

Complementarity in Dark Matter Searches

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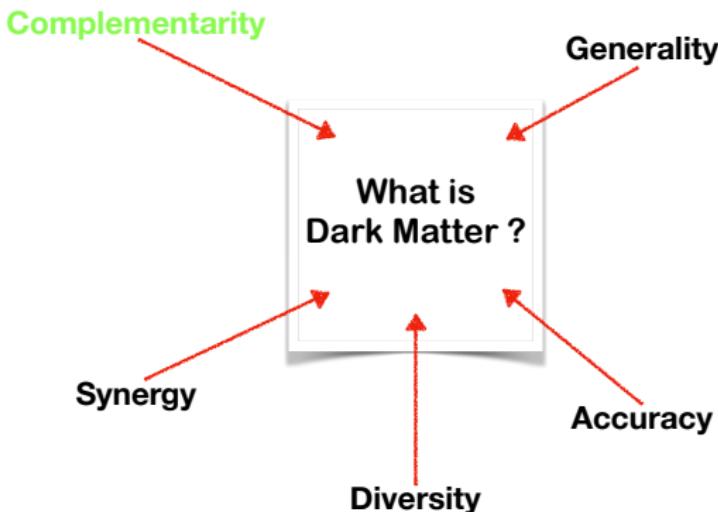
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Overview



Example: DM Direct Detection

Generality ⇒ Model independent data interpretation via EFT

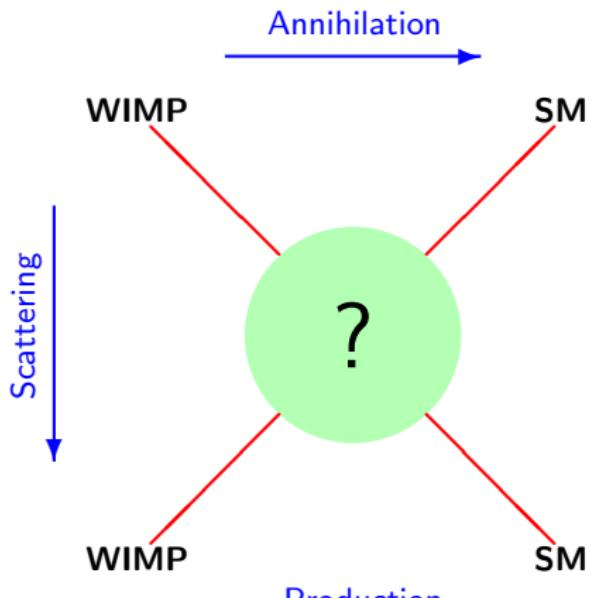
Accuracy ⇒ Astrophysical and nuclear physics uncertainties

Diversity ⇒ Alternatives to WIMP DM (e.g. Light Dark Matter)

Synergy ⇒ Collaboration with nuclear and solid state physicists

Complementarity ⇒ Identifying experimental inputs constraining DM properties unaccessible to direct detection

Complementarity in WIMP DM searches



- The LHC searches for missing transverse momentum in proton collisions
 - Direct detection experiments search for DM-nucleus scattering events
 - Indirect detection experiments search for DM pair annihilation products
- WIMP spin
— Hierarchy of constraints

Theoretical framework

- I will focus on a general class of simplified models for spin ≤ 1 DM interacting with quarks
J. B. Dent, L. M. Krauss, J. L. Newstead and S. Sabharwal, Phys. Rev. D **92**, no. 6, 063515 (2015)
S. Baum, R. Catena, J. Conrad, K. Freese and M. B. Krauss, Phys. Rev. D **97** (2018) no.8, 083002
R. Catena, J. Conrad and M. B. Krauss, Phys. Rev. D **97** (2018) no.10, 103002
- Within this framework, models can be classified in terms of WIMP and mediator spin
- Each model is characterised by 4 parameters: two masses and two coupling constants
- Each model can be mapped onto a (linear combination) of DM-nucleon interaction operators
- These operators define the non relativistic effective theory of DM-nucleon interactions (NRET)

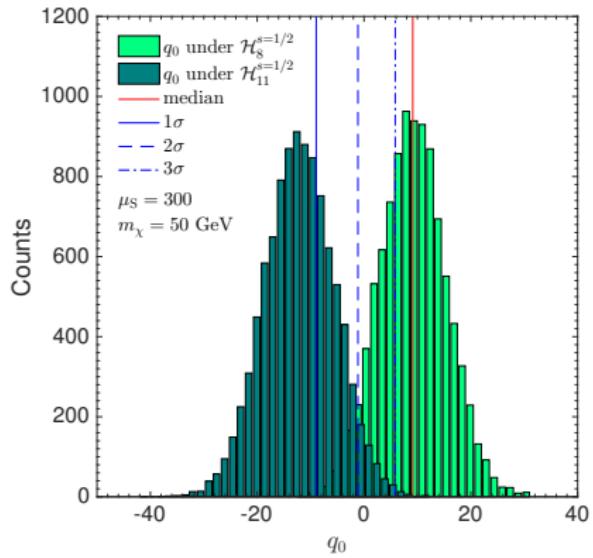
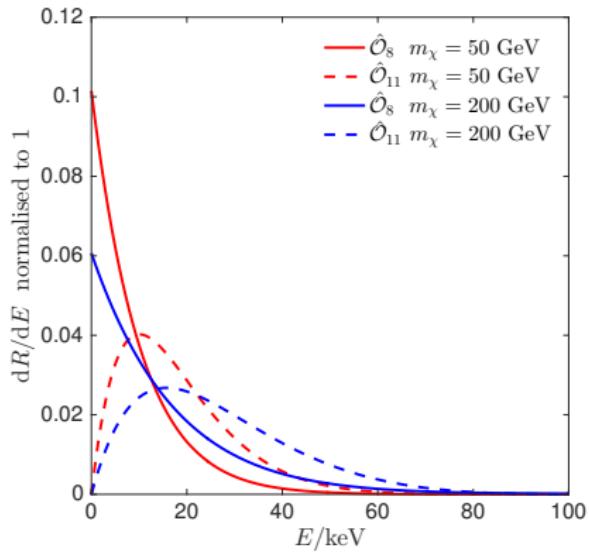
Direct Detection – LHC Complementarity

Research question

- Direct detection experiments and the LHC are complementary in probing DM models:
 - Direct detection probes coherently enhanced DM-nucleus scattering cross sections
 - The LHC probes models with momentum or velocity suppressed scattering cross sections
- Can we exploit this type of complementarity to gain insight into the DM particle spin?
- Yes, if an experiment like XENONnT will be able to detection $\mathcal{O}(100)$ signal events

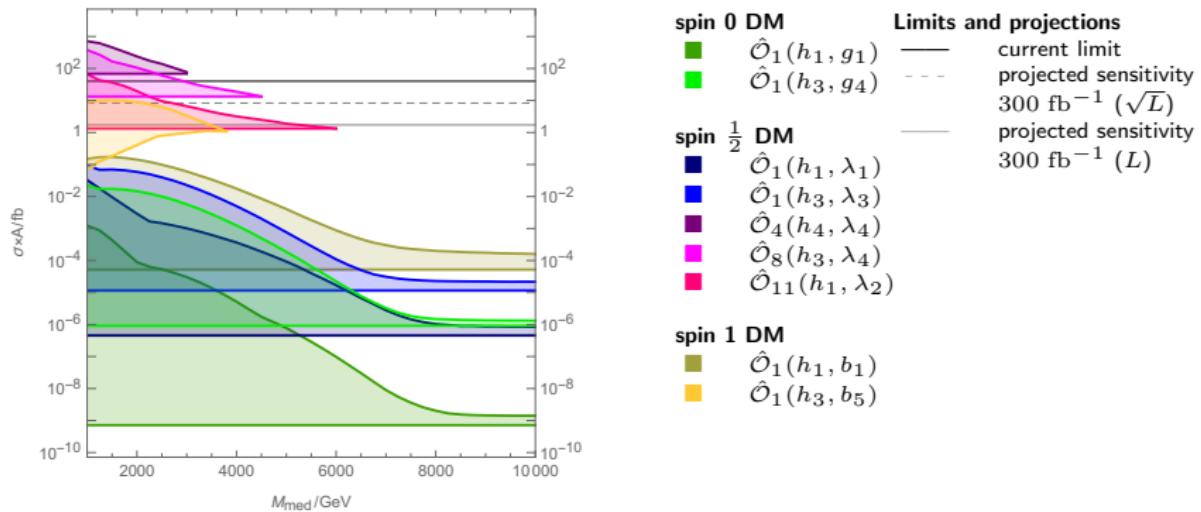
DM particle spin combining direct detection and LHC (I)

S. Baum, R. Catena, J. Conrad, K. Freese and M. B. Krauss, Phys. Rev. D **97** (2018) no.8, 083002
R. Catena, J. Conrad and M. B. Krauss, Phys. Rev. D **97** (2018) no.10, 103002



DM particle spin combining direct detection and LHC (II)

S. Baum, R. Catena, J. Conrad, K. Freese and M. B. Krauss, Phys. Rev. D **97** (2018) no.8, 083002
R. Catena, J. Conrad and M. B. Krauss, Phys. Rev. D **97** (2018) no.10, 103002



Other approaches to WIMP spin identification

- Large exposure & LHC Run 3 (this talk)

S. Baum, R. Catena, J. Conrad, K. Freese and M. B. Krauss, Phys. Rev. D **97** (2018) no.8, 083002
R. Catena, J. Conrad and M. B. Krauss, Phys. Rev. D **97** (2018) no.10, 103002

- Large exposure & directional information

R. Catena, J. Conrad, C. Dring, A. D. Ferella and M. B. Krauss, Phys. Rev. D **97** (2018) no.2, 023007

- Large exposure & polarised target materials

R. Catena, K. Fridell and V. Zema, arXiv:1810.01515 [hep-ph].

Direct Detection – Neutrino Telescopes Complementarity

Research question

- Direct detection experiments and neutrino telescopes are complementary in probing DM models:
 - Direct detection probes coherently enhanced DM-nucleus scattering cross sections
 - Neutrino telescopes probe spin-dependent DM-proton scattering cross sections
- Is there any other type of interaction that neutrino telescopes can probe better than direct detection?
- Yes, inelastic DM-nucleus interactions, for large mass splittings between incoming and outgoing DM particle

Kinematics of inelastic DM-nucleus scattering

- When there is a mass splitting δ between incoming and outgoing DM particle, one has

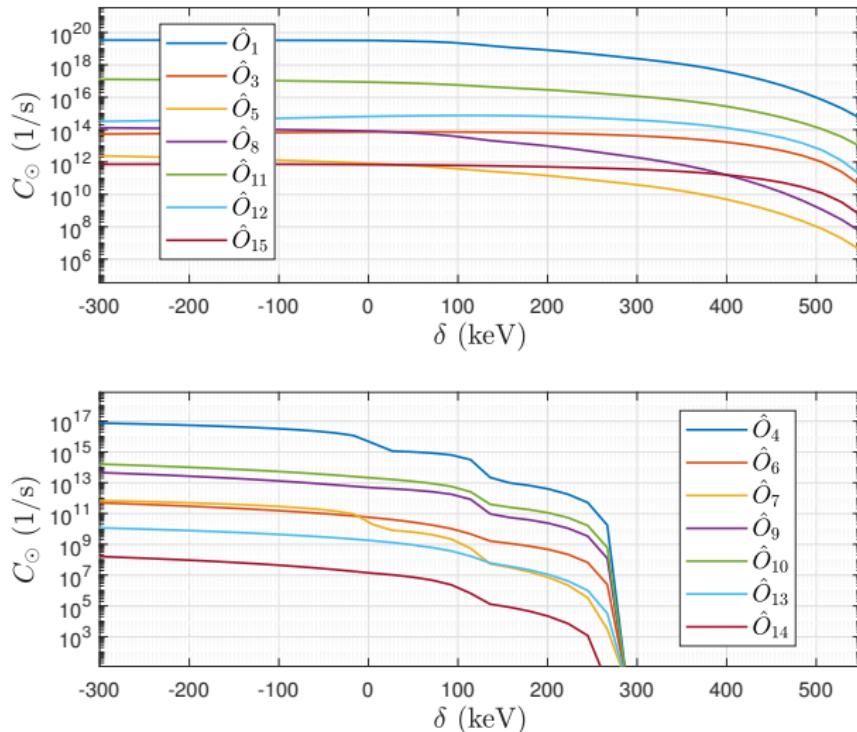
$$E_R^\pm = \frac{\mu^2}{m_T} w^2 \left(1 \pm \sqrt{1 - \frac{2\delta}{\mu w^2}} \right) - \frac{\mu}{m_T} \delta$$
$$w \geq w_{\min} \equiv \Re \sqrt{2\delta/\mu}$$

- Furthermore, when the DM particle is heavier than the target nucleus, one finds

$$E_R^- \simeq \delta; \quad w_{\min} \simeq \Re \sqrt{2\delta/m_T}$$

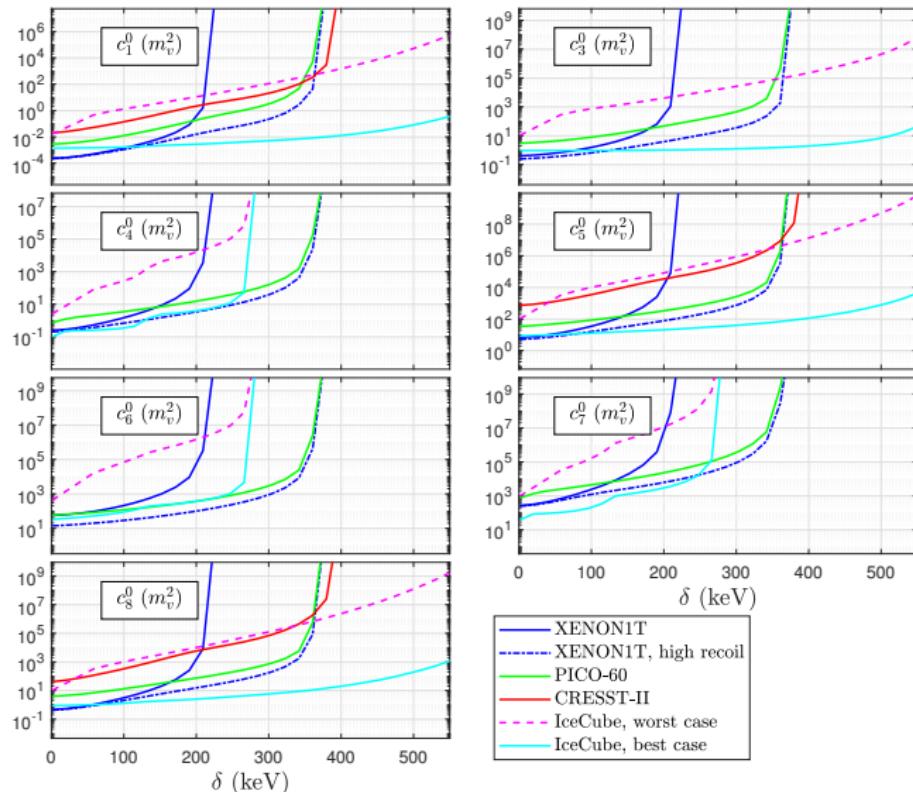
Solar capture rates

R. Catena and F. Hellström, arXiv:1808.08082



Exclusion limits from IceCube and direct detection

R. Catena and F. Hellström, arXiv:1808.08082





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Conclusion

