

Probing flavor violation and baryogenesis via primordial gravitational waves

Based on JHEP 07 (2024) 228

With Prof. Seyda Ipek, Dr. Anish Ghoshal

Oct 16, 2024

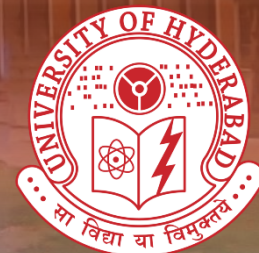


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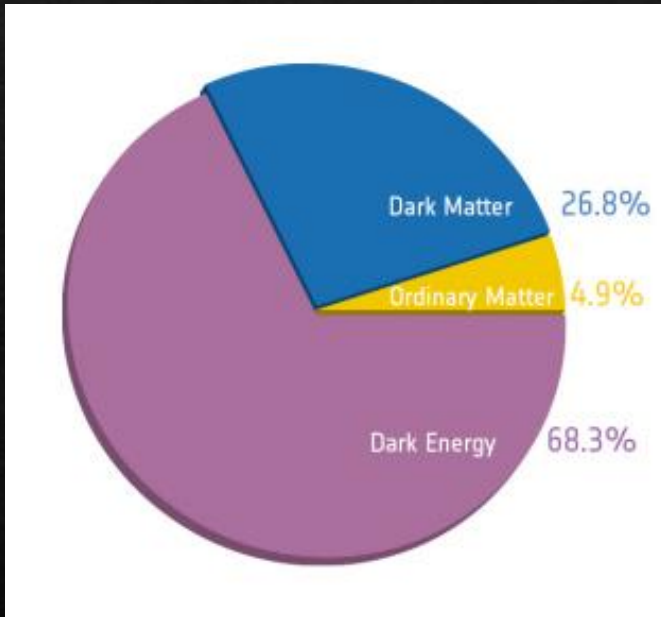
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Limitations of the Standard Model

Dark matter

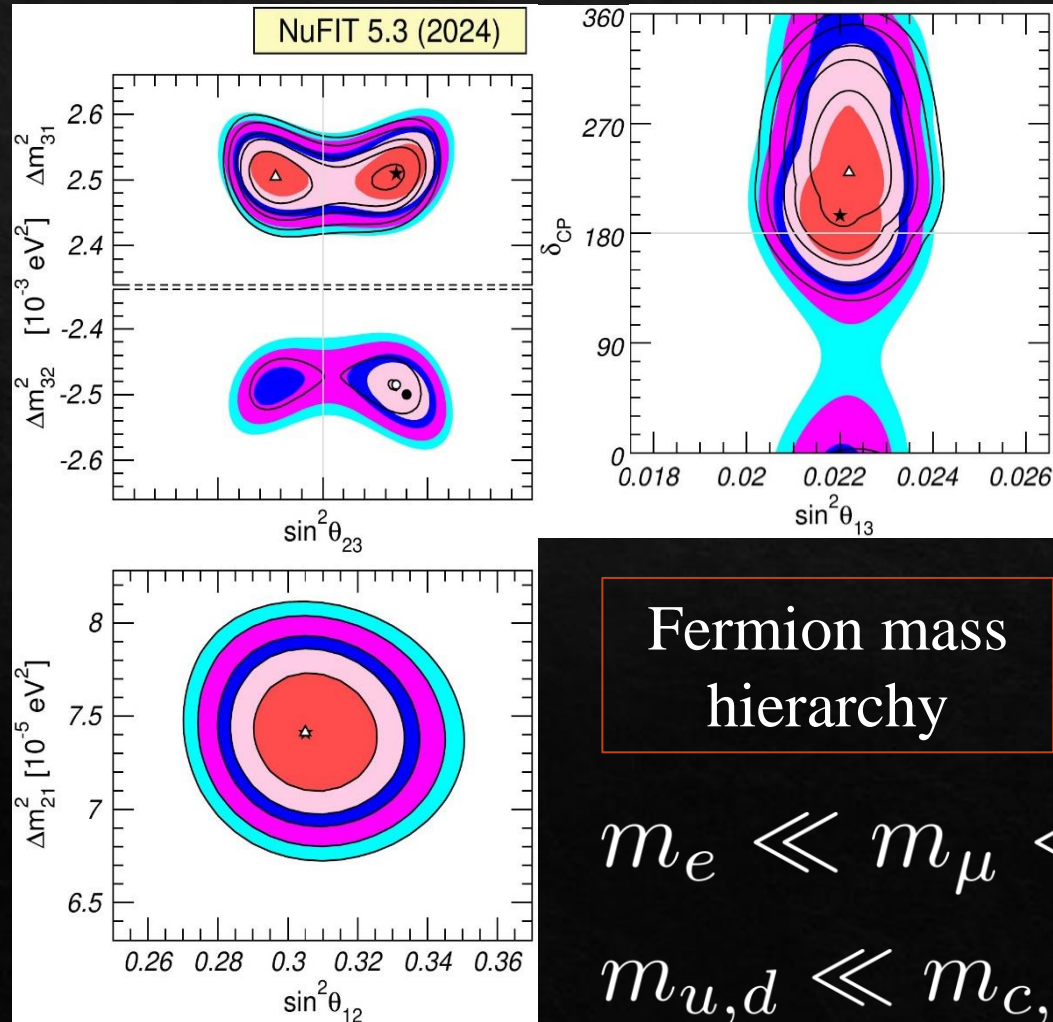


Credit: [ESA Science & Technology - Planck's new cosmic recipe](#)

Matter-antimatter asymmetry

$$\eta = \frac{n_B - n_{\bar{B}}}{n_\gamma} \sim 10^{-10}$$

Neutrino mass

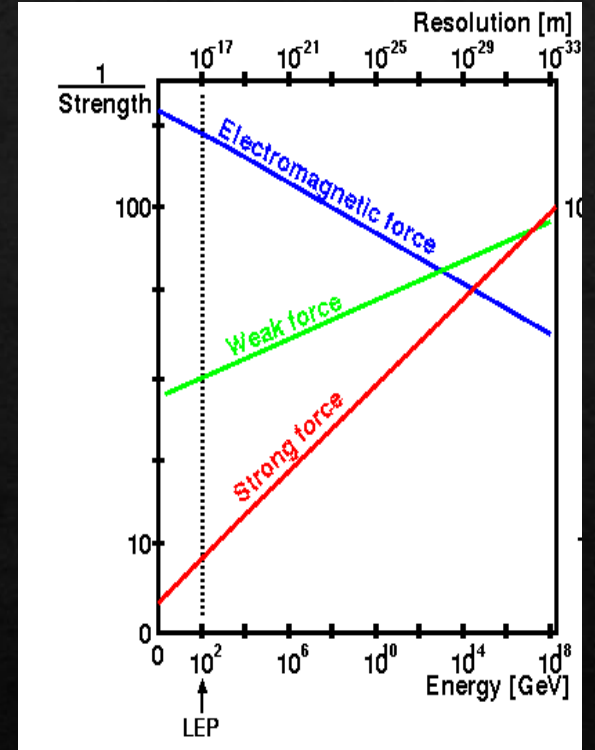


Fermion mass hierarchy

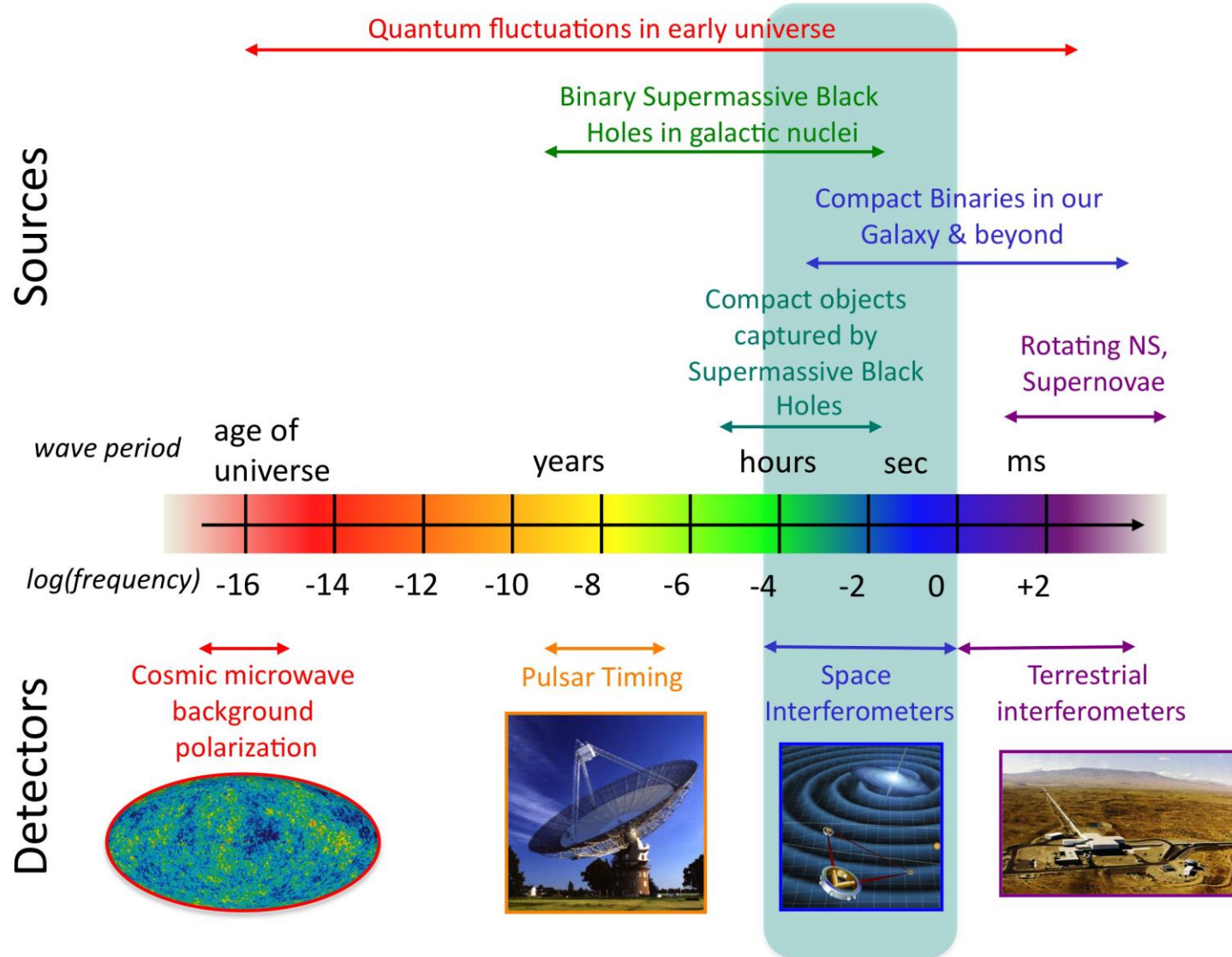
$$m_e \ll m_\mu \ll m_\tau$$

$$m_{u,d} \ll m_{c,s} \ll m_{t,b}$$

Gauge unification

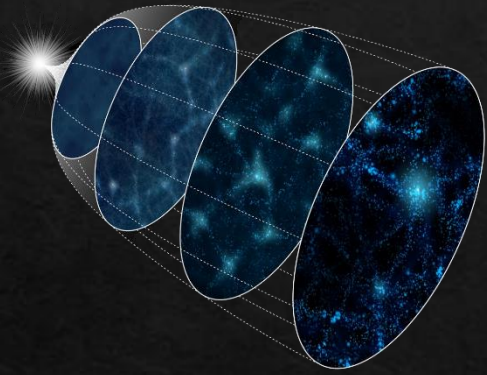


The Gravitational Wave Spectrum

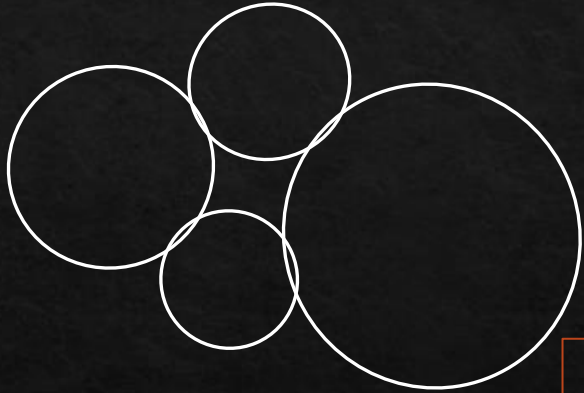


Sources of early gravitational waves

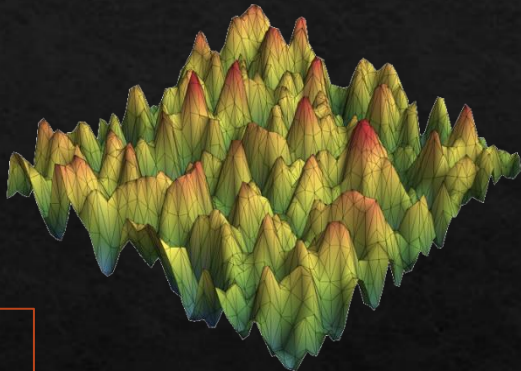
Inflation



First-order phase transition



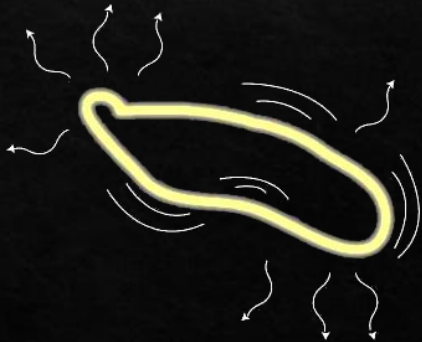
Scalar-induced gravitational waves



Domain walls



Cosmic strings

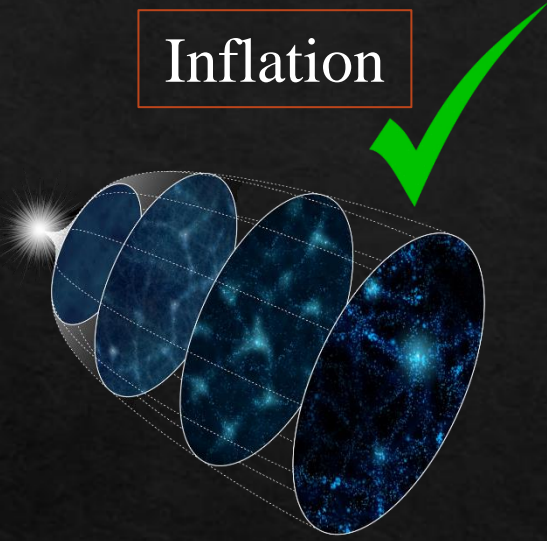


Credit: Institute of Statistical Mathematics (ISM).

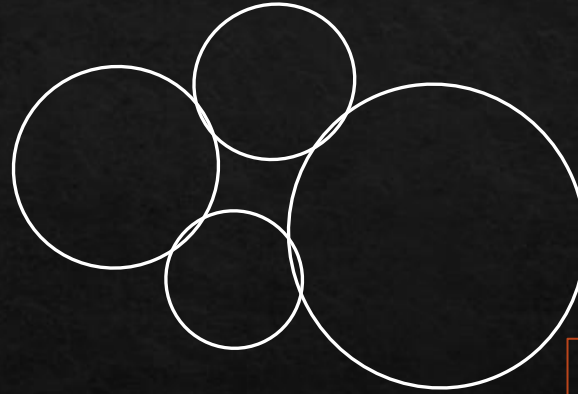
Credit: space.com

Sources of early gravitational waves

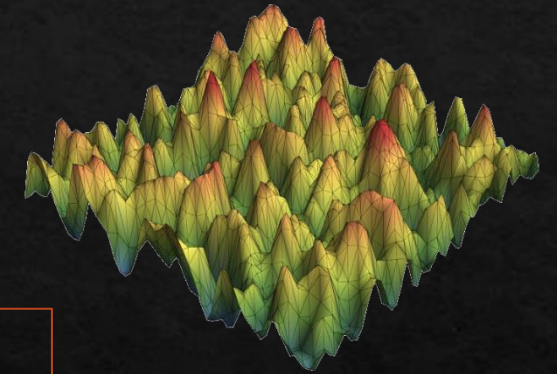
Inflation



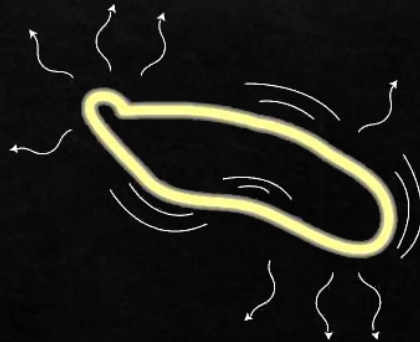
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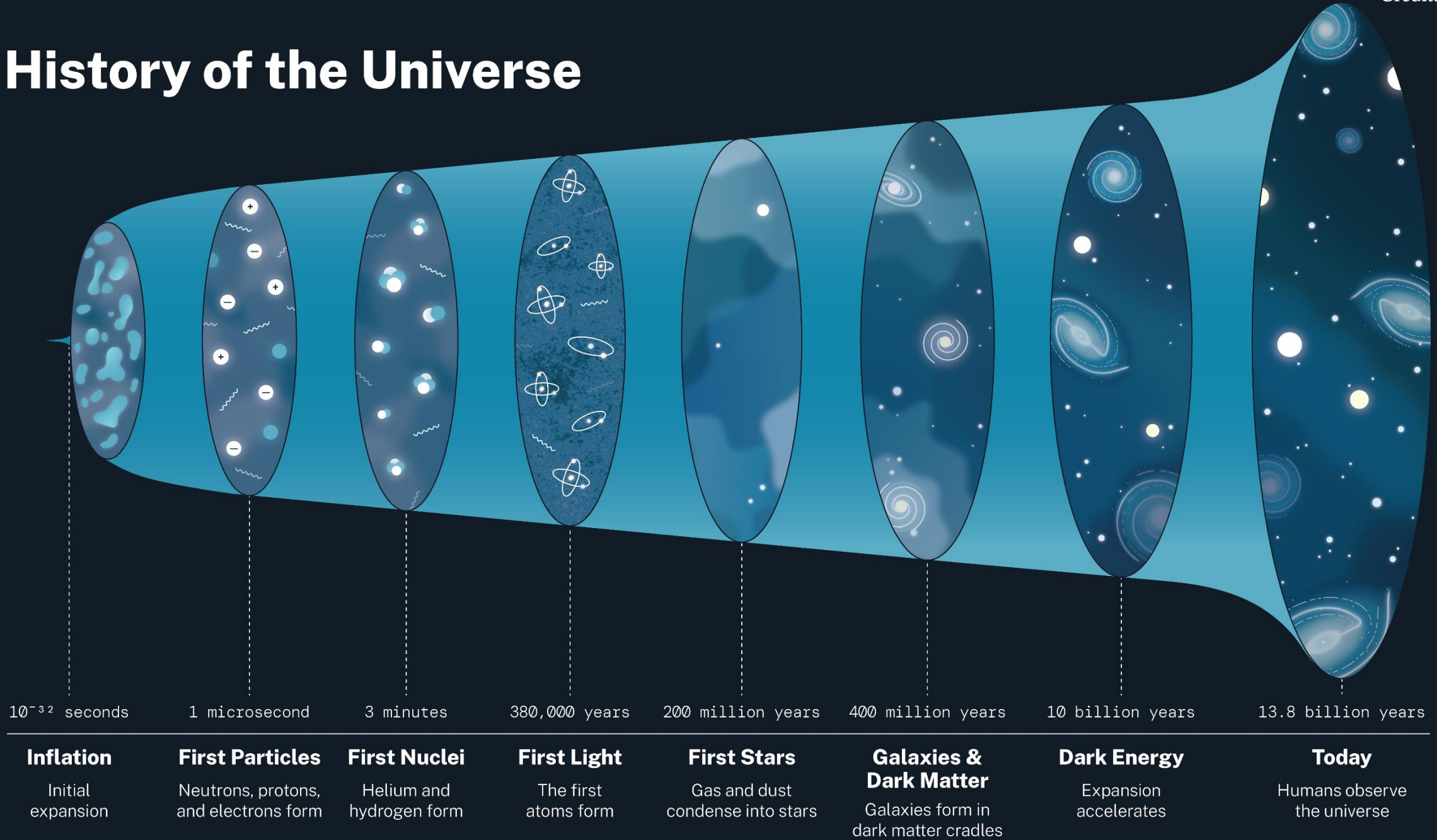
Domain walls



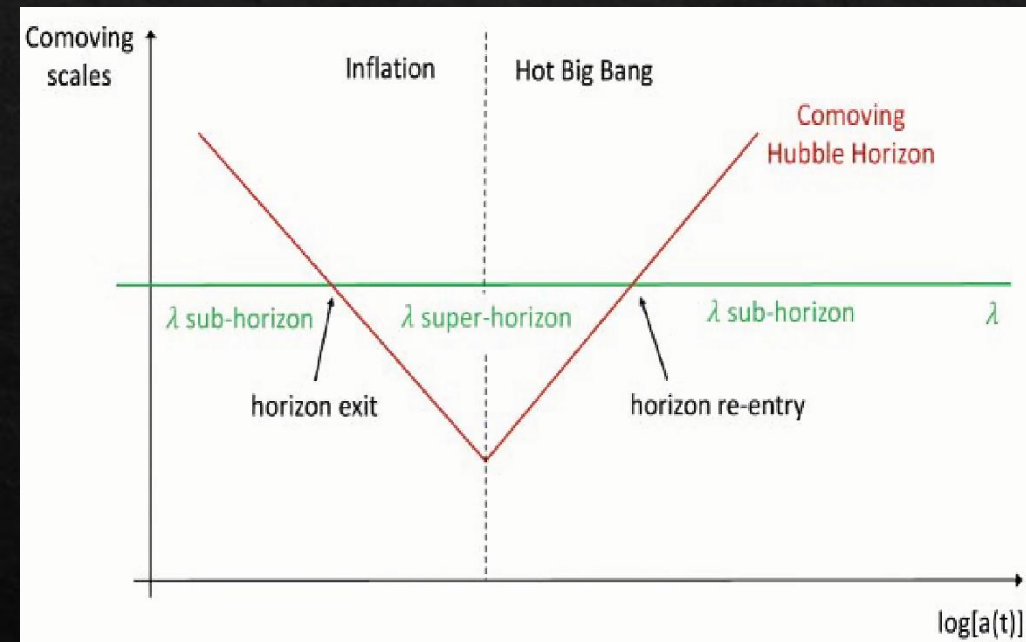


Thermal History of the Universe with Inflationary Gravitational Waves

History of the Universe

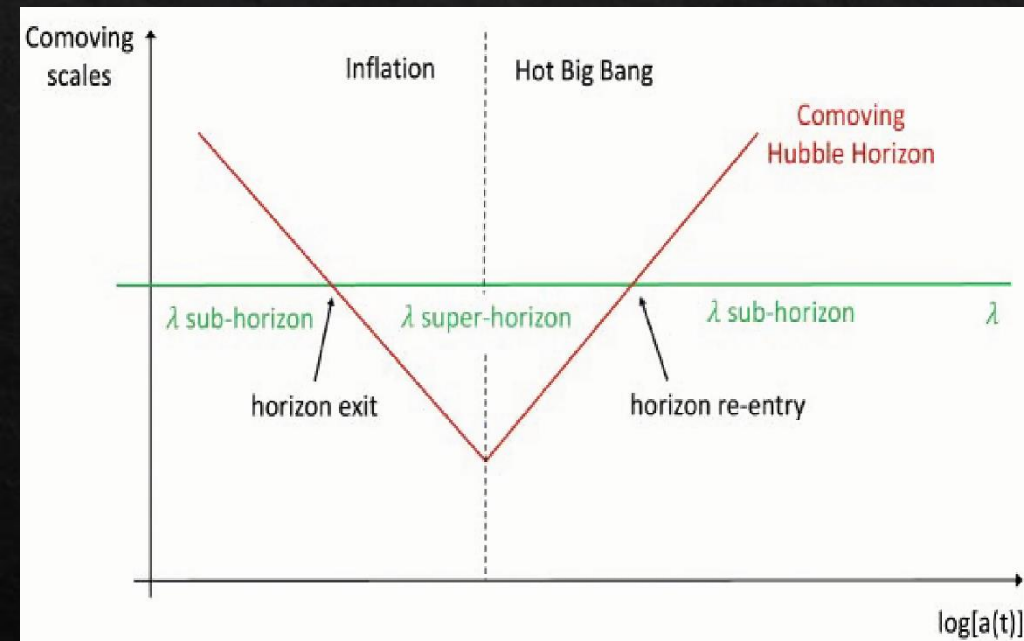


Horizon re-entry of different scales after inflation

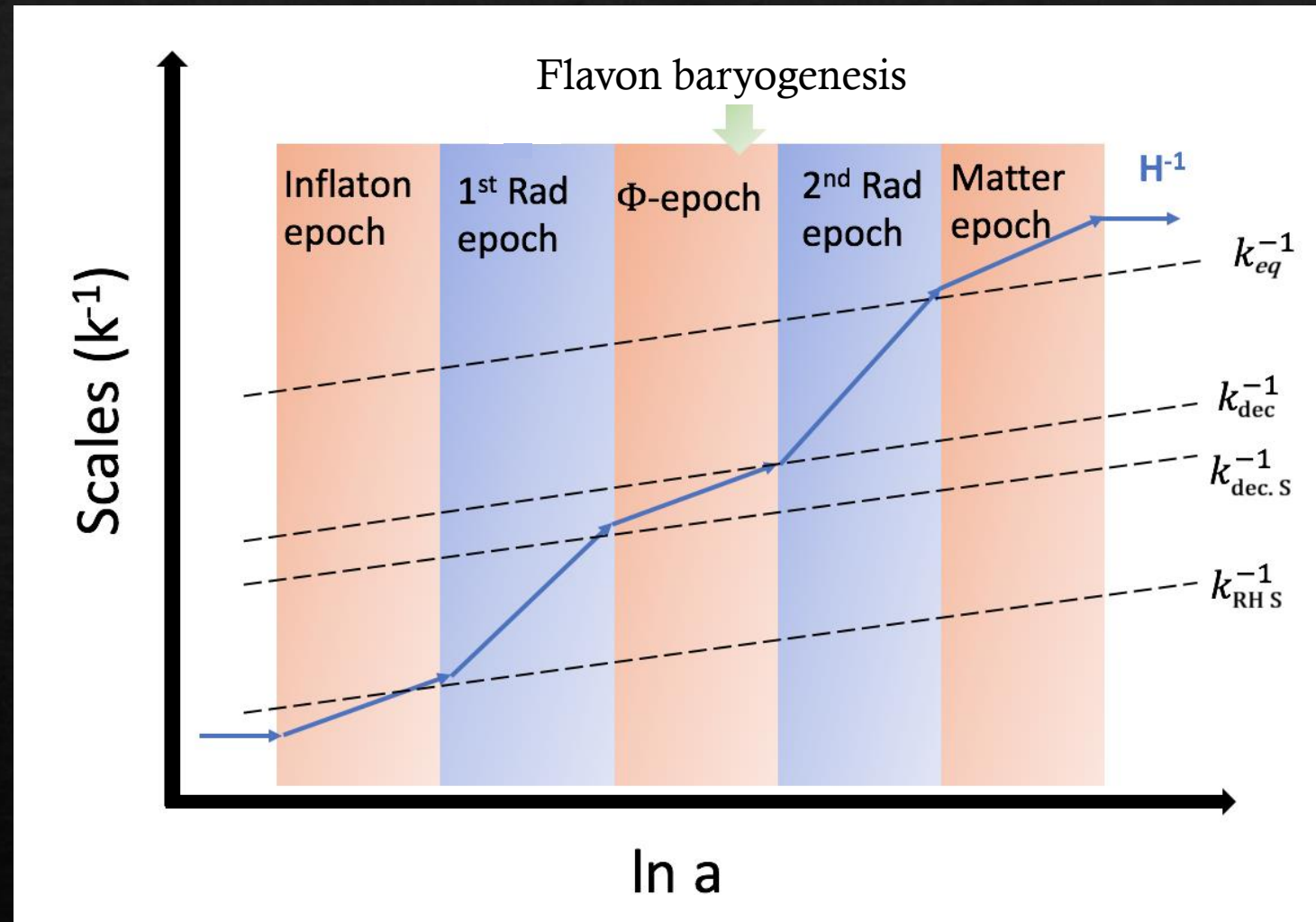


Credit: CERN.

Horizon re-entry of different scales after inflation



Credit: CERN.



S. Dutta et. al., JHEP (2022).

GW spectra from inflation

$$\Omega_{GW}(k) = \frac{1}{12} \left(\frac{k}{a_0 H_0} \right)^2 T_T^2(k) P_T^{\text{prim.}}(k)$$

$$P_T^{\text{prim.}}(k) = A_T(k_*) \left(\frac{k}{k_*} \right)^{n_T}$$

$$T_{\text{in}} = 5.8 \times 10^6 \text{ GeV} \left(\frac{106.75}{g_*(T_{\text{in}})} \right)^{1/6} \left(\frac{k}{10^4 \text{ Mpc}^{-1}} \right)$$

$$A_T(k_*) = 2.0989 \times 10^{-9} r$$

$$T_T^2(k) = \Omega_m^2 \left(\frac{g_*(T_{\text{in}})}{g_*^0} \right) \left(\frac{g_{*S}^0}{g_{*S}(T_{\text{in}})} \right)^{4/3} \left(\frac{3j_1(z_k)}{z_k} \right)^2 F(k)$$

Kuroyanagi et al JCAP 2014

Berbig et al JHEP 2023

$$F(k)_{\text{standard}} = T_1^2 \left(\frac{k}{k_{\text{eq.}}} \right) T_2^2 \left(\frac{k}{k_{\text{RH}}} \right)$$

$$F(k)_{\text{IMD}} = T_1^2 \left(\frac{k}{k_{\text{eq.}}} \right) T_2^2 \left(\frac{k}{k_{\text{dec.}}} \right) T_3^2 \left(\frac{k}{k_{\text{dec. S}}} \right) T_2^2 \left(\frac{k}{k_{\text{RH S}}} \right)$$

GW spectra from inflation

Kuroyanagi et al JCAP 2014

Berbig et al JHEP 2023

$$k_{\text{eq.}} = 7.1 \times 10^{-2} \text{ Mpc}^{-1} \cdot \Omega_m h^2$$

$$k_{\text{dec.}} = 1.7 \times 10^{14} \text{ Mpc}^{-1} \left(\frac{g_{*S}(T_{\text{dec.}})}{g_{*S}^0} \right)^{1/6} \left(\frac{T_{\text{dec.}}}{10^7 \text{ GeV}} \right)$$

$$k_{\text{RH}} = 1.7 \times 10^{14} \text{ Mpc}^{-1} \left(\frac{g_{*S}(T_{\text{RH}})}{g_{*S}^0} \right)^{1/6} \left(\frac{T_{\text{RH}}}{10^7 \text{ GeV}} \right)$$

$$k_{\text{dec. S}} = k_{\text{dec.}} \Delta^{2/3}$$

$$k_{\text{RH S}} = k_{\text{RH}} \Delta^{-1/3}$$

$$T_1^2(x) = 1 + 1.57x + 3.42x^2$$

$$T_2^2(x) = \left(1 - 0.22x^{3/2} + 0.65x^2 \right)^{-1}$$

$$T_3^2(x) = 1 + 0.59x + 0.65x^2$$

$$\Delta = \frac{s a^3|_{\text{after}}}{s a^3|_{\text{before}}}$$

Dilution factor from entropy injection



Flavon baryogenesis

Froggatt-Neilsen Mechanism

Fermion mass
hierarchy

Froggatt-Neilsen Mechanism

$$U(1)_{\text{FN}}$$

Fermion mass
hierarchy

Froggatt-Neilsen Mechanism

| | | $U(1)_{\text{FN}}$ |
|---------|----------|--------------------|
| Flavon | S | -1 |
| Fermion | ψ_i | Q_i |

Fermion mass
hierarchy

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Fermion mass
hierarchy

Effective operator: $y_{ij} \psi_i \psi_j H \left(\frac{v_S + S}{\Lambda_{\text{FV}}} \right)^{n_{ij}}$

$n_{ij} = Q_i + Q_j$
 $y_{ij} \sim \mathcal{O}(1)$

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$$\epsilon = \frac{v_S}{\Lambda_{\text{FV}}} \quad Y_{ij} = \begin{pmatrix} \epsilon^{n_{11}} & \epsilon^{n_{12}} & \epsilon^{n_{13}} \\ \epsilon^{n_{21}} & \epsilon^{n_{22}} & \epsilon^{n_{23}} \\ \epsilon^{n_{31}} & \epsilon^{n_{32}} & \epsilon^{n_{33}} \end{pmatrix}$$

Flavon baryogenesis

$$\mathcal{L} \supset \left(\frac{v_S + S}{\Lambda_{\text{FV}}} \right)^{n_i} \bar{e}_R^i \phi^* \ell_L^i + \text{h.c.}$$

Chen et al PRD 2019

$$S \rightarrow \bar{\ell}_L + e_R + \phi, \quad S^* \rightarrow \bar{e}_R + \ell_L + \phi^*$$

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$$\Gamma_{LR} \simeq 10^{-2} y_e^2 T \quad y_e \simeq 2.9 \times 10^{-6}$$

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RD: Right-handed electrons come into equilibrium at $T \sim 10^5$ GeV

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Sphalerons act only on the left-handed asymmetry at $T \sim 160$ GeV

Flavon baryogenesis

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Chen et al PRD 2019

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We are agnostic of the exact model that creates the initial flavon asymmetry. We simply assume that far below Λ_{FV} , the flavon potential preserves an approximate $U(1)_S$ symmetry that is broken explicitly by small S-number violating terms responsible for the initial asymmetry,

$$V_S = m^2 |S|^2 + \left(\begin{array}{c} \text{S-number violating terms} \\ \text{suppressed by } \Lambda_{\text{FV}} \end{array} \right)$$

Flavon baryogenesis

Boltzmann equations:

$$\frac{d\rho_S}{dt} + 3H\rho_S = -\Gamma_S\rho_S,$$

$$\Gamma_S \simeq 2.3 \times 10^{-17} \text{ GeV} \left(\frac{m_S}{\text{TeV}}\right)^3 \left(\frac{10^{10} \text{ GeV}}{\Lambda_{\text{FV}}}\right)^2,$$

$$\frac{d\rho_R}{dt} + 4H\rho_R = \Gamma_S\rho_S,$$

$$H^2 = \frac{8\pi}{3M_{\text{Pl}}^2}(\rho_S + \rho_R),$$

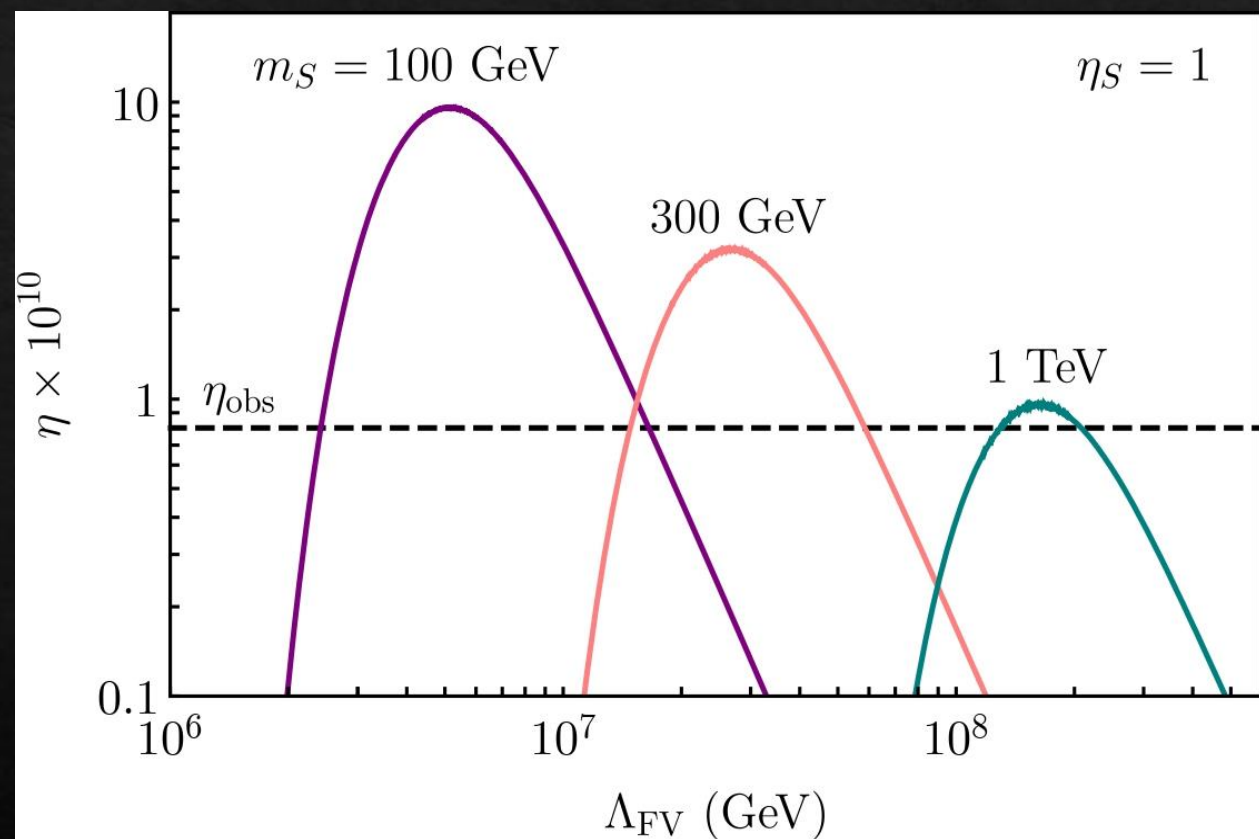
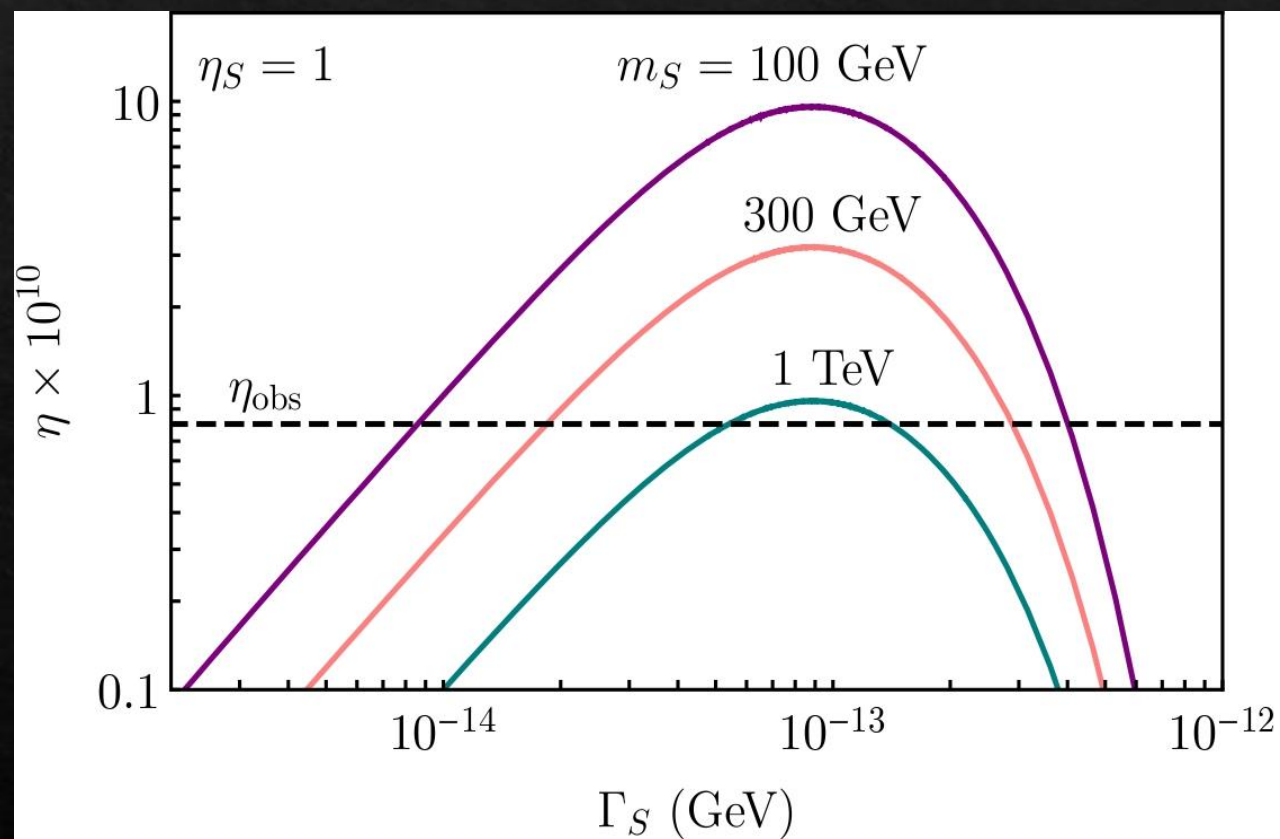
$$\frac{d\Delta_{e_R}}{dt} = -3H\Delta_{e_R} - \Gamma_{LR}\Delta_{e_R} + B_e\Gamma_S\Delta_S$$

$$\Delta_S = \eta_S \frac{\rho_S}{m_S}$$

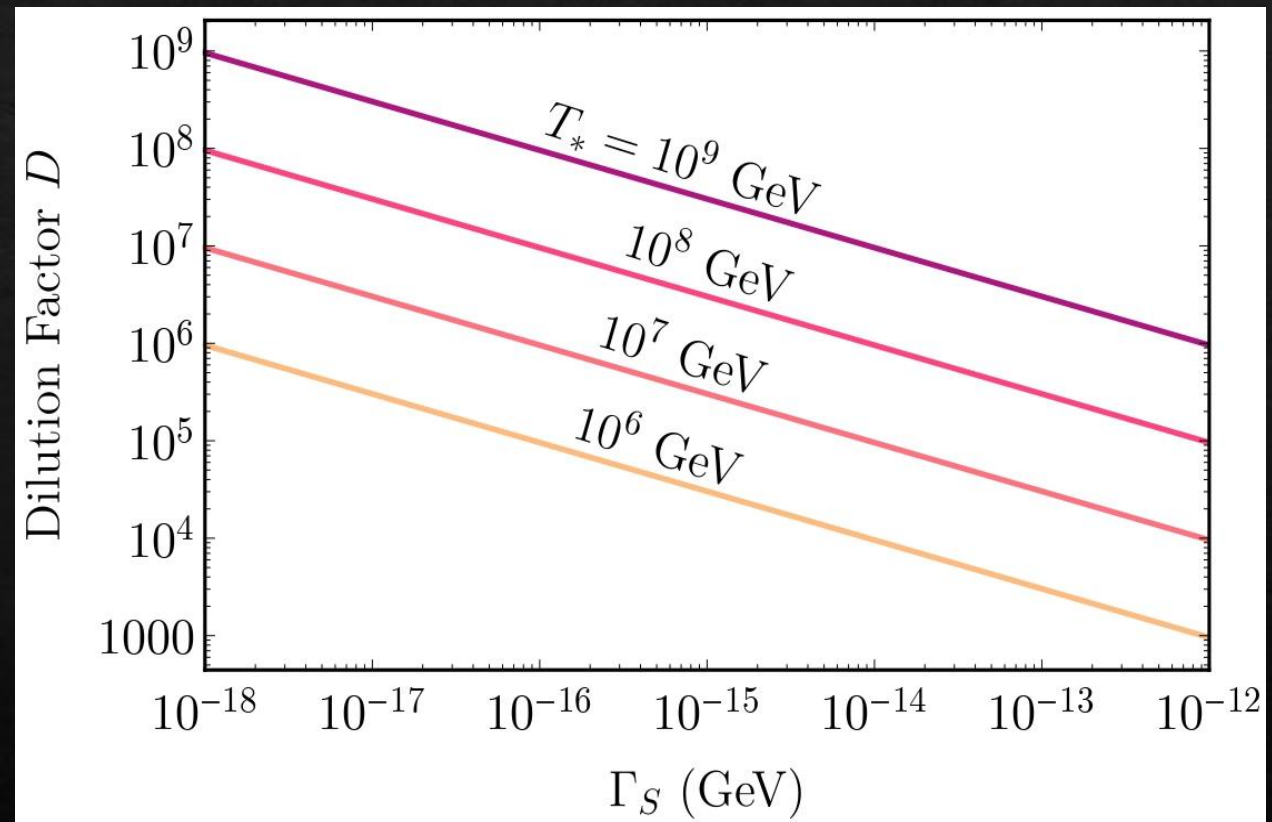
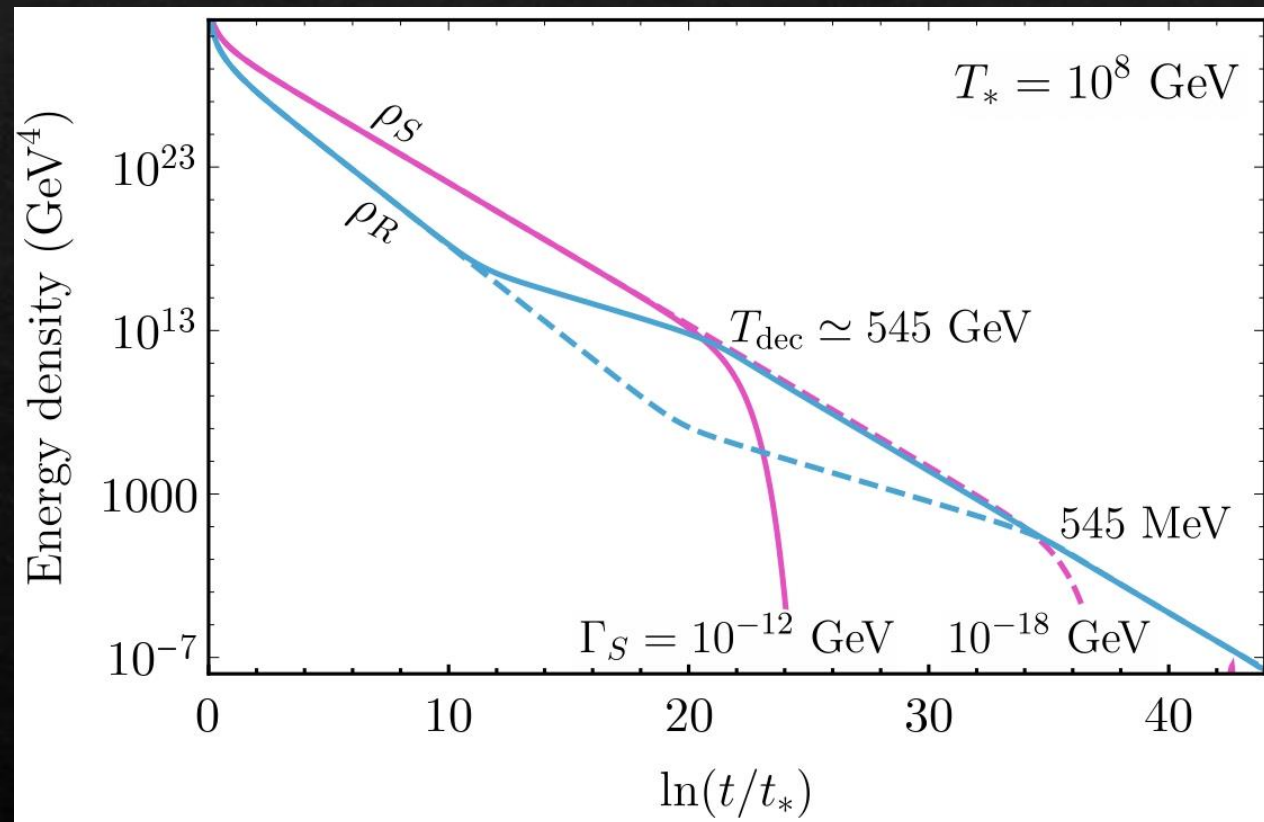
$$\eta \equiv \frac{n_B - n_{\bar{B}}}{s} \simeq \frac{198}{481} \frac{\Delta_{e_R}(T=T_{\text{EW}})}{s}$$

$$\eta_{\text{obs}} = \frac{n_B - n_{\bar{B}}}{s} \simeq 8 \times 10^{-11}$$

Flavon baryogenesis



Flavon baryogenesis

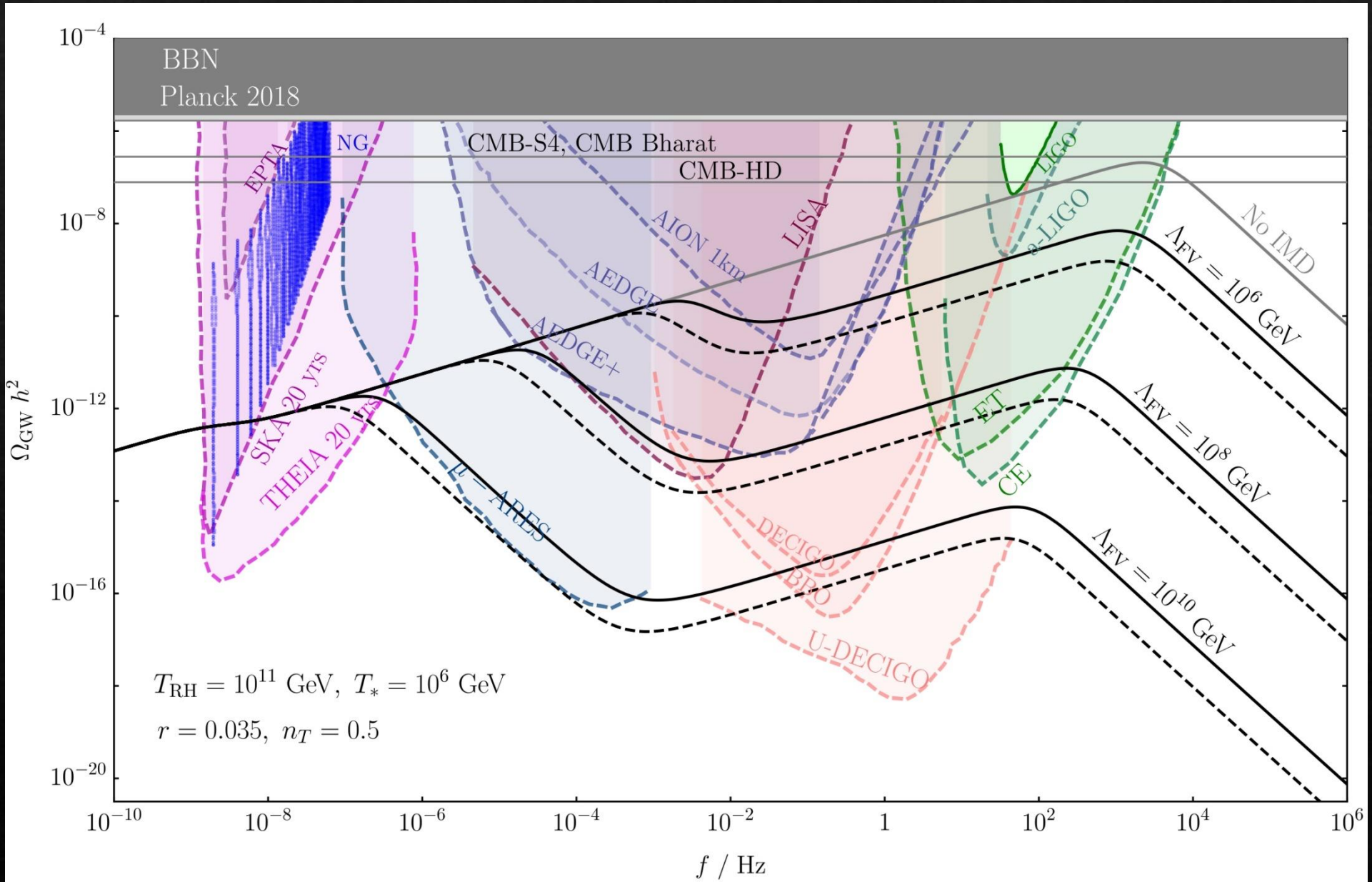


Flavon baryogenesis

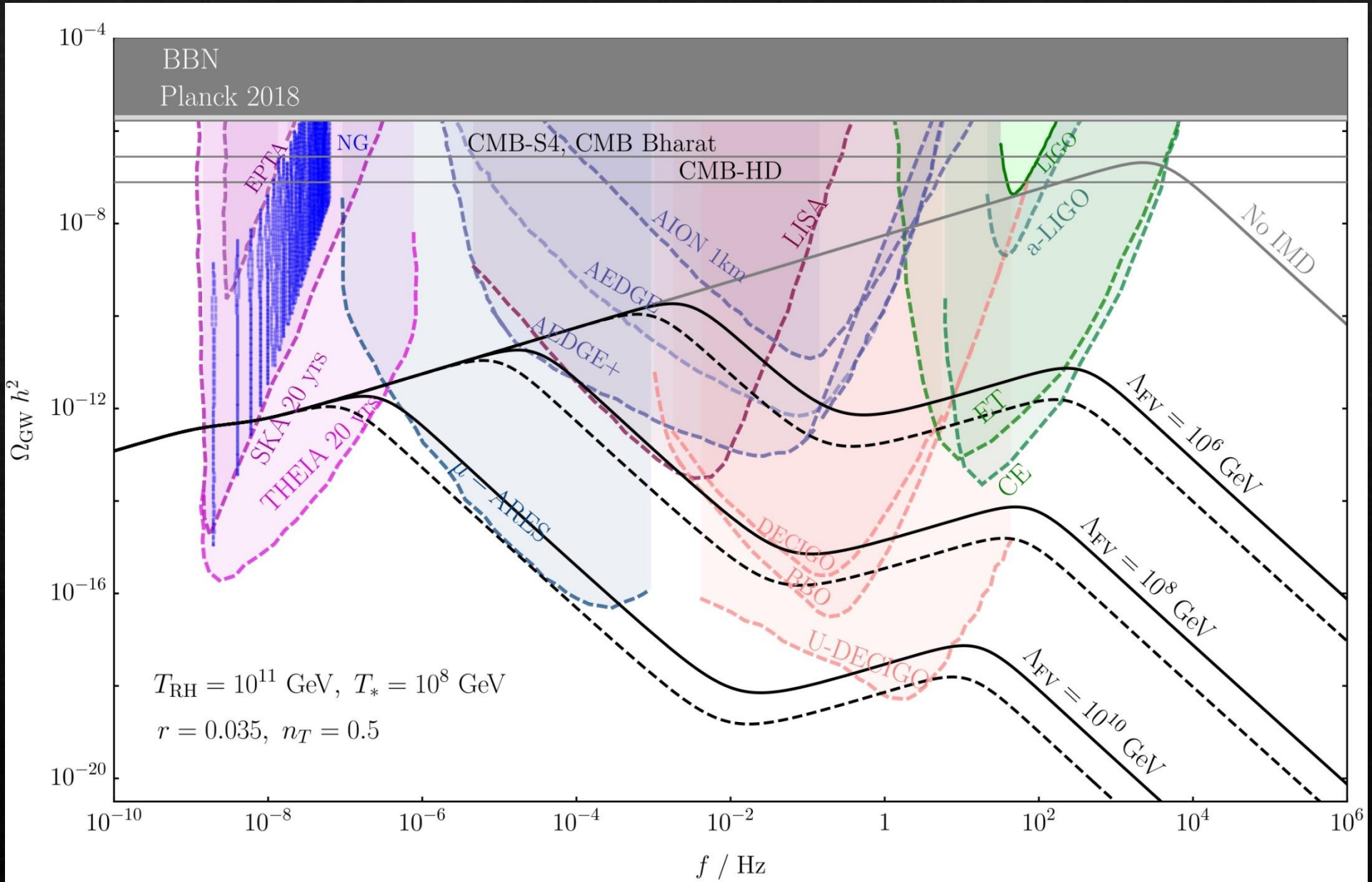
$$T_{\text{dec}} \simeq 1.8 \text{ GeV} \sqrt{\frac{\Gamma_S}{10^{-17} \text{ GeV}}} \simeq 2.7 \text{ GeV} \left(\frac{m_S}{\text{TeV}}\right)^{3/2} \left(\frac{10^{10} \text{ GeV}}{\Lambda_{\text{FV}}}\right)$$

$$D = \frac{s(T_{\text{after}})a^3(T_{\text{after}})}{s(T_{\text{before}})a^3(T_{\text{before}})} = \left(1 + 2.95 \left(\frac{2\pi^2 \langle g_*(T) \rangle}{45}\right)^{1/3} \frac{(\rho_S/s|_{\text{initial}})^{4/3}}{(M_{\text{Pl}} \Gamma_S)^{2/3}}\right)^{3/4}$$
$$\simeq 2 \times 10^6 \left(\frac{T_*}{10^6 \text{ GeV}}\right) \left(\frac{\Lambda_{\text{FV}}}{10^{10} \text{ GeV}}\right) \left(\frac{\text{TeV}}{m_S}\right)^{3/2},$$

Probing Flavon baryogenesis



Probing Flavon baryogenesis

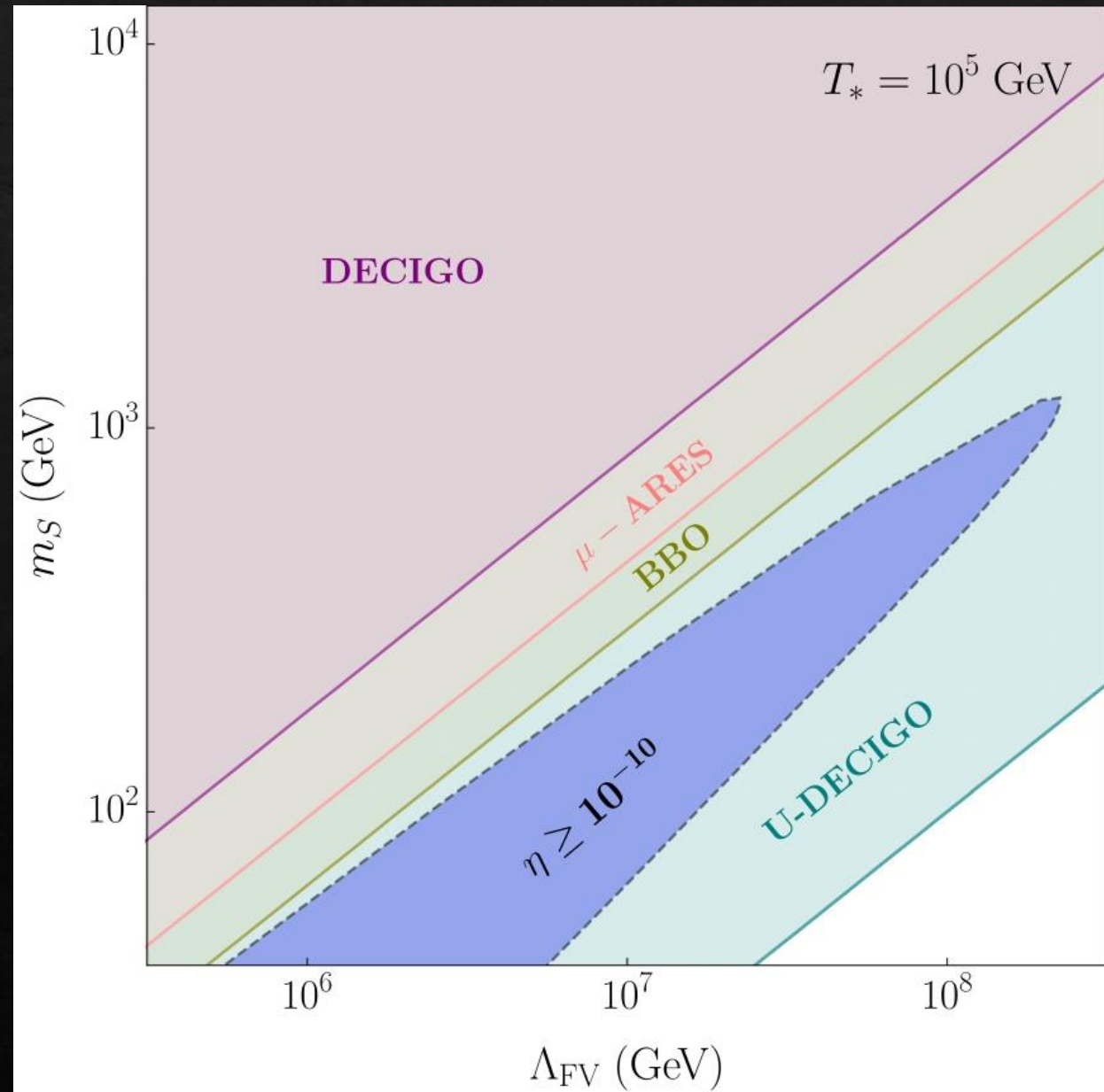
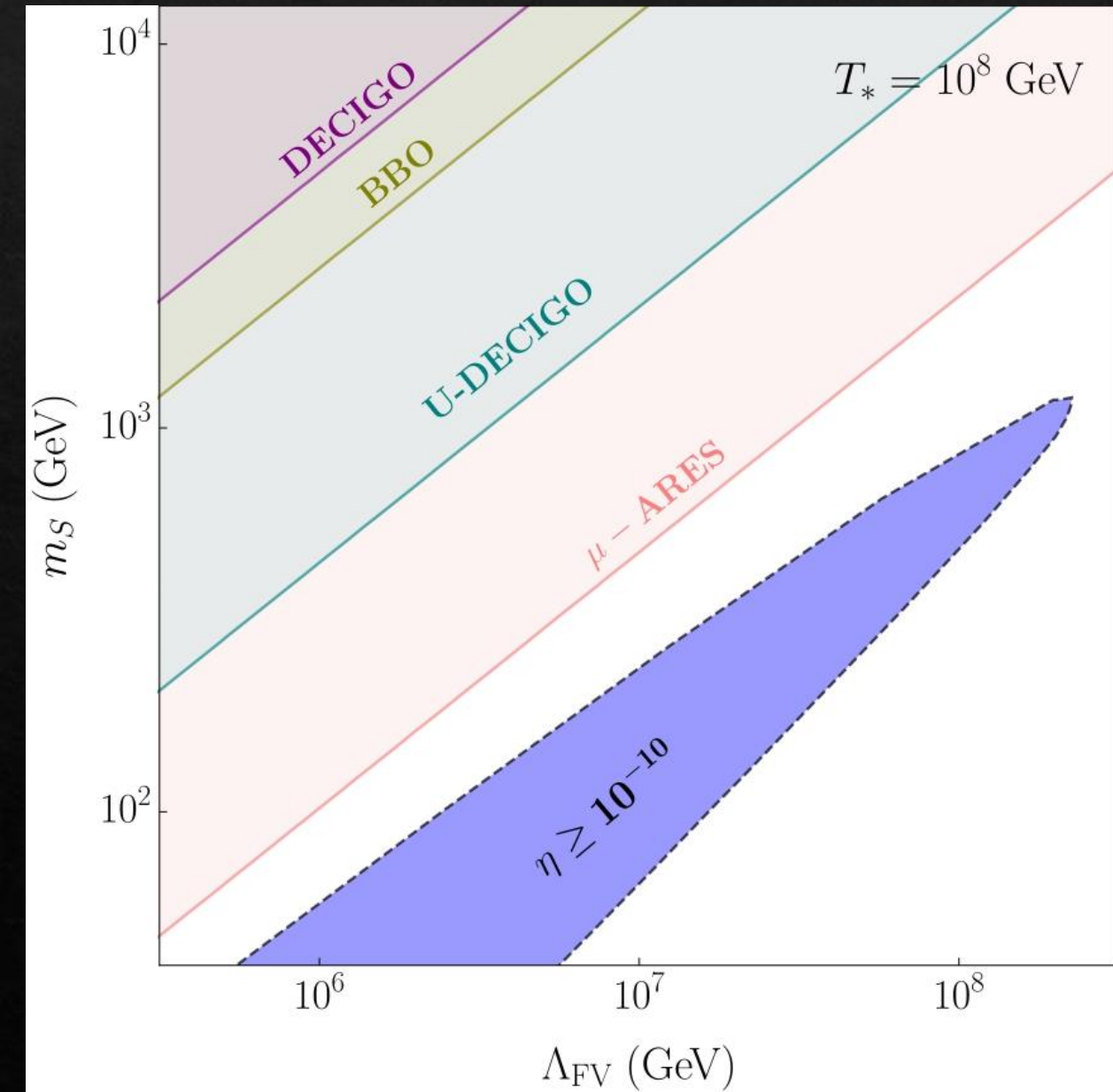


Probing Flavon baryogenesis

$$\Omega_{\text{exp}}(f)h^2 = \frac{2\pi^2 f^2}{3H_0^2} h_{\text{GW}}(f)^2 h^2$$

$$\text{SNR} \equiv \sqrt{\tau \int_{f_{\text{min}}}^{f_{\text{max}}} df \left(\frac{\Omega_{\text{GW}}(f)h^2}{\Omega_{\text{exp}}(f)h^2} \right)^2}$$

Probing Flavon baryogenesis



Conclusions

- ✓ We explore the connection between flavor violation, baryogenesis, and gravitational waves, focusing on how the dynamics of the flavon field, which explains the fermion mass hierarchy in the SM, could produce a detectable baryon asymmetry and imprint unique spectral features in primordial GWs.
- ✓ We analyze the suppression of primordial gravitational wave spectra due to flavon domination and decay, identifying model parameters for which both baryon asymmetry and GW signals are detectable by future GW detectors like U-DECIGO, BBO, LISA, ET and μ -ARES.
- ✓ As GW detector technology advances, the precision achieved could enable the high-energy physics and gravitational wave communities to test BSM mechanisms related to flavor physics, matter-antimatter asymmetry, and inflationary cosmology in unprecedented detail.

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Thank you !!