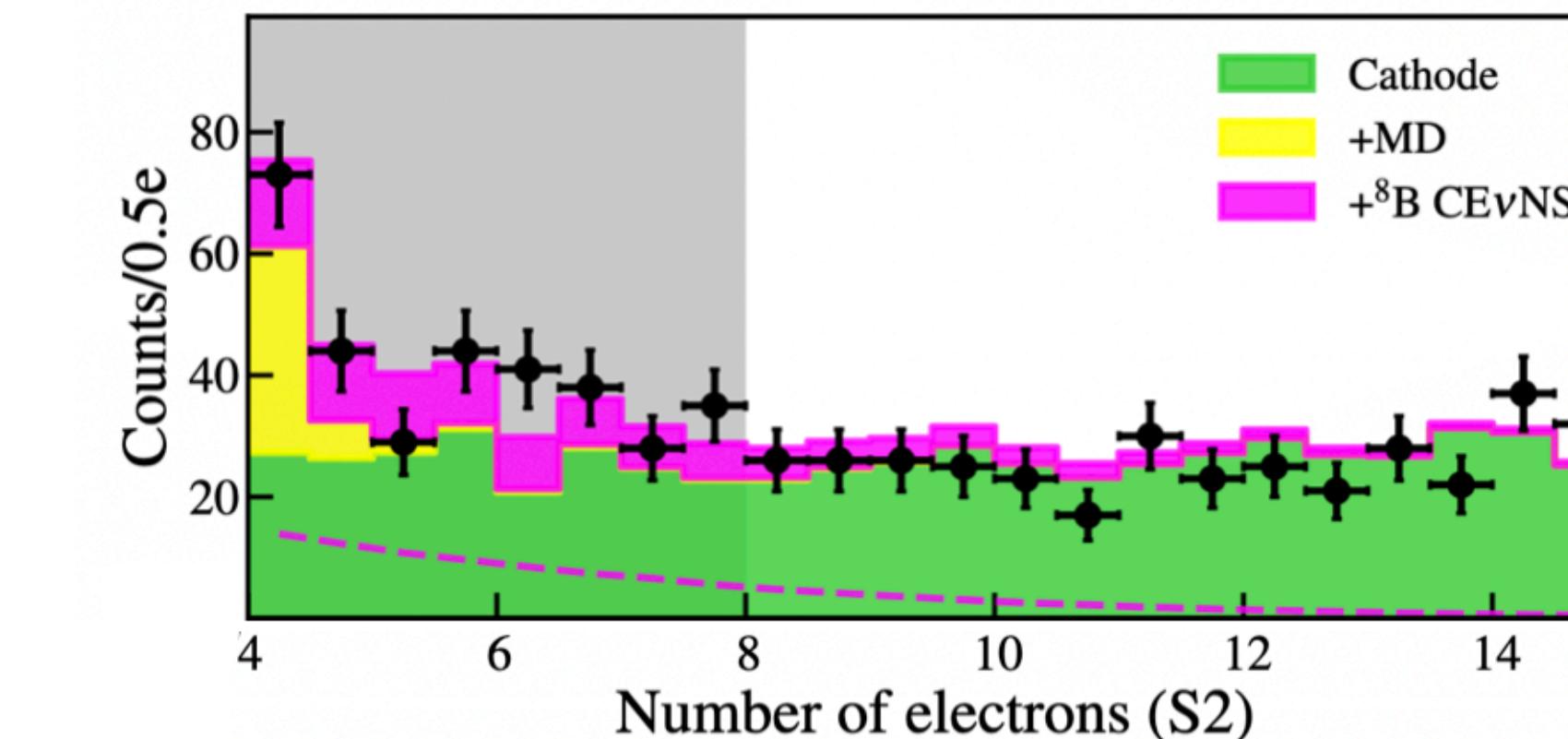
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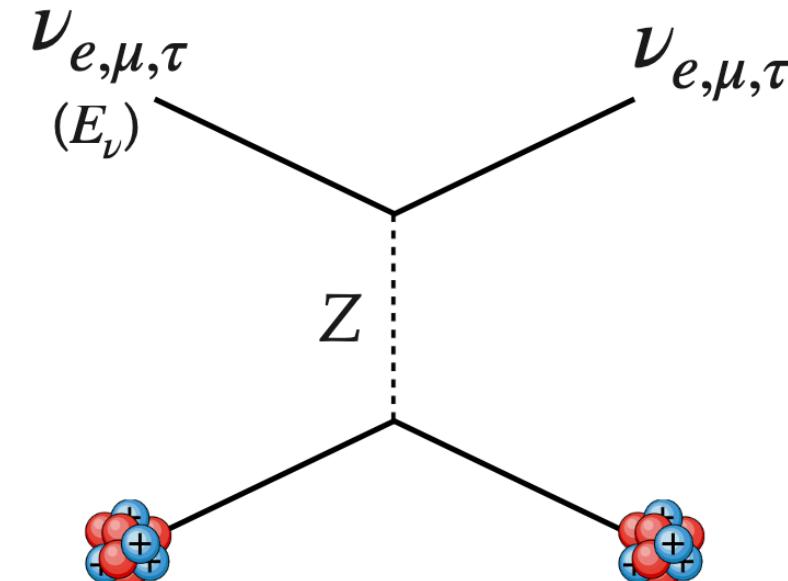
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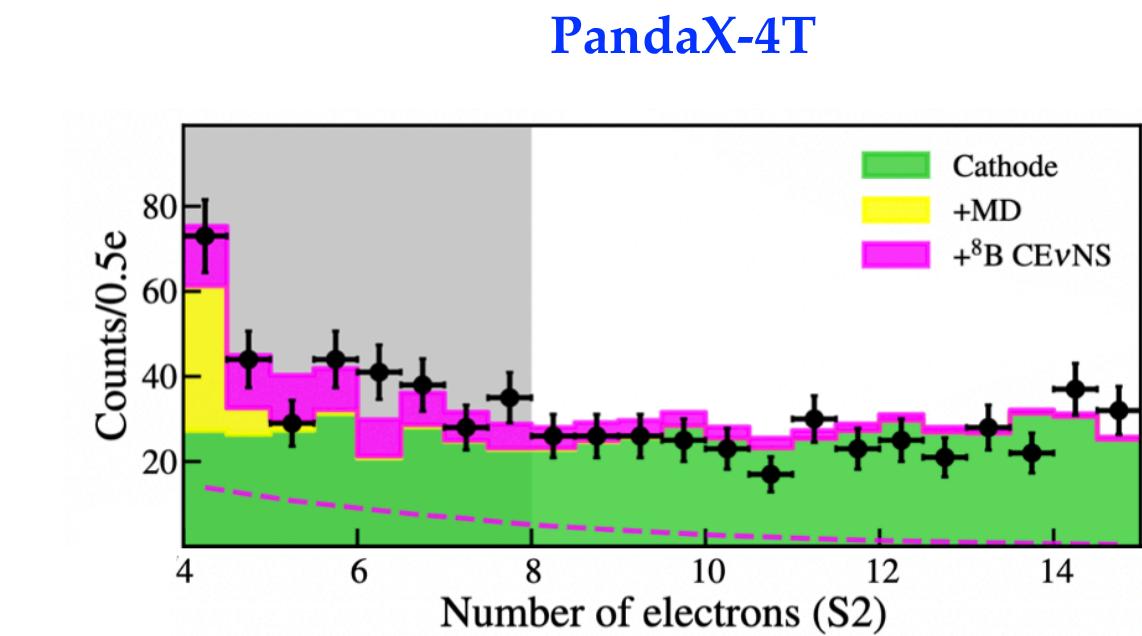
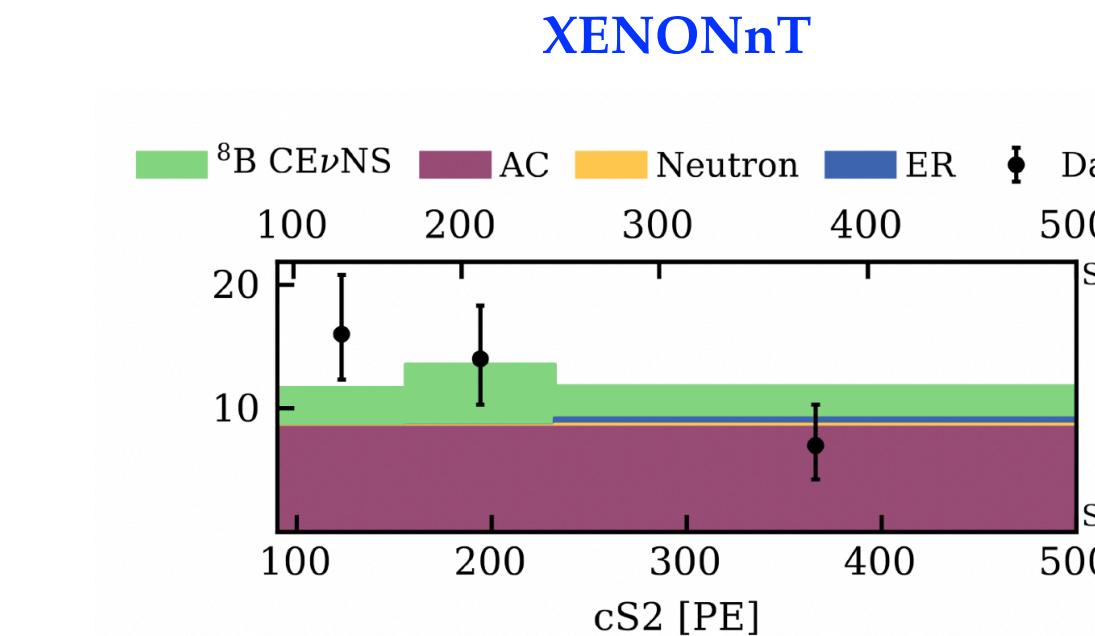
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Observing essentially the Standard Model process, can we say something new?

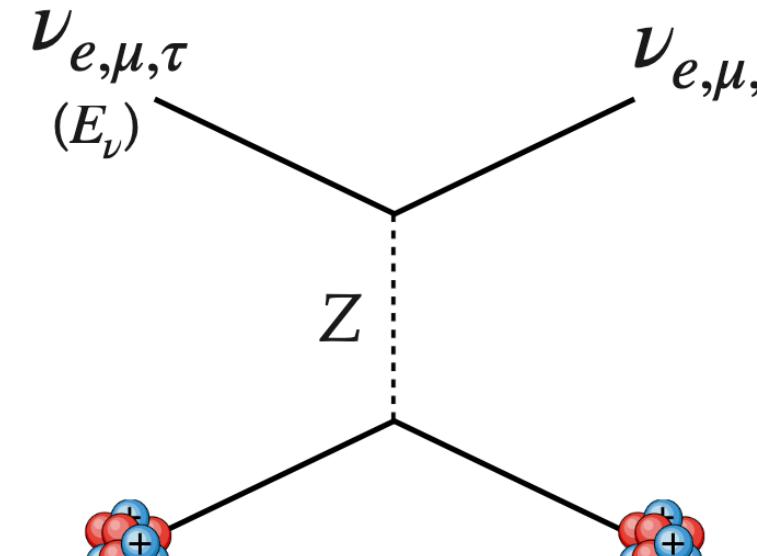
Our results: nuclear recoil



$$\frac{d\sigma}{dE_N} \propto f(\sin^2 \theta_W)$$



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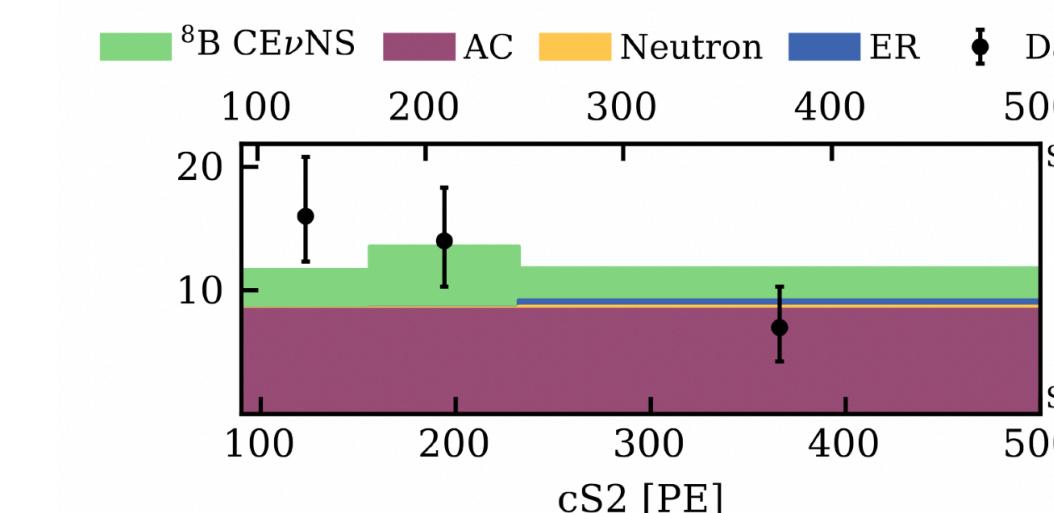


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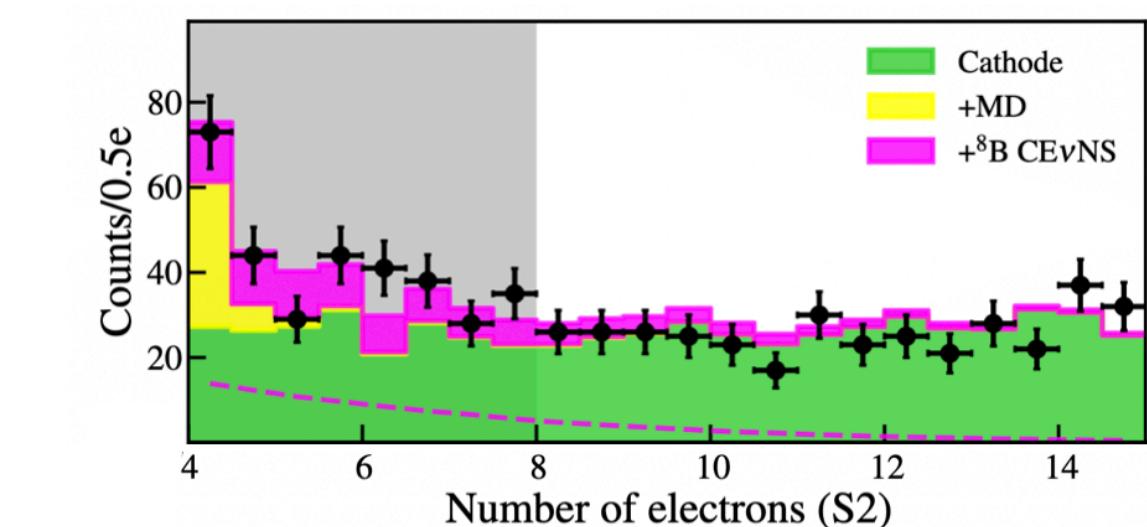
CE ν NS

Observed solar ${}^8\text{B}$ events depends on $\sin^2 \theta_W$

XENONnT

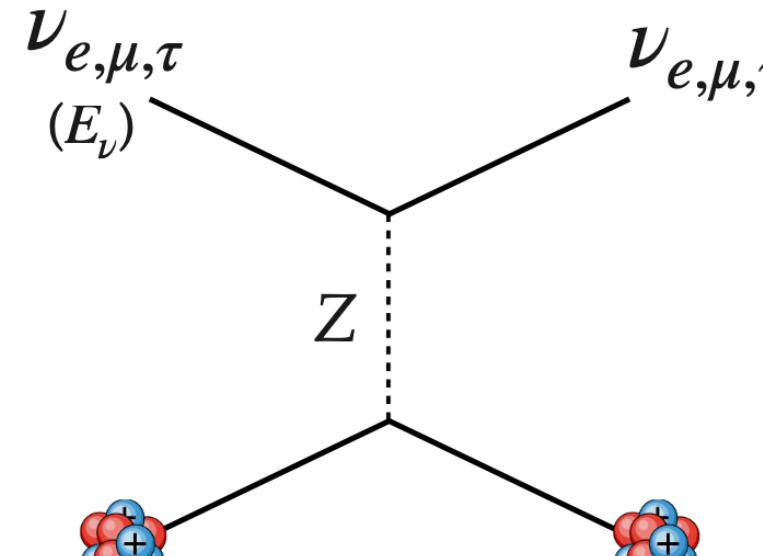


PandaX-4T



One can measure $\sin^2 \theta_W$ using these data

Our results: nuclear recoil

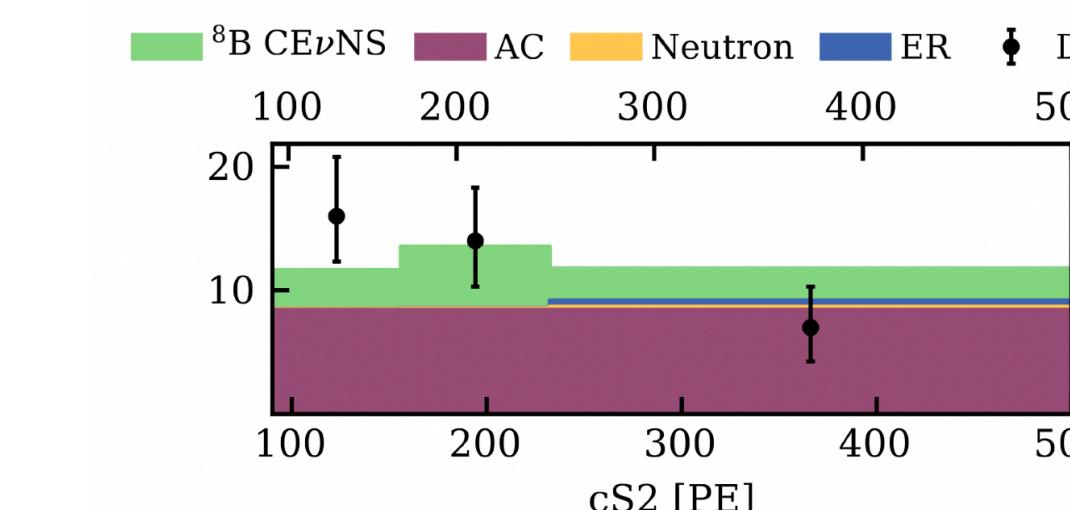


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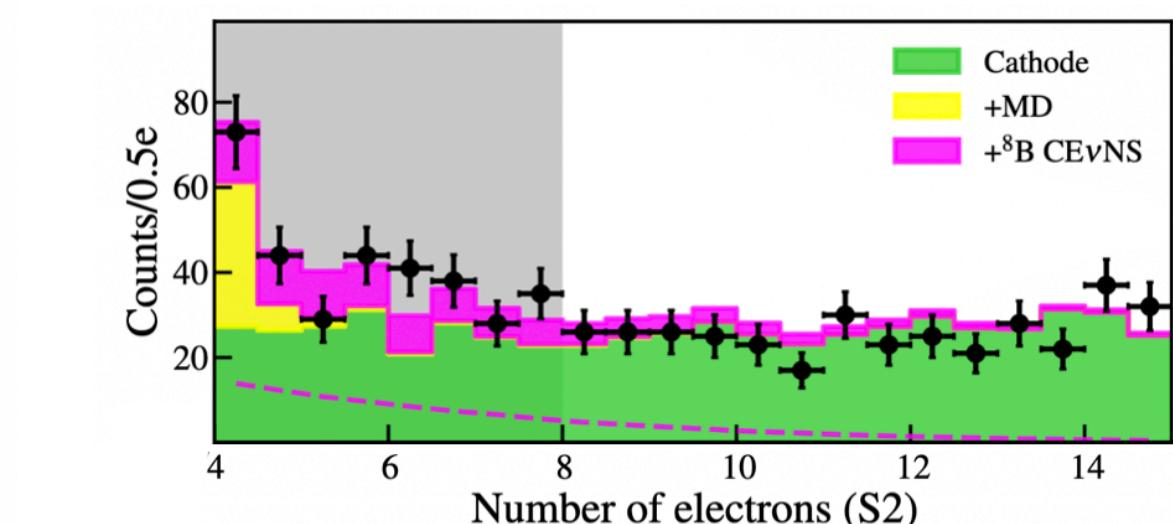
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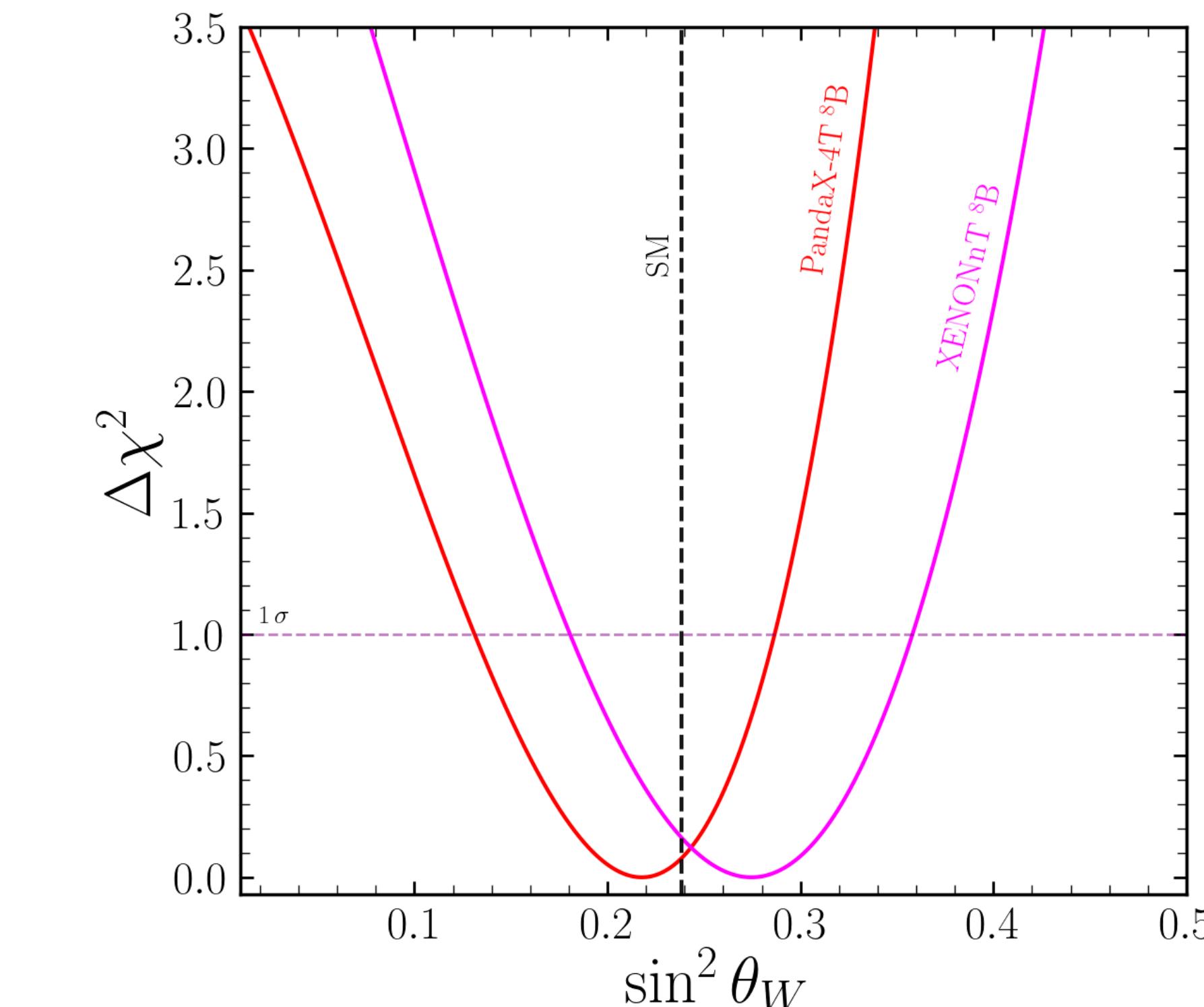
XENONnT



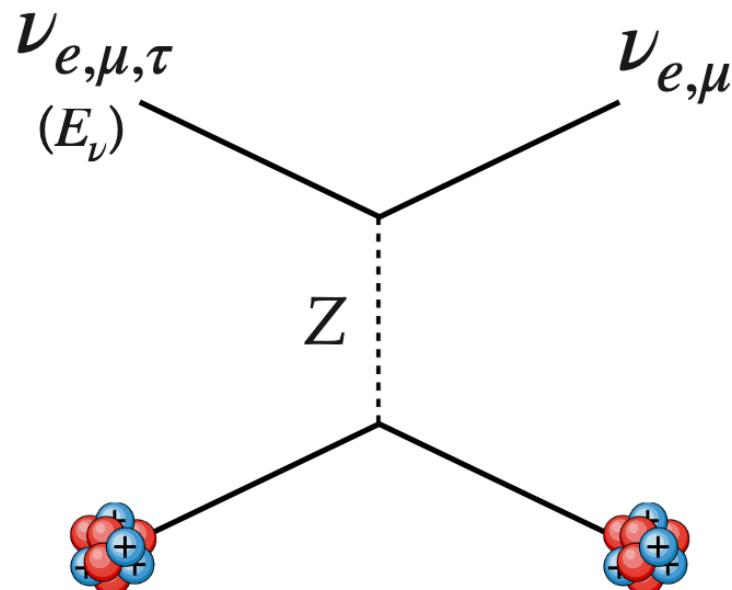
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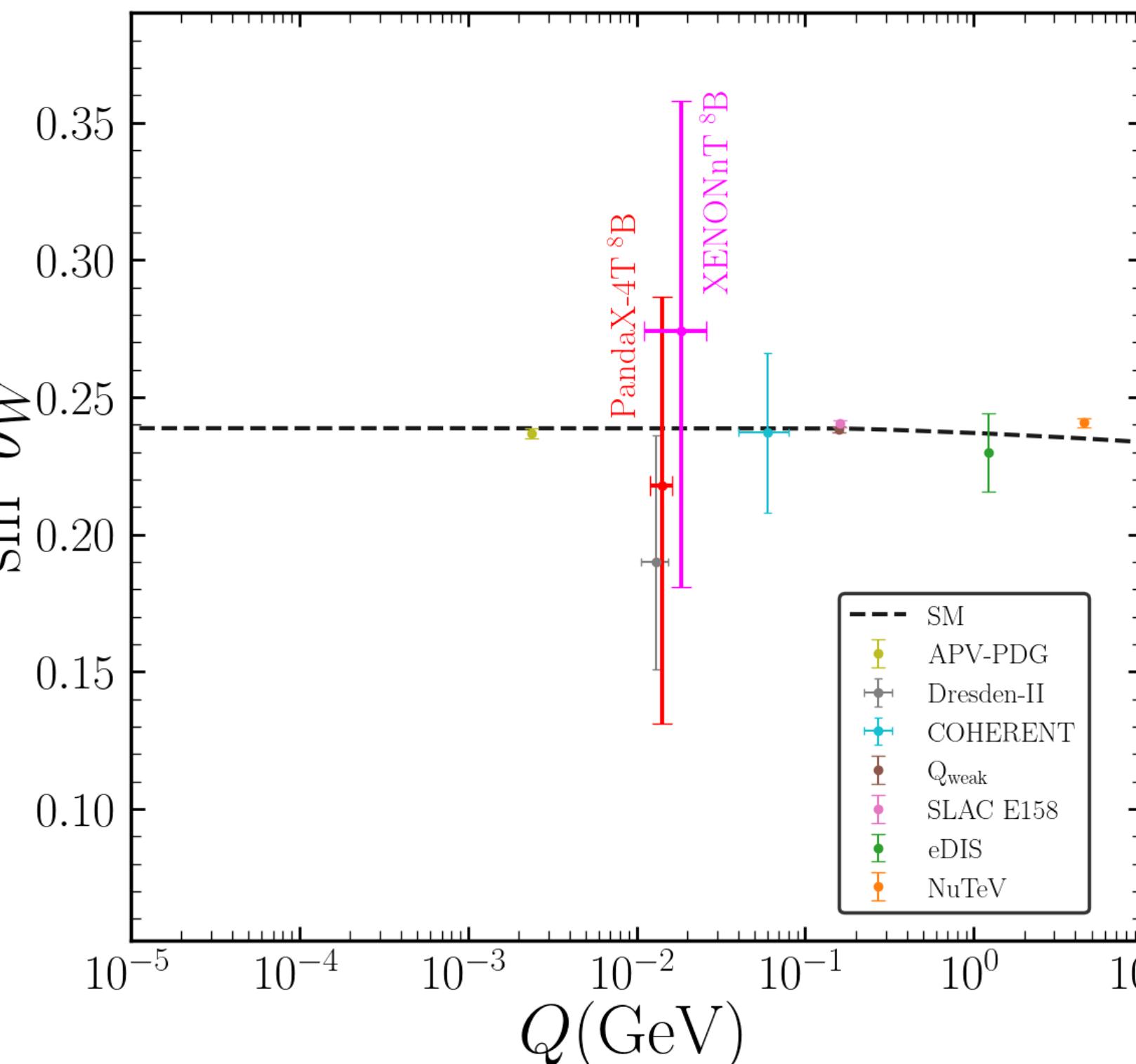
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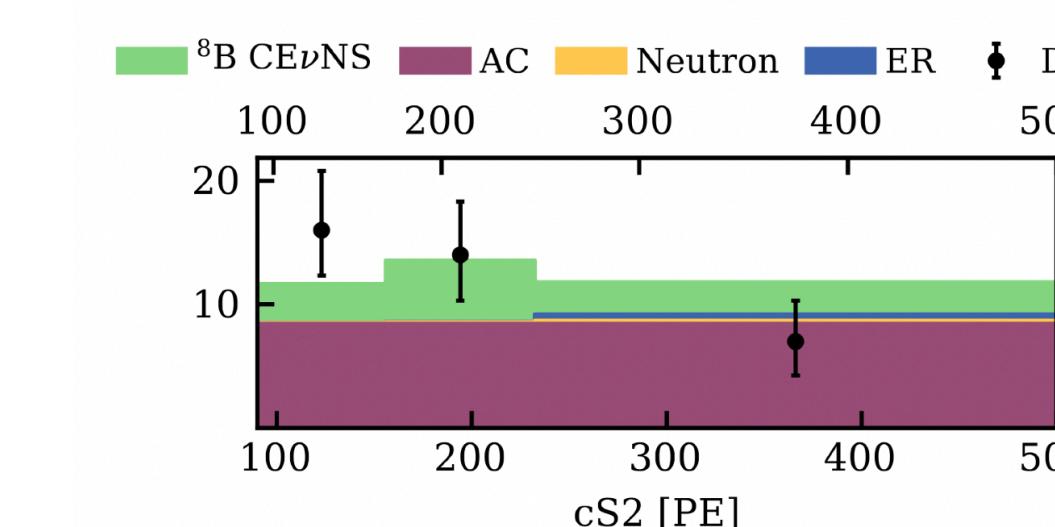
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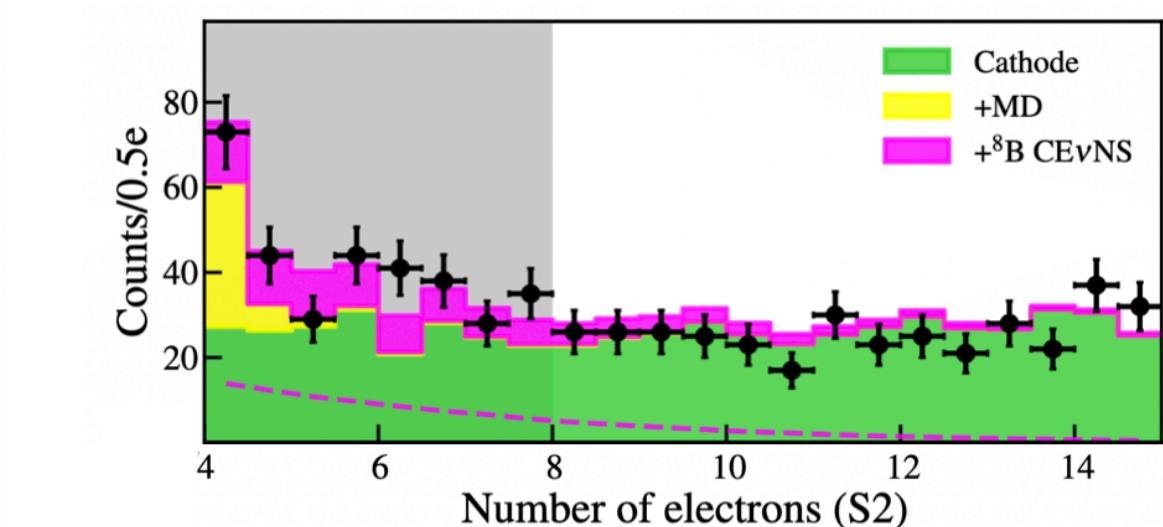


Energy scale determined from recoil energy regime

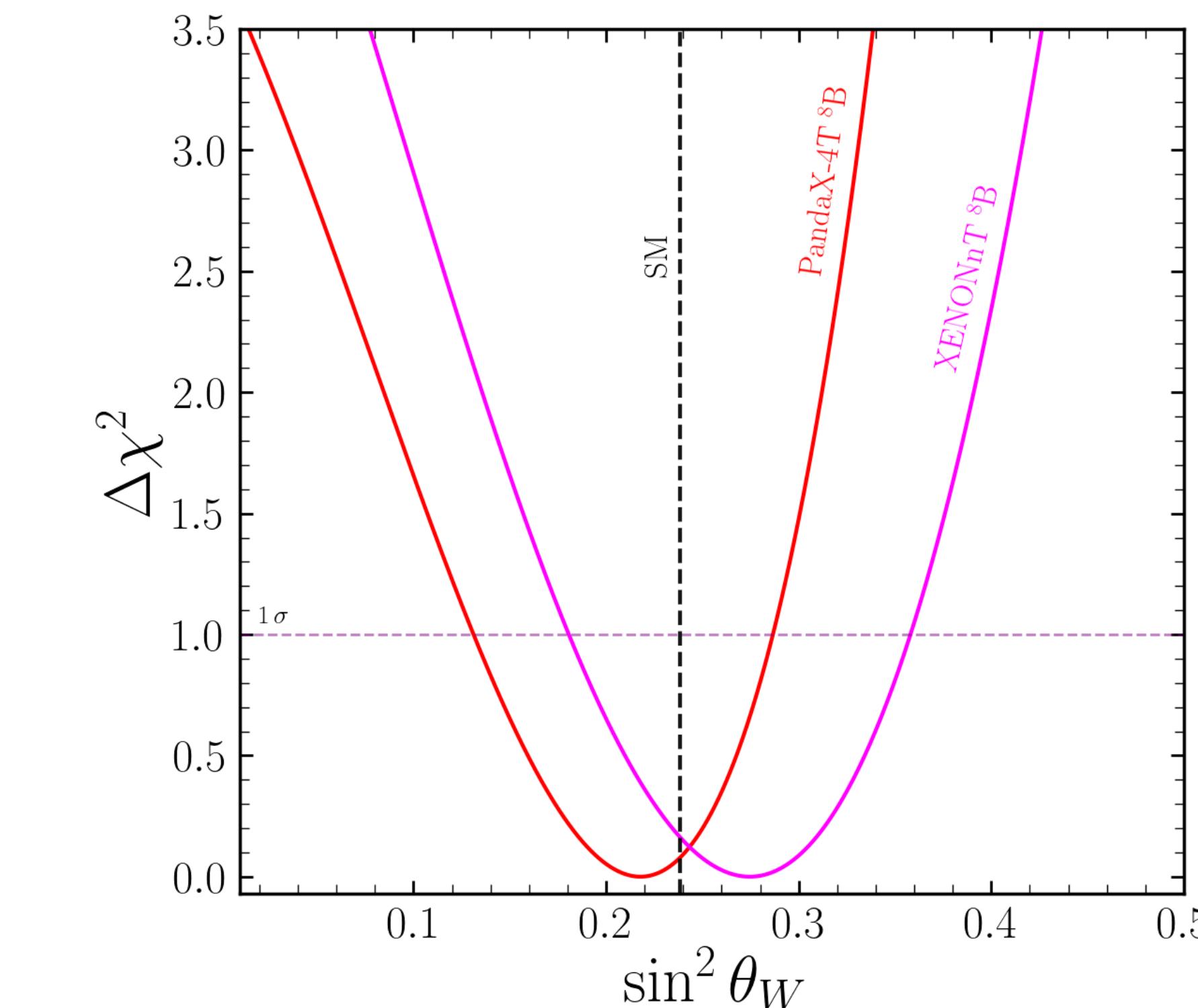
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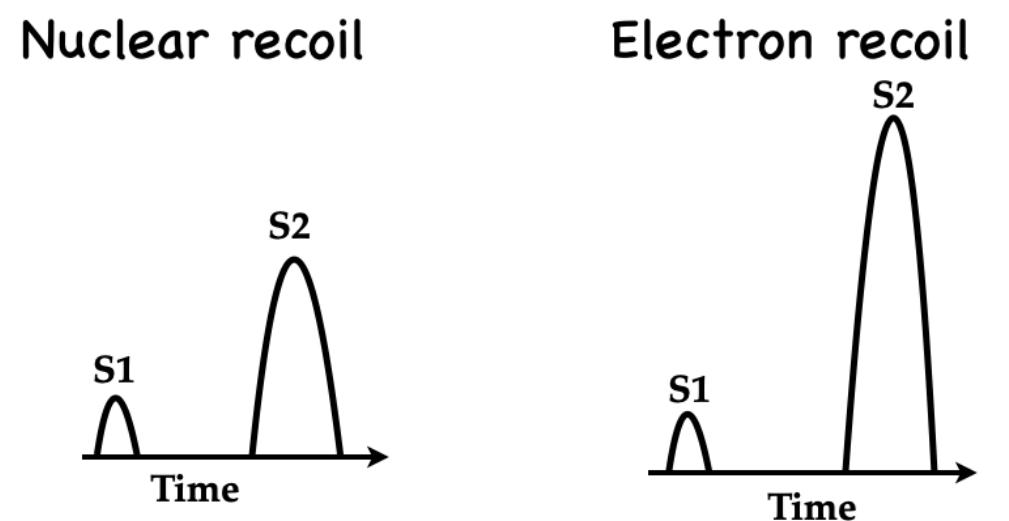


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Our results: electron recoil

S1-S2 only analysis

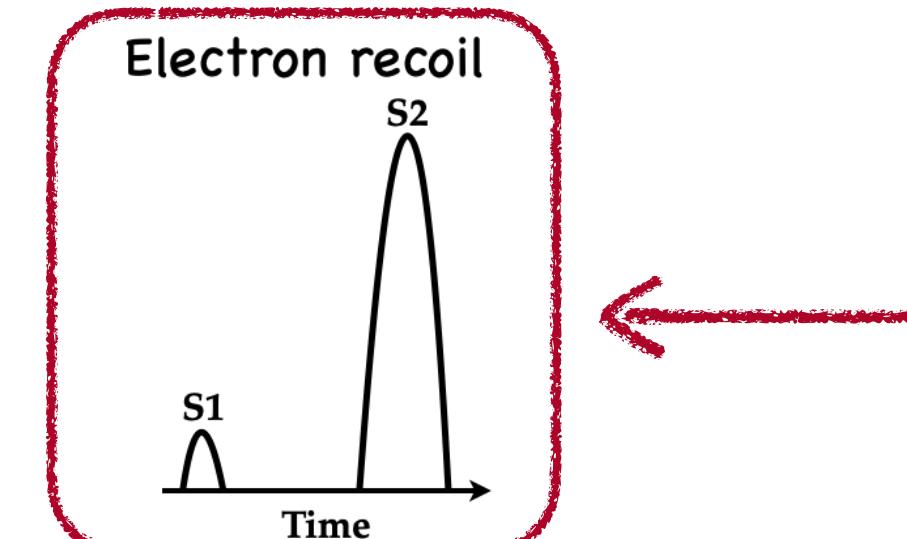
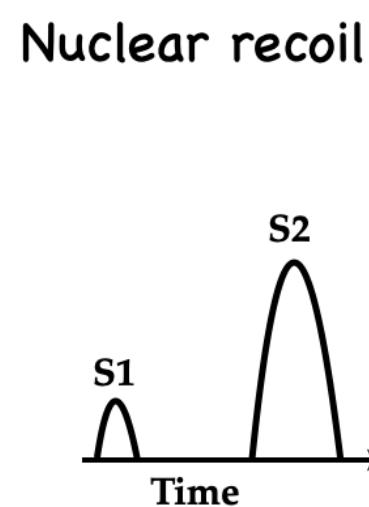


S2/S1 ratio - can distinguish - nuclear and electron recoil

$$E_{\text{recoil}} \gtrsim 0.5 \text{ keV}$$

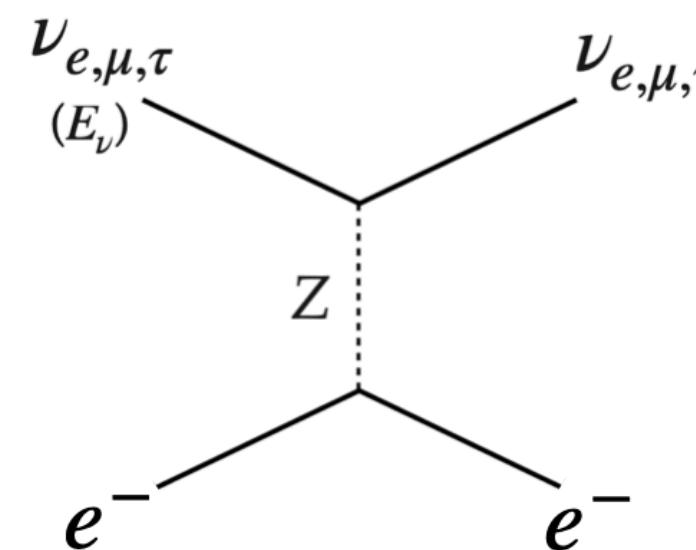
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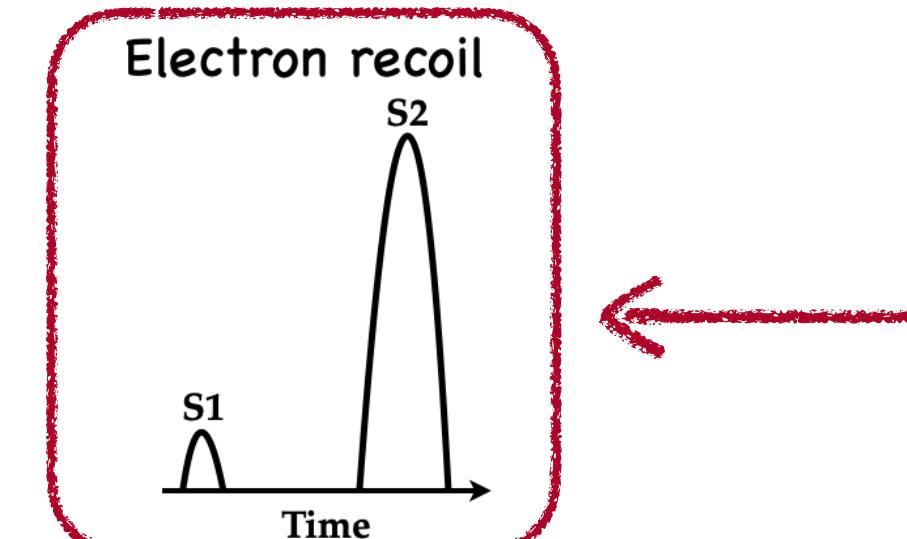
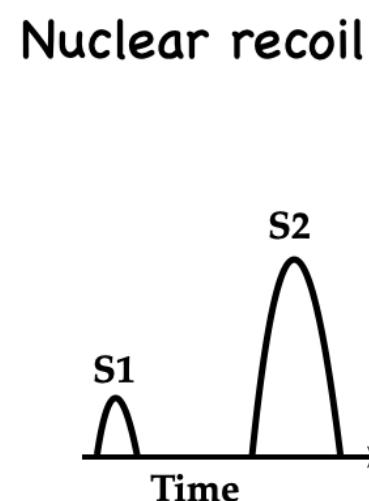
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Neutrino-electron scattering

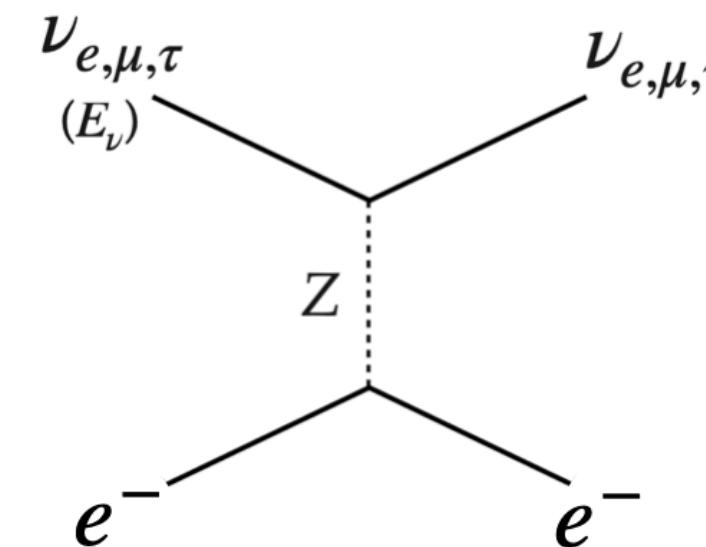
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Neutrino-electron scattering

Observed $\nu - e$ events: ~ 60

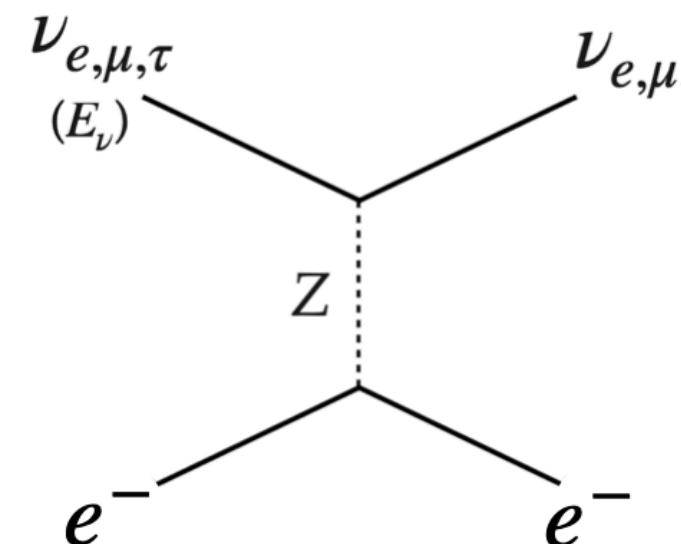
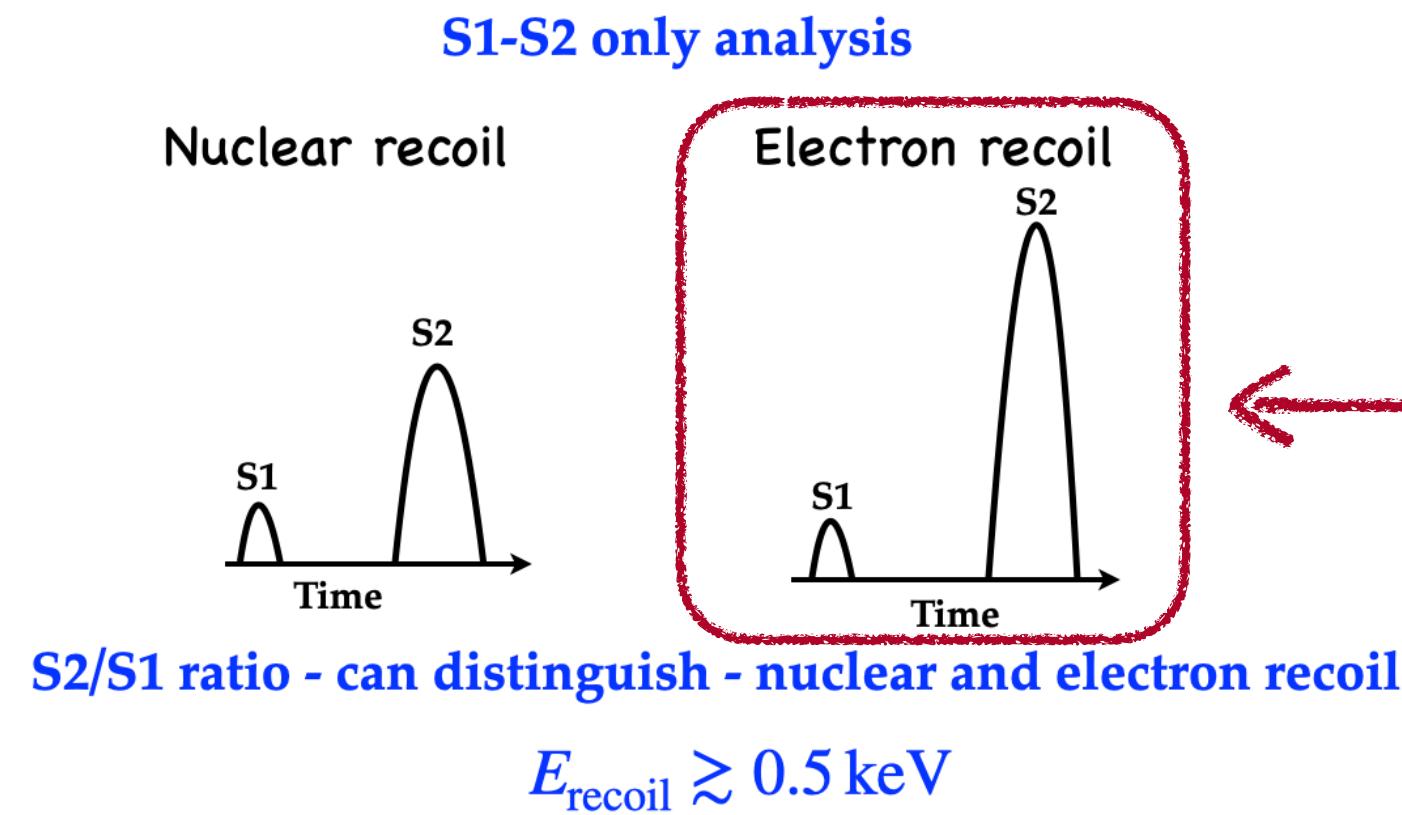
But statistically not significant due
to huge background

XENONnT 2207.11330

LZ 2307.15753

PandaX-4T 2408.07641

Our results: electron recoil

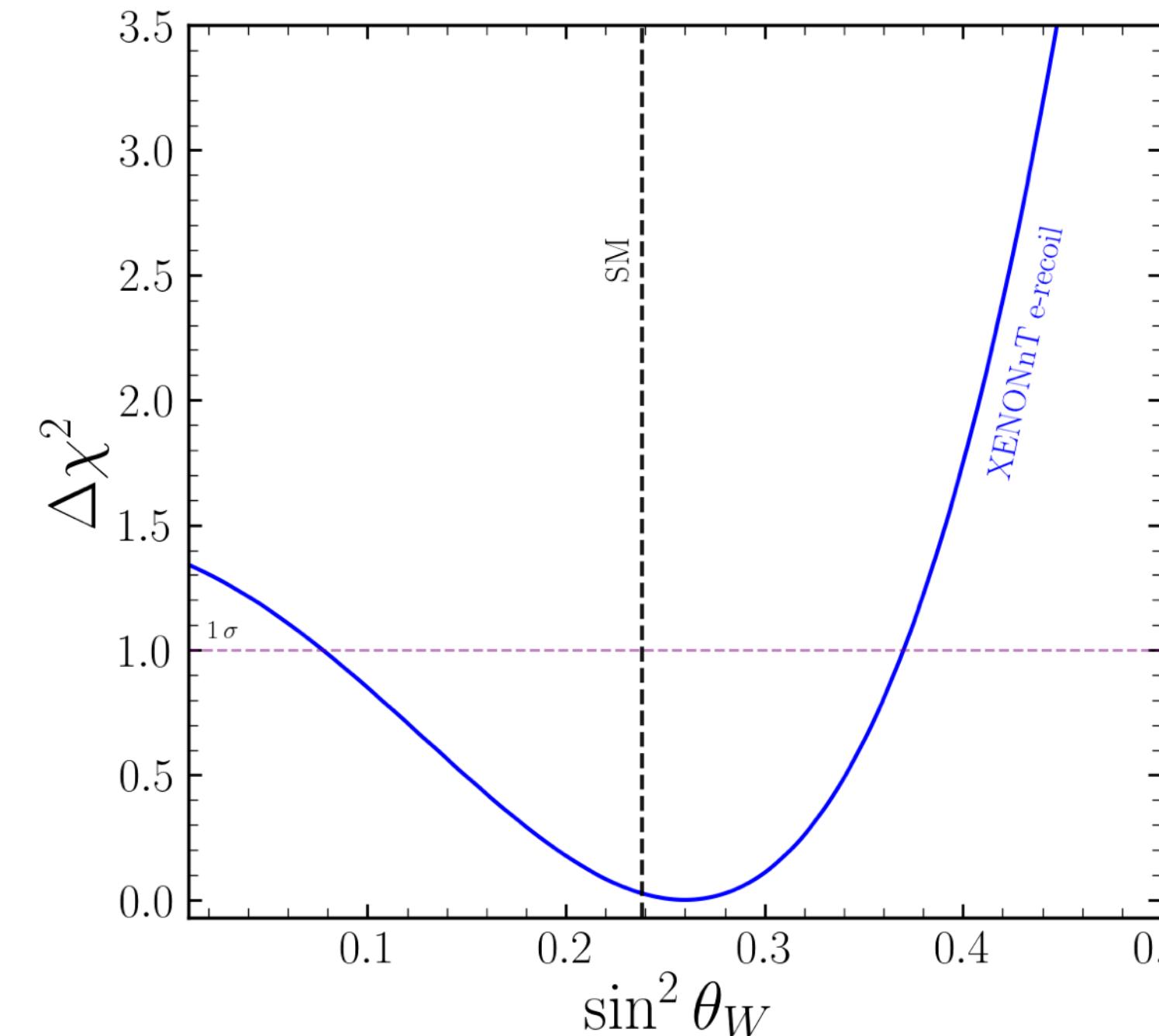
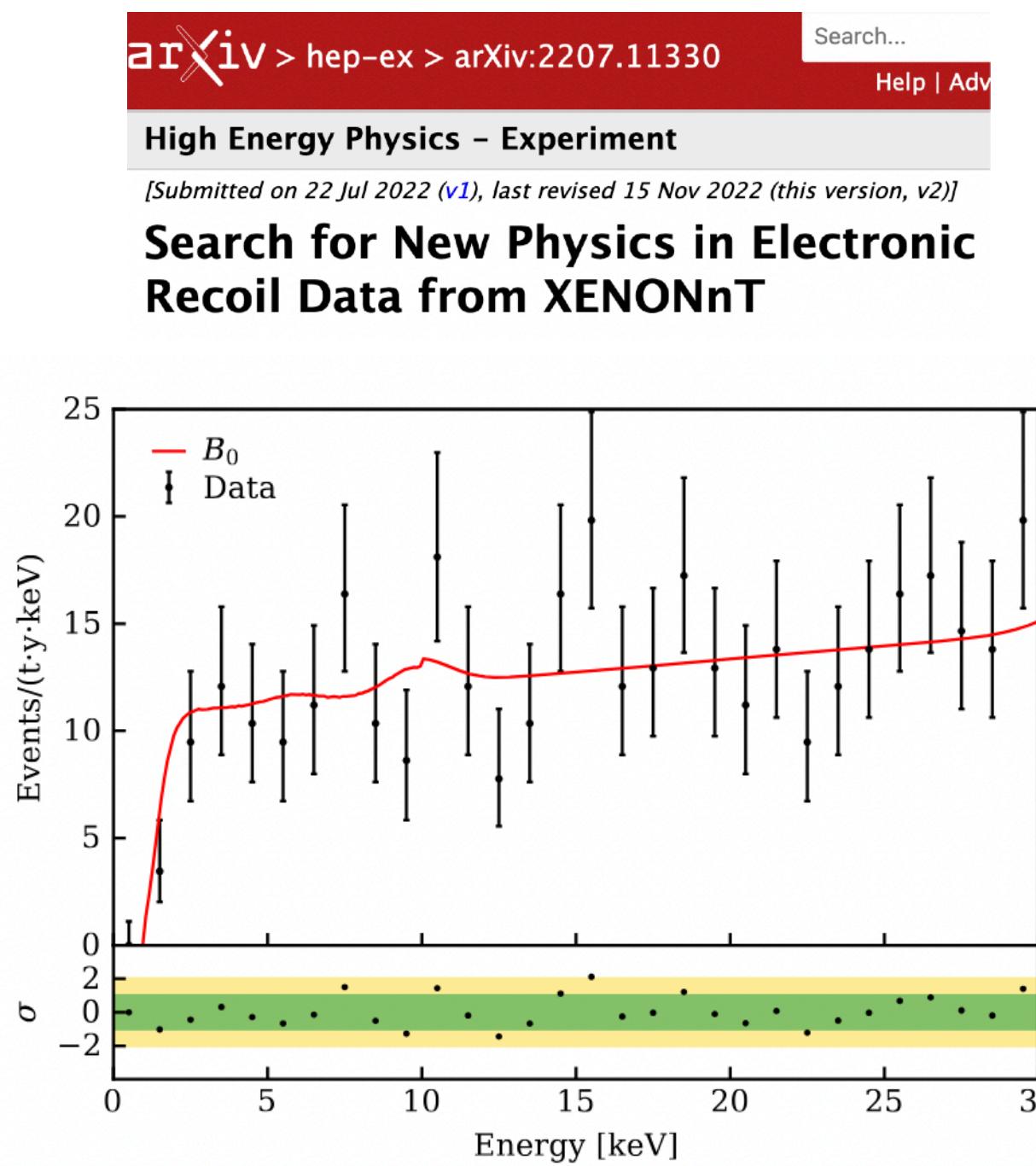


Neutrino-electron scattering

Observed $\nu - e$ events: ~ 60

But statistically not significant due
to huge background

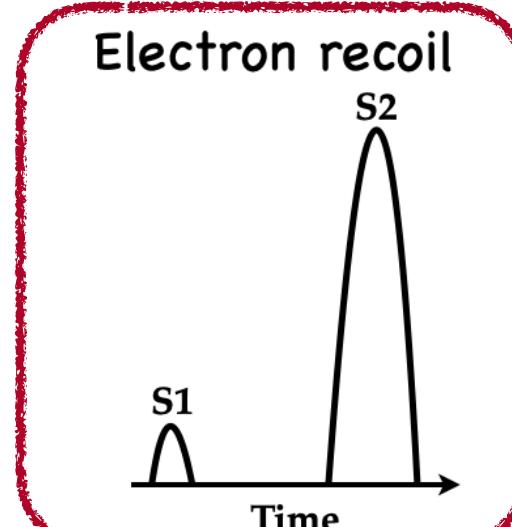
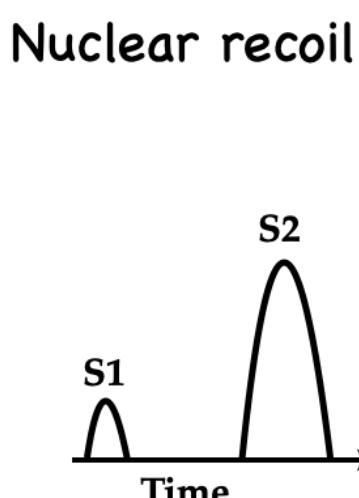
XENONnT 2207.11330
LZ 2307.15753
PandaX-4T 2408.07641



Only upper limits above 1.16σ

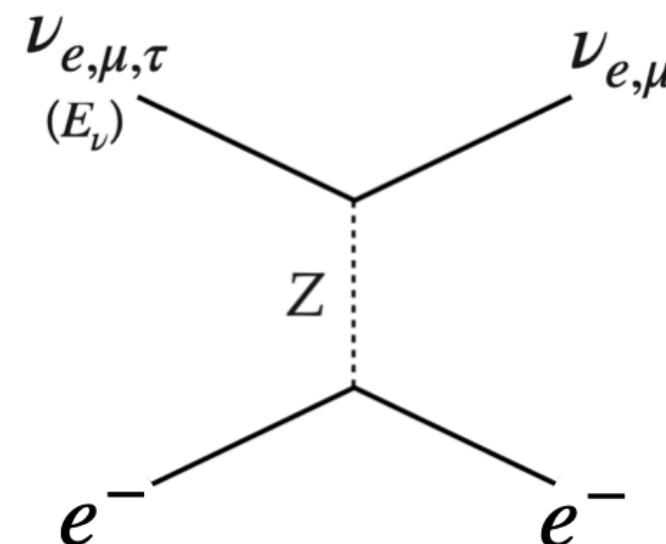
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S1-S2 only analysis



S2/S1 ratio - can distinguish - nuclear and electron recoil

$$E_{\text{recoil}} \gtrsim 0.5 \text{ keV}$$



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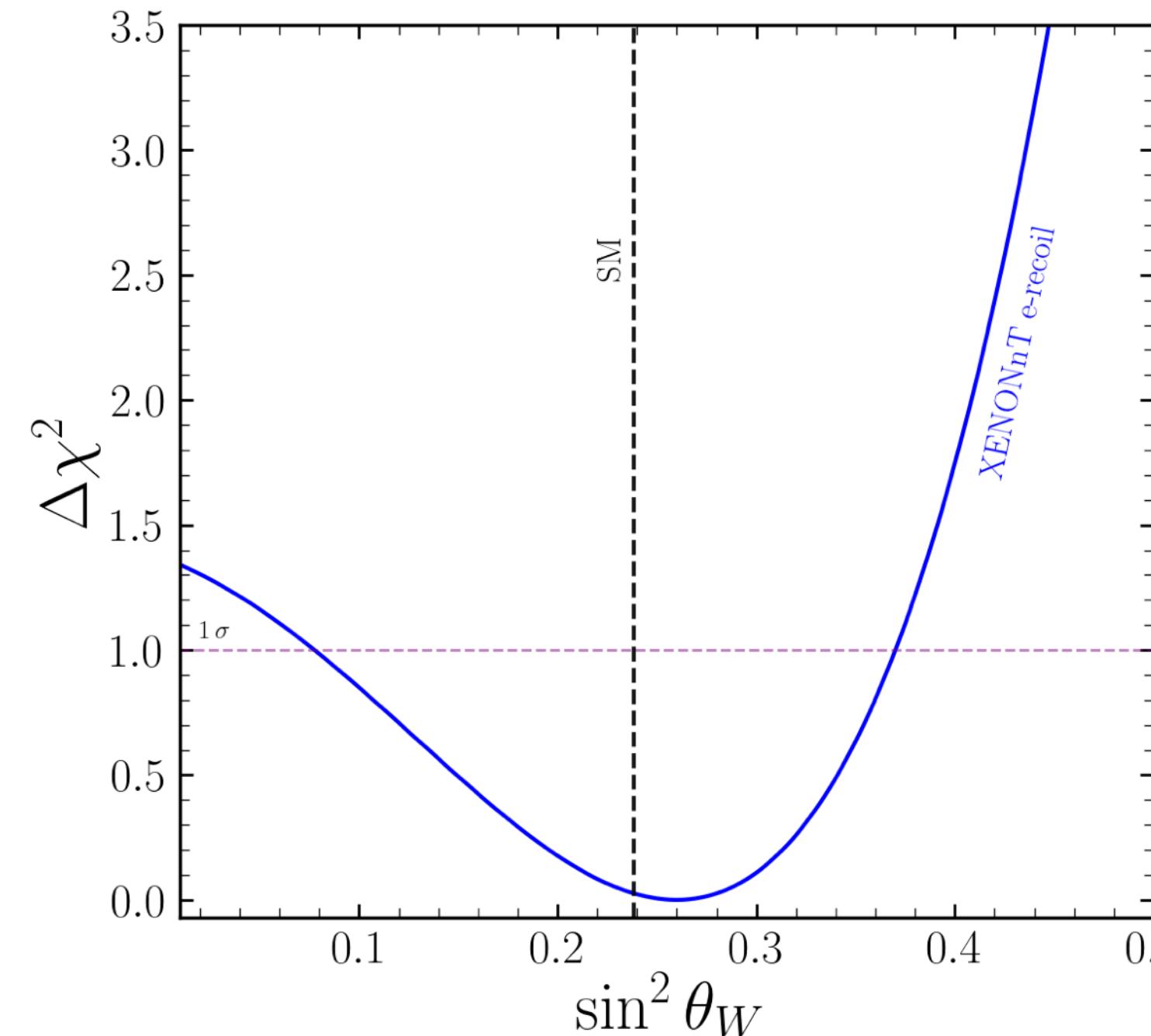
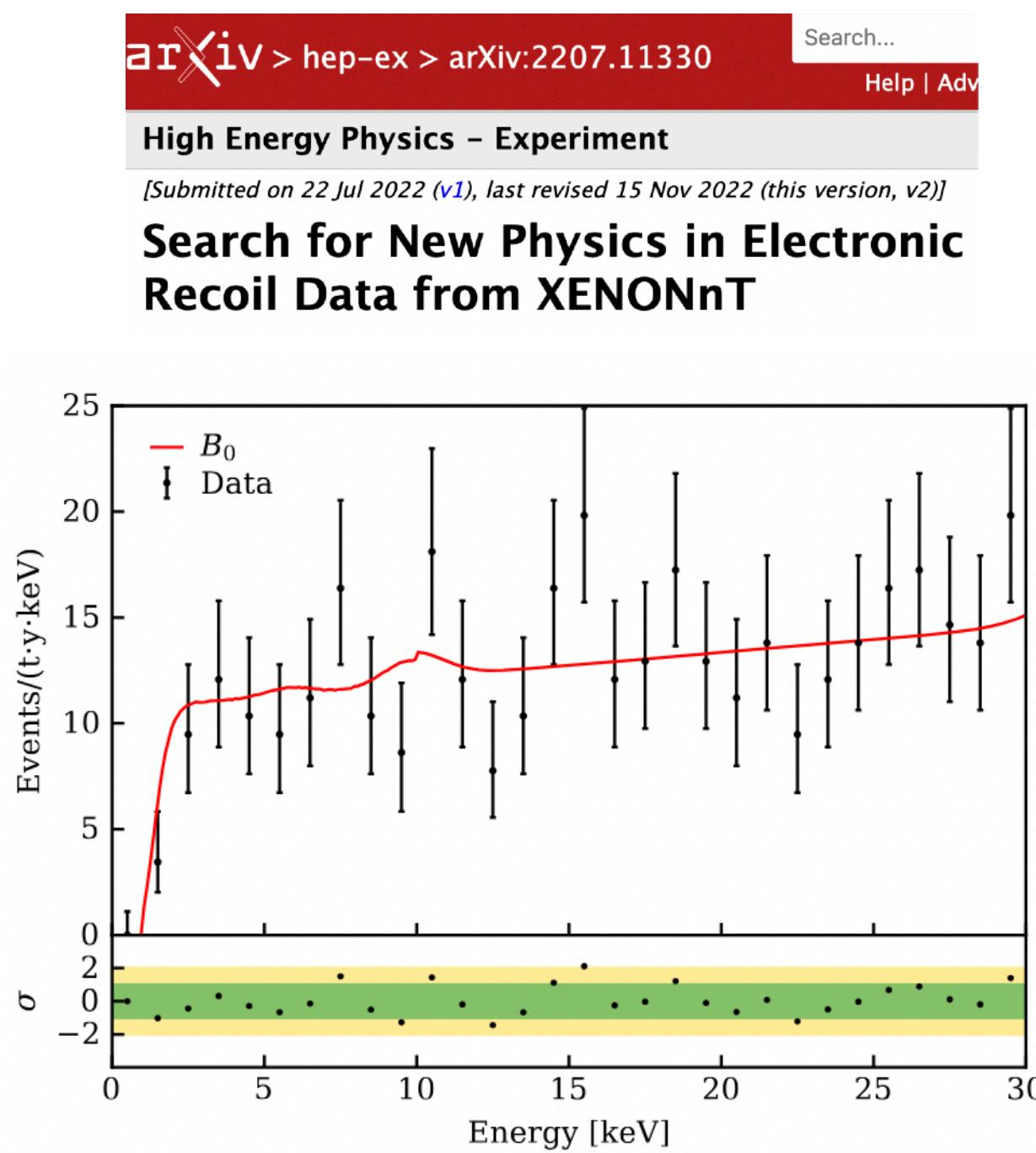
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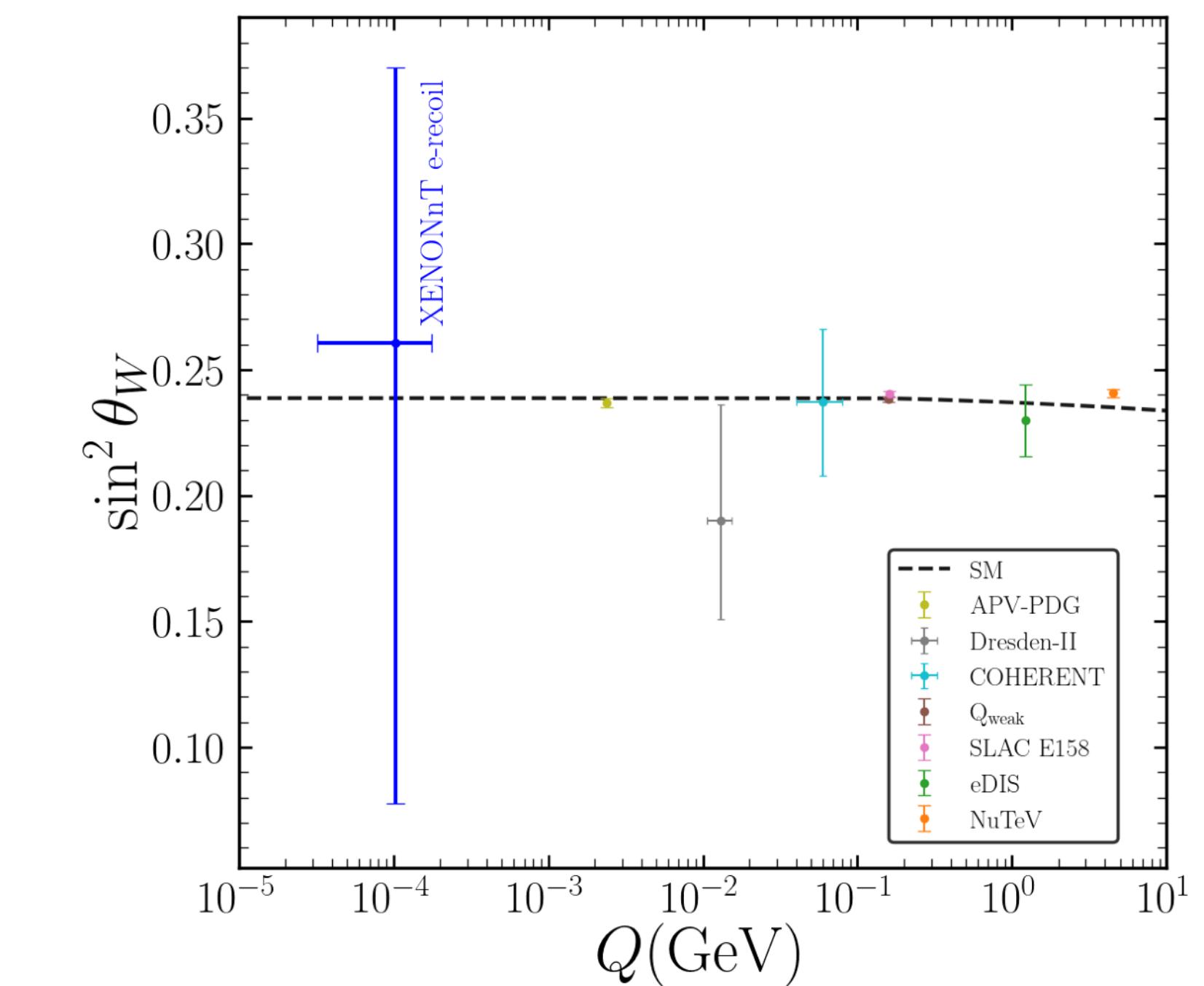
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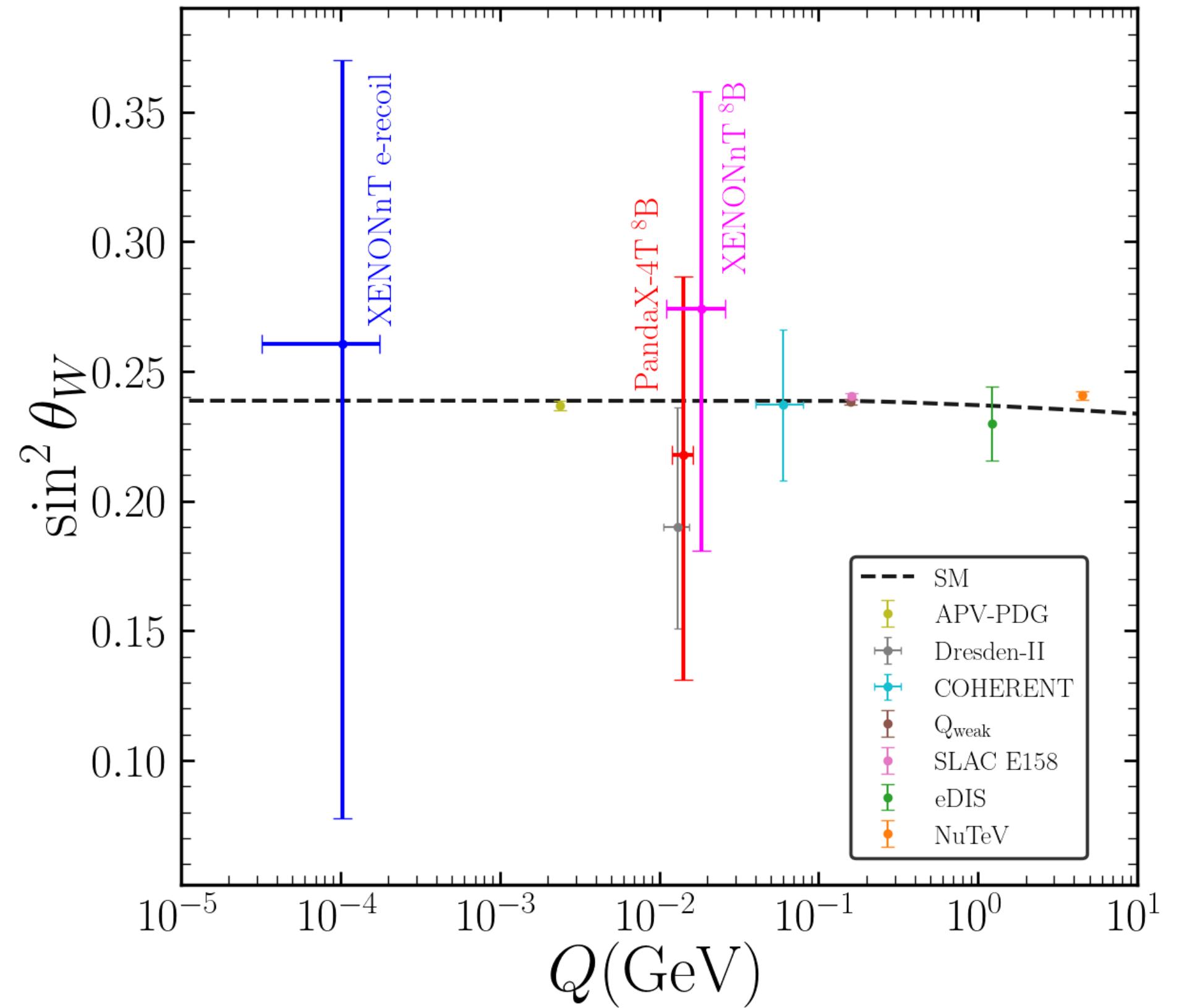
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Probing a SM parameter in
an entirely new regime

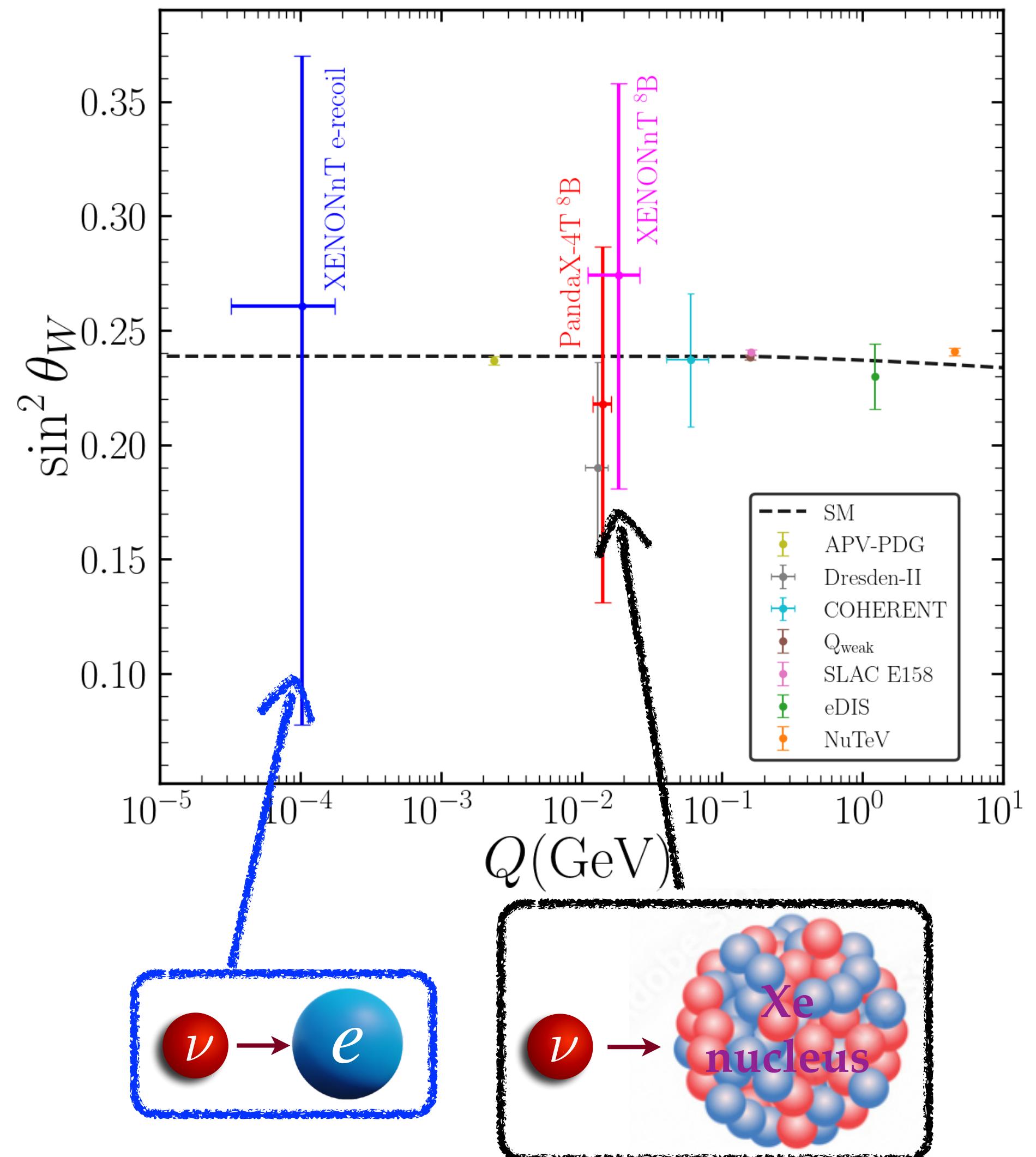
Summary

DD already probing SM in an uncharted territory



Summary

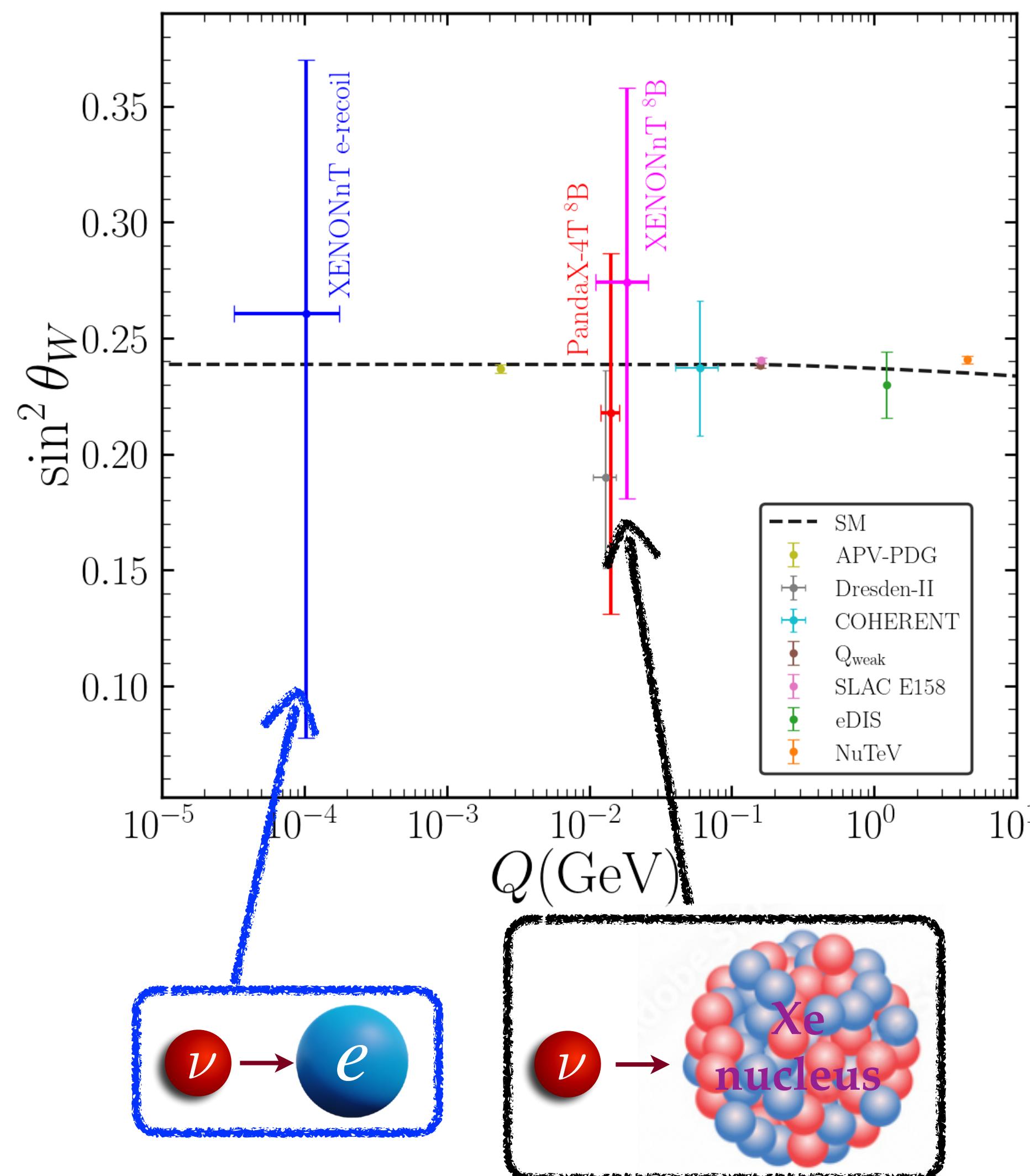
DD already probing SM in an uncharted territory



While recoil energy of our analysis similar but
heavy Xe nucleus shifts CE ν NS in ~ 10 MeV scale

Summary

DD already probing SM in an uncharted territory



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TNM, Boehm; 2409.04385

Tarak Nath Maity

Future Xe based: DARWIN/XLZ, PandaX-xT

trino observatories. With 300 ty, DARWIN would be able to achieve 0.15% precision in the pp flux measurement, ap-

DARWIN, 2006.03114

Future Ar based: DarkSide-20k

With 10 years exposure, the neutrino fog can be reached for WIMP masses around $5 \text{ GeV}/c^2$.

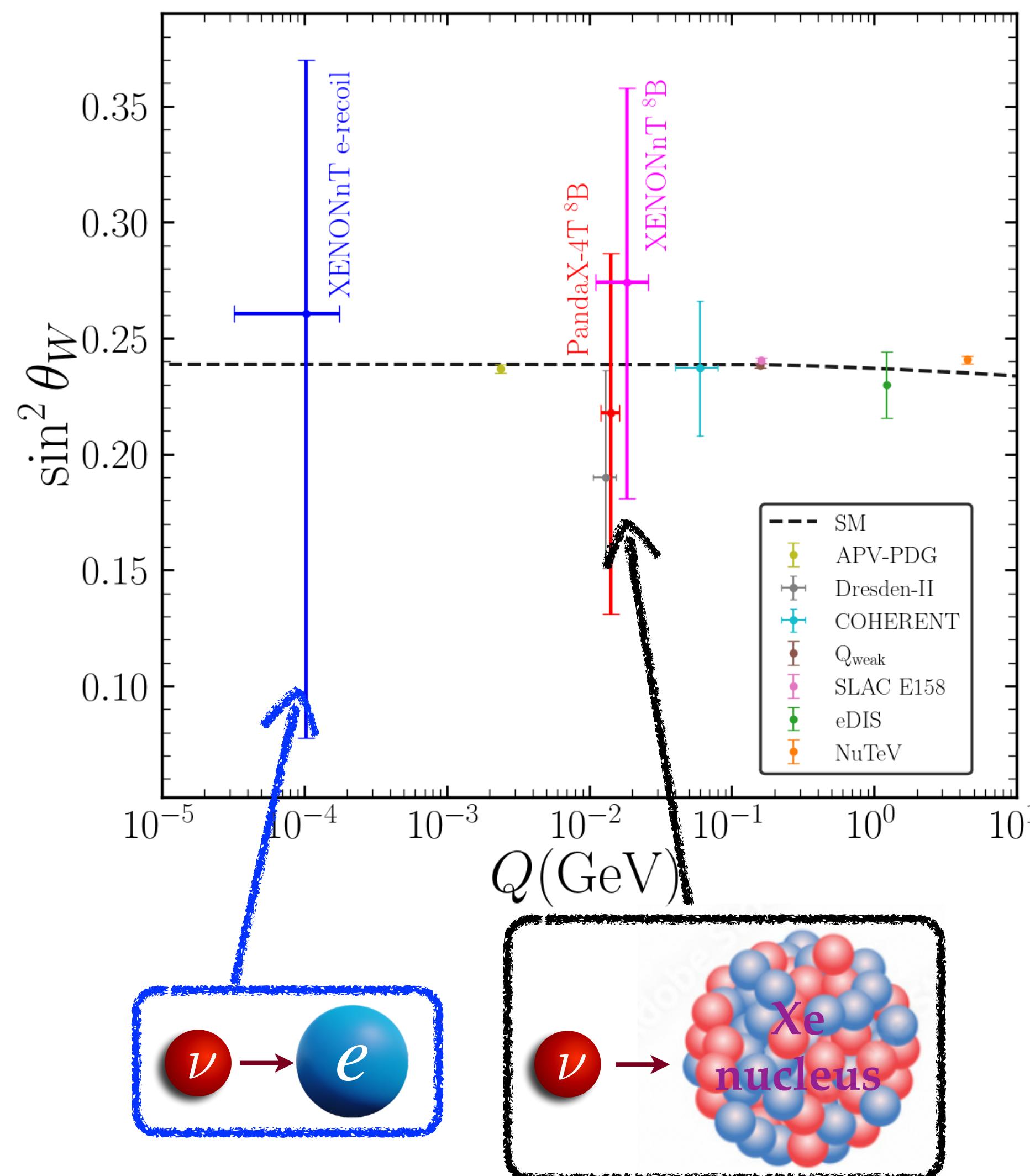
DarkSide-20k, 2407.05813

Future low threshold experiment: Oscura

Expected to see some CE ν NS events - no detailed analysis yet

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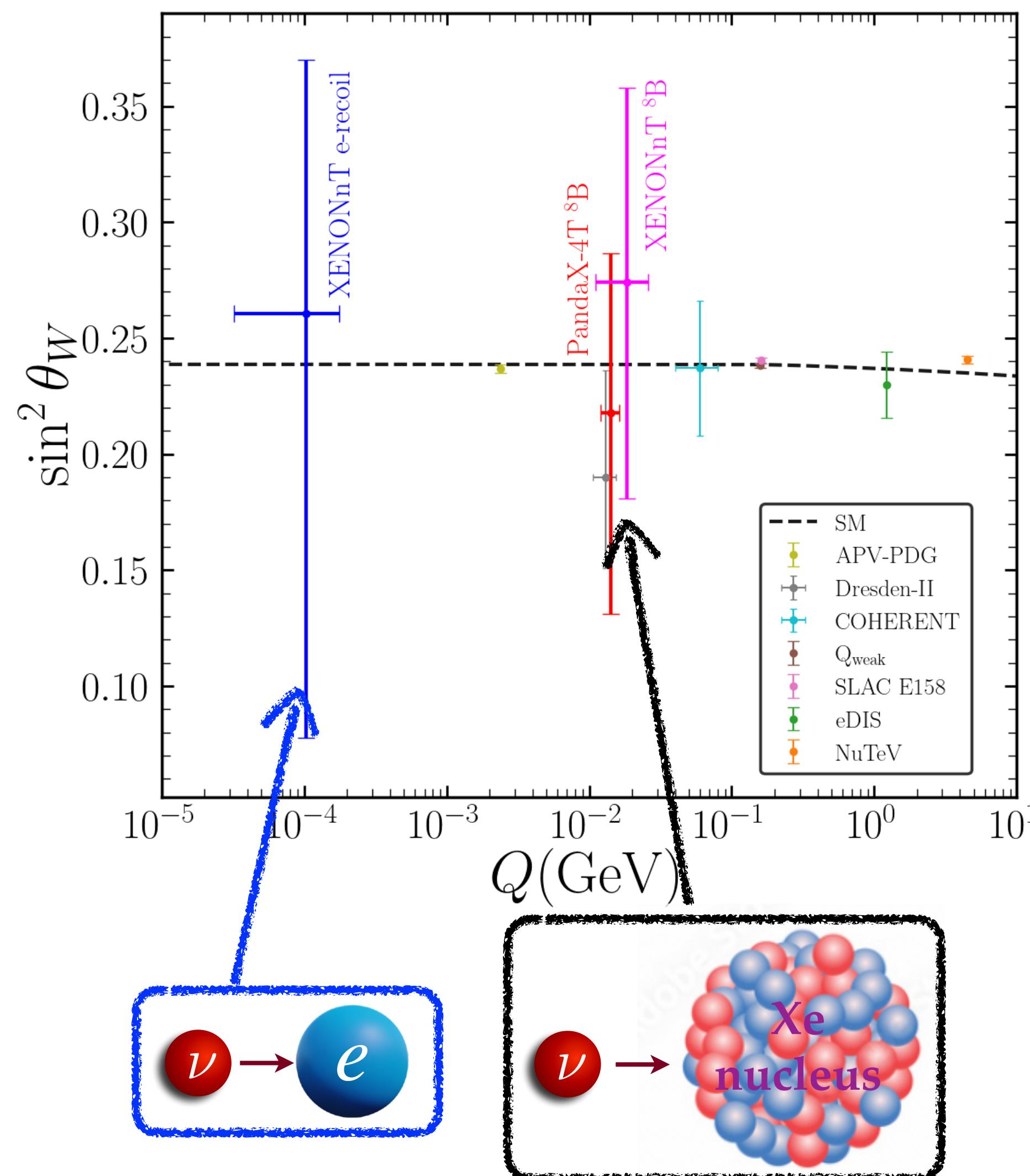
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email: tarak.maity.physics@gmail.com

Thank you

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