

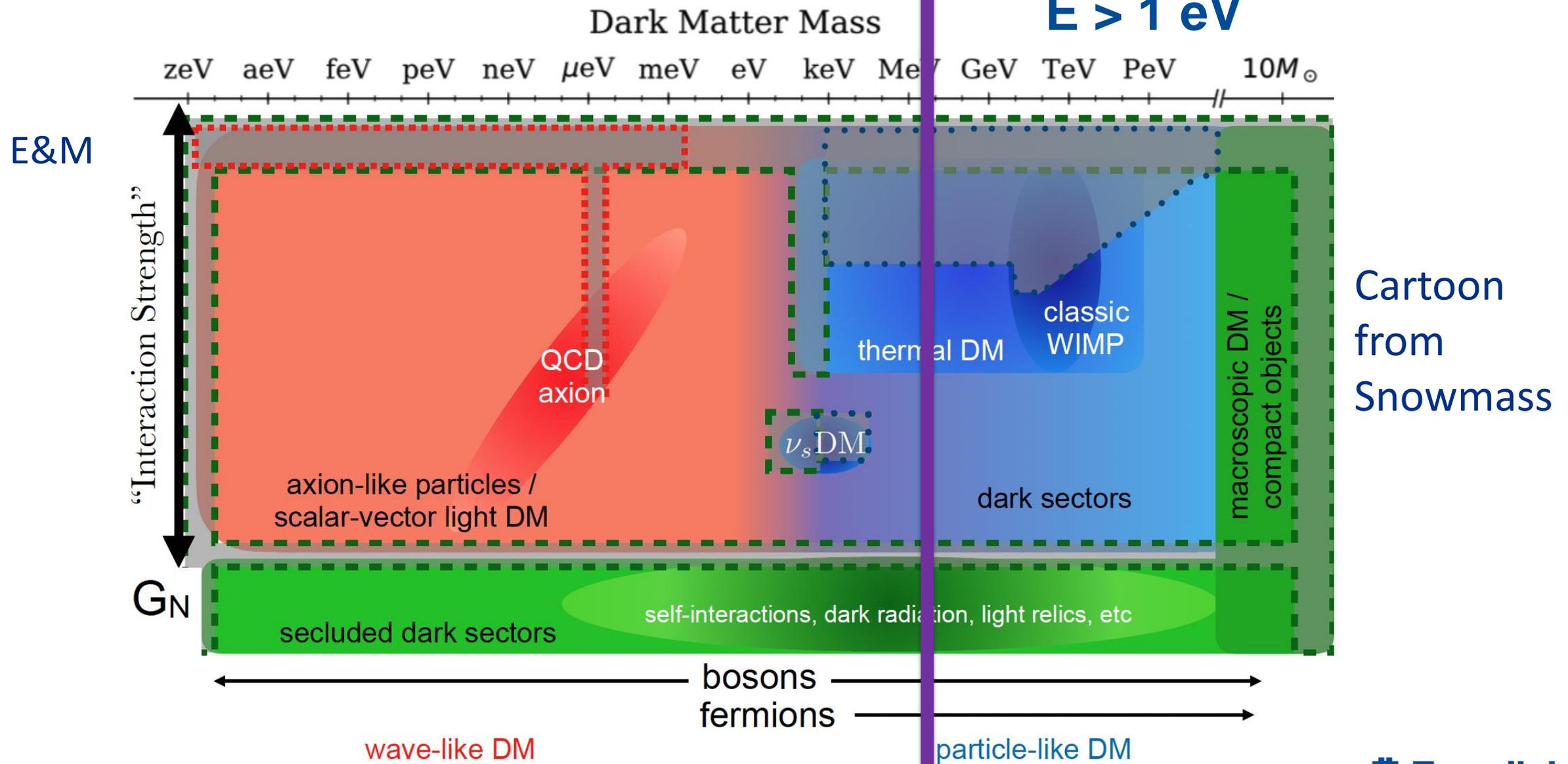
Detection of Ultraweak Forces from DM using Stimulated Emission

Aaron S. Chou

Fermilab

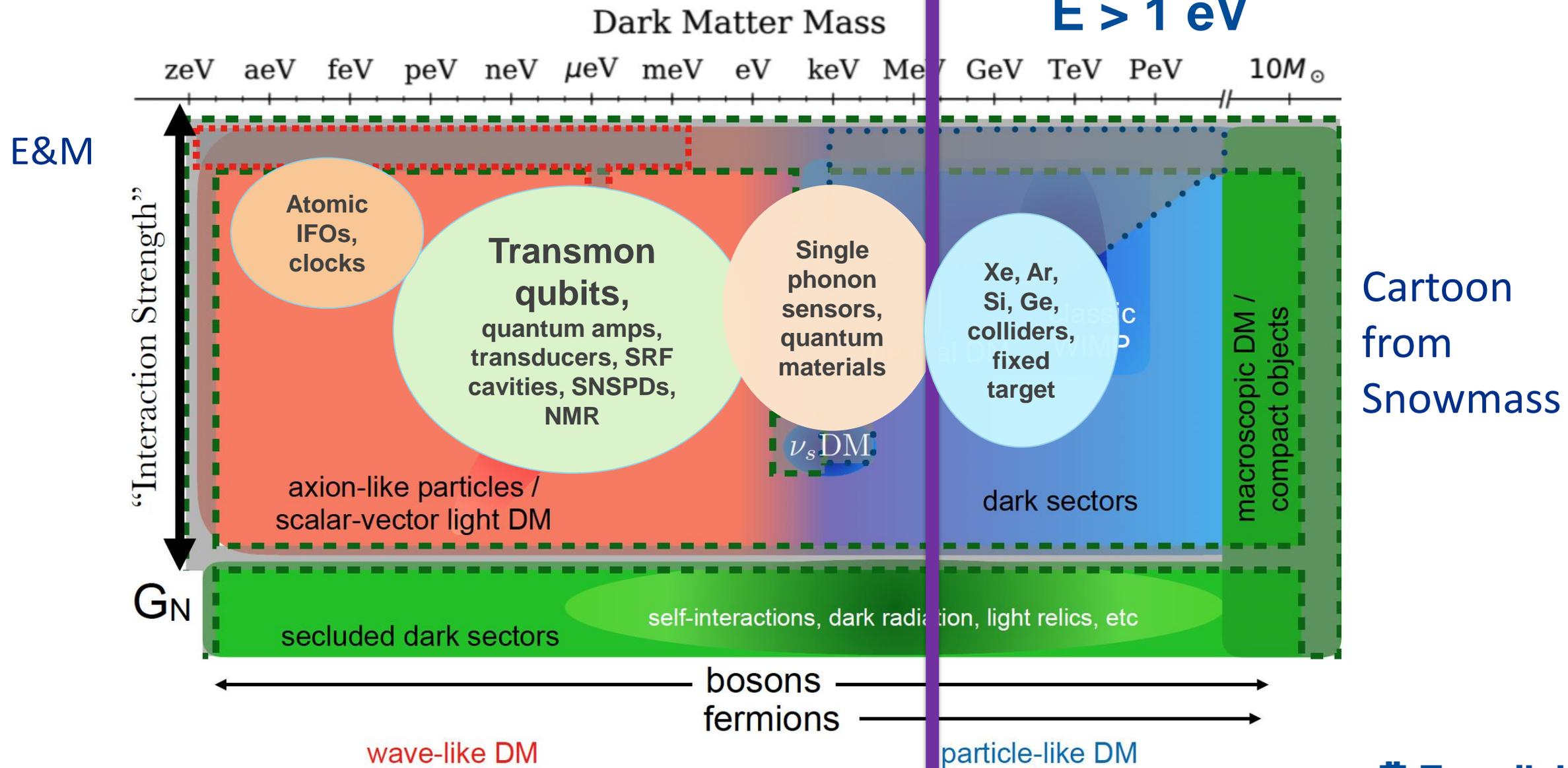
April 6, 2023

Quantum sensors: $E < 1 \text{ eV}$ ← → Conventional tech: $E > 1 \text{ eV}$



Cartoon from Snowmass

Quantum sensors: $E < 1 \text{ eV}$ ← → Conventional tech: $E > 1 \text{ eV}$



Cartoon from Snowmass

Hybrid Town Hall on Quantum Sensors for HEP

April 27 afternoon

Quantum Sensors for HEP

Apr 27 – 29, 2023
Yale University
US/Eastern timezone



Overview

Call for Abstracts

Timetable

Book of Abstracts

Registration

Participant List

Abstract submission is open for 2-slide presentations on new ideas,

i.e. not previously discussed at Snowmass or other past community meetings, and not already fundable by some existing program

<https://indico.fnal.gov/event/59102/>

The goal of this workshop is to explore the most promising directions for applying quantum sensing technologies to DOE-OHEP science targets, with a focus on sensors that could be deployed in future DOE-funded experiments. While we will provide an overview of existing DOE-OHEP quantum sensing programs for context, the workshop's main emphasis will be on novel ideas that can form the foundation of new DOE-OHEP quantum sensing programs or possibly to significantly enhance current programs. The goal is to pinpoint areas where DOE-OHEP can have a unique impact, leveraging its people, technological capabilities, and facilities. We are particularly interested in identifying new research directions not currently covered by existing funding sources and which could benefit the DOE-OHEP mission.

The in-person workshop is open to invited participants and we will have a hybrid town hall to capture ideas from the broader community. Travel and other local information can be found on the event page here: <https://campuspress.yale.edu/quantisedhep23/>.



Starts Apr 27, 2023, 9:00 AM
Ends Apr 29, 2023, 1:00 PM
US/Eastern



Yale University
Yale Quantum Institute / Wright Laboratory
17 Hillhouse Ave., 4th floor
New Haven, CT 06511
[Go to map](#)



Aaron Chou
Kathryn Zurek
Kent Irwin
Reina Maruyama

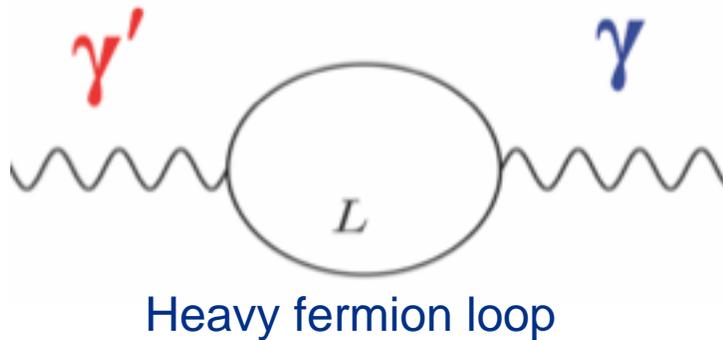


There are no materials yet.

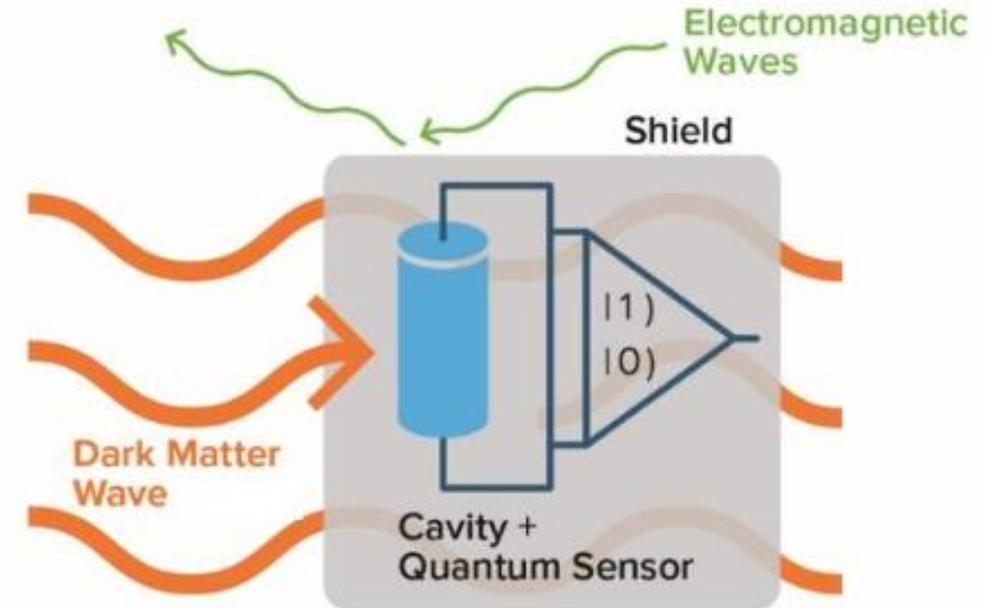
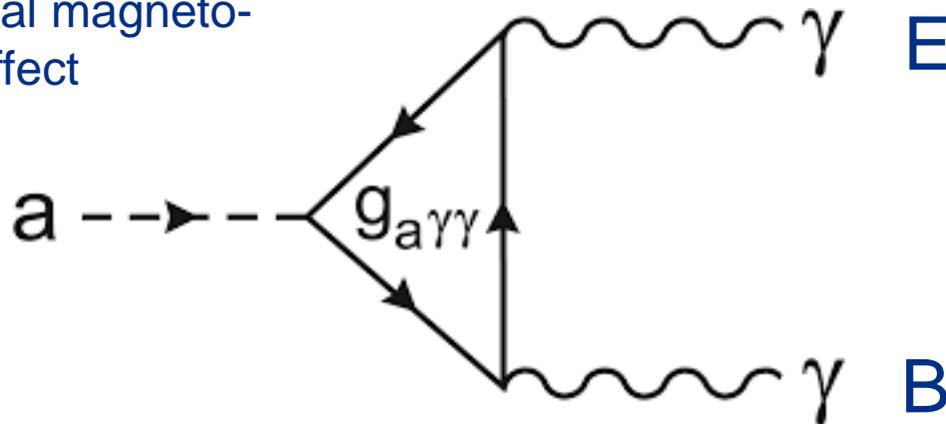


Bosonic DM waves: dark photons, spin-0 particles like the QCD axion

Flavor mixing



Topological magneto-electric effect



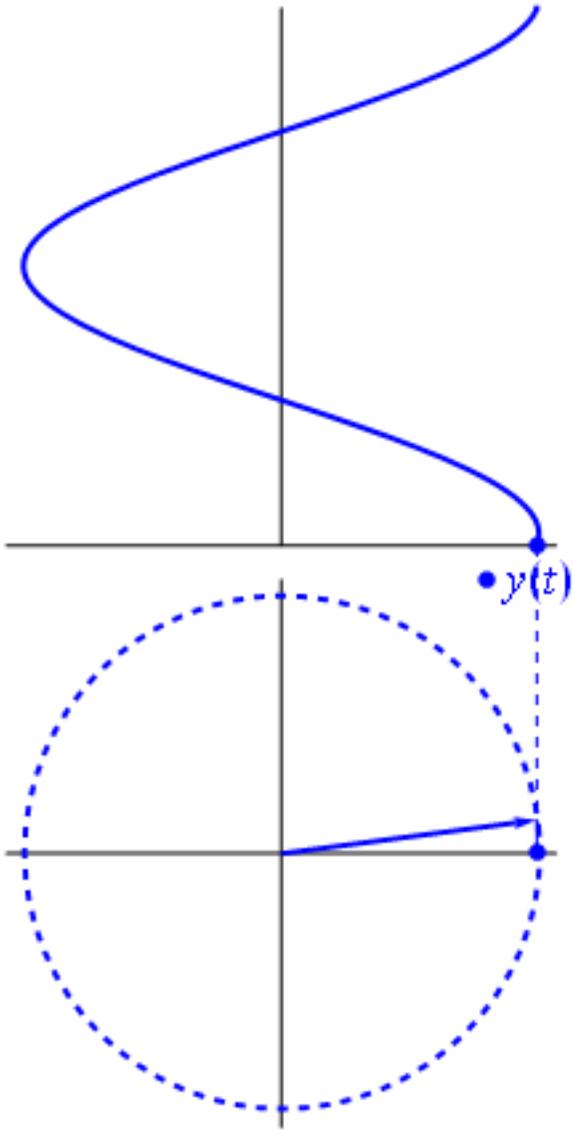
Forces from classical dark matter waves can deliver energy/momentum in the form of single photons that mysteriously appear in your well-shielded apparatus.

Energy transfer between two coupled oscillators



Weak coupling -- takes many swings to fully transfer the wave amplitude.
In real life, the number of swings is limited by coherence time and one only gets the transfer of an occasional single quantum of energy.

Sine waves are described by phasors



Quantum mechanics:

Endpoint of phasor in X, P phase space has uncertainty principle

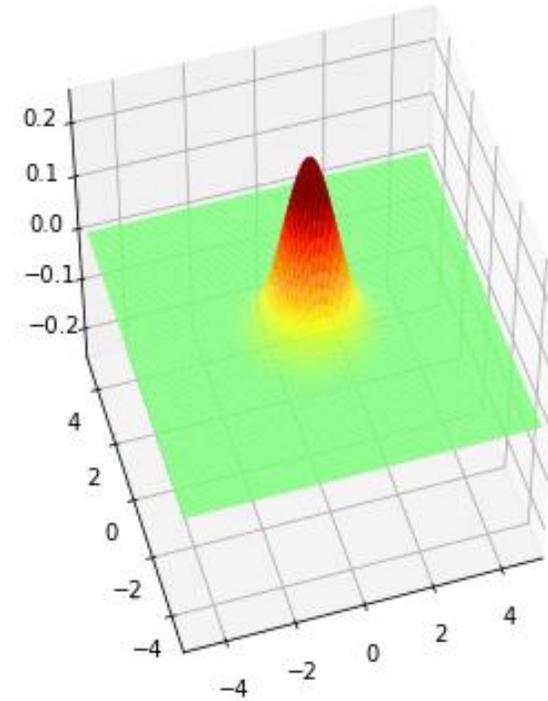
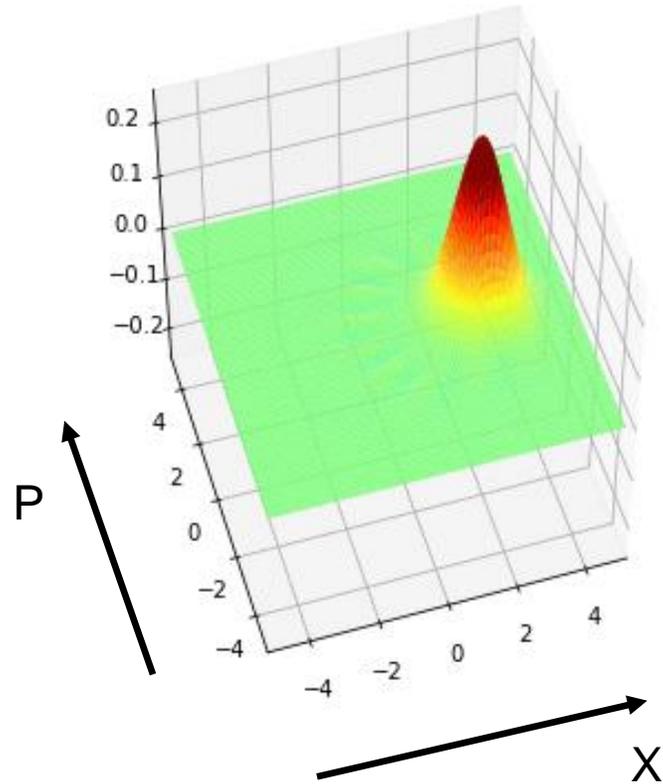
$$\Delta X \times \Delta P \geq \frac{1}{2}.$$

In polar coordinates this becomes:

$$\Delta N \times \Delta \varphi \geq \frac{1}{2}.$$

Classical pendulum system: $|\alpha = 3\rangle \otimes |\alpha = 0\rangle$

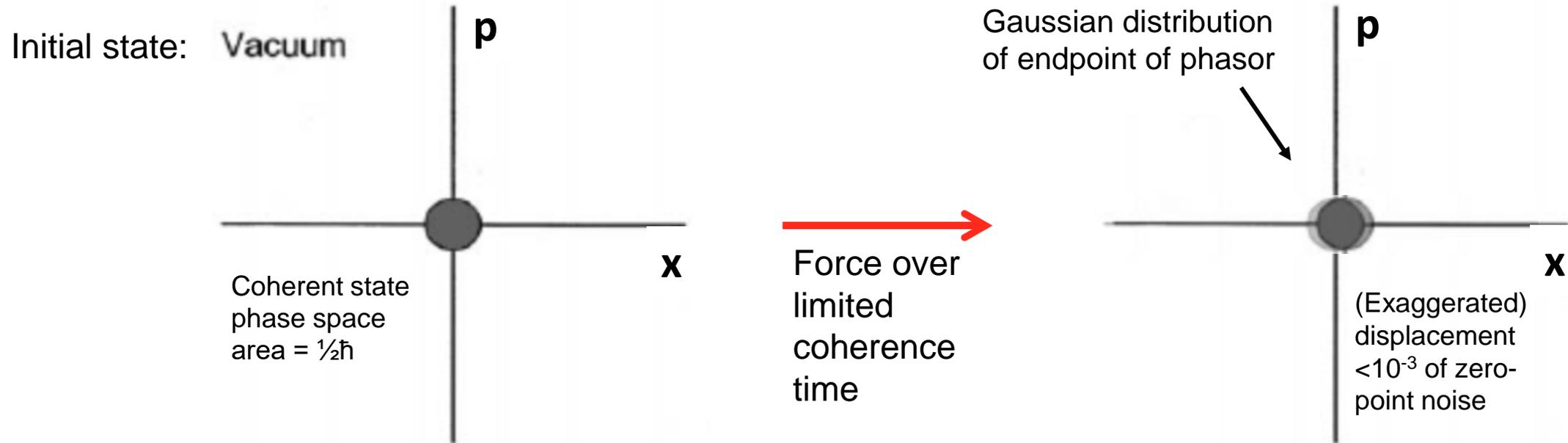
Time evolution of Wigner distributions in X-P “phasor” space.
Each gaussian blob of phase space area satisfies $\Delta X \cdot \Delta P = \frac{1}{2}$



Simulated with QuTIP

Given enough time, the two pendula swap their coherent states.
Real experiments are limited by coherence time so only get the occasional single quantum of signal

Big Problem: Signal/Noise Ratio

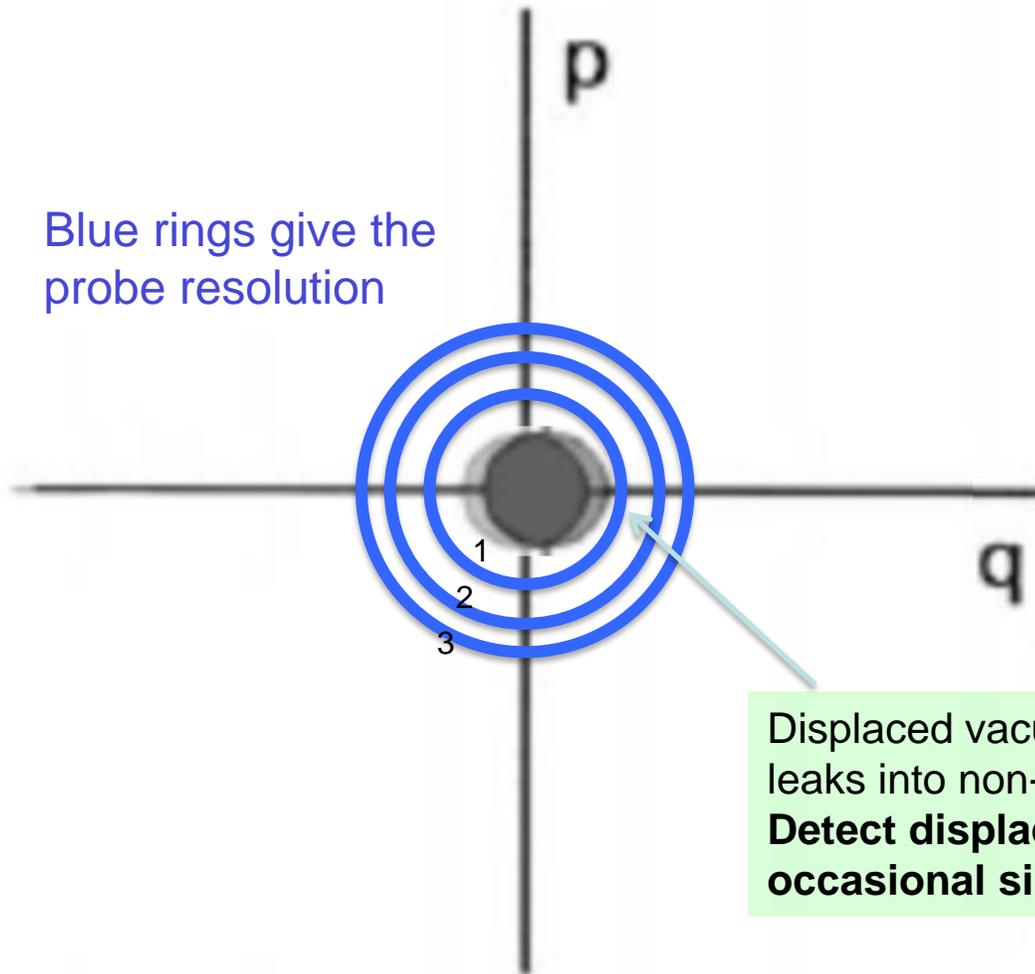


The Gaussian blob is also effectively the probe resolution when using a linear amplifier to resolve the tiny signal sine wave. Simultaneous measurement of non-commuting observables N and φ incurs the Heisenberg uncertainty principle $\Delta N \times \Delta \varphi \geq \frac{1}{2}$.

Need millions of standard-quantum-limit measurements to average away the zero-point noise to resolve the tiny displacement signal.

To reduce readout noise, use photon counting to project displacements into the Fock basis, i.e. number eigenstates

Previously we measured *both amplitude and phase*, but this is dumb since the dark matter phase is random and not interesting. Useless information obtained at high cost!

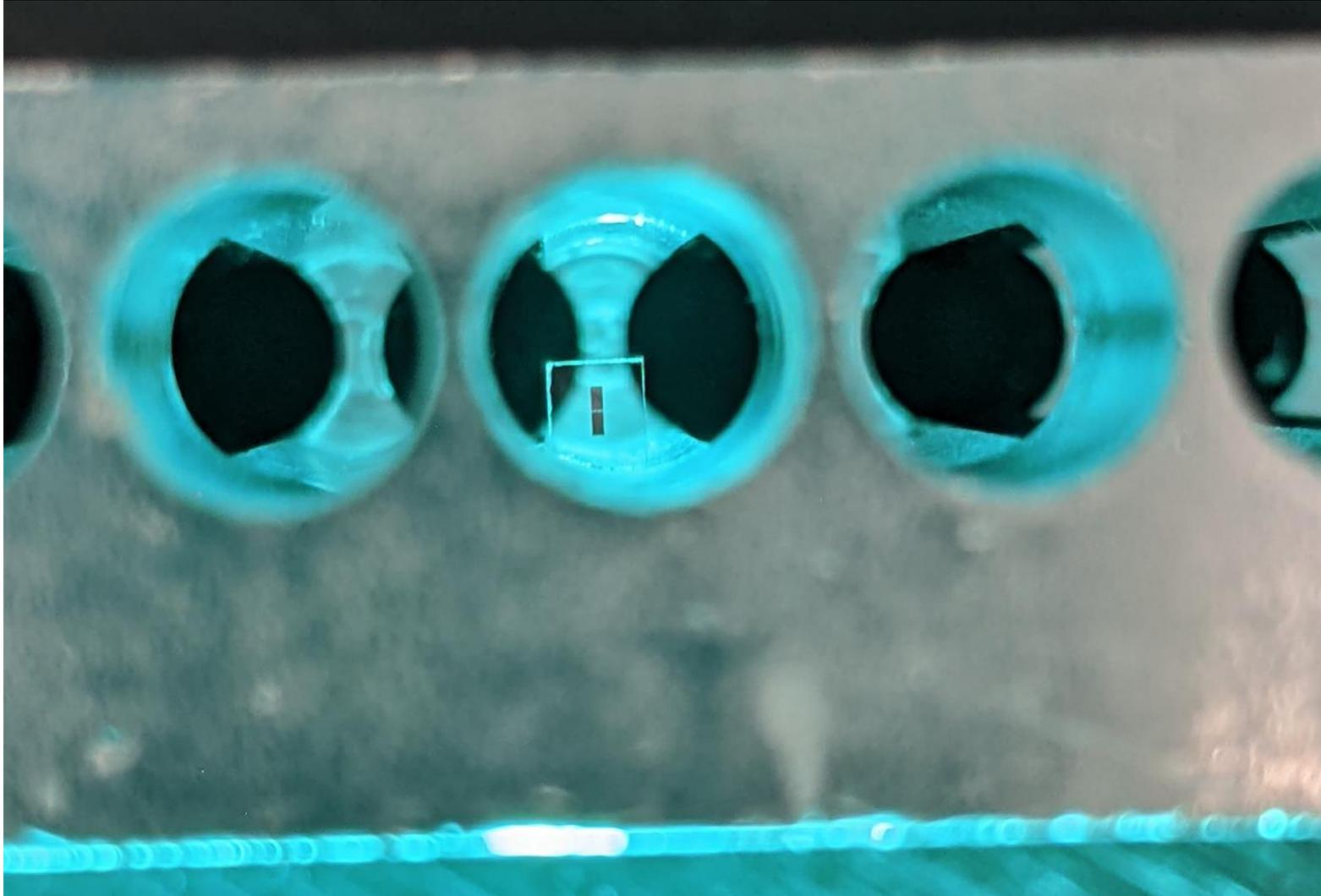


Instead, measure only displacement amplitude
→ no zero-point noise since we are not simultaneously measuring non-commuting observables
→ **Measurement noise, e.g. from dark counts can be arbitrarily low**

Displaced vacuum state exponentially leaks into non-zero Fock number.
Detect displacement by counting the occasional single photon.

Transmon qubit in 6 GHz aluminum cavity

Image credit: Akash Dixit



Cavity = quantum RAM

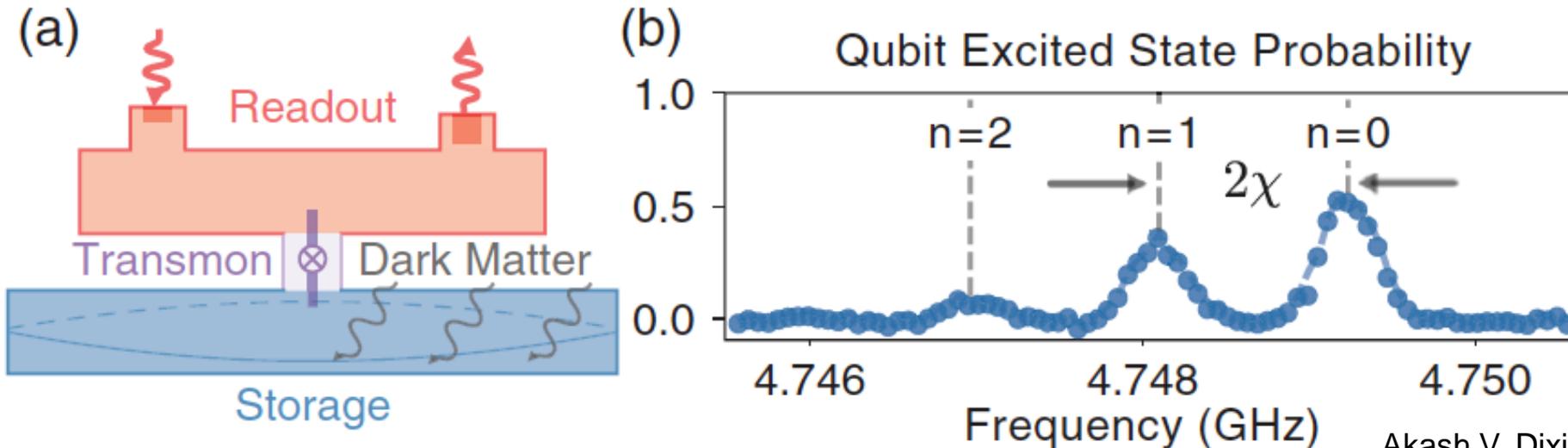
Qubit = nonlinear LC oscillator

Aaron S. Chou, IQ/PACC axion workshop,
4/6/2023

Dark matter waves scatter on walls of cavity, depositing occasional single photons whose electric field drives the qubit nonlinear. → Measure AC Stark shift of the qubit's frequency.

Detect single photons via the AC Stark shift of the qubit:

Example: Measure qubit $|g\rangle \rightarrow |e\rangle$ transition frequencies after mimicking the DM signal by weakly driving the primary cavity mode into a coherent state with $\langle n \rangle = 1$



Akash V. Dixit, et.al, Phys.Rev.Lett.
126, 141302 (2021)

The measured qubit response spectrum exhibits a distribution of resonances which are in 1-1 correspondence with the Poisson distribution of the cavity's coherent state.

- Non-destructively count photons by measuring the qubit's quantized frequency shift via this "quantum non-demolition" measurement.
- Repeat many times for high fidelity single photon detection.
- **Demonstrated reduction in noise levels to 40x below the Standard Quantum Limit**

More stupid qubit tricks: Increase signal rate by stimulated emission of photons from DM waves

Start the photon wave swinging so it can more easily accept energy from the DM.

$$\text{Power} = \overrightarrow{\text{Force}} \cdot \overrightarrow{\text{velocity}}$$



Good!



Waiting...



Oops, wrong phase

Phase offset determines the direction of energy flow.

Wait ... but we don't know the instantaneous phase of the DM wave!

What happens if we initialize the probe to a Fock state instead of a Gaussian blob?

A Fock state is a superposition of an oscillator in all possible phases of its sinusoidal motion: $\Delta N \times \Delta \varphi \geq \frac{1}{2}$
→ responds equally well to pushes at any time.

It also has definite occupation number N
→ no Poisson noise!

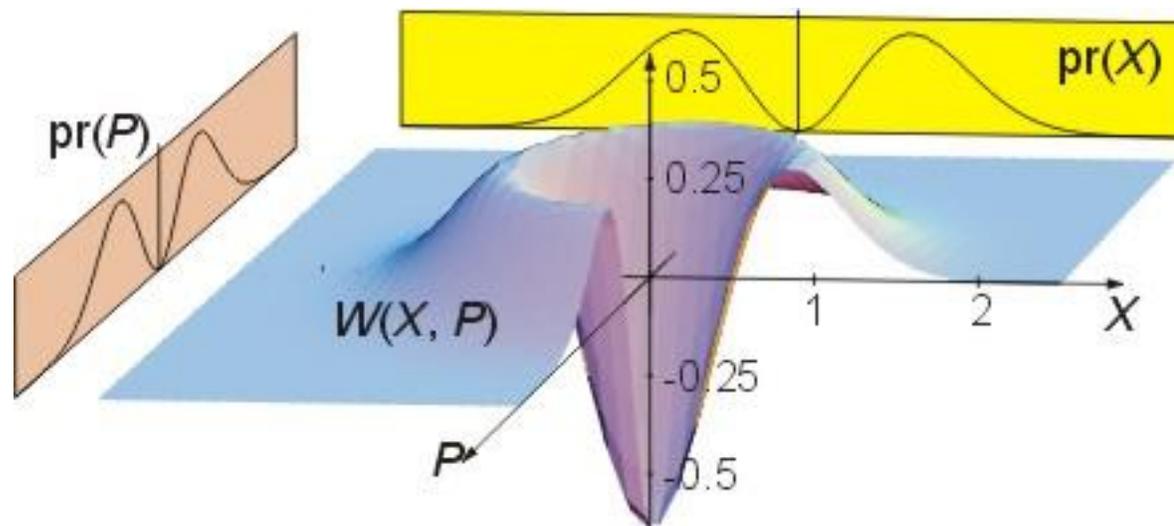
... and force \times velocity still works:

$$H_I = g(a^\dagger b + ab^\dagger) \rightarrow \langle \alpha, N+1 | H_I | \alpha, N \rangle = g\alpha\sqrt{N+1}$$



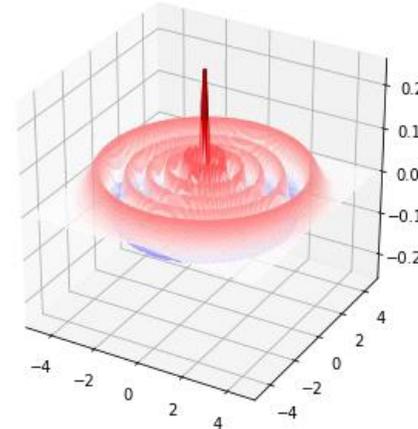
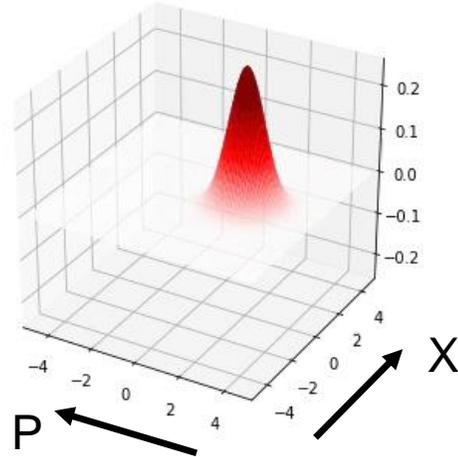
Idea from Konrad Lehnert
(JILA/U.Colorado)

N=1 Fock state



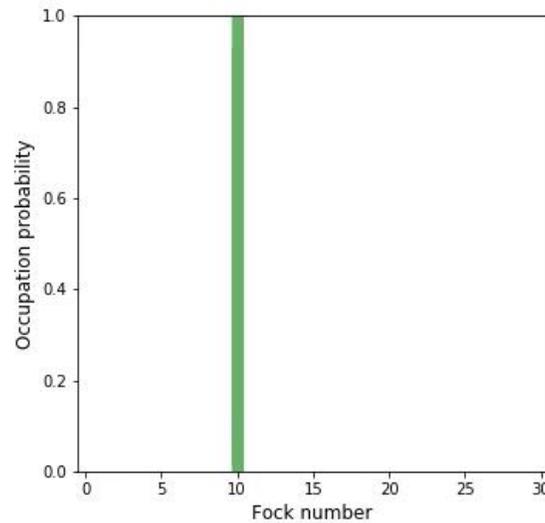
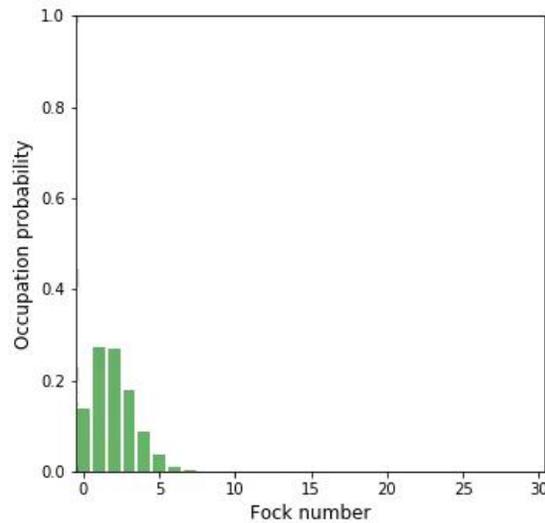
Mixing between a coherent state and a Fock state

The dark matter wave is a sine wave described by a Glauber coherent state



The cavity photon mode has been prepared as a $N=10$ Fock state, highly sensitive to environmental disturbances

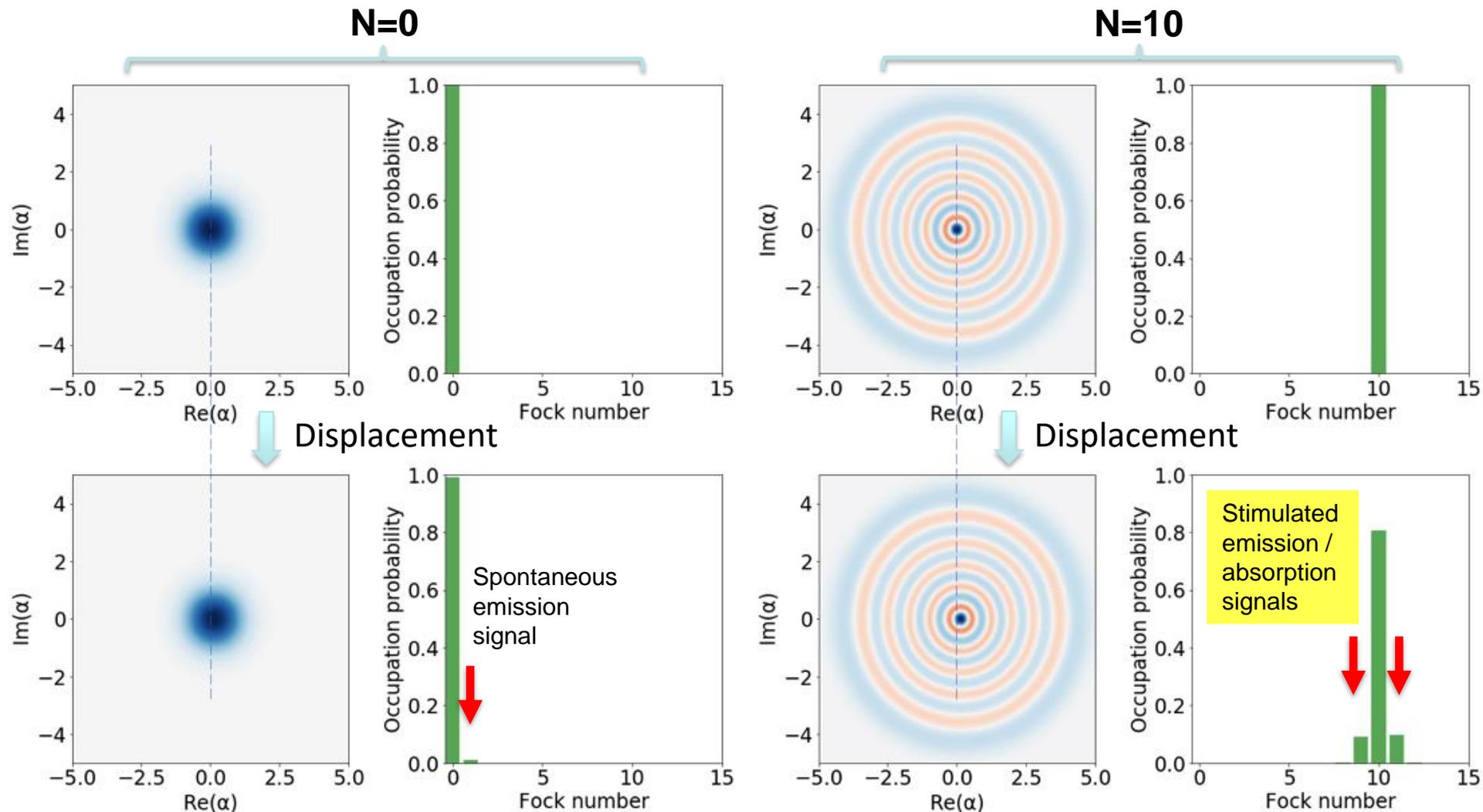
QuTiP simulation



The two oscillators swap states! This can be understood by decomposing the Fock state into a linear superposition of coherent states (which form an overcomplete basis for the phase space).

For small amplitudes α , the transfer of quanta from DM to photons is enhanced by a factor $(n+1)$

$$D(\alpha) |n\rangle \approx (1 + \alpha a^\dagger + \alpha^* a) |n\rangle = |n\rangle + \alpha\sqrt{n+1} |n+1\rangle + \alpha^*\sqrt{n} |n-1\rangle$$



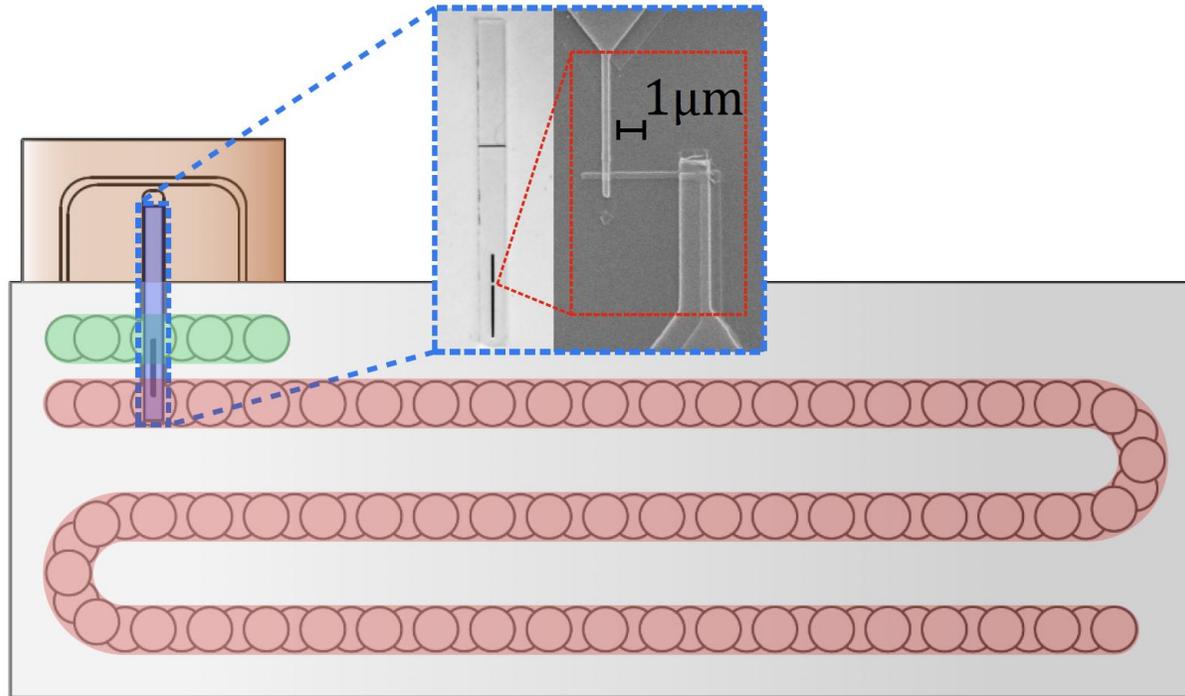
QuTiP simulation

Aaron S. Chou, IQ/PACC axion workshop, 4/6/2023

16

$$|\langle n+1 | \hat{D}(\alpha) |n\rangle|^2 \sim \alpha^2 (n+1)$$

Qubit-cavity system



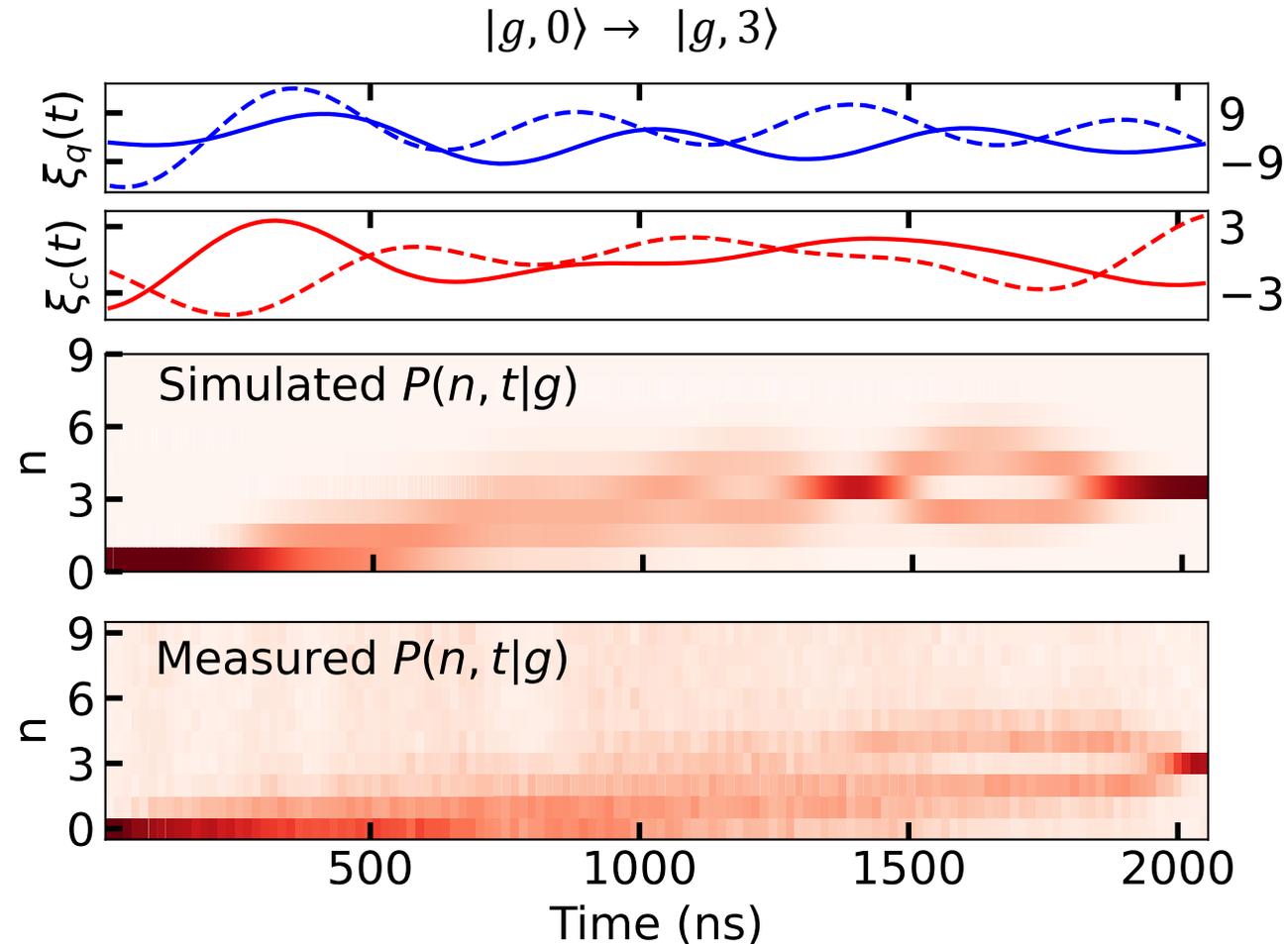
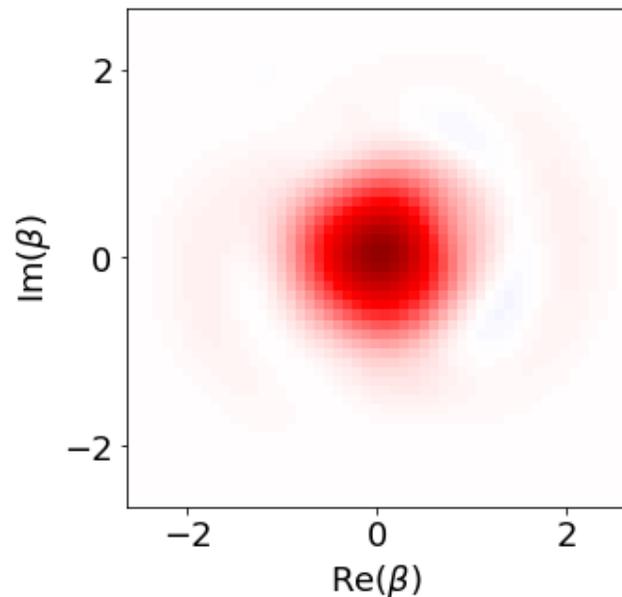
Storage	6.011 GHz
Readout	8.052 GHz
Transmon	4.749 GHz

$$\hat{\mathcal{H}} = \omega_c a^\dagger a + \omega_q \frac{\sigma_z}{2} + \chi a^\dagger a \frac{\sigma_z}{2}$$

Creating Fock states in a (non)linear system

“Optimal control” sequence of drives at both the qubit and cavity frequencies, determined by “gradient ascent pulse engineering”

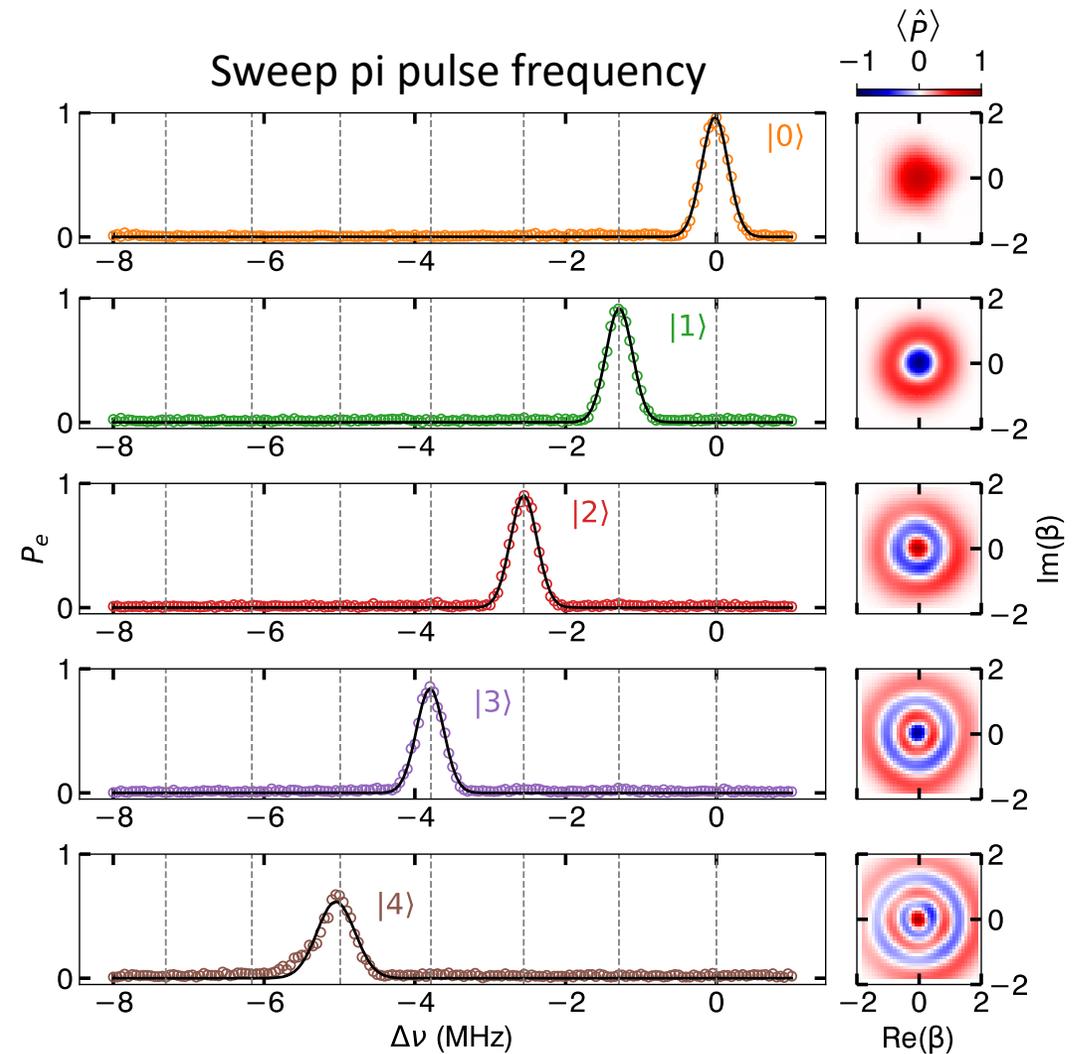
$$\hat{\mathcal{H}} = \left(\omega_c + \chi \frac{\sigma_z}{2} \right) a^\dagger a + \omega_q \frac{\sigma_z}{2}$$



Verifying the state preparation

Method 1: Qubit spectroscopy with a number resolved π pulse

Method 2: Wigner tomography to reconstruct the density matrix

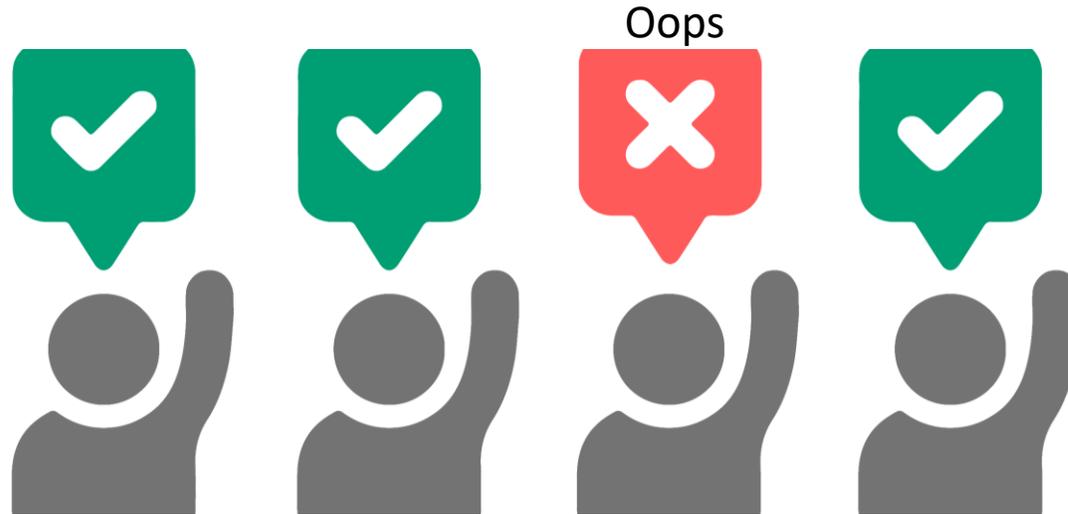


Repeated measurements to mitigate errors

$$\hat{H} = \omega_c a^\dagger a + \omega_q \frac{\sigma_z}{2} + \chi a^\dagger a \frac{\sigma_z}{2}$$

1st pi pulse if
successfully flips the
qubit, collapses cavity
wavefunction into
definite $|n\rangle$

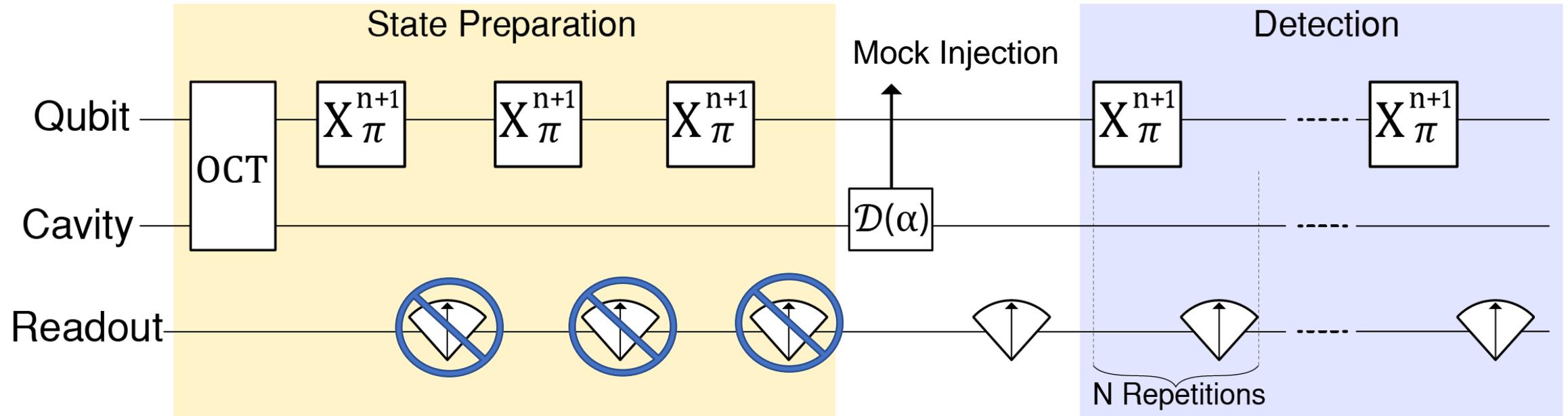
Subsequent pi pulses to
measure n are **quantum
non-demolition** since
 $[H, n]=0$



AV Dixit, AA et al.
PRL 126, 141302, 2021

Ask the qubit same question multiple times to boost the confidence and assign a probability for tagging the photon

Stimulated emission protocol

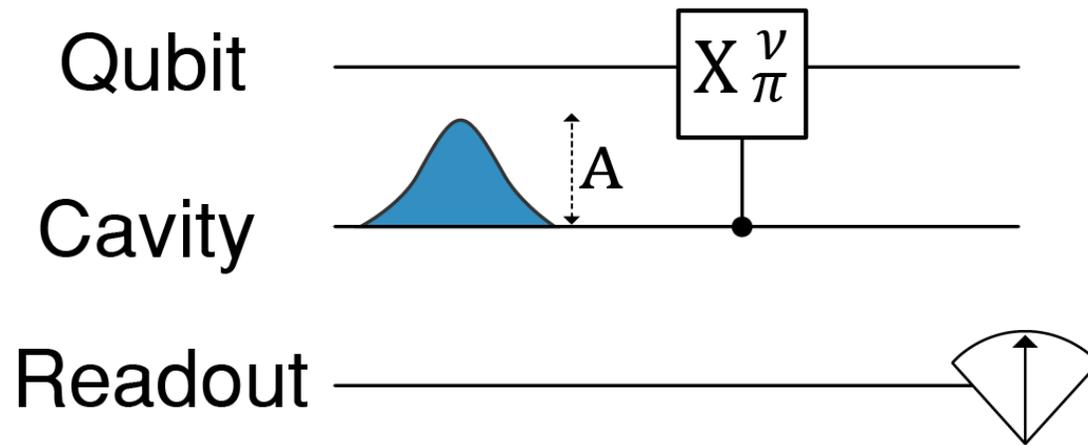


Ensures false positive probability from accidentally prepping $|n+1\rangle$; $< 3\%$, comparable to measured cavity occupation probability

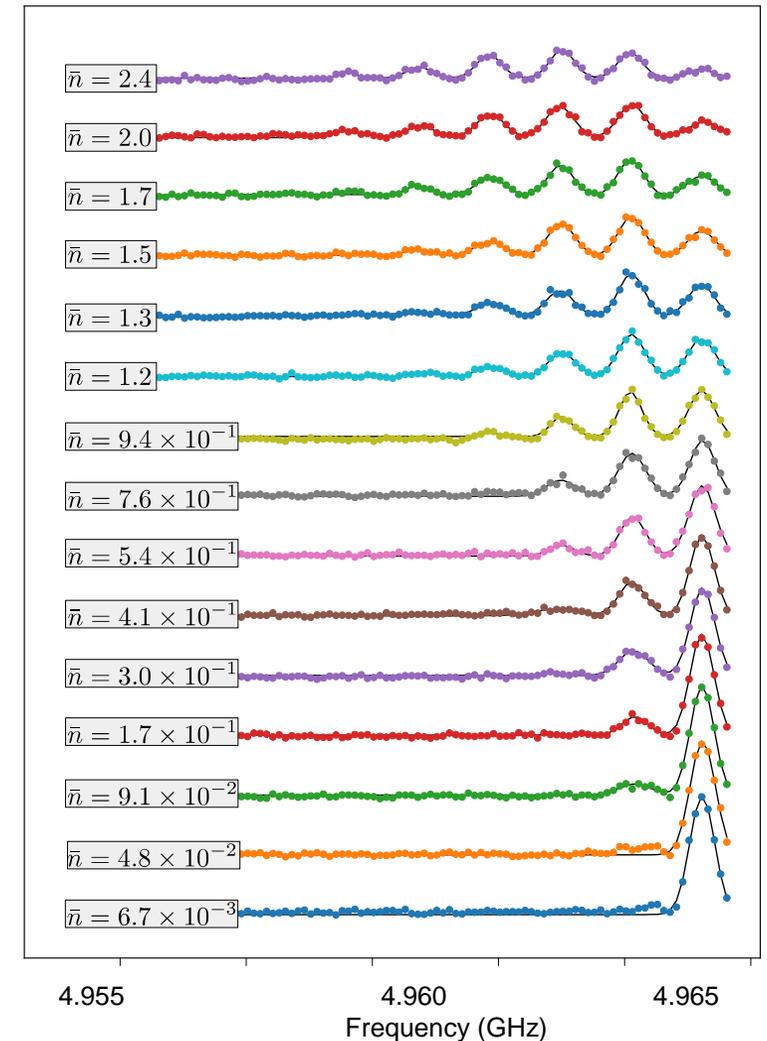
Hidden Markov model analyzes probability of observed sequence of yes's and no's from repeated pi pulses.

Absolutely calibrated photon injection

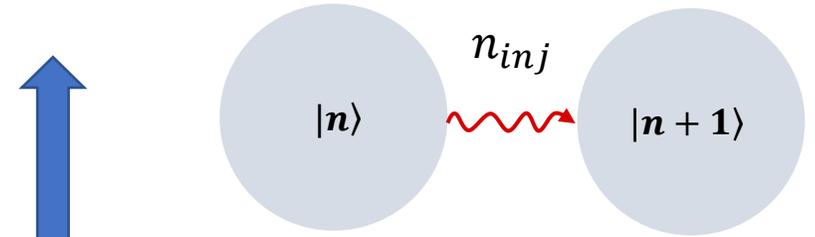
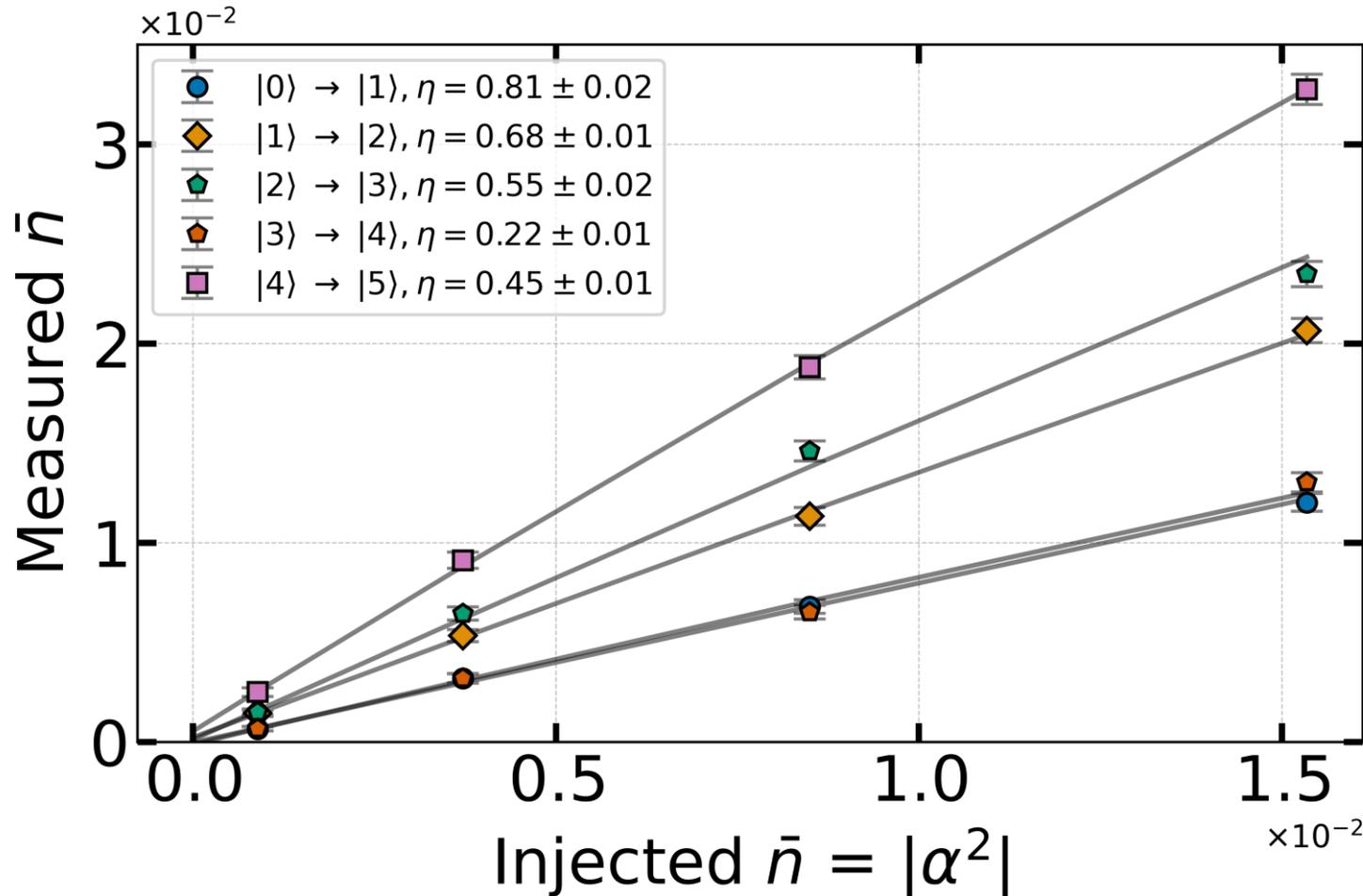
Inject sine wave in Gaussian envelope.



Measure resulting Poisson distribution of cavity occupation number



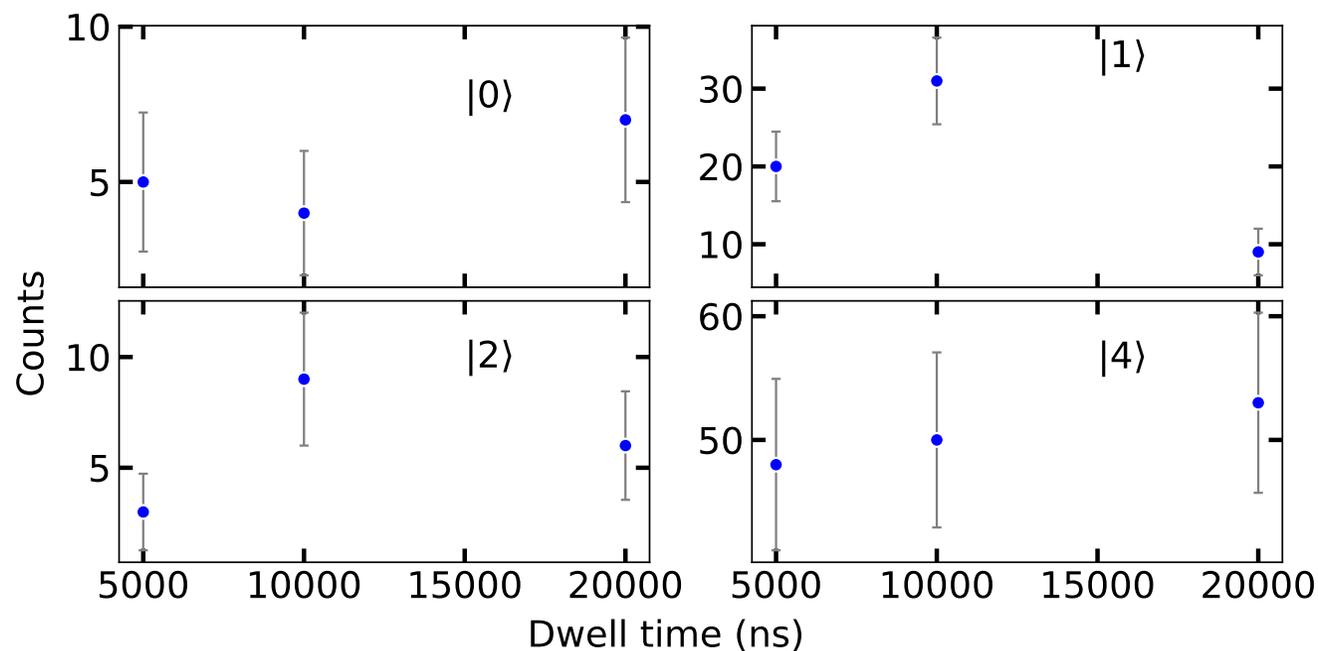
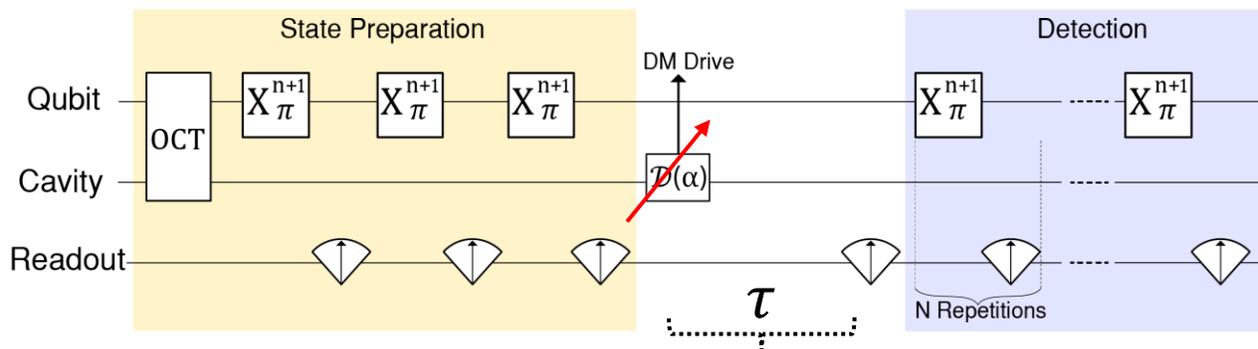
Signature of enhancement



This is the first demonstration so far showing signal enhancement with $|n\rangle = |4\rangle$ Fock state

Less than a factor of $(n+1)$ due to measurement inefficiencies, e.g. photon decays during measurement, QND has small demolition probability, etc. HMM optimized for small background rate $n_{\text{bgd}} < 1\%$

Backgrounds



~ 20000 trials for each point. $\tau \sim 10$ us, $\tau_{DM} \sim 100$ us, $\tau_{cav} \sim 1$ ms

Could be dark matter?

$$N_{meas} \propto a_0(n+1)\tau + b_0 + c_n$$

Coherent

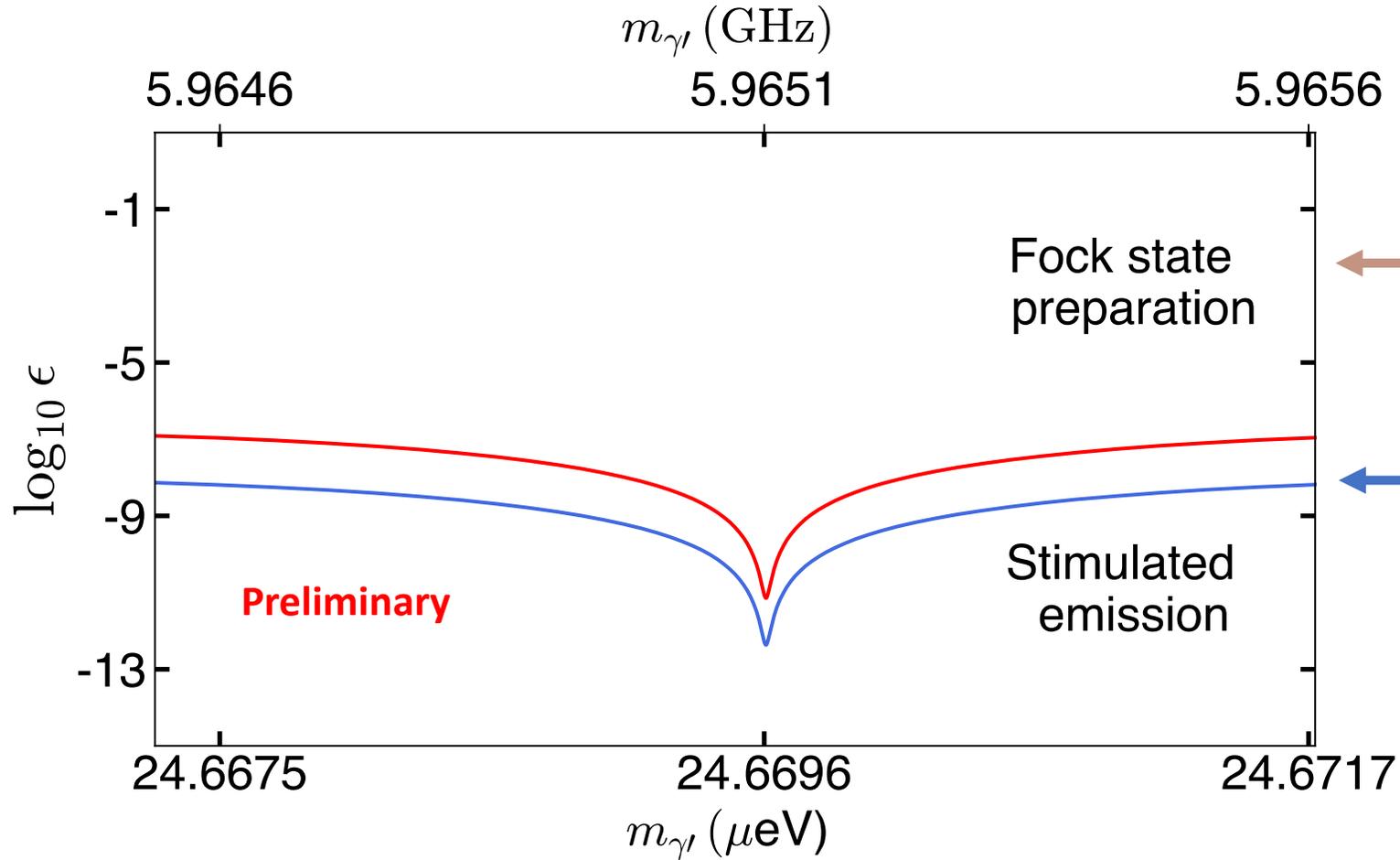
Incoherent

State prep.

$$\varepsilon \propto \sqrt{\frac{a_0}{m_{\gamma'}}$$

Hidden photon search

Not optimized. Just using data from zero injected signal.

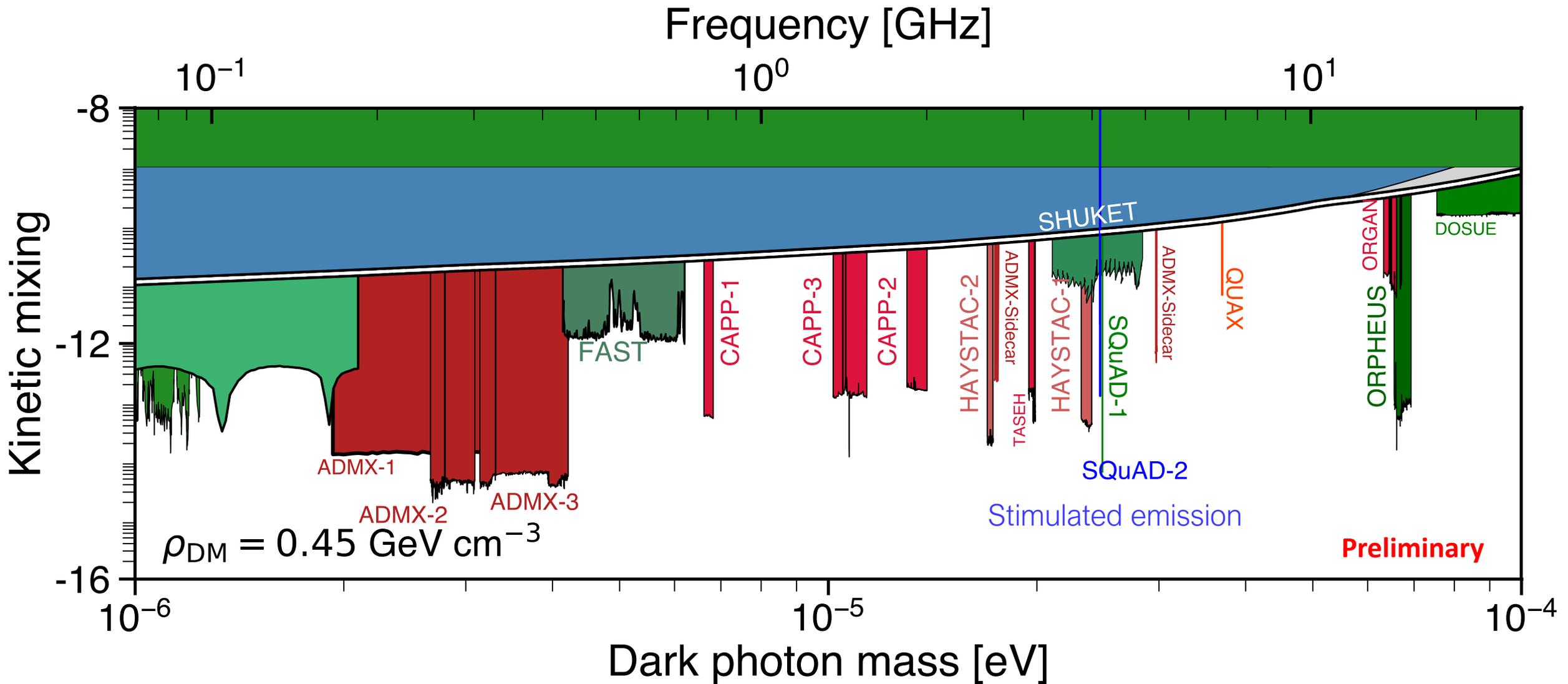


Dark matter drive is always on. If too large, this would have screwed up the Fock state prep.

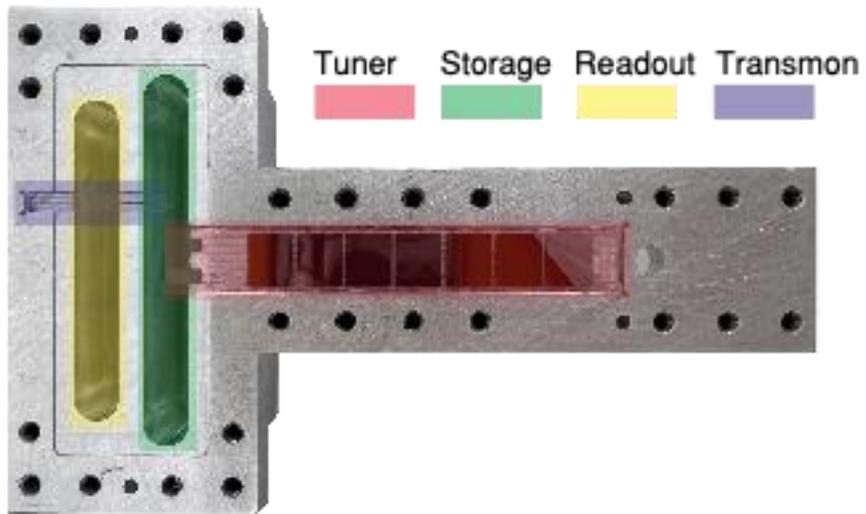
Conservative limit set by assuming all background counts are dark matter

Agrawal et. al.
(Preprint on arxiv soon)

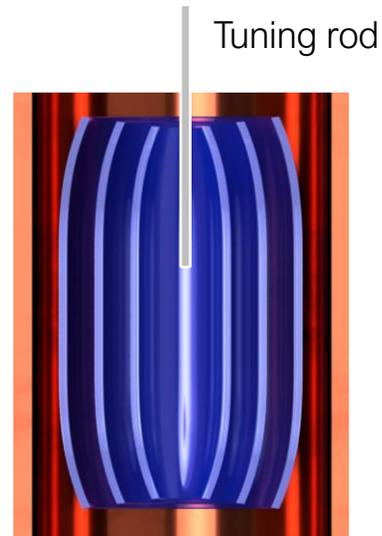
Hidden photon search



To make compatible with large B fields for axions

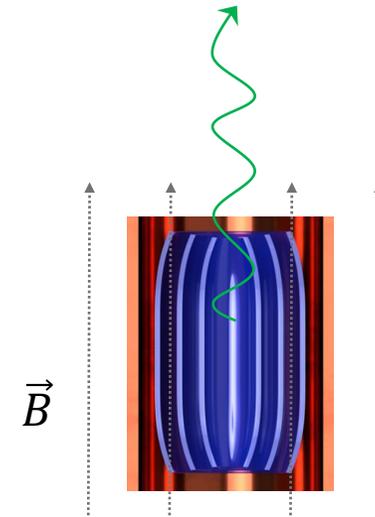
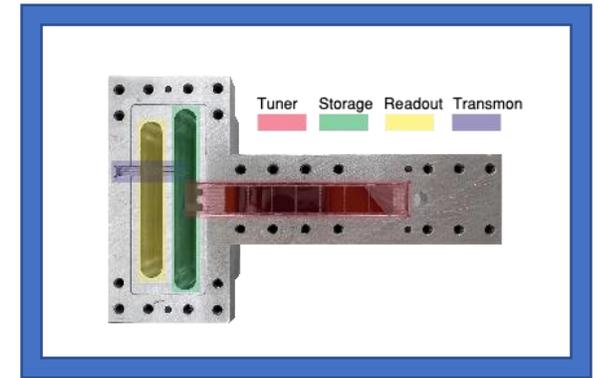


Tunable photon counter



Tunable photonic cavity

Magnetic shielding



State swap

Summary

- SNR for resonant detectors plummets at higher frequencies as cavity volume shrinks and SQL noise increases
- Photon counting measurements have no Heisenberg uncertainty limit and can achieve 10^{-2} SQL.
 - Increases scan speed by $(\text{SNR})^2 = 10^4$.
- Signal can be boosted using stimulated emission with Fock states
 - Potentially increase scan speed by factor $(n+1)$
 - Demonstrated in hidden photon search
- TBD: transport of Fock state via state swap for axion search, demonstration of improvement in scan speed

Paper coming soon!

Stimulated emission of signal photons from dark matter waves

Ankur Agrawal,^{1,2,3,*} Akash V. Dixit,^{1,2,3,†} Tanay Roy,^{1,2,‡} Srivatsan Chakram,^{1,2,4,§}
Kevin He,^{1,2} Ravi K. Naik,⁵ David I. Schuster,^{1,2,6,7} and Aaron Chou⁸

¹*James Franck Institute, University of Chicago, Chicago, Illinois 60637, USA*

²*Department of Physics, University of Chicago, Chicago, Illinois 60637, USA*

³*Kavli Institute for Cosmological Physics, University of Chicago, Chicago, Illinois 60637, USA*

⁴*Department of Physics and Astronomy, Rutgers University, Piscataway, New Jersey 08854, USA*

⁵*Computational Research Division, Lawrence Berkeley National Laboratory, Berkeley, CA, USA*

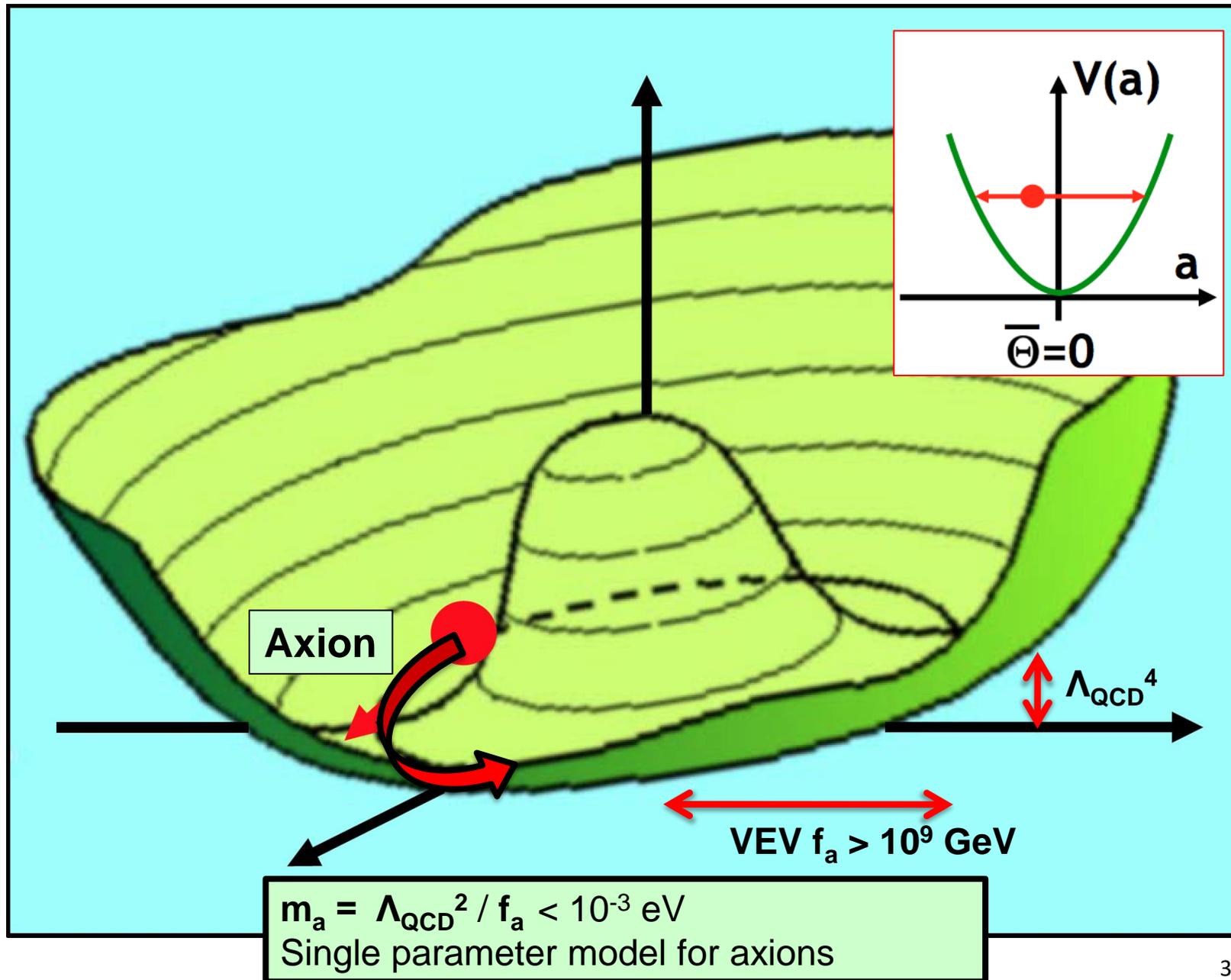
⁶*Pritzker School of Molecular Engineering, University of Chicago, Chicago, Illinois 60637, USA*

⁷*Department of Applied Physics, Stanford University, Stanford, California 94305, USA*

⁸*Fermi National Accelerator Laboratory, Batavia, Illinois 60510, USA*

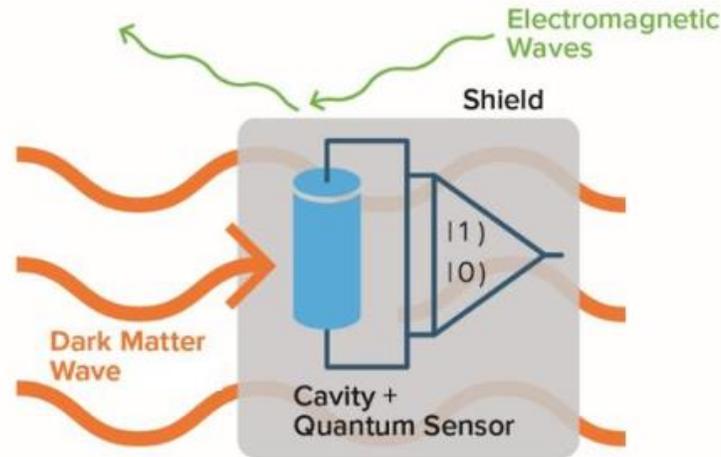
Backup slides

Axion mass = harmonic oscillator frequency



Hmmm... quantum computing platforms look just like dark matter searches:

DOE-OHEP Basic Research Needs white paper, 2018



Sensitive single-quantum devices are operated in a cryostat and/or vacuum system and well-shielded from external disturbances (heat, light, sound) in order to maximize their coherence time.

Impossible to shield from the dark matter – the DM interacts so weakly that it flies right through the walls.

**If your quantum computer crashes, it could be due to dark matter!
... but as consolation, you'll get a Nobel prize anyway for the discovery.**



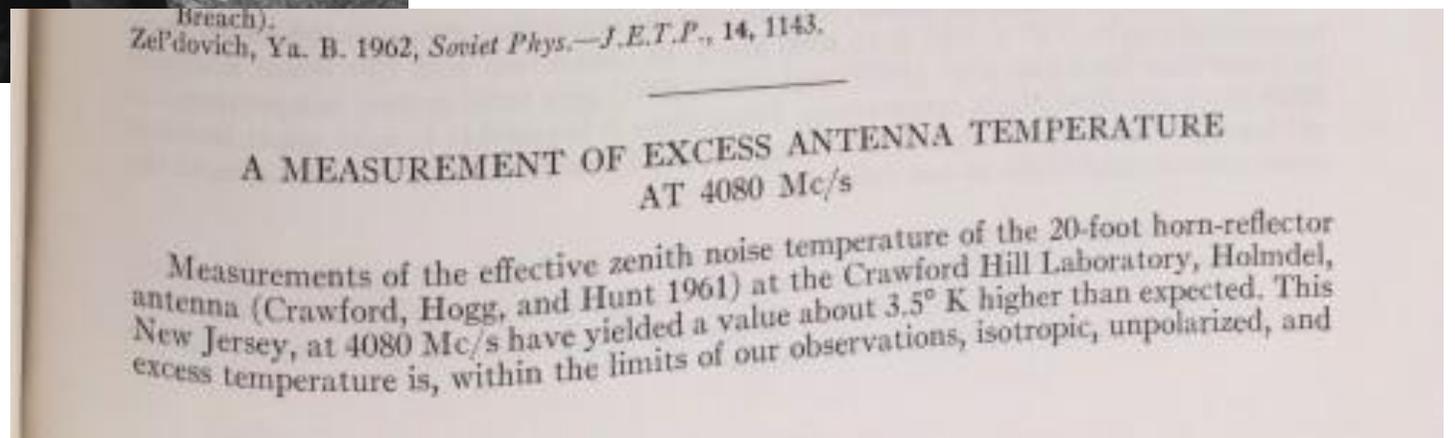
Mystery noise in your experimental apparatus is not necessarily mundane



Penzias and Wilson, 1965:
After chasing away nesting birds with a shotgun,
discovered the cosmic microwave background

- They had no idea that this was first evidence for the big bang theory
- Nobel Prize, 1978

N.B. Discovery of color photography would have to wait until the 1970's.



Need 10^5 seconds to completely convert the axion wave into a photon wave. But only have 10^{-4} s of cavity time...

“Real world values”

$$|\alpha\rangle = 10^{11}$$

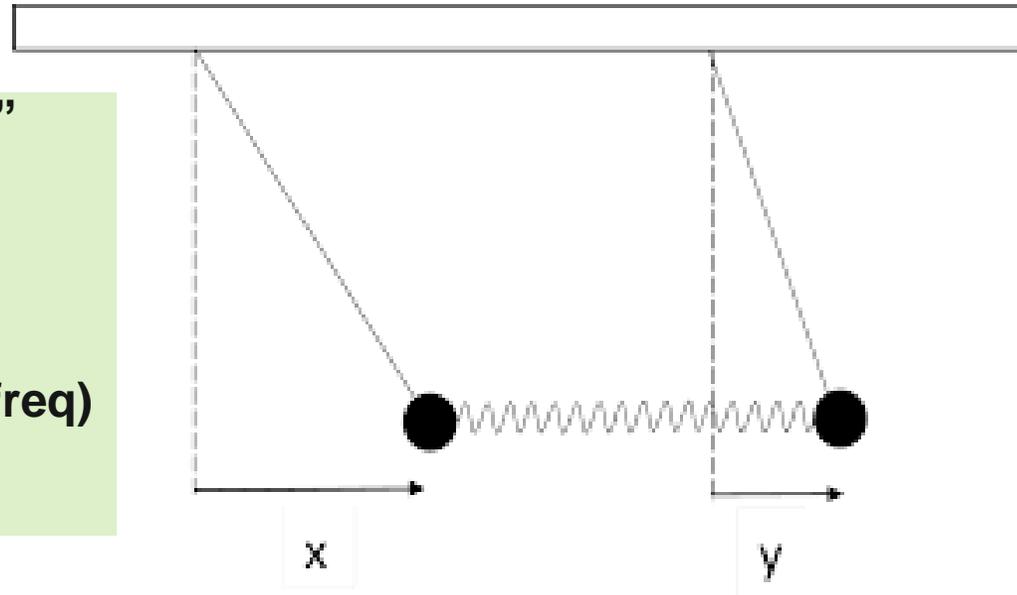
$$\frac{\omega}{2\pi} = 10^{10} \text{ Hz}$$

$$\frac{2g}{2\pi} = 10^{-5} \text{ Hz (beat freq)}$$

$$t_{\text{coherence}} = 10^{-4} \text{ s}$$

Axion = classical wave

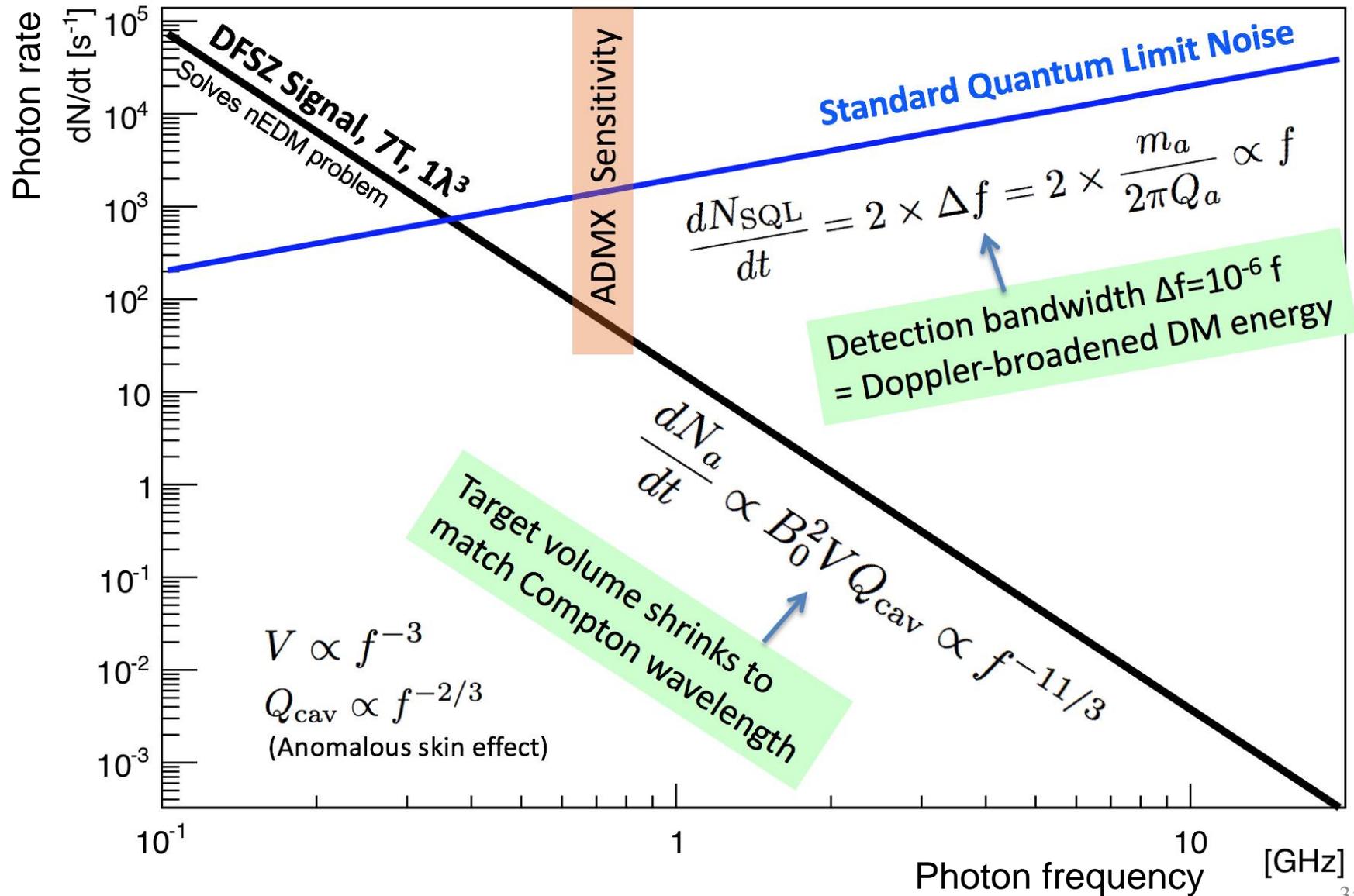
$$\theta(t) = \sqrt{\frac{2\rho_a}{\Lambda_{\text{QCD}}^4}} e^{im_a t}$$



$$H_I = \underbrace{igB\sqrt{\omega V}}_{\text{Beat frequency}} (\theta(t)a^\dagger - \theta(t)^* a)$$

Beat frequency **(derive this!)**

The predicted axion DM signal/noise ratio plummets as the axion mass increases \rightarrow SQL readout is not scalable.



The Glauber displacement operator and coherent states

$$p = i \frac{d}{dx}$$

Generates translations in position

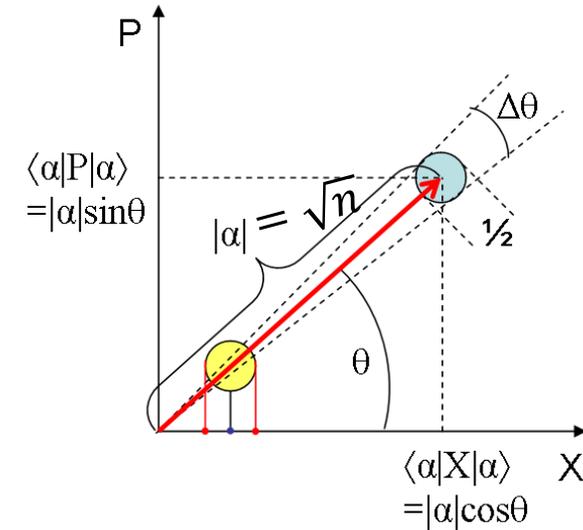
$$x = i \frac{d}{dp}$$

Generates translations in momentum

$$a^\dagger = x + ip$$

Generate translations in an arbitrary direction in x-p phase space

$$a = x - ip$$



Exponentiate differential operator to get finite translation α in complex plane:

$$\hat{D}(\alpha) = \exp(\alpha \hat{a}^\dagger - \alpha^* \hat{a})$$

Phasor of amplitude α is generated as:

$$D(\alpha) |0\rangle = |\alpha\rangle \quad \text{Classical sine wave}$$

This is an eigenstate of the annihilation operator: $a |\alpha\rangle = \alpha |\alpha\rangle$

Prove this!