YVONNE NG (UNIVERSITY OF CALIFORNIA, IRVINE) ON BEHALF OF THE ATLAS COLLABORATION FOR SILAFAE 2018 DARK MATTER SEARCHES WITH THE ATLAS DETECTOR



29nd November, 2018







INDICATIONS FOR DARK MATTER:

- Indications for new, unknown matter outside of SM
- Dark matter candidates offered by SM extensions:
 - Axions
 - Sterile neutrinos
 - Weakly interacting Massive Particles (W.I.M.P)





Galactic rotational curves



Bullet clusters collisions





EVIDENCE FOR DARK MATTER:

- **Evidence** for new, unknown matter outside of SM
- Dark matter candidates offered by SM extensions:
 - Axions
 - Sterile neutrinos

Weakly interacting Massive Particles (W.I.M.P)

YVONNE NG @ UCI







HOW CAN WE FIND DARK MATTER?

Three different approaches to look for DM









DARK MATTER DETECTION THROUGH COLLIDER PRODUCTION

- Or we can... produce them!
- Collider production:
 - Strength: DM/SM interactions and decay channels
 - Weakness: No information about the found "Dark matter" 's lifetime, to be complemented by DD/ID.





CLASSIC COLLIDER DARK MATTER SEARCHES: (A.K.A: MONO-X)

- DM invisible in the ATLAS detector:
 - Look for MET + something
- Mono-X DM search signature :



E.G: Mono-jet, mono-photon, mono-W



MONO-PHOTON AND MONO-JET

Mono-Photon





Challenges: Precise object calibration, SM background prediction and MET measurement.

Fake particles vetoed, various control regions used to measure SM background prediction.





DARK MATTER SEARCH IN THE ATLAS EXPERIMENT

MONO-BOSON (W,Z,Z')



Jet substructure to improve sensitivity







MONO-HIGGS(VISIBLE)











DARK MATTER SEARCHES THROUGH MEDIATOR

- decay back into quarks."
- Can we look for the simplified force mediator instead?

Mono-X analyses



For DM to be made from quarks, it means that there is a process that can

Two Body Decays





DARK MATTER SEARCHES THROUGH MEDIATOR

decay back into quarks."



Look for a bump on top of SM background!

For DM to be made from quarks, it means that there is a process that can

DIJET AND DI-B JETS

Little is known about the force mediator signatures are preferred.



Little is known about the force mediators, based on different assumptions, different





TRIGGER LEVEL ANALYSIS (TLA)





TRIGGER LEVEL ANALYSIS (TLA) RESULTS

- (Phys. Rev. Lett. 121 (2018)
 081801)
- Covering a lower mass range than dijet





CAN WE GO EVEN LOWER IN THE Z' SEARCH MASS?

Dijet/TLA:



2. Other jet $pT \approx trigger$



Average net PT= 0

1. A photon at trigger





DIJET-ISR

Dijet-ISR Resolved (ATLAS-CONF-2016-070)









IS THAT ALL? OTHER DARK MATTER SEARCHES IN THE ATLAS DETECTOR

Other Mono-X signatures:

- Mono-H to invisible particles
- DM with heavy flavor quarks
- Other two-body decays:
 - di-leptons
- Supersymmetry (SUSY)
 - Many SUSY searches that looks for grad DM.

Many SUSY searches that looks for gravitino, neutralino etc are natural candidates of



DARK MATTER SEARCH IN THE ATLAS EXPERIMENT

DARK MATTER SUMMARY



Excluded the Z' Mass Range from 50GeV to 5000GeV+

Extra exclusion phase space from the mono-X and dijet searchess



Dijet (s = 13 TeV, 37.0 fb⁻¹ Dijet TLA VS = 13 TeV, 29.3 fb⁻¹ Dijet + ISR 15 = 13 TeV, 15.5 fb

s = 13 TeV 36.1 fb⁻¹

s = 13 TeV, 36.1 fb⁻¹

E_x^{miss}+y **s** = 13 TeV, 36.1 fp⁻¹ Eur. Phys. J. C 77 (2017) 393 E____+jet 13 TeV, 36.1 fb E_T^{miss}+Z(II) **[**S = 13 TeV, 36.1 fb⁻¹ E_T +V(had) 45 = 13 TeV, 36.1 fb

DM SEARCHES IN ATLAS

ATLAS DM searches makes model assumptions, it is complementary to ID + DD in the overall DM search.

10 σ_{SI} (χ-nucleon) [cm²]





CONCLUSION

- No signs of DM yet! But in ATLAS, we have:
 - More data than ever
 - Continuous advancements in our search techniques
 - Expect to gain sensitivity in all variable phase space
- Future of DM searches in ATLAS:
 - Long lived particle signature searches
 - Moving towards more detailed models



"We're quietly confident that it smells of cinnamon."



DARK MATTER SEARCH IN THE ATLAS EXPERIMENT





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DARK MATTER



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DARK MATTER SEARCH IN THE ATLAS EXPERIMENT







MONO-HIGGS(INVISIBLE)

Mono-Higgs(->Invisible Decay)

(arXiv:1807.11471/HIGG-2018-54)



