

Angeles, De Angelis, Forshaw, Plätzer, Seymour – JHEP 05 (2018) 044 Gieseke, Kirchgaesser, Plätzer, Siodmok – arXiv:1808.06770 De Angelis, Forshaw, Plätzer – arXiv:181y.xxxxx Forshaw, Holguin, Plätzer – arXiv:181y.xxxxx

More on soft gluons beyond leading colour

Simon Plätzer

Particle Physics, University of Vienna

at the HARPS Meeting Genova | 13 September 2018



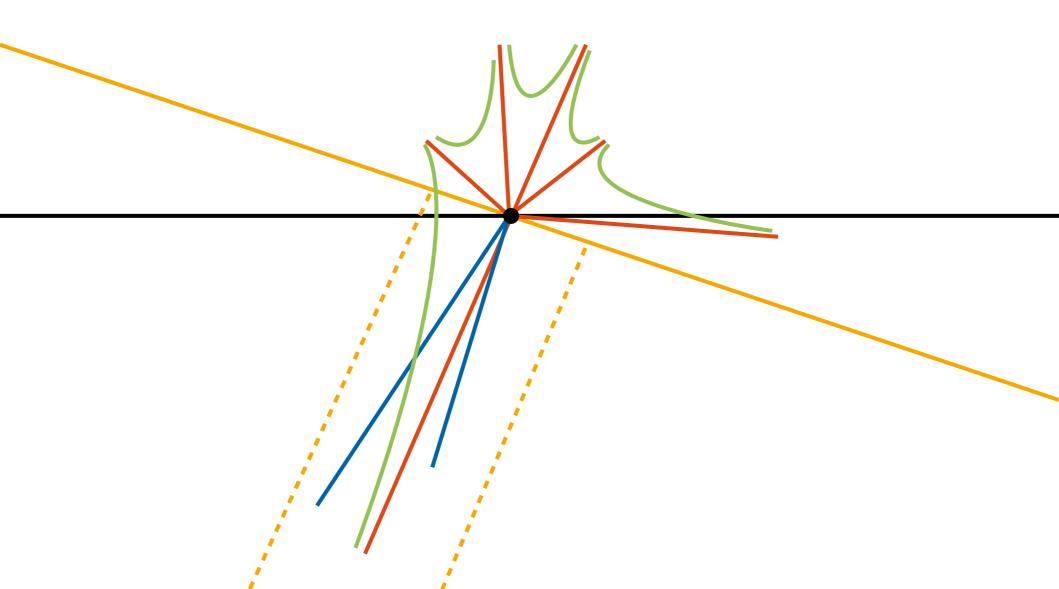


State of the art



Resummation of non-global observables is with dipole cascades in the large-N limit.

[Dasgupta, Salam – Phys.Lett. B512 (2001) 323] [Balsinger, Becher, Shao – arXiv:1803.07045]

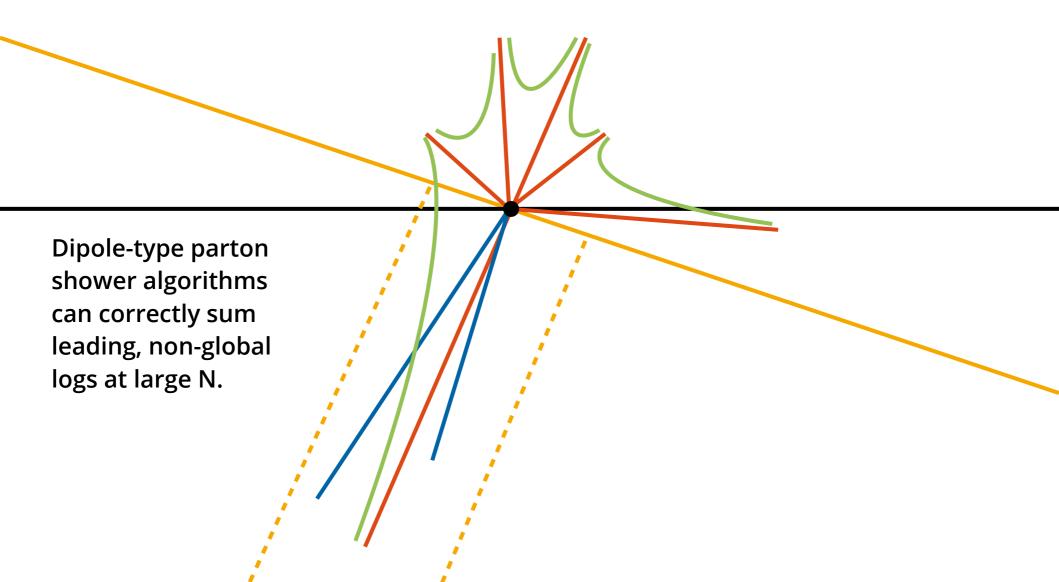


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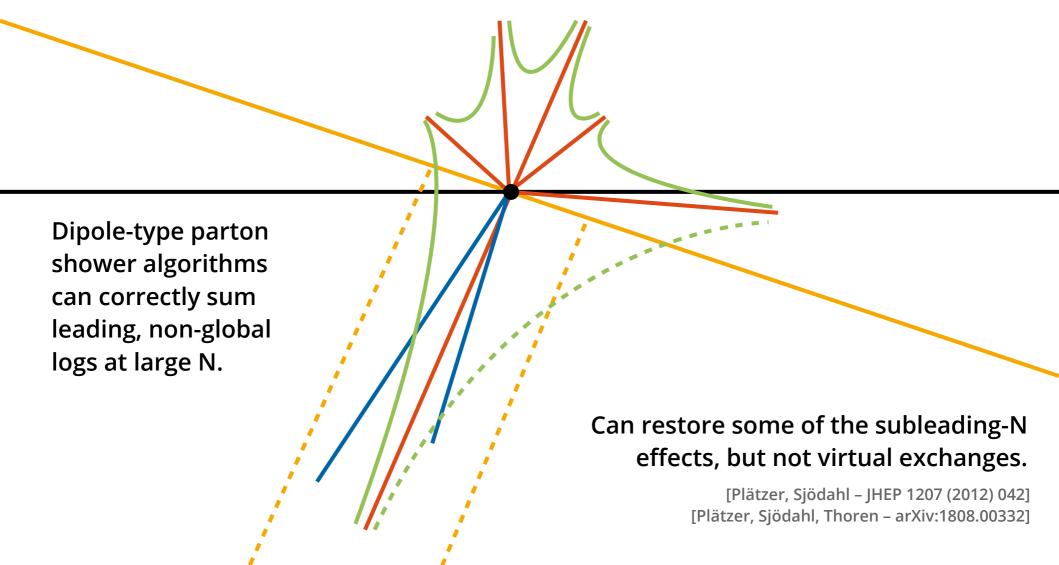


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Strategy and Goals



Seek a framework to systematically address and improve **a new kind of parton shower algorithm**, not relying on ad-hoc constructions, treating colour exactly as far as possible.

[also see Nagy, Soper]

Non-global observables are a unique playground: At large-N they provide a clean way of deriving a dipole-type parton shower, but the origin of the method used is much more general.

$$\sigma = \sum_n \int \mathrm{Tr} \left[\mathbf{A}_n(\mu) \right] \ u(p_1,...,p_n) \ \mathrm{d}\phi_n$$

$$\mathbf{A}_n(\mu) = |\mathcal{M}_n(\mu)\rangle \langle \mathcal{M}_n(\mu)|$$
 Evolved `density operator` Observable Phase space



[Angeles, De Angelis, Forshaw, Plätzer, Seymour – JHEP 05 (2018) 044]

Successive virtual evolution/emission combinations down to an infrared cutoff, which will need to be removed at the end. Observable value itself can act as a cutoff, if fully inclusive for gluon energies below this scale.

$$\mathbf{A}_n(E) = \mathbf{V}(E, E_n) \mathbf{D}_n \mathbf{A}_{n-1}(E_n) \mathbf{D}_n^{\dagger} \mathbf{V}^{\dagger}(E, E_n) \theta(E - E_n)$$

$$\mathbf{V}_n(E,Q) = \operatorname{P} \exp \left(-\int_E^Q \frac{\mathrm{d}q}{q} \mathbf{\Gamma}_n(q)\right) \qquad \mathbf{D}_n = \sum_{i=1}^{n-1} \frac{p_i \cdot \epsilon^*(p_n)}{p_i \cdot p_n} \mathbf{T}_i$$

Non-probabilistic evolution at the amplitude level, keeping full colour structure, virtual corrections encoded in anomalous dimension

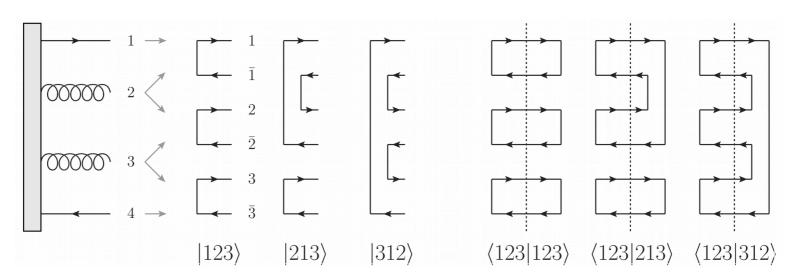
$$\mathbf{\Gamma}_n = \frac{\alpha_s}{\pi} \sum_{i < j} \int d\Omega \frac{n_i \cdot n_j}{n_i \cdot n \ n \cdot n_j} (-\mathbf{T}_i \cdot \mathbf{T}_j)$$

Colour Flows



Express amplitudes in combinations of fundamental/anti-fundamental indices:

$$|\mathcal{M}\rangle = \sum_{\sigma} \mathcal{M}_{\sigma} |\sigma\rangle \qquad |\sigma\rangle = \begin{vmatrix} 1 & \cdots & n \\ \sigma(1) & \cdots & \sigma(n) \end{vmatrix} = \delta_{\bar{\alpha}_{\sigma(1)}}^{\alpha_1} \cdots \delta_{\bar{\alpha}_{\sigma(n)}}^{\alpha_n}$$



Non-orthogonal basis:

$$1 = \sum_{\sigma} |\sigma\rangle[\sigma] \qquad [\sigma|\tau\rangle = \langle\sigma|\tau] = \delta_{\tau\sigma} \qquad \text{Tr}[\mathbf{A}] = \sum_{\tau,\sigma} [\tau|\mathbf{A}|\sigma]\langle\sigma|\tau\rangle$$

Also overcomplete ... but computationally very handy: It's all about permutations.

Resumming in Colour Space



Evolution operator in colour flow basis:

$$[\tau|e^{\mathbf{\Gamma}}|\sigma\rangle = \sum_{k=0}^{\infty} \frac{(-1)^{l}}{N^{l}} \sum_{k=0}^{l} \frac{(-\rho)^{k}}{k!} \sum_{\sigma_{0},...,\sigma_{l-k}} \delta_{\tau\sigma_{0}} \delta_{\sigma_{l-k}\sigma} \prod_{\alpha=0}^{l-k-1} \Sigma_{\sigma_{\alpha}\sigma_{\alpha+1}} R(\{\sigma_{0},...,\sigma_{l-k}\})$$

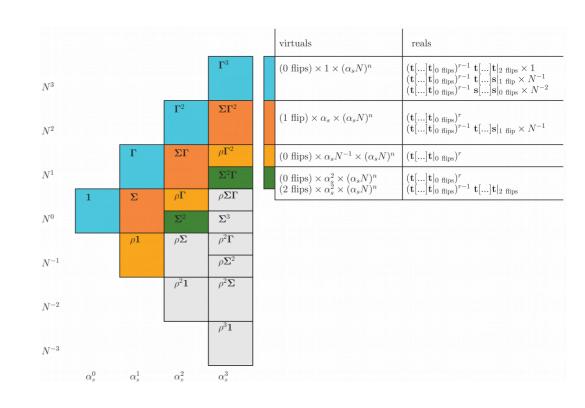
$$= \delta_{\tau\sigma} \left(e^{-N\Gamma_{\sigma}} + e^{-N\Gamma_{\sigma}} \frac{\rho}{N}\right) - \frac{1}{N} \frac{e^{-N\Gamma_{\tau}} - e^{-N\Gamma_{\sigma}}}{\Gamma_{\tau} - \Gamma_{\sigma}} \Sigma_{\tau\sigma} + \text{NNLC}$$

[Plätzer - EPJ C 74 (2014) 2907]

Sum terms enhanced by $\alpha_s N$ to all orders, insert perturbations in 1/N.

Take into account real emission contributions and the final suppression by the scalar product matrix element.

$$\operatorname{Tr}\left[\mathbf{A}\right] = \sum_{\tau,\sigma} [\tau |\mathbf{A}|\sigma] \langle \sigma | \tau \rangle \qquad \langle \sigma | \tau \rangle = N^{n-\#(\tau,\sigma)}$$



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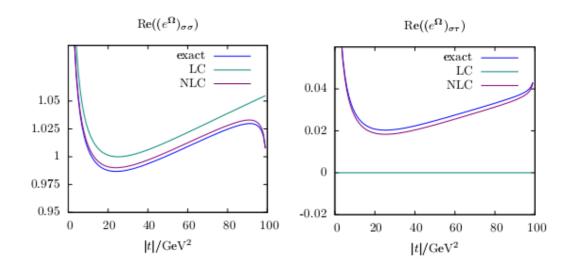
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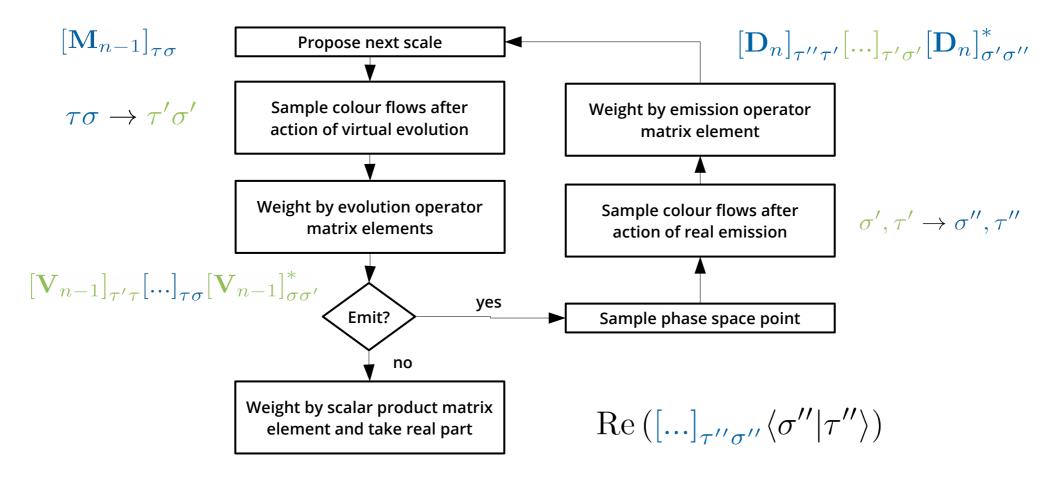


Basic algorithm



[De Angelis, Forshaw, Plätzer – arXiv:181y.xxxxx]

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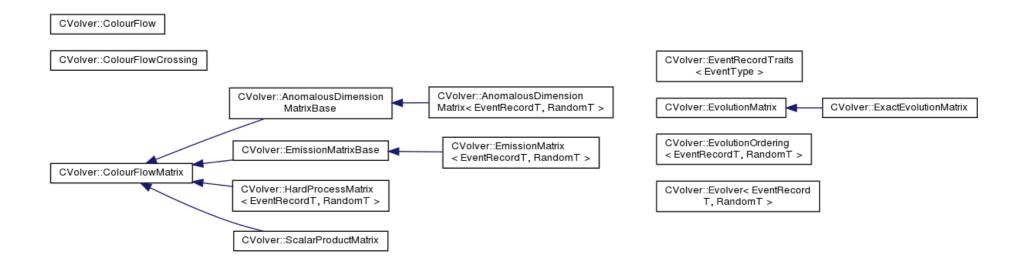
Exploring Colours with CVolver



[De Angelis, Forshaw, Plätzer – arXiv:181y.xxxxx]

A framework to solve multi-differential evolution equations in colour space. Concise, simple, and light-weight code structure.

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Dedicated Monte Carlo algorithms to sample colour structures.

Plugin approach can accommodate anything from (N)GLs to full parton showers.

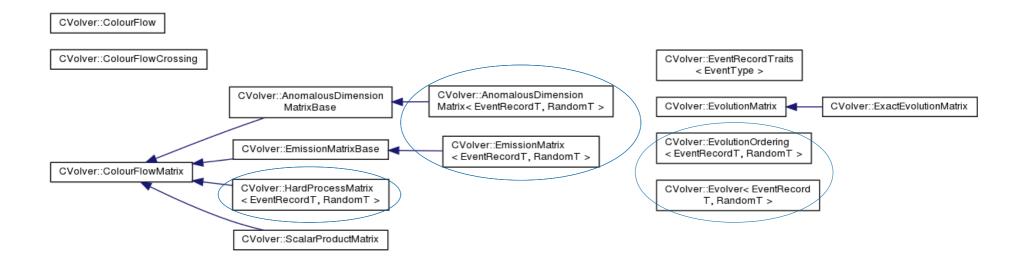
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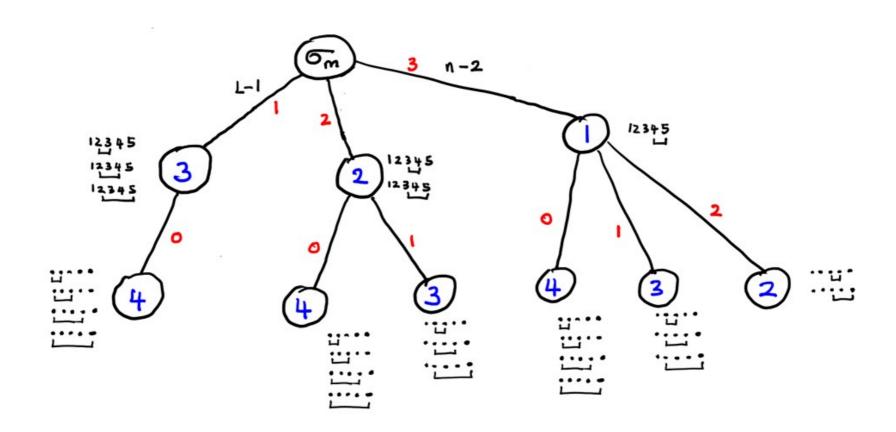
Importance Sampling in Colour Space



Importance sampling in colour space rules: $\#(au, au')\sim 1/N^{\#(au, au')}$

$$\#(\tau, \tau') \sim 1/N^{\#(\tau, \tau')}$$

Enumerate and address permutations with fixed cycle length:

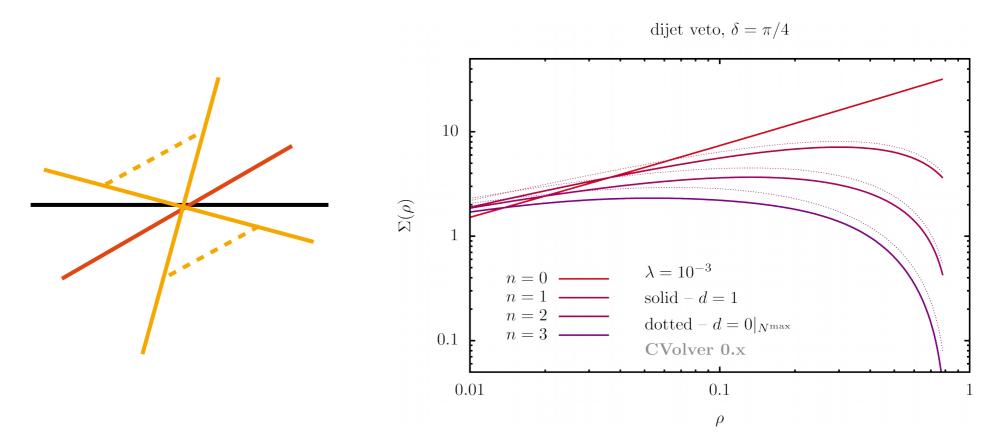


Numerical Results from CVolver



[De Angelis, Forshaw, Plätzer – arXiv:181y.xxxxx]

Code is differential for a large class of (non-global) observables. Example: Cone-dijet veto cross section.



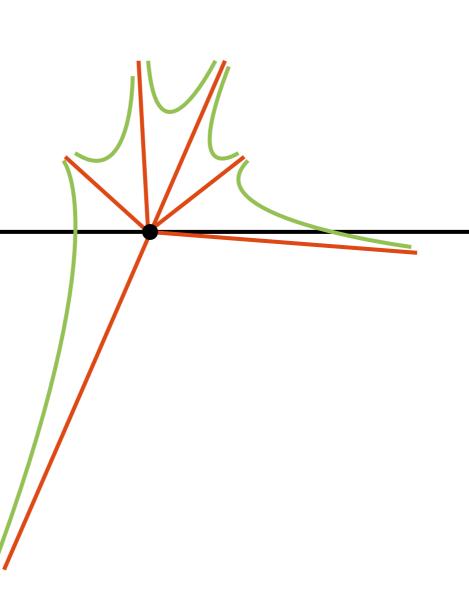
1/N breakdowns possible, scales up to several 10s of emissions for d=2.



The cluster model is based on a single colour flow after the shower has stopped.

Essentially no evolution is considered, just decays.

[Gieseke, Kirchgaesser, Plätzer, Siodmok – arXiv:1808.06770]

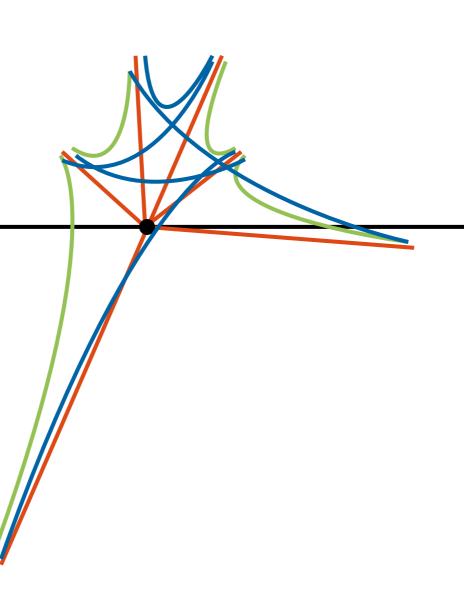




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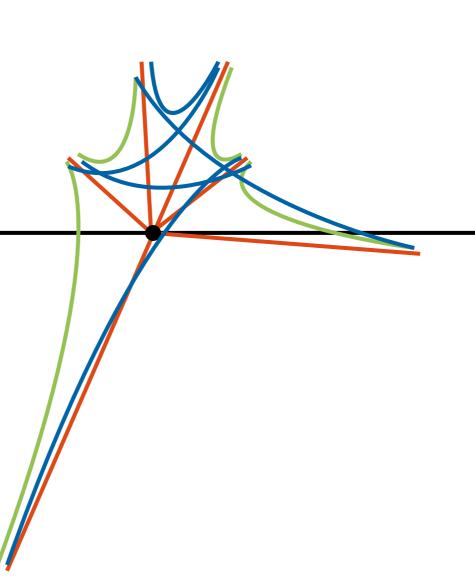
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View as an evolving amplitude, driven by a single initial colour flow:

$$|\mathcal{M}\rangle = e^{\mathbf{\Gamma}}|\text{clusters}\rangle$$

$$P_{\text{reco}} \sim |\langle \text{clusters'}|\mathcal{M}\rangle|^2$$

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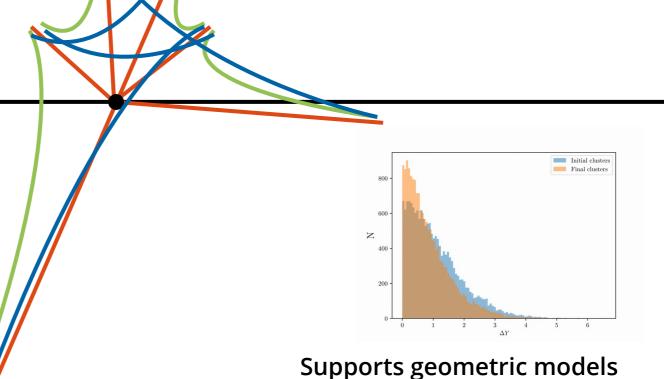
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[Gieseke, Kirchgaesser, Plätzer – EPJ C 78 (2018) 99]

Summary & Outlook



A framework for amplitude-level evolution at full colour, including a Monte Carlo method to produce actual `events'.

Serves to host a new approach to parton shower algorithms, which are systematically improvable, and first steps in that direction are underway.

Resummation of non-global observables requires the very same methods, and so provide an ideal playground to quantify these effects in a controlled way.

Details of colour expansion under study.

Beyond the perturbative evolution this provides us with unique insight into colour reconnection models.