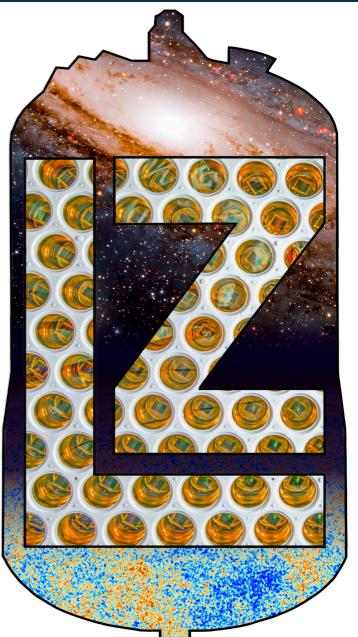




Status and Results from the LUX-ZEPLIN Experiment

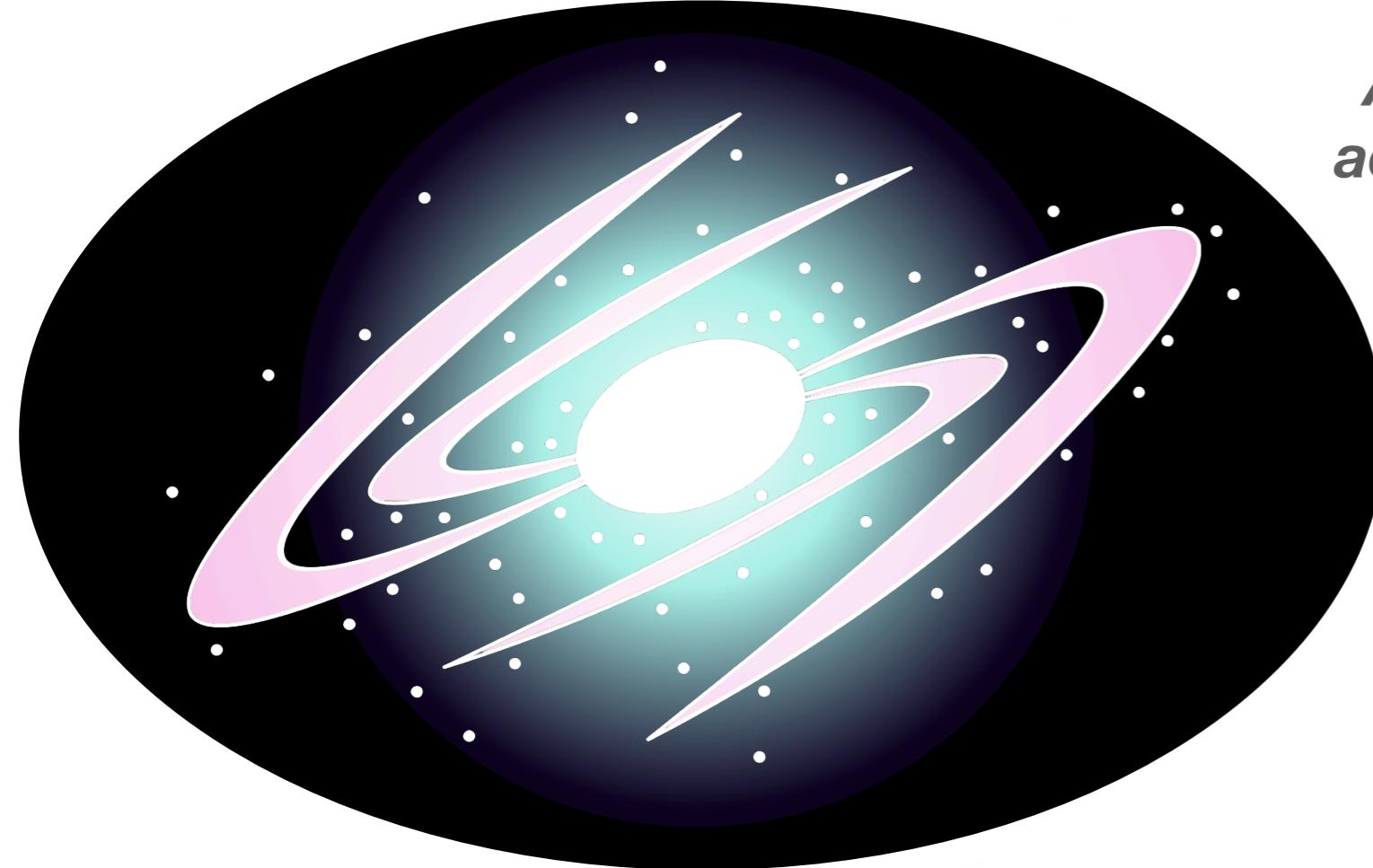
Joe McLaughlin
University College London
On behalf of the LZ Collaboration
Cosmology 2023



Contents

The LZ Experiment
First Results with 60 Live Days
Low Energy ER Searches
In The Pipeline
Conclusion

Some Context



*Assuming dark matter made up of
Weakly Interacting Massive Particles (WIMPs)*

- GeV–TeV mass scale
- Extremely weak coupling to baryonic matter
- e.g. SUSY neutralino, Kaluza-Klein dark matter

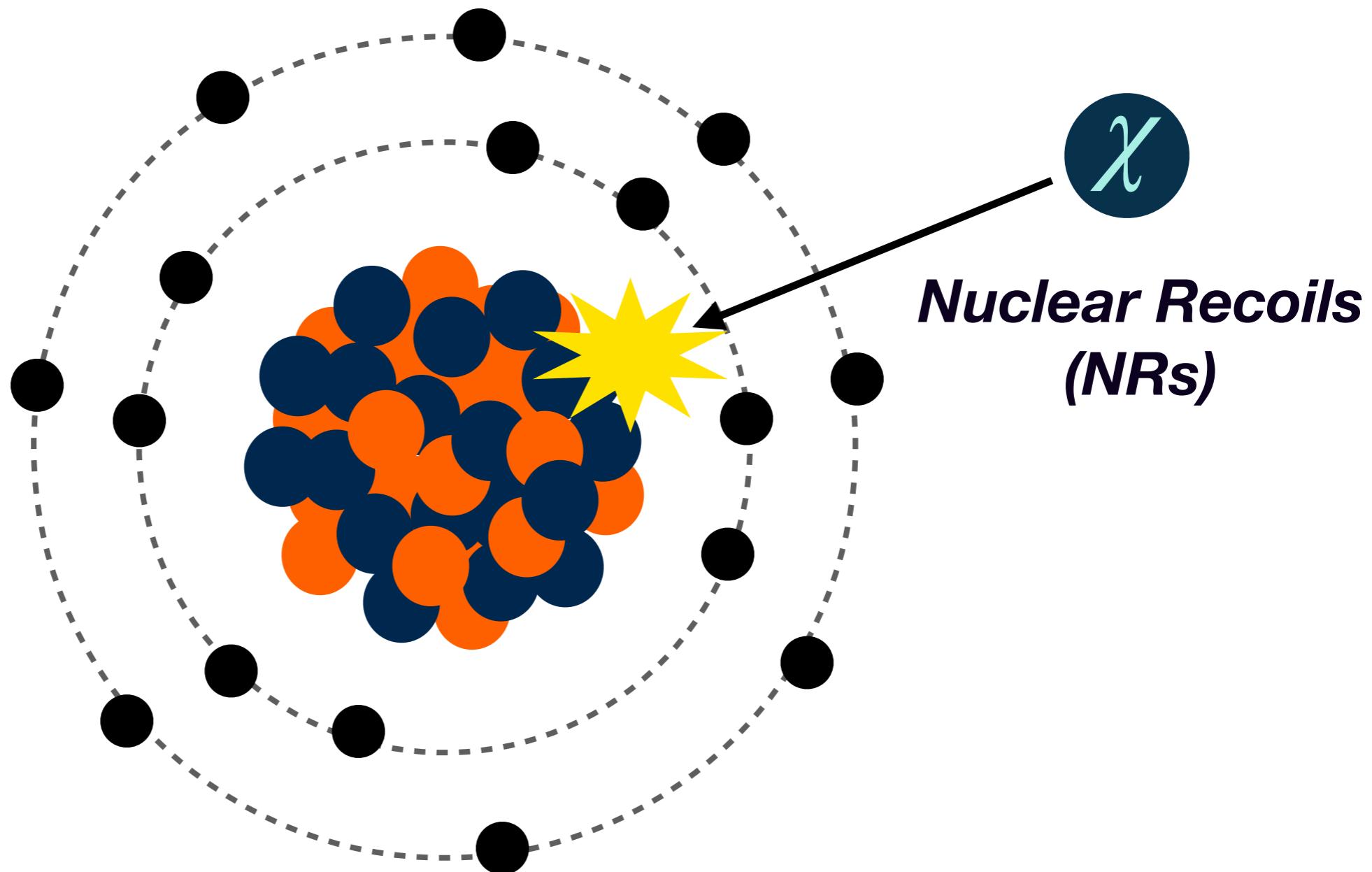
*Assuming dark matter is distributed
according to the Standard Halo Model*

- Isothermal sphere of DM, $\rho \propto r^{-2}$
- Local density $\rho_0 \sim 0.3 \text{ GeV/cm}^3$
- Truncated Maxwell-Boltzmann velocity distribution
- Characteristic velocity $v_0 = 220 \text{ km/s} \rightarrow \text{non-relativistic!}$



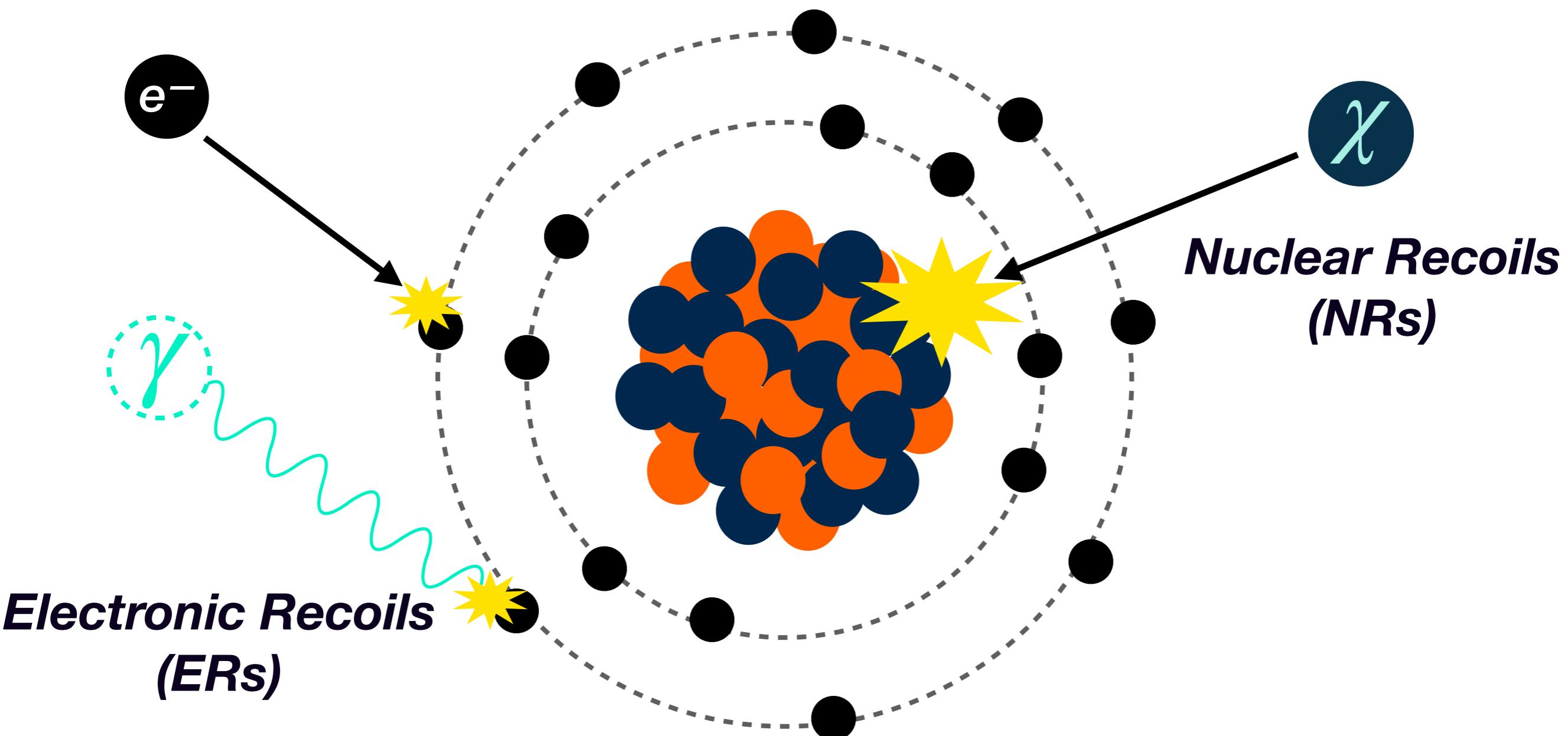
Some Context

Assuming dark matter can interact with baryonic matter, virtually always via elastic collisions with nuclei



Some Context

Assuming dark matter can interact with baryonic matter, virtually always via elastic collisions with nuclei



Lighter particles, such as electrons or photons, kinematically limited to scatter with atomic electrons

The LZ Experiment

LUX-ZEPLIN (LZ) is...

- 250 scientists, engineers, and technical staff
- 37 institutions
- 5 countries (US, UK, Portugal, Korea, Australia)
- Follow us! [@lzdarkmatter](https://twitter.com/lzdarkmatter) X  
- Our website: <https://lz.lbl.gov/>



The LZ Experiment

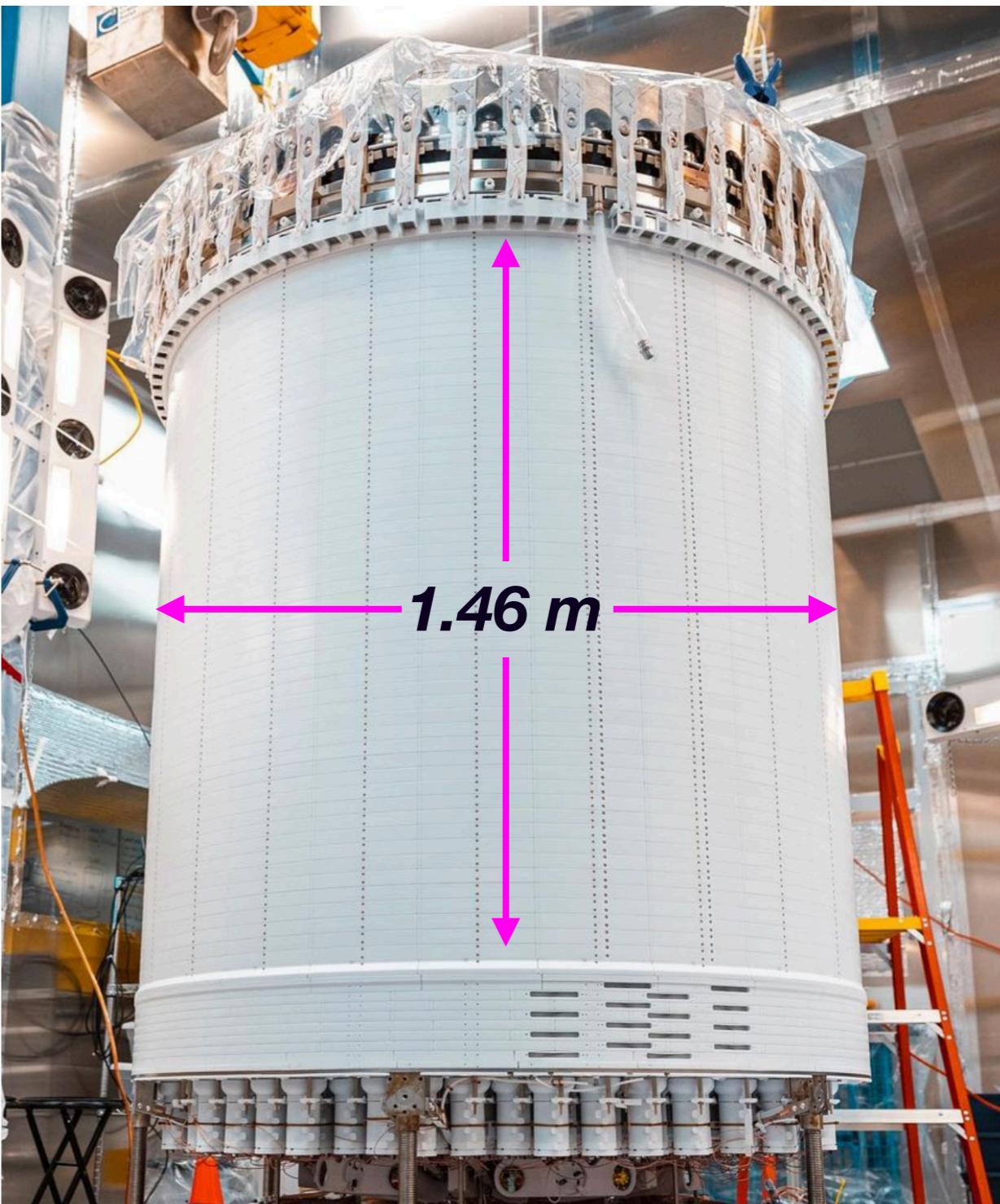
LUX-ZEPLIN (LZ) is...

A dual-phase time projection chamber (TPC) filled with liquid xenon (LXe) as a WIMP scattering target

10 tonnes total mass

7 tonnes active mass

5.5 tonnes fiducial mass



The LZ Experiment

LUX-ZEPLIN (LZ) is...

A dual-phase time projection chamber (TPC) filled with liquid xenon (LXe) as a WIMP scattering target

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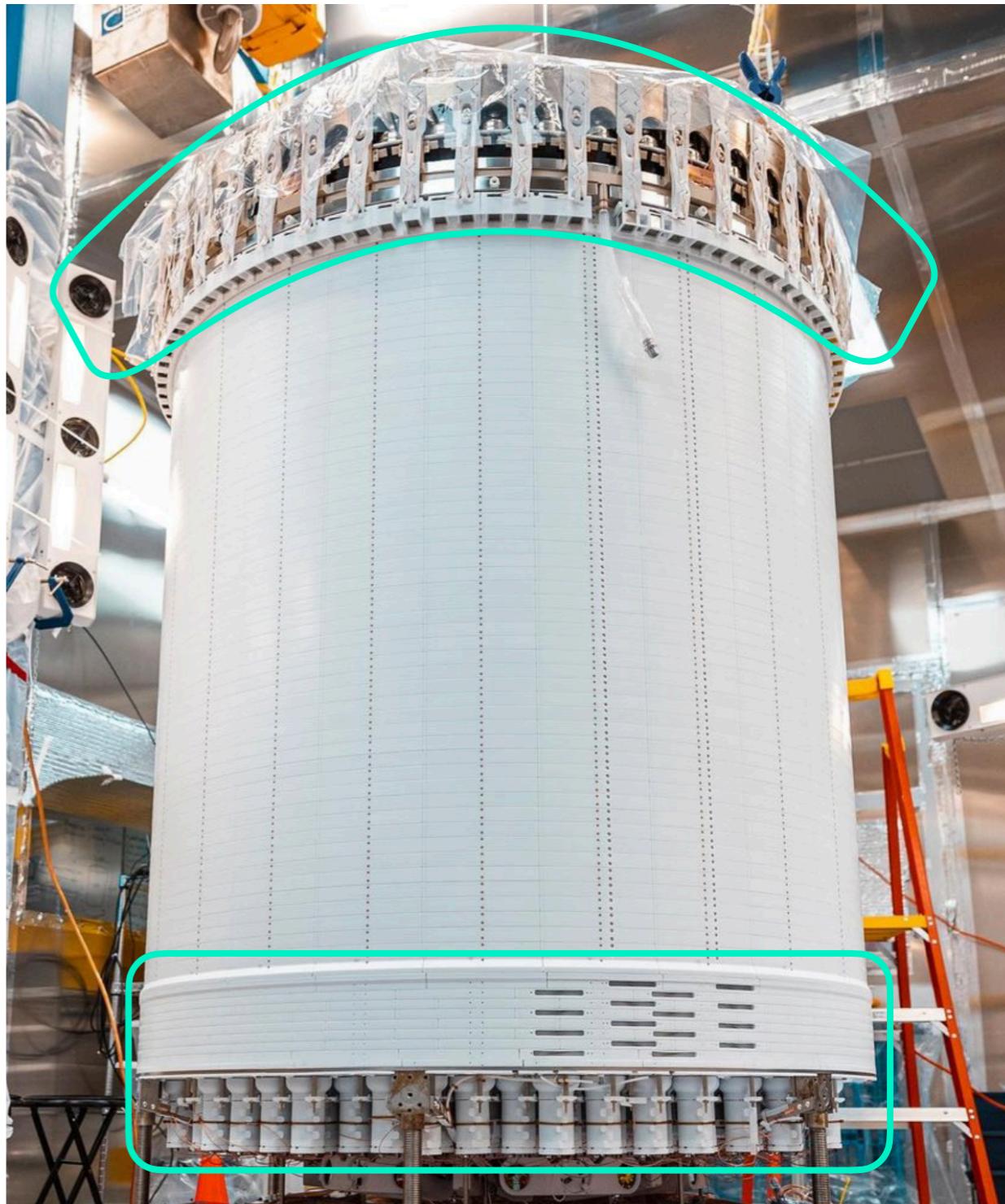
7 tonnes active mass

5.5 tonnes fiducial mass

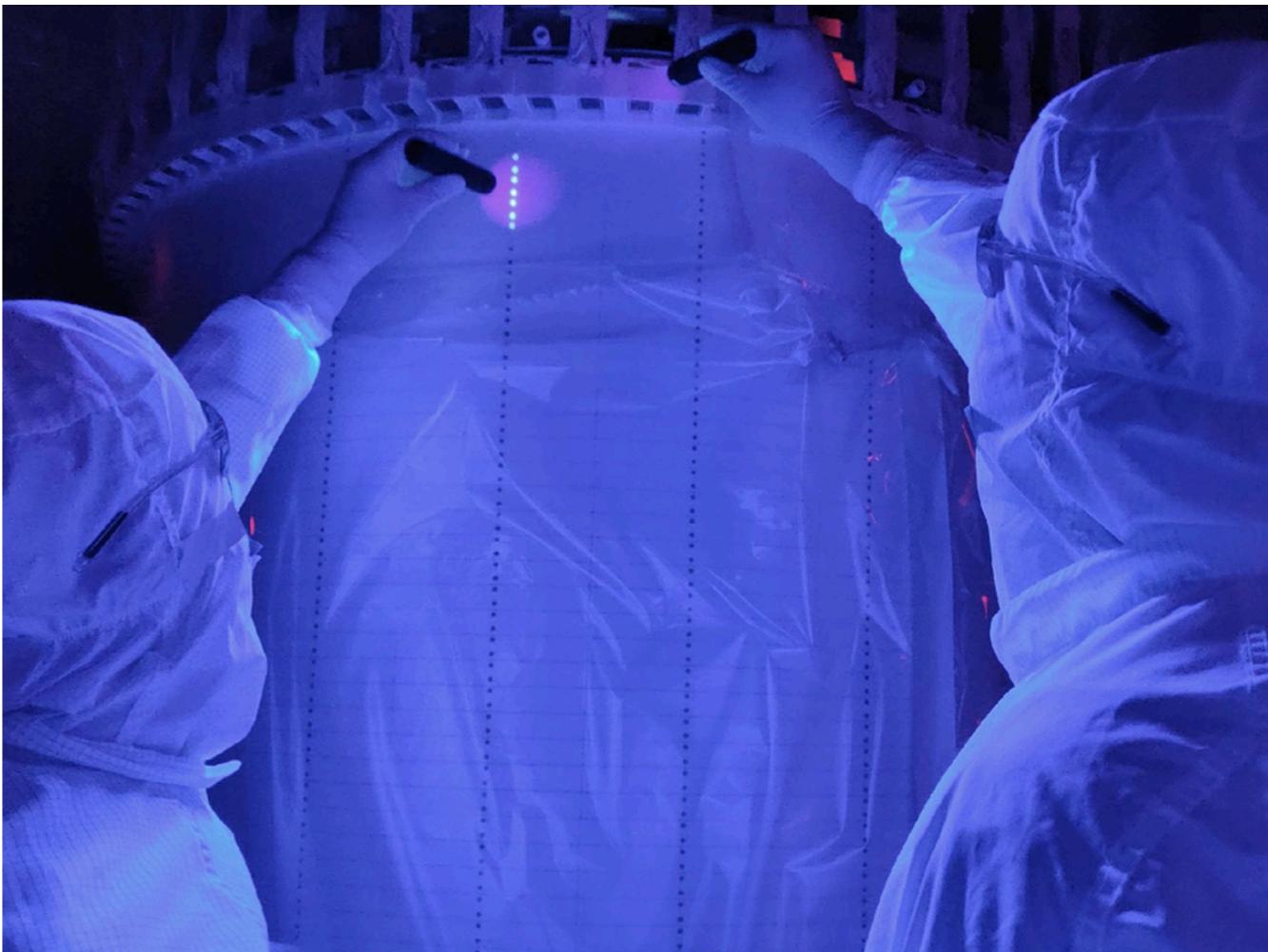
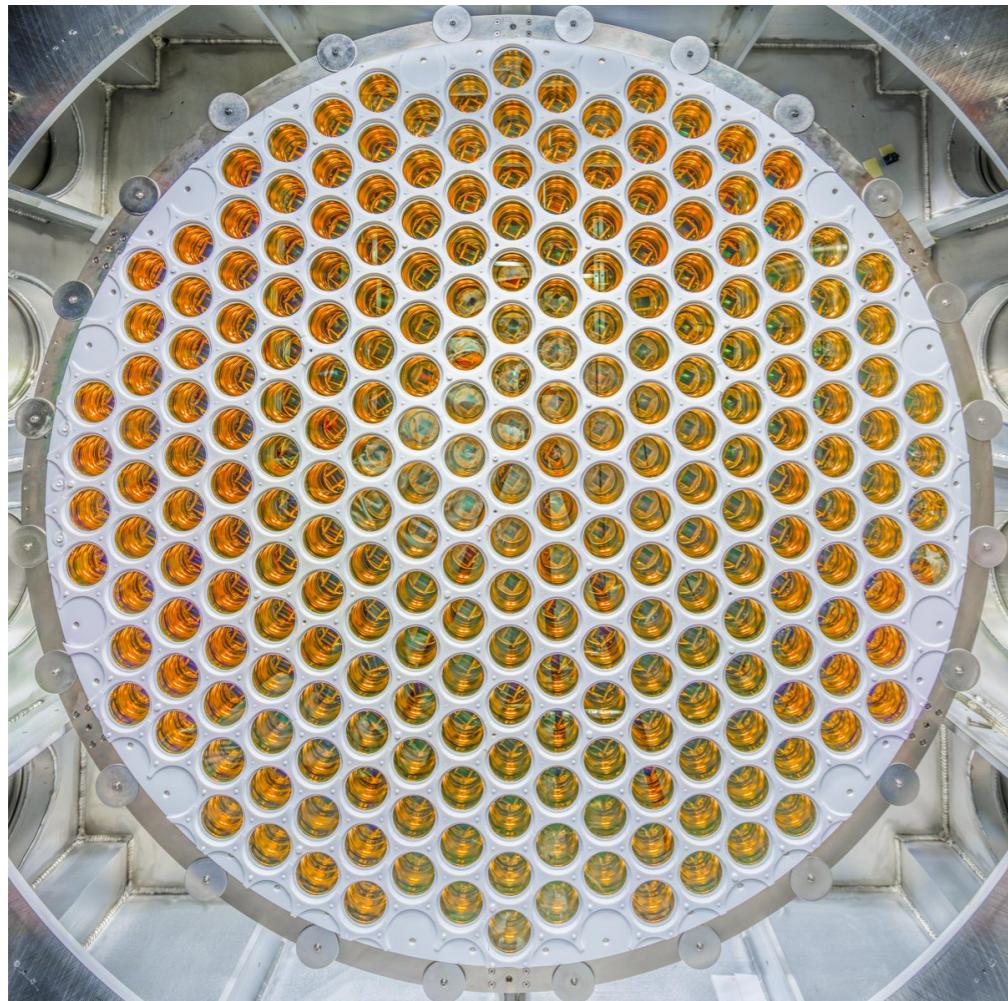
Instrumented with Hamamatsu R11410-22 3" Photomultiplier Tubes (PMTs) on top and bottom arrays and high reflectivity PTFE walls

Sensitivity in VUV range

494 PMTs in total



The LZ Experiment

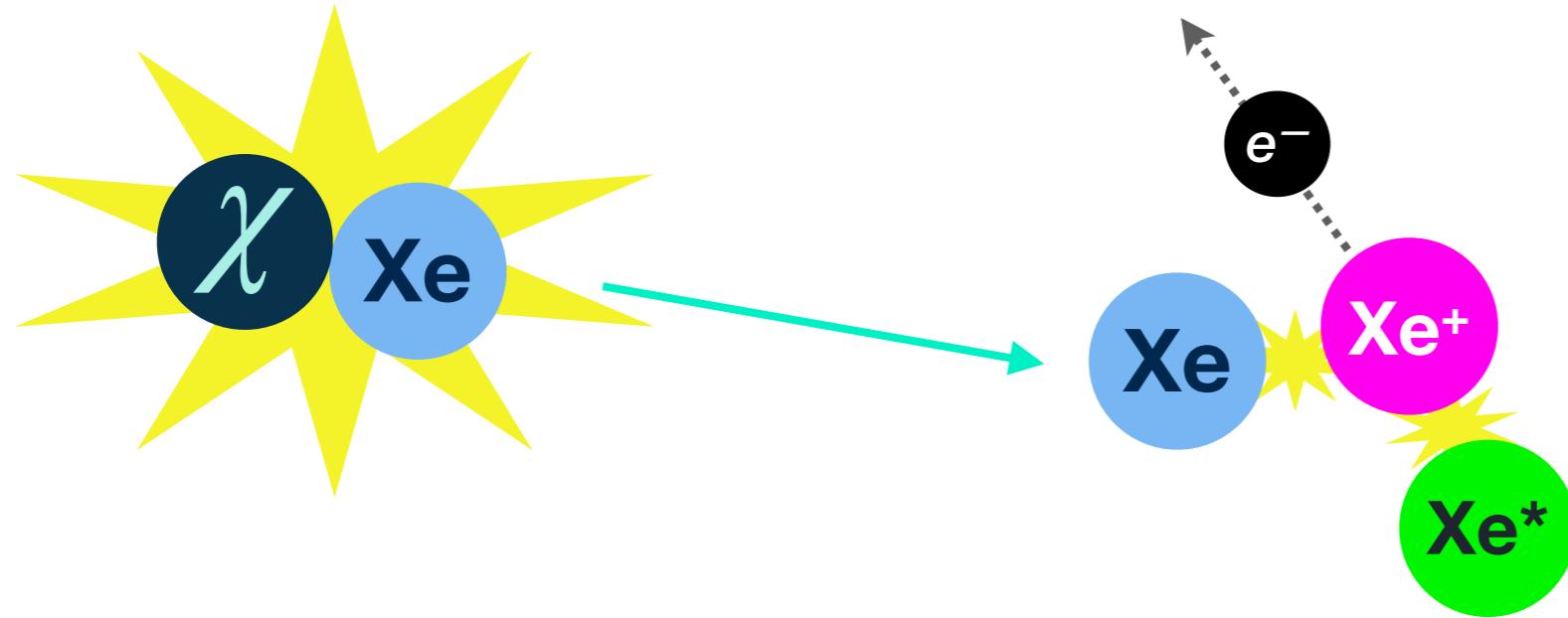


The LZ Experiment



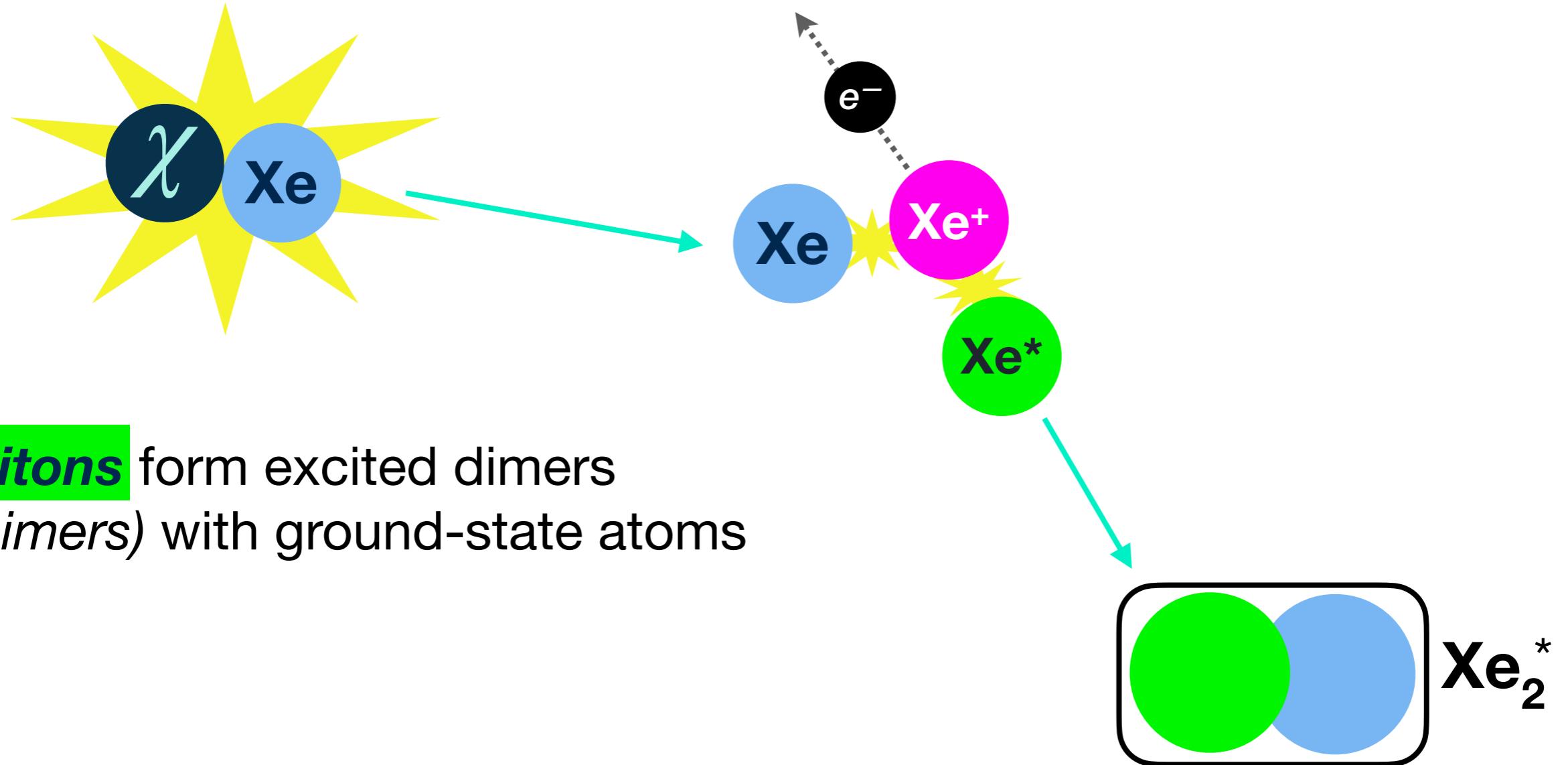
1 mile under the Black Hills of South Dakota

The LZ Experiment

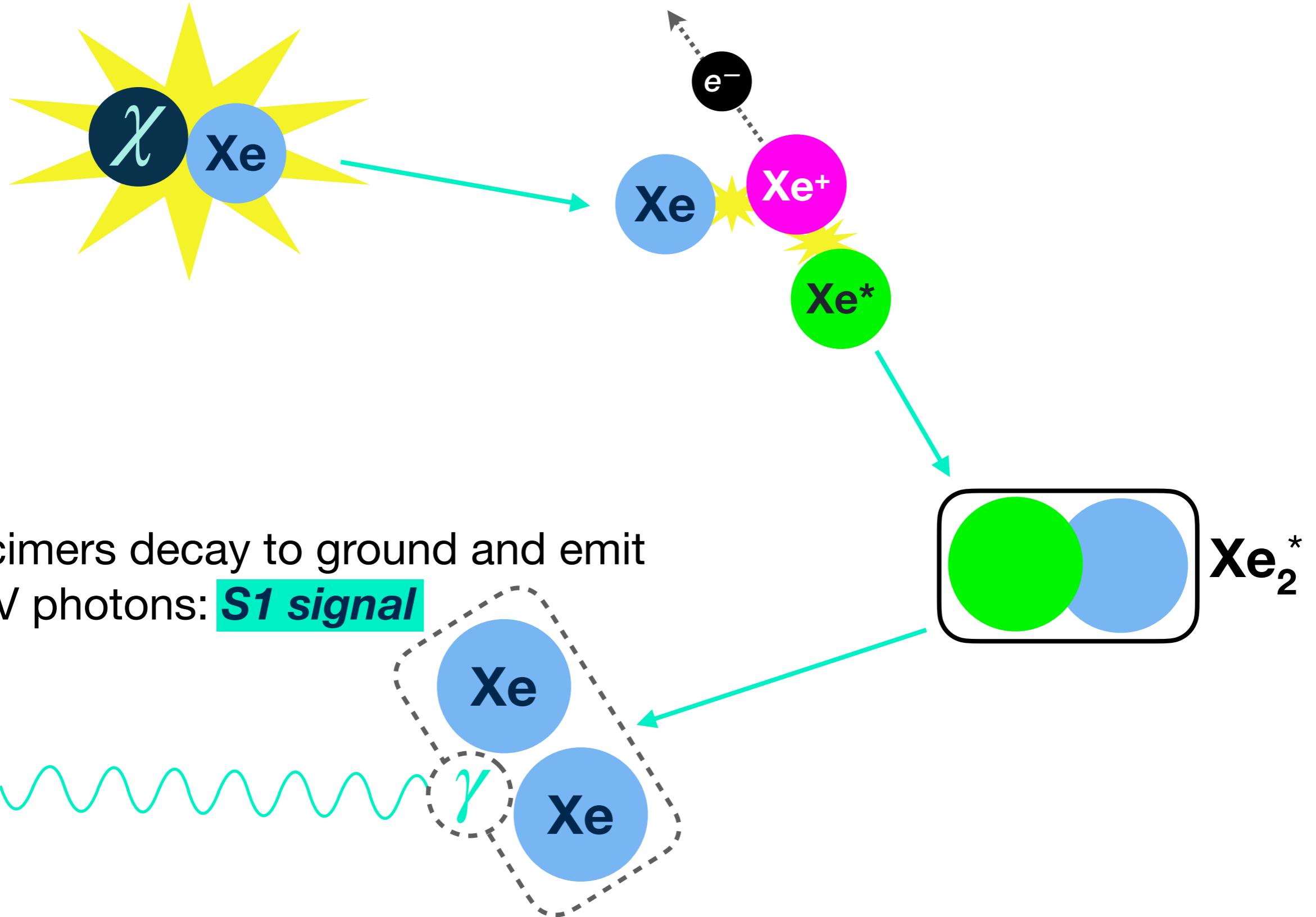


Atoms in the track of a recoiling nucleus
become **excitons** or **ions**

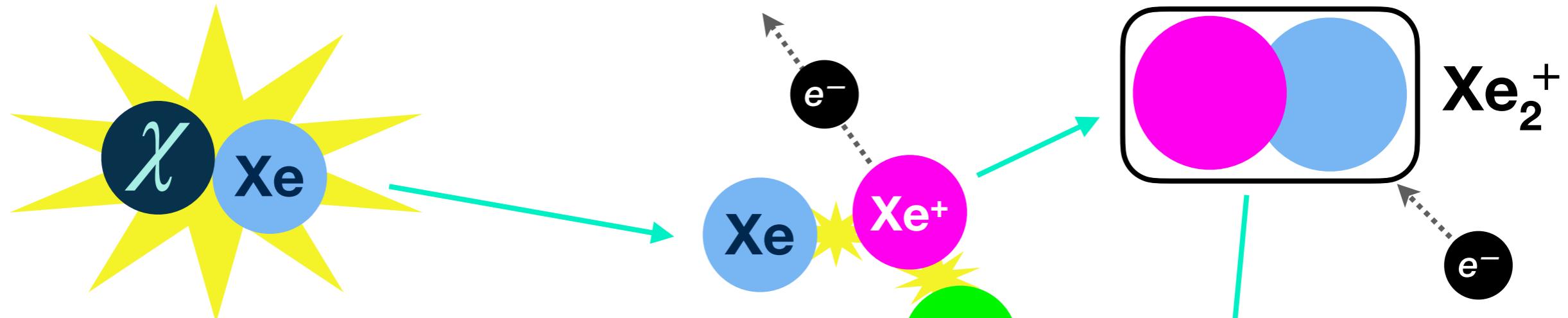
The LZ Experiment



The LZ Experiment

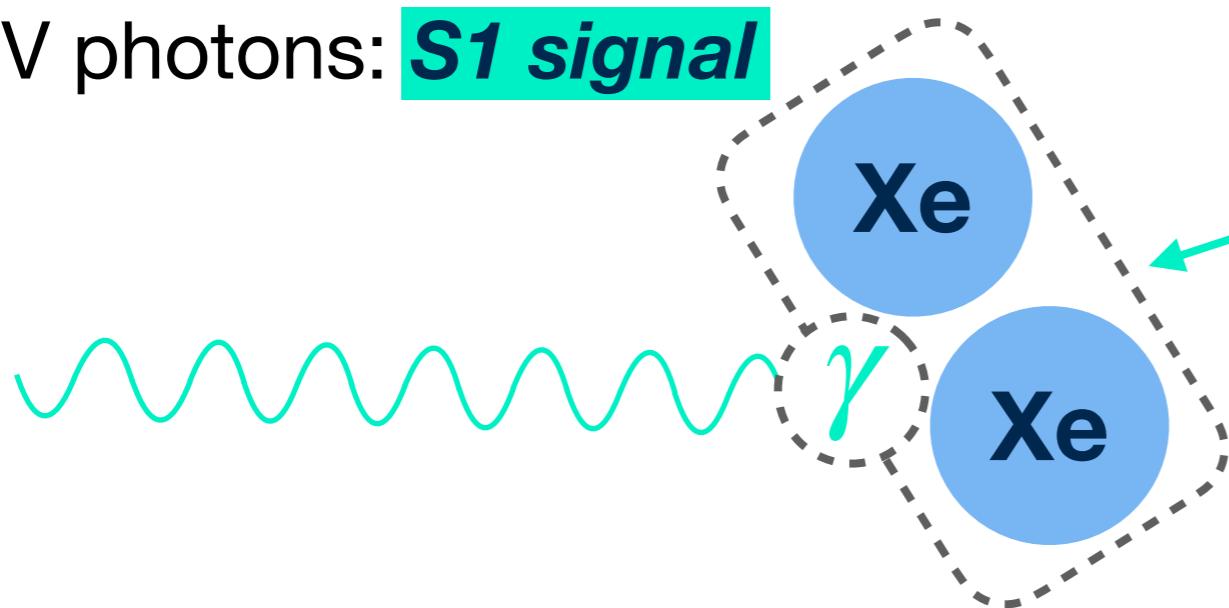


The LZ Experiment

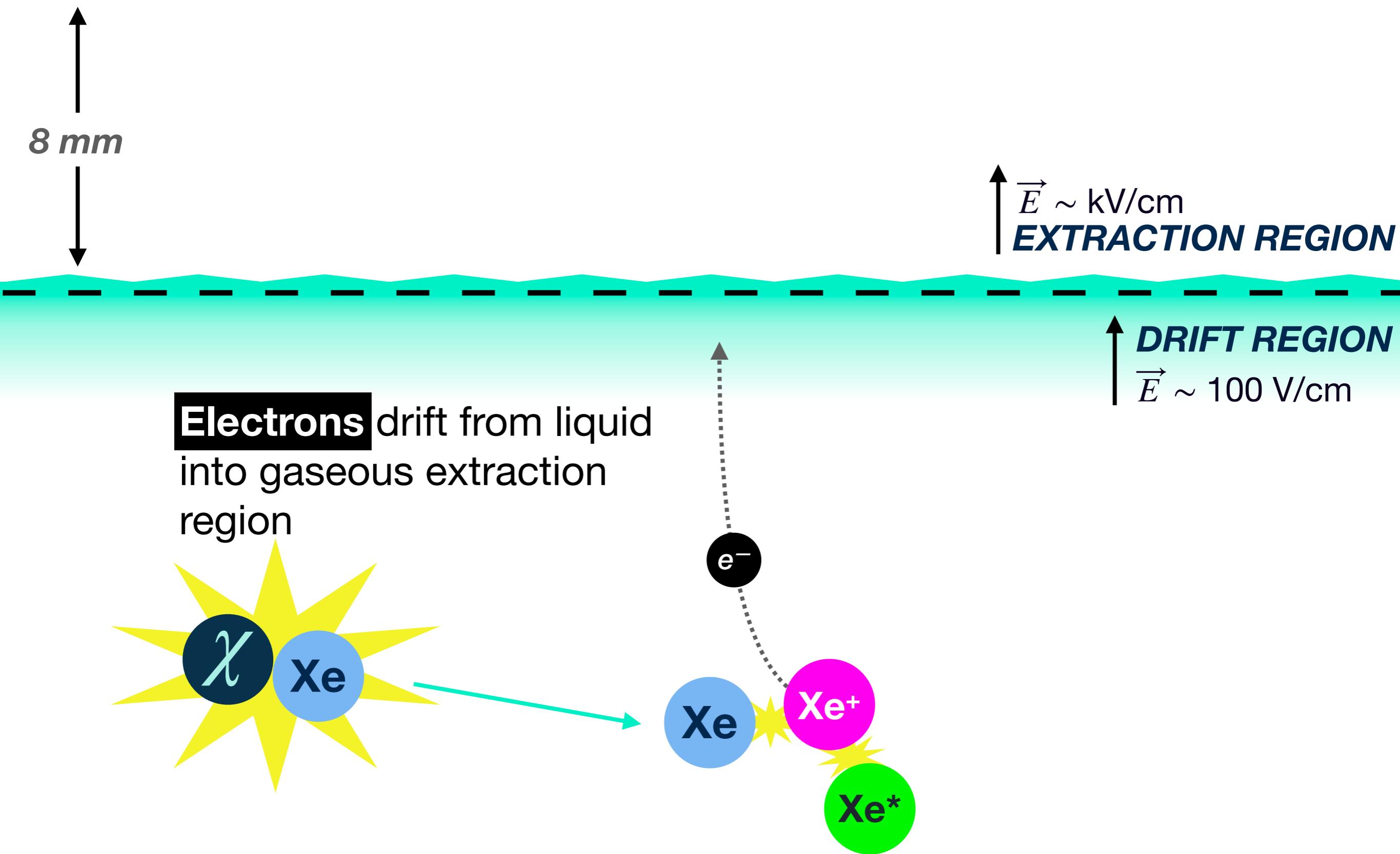


Ions form charged dimers, can recombine to become excimers

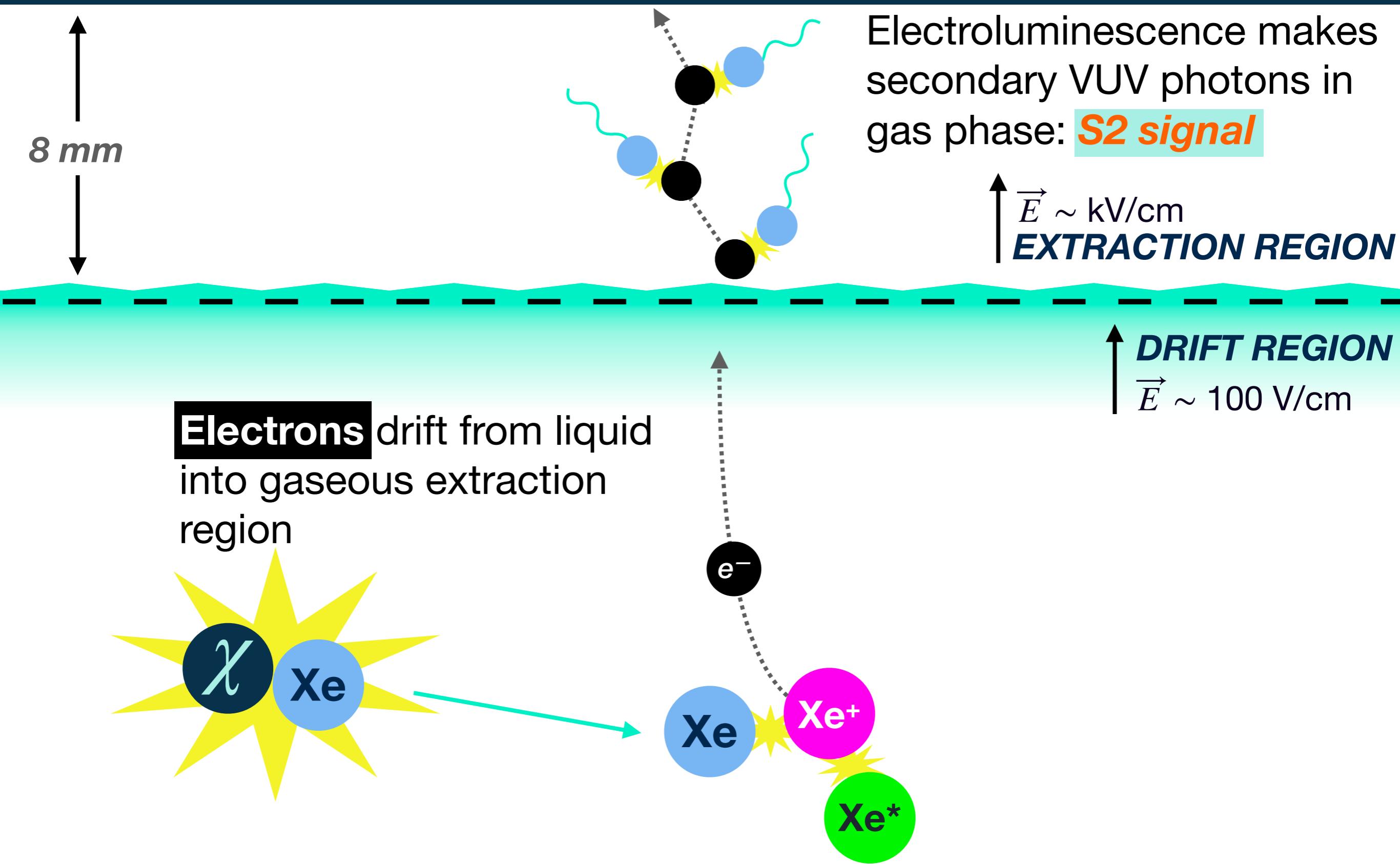
Excimers decay to ground and emit VUV photons: **S1 signal**



The LZ Experiment

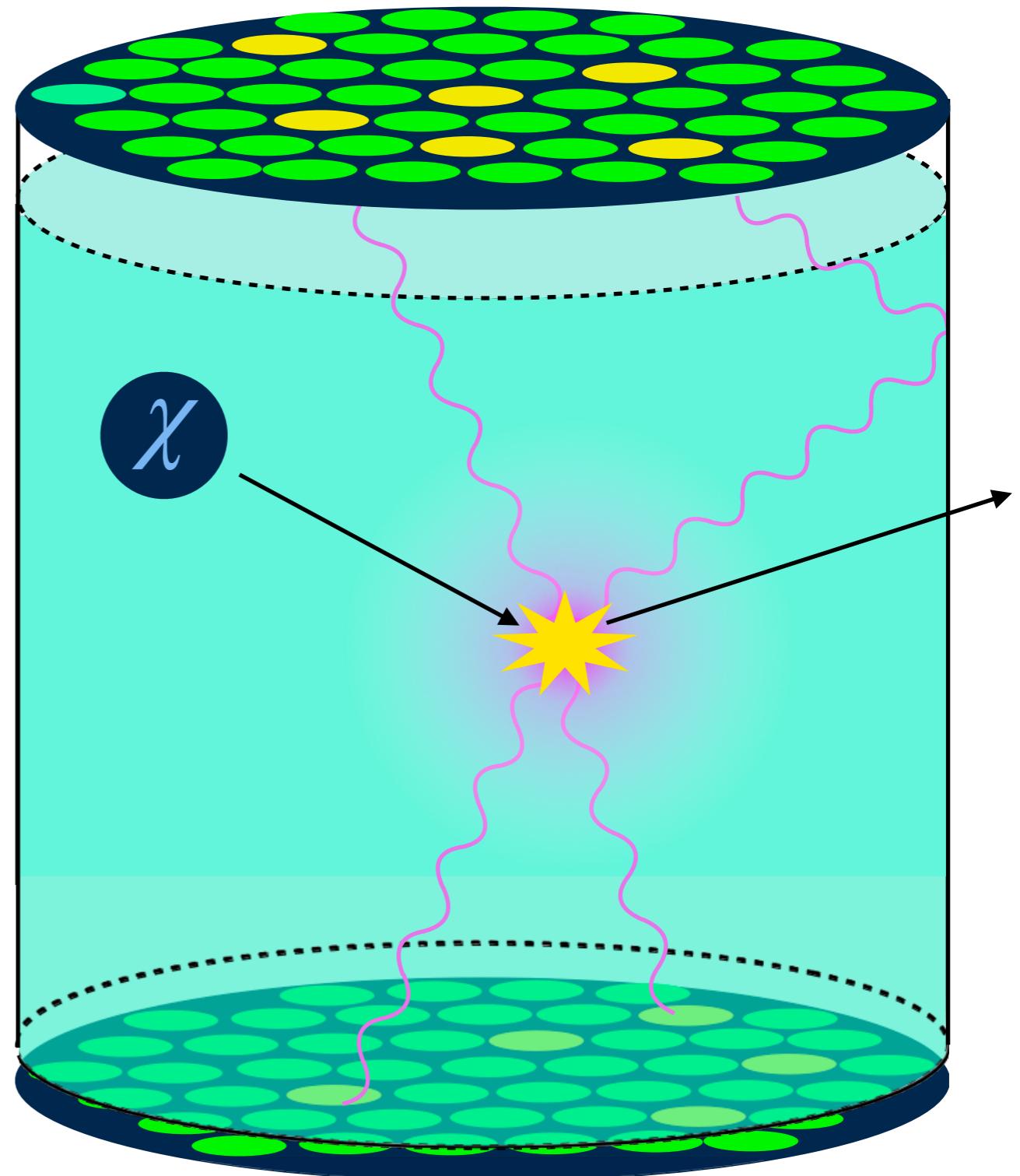


The LZ Experiment



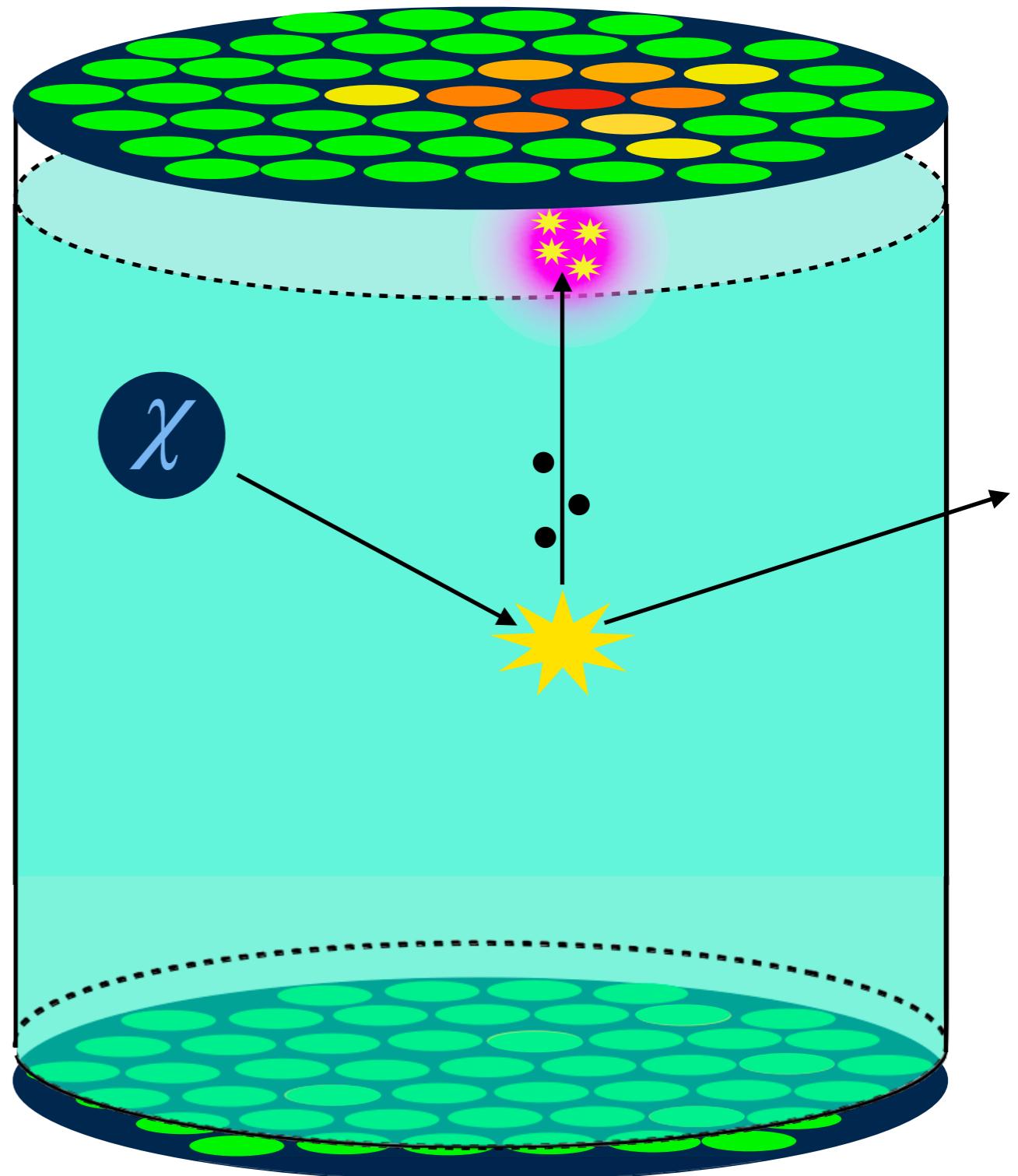
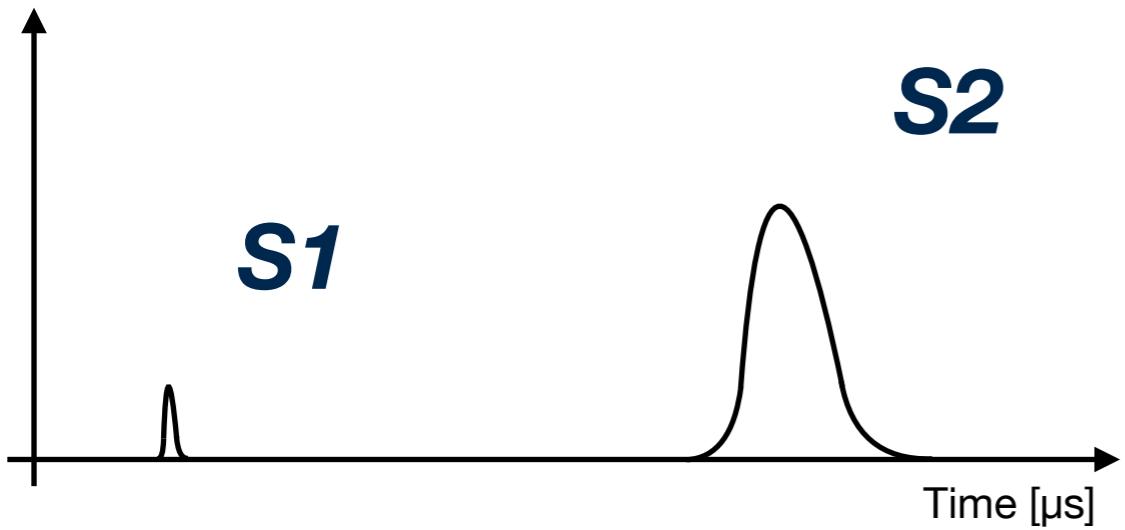
The LZ Experiment

A typical signal



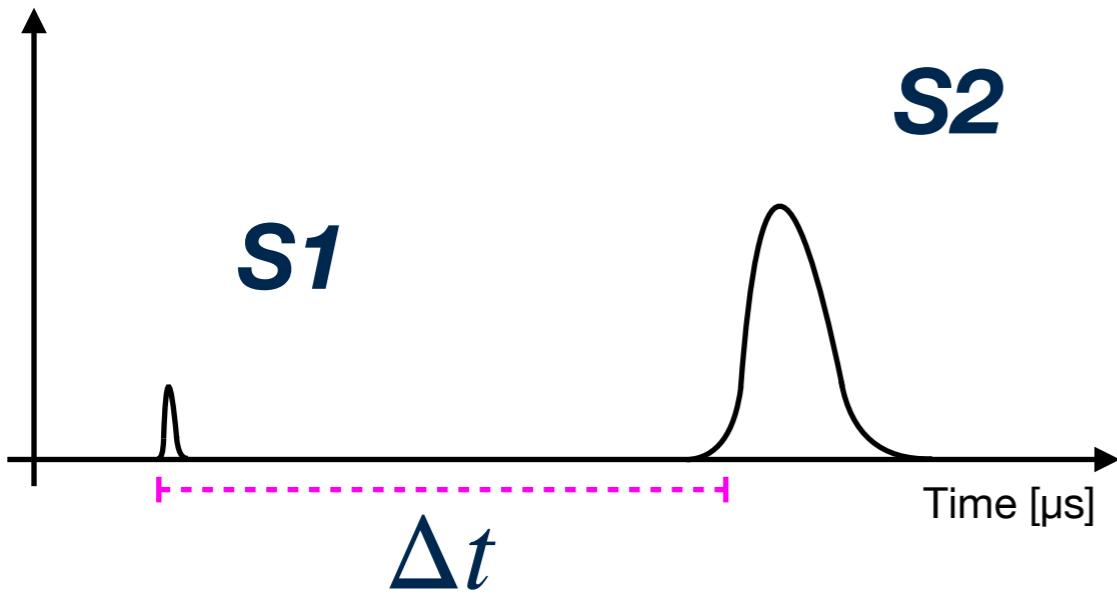
The LZ Experiment

A typical signal



The LZ Experiment

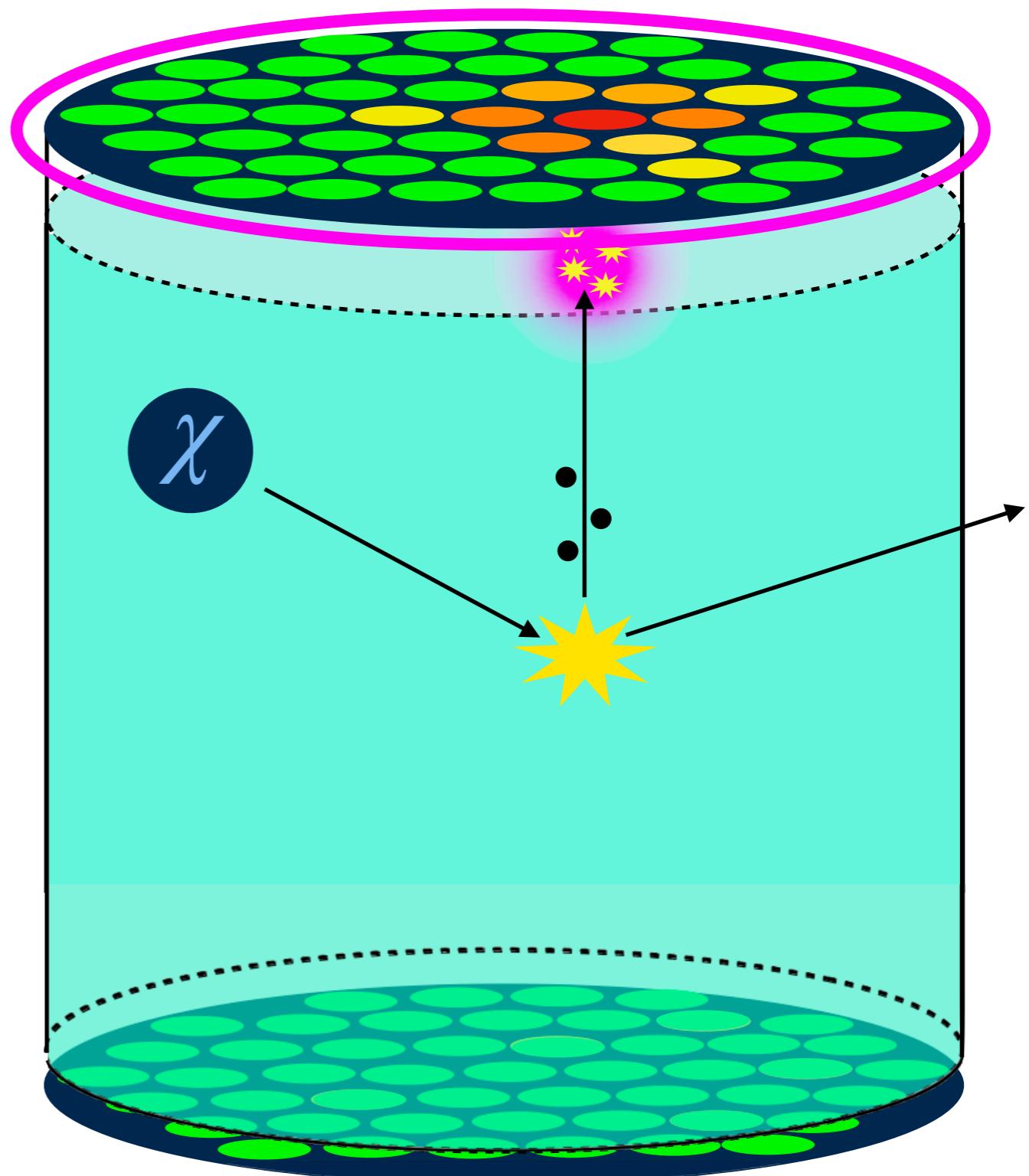
A typical signal



*Hit pattern tells us radial position
Δt tells us vertical position*

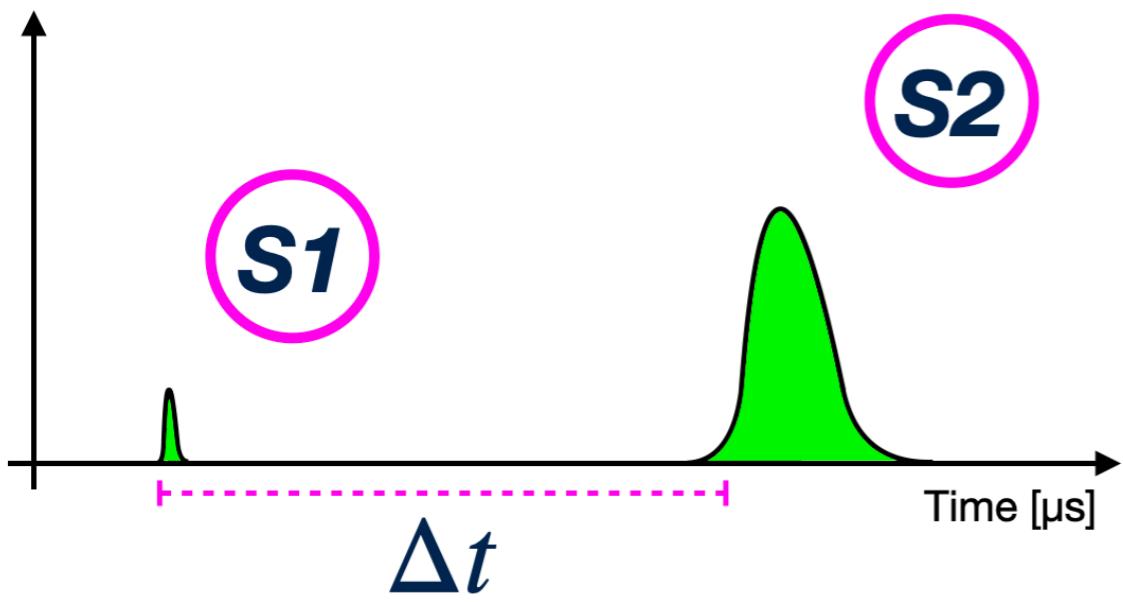
Resolution ~ mm

Critical for fiducialization!



The LZ Experiment

A typical signal

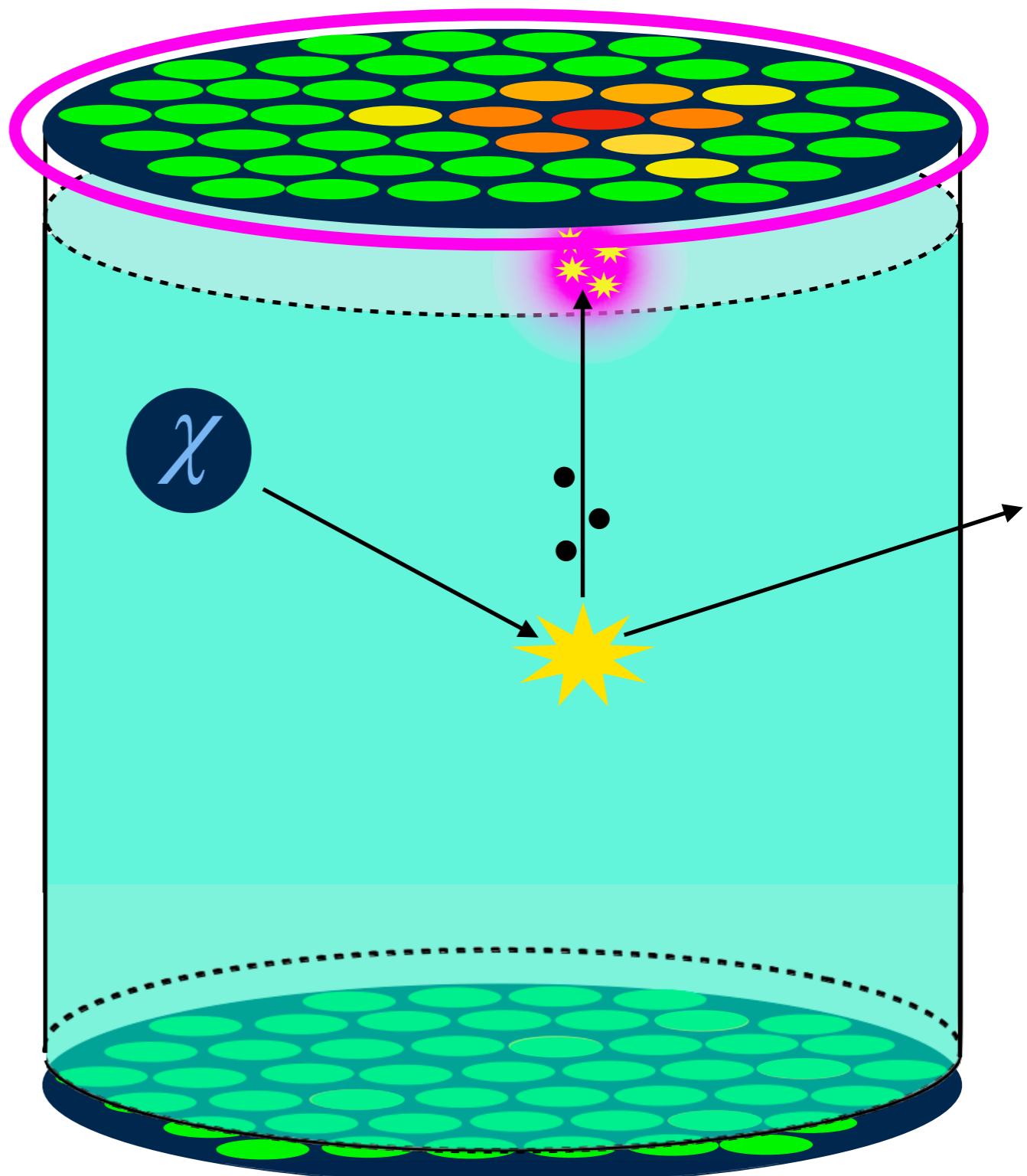


Hit pattern tells us radial position

Δt tells us vertical position

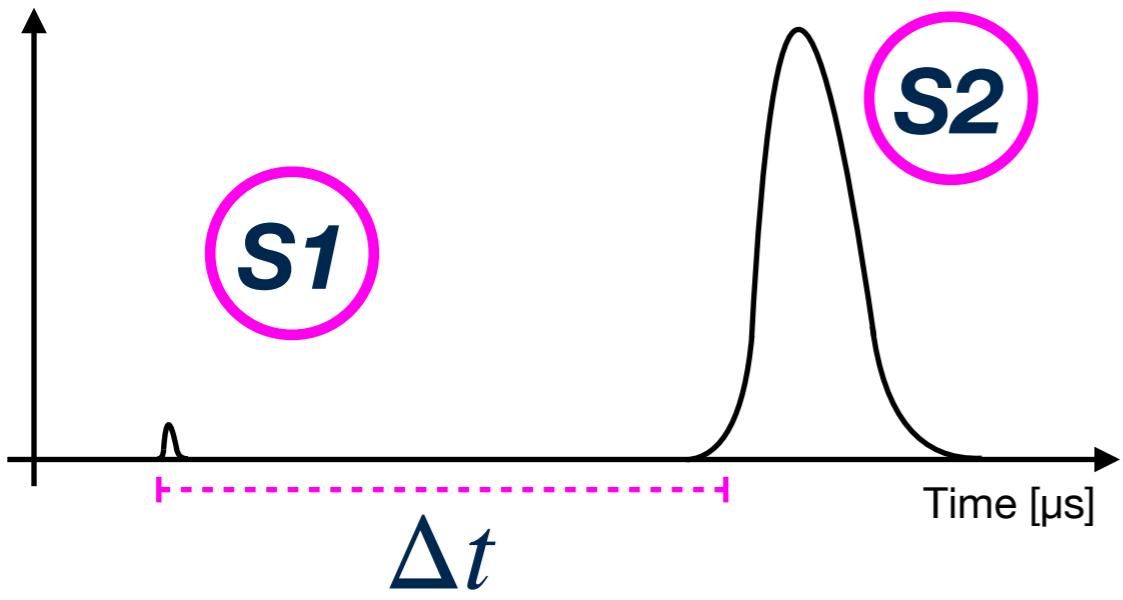
S1+S2 tells us the collision energy

*Expressed in ‘photons detected’
(phd)*



The LZ Experiment

A typical signal

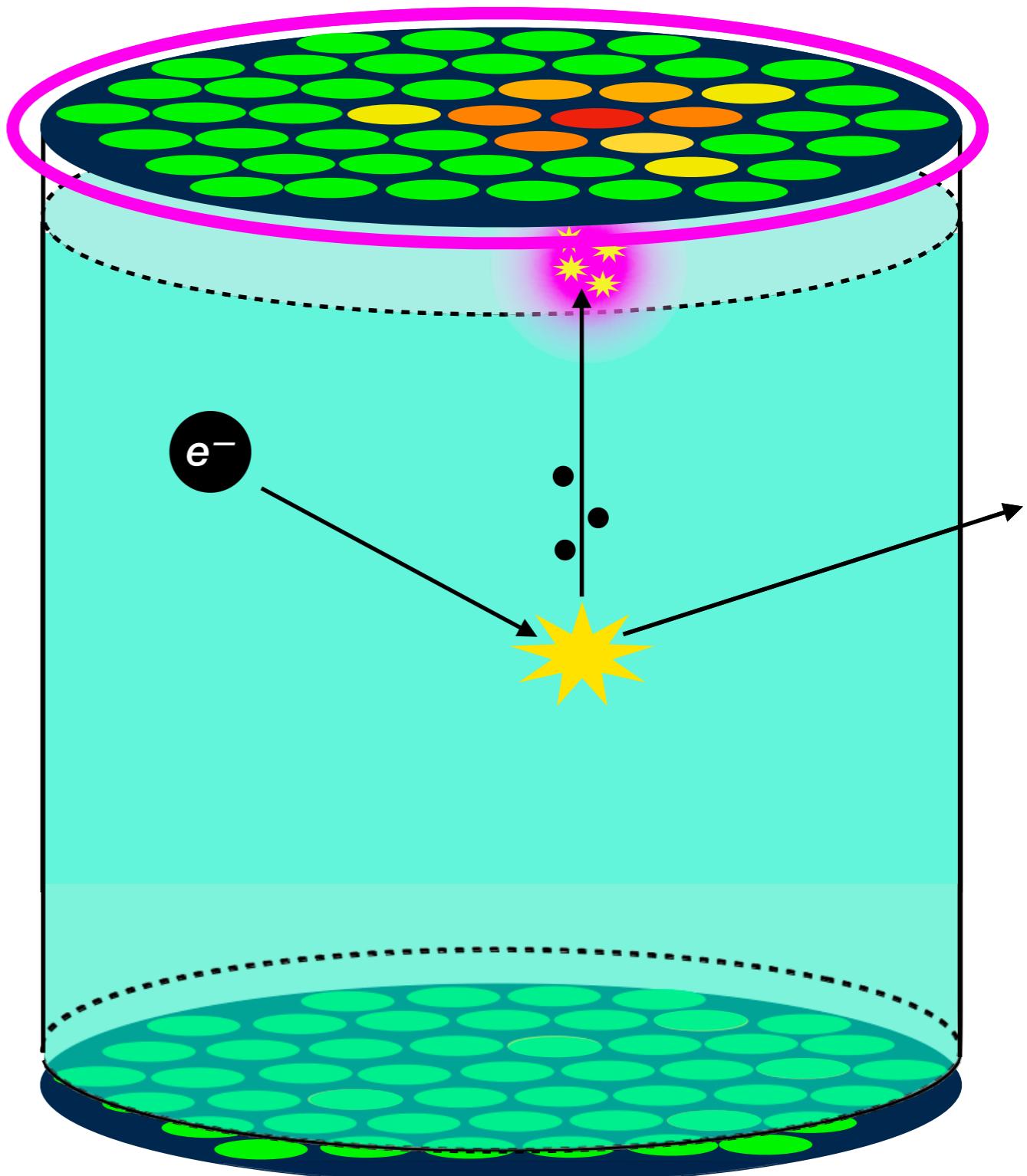


Hit pattern tells us radial position

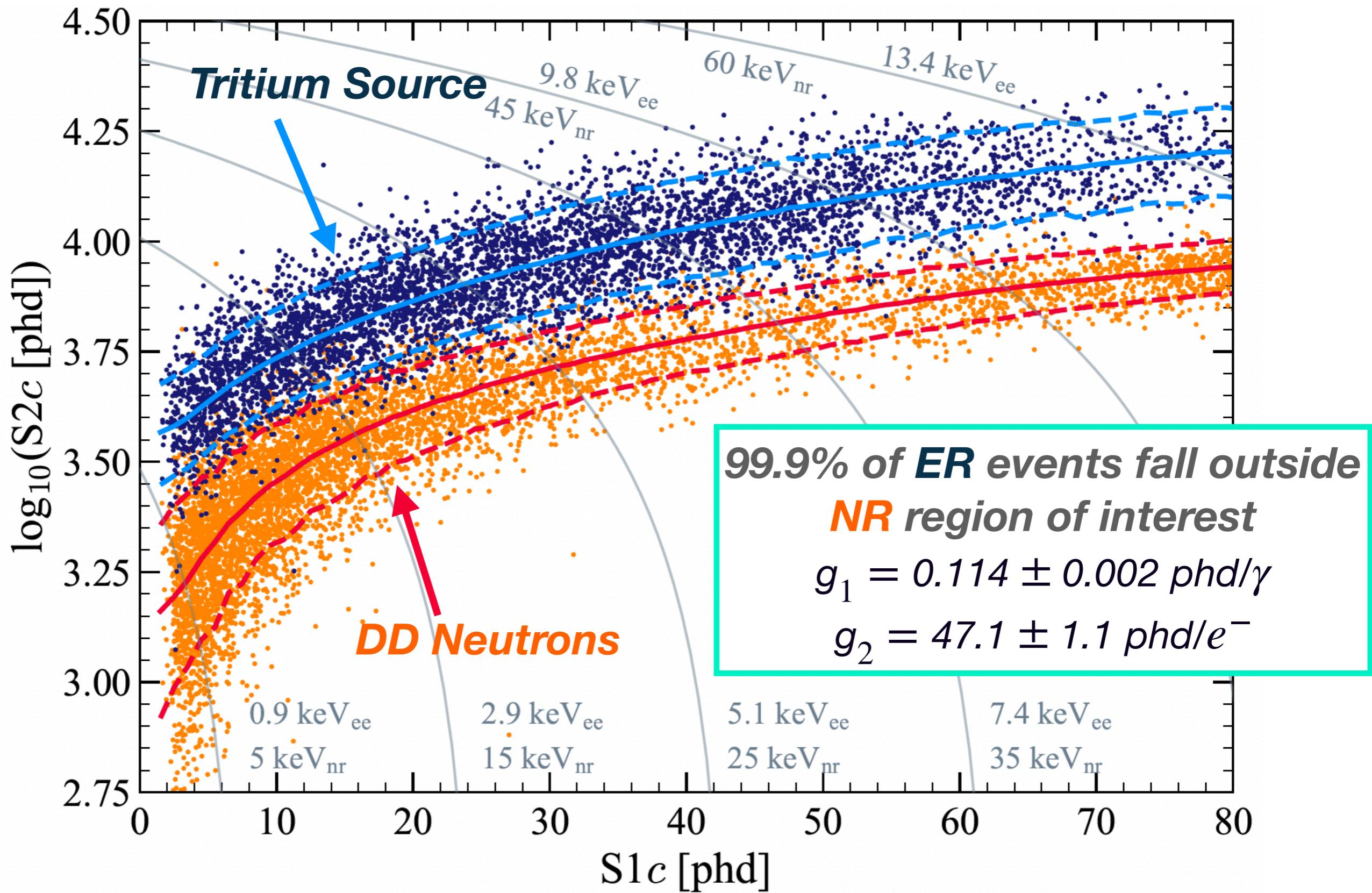
Δt tells us vertical position

$S_1 + S_2$ tells us the collision energy

$S_1:S_2$ ratio tells us the interaction type

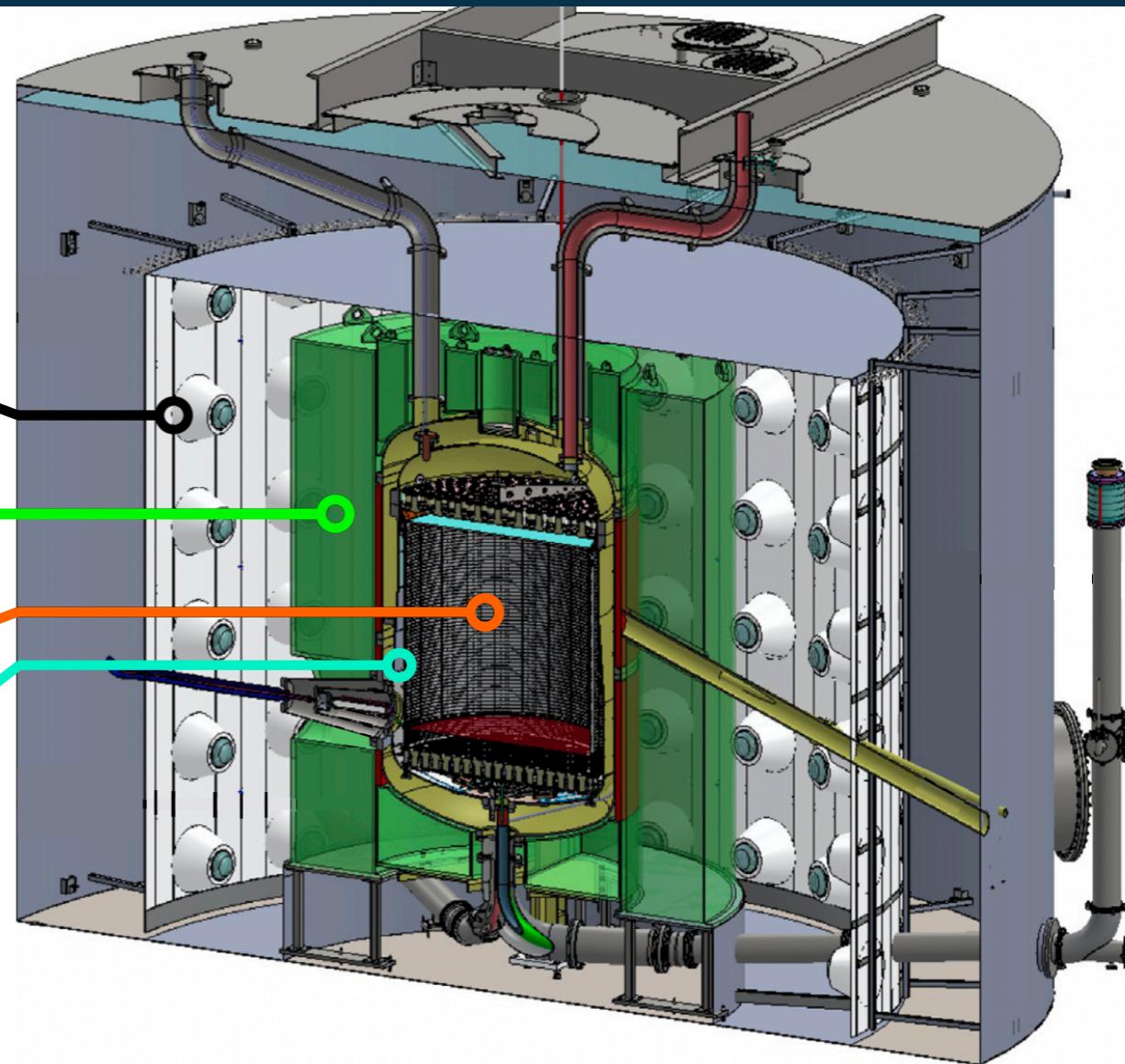
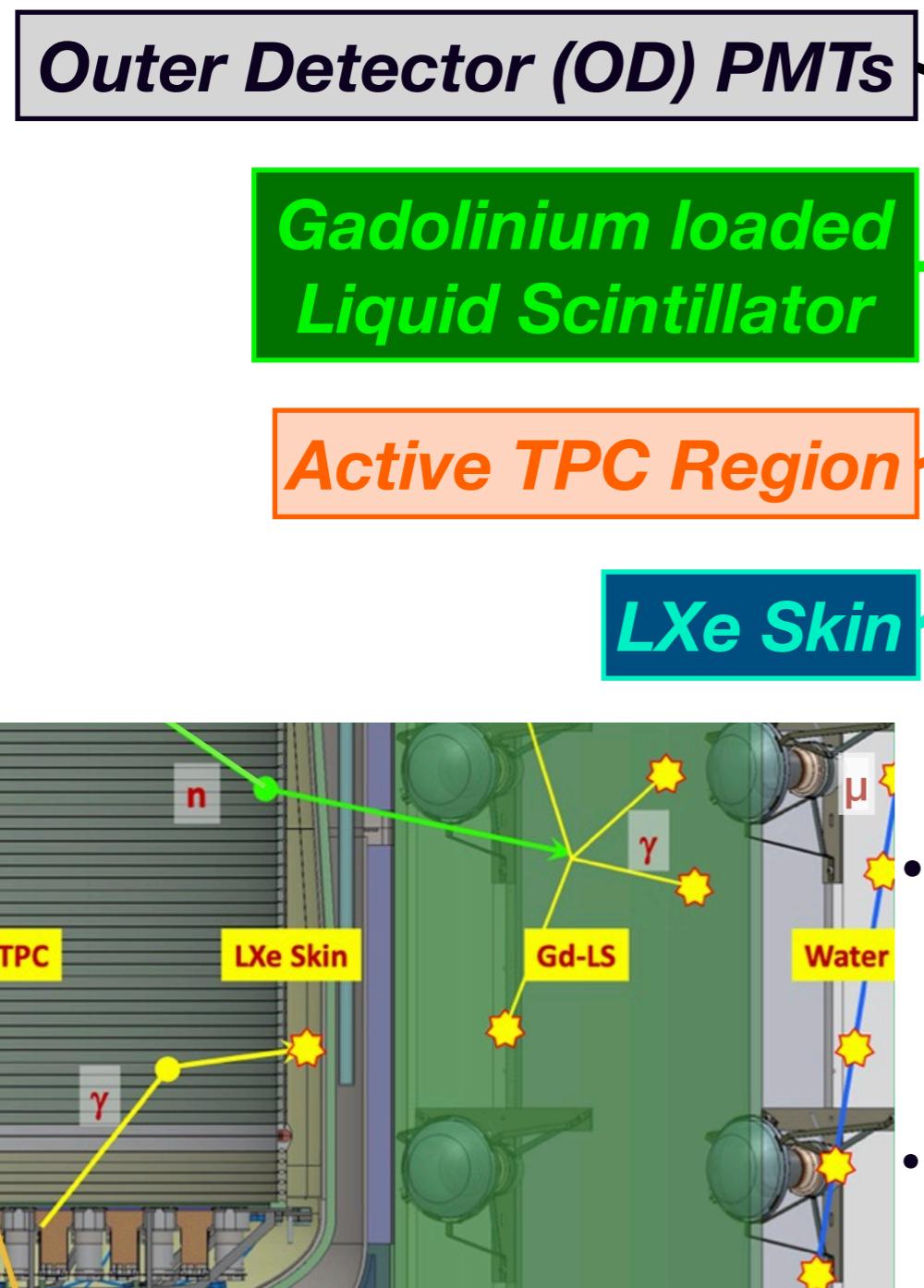


The LZ Experiment



The LZ Experiment

Anti-Coincidence Veto System



- 17 tonnes **Gd-loaded scintillator** in OD
 - High thermal neutron capture cross-section
 - Release of ~8 MeV γ -rays from neutron capture
- **LXe Skin** detector effective at tagging γ -rays from internal TPC decays, OD neutron captures

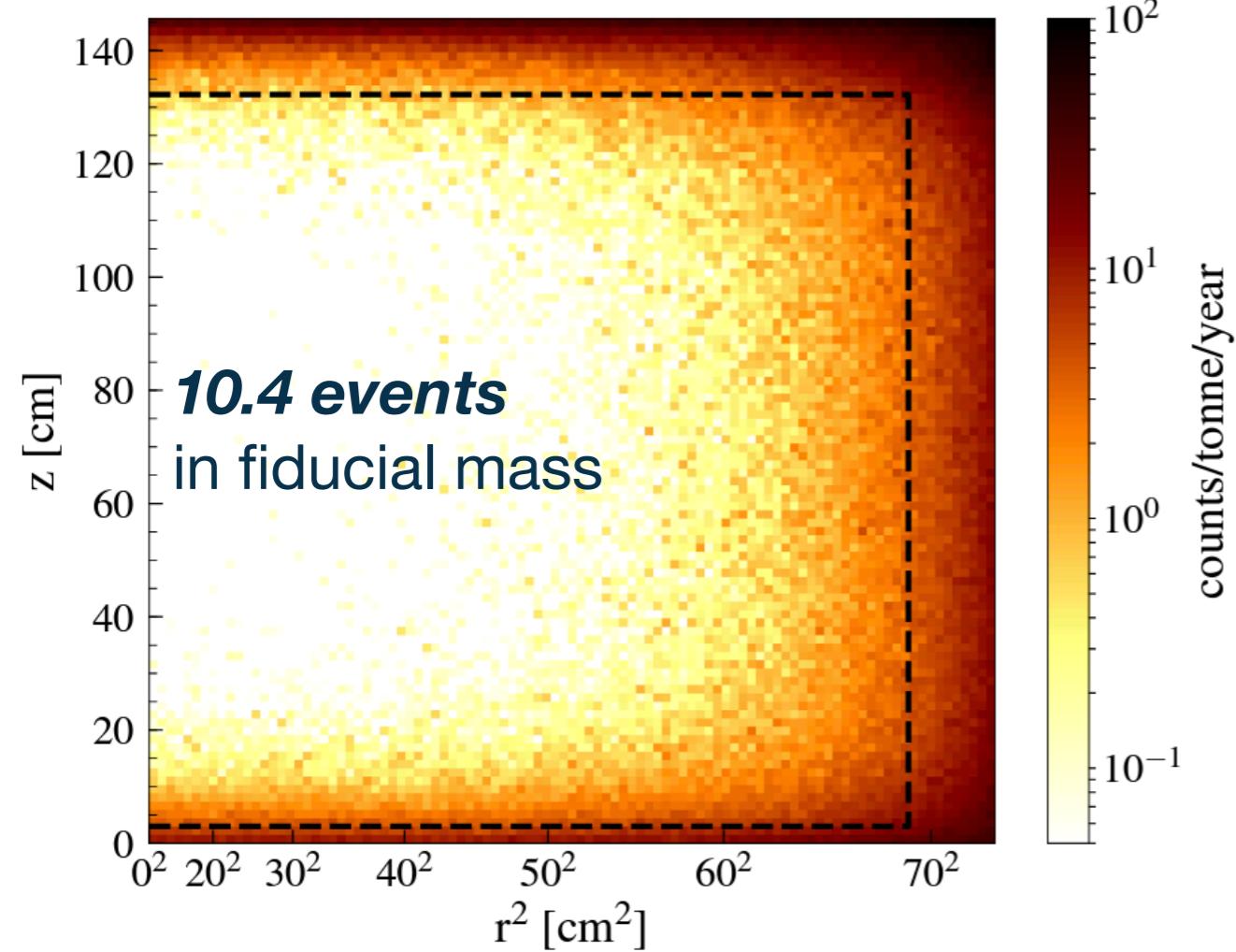
The LZ Experiment

Anti-Coincidence Veto System

- Simulations show anti-coincidence veto reduces NR backgrounds in 1000 day run by factor of 10

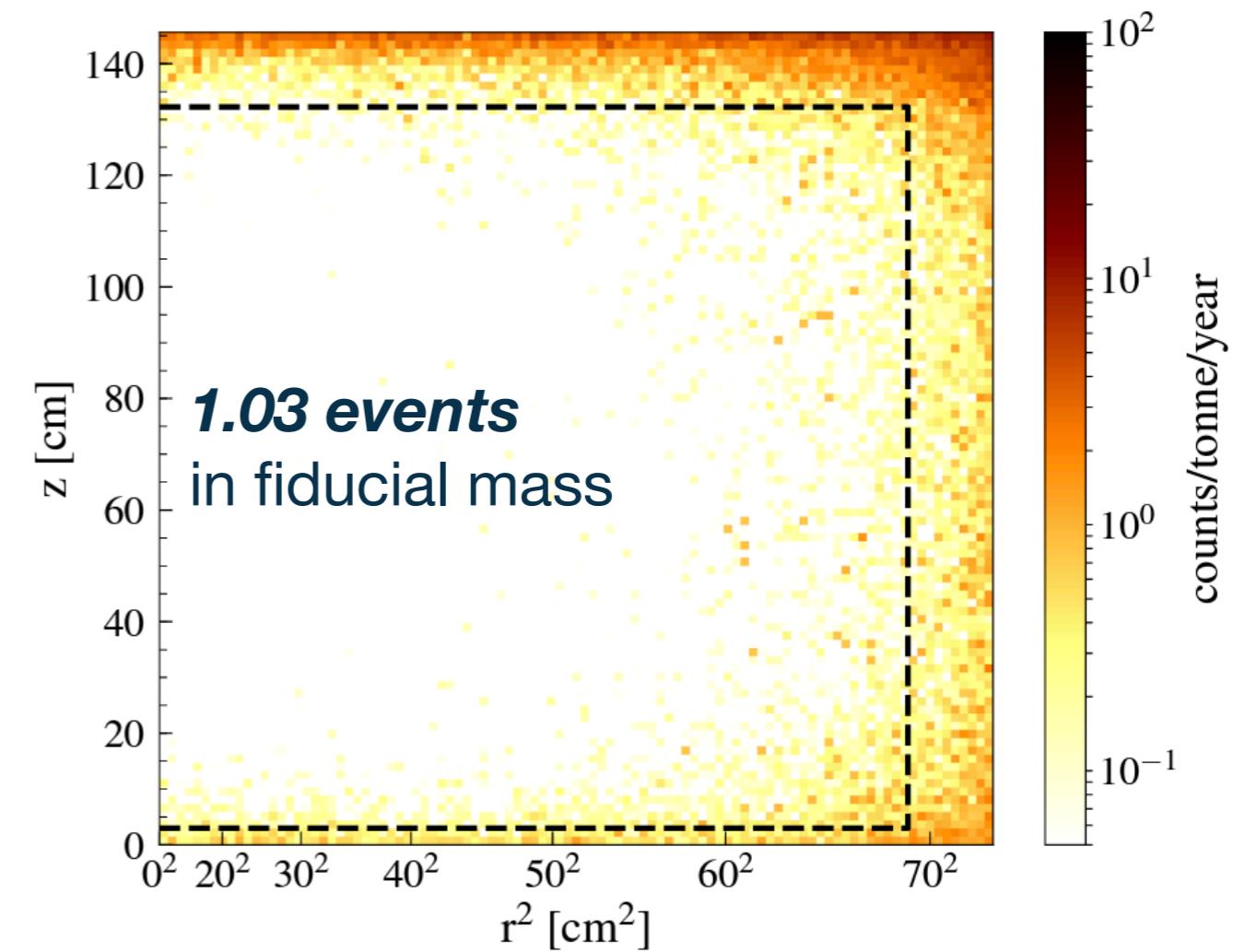
6-30 keV nuclear recoils

before anti-coincidence veto



6-30 keV nuclear recoils

after anti-coincidence veto



First Results (SR1)

PHYSICAL REVIEW LETTERS 131, 041002 (2023)

Editors' Suggestion

Featured in Physics

First Dark Matter Search Results from the LUX-ZEPLIN (LZ) Experiment

J. Aalbers,^{1,2} D. S. Akerib,^{1,2} C. W. Akerlof,³ A. K. Al Musalhi,⁴ F. Alder,⁵ A. Alqahtani,⁶ S. K. Alsum,⁷ C. S. Amarasinghe,³ A. Ames,^{1,2} T. J. Anderson,^{1,2} N. Angelides,^{5,8} H. M. Araújo,⁸ J. E. Armstrong,⁹ M. Arthurs,³ S. Azadi,¹⁰ A. J. Bailey,⁸ A. Baker,⁸ J. Balajthy,¹¹ S. Balashov,¹² J. Bang,⁶ J. W. Bargemann,¹⁰ M. J. Barry,¹³ J. Barthel,¹⁴ D. Bauer,⁸ A. Baxter,¹⁵ K. Beattie,¹³ J. Belle,¹⁶ P. Beltrame,^{5,17} J. Bensinger,¹⁸ T. Benson,⁷ E. P. Bernard,^{13,19} A. Bhatti,⁹



(Received 18 July 2022; revised 6 March 2023; accepted 7 June 2023; published 28 July 2023)

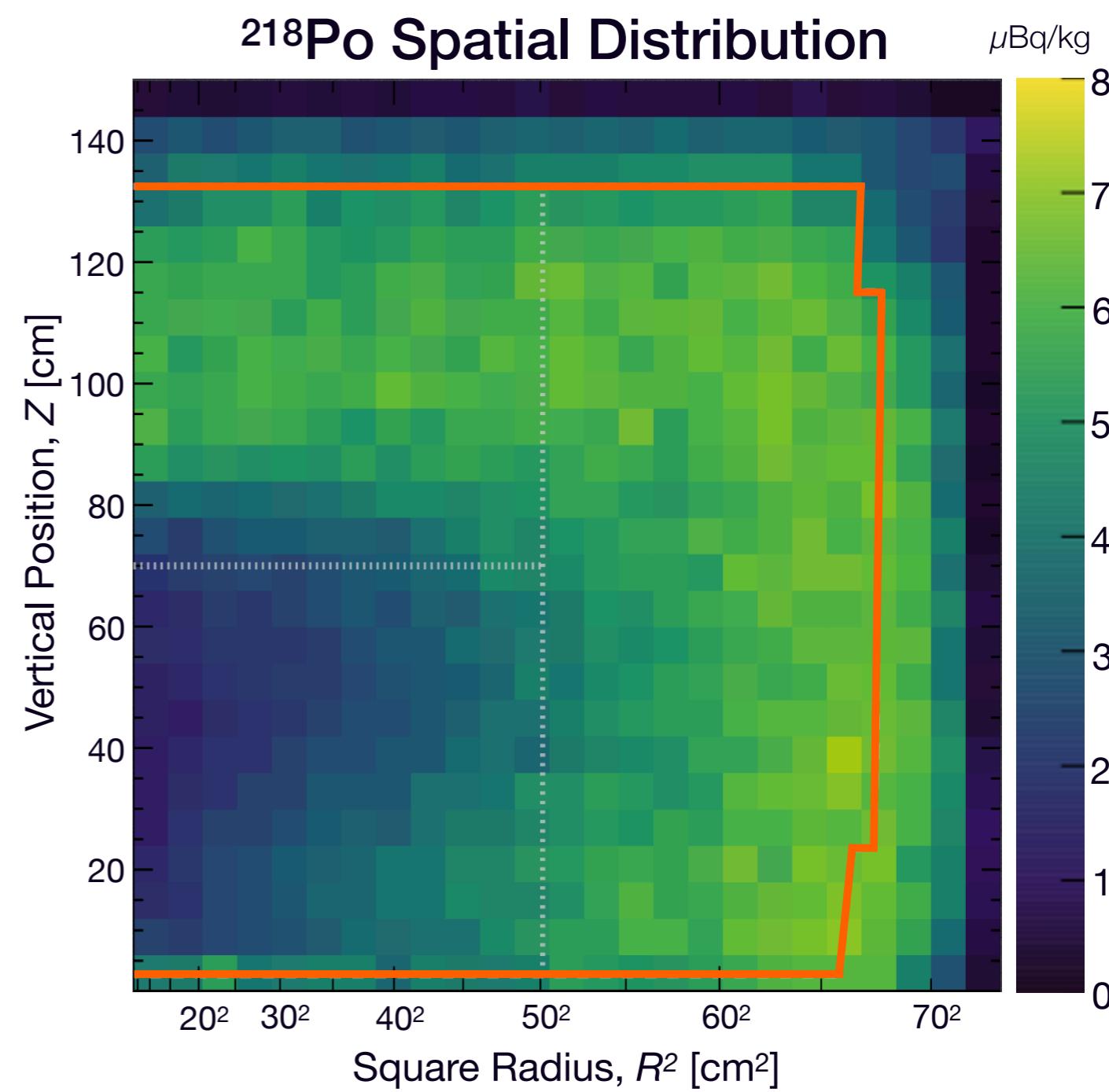
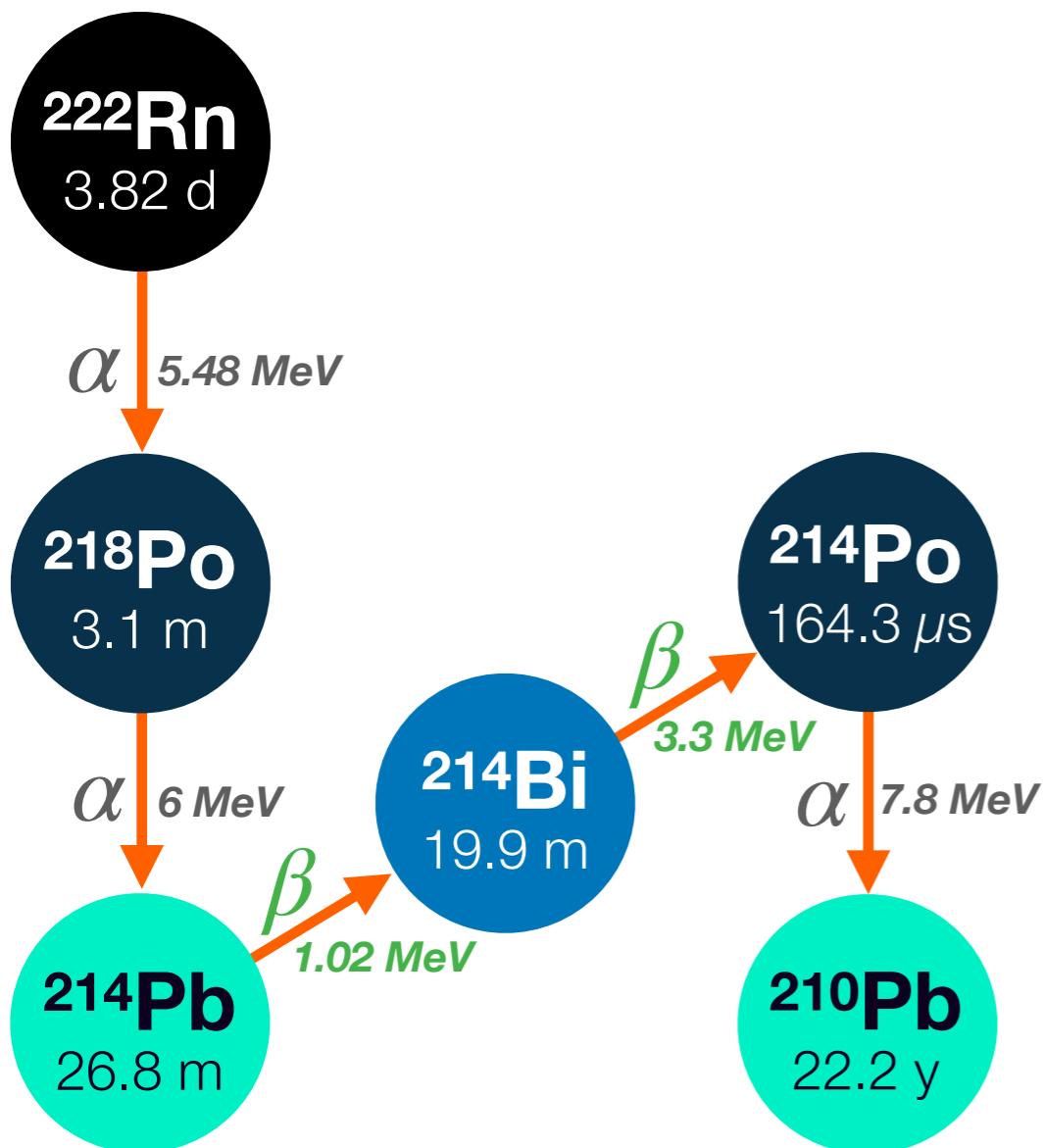
The LUX-ZEPLIN experiment is a dark matter detector centered on a dual-phase xenon time projection chamber operating at the Sanford Underground Research Facility in Lead, South Dakota, USA. This Letter reports results from LUX-ZEPLIN's first search for weakly interacting massive particles (WIMPs) with an exposure of 60 live days using a fiducial mass of 5.5 t. A profile-likelihood ratio analysis shows the data to be consistent with a background-only hypothesis, setting new limits on spin-independent WIMP-nucleon, spin-dependent WIMP-neutron, and spin-dependent WIMP-proton cross sections for WIMP masses above 9 GeV/c². The most stringent limit is set for spin-independent scattering at 36 GeV/c², rejecting cross sections above 9.2×10^{-48} cm² at the 90% confidence level.

DOI: [10.1103/PhysRevLett.131.041002](https://doi.org/10.1103/PhysRevLett.131.041002)

SR1 : Backgrounds

Radon emanates from detector materials into the xenon

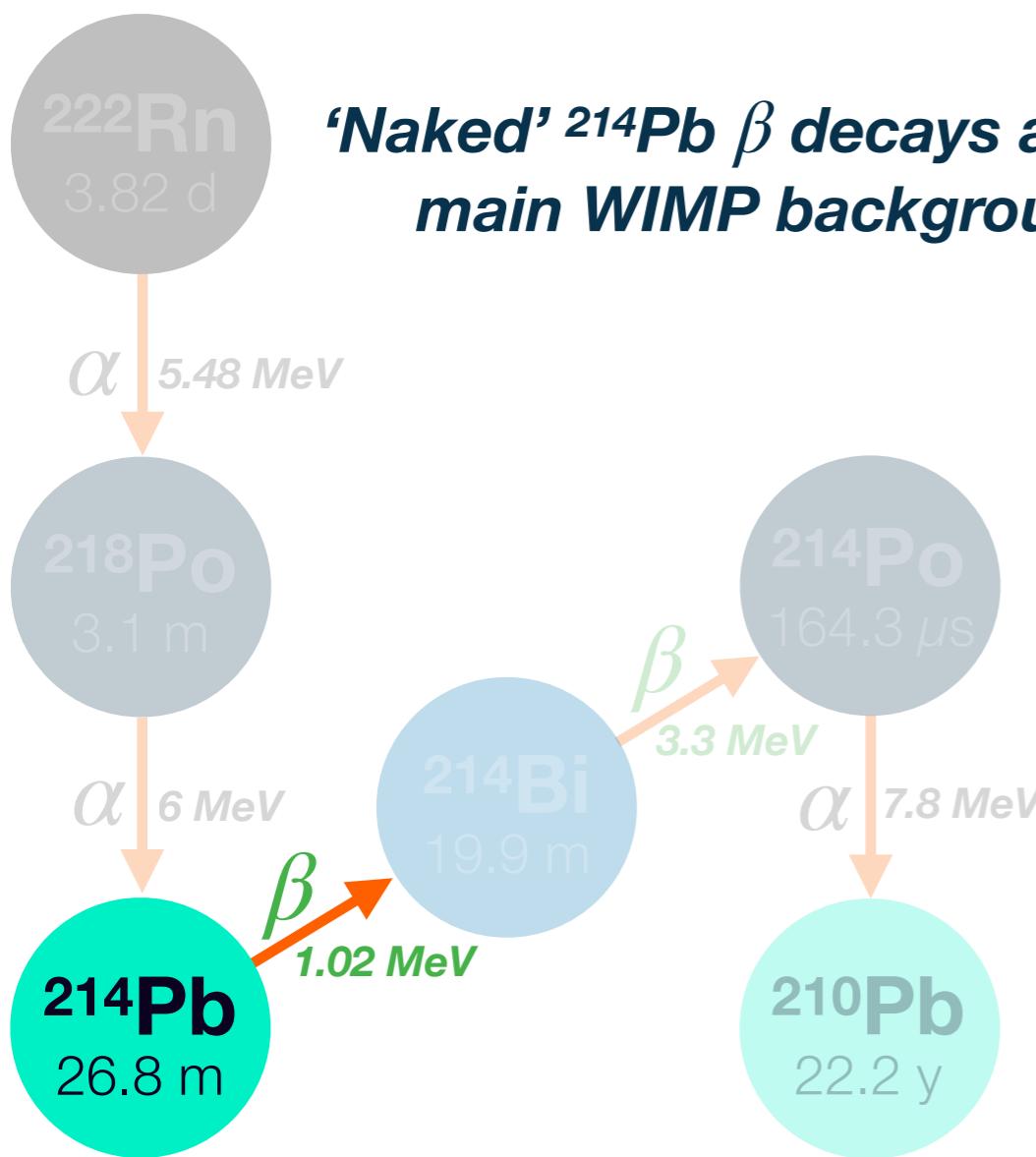
Non-uniform position distribution due to xenon flow and charged ion movement



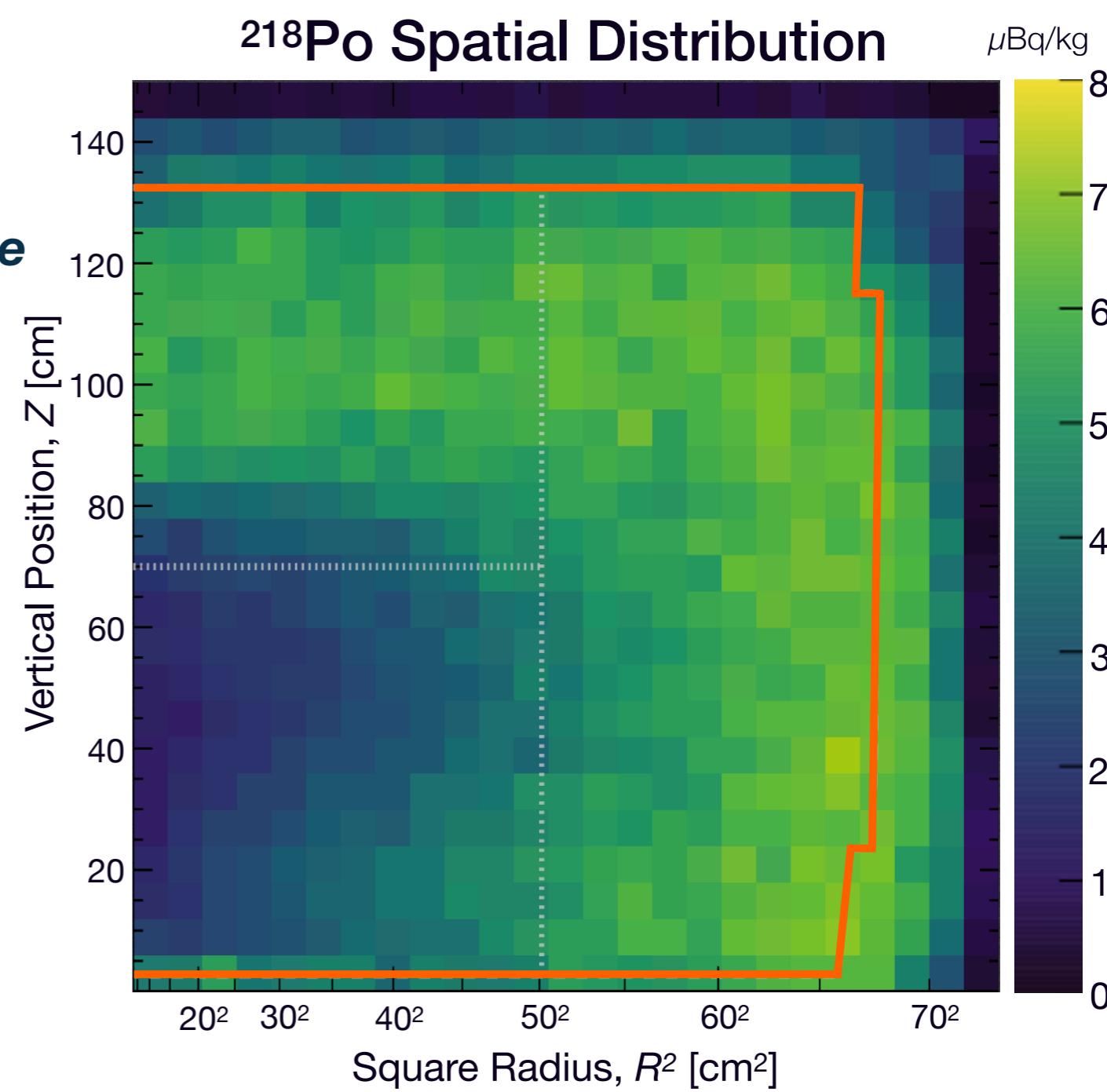
SR1 : Backgrounds

Radon emanates from detector materials into the xenon

Non-uniform position distribution due to xenon flow and charged ion movement

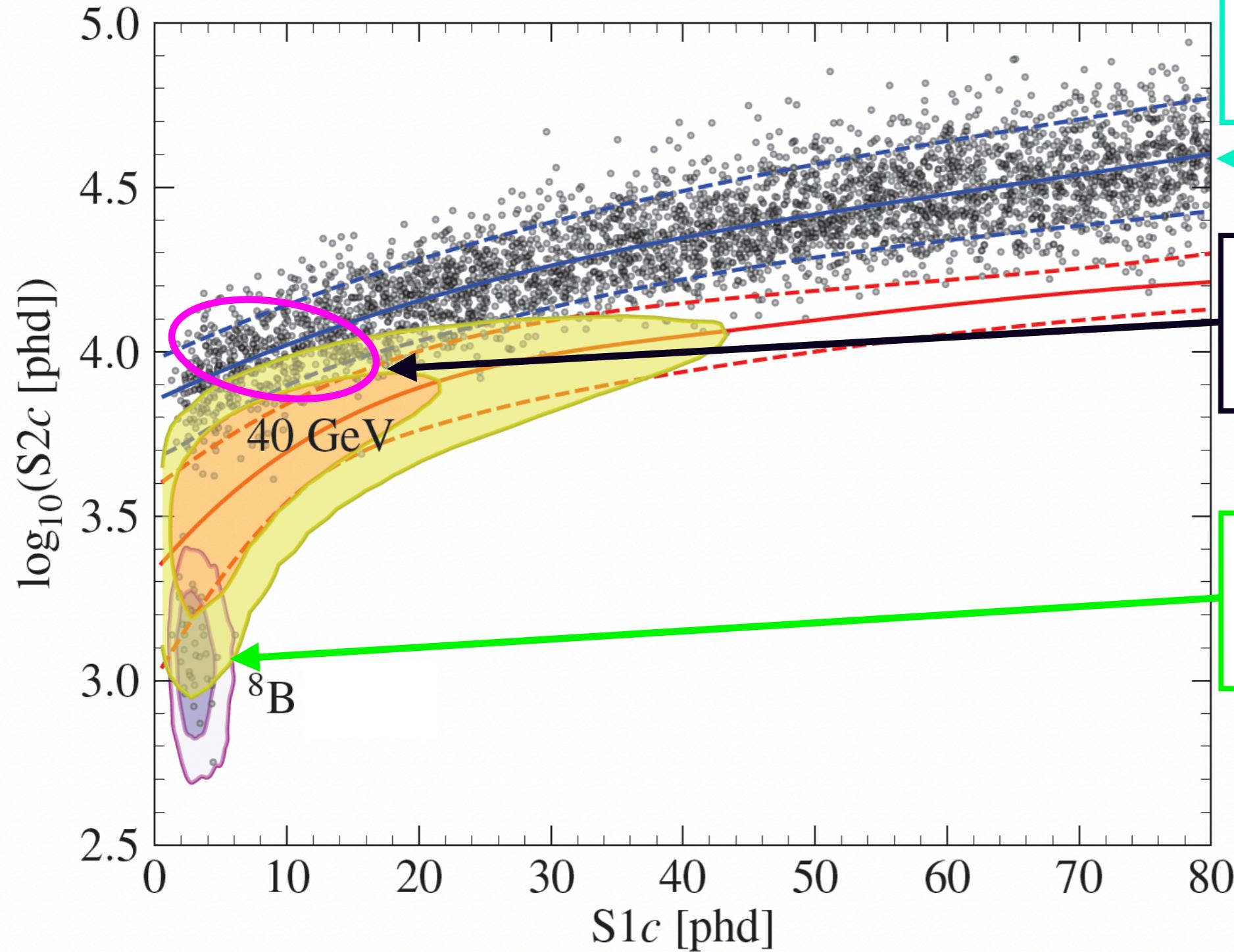


'Naked' 214Pb β decays are the main WIMP background



SR1 : Backgrounds

Other backgrounds considered:



^{212}Pb (β), ^{85}Kr (β), ^{124}Xe ($\varepsilon\varepsilon$),
 ^{136}Xe ($\beta\beta$), ^{127}Xe (ε),
CC solar neutrinos...

Cosmogenically activated
 $^{37}\text{Ar} (\varepsilon)$ should live around
here

Coherent elastic neutrino-
neutron scattering (CE ν NS)
from ^8B solar neutrinos

SR1 : Backgrounds

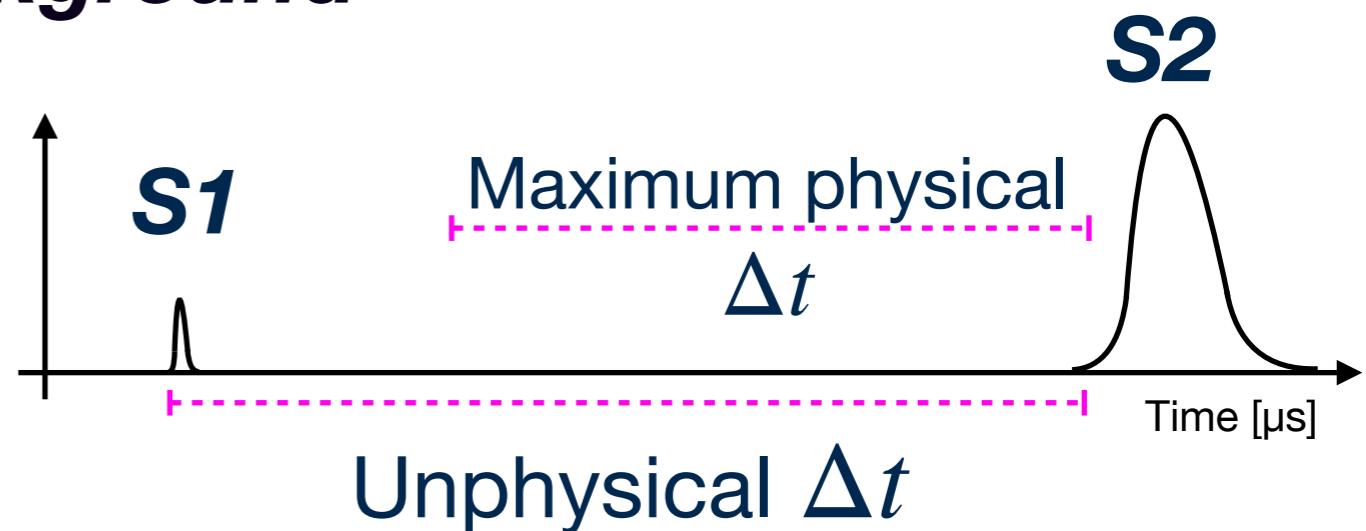
Accidental Coincidence Background

- Unrelated S1s & S2s can accidentally combine to produce single scatter events

SR1 : Backgrounds

Accidental Coincidence Background

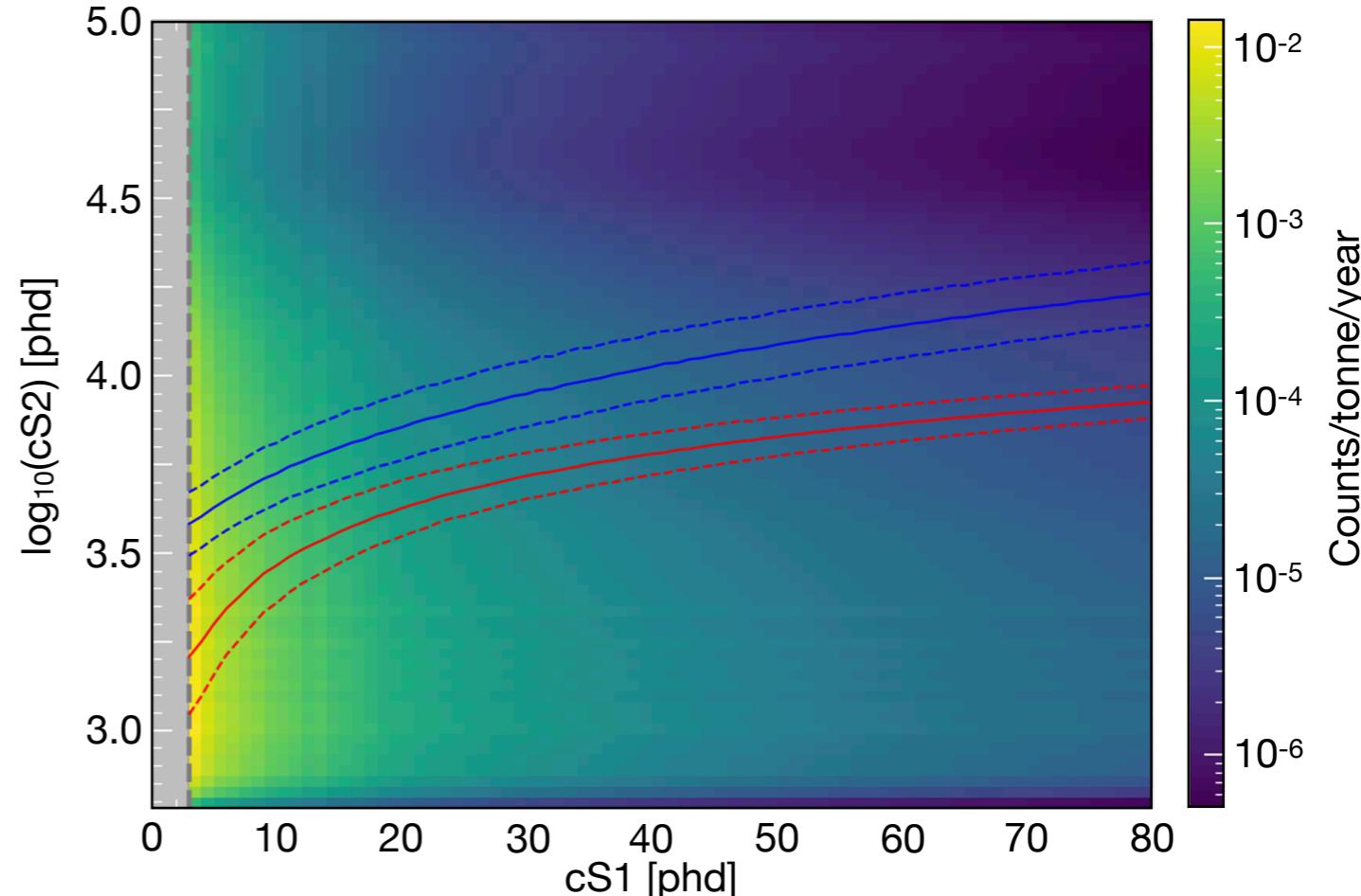
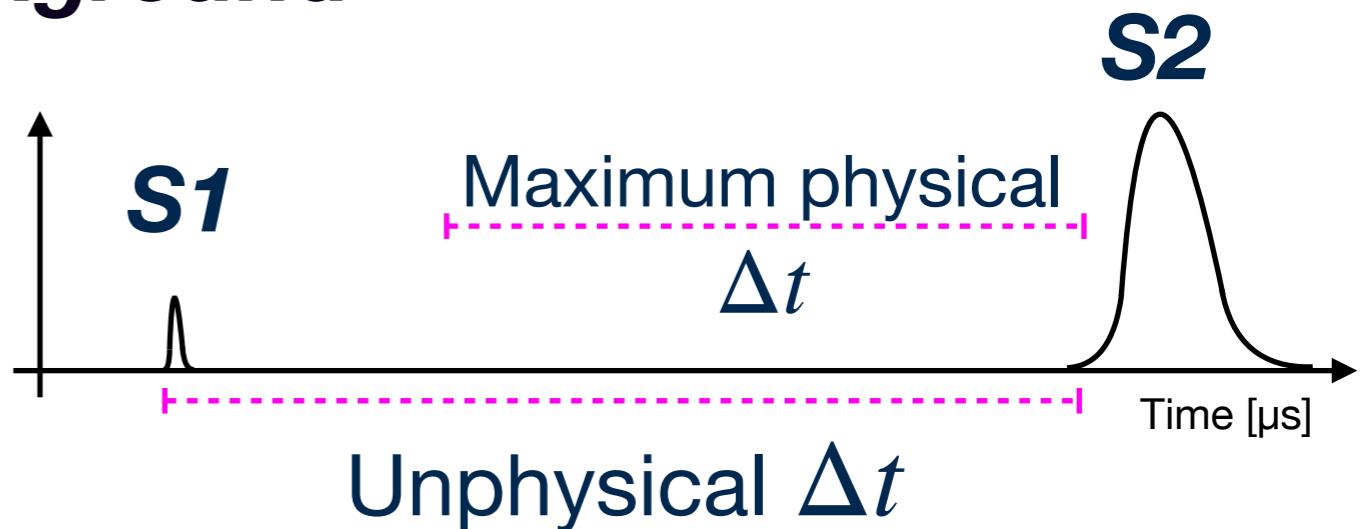
- Unrelated S1s & S2s can accidentally combine to produce single scatter events
- **Rate**: population of definite accidental events with drift time >1 ms



SR1 : Backgrounds

Accidental Coincidence Background

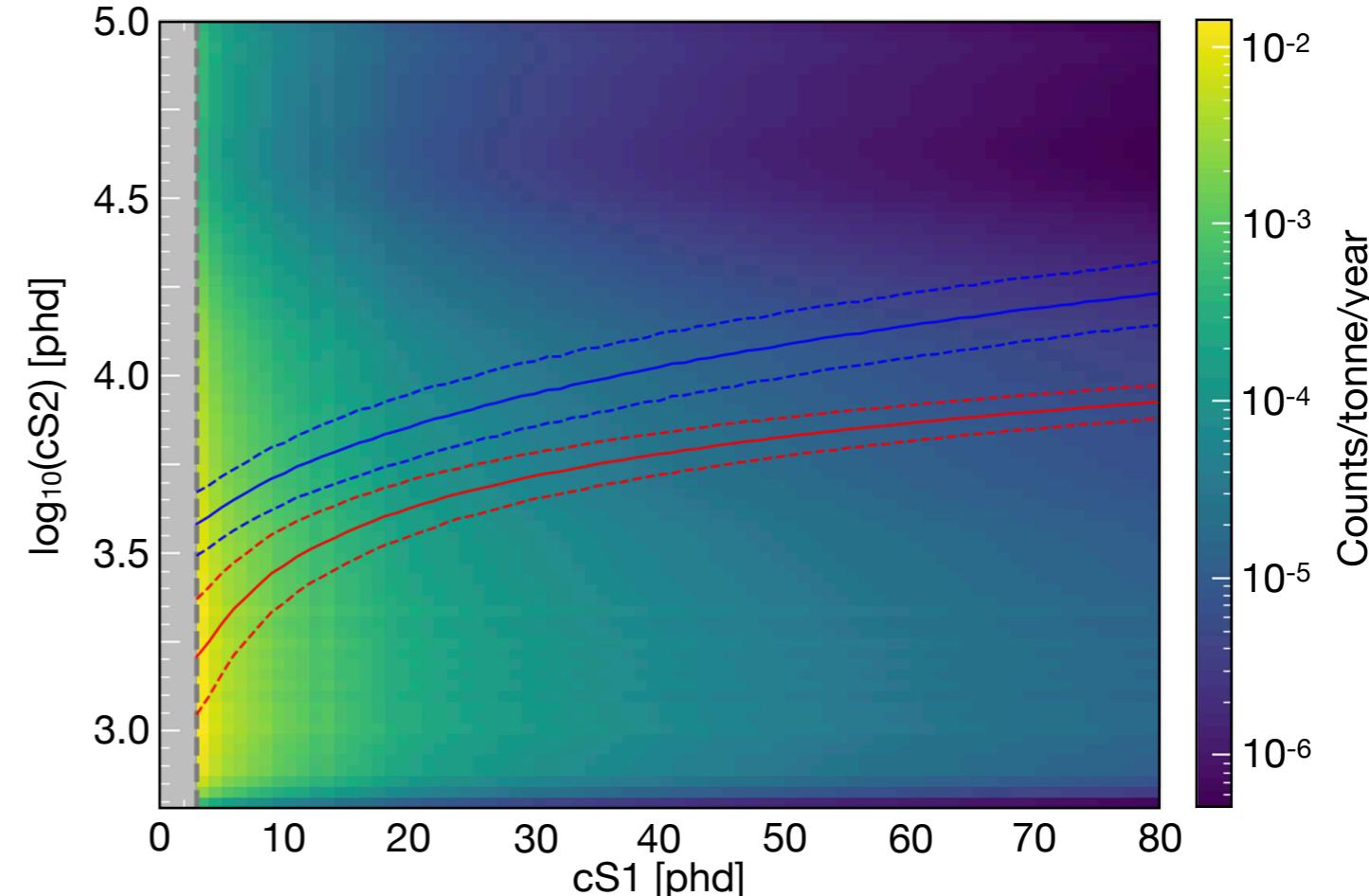
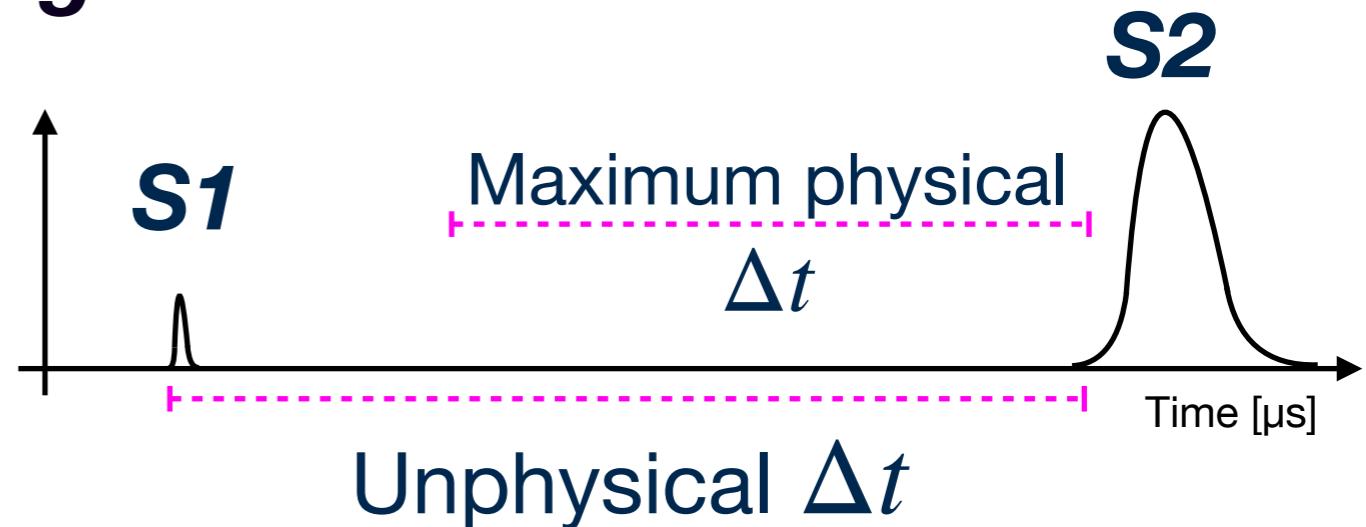
- Unrelated S1s & S2s can accidentally combine to produce single scatter events
- **Rate**: population of definite accidental events with drift time >1 ms
- **Distribution**: fake events constructed from lone S1 & S2 pulse waveforms



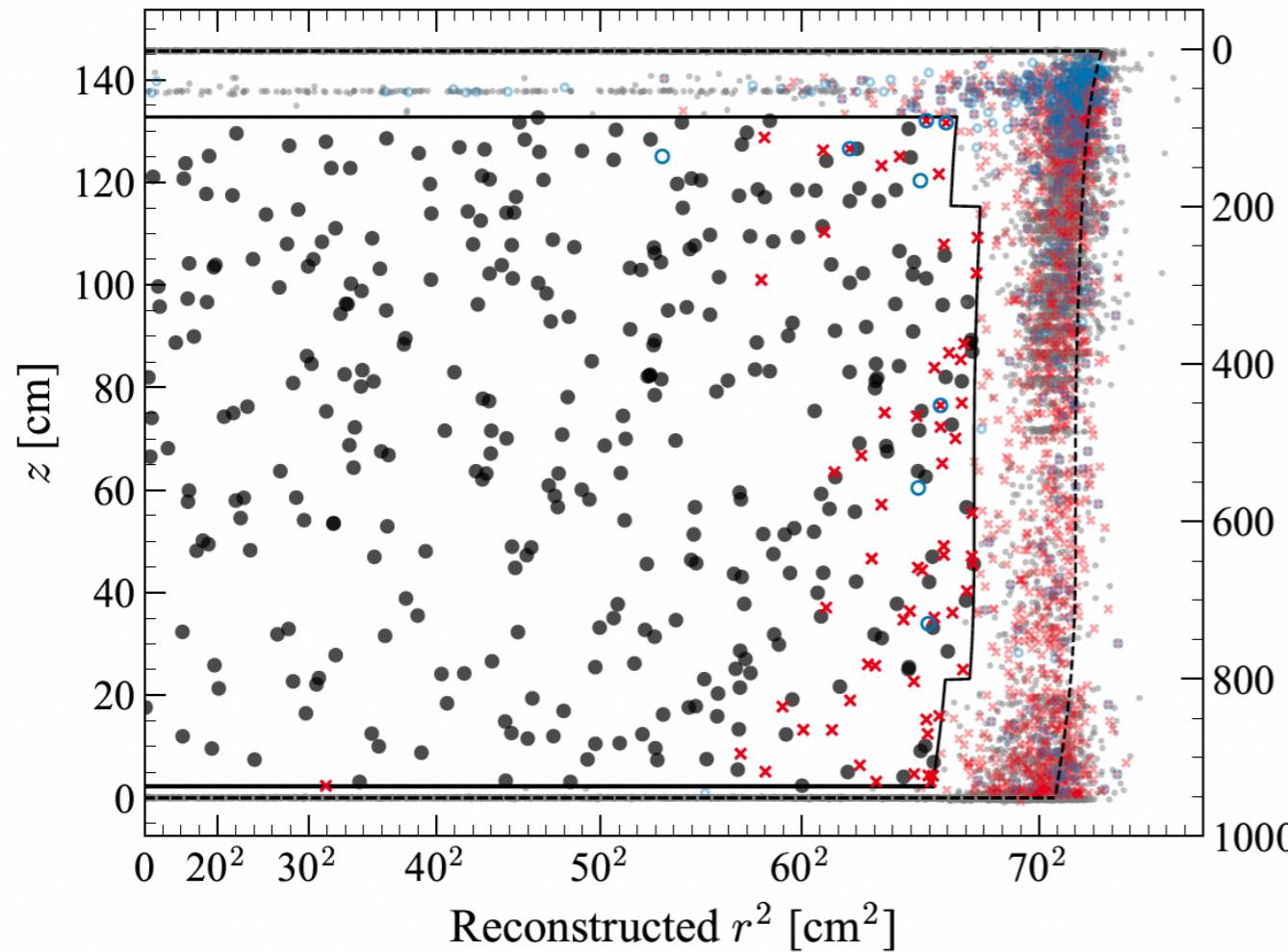
SR1 : Backgrounds

Accidental Coincidence Background

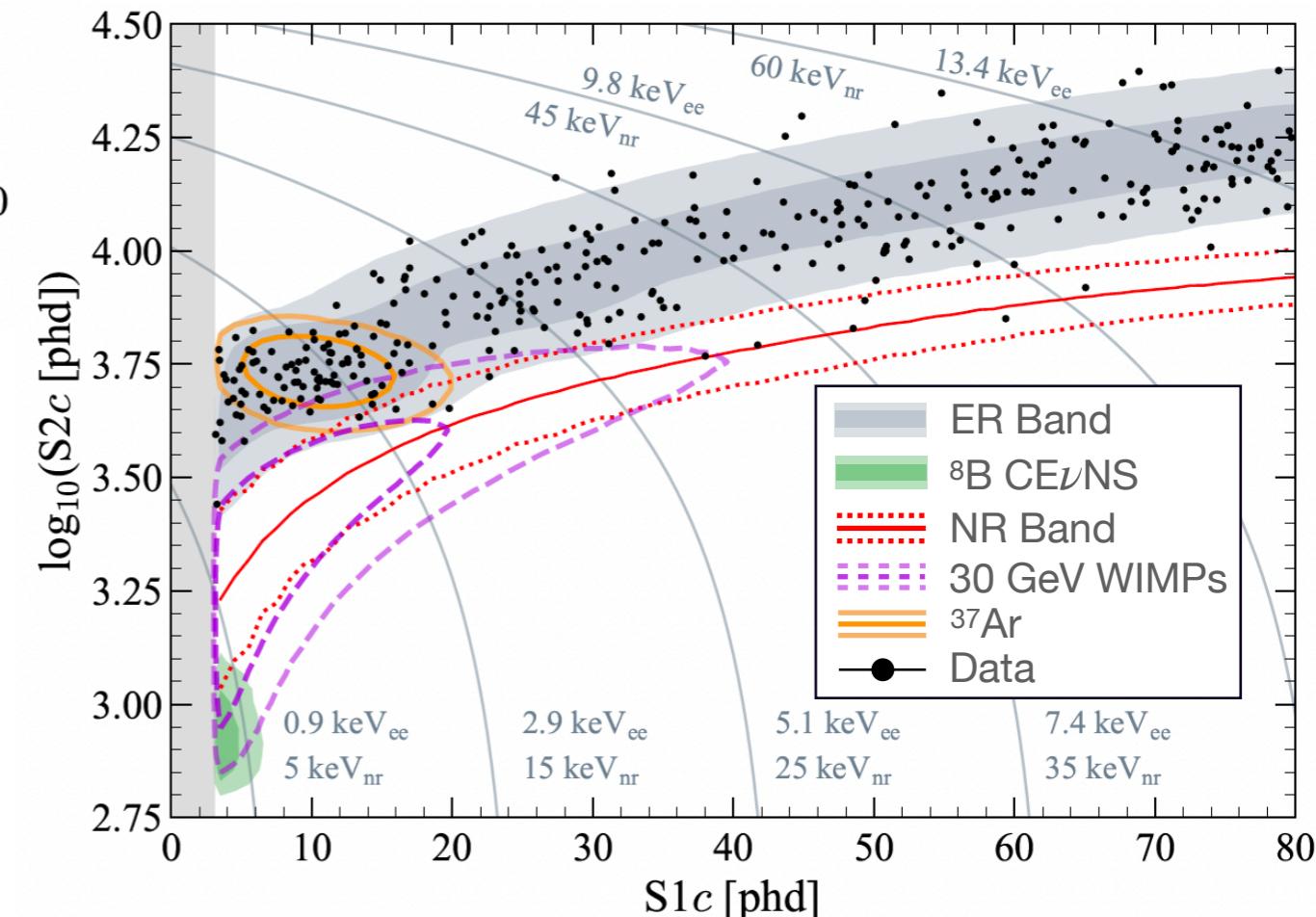
- Unrelated S1s & S2s can accidentally combine to produce single scatter events
- **Rate**: population of definite accidental events with drift time >1 ms
- **Distribution**: fake events constructed from lone S1 & S2 pulse waveforms
- Analysis cuts developed to combat observed pulse/event pathologies
 - >99.5% efficiency in removing accidentals
 - SR1 WIMP search counts: 1.2 ± 0.3



SR1 : Results



- **Inside Fiducial Volume**
- **Outside Fiducial Volume**
- ✗ **Tagged by LXe Skin**
- **Tagged by OD PMTs**

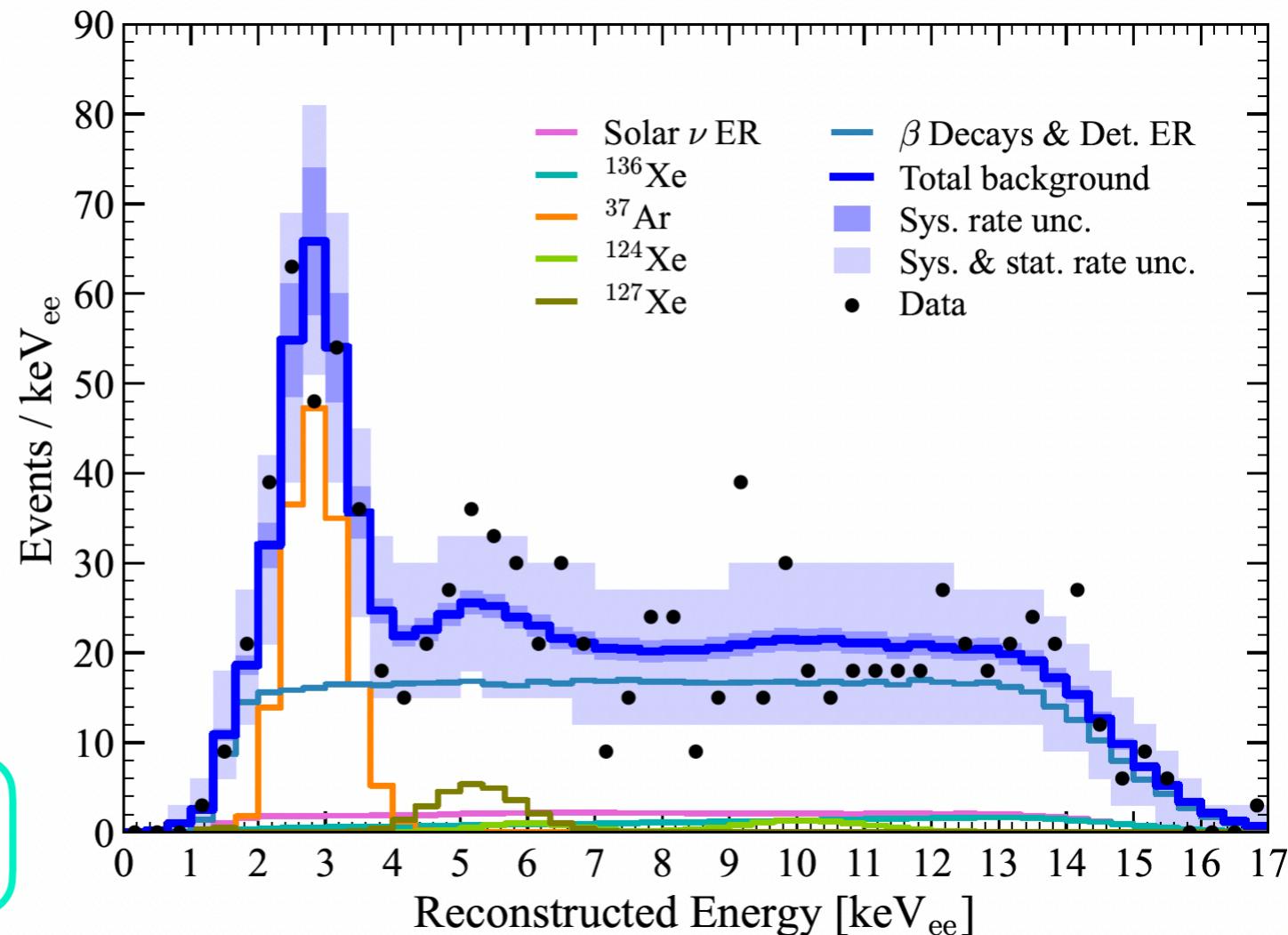


- 335 events remaining after analysis cuts
- Statistical inference with Profile Likelihood Ratio (PLR) method

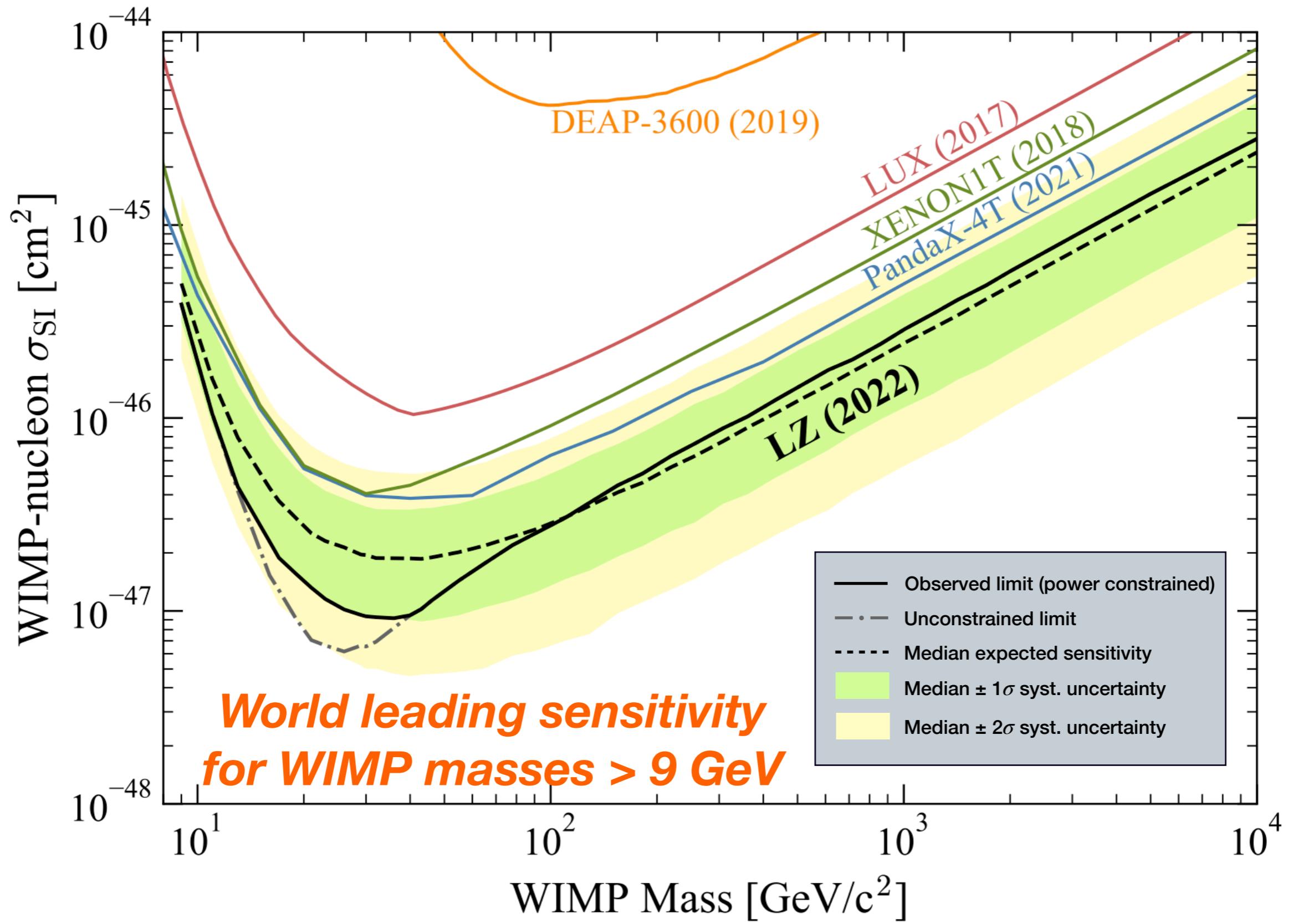
SR1: Results

Data are consistent with background-only hypothesis

Source	Expected Events	Fit Result
β decays + Det ER	215 ± 36	222 ± 16
ν ER	27.1 ± 1.6	27.2 ± 1.6
^{127}Xe	9.2 ± 0.8	9.3 ± 0.8
^{124}Xe	5.0 ± 1.4	5.2 ± 1.4
^{136}Xe	15.1 ± 2.4	15.2 ± 2.4
$^8\text{B} \text{ CE}\nu\text{NS}$	0.14 ± 0.01	0.15 ± 0.01
Accidentals	1.2 ± 0.3	1.2 ± 0.3
Subtotal	273 ± 36	280 ± 16
^{37}Ar	[0, 288]	$52.5^{+9.6}_{-8.9}$
Detector neutrons	$0.0^{+0.2}$	$0.0^{+0.2}$
30 GeV/c ² WIMP	...	$0.0^{+0.6}$
Total	...	333 ± 17
Total Observed Events		335



SR1 : Results



Low Energy ERs

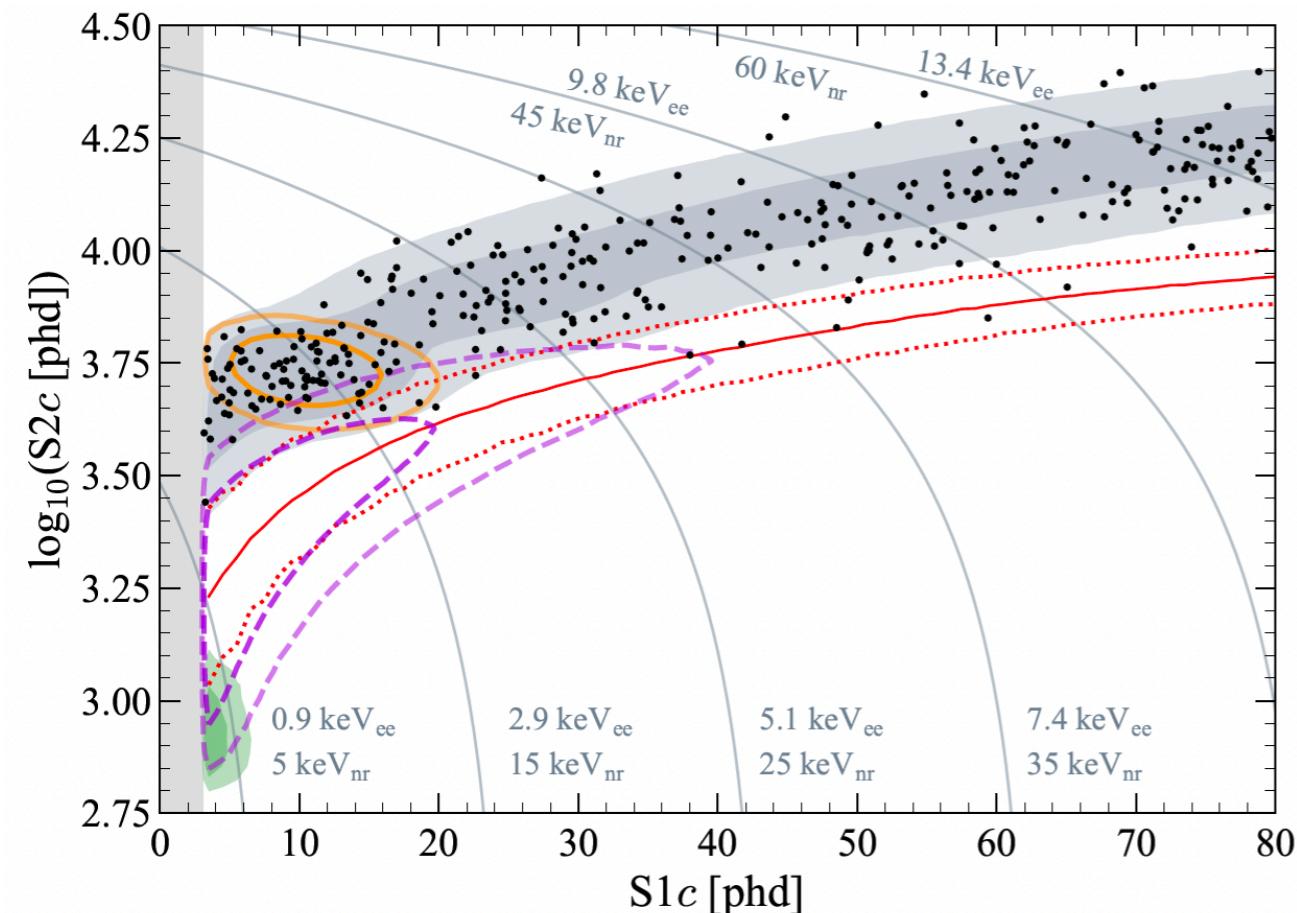
A search for new physics in low-energy electron recoils from the first LZ exposure

J. Aalbers,^{1, 2} D.S. Akerib,^{1, 2} A.K. Al Musalhi,³ F. Alder,⁴ C.S. Amarasinghe,⁵ A. Ames,^{1, 2} T.J. Anderson,^{1, 2} N. Angelides,⁶ H.M. Araújo,⁶ J.E. Armstrong,⁷ M. Arthurs,^{1, 2} A. Baker,⁶ S. Balashov,⁸ J. Bang,⁹ J.W. Bargemann,¹⁰ A. Baxter,¹¹ K. Beattie,¹² P. Beltrame,⁴ T. Benson,¹³ A. Bhatti,⁷ A. Biekert,^{12, 14} T.P. Biesiadzinski,^{1, 2} H.J. Birch,⁵ G.M. Blockinger,¹⁵ B. Boxer,¹⁶ C.A.J. Brew,⁸ P. Brás,¹⁷ S. Burdin,¹¹ M. Buuck,^{1, 2} M.C. Carmona-Benitez,¹⁸ C. Chan,⁹ A. Chawla,¹⁹ H. Chen,¹² J.J. Cherwinka,¹³ N.I. Chott,²⁰ M.V. Converse,²¹ A. Cottle,^{3, 22} G. Cox,^{18, 23} D. Curran,²³ C.E. Dahl,^{22, 24} A. David,⁴ J. Delgaudio,²³ S. Dey,³

The LUX-ZEPLIN (LZ) experiment is a dark matter detector centered on a dual-phase xenon time projection chamber. We report searches for new physics appearing through few-keV-scale electron recoils, using the experiment's first exposure of 60 live days and a fiducial mass of 5.5 t. The data are found to be consistent with a background-only hypothesis, and limits are set on models for new physics including solar axion electron coupling, solar neutrino magnetic moment and millicharge, and electron couplings to galactic axion-like particles and hidden photons. Similar limits are set on weakly interacting massive particle (WIMP) dark matter producing signals through ionized atomic states from the Migdal effect.

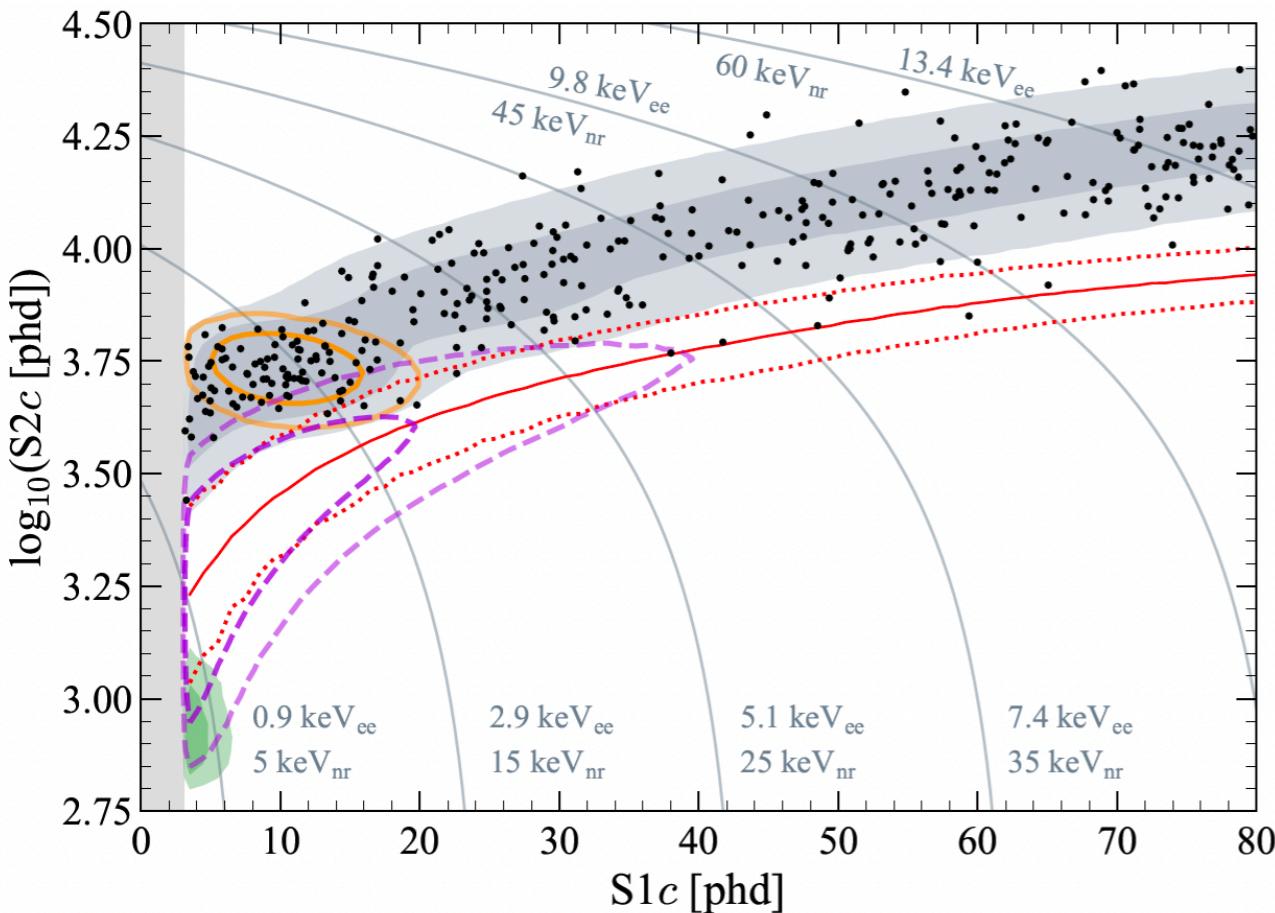
arXiv preprint: [2307.15753](https://arxiv.org/abs/2307.15753)

Low Energy ERs



Anything interesting happening here?

Low Energy ERs

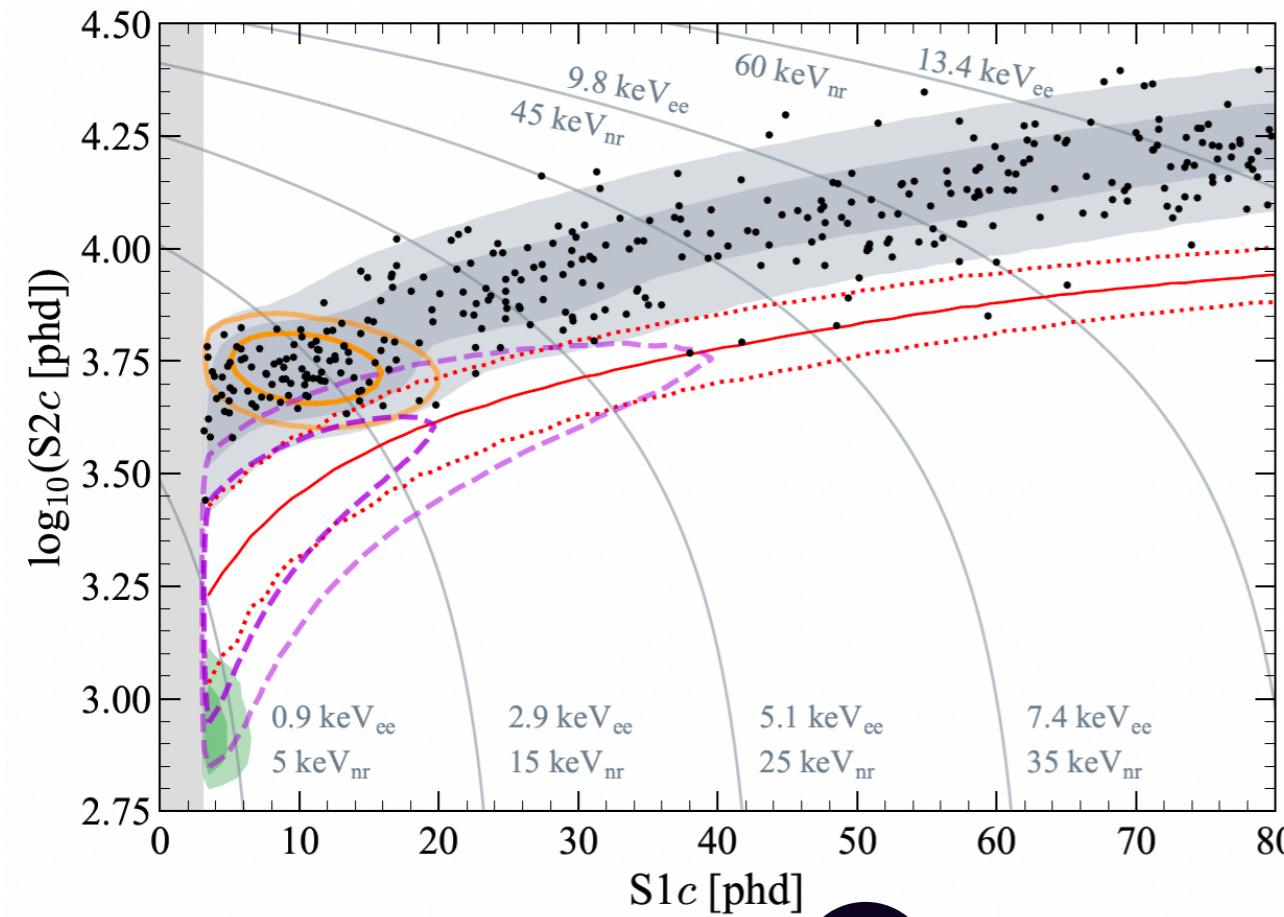


Neutrino Electromagnetic Moments

The diagram shows a stylized symbol for a neutrino (ν) enclosed in a circle. Two arrows point away from the circle: one cyan arrow points upwards, labeled μ_ν (neutrino magnetic moment), and one magenta arrow points downwards, labeled q_ν (neutrino millicharge).

$$\frac{d\sigma}{dE_r} = \frac{d\sigma_{EW}}{dE_r} + \mu_\nu^2 \mathcal{O}(E_r^{-1}) + q_\nu^2 \mathcal{O}(E_r^{-2})$$

Low Energy ERs



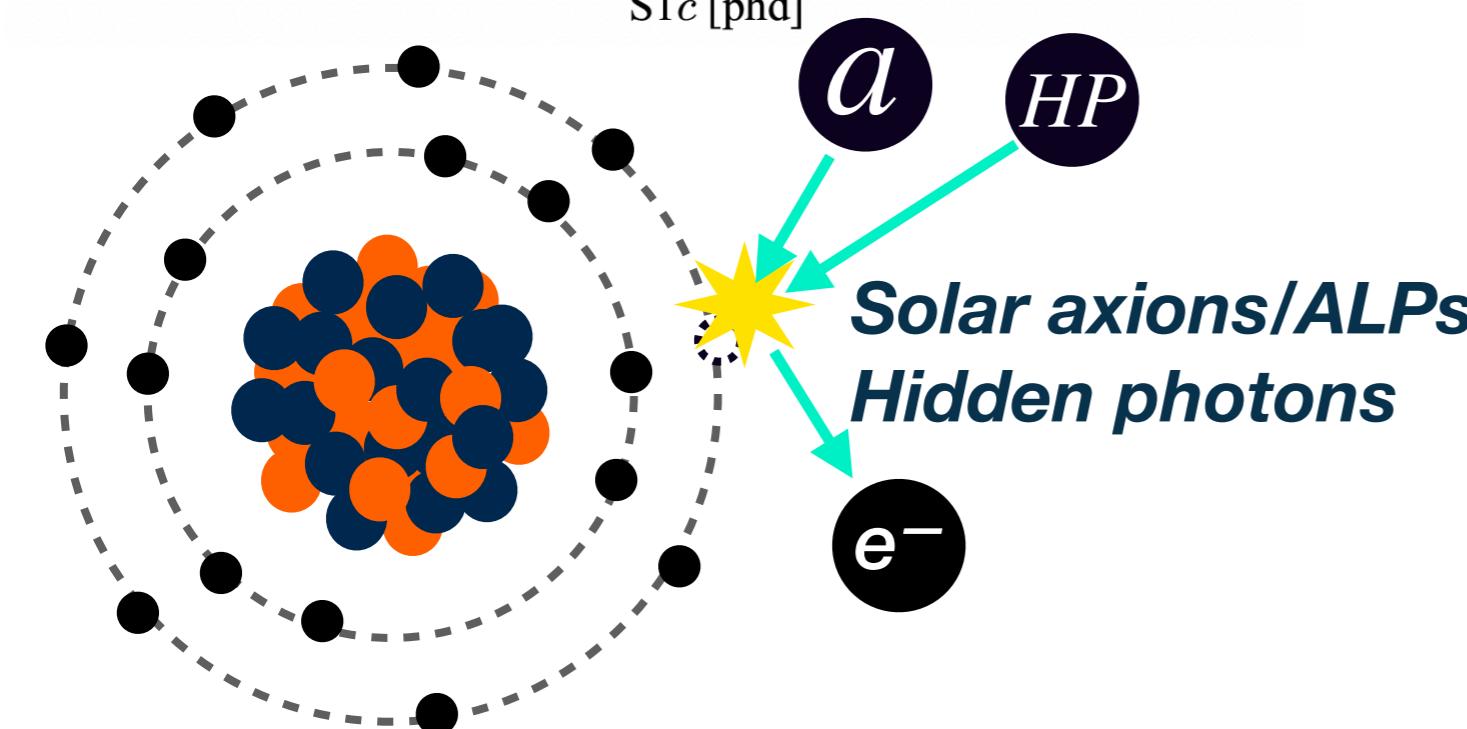
Neutrino Electromagnetic Moments



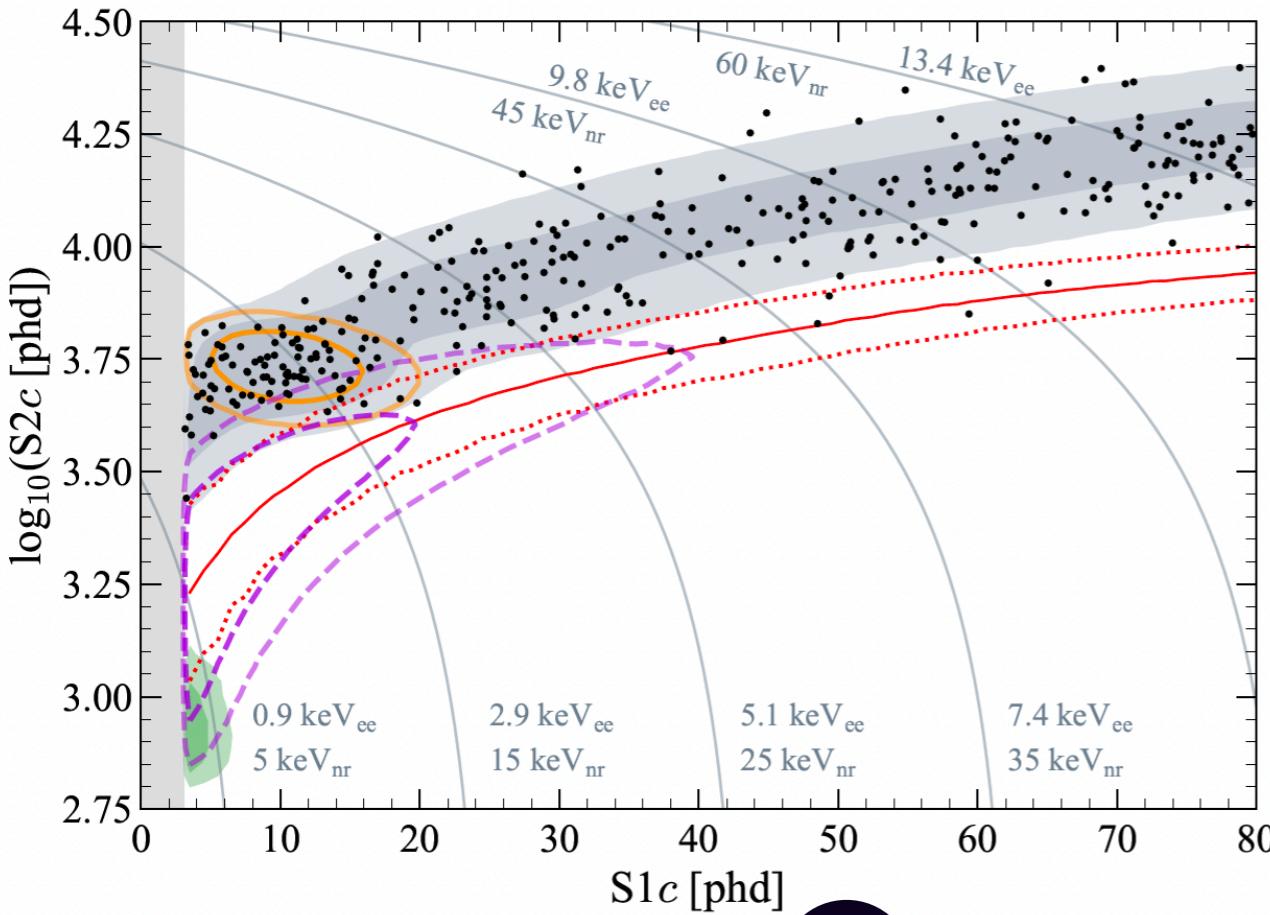
μ_ν (neutrino magnetic moment)

q_ν (neutrino millicharge)

$$\frac{d\sigma}{dE_r} = \frac{d\sigma_{EW}}{dE_r} + \mu_\nu^2 \mathcal{O}(E_r^{-1}) + q_\nu^2 \mathcal{O}(E_r^{-2})$$



Low Energy ERs



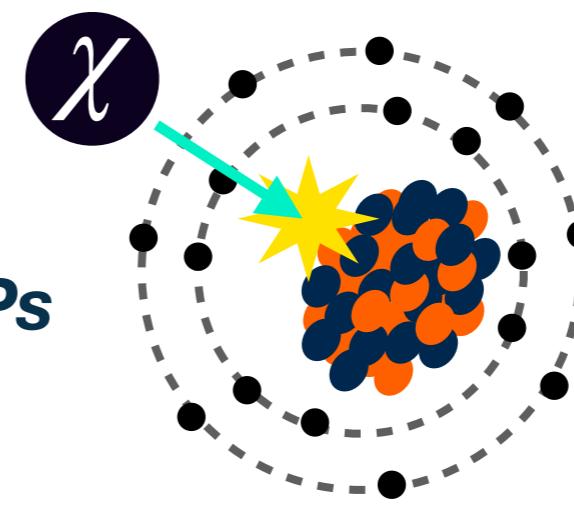
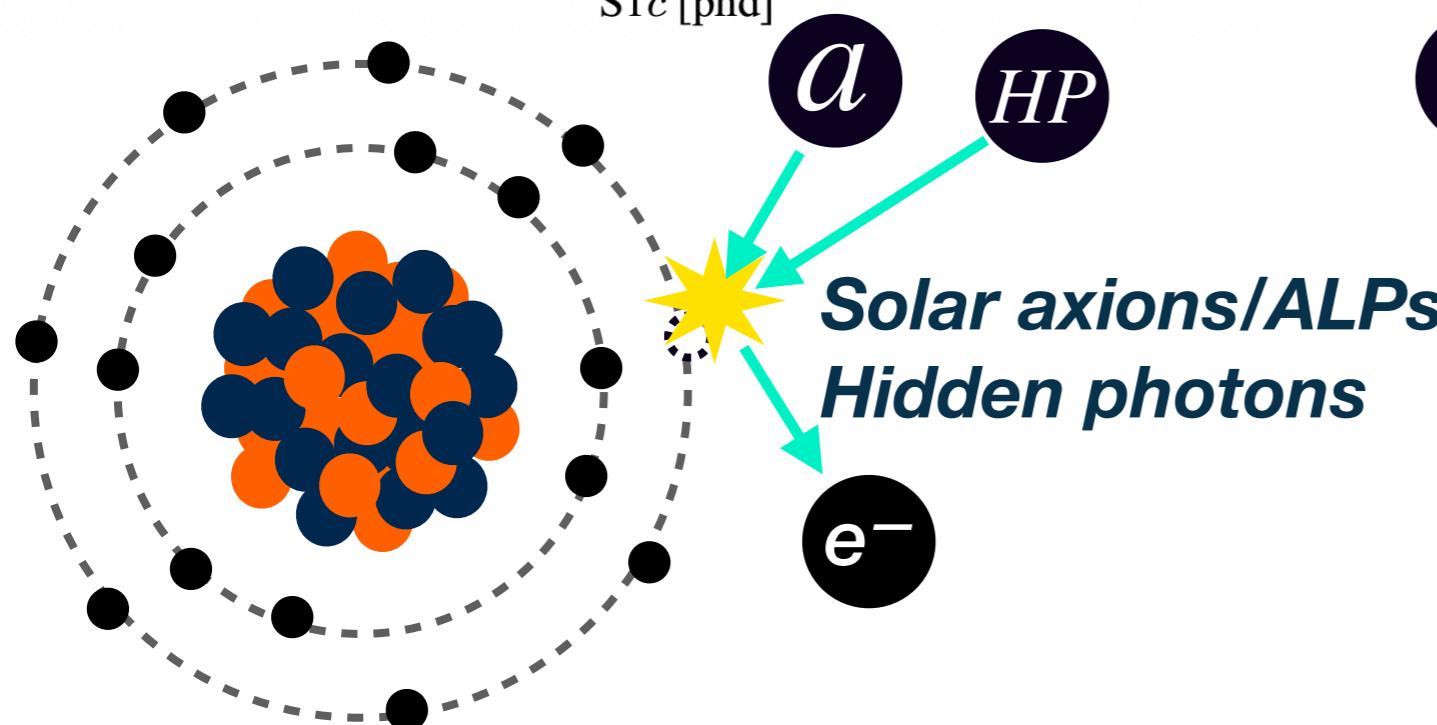
Neutrino Electromagnetic Moments



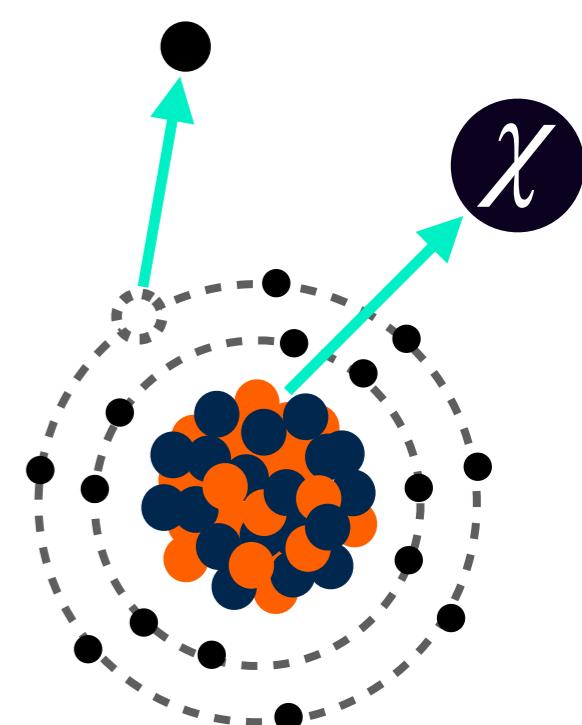
μ_ν (neutrino magnetic moment)

q_ν (neutrino millicharge)

$$\frac{d\sigma}{dE_r} = \frac{d\sigma_{EW}}{dE_r} + \mu_\nu^2 \mathcal{O}(E_r^{-1}) + q_\nu^2 \mathcal{O}(E_r^{-2})$$



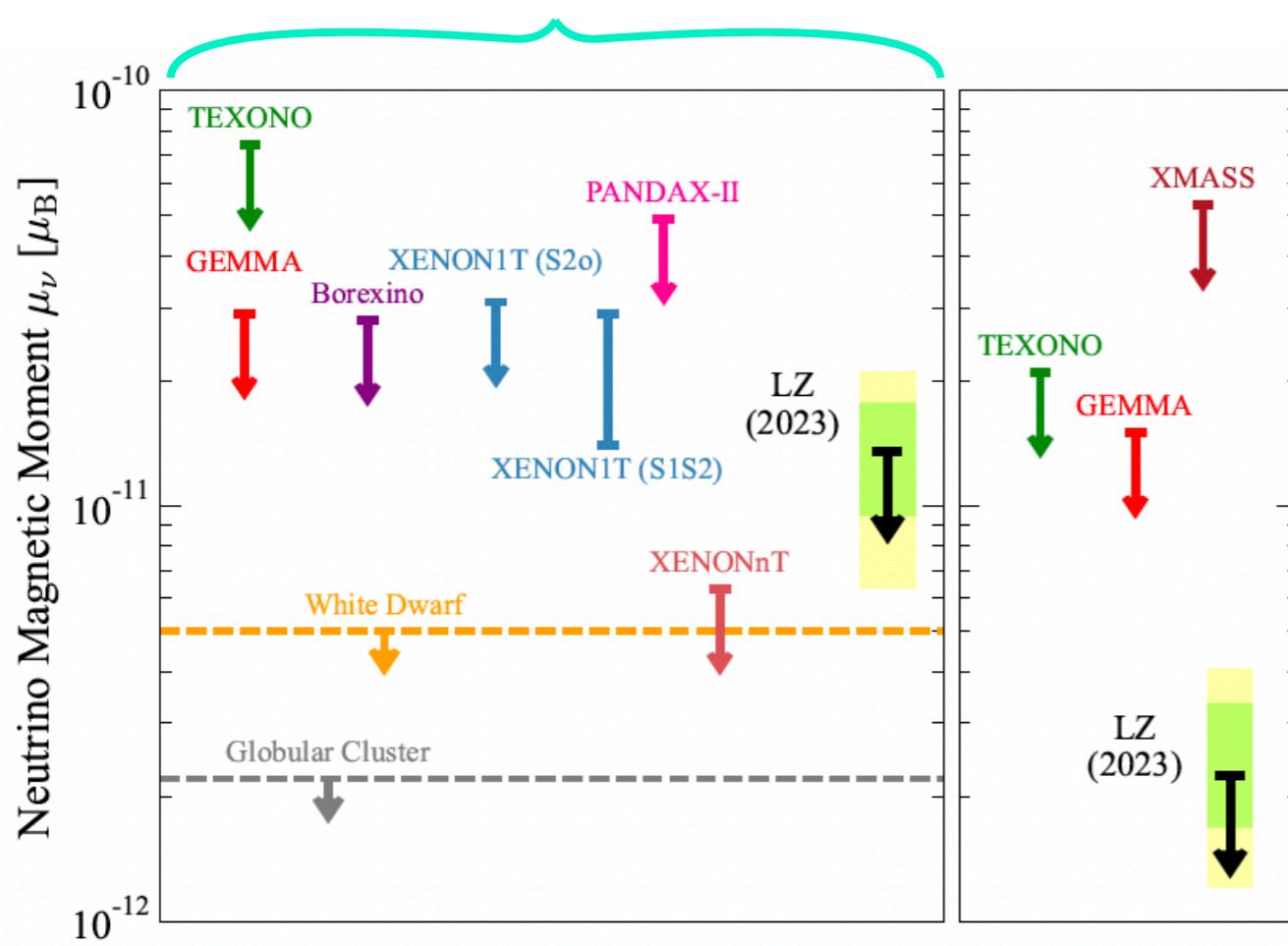
Migdal Effect



Low Energy ERs

**Neutrino magnetic moment
most stringently constrained by
astrophysical observations**

LZ upper limit: $1.36 \times 10^{-11} \mu_B$



Neutrino Electromagnetic Moments

μ_ν (neutrino magnetic moment)

q_ν (neutrino millicharge)



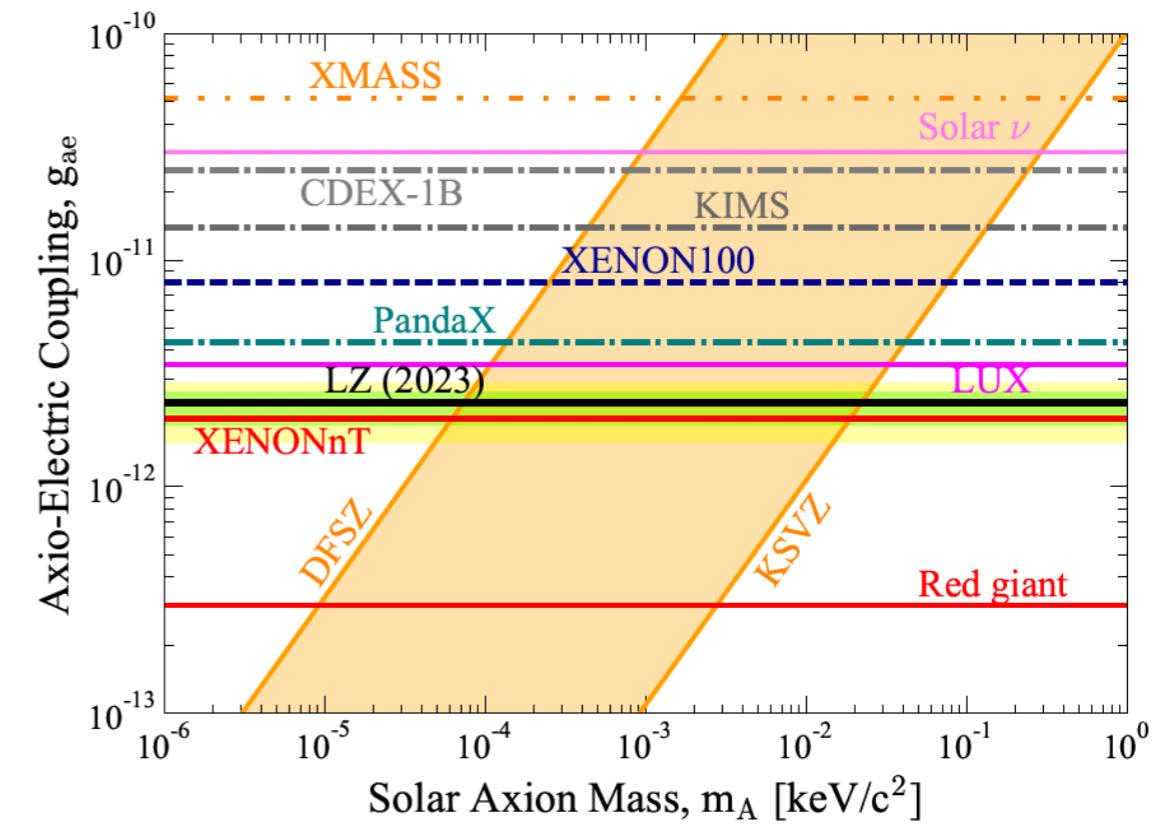
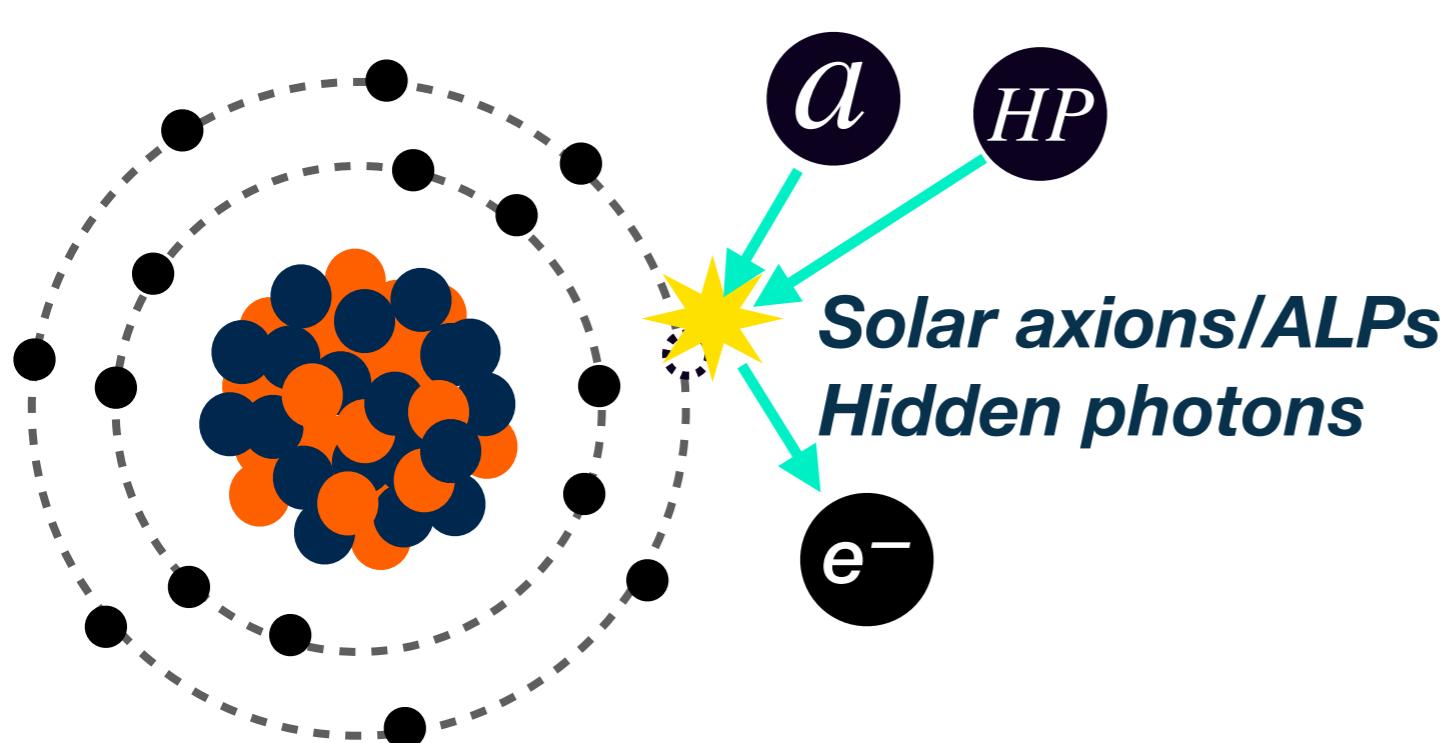
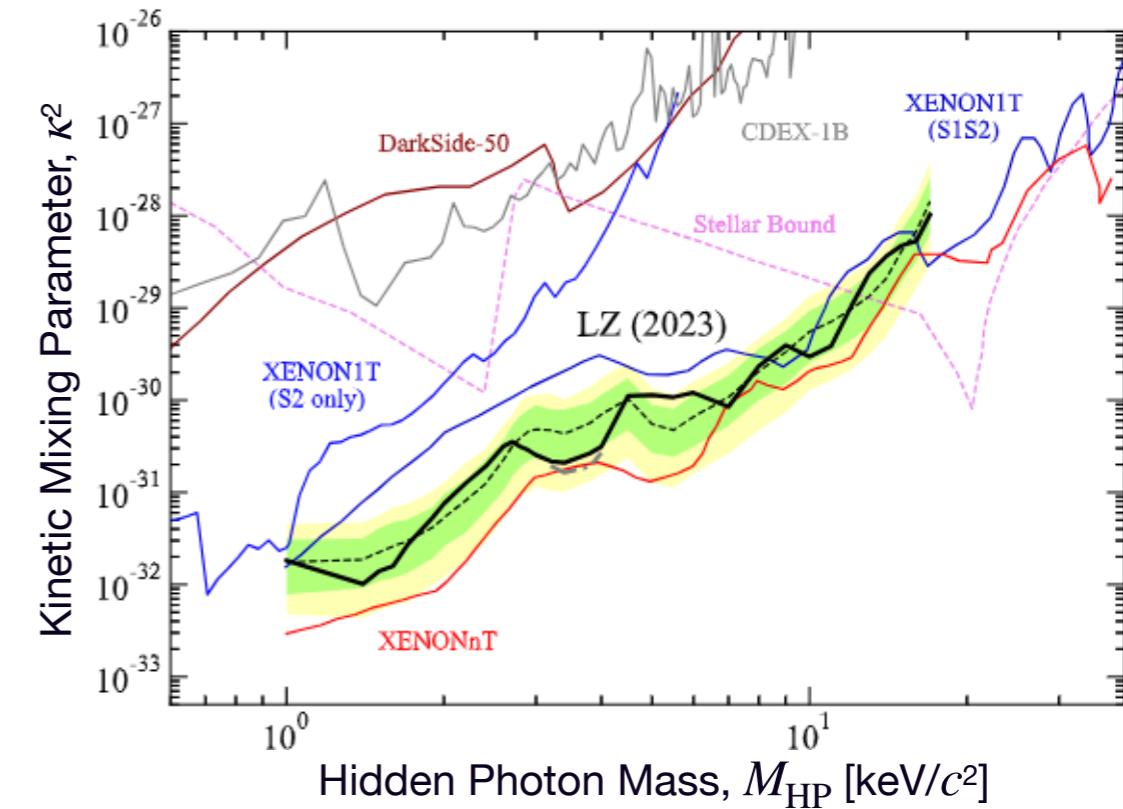
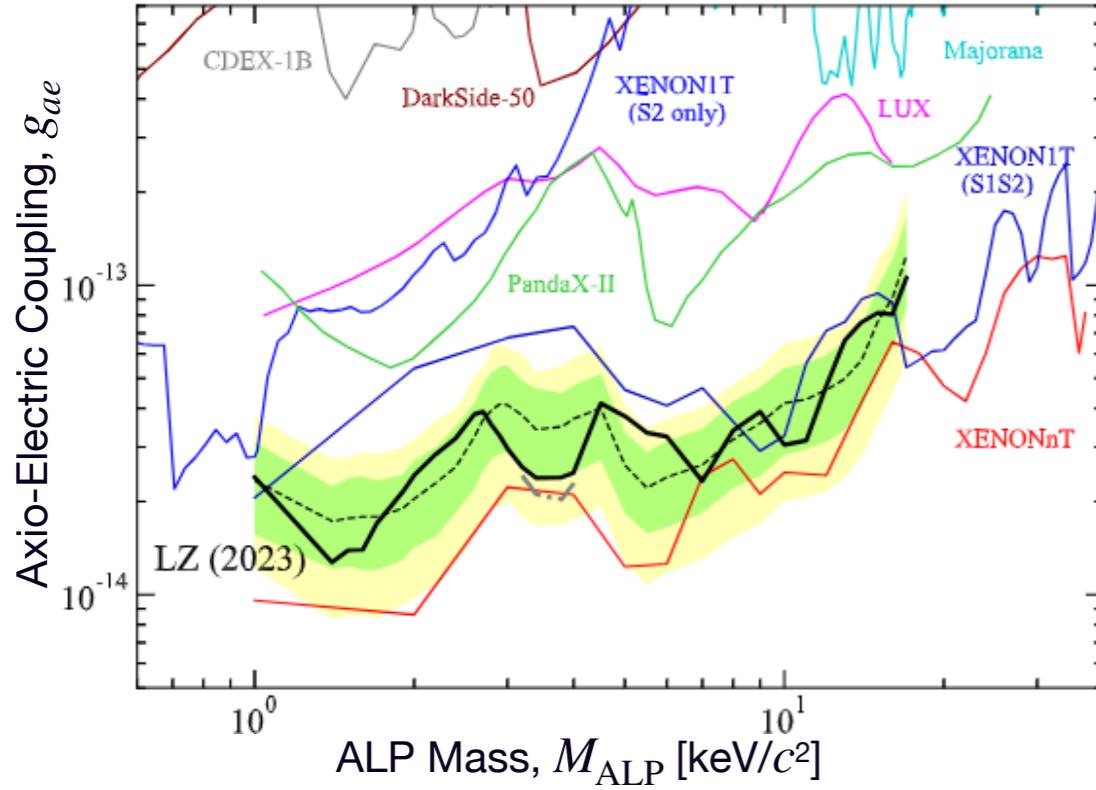
$$\frac{d\sigma}{dE_r} = \frac{d\sigma_{EW}}{dE_r} + \mu_\nu^2 \mathcal{O}(E_r^{-1}) + q_\nu^2 \mathcal{O}(E_r^{-2})$$

**Leading neutrino millicharge
observed**

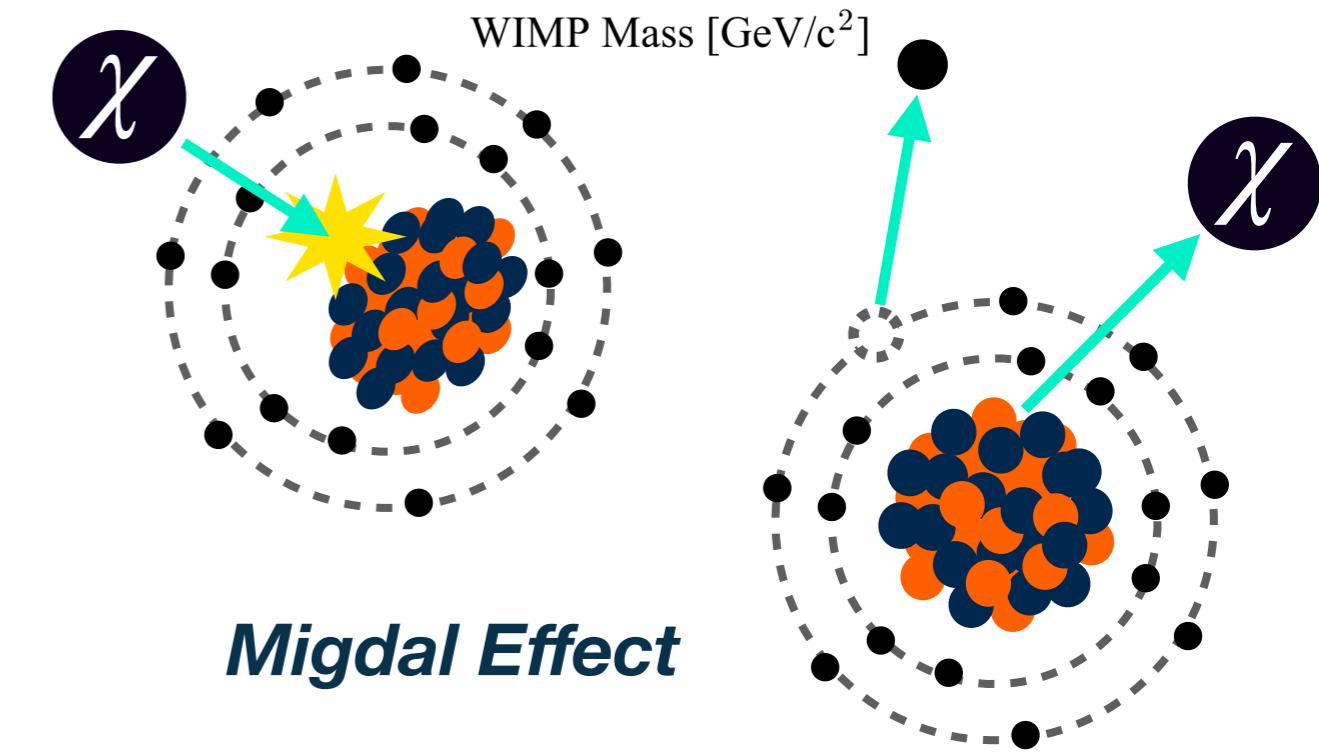
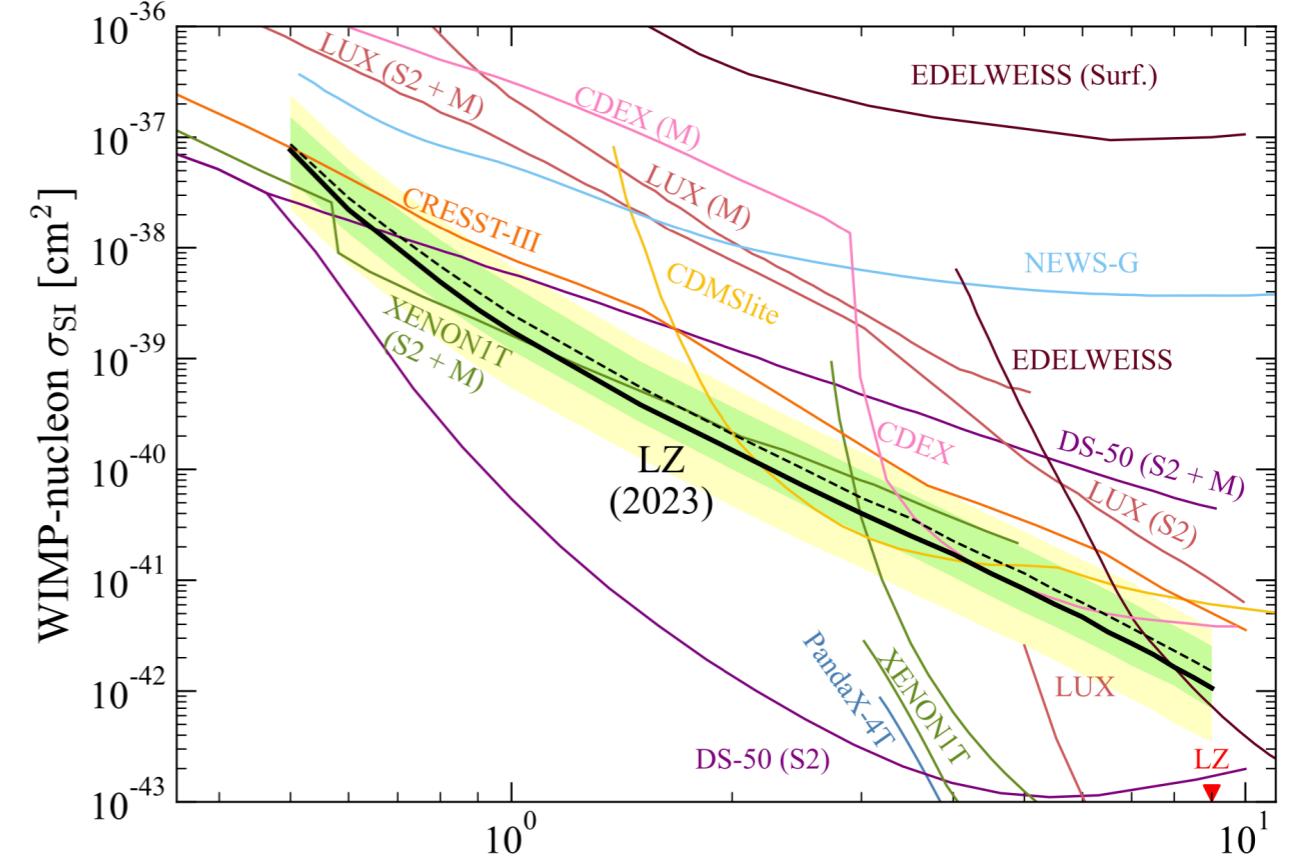
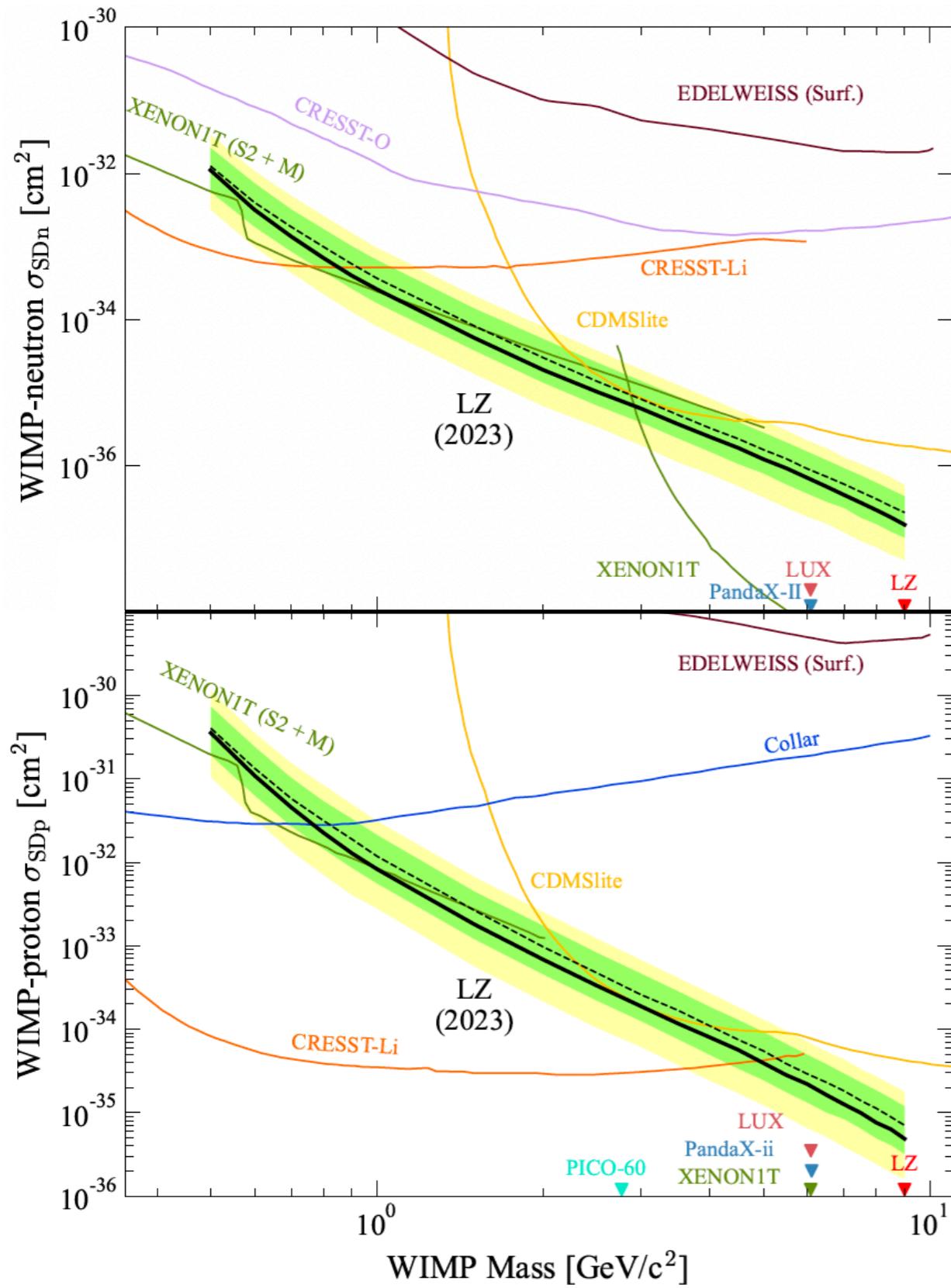
LZ upper limit: $2.24 \times 10^{-13} e_0$

$e_0 \equiv$ electron charge
 $\mu_B \equiv$ Bohr magneton

Low Energy ERs



Low Energy ERs



In The Pipeline

Nonrelativistic Effective Field Theory

Searches utilizing a model-agnostic framework for WIMP-nucleon interactions in an extended energy region

Ultra Heavy Dark Matter Search

Search for Planck scale ($M \sim 10^6$ GeV) high multiplicity interactions in LZ detector. Multiply Interacting Massive Particles—MIMPs

Radon Tagging Veto

An investigation into mitigating the ^{214}Pb background using a tagged coincidence veto with ^{222}Rn alpha decays

Muon Flux Measurements

Characterization of the muon rate during SR1, with comparison to current simulation models

Conclusion

First results leading the way for WIMPs

*Most stringent cross-section upper limits for SI scatters for $m > 9$ GeV.
6% of total planned exposure*

Low energy ER searches

LZ sets world leading upper limits in parts of SD-neutron Migdal parameter space, as well as for neutrino millicharge

Onward to 1000 live days

LZ continues to take data for a planned exposure of 1000 live days, pushing into discovery territory for WIMPs and with significantly improved expected sensitivity for ER physics signals



Thank you!

