

# Physics opportunities with kaon decay-at-rest neutrinos: search for sterile neutrino and non-standard interactions

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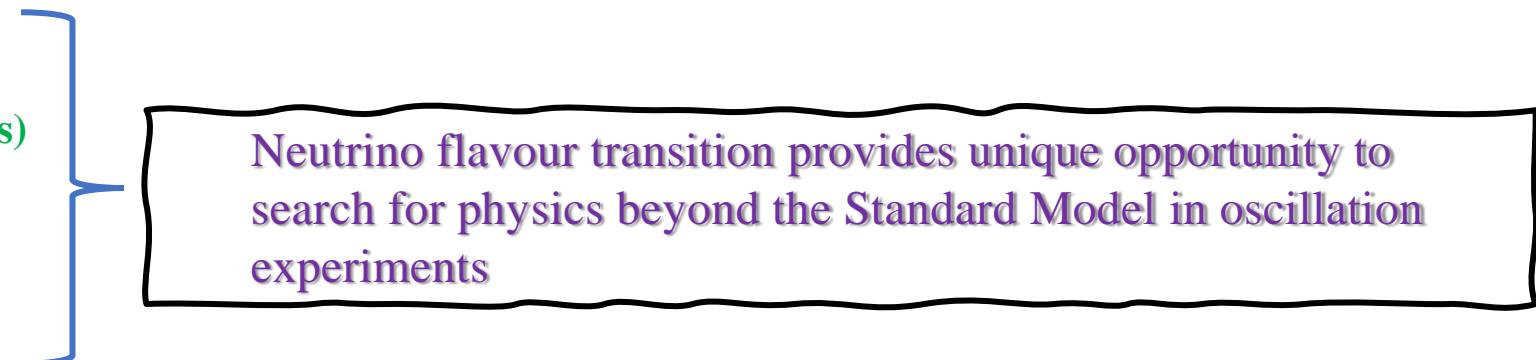
# Introduction: neutrino oscillation and new physics

- Neutrino flavour oscillation arises from mixing between flavour states ( $\nu_e, \nu_\mu, \nu_\tau$ ) and mass eigenstates ( $\nu_1, \nu_2, \nu_3$ ) of neutrinos.
- The 3-flavour oscillation probability depends upon 3 mixing angles ( $\theta_{12}, \theta_{13}, \theta_{23}$ ), 2 independent mass squared differences ( $\Delta m_{21}^2, \Delta m_{31}^2$ ) and 1 CP phase  $\delta_{CP}$ .
- In three flavour standard neutrino oscillation picture the important unknown parameters are
  1. Sign of  $\Delta m_{31}^2$  (Neutrino mass ordering)
  2.  $\delta_{CP}$  (CP violating phase)
  3. Octant of  $\theta_{23}$  ( $\theta_{23} > 45^\circ$  or  $\theta_{23} < 45^\circ$ )
- While the standard three flavour oscillation framework is firmly established by current data; **subdominant effects can not be ruled out completely**
- Current and future neutrino oscillation experiments are aimed to measure these parameters. But...
- There are several “**New Physics**” scenarios which can significantly impact the determination of these unknowns

<http://dx.doi.org/10.21468/SciPostPhysProc.2.001>

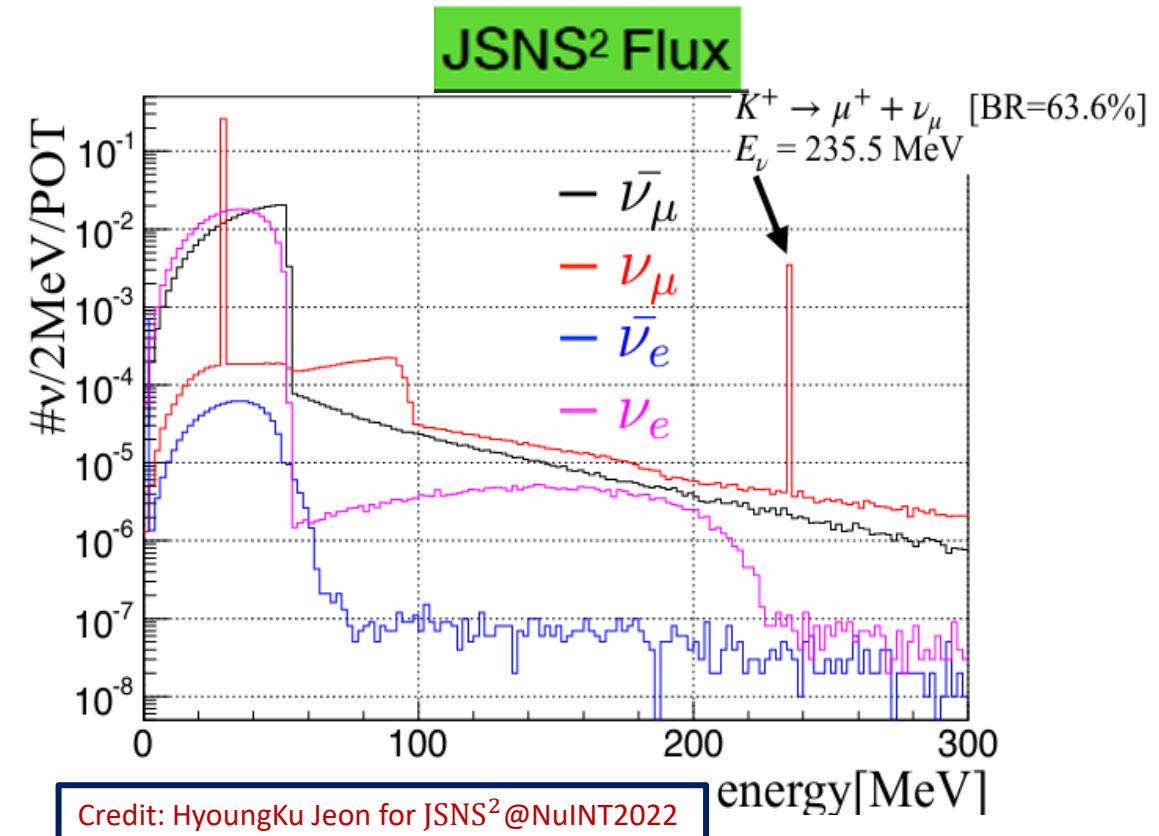
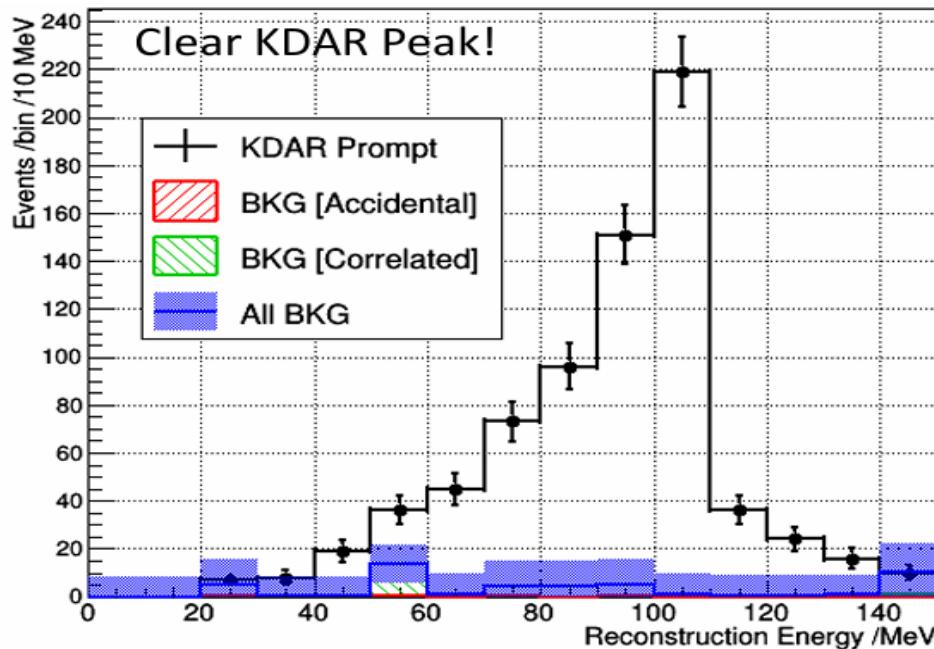
## ► New Physics:

1. Sterile neutrinos
2. Non-standard interactions (NSIs)
3. Neutrino decoherence and decay
4. Unitarity violation
5. LIV/CPT, etc...



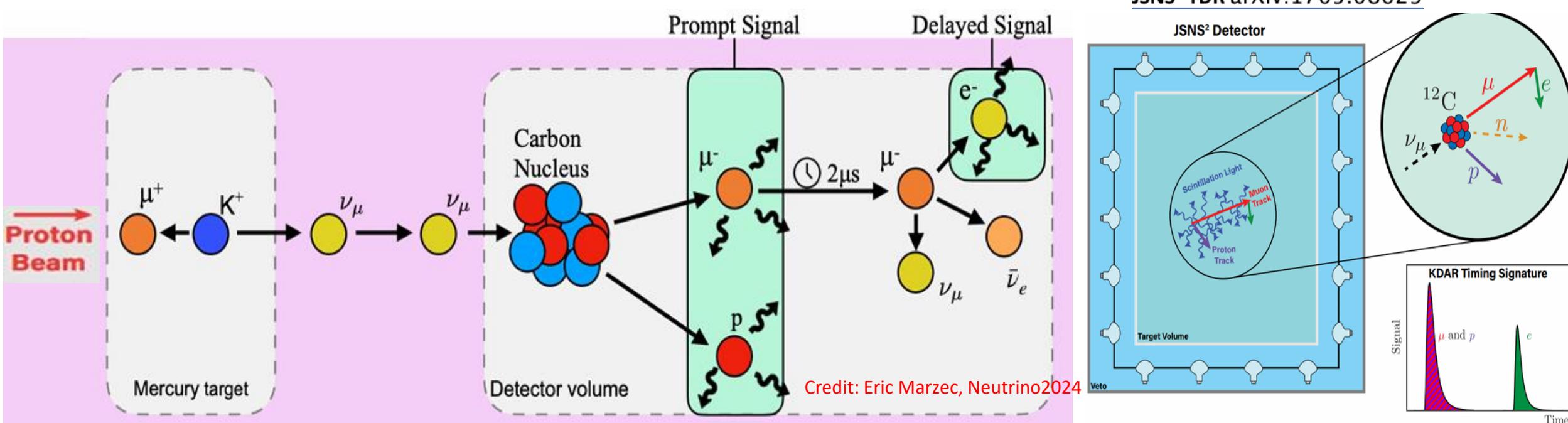
# Neutrinos from kaon-decay-at-rest (KDAR)

- Two-body decay of charged kaons at rest produce mono-energetic beam of muon neutrinos at  $\sim 236$  MeV
- Because of their KNOWN energy KDAR neutrinos are ideal for a cross section measurement
- MiniBOONE (PRL 120, 141802), and JSNS<sup>2</sup> experiments (arXiv:1705.08629) have observed KDAR neutrinos so far
- We use data from JSNS<sup>2</sup> (0.7-0.8 MW beam)
- 730 muon events including 692 signal + 38 background



# JSNS<sup>2</sup> as source for KDAR neutrino signal

- The J-PARC Sterile Neutrino Search at the J-PARC Spallation Neutron Source (JSNS<sup>2</sup>) experiment will produce such types of neutrinos with decay-at-rest processes of pions, muons, and kaons.
- Primary aim of the experiment: Probe **sterile neutrinos with  $\Delta m^2 \sim 1 \text{ eV}^2$  from  $\nu_\mu \rightarrow \nu_e$**  oscillations at a short baseline (24 meters)
- **17 t Gd-loaded liquid scintillator detector**
- **Coincident signal** between initial neutrino interaction and subsequent decay provides **excellent background rejection**



# Neutrino-matter interactions: Standard (SI) and non standard (NSI)

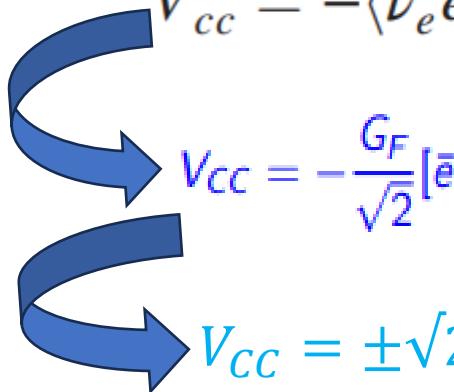
- In the SM there are two ways of interacting neutrinos with matter; **Charged Current** and **Neutral Current**
- The **Charged Current** Lagrangian is given by

$$\mathcal{L}_{\text{cc}}^{\text{eff}} = -\frac{4G_F}{\sqrt{2}} [\bar{\nu}_e(p_3)\gamma_\mu P_L \nu_e(p_2)][\bar{e}(p_1)\gamma^\mu P_L e(p_4)].$$

Flavour-dependent

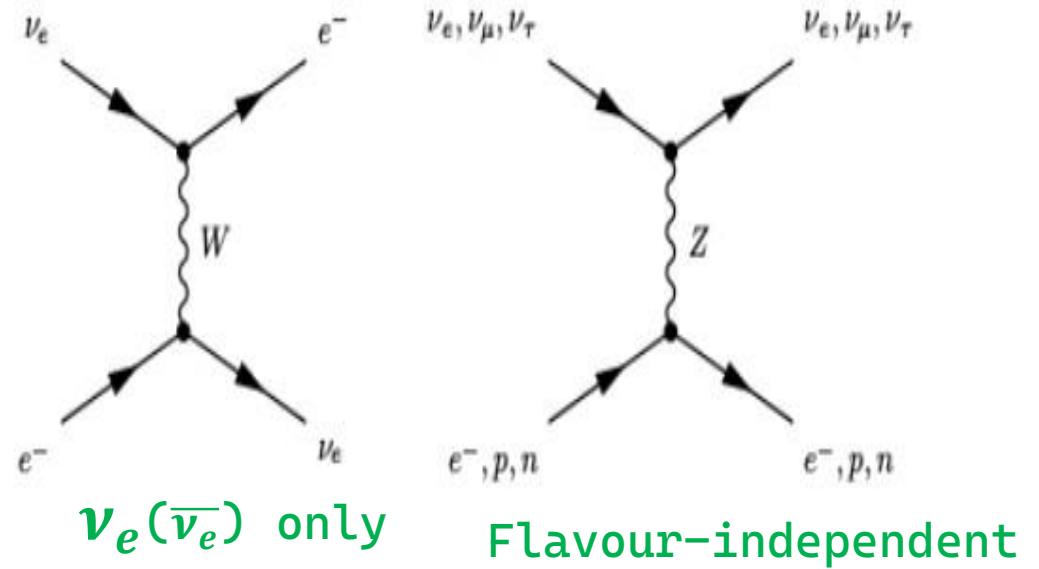


$$V_{cc} = -\langle \nu_e e(p_e, s_e) | \mathcal{L}_{\text{eff}}^{\text{cc}} | \nu_e e(p_e, s_e) \rangle$$



$$V_{CC} = -\frac{G_F}{\sqrt{2}} [\bar{e}\gamma^\mu(1-\gamma_5)\nu_e][\bar{\nu}_e\gamma_\mu(1-\gamma_5)e]$$

$$V_{CC} = \pm\sqrt{2}G_F N_e$$



$\nu_e(\bar{\nu}_e)$  only

Flavour-independent

Similar term for  $V_{NC}$  but due to flavour universal it does not affect neutrino oscillation

Affect neutrino oscillation significantly

# Neutrino-matter interactions: Standard (SI) and non standard (NSI)

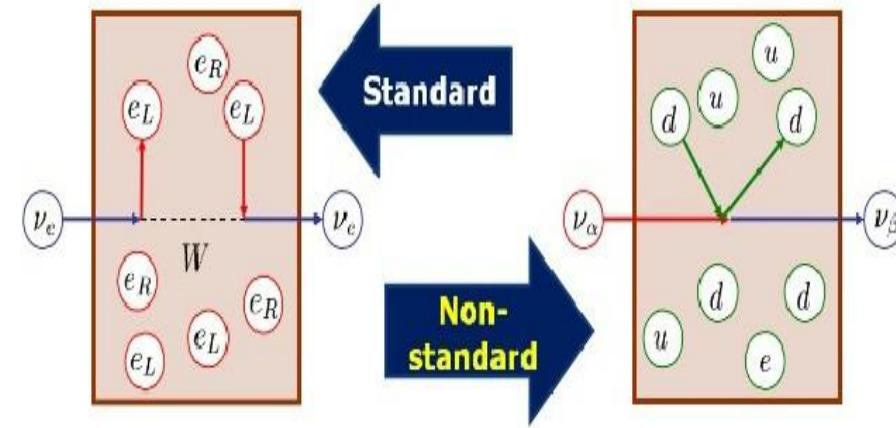
- Non-Standard neutrino interactions are the new interactions and couplings between neutrinos and matter fermions beyond those in the SM. It could be responsible for sub-leading effects in neutrino oscillation.

## ► NSI Lagrangian:

10.1103/PhysRevD.78.053007

$$\mathcal{L}_{NSI}^{NC} = -2\sqrt{2}G_F \sum_{f,P,\alpha,\beta} \epsilon_{\alpha\beta}^{f,P} (\bar{\nu}_\alpha \gamma^\mu P_L \nu_\beta)(\bar{f} \gamma_\mu P_f),$$

- Modify the neutrino coherent-forward scattering with matter over long-baselines aka “**Propagation NSI**”  $\propto E_\nu, \rho$



Credits : T. Ohlsson, MPIK'09

- NSIs which affect neutrino production or detection involve **Charged Current** processes

$$\mathcal{L}_{NSI}^{CC} = -2\sqrt{2}G_F \sum_{f,P,\alpha,\beta} \epsilon_{\alpha\beta}^{f,P} (\bar{\nu}_\alpha \gamma^\mu P_L l_\beta)(\bar{f} \gamma_\mu P_f'),$$

For most experiments, neutrinos are produced from pion decay and detected through their interactions with nucleons, i.e. they are sensitive to the source/detector NSI parameters  $\epsilon_{\alpha\beta}^{ud}$

In this work we use neutrinos from kaon decay to probe a different family of NSI parameters:  $\epsilon_{\alpha\beta}^{us}$

# Neutrino oscillations with source NSI

- In the SM, interactions of charged leptons with neutrinos are flavour-diagonal, i.e.  $|\nu_\alpha^S\rangle = |\nu_\alpha\rangle$ .
- However, the inclusion of CC-NSI can alter this and the neutrino produced in association with the charged lepton  $l_\alpha$  can also have an admixture of other flavour  $\nu_\beta$ , i.e.  $|\nu_\alpha^S\rangle = \sum_\beta (\delta_{\alpha\beta} + \epsilon_{\alpha\beta}^{ff'}) |\nu_\beta\rangle$
- Source NSIs induce non-unitarity: Non-trivial normalization of the states
- For KDAR neutrinos  $|\nu_\mu^S\rangle$  will be modified as  $|\nu_\mu^S\rangle = (\mathbb{I} + \epsilon^{ff'})_{\mu\alpha} |\nu_\alpha\rangle = (\mathbb{I} + \epsilon^{ff'})_{\mu\alpha} U_{\alpha i} |\nu_i\rangle$
- The source NSI parameters relevant for this work
- For very short baselines ( $L/E \ll 1$ ), only (standard) survival amplitudes contribute, giving rise to ‘zero-distance flavour conversion’

$$\epsilon_{\alpha\beta}^{us} = \begin{bmatrix} \epsilon_{ee}^S & \epsilon_{e\mu}^S & \epsilon_{e\tau}^S \\ \epsilon_{\mu e}^S & \epsilon_{\mu\mu}^S & \epsilon_{\mu\tau}^S \\ \epsilon_{\tau e}^S & \epsilon_{\tau\mu}^S & \epsilon_{\tau\tau}^S \end{bmatrix}$$

Oscillation Probabilities



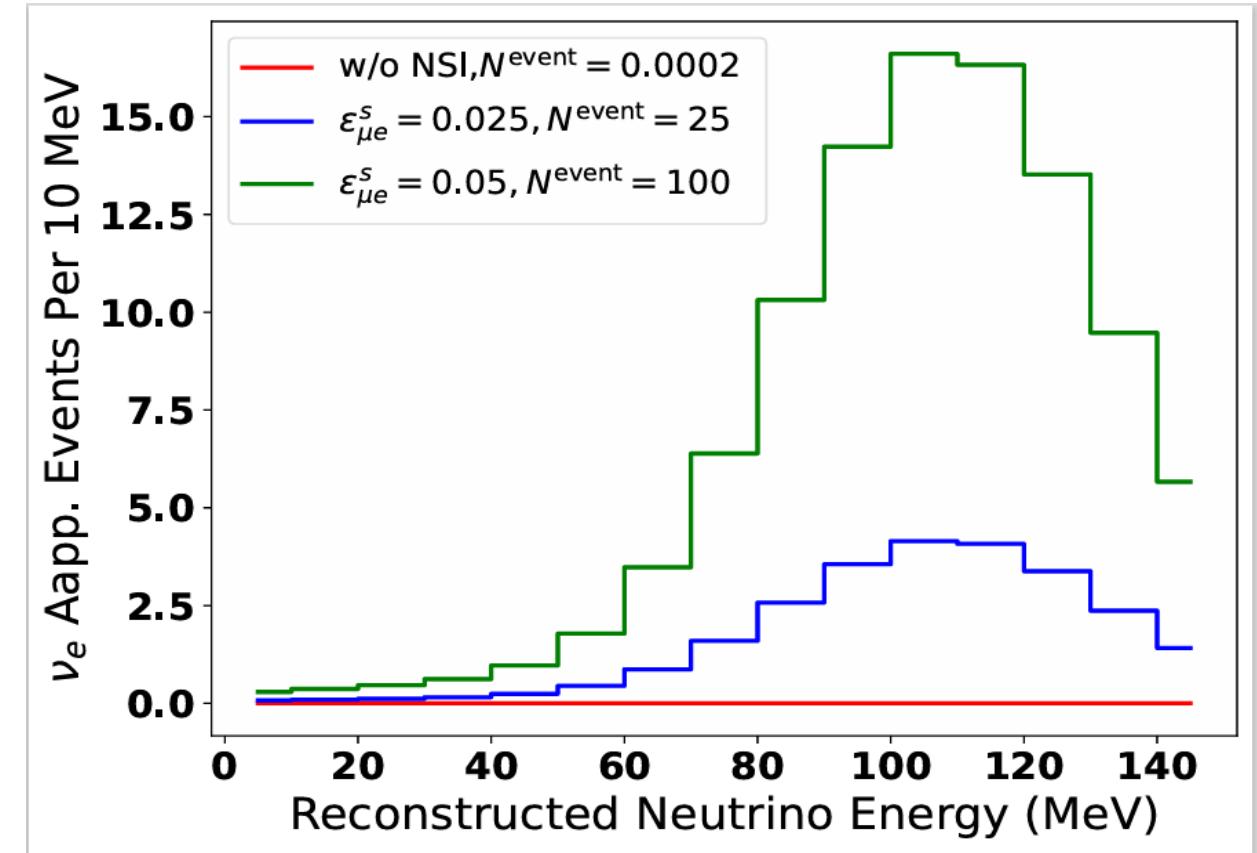
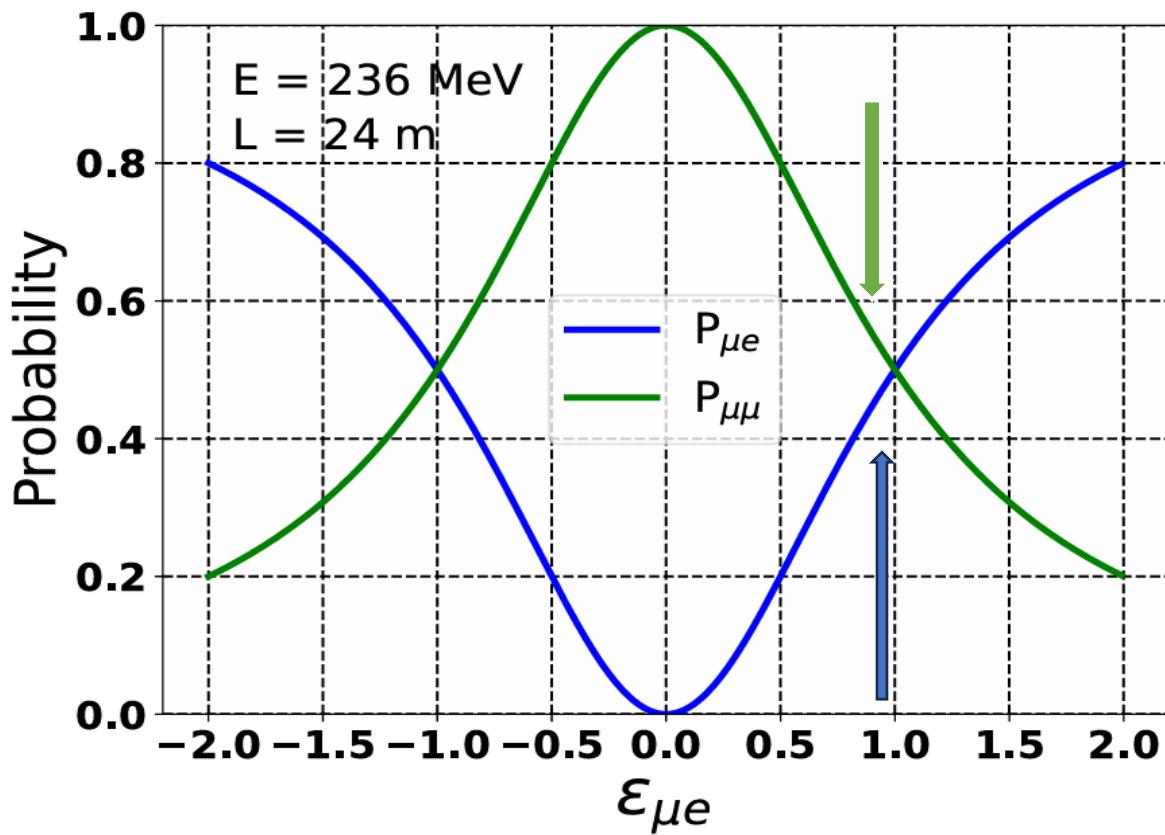
$$P_{\mu\alpha} = |\sum_\beta (\mathbb{I} + \epsilon^{us})_{\mu\beta} \mathcal{A}_{\beta\alpha}^{SM}|^2 \text{ (up to a normalization factor)}$$

$$P_{\mu e} = |\epsilon_{\mu e}^{us}|^2 / (|\epsilon_{\mu e}^{us}|^2 + |1 + \epsilon_{\mu\mu}^{us}|^2)$$

$$P_{\mu\mu} = |1 + \epsilon_{\mu\mu}^{us}|^2 / (|\epsilon_{\mu e}^{us}|^2 + |1 + \epsilon_{\mu\mu}^{us}|^2)$$

# Results: oscillation probability and event spectrum

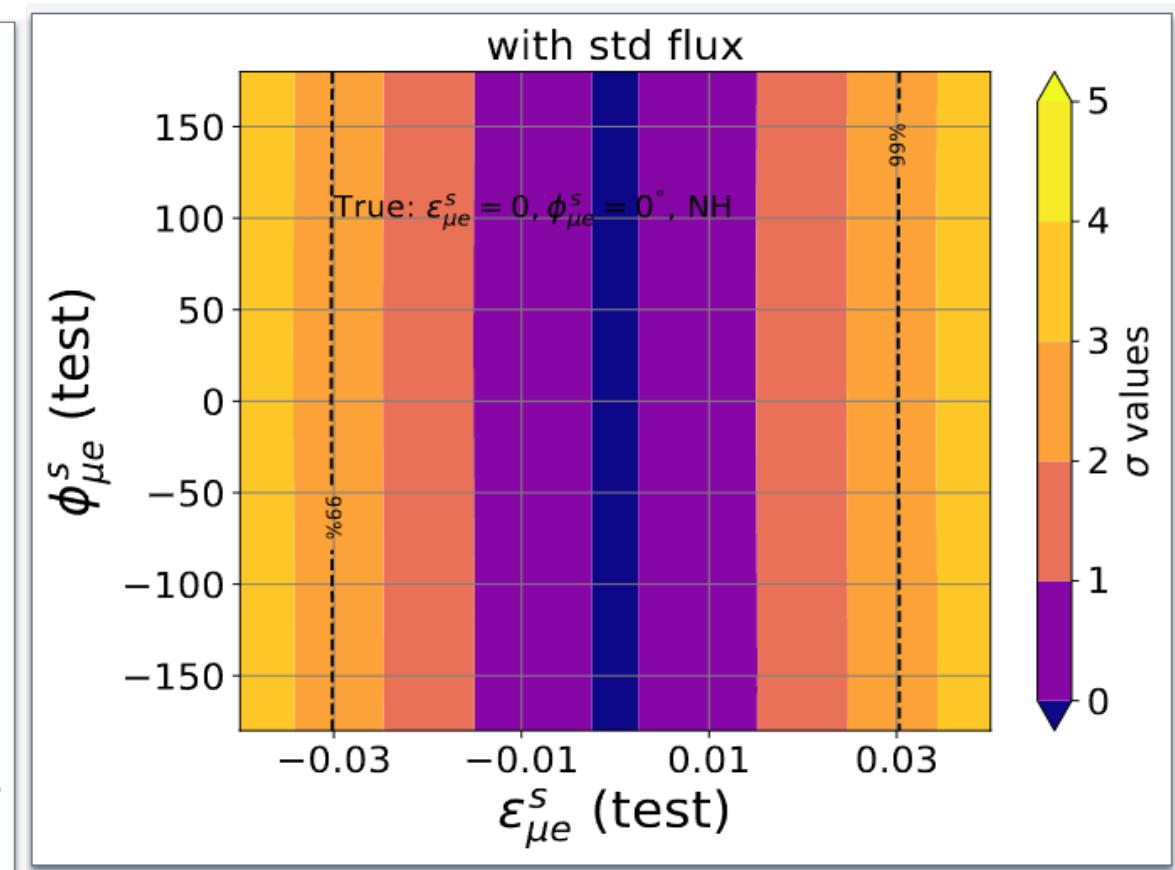
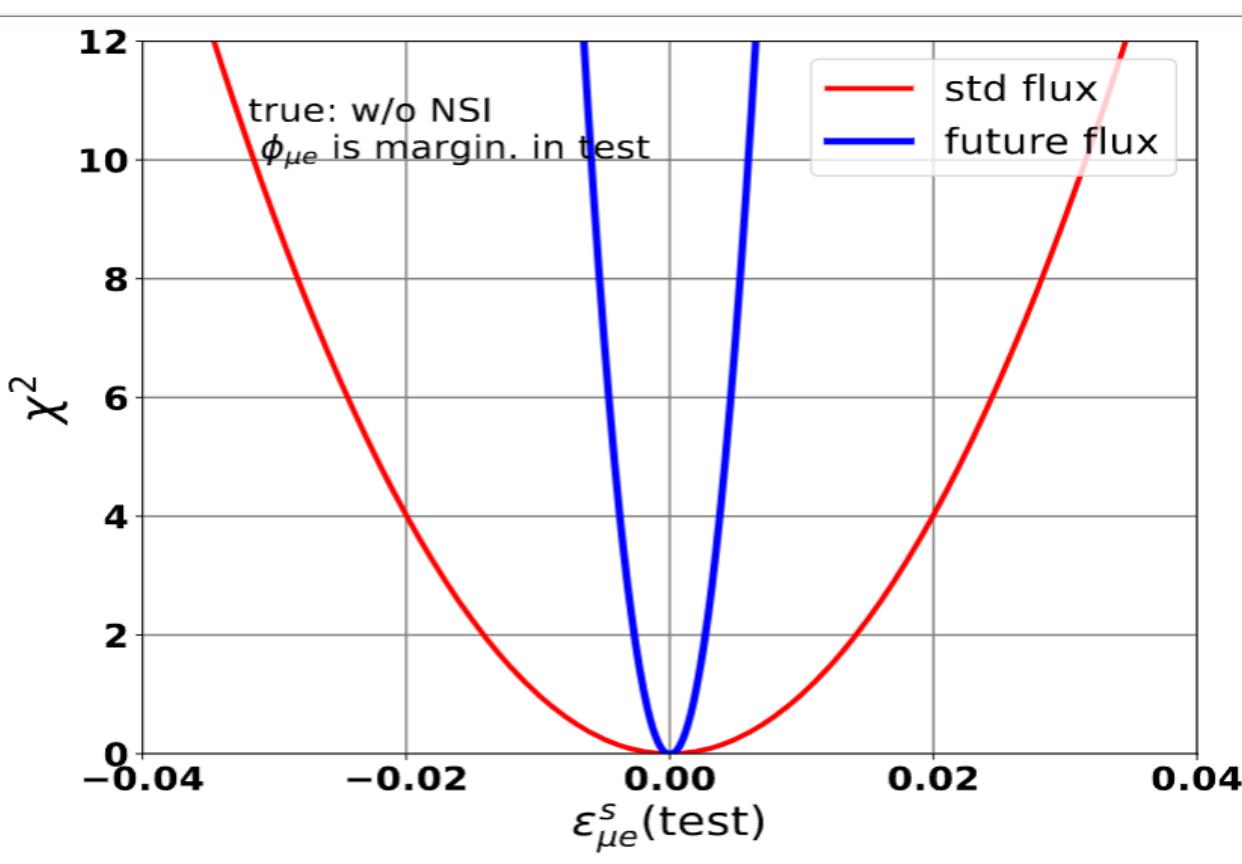
- We use GLoBES to compute the event spectrum and sensitivity of JSNS<sup>2</sup> by implementing our own NSI probability engine.
- Only one NSI parameter is considered at a time.



- The effect of only  $\epsilon_{\mu \mu}^{us}$  is wiped out due to the probability normalization by  $\nu_\mu$  disappearance events

# Results: source NSI constraint from KDAR

- We present the result for both Current data (calibrated to 730  $\nu_\mu$  events) and future data (calibrated to 40,000  $\nu_\mu$  events)



The bounds by JSNS<sup>2</sup> on NSI parameter  $|\epsilon_{\mu e}^{us}| < 0.03$  (0.005) at 99% C.L. with current (future) statistics

# Search for Sterile neutrinos from KDAR

- Over the past few decades, several anomalous results have been observed in experiments involving the production and detection of neutrinos over short baselines (less than 1 km). **To explain these anomalies, sterile neutrino oscillations with a mass of around 1 eV have been proposed as a key solution.**
- The short-baseline oscillation behaviour of KDAR neutrinos will be altered in the presence of eV-scale sterile neutrino: new mixing angles, phases, mass-squared difference

In “3+1 model”

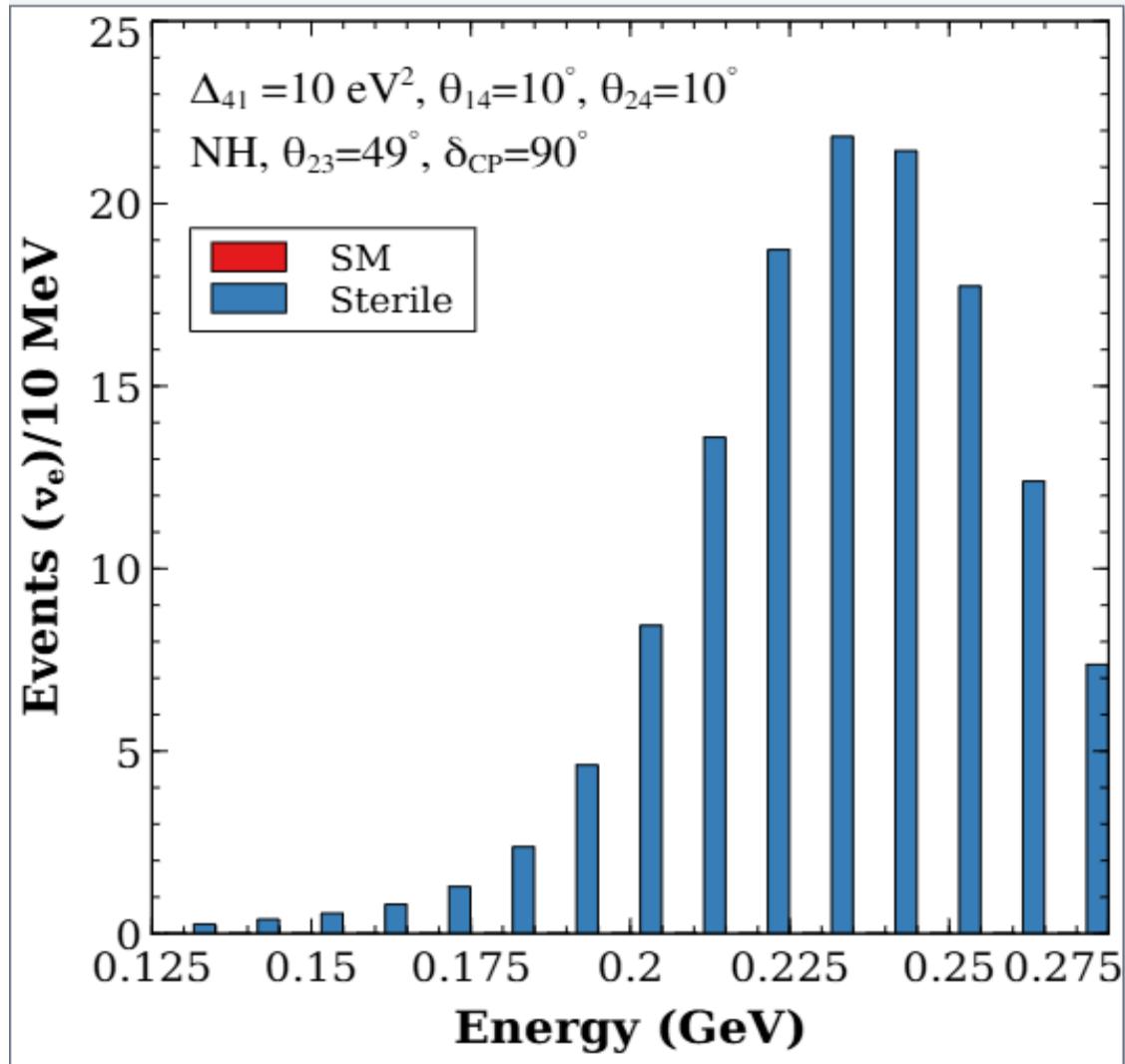
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{pmatrix}$$

In the short-baseline limit  $\left(\frac{\Delta m_{21}^2 L}{E} \ll 1, \frac{\Delta m_{31}^2 L}{E} \ll 1\right)$ , where standard oscillations are suppressed, the  $\nu_\mu \rightarrow \nu_e$  probability is

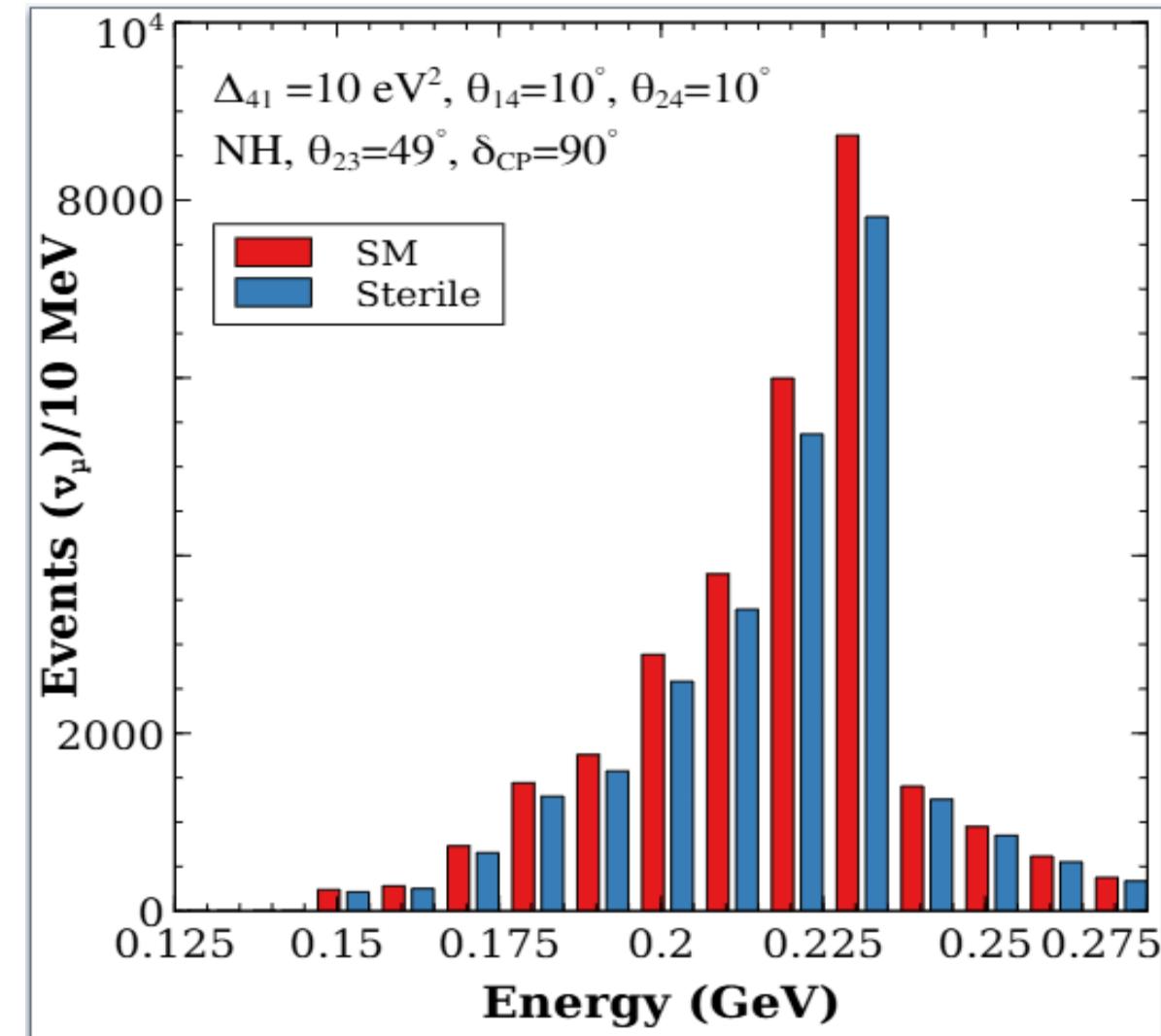
$$P_{\mu e}^{\text{sbl}} \simeq 4|U_{e4}|^2|U_{\mu 4}|^2 \sin^2 \frac{\Delta m_{41}^2 L}{4E}$$

We compare the **KDAR neutrino** spectra with standard oscillations versus with sterile neutrinos to put bounds on the sterile parameter space

# KDAR event spectra at JSNS<sup>2</sup> with sterile neutrino

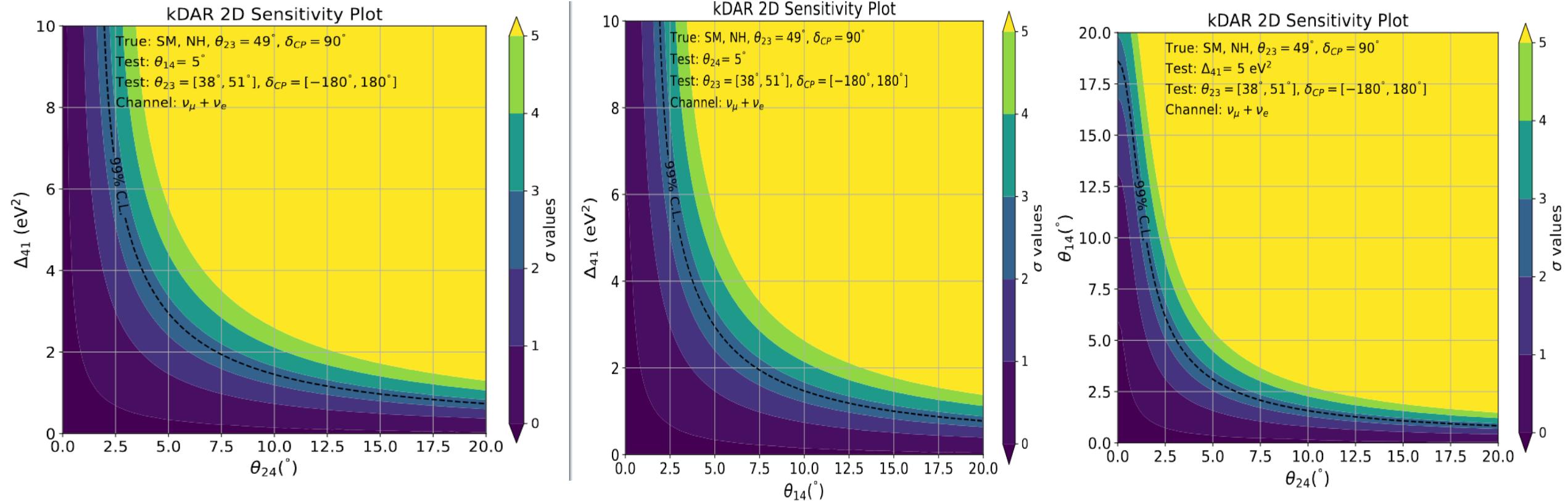


Appearance



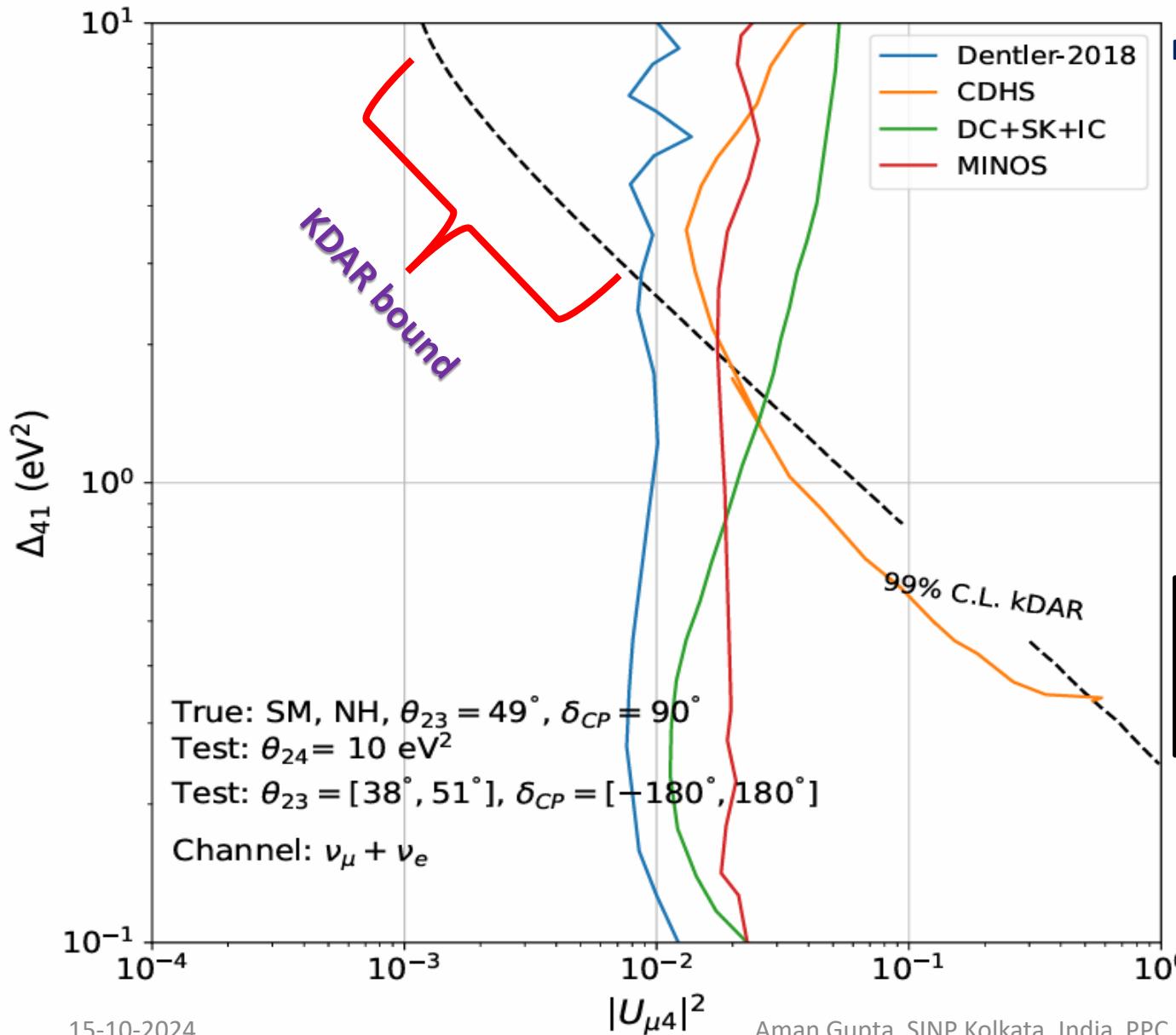
Disappearance

# Results: constraints on sterile neutrino parameter space



Sensitivity of JSNS<sup>2</sup> (with KADAR only data) in constraining  $\theta_{14}$ ,  $\theta_{24}$  and  $\Delta m_{41}^2$  from appearance and disappearance channel data

# Results: constraints on sterile neutrino parameter space



M. Dentler et al., JHEP 08 (2018) 010

Constraints on the active-sterile mixing matrix element  $|U_{\mu 4}^2|$  from future KDAR data (black, dotted), overlaid on bounds from CDHS, MINOS, and atmospheric neutrino data

The best bounds (as small as  $|U_{\mu 4}^2| \sim 10^{-3}$ ) are obtained for  $\Delta m_{41}^2 > 2$  eV $^2$  which is consistent with the baseline and energy of this experiment

# Concluding Remarks

- The possibility of new physics searches such as **source NSI** and **sterile neutrino** have been explored exploiting **KDAR neutrino facility at JSNS<sup>2</sup>** experiment
- Unlike propagation NSI, source (or production) NSI is independent of matter potential and neutrino energy and can give rise to '**zero-distance flavour conversion**'
- **Constraints on the non-standard coupling, for the first time in *us* sector (strange quark) have been obtained:**

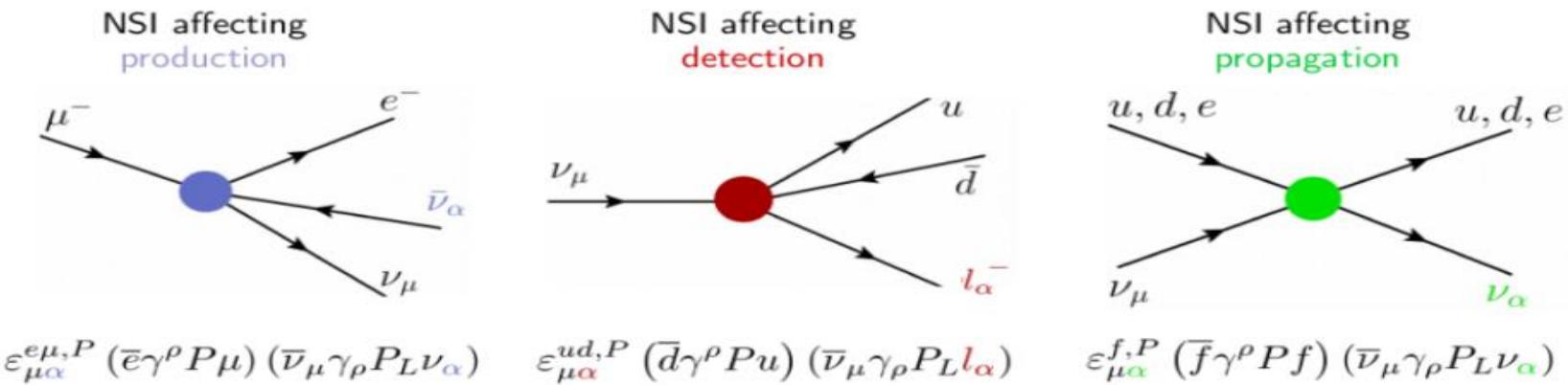
$$|\varepsilon_{\mu e}^{us}| < 0.03 \text{ (0.005)} \text{ at 99% C.L. with current (future) statistics}$$

- We also find that with the **JSNS<sup>2</sup>** experiment and future KDAR only data  
***Active sterile mixing can be probed down to  $|U_{\mu 4}^2| \sim 10^{-3}$  for  $\Delta m_{41}^2 \sim 10 \text{ eV}^2$***
- Monoenergetic 236 MeV neutrinos from kaon decay-at-rest, can also be used to study neutrino-nucleus cross-section

Thank you ☺



# NSI at Production and Detection Level



Credits: P. Coloma, Fermilab'17

Neutrino states at sources and detectors:

$$|\nu_\alpha^s\rangle = |\nu_\alpha\rangle + \sum_{\beta=e,\mu,\tau} \varepsilon_{\alpha\beta}^s |\nu_\beta\rangle = (1 + \varepsilon^s) U |\nu_m\rangle$$

$$\langle\nu_\beta^d| = \langle\nu_\beta| + \sum_{\alpha=e,\mu,\tau} \varepsilon_{\alpha\beta}^d \langle\nu_\alpha| = \langle\nu_m| U^\dagger [1 + (\varepsilon^d)^\dagger]$$

Superpositions of pure orthonormal flavor eigenstates

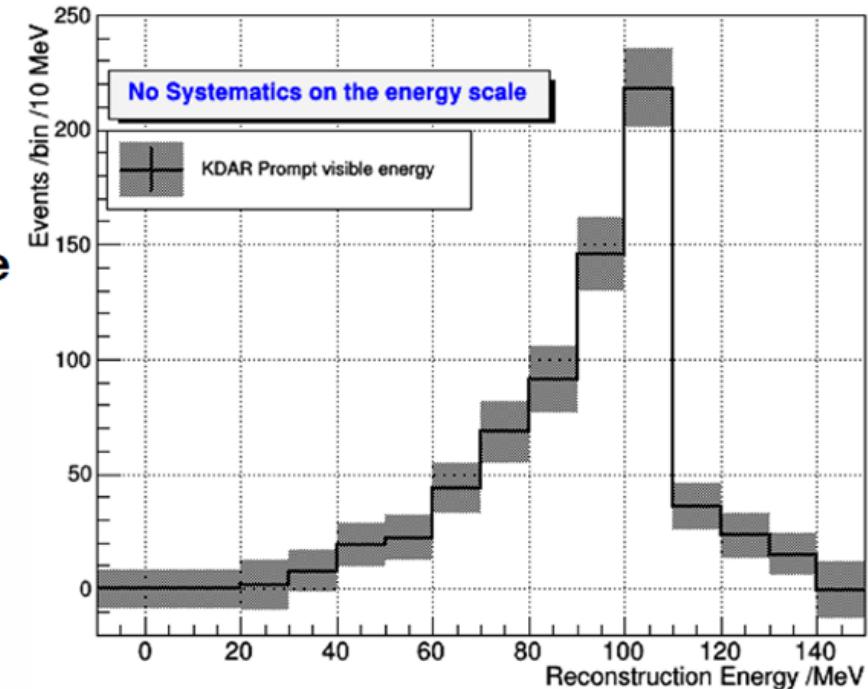
Grossman (1995); Gonzalez-Garcia *et al.* (2001); Bilenky, Giunti (1993); Meloni *et al.* (2010)

# First clear KDAR signal

## (Toward first precise KDAR measurement)

- KDAR peak is clearly seen
- High purity (95%) KDAR signal
  - Background: 5.2 %
- Note that **the systematics on the energy scale are not included yet.**

BKG ID	Correlated/ Accidental	BKG (# of events)	
1	Correlated	36.6 +- 34.8	5.0 +- 5.1%
2	Accidental	1.5 +- 0.1	0.2 +- 0.01%
KDAR Candidates : 730 events		38.1 +- 38.4	5.2 +- 5.3%



# KDAR Backgrounds

- Dominant background source is pion decay-in-flight (DIF) neutrinos
- DIF background spectral shape estimated with MC
- Both NuWro & GiBUU event generators are used for DIF background simulation
- Normalization estimated using the kinematically disallowed ( $>150\text{MeV}$ ) portion of the KDAR spectrum

