



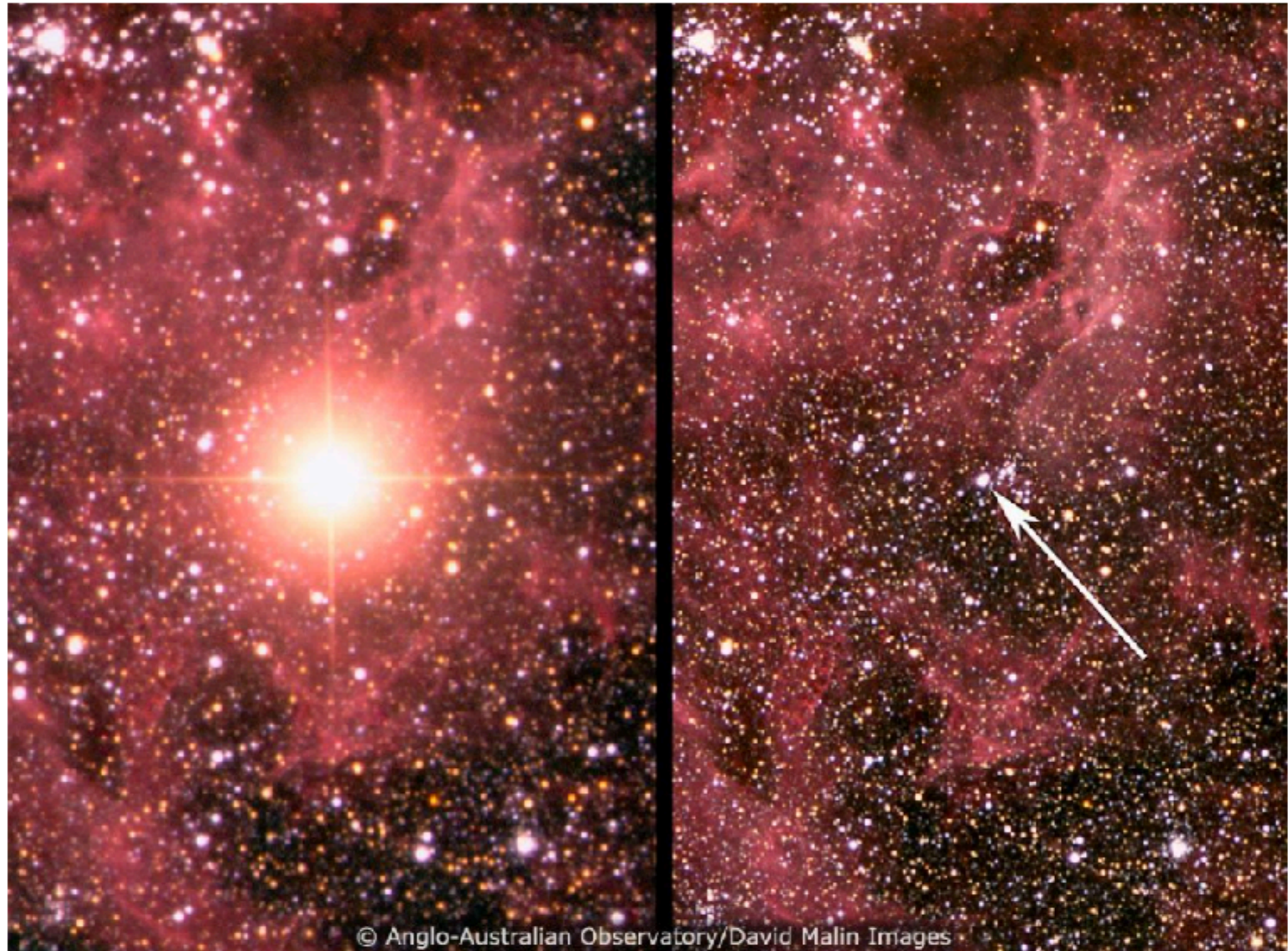
# **Possible Neutrino Signature of Hadron-quark Phase Transition in Failing Core-collapse Supernovae**

**Shuai Zha and Evan O'Connor**  
Stockholm University  
20201123



# Supernova: an example

SN 1987A at Large Magellanic Cloud, ~160,000 light years away

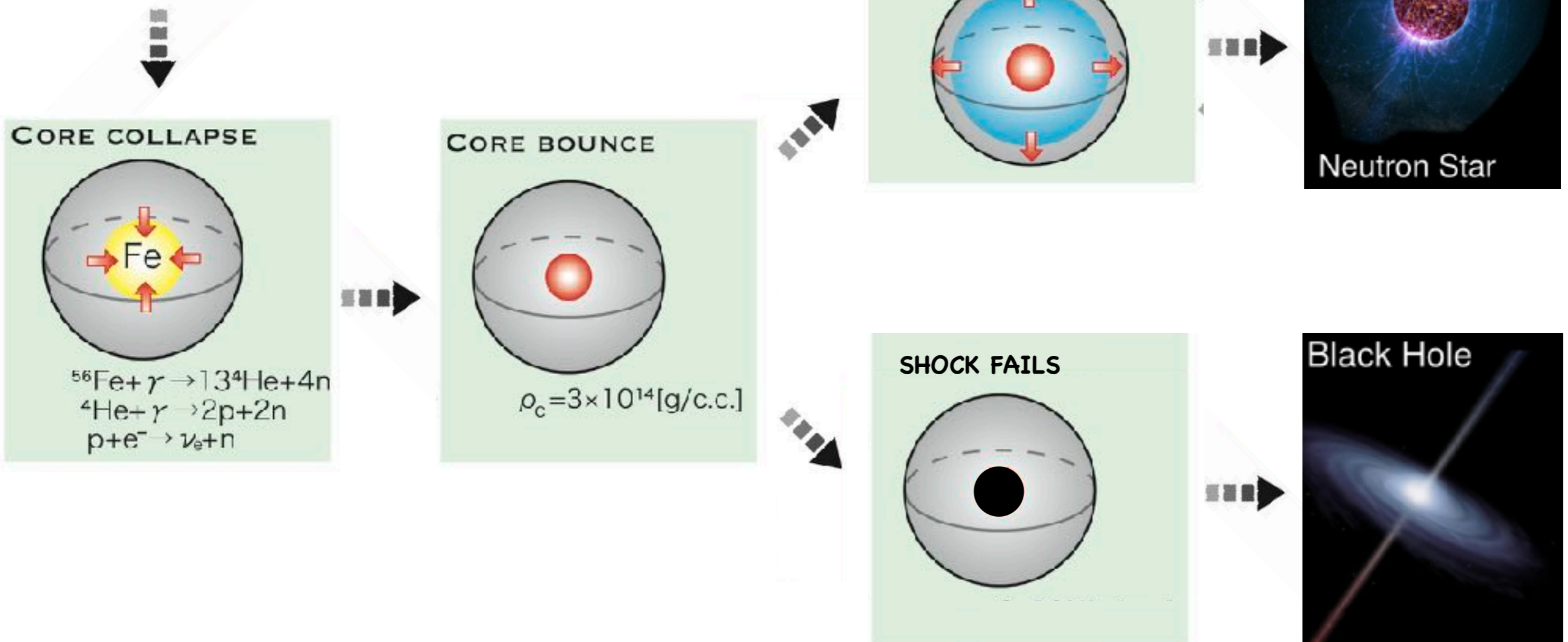


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# Core-collapse Supernova (CCSN)

Massive star  
 $\sim 8-100 M_{\odot}$



Failing CCSN

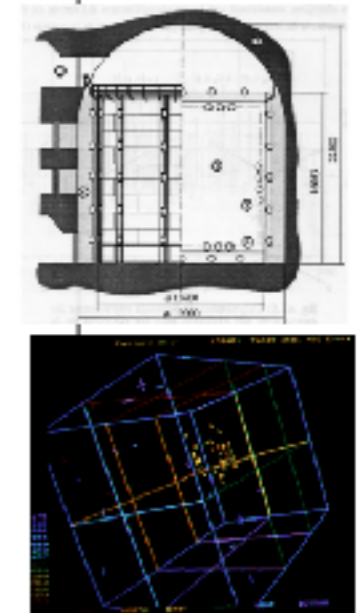
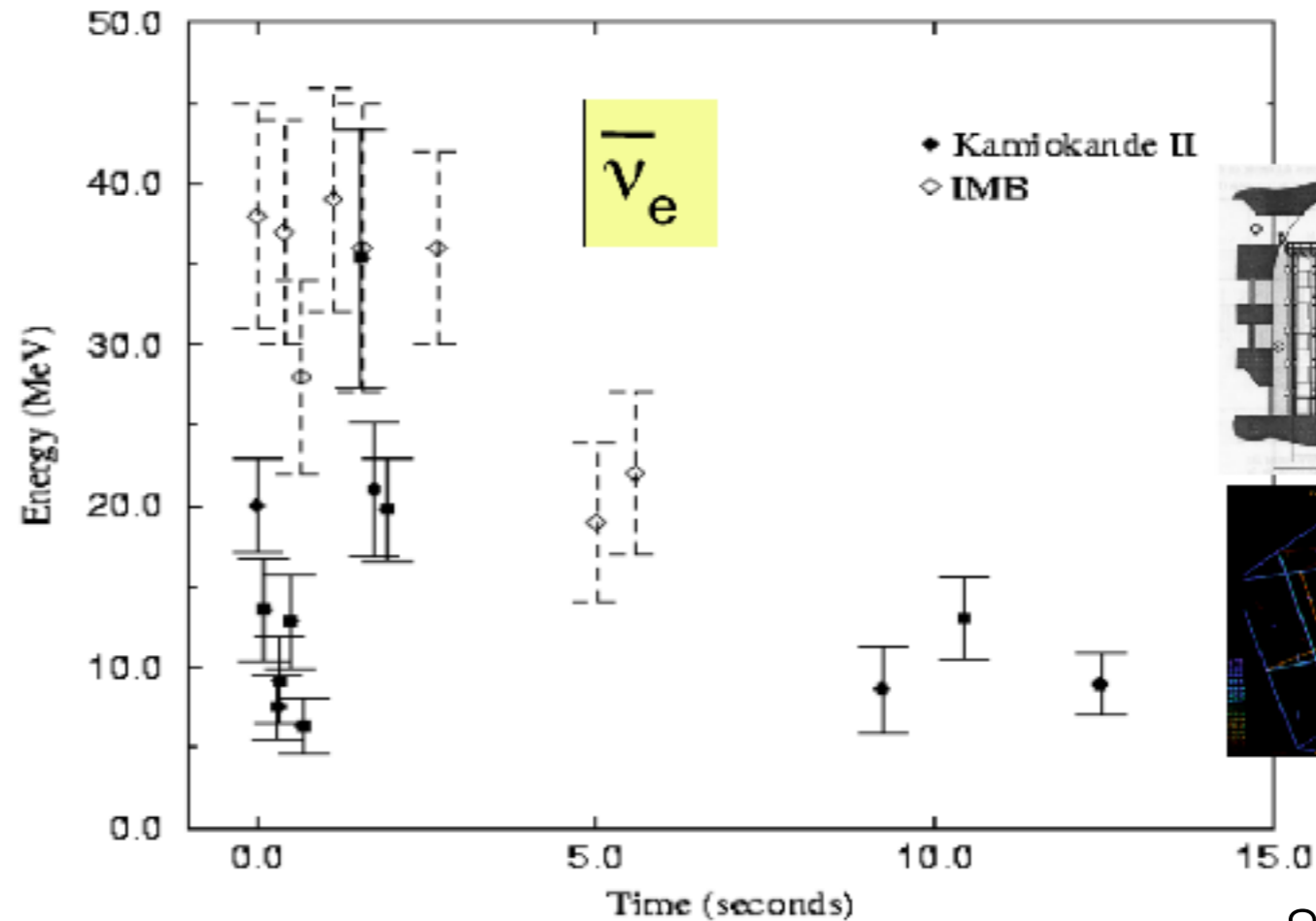
# Neutrino signal of CCSN

A stellar collapse releases  $\sim 10^{53}$  erg gravitational energy

- $\sim 99\%$  is carried away by neutrinos
- SN explosion energy  $\sim 10^{51}$  erg



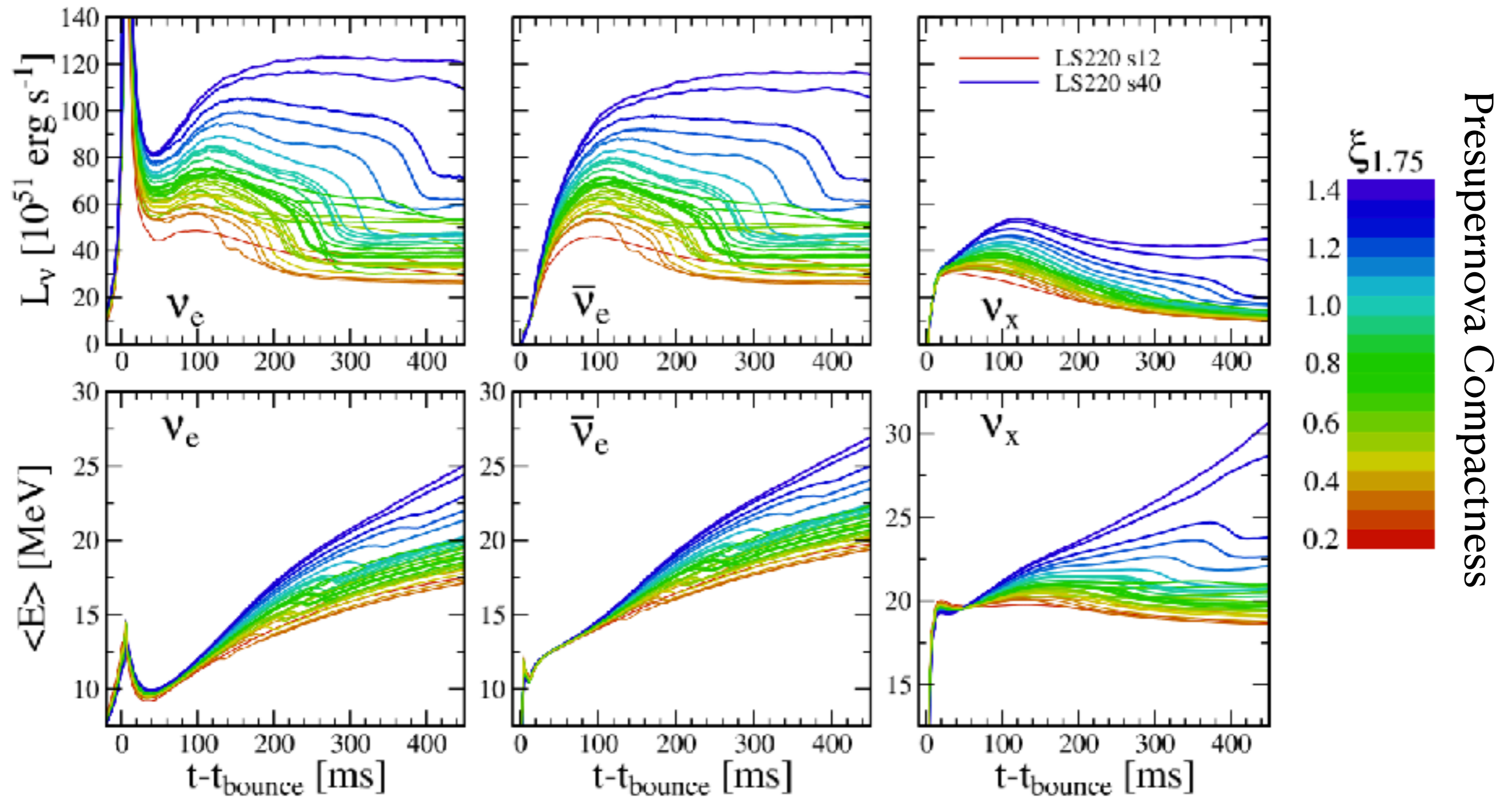
SN1987A (ESA/Hubble)



Scholberg 2017

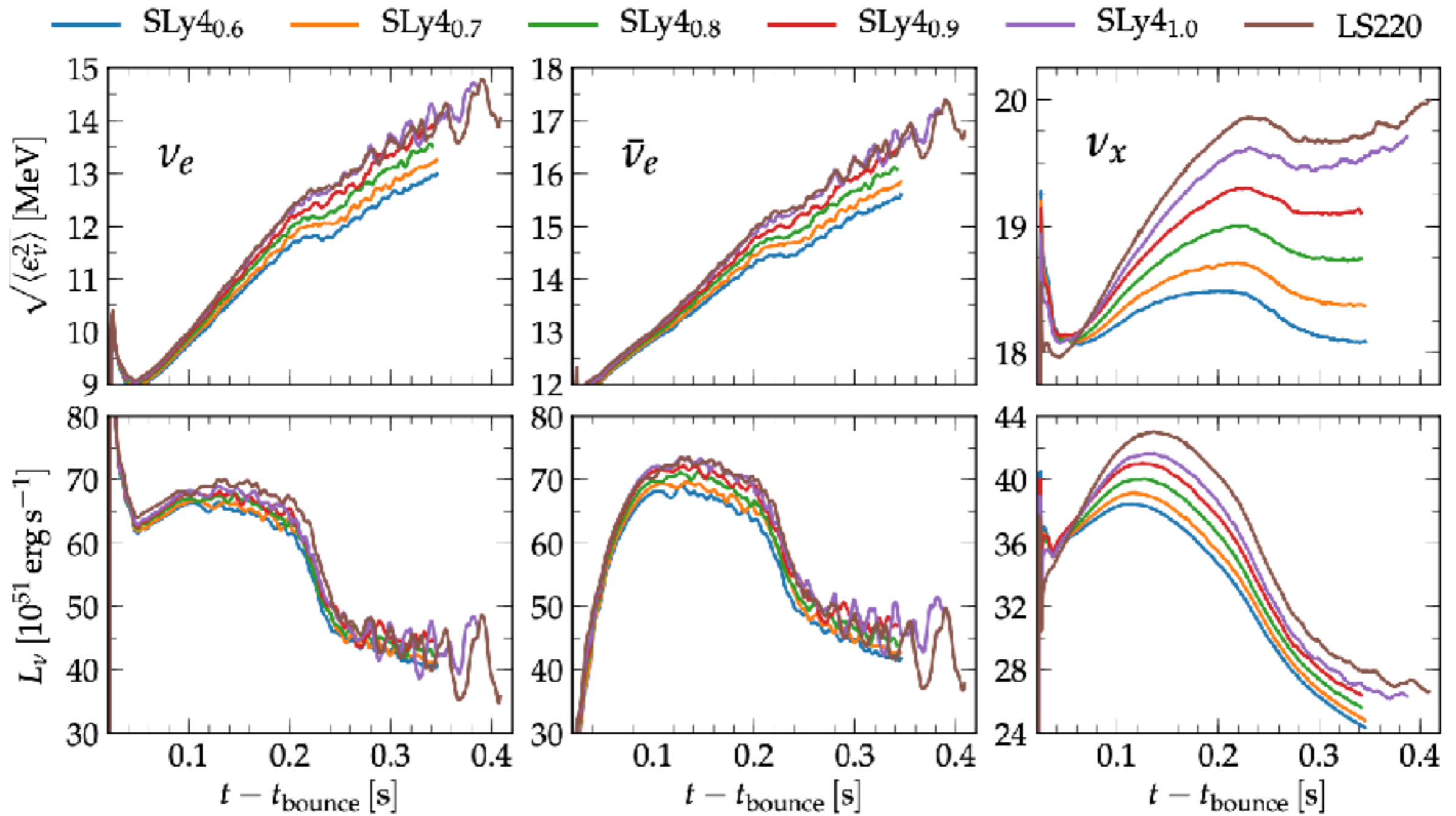
- Confirmed baseline model, i.e., SN II  $\rightarrow$  core-collapse to a proto-NS
- Better statistics needed to probe progenitor properties, validate explosion mechanism, constrain neutrino properties, etc.

## Dependence on progenitor star

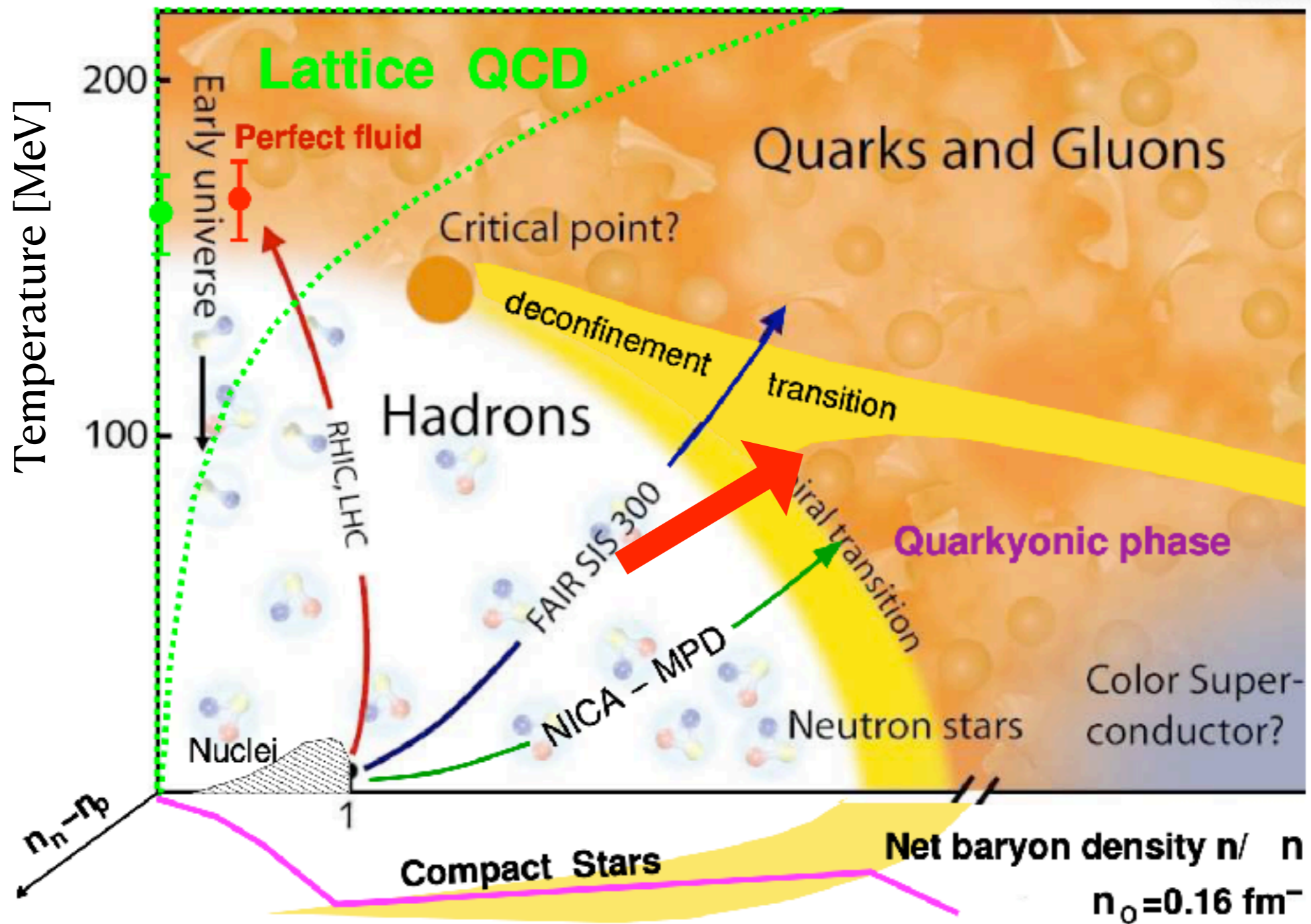




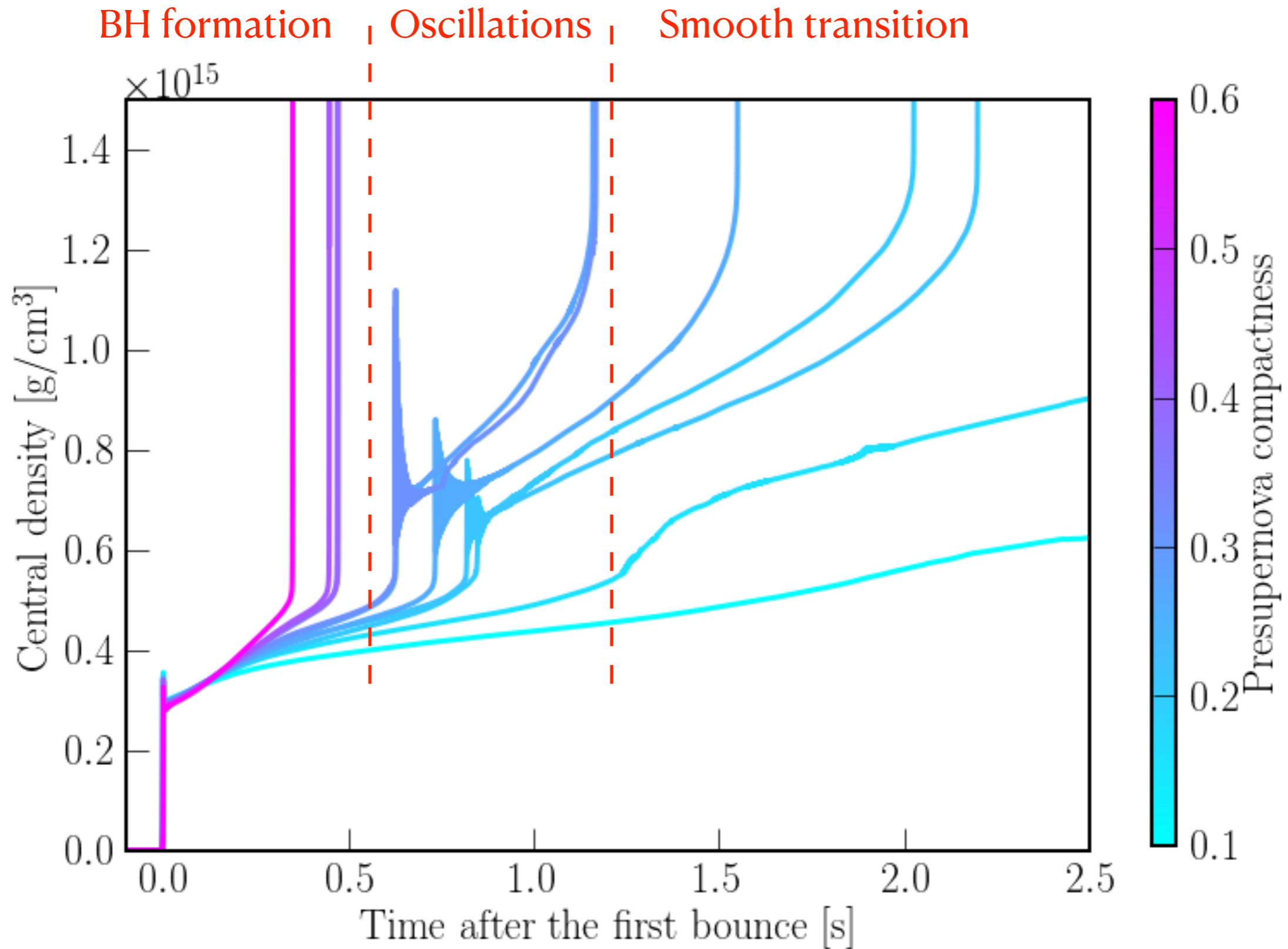
# Dependence on Equation of State (EoS)



# Hadron-quark phase transition



# Consequence of phase transition in CCSN

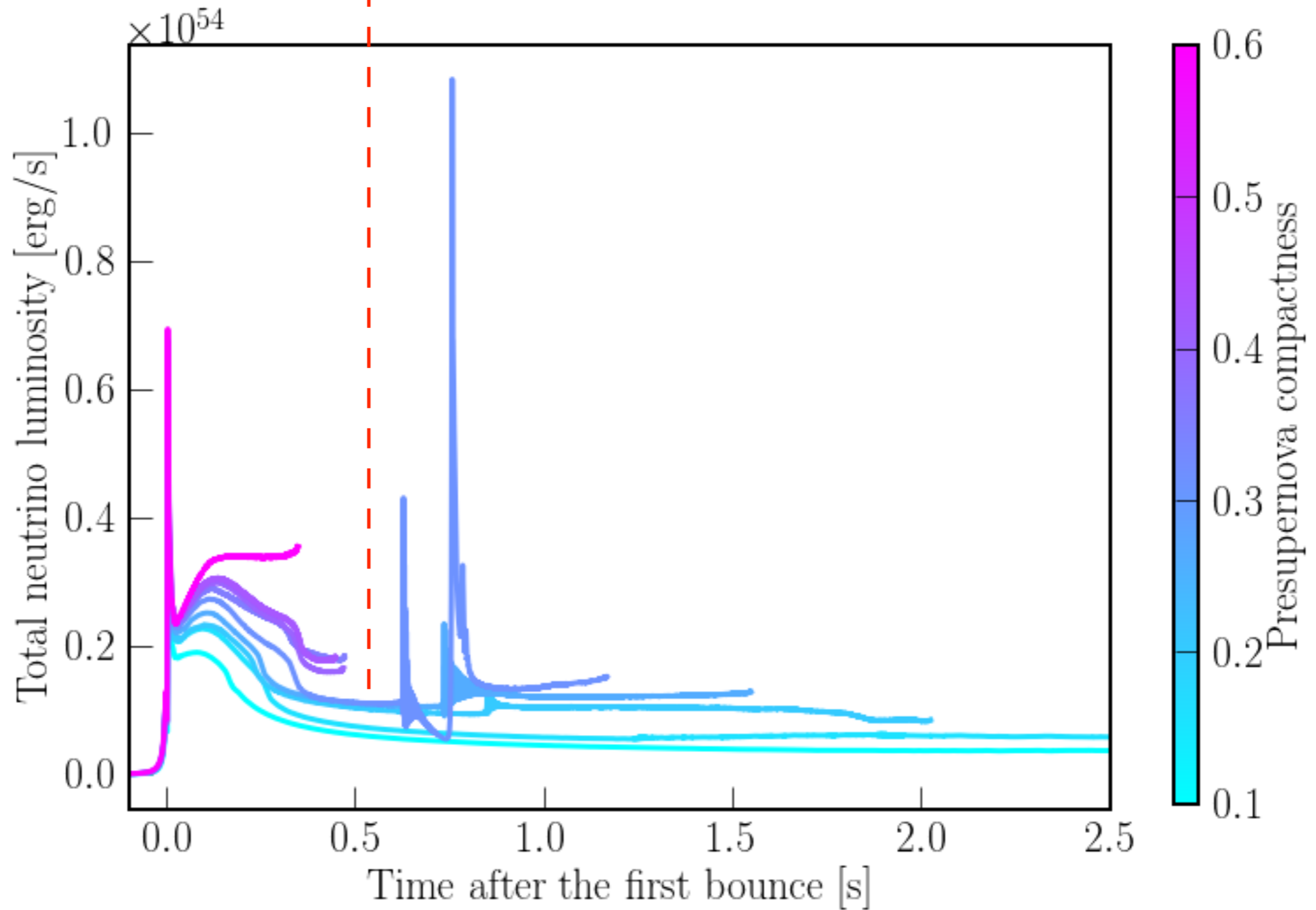




# Neutrino signal

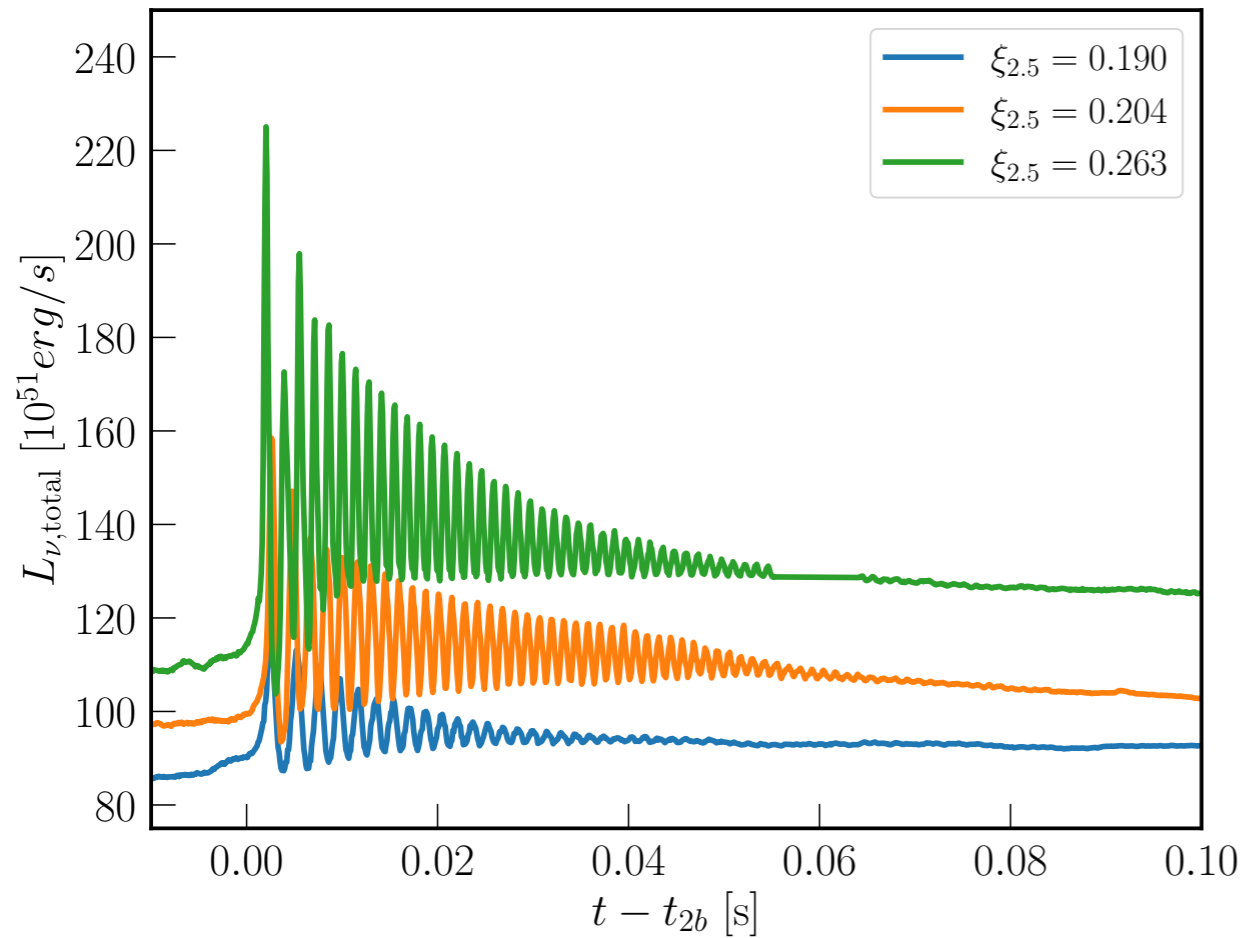
Shut off at BH formation

Oscillations after 2nd collapse



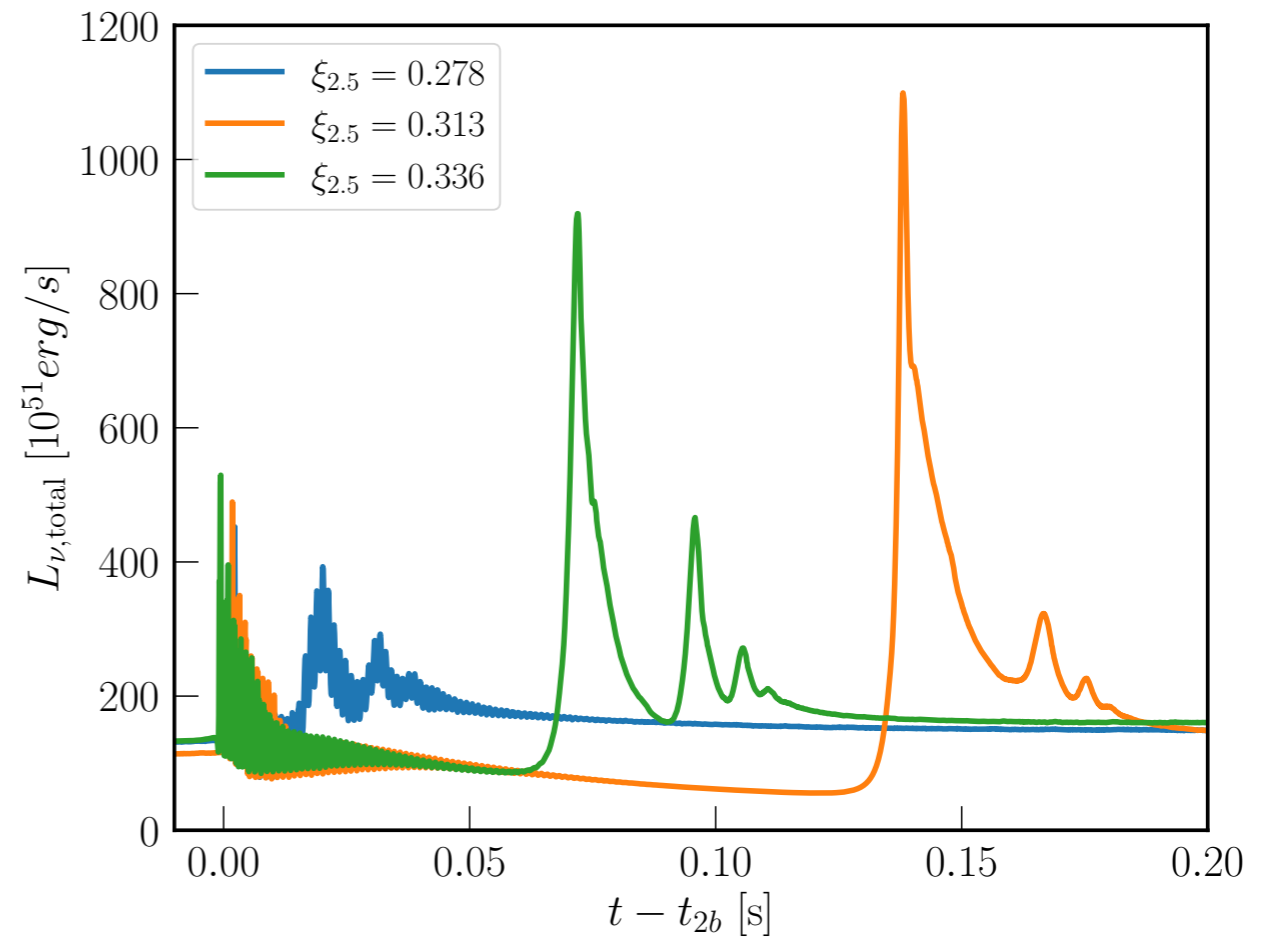
# Two classes

Less compact



~ 1 ms period, last for ~50 ms

More compact



+ secondary bursts

# Summary

- Hadron-quark phase transition leads to collapse of protoneutron star in core-collapse supernovae.
- The collapse and oscillations of protoneutron star emit oscillatory neutrino signal with  $\sim 1$  ms period.

Thank you!  
Tack!